HW Assignment 1

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(1) Curtis (1.6): An 80 kg man and a 50 kg woman stand 0.5 m from each other. What is the force of gravitational attraction between the couple?

$$F = rac{Gm_1m_2}{r_{12}^2} egin{displayspicture} G = 6.67259 imes 10^{-11} rac{ ext{m}^3}{ ext{kg - sec}^2} \ m_1 = 80 ext{ kg} \ m_2 = 50 ext{ kg} \ r_{12} = 0.5 ext{ m} \ \end{bmatrix} \ = rac{6.67259 imes 10^{-11})(80)(50)}{(0.5)} \ = \boxed{1.068 \ \mu N}$$

(2) Curtis (1.8): If a person's weight is W on the surface of the earth, calculate what it would be, in terms of W, at the surface of a) The moon; b) Mars; c) Jupiter.

On surface of Earth:

$$F=W=rac{Gm_1m_E}{r_E^2}=m_1\left(rac{Gm_E}{r_E^2}
ight)$$

Knowing that $G=6.67259 imes10^{-11}$ ${
m m^3\over kg$ - ${
m sec}^2}$, $m_E=5.9722 imes10^{24}$ kg, and $r_E=6.3781 imes10^6$ ${
m m}$:

$$egin{align} W &= m_1 \left(rac{(6.67259 imes 10^{-11})(5.9722 imes 10^{24})}{(6.3781 imes 10^6)^2}
ight) = m_1 (\ rac{9.796 \ rac{ ext{m}}{ ext{s}^2}}{9.796 \ rac{ ext{m}}{ ext{s}^2}}) \ \end{array}$$

a) On the moon;

$$g_{moon} = 1.62 \ rac{\mathrm{m}}{\mathrm{s}^2}$$

$$egin{align} \Rightarrow W_{moon} = m_1 g_{moon} = \left(rac{W}{9.796 rac{ ext{m}}{ ext{s}^2}}
ight) \left(1.62 rac{ ext{m}}{ ext{s}^2}
ight) \ = \boxed{0.165W} \end{aligned}$$

b) On Mars;

$$g_{Mars} = 3.71 \ rac{\mathrm{m}}{\mathrm{s}^2}$$

$$egin{align} \Rightarrow W_{Mars} = m_1 g_{Mars} = \left(rac{W}{9.796 rac{ ext{m}}{ ext{s}^2}}
ight) \left(3.71 rac{ ext{m}}{ ext{s}^2}
ight) \ = \boxed{0.379W} \end{aligned}$$

c) On Jupiter;

$$g_{Jupiter} = 24.79 \; rac{\mathrm{m}}{\mathrm{s}^2}$$

$$egin{aligned} &\Rightarrow W_{Jupiter} = m_1 g_{Jupiter} = \left(rac{W}{9.796 rac{ ext{m}}{ ext{s}^2}}
ight) \left(24.79 rac{ ext{m}}{ ext{s}^2}
ight) \ &= \boxed{2.53W} \end{aligned}$$

(3) Prove that equations (3.3) and (3.4) given in the class notes are equivalent to (3.2).

$$m_i rac{dar{r}_i}{dt^2} = G m_i \sum_{j=1}^{N} rac{m_j}{r_{ij}} (ar{r}_j - ar{r}_i) \qquad (3.2)$$

$$m_i \frac{d\bar{r}_i}{dt^2} = \nabla_i U \tag{3.3}$$

$$U = \frac{G}{2} \sum_{i=1}^{N} \sum_{j=1}^{N} \frac{m_i m_j}{r_{ij}}$$
 (3.4)

From (3.4), we can get rid of the i summation because we only want the force of attraction on mass m_i , and not on every possible pair. To prove this - at some point in expanding the double summation, we will get similar terms that add up. e.g.

$$\left[rac{m_1m_2}{r_{12}} = rac{m_2m_1}{r_{21}} = 2\left[rac{m_1m_2}{r_{12}}
ight]$$

Since this will be true for each pair, and by keeping i = 1, we get

$$egin{aligned} U &= rac{G}{2} \sum_{j=1}^{N}^{'} 2 \left[rac{m_i m_j}{r_{ij}}
ight] \ &= G m_i \sum_{j=1}^{N}^{'} rac{m_j}{r_{ij}} \end{aligned}$$

Now we can plug the force function U into (3.3)

$$m_i rac{dar{r}_i}{dt^2} =
abla_i \left(Gm_i \sum_{j=1}^{N} rac{m_j}{r_{ij}}
ight)$$

To evaluate the gradient, we must take the partial of U with each direction

$$abla_i U = rac{\partial U}{\partial x_i} \hat{i} \ + rac{\partial U}{\partial y_i} \hat{i} \ + rac{\partial U}{\partial z_i} \hat{i}$$

Let us evaluate partial x first. Since G, m_i , and m_j are independent of direction, we can group them outside of the gradient

$$egin{align} rac{\partial U}{\partial x_i}\hat{i} &= rac{\partial}{\partial x_i}igg(Gm_i\sum_{j=1}^{N}^{'}rac{m_j}{r_{ij}}igg)\hat{i} \ &= Gm_i\sum_{j=1}^{N}^{'}m_jrac{\partial}{\partial x_i}igg(rac{1}{r_{ij}}igg)\hat{i} \ \end{cases}$$

We know that

$$egin{aligned} r_{ij} &= \mid ar{r}_j - ar{r}_i \mid \ &= ((ar{r}_j - ar{r}_i)^2)^{1/2} \ &= ((x_j - x_i)^2 + (y_j - y_i)^2 + (z_j - z_i)^2)^{1/2} \end{aligned}$$

