

Radiative Corrections -The SIMC Way

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Introduction

Electrons radiate in the presence of a nuclear field due to changes in their velocity brought about by coulomb interaction and radiation resulting from such deceleration of the electron is called Bremsstrahlung. The incoming and outgoing electrons can interact with the coulomb field of the nucleus involved in the scattering process, this results in emission and reabsorption of virtual photons and emission of real soft photons. Such processes are known as internal Bremsstrahlung. The electrons can also interact with the coulomb field of a nucleus other than the one involved in the scattering process and thereby radiate photons. These radiations are known as external Bremsstrahlung. In an experiment involving electrons scattering off some target, these radiative processes have a twofold effect on the data. Firstly the cross section of the process is modified and secondly the kinematics (energy, momentum, angle) of the electron are also changed. Although, these are real physical processes, they are experiment specific, and so most theoretical calculations do not take these effects into account. Thus, in order to get to the underlying physics, and also to directly compare with theoretical calculations, one needs to unfold these radiative processes from the data. The procedure for doing such radiative corrections was first derived by J. Schwinger [1], which were later modified by Mo and Tsai [3]. Radiative correction formulas for coincidence $(e, e'p)$ reactions were calculated from the Mo and Tsai formulation by Makins et. al. [2]. The derivation of these formulae is very well described in references [2] and [4], so I will not venture to write down the full derivation at this time. The goal of this paper is to elaborate on how the above mentioned formulae are used in the PWIA Monte Carlo SIMC.

Radiative Correction for Coincidence $(e, e'p)$ Reactions

The cross-section for radiating energy E_e along the incoming electron direction \hat{k} , $E_{e'}$ along the scattered electron direction \hat{k}' , E_p along the direction of the scattered proton \hat{p} and also radiating any number of soft photons with energy less than ΔE , calculated to all orders, is given by :

$$\frac{d\sigma}{d\Omega_e dE_e dE_{e'} dE_{p'}} = \frac{d\sigma}{d\Omega_e}|_{ep}(1 - \delta_{hard}) \frac{\lambda_e \lambda_{e'} \lambda_{p'}}{(\sqrt{k k'})^{\lambda_e} (\sqrt{k k'})^{\lambda_{e'}} (\sqrt{M p^{o'}})^{\lambda_{p'}}} \times \frac{1}{E_e^{1-\lambda_e} E_{e'}^{1-\lambda_{e'}} E_{p'}^{1-\lambda_{p'}}} \quad (1)$$

Here the total energy radiated is $\omega_{total} = E_e + E_{e'} + E_{p'}$, and the λ s are the angular distribution functions of the photons radiated in the three directions. The Angular distribution of the radiation is approximated as :

$$A_{peaking}(\hat{\omega}) = \lambda_e \delta(\hat{\omega} - \hat{k}) + \lambda_{e'} \delta(\hat{\omega} - \hat{k}') + \lambda_{p'} \delta(\hat{\omega} - \hat{p}') \quad (2)$$

This simple approach to the angular distribution is also known as the 'extended peaking approximation'. The λ s are given by the following expressions:

$$\lambda_e = \frac{\alpha}{\pi} [\ln(\frac{4k^2}{m^2} - 1)] + \frac{\alpha}{\pi} [2\ln(\frac{k}{k'} + \ln(\frac{1 - \cos\theta_e}{2})] \quad (3)$$

$$\lambda_{e'} = \frac{\alpha}{\pi} [\ln(\frac{4k'^2}{m^2} - 1)] + \frac{\alpha}{\pi} [2\ln(\frac{k}{k'} + \ln(\frac{1 - \cos\theta_{e'}}{2})] \quad (4)$$

$$\lambda_{p'} = \frac{\alpha}{\pi} [\ln(\frac{p'^o + |p'|}{p'^o - |p'|}) - 2] \quad (5)$$

k, k' and p' are the magnitude of the incident electron momentum, scattered electron momentum and the scattered proton momentum respectively and p'^o is the proton energy. δ_{hard} is the contribution from the second order virtual photon radiation to the vertex corrections, and is given by:

$$\delta_{hard} = 2\alpha [-\frac{3}{4\pi} \ln(-q^2/m^2) + \frac{1}{\pi} - \sum_i \delta_i^{vp}(q^2)], \quad (6)$$

