

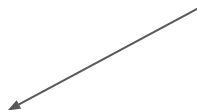
Software Tutorial

Streamlining PWA Workflow and
Regularization

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Regularization and PWA

$$\mathcal{I}(\theta, \xi)$$



Bin in mass describe only angular
dependence

“Mass independent fit”

α_m “Production coefficients” in mass bin
m

Intensity Distribution

θ = Angular dependence

ξ = (mass, ...) dependence

For a more detailed
writeup see

[GlueX Doc: 6826](#)

Regularization and PWA

α_m Parameter vector (real/imaginary parts of amplitudes)
Dimensionality $\sim (2 \times N_{\text{partial waves}})$

$P(\alpha_m)$ We wish to infer the probability distribution

X_m Observed data samples

$P(X_m|\alpha_m) = \prod_j P(x_{m,j}|\alpha_{m,j})$ Likelihood (product of probs) of observing data $\{j\}$

$\hat{\alpha}_m = \arg \min_{\alpha_m} [-\log P(X_m|\alpha_m)]$

Best guess by maximizing the likelihood (or minimizing the negative log-likelihood)

If we believe some free parameters are small then we can apply pressure on the model -> Regularization

Lasso Regularization (L1 norm of the parameter vector) penalizing non-zero values

$$\hat{\theta} = \arg \min_{\theta} \left[-\log P(X|\theta) + \lambda \sum_{i=1}^D |\theta_i| \right]$$

So far, these concepts live in the Frequentist framework for Statistics

Alternatively we can apply pressure by multiplying the likelihood by a prior distribution

Bayesian Framework

$$P(\theta|X) \propto P(X|\theta)P(\theta)$$

Posterior Distribution

Likelihood

Prior
Distribution


Normalization constant is generally difficult to evaluate as it is a high dimensional integral

$$Z = P(X) = \int P(X|\theta)P(\theta)d\theta$$

Markov Chain Monte Carlo performs monte carlo integration to approximate expectation values, i.e. marginalization

$$P(\theta_i|X) = \int P(\theta|X)d\theta_{-i} = \frac{1}{Z} \int P(\theta|X) \prod_{j \neq i} d\theta_j \approx \text{Histogram}\{\theta_i^{(s)}\}_{s=1}^S$$

Regularization and PWA

$$\mathcal{I}(\theta, \xi)$$
Two arrows originate from the parameters of the equation. One arrow points from θ to the text 'Bin in mass describe only angular dependence'. The other arrow points from ξ to the text 'Alternatively we can describe mass dependence'.

θ = Angular dependence

ξ = (mass, t, ...) dependence

Bin in mass describe only angular
dependence
“Mass independent fit”

Alternatively we can describe mass
dependence
“Mass dependent fit”

α_m “Production coefficients” in mass bin
m

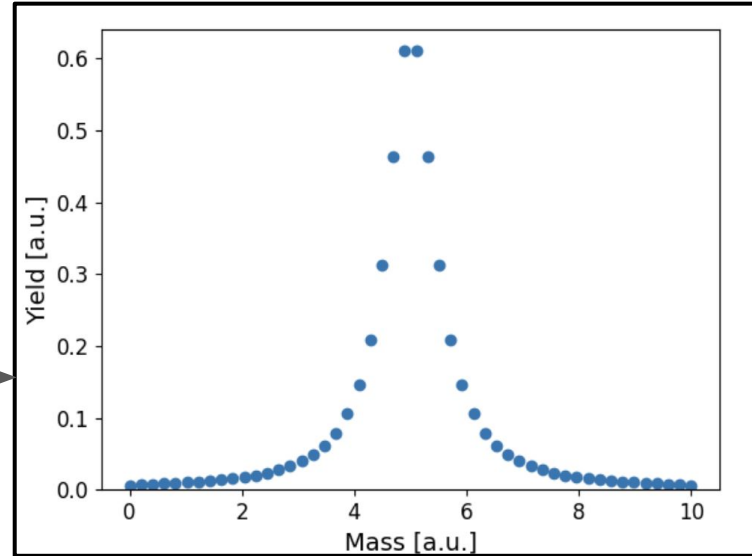
Regularization | Mass Dependent Fits

$$P(\alpha) = P(\alpha_0, \dots, \alpha_m)$$

“mass independent” fits across all mass bins
Dimensionality $\sim (M \text{ bins}) \times (2 \times N_{\text{partial waves}})$

