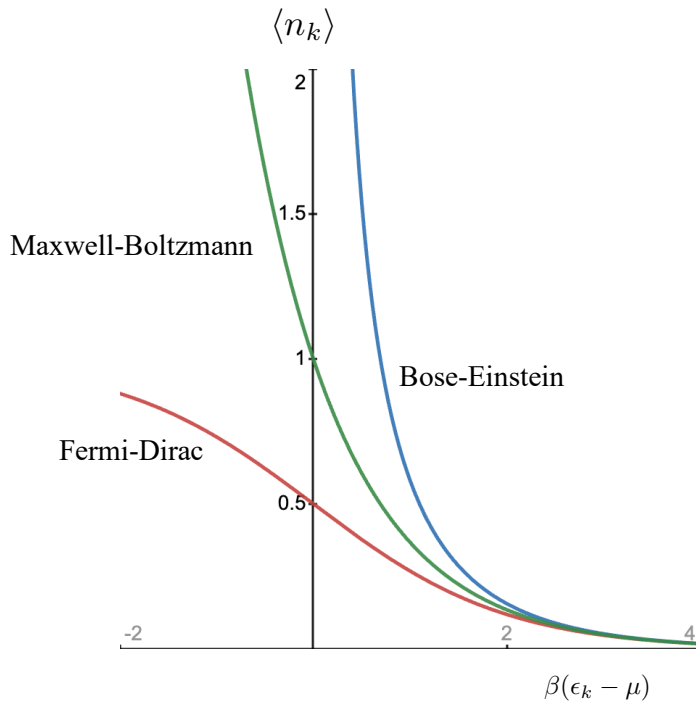


## Quantum gases



Fermi - Dirac distribution:

$$\langle n_k \rangle = \frac{1}{e^{\beta(\epsilon_k - \mu)} + 1}$$

Bose - Einstein distribution:

$$\langle n_k \rangle = \frac{1}{e^{\beta(\epsilon_k - \mu)} - 1}$$

Maxwell - Boltzmann distribution:

$$\langle n_k \rangle = e^{-\beta(\epsilon_k - \mu)}$$

### Fermi gas (ideal case)

The average number of particles can now be written as

$$\langle N \rangle = \sum_k \frac{1}{\zeta^{-1} e^{\beta \epsilon_k} + 1}$$

As we approximated for Bosons,  $\langle N \rangle$  can be expressed as

$$\langle N \rangle = \frac{4 \pi V}{h^3} \int_0^\infty \frac{p^2}{\zeta^{-1} e^{\beta p^2/(2m)} + 1} dp$$

Choosing a new variable  $t = \beta p^2/(2m)$ , we get

$$\langle N \rangle = \frac{2 V}{\sqrt{\pi} h^3} (2\pi m k_B T)^{3/2} \int_0^\infty \frac{\sqrt{t}}{\zeta^{-1} e^t + 1} dt$$

Where  $\lambda = h/\sqrt{2\pi m k_B T}$  and the integral is defined as

$$f_\nu(\zeta) = \frac{1}{\Gamma(\nu)} \int_0^\infty \frac{t^{\nu-1}}{\zeta^{-1} e^t + 1} dt$$

Here,  $\Gamma(\nu)$  is the gamma function. Thus, number of particles per unit volume can be written as

$$\frac{\langle N \rangle}{V} = n = \frac{1}{\lambda^3} f_{3/2}(\zeta)$$

Where  $n$  is particle density.