where \sum_i sums over the different flavors of leptons with mass m_i and,

$$\delta_i^{vp} = \frac{1}{3\pi} [-\frac{5}{3} + \ln(-q^2/m_i^2)] \quad (7)$$

All of the above expressions are for what is known as internal Bremsstrahlung, however photons are also emitted when the electrons are in the field of nuclei other than those involved in the hard scattering process. This is known as external Bremsstrahlung. The proton being massive emits negligible amounts of external radiation. Both E_i^{int} and E_i^{ext} are emitted in the same direction, thus if the internal and external Bremsstrahlung are added together then we can write the cross-section in terms of E_i and E_f radiated along k and k' as :

$$\frac{d\sigma}{d\Omega_e dE_i^{int} dE_i^{ext} dE_f^{int} dE_f^{ext} dE_{p'}} = \frac{d\sigma}{d\Omega_e}|_{ep}(1 - \delta_{hard}) \frac{1}{\Gamma(1 + bt_i)} \frac{1}{\Gamma(1 + bt_f)} \\ \times \frac{(bt_i + \lambda_i)}{k^{bt_i}(\sqrt{k k'})^{\lambda_i}} \frac{(bt_f + \lambda_f)}{k'^{bt_f}(\sqrt{k k'})^{\lambda_f}} \frac{dE_i}{E_i^{1-\lambda_i-bt_i}} \frac{dE_f}{E_f^{1-\lambda_f-bt_f}} \Phi_i^{ext}(E_i) \Phi_f^{ext}(E_f) \quad (8)$$

Where the function Φ^{ext} is a correction for external radiation which have large photon energies and has the form:

$$\Phi_i^{ext}(E_i) = 1 - \frac{bt_i}{bt_i + \lambda_i} \frac{E_i}{k_i} \quad (9)$$

and

$$b = 1/9(12 + \frac{Z + 1}{ZL_1 + L_2}) \quad (10)$$

$$L_1 = \ln(184.15) - \frac{1}{3}\ln(Z) \quad (11)$$

$$L_2 = \ln(1194.) - \frac{2}{3}\ln(Z) \quad (12)$$

The Radiative Correction Procedure in The PWIA Monte Carlo SIMC

In the Monte Carlo SIMC the photon energies $E_e, E_{e'}$ and $E_{p'}$ along the incoming electron, the outgoing electron and the proton directions respectively, are generated separately. The total energy radiated is the vector sum as given by equation 2. This comes about because in equation 1 above we notice that the energy and angular distribution of radiation in the three directions factorize into three independent functions.

The shape of each of these distributions has the form:

$$\frac{1}{\Gamma(1 + bt)} \frac{bt + \lambda}{k^{bt}(\sqrt{k k'})^\lambda} \frac{dE}{E^{1-\lambda-bt}} \quad (13)$$

If we rename $bt + \lambda$ as g and $\frac{bt+\lambda}{\Gamma(1+bt)} \frac{1}{k^{bt}(\sqrt{k k'})^\lambda}$ as C we get the simple form:

$$C * E^{g-1} dE \quad (14)$$

Hence we can use this simple form to generate the energy radiated in a given direction between limits E_{max} and E_{min} . The generating function must normalize to 1 between these limits so we have

$$N * \int_{E_{min}}^{E_{max}} (E^{g-1}) dE = 1 \quad (15)$$

or

$$N = \frac{g}{E_{max}^g - E_{min}^g} \quad (16)$$

Thus for each of the radiation tails, the energy radiated in that particular direction is randomly generated in the range $E_{max} - E_{min}$ using the generating function or energy shape:

$$G = \frac{gE^{g-1}}{E_{max}^g - E_{min}^g} \quad (17)$$

The limits E_{max} and E_{min} are determined from the limits of the model spectral function, the limits on the energy and momenta of the incident and scattered particles determined from the spectrometer acceptance and the randomly generated energy and momenta of the incident and scattered particles.

Once the energy radiated in each of the tails is known, the next step is to use these energies to modify the momentum and energy of the incident and scattered particles involved in the reaction. This is done for each event by subtracting off the radiated energy from the randomly generated vertex energies (energy of the particles at the reaction vertex).