If isolated resonance than Breit-Wigner
good description - needs only 2
parameters

$$\frac{1}{E - M_0 - i\Gamma_0/2}$$



More parameters needed if modeling background and approaches $(M \text{ bins}) \times (2 \times N_{\text{partial waves}})$ if using piecewise description

Regularization | Bayesian Mass Dependence

Posterior distribution of production coefficients across all mass bins

$$P(\alpha|X) \propto \prod_j P(x_j|\alpha)P(\alpha)$$

Physical Constraints

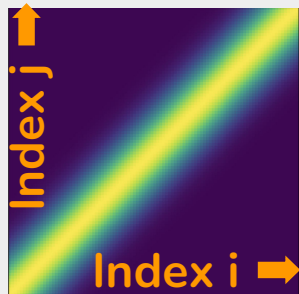
- Enforce smoothness by considering correlations between mass bins
- How might this be achieved?

Regularization | Information Field Theory

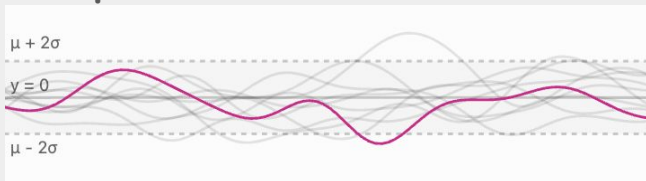
$k(\alpha, \alpha')$ Kernel function - correlation between production coefficients in a pair of kinematic bins

Radial Basis
Function Kernel

$$\sigma^2 \exp\left(-\frac{\|t-t'\|^2}{2l^2}\right)$$

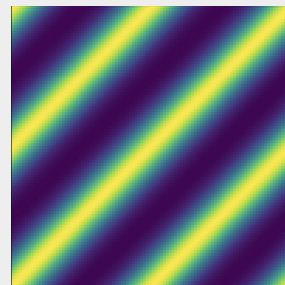


Samples drawn from Kernel

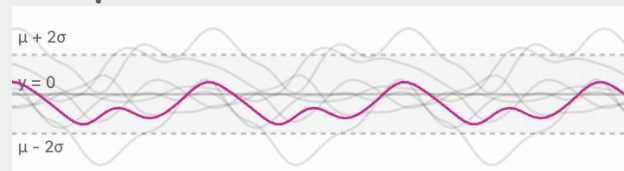


Periodic Kernel

$$\sigma^2 \exp\left(-\frac{2 \sin^2(\pi|t-t'|/p)}{l^2}\right)$$



Samples drawn from Kernel



[Distill.pub](https://distill.pub/): A Visual Exploration of Gaussian Processes

Regularization | Information Field Theory

If bin spacing is uniform (By Wiener–Khinchin theorem) it is sufficient to infer only the power spectral density

Example:

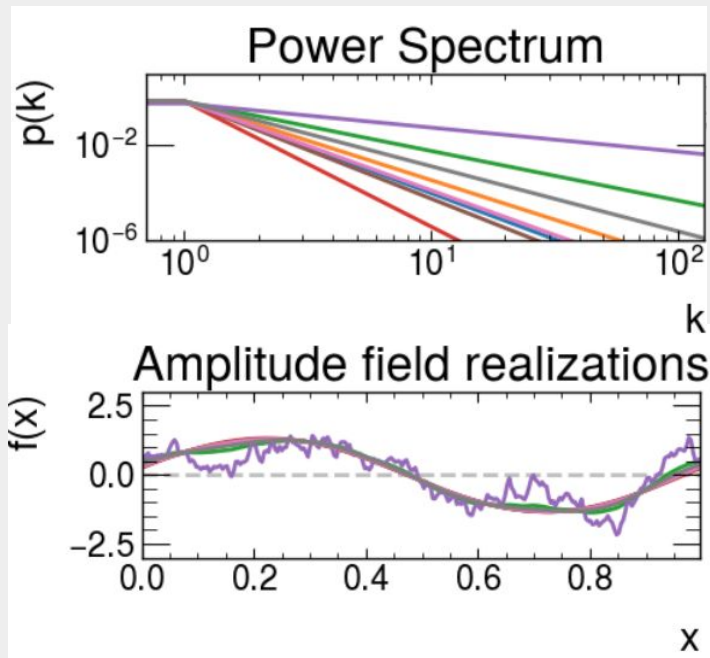
Power spectrum
slope parameter

~

LogNormal(-6, 3)

Draw 8 samples

[NIFTy Correlated Field Demo](#)



Regularization | Information Field Theory

$$P(\alpha|X) \propto \prod_j P(x_j|\alpha)P(\alpha)$$

$P(\alpha)$ Smoothness enforced by the prior distribution of power spectral density for each partial wave

For a more detailed writeup see
[GlueX Doc: 6826](#)

Motivation

- Partial wave analysis can be complicated (fitting, bookkeeping, result handling, ...)
- PyAmpTools sacrifices flexibility for ease of use and focuses on Python as the core language
- **Features:**
 - Scientific Python ecosystem is massive (achieve most of the benefits of low level languages without the hassle)
 - Optimization uses gradients provided by JAX automatic differentiation (scales better than numerical diff.)
 - Hooks into several optimization frameworks spanning MLE, MCMC, to Variational Inference
 - iminuit, scipy, numpyro, iftpwa ← **Information Field Theory framework**
 - YAML file is used to configure the analysis for consistency and automation
 - Fitting is generally done through the command line

Loading the container

On the JLab farm type this in

...

```
apptainer exec --contain --writable-tmpfs \  
  --bind /my/working/directory \  
  --bind /scratch \  
  --bind ~/.cache/fontconfig \  
  --env BASH_ENV=/dev/null \  
  /w/halld-scsshelf2101/lng/WORK/PyAmpTools/pyamptools.sif bash
```

...

```
source /etc/bash.bashrc
```

PyAmpTools| Command line

- Fitting is generally done through the command line

`bash$ pa -h`

Dispatch pyamptools commands. Select a command from the Commands section below. Remaining arguments will be passed to the selected command.

optional arguments:

`-h, --help` show this help message and exit
`-f, --files` show command file locations and exit

Commands:

<code>* nentries</code>	Print entries in List[ROOT files] (* wildcard), flag for branch integration
<code>* fit</code>	(AmpTools) Perform a set of MLE fits given an amptools config file
<code>* fitfrac</code>	(AmpTools) Extract fit fractions from a given amptools FitResults file
<code>* gen_amp</code>	(AmpTools) Generate data for a given configuration file
<code>* gen_vec_ps</code>	(AmpTools) Generate vector-pseudo scalar data for a given configuration file
<code>* ift_pkl_summary</code>	Summarize the contents of an IFT results pickle file
<code>* calc_ps</code>	[In Development] Calculate the phase space factor for IFT fits

==== YAML based commands below (takes a YAML file argument to configure setup) ====

<code>* from_default</code>	Copy default yaml to a given location
<code>* run_priorSim</code>	Draw sample from NIFTy prior, generate simulated data, and split into kinematic bins
<code>* run_cfgGen</code>	(AmpTools) Generate an AmpTools fit configuration file
<code>* run_divideData</code>	Divide data into kinematic bins (separate folders)
<code>* run_processEvents</code>	(AmpTools) Process bins: dump AmpVecs + NormInts to pkl files
<code>* run_fit</code>	(AmpTools) Run MLE fits over kinematic bins using AmpTools
<code>* run_mle</code>	Run MLE fits over kinematic bins using variety of optimizers (minuit, lbfgs, ...)
<code>* run_mcmc</code>	Run MCMC fits over kinematic bins using numpyro NUTS sampler
<code>* run_ift</code>	Run IFT fit over kinematic bins
<code>* run_momentInverter</code>	Run moment inverter
<code>* run_resultMan</code>	Run result manager commands, make all default plots
<code>* dash_ift_cmp</code>	Compare multiple IFT fits (intensity and phase) using dash package