### Equation of state (Fermions)

$$\begin{aligned}\frac{PV}{k_B T} &= \sum_k \ln[1 + \zeta e^{-\beta \epsilon_k}] \\ \frac{PV}{k_B T} &= \frac{4 \pi V}{h^3} \int_0^\infty p^2 \ln[1 + \zeta e^{-\beta p^2/(2m)}] dp \\ \frac{PV}{k_B T} &= \frac{2 V}{\sqrt{\pi} h^3} (2 \pi m k_B T)^{3/2} \int_0^\infty \sqrt{t} \ln[1 + \zeta e^{-t}] dt\end{aligned}$$

The integration can be done by parts to obtain

$$\frac{PV}{k_B T} = \frac{2 V}{\sqrt{\pi} h^3} \frac{1}{\lambda^3} \left[ \frac{t^{3/2}}{3/2} \ln[1 + \zeta e^{-t}] - \frac{2}{3} \int_0^\infty \frac{t^{3/2} \zeta e^{-t}}{1 + \zeta e^{-t}} dt \right]$$

$$\frac{P}{k_B T} = \frac{1}{\lambda^3} \frac{1}{\Gamma(5/2)} \int_0^\infty \frac{t^{\frac{5}{2}-1}}{\zeta^{-1} e^t + 1} dt$$

$$\frac{P}{k_B T} = \frac{1}{\lambda^3} f_{5/2}(\zeta)$$

### Alternative expressions for $g_\nu(\zeta)$ and $f_\nu(\zeta)$

$$\begin{aligned}g_\nu(\zeta) &= \sum_{k=1}^\infty \frac{\zeta^k}{k^\nu} \\ f_\nu(\zeta) &= \sum_{k=1}^\infty (-1)^{k+1} \frac{\zeta^k}{k^\nu}\end{aligned}$$

### Energy density

$$\langle E \rangle = \sum_k \epsilon_k \langle n_k \rangle$$

The general expression is

$$\frac{\langle E \rangle}{V} = \frac{1}{V} \sum_k \frac{\epsilon_k}{\zeta^{-1} e^{\beta \epsilon_k} \pm 1}$$

Where + symbol corresponds to fermions and - symbol corresponds to bosons.

$$\frac{\langle E \rangle}{V} = \frac{4\pi}{h^3} \int_0^\infty p^2 \frac{\epsilon(p)}{\zeta^{-1} e^{\beta \epsilon(p)} \pm 1} dp$$

This can be shown as, solving the integrals as above,

$$\frac{\langle E \rangle}{V} = \frac{3}{2} P$$

Implies,

$$P = \frac{2}{3} \frac{\langle E \rangle}{V}$$

This expression is true for both fermions and bosons.

### Comparison between bosons and fermions

Bosons	Fermions
$\frac{\langle N \rangle}{V} = \frac{1}{\lambda^3} g_{3/2}(\zeta)$	$\frac{\langle N \rangle}{V} = \frac{1}{\lambda^3} f_{3/2}(\zeta)$
$\frac{P}{k_B T} = \frac{1}{\lambda^3} g_{5/2}(\zeta)$	$\frac{P}{k_B T} = \frac{1}{\lambda^3} f_{5/2}(\zeta)$
$\frac{\langle E \rangle}{V} = \frac{3}{2} P$	$\frac{\langle E \rangle}{V} = \frac{3}{2} P$
$\frac{\langle E \rangle}{\langle N \rangle} = \frac{3}{2} k_B T \frac{g_{5/2}(\zeta)}{g_{3/2}(\zeta)}$	$\frac{\langle E \rangle}{\langle N \rangle} = \frac{3}{2} k_B T \frac{f_{5/2}(\zeta)}{f_{3/2}(\zeta)}$

### Equation of state (Virial expansion)

At low densities, i.e.,  $\zeta \ll 1$ , the equation of state can be expressed as

$$\frac{P}{k_B T} = \frac{1}{\lambda^3} \left\{ \frac{f_{5/2}(\zeta)}{2} \right\}$$

Where  $f_{\frac{5}{2}}(\zeta)$  corresponds to fermions and  $g_{5/2}(\zeta)$  corresponds to bosons. Using the alternative expressions  $g_v(\zeta)$  and  $f_v(\zeta)$ , the equation of state can be written as

$$\frac{P}{k_B T} = \frac{-1}{\lambda^3} \left( -\zeta \pm \frac{\zeta^2}{2^{\frac{5}{2}}} \pm \dots \dots \right)$$