Next the radiation weight is calculated for each event which is then assigned to that event. The radiation weight is the probability of radiating soft or hard photons of a given energy. The radiation weight had three components. The first component is the probability of emitting a photon which has the correct radiative tail shape, this comes from equation 13 and the generating function (equation 17). From these equations we get the weight for each of the three tails as:

$$W_{rad}^i = C/g * ((E_{max}^i)^g - (E_{min}^i)^g); i = e, e', p' \quad (18)$$

The product of the three weights for the three tails give us W_{rad}^{soft} .

$$W_{rad}^{soft} = W_{rad}^e W_{rad}^{e'} W_{rad}^{p'} \quad (19)$$

The second component is the multiplicative correction factor due to external radiation Φ_i^{ext} , as described in equation 9. This too is calculated for each tail. The final component is the due to the vertex corrections and is given by $(1 - \delta_{hard})$. So finally the product of all these little pieces gives us the radiation weight for an event.

$$W_{rad}^{event} = W_{rad}^{soft} \Phi_e^{ext} \Phi_{e'}^{ext} \Phi_{p'}^{ext} (1 - \delta_{hard}) \quad (20)$$

When the data is binned in terms of E_m and P_m , we have accounted for events which radiated into a particular bin by modifying the vertex. The E_m and P_m was changed by exactly the total radiated energy, hence they contribute to bins they radiated into and not the ones they would have if

there was no radiation. In addition the radiation weight assigned to each event accounts for events which radiated out of the bin. These two features together constitute the radiative correction procedure of the Monte Carlo SIMC. Thus using the described procedure we can generate radiated spectra with SIMC. This method obtains correct multi-photon angular distributions and hence is also known as the 'multi-photon' technique. However, one must remember that it does involve the peaking approximation at the single photon level.

As an illustration of how the procedure works, Fig 1 compares an unradiated missing energy vs missing momentum spectrum with a radiated spectrum for Hydrogen target. One can clearly identify the radiative tails in this figure. In Fig 2 the hydrogen missing energy spectrum is compared with the same calculated using the Monte Carlo.

Conclusions

The procedure described above is one of the three radiative correction procedures which are available in SIMC. The procedure described is the one currently in use. Various comparisons of the different procedures were done by the authors and they found agreement to be better than 1% within the different procedures. Since the above procedure approximates the angular distribution most effectively, it is the procedure of choice.

REFERENCES

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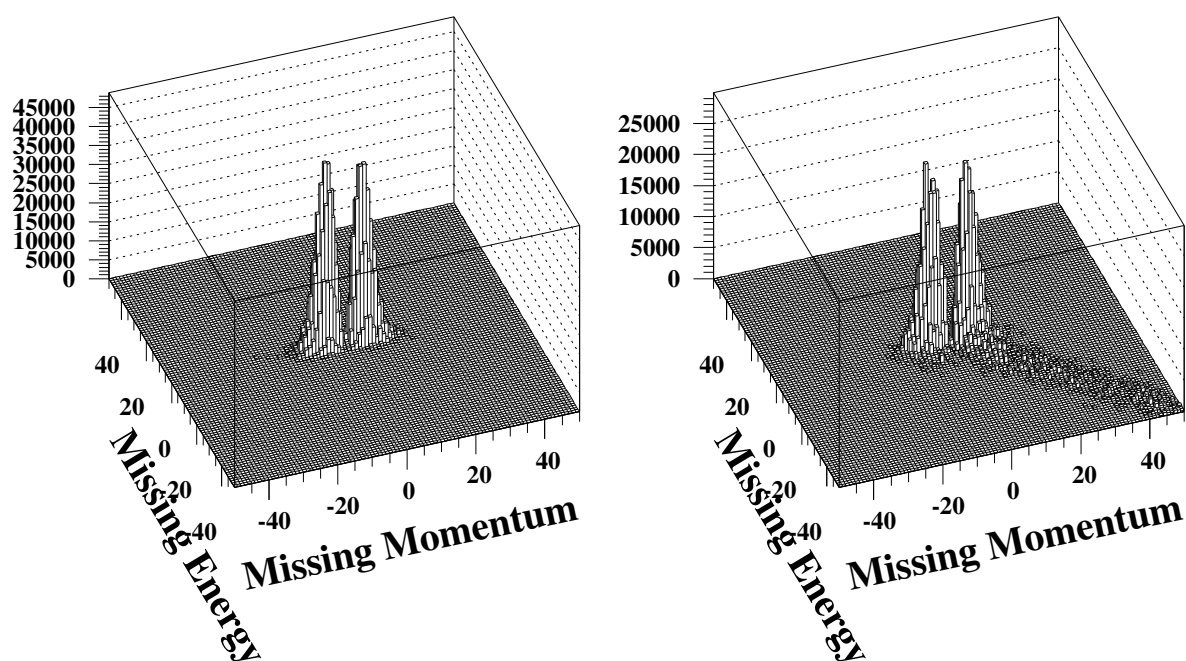


FIGURE 1. Plot of missing energy vs missing momentum for hydrogen with and without radiation.