YAML Commands of form
`pa run_ift main.yaml`

Generate/Format
Data

Fit the Data

Plot the data

YAML file contains all the knobs

```
defaults_location: null
base_directory: /w/halld-scsshelf2101/lng/WORK/PyAmpTools9/demos/RESULTS # base directory for
data_folder: ${base_directory}/DATA_SOURCES # folder for data sources that will be binned
n_processes: 4 # global number of processes to generally use
polarizations:
  "000": 1.0 # polarization magnitude in each orientation
waveset: Sp0+_Sp0-_Dp2+_Dp2- # underscore separated list of waves to use
phase_reference: Sp0+_Sp0- # reference wave in each reflectivity sector
reaction: Beam Proton Pi0 Eta # amptools reaction scheme
daughters: # Daughter masses
  Pi0: 0.135
  Eta: 0.548
min_mass: 1.04 # minimum mass to consider
max_mass: 1.72 # maximum mass to consider
n_mass_bins: 17 # number of mass bins to use
min_t: 0.0 # minimum t to consider
max_t: 1.0 # maximum t to consider
n_t_bins: 1 # number of t bins to use
acceptance_correct: true # whether to apply acceptance corrections
datareader: ROOTDataReader # data reader to use
coordinate_system: cartesian # ['cartesian', 'polar'], dont use polar
bins_per_group: 1 # create group dirs grouping bins allowing nifty to handle finer bins
merge_grouped_trees: true # remerge the trees in each group
constrain_grouped_production: false # if not remerging, we can choose to constrain amplitude
real_waves: "" # same form as waveset, define which waves are purely real
fixed_waves: "" # same form as waveset, define which waves are fixed
add_amp_factor: "" # Add an amplitude factor to every amplitude. For example, 0omegaDalitz 0.
append_to_decay: "" # append this string to the decay amplitude, i.e. 'omega3pi' can be appe
append_to_cfg: "" # append this string to the AmpTools configuration file
```

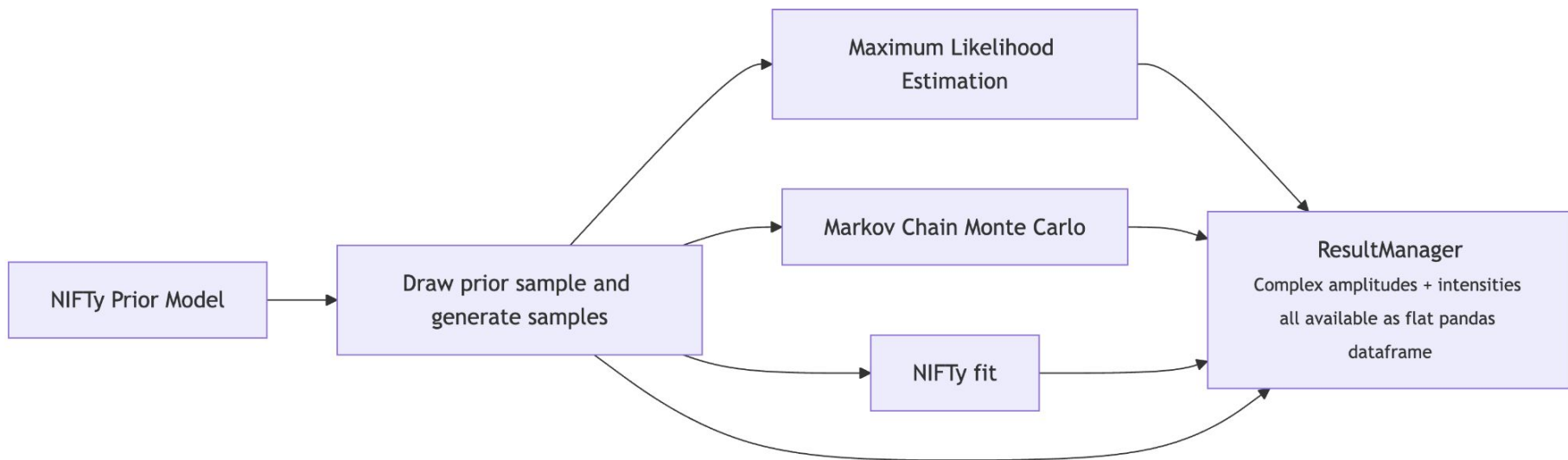
For vector pseudoscalar
just change waves to
another format:
1Sp1-
1Sp1+