So that partial x is equal to

$$egin{aligned} &=rac{\partial}{\partial x_i} \Big(((x_j-x_i)^2+(y_j-y_i)^2+(z_j-z_i)^2)^{-1/2} \Big) \hat{i} \ &= \left[-rac{1}{2} ((x_j-x_i)^2+(y_j-y_i)^2+(z_j-z_i)^2)^{-3/2}
ight] [2(x_j-x_i)] [-1] \hat{i} \ &= rac{(x_j-x_i)}{((x_j-x_i)^2+(y_j-y_i)^2+(z_j-z_i)^2)^{3/2}} \ &= rac{(x_j-x_i)}{r_{ij}^3} \hat{i} \end{aligned}$$

Similarly, for partial y and partial z

$$egin{align} rac{\partial}{\partial y_i}igg(rac{1}{r_{ij}}igg)\hat{j} &=rac{(y_j-y_i)}{r_{ij}^3}\hat{j} \ rac{\partial}{\partial z_i}igg(rac{1}{r_{ij}}igg)\hat{k} &=rac{(z_j-z_i)}{r_{ij}^3}\hat{k} \ \end{aligned}$$

We can take out the r_{ij}^3 term since it is the same in all 3 components of the gradient. We are left with

$$Gm_i \sum_{j=1}^{N} rac{m_j}{r_{ij}^3} \underbrace{\left((x_j - x_i)\hat{i} \ + (y_j - y_i)\hat{j} \ + (z_j - z_i)\hat{k}
ight)}_{(ar{r}_j - ar{r}_i)} = \left[Gm_i \sum_{j=1}^{N} rac{m_j}{r_{ij}^3} (ar{r}_j - ar{r}_i)
ight]$$

which is identical to the right hand side of equation (3.2)

(4) Prove that the force function U given by equation (3.4) in the class notes is equal to the total work done by the gravitational forces in assembling a system of N point masses from a state of infinite dispersion to a given configuration.

$$U = rac{G}{2} \sum_{i=1}^{N} \sum_{j=1}^{N} rac{m_i m_j}{r_{ij}} \qquad (3.4)$$

Total work is given by

$$-\int F*dr$$

Force due to gravitational attraction on mass m_i

$$F_i = rac{Gm_im_j}{r_{ij}^2}$$

Since we want to bring point masses from infinity to a given config, we do

$$egin{align} -\int_{\infty}^{r_{ij}} F \, dr &= -\int_{\infty}^{r_{ij}} rac{Gm_i m_j}{r^2} \, dr \ &= -Gm_i m_j \int_{\infty}^{r_{ij}} r^{-2} \, dr \ &= -Gm_i m_j iggl[-rac{1}{r} iggr]_{\infty}^{r_{ij}} \ &= -Gm_i m_j \left[iggl(-rac{1}{r_{ij}} iggr) + iggl(-rac{1}{\infty} iggr) iggr] \ &= rac{Gm_i m_j}{r_{ij}} \ \end{aligned}$$

This is only for 2 masses. If we want N masses, we include summations. Similar to problem (3), we will have a summation iterating the i^{th} mass and another iterating the j^{th} mass (where $j \neq i$). As previously stated, this type of notation will lead to accounting each pair of masses twice, hence dividing the entire equation by 2

$$\Rightarrow oggl| rac{G}{2} \sum_{i=1}^N \sum_{j=1}^{N^{'}} rac{m_i m_j}{r_{ij}} iggr|$$

which is identical to equation (3.4)

(5) Prove equation (4.7) given in the class notes.

$$rac{ar{d}_j}{d_j^3} + rac{ar{
ho}_j}{
ho_j^3} = -
abla \left(rac{1}{d_j} - rac{1}{
ho_j^3}ar{r}\cdotar{
ho}_j
ight) \qquad (4.7)$$

We know that

$$ar{r}=ar{r}_2-ar{r}_1$$

If we set m_1 to 0 in all directions (origin), then $ar{r}=0$

$$egin{aligned} ar{r} &= ar{r}_2 - (0) = ar{r}_2 \ ar{
ho}_j &= ar{r}_j - (0) = ar{r}_j \ ar{d}_j &= ar{r}_2 - ar{r}_j \end{aligned}$$

We can separate this gradient into

$$-
abla\left(rac{1}{d_{j}}
ight)-
abla\left(-rac{1}{
ho_{j}^{3}}ar{r}\cdotar{
ho}_{j}
ight)$$

The left part will be

$$egin{align} &= -
abla \left(((x_2-x_j)^2 + (y_2-y_j)^2 + (z_2-z_j)^2)^{-1/2}
ight) \ &= rac{1}{d_j^3} ((x_2-x_j)\hat{i} + (y_2-y_j)\hat{j} + (z_2-z_j)\hat{k}) \ \hline rac{ar{d}_j}{d_j^3} \end{array}$$

For the right side, we have

$$ar{r}\cdotar{
ho}_j=(x\cdot x_j\ \hat{i})+(y\cdot y_j\ \hat{j})+(z\cdot z_j\ \hat{k})$$

So that

$$egin{split} &= -
abla \left(-rac{(x \cdot x_j \ \hat{i}) + (y \cdot y_j \ \hat{j}) + (z \cdot z_j \ \hat{k})}{\left((x_j^2 + y_j^2 + z_j^2)^{1/2}
ight)}
ight) \ &= rac{1}{
ho_j^3} (x_j \hat{i} + y_j \hat{j} + z_j \hat{k}) \ &= \left[rac{ar{
ho}_j}{
ho_j^3}
ight] \end{split}$$

The steps in taking the derivative were skipped due to similarity of the derivate taken in question (3). Combining the left and right part we get

$$egin{aligned} rac{ar{d}_j}{d_j^3} + rac{ar{
ho}_j}{
ho_j^3} \end{aligned}$$