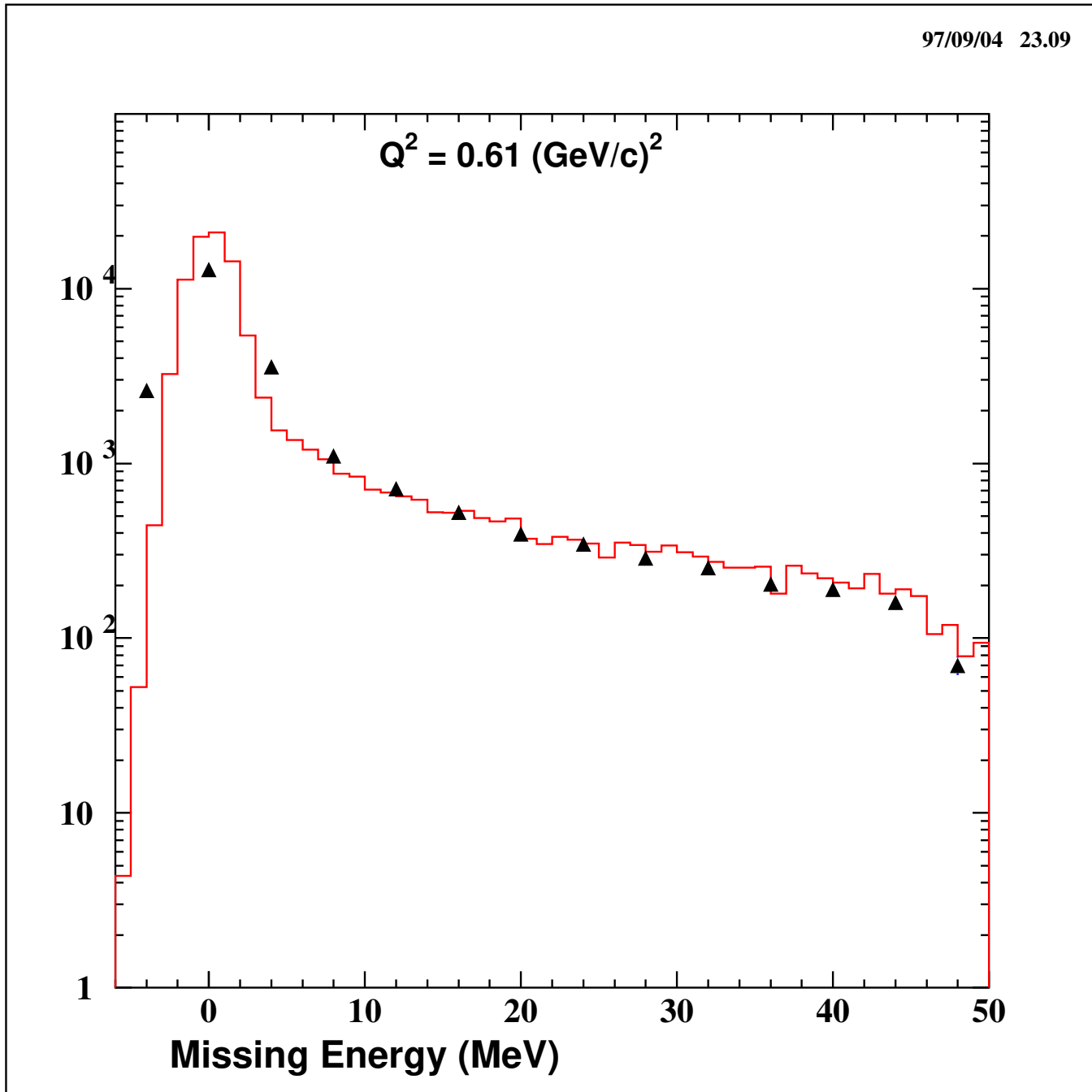


FIGURE 2. Plot of the missing energy spectrum of hydrogen, solid line is the calculation using Monte Carlo SIMC and triangles are data, with statistical errors only.