Also see [Here](#)

Lots more knobs not shown but defaults can be pulled and most work decently well (of course you need dataset, waveset, binning, ...)

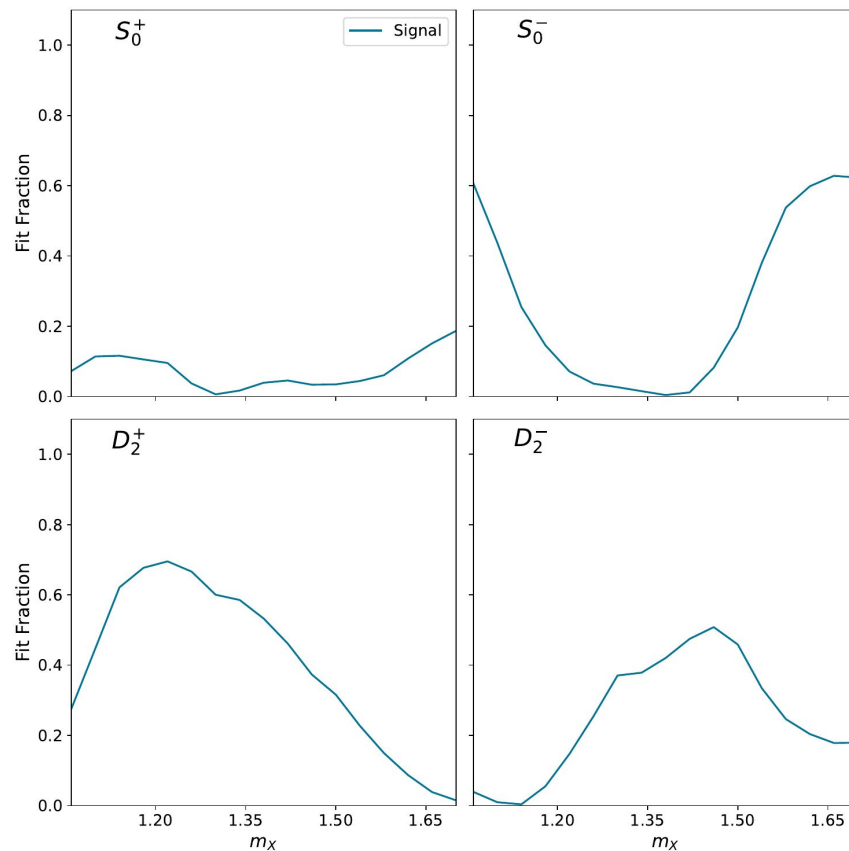
Tutorial workflow

Since IFT model is a Bayesian framework we can draw samples from the prior and generate MC simulations to rapidly iterate on Input/Output studies and compare optimization frameworks with this toolkit