Here + symbol corresponds to fermions and - symbol corresponds to bosons. Similarly,

$$\frac{\langle N \rangle}{V} = \frac{1}{\lambda^3} \left\{ \frac{f_{3/2}(\zeta)}{2} \right\} = \frac{-1}{\lambda^3} \left( -\zeta \pm \frac{\zeta^2}{2^{\frac{3}{2}}} \pm \dots \dots \right)$$

This implies, the ratio of  $\frac{P}{k_B T}$  and  $\frac{\langle N \rangle}{V}$  will give us

$$\begin{aligned}\frac{P}{k_B T} &= \frac{\langle N \rangle}{V} \frac{\left( -\zeta \pm \frac{\zeta^2}{2^{\frac{5}{2}}} \pm \dots \right)}{\left( -\zeta \pm \frac{\zeta^2}{2^{\frac{3}{2}}} \pm \dots \right)} \\ \frac{P}{k_B T} &= \frac{\langle N \rangle}{V} \left( -1 \pm \frac{\zeta}{2^{\frac{5}{2}}} \right) \left( -1 \pm \frac{\zeta}{2^{\frac{3}{2}}} \right)^{-1} \\ \frac{P}{k_B T} &= \frac{\langle N \rangle}{V} \left( -1 \pm \frac{\zeta}{2^{\frac{5}{2}}} \right) \left( -1 \mp \frac{\zeta}{2^{\frac{3}{2}}} \right) \\ \frac{P}{k_B T} &= \frac{\langle N \rangle}{V} \left( 1 \pm \frac{\zeta}{2^{\frac{3}{2}}} \mp \frac{\zeta}{2^{\frac{5}{2}}} \right) \\ \frac{P}{k_B T} &= \frac{\langle N \rangle}{V} \left( 1 \pm \frac{\zeta}{2^{\frac{5}{2}}} \right)\end{aligned}$$

In the above expression,  $\pm \zeta/2^{5/2}$  term is the quantum correction to the classical ideal gas. Here, + symbol corresponds to fermions and - symbol corresponds to bosons.

For small  $\zeta$ , we can write  $\frac{\langle N \rangle}{V} = \frac{\zeta}{\lambda^3}$ . This implies,  $\zeta = \frac{\langle N \rangle}{V} \lambda^3$ . So, in this limit the equation of state reads,

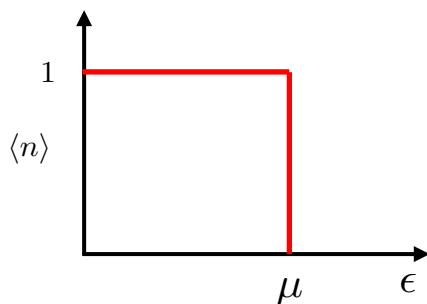
$$P V = \langle N \rangle k_B T \left( 1 \pm \frac{\langle N \rangle}{V} \frac{\lambda^3}{2^{5/2}} \right)$$

## Electrons in a conductor

The average occupation number of fermions, in a particular energy level, is given by the expression

$$\langle n \rangle = \frac{1}{e^{\beta(\epsilon - \mu)} + 1}$$

In the  $T \rightarrow 0$ , limit



$$e^{\beta(\epsilon - \mu)} = \begin{cases} \infty & \epsilon > \mu \\ 0 & \epsilon < \mu \end{cases}$$

$$\langle n \rangle = \begin{cases} 0 & \epsilon > \mu \\ 1 & \epsilon < \mu \end{cases}$$

All the states with  $\epsilon < \mu$  are filled and all the states  $\epsilon > \mu$  are empty. This is the  $T = 0$  ground state of the fermi gas. We there see that  $\mu (T = 0)$  is the highest occupied energy state. One calls this energy the “Fermi - energy”  $\epsilon_F = \mu (T = 0)$ .

