

# Physics Beyond

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## Kinematics

### How to describe motion? (mathematically)

- In maths we need a clear and precise definition.
- Motion is change in position over time.
  - ↪ What is space?
  - ↪ What is time?
- How to describe motion?
  - ↪ Need a reference point / origin.
  - ↪ Intuition about vectors.
    - ↪ Rule of displacement from one point to another along a straight line
    - ↪ Vector - from Latin "vehere" - to carry

To describe motion we need  $f(t) \rightarrow x$  and  $g(t) \rightarrow y$ .

$$\begin{aligned}\vec{x} : \mathbb{R} &\rightarrow \text{Vector Space} \\ t &\mapsto \vec{x}(t)\end{aligned}$$

### What is vector space?

- Euclidean 3-dimensional space.
- For example, vector space of a 3 tuple of reals can be written as  $\mathbb{R}^3$

→ We say  $V$  is a vector space if:

1.  $\mathbb{R} \cdot V \rightarrow V$
2. Addition is commutative, associative, and has neutral element  $\vec{0}$
3.  $(\lambda + \mu)\vec{v} = \lambda\vec{v} + \mu\vec{v}$        $\lambda, \mu \in \mathbb{R}$        $\vec{v} \in V$
4.  $(\lambda\mu)\vec{v} = \lambda(\mu\vec{v})$
5.  $1 \cdot \vec{v} = \vec{v}$

→ We need to introduce a coordinate system.

In 2 dimensions

$$\vec{e}_1 = \begin{bmatrix} 1 \\ 0 \end{bmatrix}$$

$$\vec{e}_2 = \begin{bmatrix} 0 \\ 1 \end{bmatrix}$$

In n dimensions

$$\vec{e}_i = \begin{bmatrix} 0 \\ 0 \\ \vdots \\ 1 \\ \vdots \\ 0 \end{bmatrix}$$

Where 1 is at index  $i$

→ We can describe using unitary vectors:

Two sets  $A, B$        $A \times B = \{(a, b) \mid a \in A, b \in B\}$   
 $A \times A = A^2$  Where  $\times$  is the cartesian product

## What is happening mathematically?

We have constructed a mapping  $V \rightarrow \mathbb{R}^2$ ; it is  $1 \rightarrow 1$   
 $\therefore$  For any vector, a pair of reals visa versa.

$$\vec{x} \mapsto (x_1, x_2) \quad s.t \quad \vec{x} = x_1\vec{e}_1 + x_2\vec{e}_2$$

1. Note that for a basis,  $\vec{e}_1, \vec{e}_2$  this mapping is one to one and onto (bijection)  
 $\vec{x}(f) = (ae_1, be_2) \leftarrow$  Linear combination of  $\vec{e}_1$  and  $\vec{e}_2$

2. This means that we have an inverse mapping.

$$\begin{aligned} \mathbb{R} &\rightarrow E \\ \begin{bmatrix} \lambda \\ \mu \end{bmatrix} &\mapsto \lambda e_1 + \mu e_2 \end{aligned}$$

Coordinate systems translate  $\mathbb{R}^2$  to vector E

3. This mapping depends on the chosen coordinate system.

→ Coordinate system: (origin, two basis vectors)

↪  $(\vec{o}, \vec{e}_1, \vec{e}_2)$

↪  $\vec{e}_1, \vec{e}_2$  are linearly independent.

↪ We assume  $\vec{e}_1 \perp \vec{e}_2 \rightarrow$  Is orthogonal (meet at 90 deg /  $\frac{\pi}{2}$ )

We have an improvement of our description:

$$\begin{aligned} \vec{x} : \mathbb{R} &\rightarrow \mathbb{R}^2 \\ t &\mapsto \vec{x}(t) \qquad \vec{x}(t) = \begin{bmatrix} x_1(t) \\ x_2(t) \end{bmatrix} \qquad x_1, x_2 : \mathbb{R} \rightarrow \mathbb{R} \end{aligned}$$

## Operators

$$+ : \mathbb{R}^n, \mathbb{R}^n \rightarrow \mathbb{R}^n$$

$$\left( \begin{bmatrix} x_1 \\ \vdots \\ x_n \end{bmatrix}, \begin{bmatrix} y_1 \\ \vdots \\ y_n \end{bmatrix} \right) \mapsto \begin{bmatrix} x_1 + y_1 \\ \vdots \\ x_n + y_n \end{bmatrix}$$

$$\cdot : \mathbb{R}, \mathbb{R}^n \rightarrow \mathbb{R}^n$$

$$\left( \lambda, \begin{bmatrix} x_1 \\ \vdots \\ x_n \end{bmatrix} \right) \mapsto \begin{bmatrix} \lambda x_1 \\ \vdots \\ \lambda x_n \end{bmatrix}$$

$$\mathbb{R}^x = \{f : x \rightarrow \mathbb{R}\} \quad \text{Where } x \text{ is any set and } f \text{ is a function}$$

# Maths foundations

Current foundation: Set theory. What does that mean?