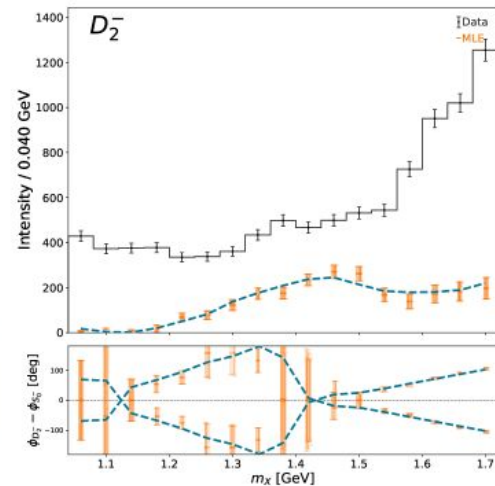
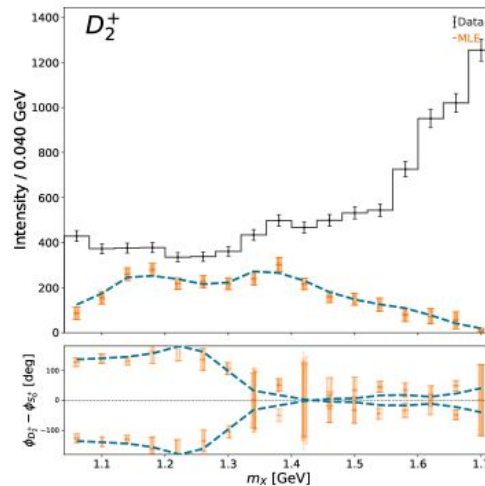
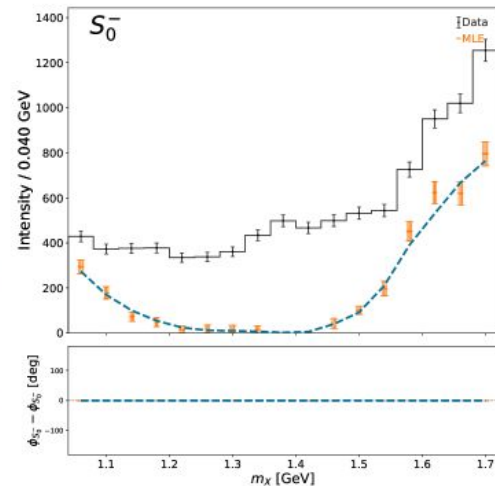
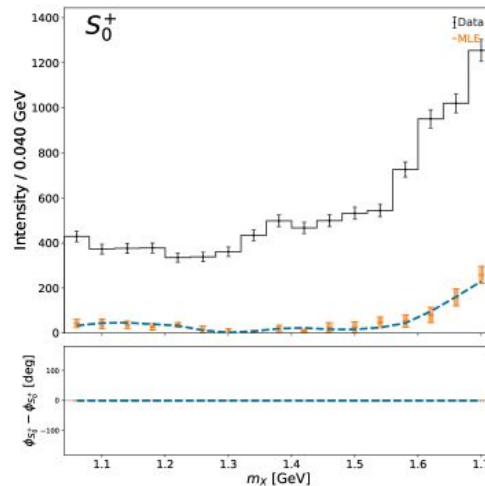
Draw prior sample and
generate samples

```
pa run_priorSim <yaml_file>
```



