# High precision perturbative QCD predictions for Higgs production at the LHC

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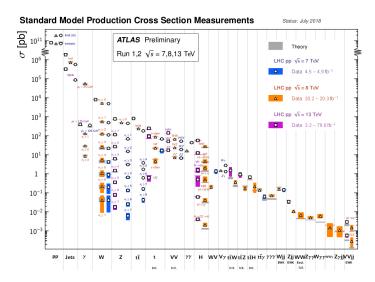


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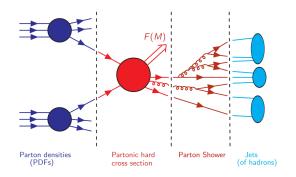
## Outline

- QCD in a nutshell
  - Factorization in QCD
  - Partonic cross section
  - Perturbative QCD
- Resummation in QCD
  - Higher order corrections
  - Resummation of  $q_{\perp}$
- 4 HTurbo
  - Higgs production at the LHC: HRes and HqT
  - HTurbo: Fast predictions for Higgs production
  - Results & Conclusions

### LHC physics



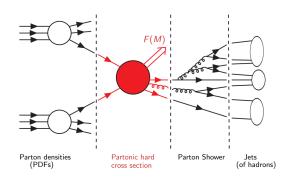
#### Factorization theorem



Compute cross sections is a Hard problem → QCD Factorization

$$\sigma^{F}(p_{1}, p_{2}) = \int_{0}^{1} dx_{1} dx_{2} f_{\alpha}(x_{1}, \mu_{F}^{2}) * f_{\beta}(x_{2}, \mu_{F}^{2}) * \hat{\sigma}_{\alpha\beta}^{F}(x_{1}p_{1}, x_{2}p_{2}, \alpha_{s}(\mu_{R}^{2}), \mu_{F}^{2})$$

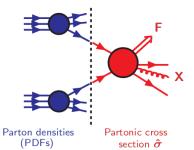
#### Partonic cross section



- PDFs  $f_{\alpha}(x_i, \mu_F^2)$  absorb the non perturbative effects, evaluated at  $\mu_F$
- Partonic  $\hat{\sigma}_{\alpha\beta}^{F}$  can be computed as perturbative series in  $\alpha_s$

#### Perturbative QCD

- Born cross section as LO value of a perturbative series
- $\sigma^{(1)}, \sigma^{(2)}, \sigma^{(3)}$  are the NLO, NNLO, N3LO corrections

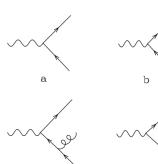


$$\hat{\sigma} = \sigma^{\mathtt{Born}} \left( 1 + \frac{\alpha_s}{2\pi} \sigma^{(1)} + \left( \frac{\alpha_s}{2\pi} \right)^2 \sigma^{(2)} + \left( \frac{\alpha_s}{2\pi} \right)^3 \sigma^{(3)} + \ldots \right)$$

Leading order predictions can strongly depend on the renormalization and factorization scales  $\rightarrow$  Need higher order corrections!

### Higher order corrections

- Higher order corrections are not an easy task due to infrared (IR) singularities
- Final state radiations cancel by combining real and virtual contributions
- Initial state radiations factorized inside the PDFs, then IR-free, finite  $\hat{\sigma}$



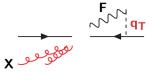
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 $q_{\perp}$  resummation

Study the differential  $q_{\perp}$  distribution

$$h_1(p_1) + h_2(p_2) \longrightarrow F(M, \mathbf{q}_{\perp}) + X$$

$\alpha_S L^2$	$\alpha_{\mathcal{S}}L$	 $\mathcal{O}(\alpha_{\mathcal{S}})$
$\alpha_S^2 L^4$	$\alpha_S^2 L^3$	 $\mathcal{O}(\alpha_S^2)$
$\alpha_S^n L^{2n}$	$\alpha_S^n L^{2n-1}$	 $\mathcal{O}(\alpha_S^n)$
dominant logs		 



Truncated fixed order predictions  $\rightarrow$  divergent  $\alpha_s^n \ln^m(M^2/q_\perp^2)$  appear

 $q_{\perp}$  resummation

Replace partonic  $q_{\perp}$  distribution as follows

$$\frac{d\hat{\sigma}_{ab}}{dq_{\perp}^2} \rightarrow \left[\frac{d\hat{\sigma}_{ab}^{(\rm res.)}}{dq_{\perp}^2}\right]_{\rm l.a.} + \left[\frac{d\hat{\sigma}_{ab}^{(\rm fin.)}}{dq_{\perp}^2}\right]_{\rm f.o.} \quad , \, {\rm such \, \, that}$$

$$\begin{split} &\int dq_\perp^2 \frac{d\hat{\sigma}_{ab}^{(\mathrm{res.})}}{dq_\perp^2} \sim \sum \alpha_s^n \log^m \frac{M^2}{Q_\perp^2} \quad \mathrm{for} \quad \ q_\perp \to 0 \\ &\int dq_\perp^2 \frac{d\hat{\sigma}_{ab}^{(\mathrm{fin.})}}{dq_\perp^2} \sim 0 \quad \mathrm{for} \quad \ q_\perp \to 0 \end{split}$$

Resummed and finite components need to be matched at some logarithmic accuracy (LL, NLL, NNLL, ...)