→ Everything must be related back to sets.

→ Relatively modern idea (2nd half of 20th century)

## Examples

1. Numbers:

$0 = \emptyset$  - empty set

$1 = \{\emptyset\}$  - set of the empty set

$2 = \{\emptyset, \{\emptyset\}\}$  - set of the empty set and a set of the empty set

$\vdots$

This means that natural numbers can be modelled using set theory.

→ Model ideas = helps us to understand abstract ideas.

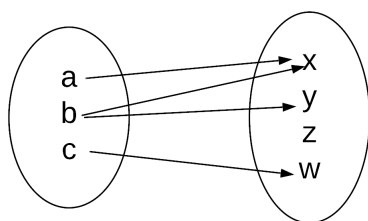
2. Ordered pair:  $(x, y)$  - order matters

How can we model this using sets? (sets are unordered)

$\{x, y\}$  doesn't work as unordered.

$\{x, \{x, y\}\}$  could work for example.

3. Relation



Relation  $\mathbb{R} \subseteq A \times B$

$(A \times B = \{(a, b) \mid a \in A, b \in B\})$

Example 1:

$x \leq y \in \mathbb{R}$

$\leq \subseteq \mathbb{R} \times \mathbb{R}$

$(x, y) \in \leq$

Example 2 - Equivalence relation:

(Example of vectors)

Vectors:

- (a) Idea: Rule of displacement of things along a straight line by a certain length (physics)
- (b) Representation: Directed line segment

How to model a vector?

Consider the set  $L$  of all line segments in the plane

- Subdivide  $L$  into disjointed subsets  $[\vec{AB}]$  of line segments parallel, of same length and same orientation as  $\vec{AB}$   
 $\vec{v} = [\vec{AB}]$  is a (model of a) vector.
- Define a relation  $\sim \subseteq L \times L$  ( $\sim$  means equivalent to)

$$\begin{aligned}\vec{AB} \sim \vec{CD} : & \iff \text{They represent the same vector} \\ & \iff [\vec{AB}] = [\vec{CD}] \\ & \iff \vec{AB} \parallel \vec{CD}, \|\vec{AB}\| = \|\vec{CD}\|\end{aligned}$$

In general, these three properties define what is called an equivalence relation.

#### 4. Functions / maps / mapping

$$\rho : A \rightarrow B \quad x \mapsto f(x)$$

Functions can be considered a type of relation. A relation is called a function if it is right-unique and left-total.

$$\begin{aligned}(x, y) \wedge (x, z) & \Rightarrow y = z \\ vx & \in A \\ \exists y & \in B \\ (x, y) & \in f \\ y & = f(x)\end{aligned}$$

Example:

$$\begin{aligned}\rho : \mathbb{R} & \rightarrow \mathbb{R} & x & \mapsto x^2 \\ \rho & \subseteq \mathbb{R} \times \mathbb{R} & \rho & = \{(x, y) \in \mathbb{R}^2 \mid y = x^2\}\end{aligned}$$

The relation is the graph of the function.

## Motion

Two areas:

- Kinematics (mathematical modelling)
- Dynamics (what causes motion)

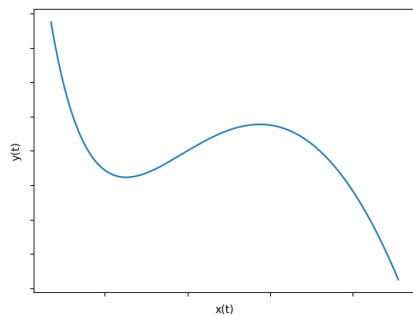
Def: Motion is the change in position in time.

position  $\rightarrow$  space                      time  $\rightarrow$  time

$\leadsto$  Space + time too abstract  $\rightarrow$  need to simplify

Space  $\mapsto$  Euclidean geometry

Time  $\mapsto$  Use  $\mathbb{R}$  (Time is just a parameter)