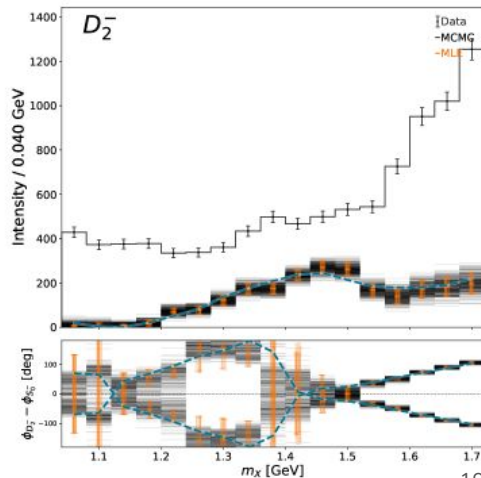
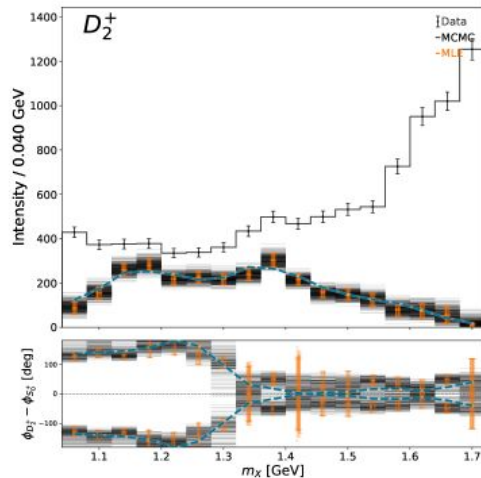
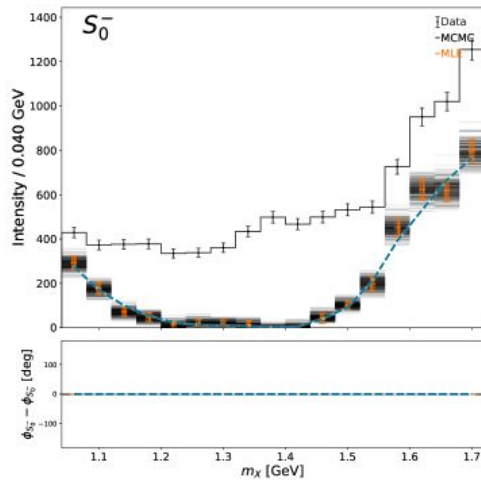
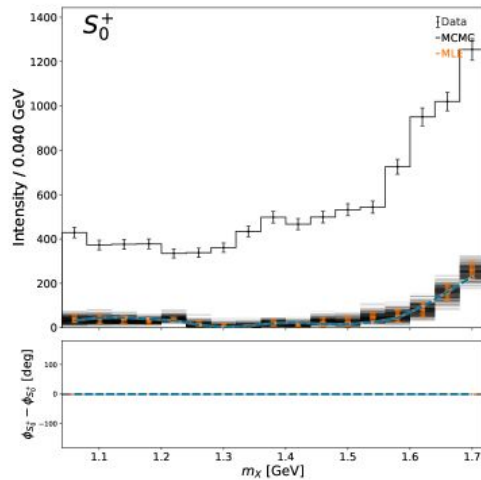
Maximum Likelihood Estimation

```
pa run_mle <yaml_file>
```



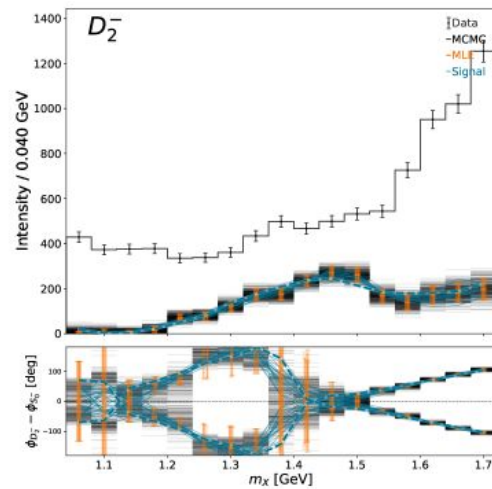
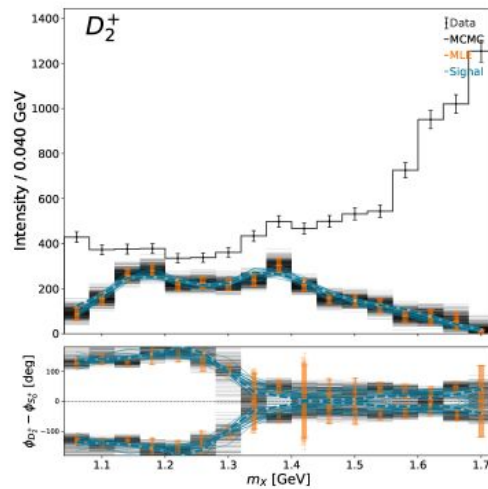
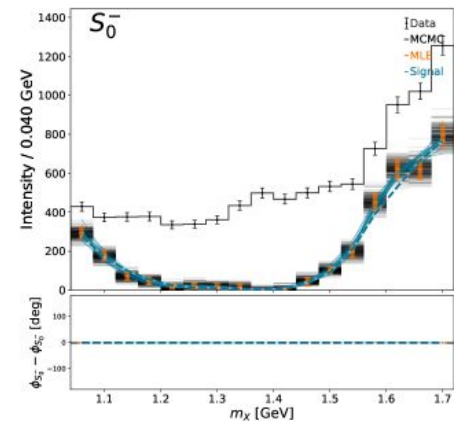
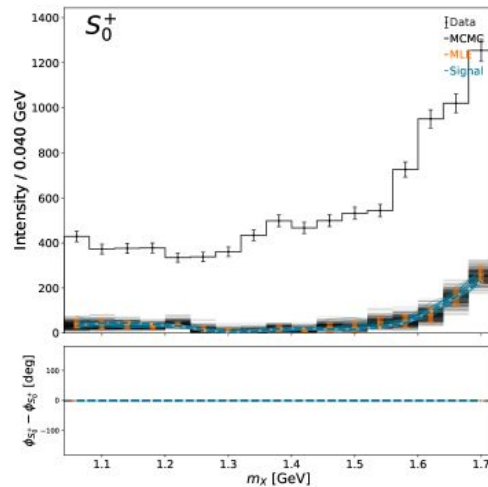
Markov Chain Monte Carlo