 $q_{\perp}$  resummation

Resummation holds in impact parameter space b

$$rac{d\hat{\sigma}_{ab}^{(\mathrm{res.})}}{dq_{\perp}^2} = rac{\mathit{M}^2}{\hat{\mathsf{s}}} \int db \; rac{b}{2} \; J_0(bq_{\perp}) \; \mathcal{W}_{ab}(b, M)$$

Which is expressed in Mellin space (with respect to  $z = M^2/\hat{s}$ )

$$W_N(b, M) = \mathcal{H}_N(\alpha_s) \times \exp\{\mathcal{G}_N(\alpha_s, L)\}$$
 being  $L \equiv \log(M^2 b^2)$ 

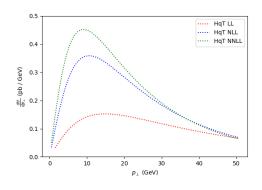
- Resummed effects exponentiated in the Sudakov factor  $\mathcal{G}_N(\alpha_s, L)$
- Process-dependence factorized in the hard factor  $\mathcal{H}_N(\alpha_s)$

HTurbo: Fast predictions for Higgs production

# HqT and HRes

## Predictions for Higgs $q_{\perp}$ distribution

- HqT and HRes [de Florian, G.F., Grazzini, Tommasini] produce q<sub>1</sub> distributions
- Higher order corrections require high computation times
- Codes producing fast predictions are needed for precision era of the LHC



## **HTurbo**

#### Starting point DYTurbo

• Start from **DYTurbo** [Camarda et al.] ref. at 1910.07049, producing  $q_{\perp}$  distribution fro Drell-Yan  $(q\bar{q} \rightarrow l^+ l^-)$ 

C++ interface rather that Fortran of HRes or HqT

- Set LO amplitude  $gg \rightarrow H$
- Set Sudakov and Hard coefficients for Higgs production
- Implement PDF evolution
- Compare with HRes and HqT

State of the art is just up to NNLL!

## **HTurbo**

## Starting point DYTurbo

- ullet Set LO amplitude to be gg o H
- Set Sudakov and Hard coefficients for Higgs production
- Compare with **HRes** and **HqT**

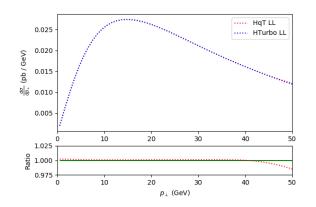
$$\mathcal{G}_{N}(\alpha_{s}, L) = L g^{(1)}(\alpha_{s}L) + g^{(2)}(\alpha_{s}L) + \frac{\alpha_{s}}{\pi}g^{(3)}(\alpha_{s}L) + \dots$$
$$\mathcal{H}_{N}(\alpha_{s}) = 1 + \frac{\alpha_{s}}{\pi}\mathcal{H}^{(1)} + \left(\frac{\alpha_{s}}{\pi}\right)^{2}\mathcal{H}^{(2)} + \dots$$

LL(
$$\sim \alpha_s^n L^{n+1}$$
):  $g^{(1)}$ ,  $\hat{\sigma}^{(0)}$   
NLL( $\sim \alpha_s^n L^n$ ):  $g^{(2)}$ ,  $\mathcal{H}^{(1)}$ 

NNLL(
$$\sim \alpha_s^n L^{n-1}$$
):  $g^{(3)}$ ,  $\mathcal{H}^{(2)}$ 

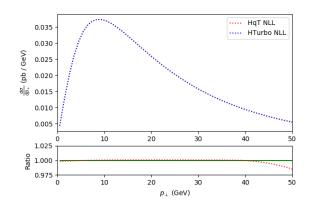
Start by building predictions up to NNLO, then add  $N^3LL$ 

#### Comparison HTurbo and HqT - LL



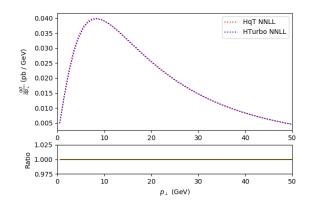
- ullet HTurbo  $q_{\perp}$  distribution matches HRes and HqT at LL
- ullet Excellent numerical agreement up to the 1/1000 level

#### Comparison HTurbo and HqT - NLL



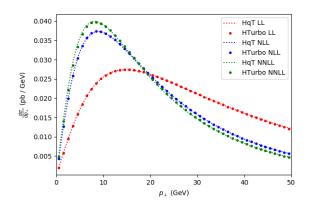
- HTurbo  $q_{\perp}$  distribution matches HRes and HqT at NLL
- Agreement obtained by switching off PDF evolution

#### Comparison HTurbo and HqT - NNLL



- HTurbo  $q_{\perp}$  distribution matches HRes and HqT at NNLL
- Agreement obtained by switching off PDF evolution

#### Comparison HTurbo and HqT - all orders



- Higher orders lead to more reliable predictions √
- Agreement up to NNLL  $\longrightarrow$  ready for N<sup>3</sup>LL

# Summary & Conclusions

- Fast predictions are required towards the precision era of the LHC
- $oldsymbol{0}$   $q_{\perp}$  distributions from HTurbo match HRes and HqT up to NNLL
- HTurbo is much faster than any of the existing Higgs codes
- Next steps: Implement PDF evolution and N<sup>3</sup>LO distributions

## Thank you!



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