Say this is the graph of  $p(t)$

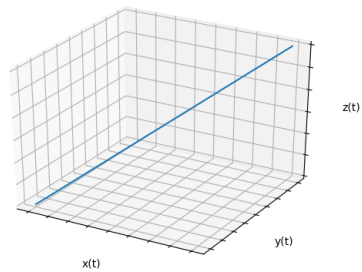
$p : \mathbb{R} \rightarrow \text{Space}$

$\leadsto$  Vectors + coordinates (reference frame)

$$x : \mathbb{R} \rightarrow \mathbb{R}^3$$

$$\vec{x} = x_1 e_1 + x_2 e_2 + x_3 e_3$$

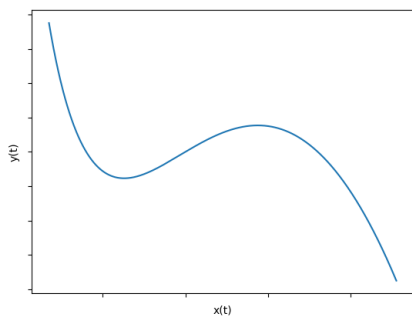
$$\vec{x} = \begin{bmatrix} x_1 \\ x_2 \\ x_3 \end{bmatrix}$$



$\rightsquigarrow$  Algebra of geometric vectors translates directly into algebra of tuples.

$$\vec{x} + \vec{y} \longleftrightarrow \begin{bmatrix} x_1 \\ x_2 \\ x_3 \end{bmatrix} + \begin{bmatrix} y_1 \\ y_2 \\ y_3 \end{bmatrix} = \begin{bmatrix} x_1 + y_1 \\ x_2 + y_2 \\ x_3 + y_3 \end{bmatrix}$$

$$\lambda \vec{x} \longleftrightarrow \lambda \begin{bmatrix} x_1 \\ x_2 \\ x_3 \end{bmatrix} = \begin{bmatrix} \lambda x_1 \\ \lambda x_2 \\ \lambda x_3 \end{bmatrix}$$



$$\vec{x} : \mathbb{R} \rightarrow \mathbb{R}^3$$

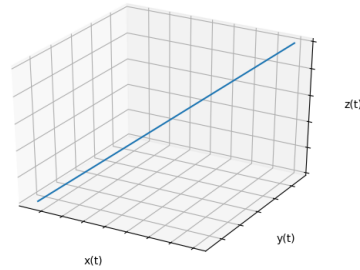
$$t \mapsto \vec{x}(t)$$

$$\vec{x}(t) = \begin{bmatrix} x_1(t) \\ x_2(t) \\ x_3(t) \end{bmatrix}$$

## Exaples

1. Motion along a line (with constant velocity)

$$\begin{aligned}\vec{x}(t) &= \vec{x}_0 + t\vec{v} \\ \begin{bmatrix} x_1(t) \\ x_2(t) \\ x_3(t) \end{bmatrix} &= \begin{bmatrix} x_{01} \\ x_{02} \\ x_{03} \end{bmatrix} + t \begin{bmatrix} v_1 \\ v_2 \\ v_3 \end{bmatrix} \\ &= \begin{bmatrix} x_{01} + tv_1 \\ x_{02} + tv_2 \\ x_{03} + tv_3 \end{bmatrix}\end{aligned}$$



2. Circular motion

$$\vec{x}(t) = r \begin{bmatrix} \cos(\rho t) \\ \sin(\rho t) \end{bmatrix}$$

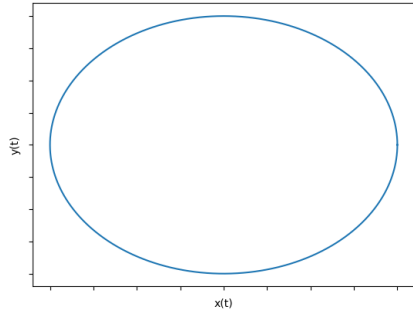
Where  $\rho$  some function and  $r$  is the radius.