pa run_mcmc <yaml_file>



NIFTy fit

pa run_ift <yaml_file>



ResultManager

Complex amplitudes + intensities

all available as flat pandas

dataframe

```
from pyamptools.utility.resultManager import ResultManager
yaml_file = "/w/halld-scshef2101/lng/WORK/PyAmpTools9/0THE
resultManager = ResultManager(yaml_file)
resultManager.attempt_load_all()
```

All results loaded as a flat pandas dataframe (intensities, complex amplitudes, moments , etc...)

```
resultManager.mle_results.head(1)
```

✓ 0.1s

Python

	mass	initial_likelihood	likelihood	sample	Sp0+_amp	Sp0-_amp	Dp2+_amp	Dp2-_amp	intensity	intensity_err	...	H2(1,1)	H2(2,1)	H2(2,2)	H2(3,1)	H2(3,2)	H2(3,3)	H2(4,1)
0	1.06	658.275784	361.181041	8	20.106399+0.000000j	59.770623+0.000000j	-25.079330+21.531924j	3.806061-16.484035j	424.992813	29.15439	...	0.0	0.0	0.061104	0.0	0.0	0.0	0.0

1 rows x 79 columns

+ Code

+ Markdown

```
resultManager.mcmc_results.head(1)
```

✓ 0.1s

Python

	intensity	Sp0+	Sp0-	Dp2+	Dp2-	Sp0+_amp	Sp0-_amp	Dp2+_amp	Dp2-_amp	mass	...	H2(1,1)	H2(2,1)	H2(2,2)	H2(3,1)	H2(3,2)	H2(3,3)	H2(4,1)
0	447.726497	9.040142	299.052373	116.342458	23.328655	10.63292+0.00000j	61.450239+0.000000j	-37.406014-8.087889j	3.974760-16.600252j	1.06	...	0.0	0.0	0.050858	0.0	0.0	0.0	0.0

1 rows x 54 columns

```
resultManager.ift_results[0].head(1)
```

✓ 0.1s

Python

	tprime	mass	sample	Sp0+_amp	Sp0-_amp	Dp2+_amp	Dp2-_amp	Sp0+	Sp0-	Dp2+	...	H2(1,1)	H2(2,1)	H2(2,2)	H2(3,1)	H2(3,2)	H2(3,3)	H2(4,1)
0	0.5	1.06	0	18.905427-0.000000j	60.91196-0.00000j	-32.059605-13.029281j	1.409181+22.108157j	28.578744	293.836164	95.129854	...	0.0	0.0	0.05376	0.0	0.0	0.0	0.0

1 rows x 55 columns

Have time? Click this
[Tutorial Link](#)