The total occupation number,

$$\langle N \rangle = \sum_{\substack{\text{States with} \\ \epsilon < \epsilon_F}} \frac{1}{e^{\beta(\epsilon_F - \mu (T=0))} + 1}$$

Here,  $\epsilon_F - \mu (T = 0) = 0$ . So,

$$\langle N \rangle = \sum_{\substack{\text{States with} \\ \epsilon < \epsilon_F}} 1$$

This can be simplified using the integral

$$\langle N \rangle = \frac{4 \pi V}{h^3} \int_0^{\sqrt{2 m \epsilon_F}} p^2 dp$$

$$\langle N \rangle = \frac{4 \pi V}{3 h^3} (2 m \epsilon_F)^{3/2}$$

$$\frac{\langle N \rangle}{V} = n = \frac{4 \pi}{3 h^3} (2 m \epsilon_F)^{3/2}$$

$$n = \frac{4 \pi}{3} \left( \frac{2 m \epsilon_F}{h^2} \right)^{3/2}$$

Thus, the fermi energy is given by

$$\epsilon_F = \left( \frac{3 n}{4 \pi} \right)^{2/3} \frac{h^2}{2 m}$$

The corresponding fermi temperature is

$$T_F = \frac{\epsilon_F}{k_B}$$

For electrons in a metal  $T_F \approx 10,000 K$ . So, electrons in a metal always in degenerate limit, i.e.  $T < T_F$ .

## Finite temperature

In the limit,  $k_B T \ll \epsilon_F$ , we get

$$k_B T \ll \left( \frac{3 n}{4 \pi} \right)^{2/3} \frac{h^2}{2 m}$$

After re-arranging the terms,

$$\begin{aligned}
\frac{2\pi m k_B T}{h^2} &\ll \pi \left( \frac{3n}{4\pi} \right)^{2/3} \\
\left( \frac{2\pi m k_B T}{h^2} \right)^{3/2} &\ll \pi^{3/2} \frac{3n}{4\pi} \\
\left( \frac{\sqrt{2\pi m k_B T}}{h} \right)^3 &\ll \frac{3n\sqrt{\pi}}{4} \\
\frac{1}{n\lambda^3} &\ll \frac{3\sqrt{\pi}}{4} \\
n\lambda^3 &\gg \frac{4}{3\sqrt{\pi}}
\end{aligned}$$

This implies,  $n\lambda^3 \gg 1$  is called the low  $T$  or high-density limit. Similarly,  $n\lambda^3 \ll 1$  is called the high  $T$  or classical limit.

### Energy in the degenerate limit ( $T = 0$ )

Consider the expression for the number density,

$$n = \frac{4\pi}{3} \left( \frac{2m\epsilon_F}{h^2} \right)^{3/2} =$$

Using this relation, we can define the density of state, i.e. number of states with energy  $\epsilon$  per unit energy per volume  $g(\epsilon)$ . The density of states,  $g(\epsilon) = \frac{n}{\epsilon} = C\sqrt{\epsilon}$ . Where the constant

$$C = \frac{2}{\pi} \left( \frac{2\pi m}{h^2} \right)^{3/2}$$

Now, we will calculate the following quantity,

$$\begin{aligned}
\frac{\langle E \rangle}{V} &= \int_0^{\epsilon_F} d\epsilon g(\epsilon) \epsilon \\
&= C \int_0^{\epsilon_F} d\epsilon \sqrt{\epsilon} \epsilon \\
\frac{\langle E \rangle}{V} &= \frac{2C}{5} \epsilon_F^{5/2}
\end{aligned}$$

In terms of the constant  $C$  the density of particles is

$$n = \frac{\langle N \rangle}{V} = \frac{2C}{3} \epsilon_F^{3/2}$$

This implies, the energy per particle is

$$\frac{\langle E \rangle}{\langle N \rangle} = \frac{3}{5} \epsilon_F$$