If it is uniform:

$$\begin{aligned}\vec{x}(t) &= r \begin{bmatrix} \cos(\omega t) \\ \sin(\omega t) \end{bmatrix} \\ \omega &\in \mathbb{R} \setminus \{0\} \qquad \omega = \frac{2\pi}{T}\end{aligned}$$

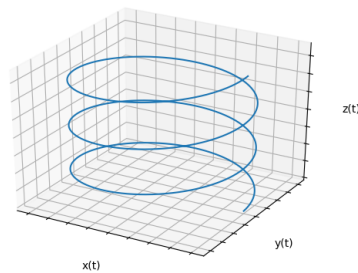
Where  $T$  is the time period.





3. Uniform spiral It is just like a circle, but with linear motion in the z-axis.

$$\vec{x}(f) = \begin{bmatrix} r \cos \omega t \\ r \sin \omega t \\ vt \end{bmatrix}$$



## Kinematics - reminder

- ~> Done with modelling motion mathematically?
  - ↪ How to study the physics of motion?
- ~> What is the natural state of motion?
  - Without physical interaction.
- ~> At rest? (Aristotle) ×
- ~> Uniform motion (Newton)
  - ↪ Why?

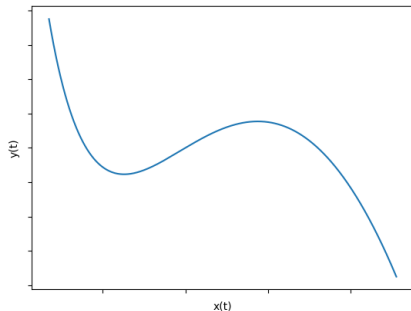
- ↔ Changing reference frame to a moving one → object at rest relatively.
- ↔ But we cannot answer/ decide what the natural state of motion is in this stage.

Observation: State of motion changes due to physical interaction → Lets model physical interaction with forces. What is the change in motion due to force?

- ↪ The shape of the trajectory is curved
- ↪ We don't know how to model that!
- ↪ Back to maths

## Kinematics 2

### Modelling change in motion



Take two points on the above curve,  $\vec{x}(t_0)$  and  $\vec{x}(t)$ . The change of displacement is  $\Delta\vec{x}$ . We know that velocity is change in displacement over change in time so we can find that using:

$$\begin{aligned}\vec{v} &= \frac{\Delta\vec{x}}{\Delta t} \\ &= \frac{\vec{x}(t) - \vec{x}(t_0)}{t - t_0}\end{aligned}$$

This gives us the average velocity between  $t_0$  and  $t$ . We need to be able to find the velocity at every point in time  $t_0$

$$\begin{aligned}\vec{v} &: \mathbb{R} \rightarrow \mathbb{R}^3 \\ t &\mapsto \text{velocity at } t\end{aligned}$$

Problem: We need a second time  $t = t_0 + h$  for  $\vec{v}(t_0)$ , but there is no natural choice for  $h$ . We notice for more precision we need to make  $h$  as small as possible!

$$\begin{aligned} & \lim_{h \rightarrow 0} \frac{\vec{x}(t+h) - \vec{x}(t)}{h} \\ &= \frac{d\vec{x}}{dt} \text{ Leibniz notation} \\ &= \dot{\vec{x}}(t) \text{ Newton notation} \end{aligned}$$

Differentiation on vectors is just element wise:

$$\frac{d}{dt} \begin{bmatrix} R \cos \omega t \\ R \sin \omega t \end{bmatrix} = \omega R \begin{bmatrix} -\sin \omega t \\ \cos \omega t \end{bmatrix}$$

Limits create paradoxes: Take  $\lim_{h \rightarrow 0} \frac{\vec{x}(t+h) - \vec{x}(t)}{h}$ . This works mathematically but causes a conceptual problem: How can something change at one point in time?

$\rightsquigarrow$  Solution: Use infinitesimals

$\rightarrow$  Numbers that are not provably not 0. Idea: If we zoom in on any curve enough, begins to look like a straight line: e.g. Earth looks flat.

How can we model this mathematically?

$\rightsquigarrow$  Around each point  $\vec{x}(t_0)$  the curve is made up of infinitesimal line segments.

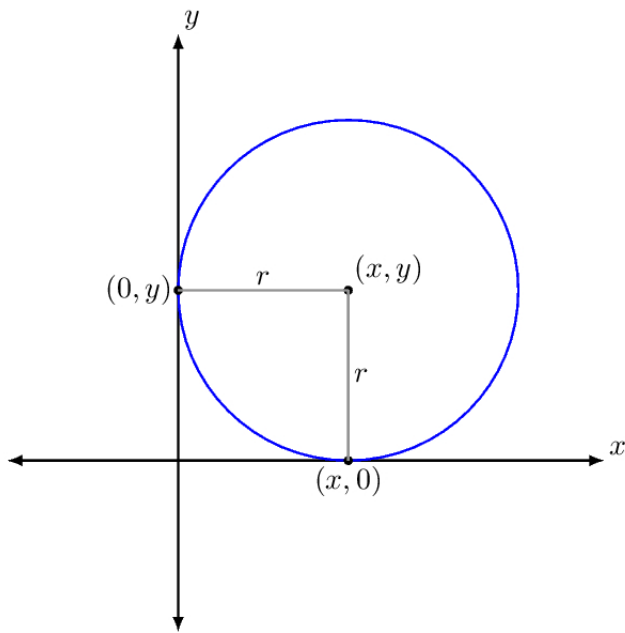
Geometrically: A tangent at  $\vec{x}(t_0)$  touches the curve at  $\vec{x}(t_0)$

$\Leftrightarrow$  We take this to mean that it intersects with the curve at an infinitesimal line segment

$\Leftrightarrow$  Infinitesimal line segment is the intersection of the tangent and the curve.

Problem: Do not have a rigorous definition of a tangent at a curve!

$\rightarrow$  But we know: Every line that intersects the circle in one points is a tangent  $\rightsquigarrow$  Use that!



Take the radius,  $r$ , to be 1 and the center to be  $(0, 1)$ .

Equation of circle:

$$c : (x - 0)^2 + (y - 1)^2 = 1$$

If we take our tangent,  $t$  to be the  $y$  axis:

$$t : y = 0$$

Sub into  $c$ :

$$x^2 + 1 = 1$$

$$\Rightarrow x^2 = 0$$

$$\therefore c \cap t = \{(d, 0) \mid d^2 = 0\}$$

$\rightsquigarrow$  Define first order infinitesimals as  $D = \{d \in R \mid d^2 = 0\}$  where  $R = \mathbb{R} \cup \{\text{infinitesimal numbers}\}$  Are such infinitesimals going to solve problems?

$$\begin{aligned}
 f(x) &= x^2 \\
 \Rightarrow f(x+d) & \qquad d \in D \\
 &= (x+d)^2 \\
 &= x^2 + 2dx + d^2 \\
 &= x^2 + 2dx \\
 &= f(x) + f'(x) \cdot d
 \end{aligned}$$

$$\begin{aligned}
 f(x) &= x^3 \\
 \Rightarrow f(x+d) & \\
 &= (x+d)^3 \\
 &= x^3 + 3x^2d + 3xd^2 + d^3 \\
 &= x^3 + 3x^2d \\
 &= f(x) + f'(x) \cdot d
 \end{aligned}$$

## Infinitesimals

The derivative is used to solve the tangent problem. We can see that infinitesimals solve this too.

Examples: sine and cosine

Take a unit circle:

$$\begin{aligned}
 \sin x &= \sqrt{1 - \cos^2 x} \\
 \Rightarrow \sin d &= \sqrt{1 - \cos^2 d}
 \end{aligned}$$

Using and arc of angle  $\alpha$  and arc length  $l$

$$l = 2\pi R \cdot \frac{\alpha}{2\pi}$$

$$= R\alpha$$

$$\text{let } \alpha = d$$

$$\Rightarrow l = d$$

$\therefore$  We are moving on a straight line and instead of an arc it is a right angled triangle

$$\Rightarrow \sin d = d$$

$$\cos d = \sqrt{1 - \sin^2(d)}$$

$$= \sqrt{1 - d^2}$$

$$= 1$$

$$\Rightarrow \sin(0 + d) = \sin(0) + \sin'(0) \cdot (d)$$

$$\Rightarrow d = \sin'(0) \cdot d$$

We would want this to be 1

$$\cos d = \cos(0 + d)$$

$$= \cos(0) + \cos'(0) \cdot d$$

$$= 1 + \cos'(0) \cdot d$$

$$1 = 1 + \cos'(0) \cdot d$$

$$\cos'(0) \cdot d = 0$$

$$\cos'(0) = 0$$

## Limits

$$\cos'(x) = \lim_{h \rightarrow 0} \frac{\cos(x + h) - \cos(x)}{h}$$

$$\sin'(x) = \lim_{h \rightarrow 0} \frac{\sin(x + h) - \sin(x)}{h}$$