

# HPEC (High Performance Embedded Computing)

## Overview



A black and white photograph of Dr. Emmett Brown from the movie Back to the Future. He is an older man with wild, white hair, wearing a dark jacket over a patterned shirt. He is looking off to the side with a concerned expression. In the background is a chalkboard with handwritten text. The word "Past" is written on the left, and "1985" is written on the right, connected by a horizontal line. A vertical line also crosses the horizontal one. At the bottom left, the word "THERMUS" is partially visible, and at the bottom right, the word "BEEF" is partially visible.

"Roads? Where we're going,  
we don't need roads"

Dr. Emmett Brown

# But, meanwhile we have a deadly schedule

*What you'll have to do !*

6 Labs x 4H00



- ☐ 6 (or 7, or even 8 !) Labs supervised by your favorite teacher
- ☐ **You'll** work in pair
- ☐ No daily report
- ☐ Last Lab :
  - Final report (~10-15 pages) + source code
  - Demo/presentation with final report submission (~15 min)

# Goal of this course

*What you'll have to do !*

Implement a **math intensive application** on an embedded System-on-Chip (SoC). But not only...



This application will be **visual** and **interactive** !



GALAXEIRB

# High-level Galaxies Simulation Modeling

[http://en.wikipedia.org/wiki/N-body\\_simulation](http://en.wikipedia.org/wiki/N-body_simulation)

[http://www.kof.zcu.cz/st/dis/schwarzmeier/galaxy\\_interactions.html](http://www.kof.zcu.cz/st/dis/schwarzmeier/galaxy_interactions.html)



“Under the milky way tonight...”

The Church

# High-level Galaxies Simulation Modeling

Overview

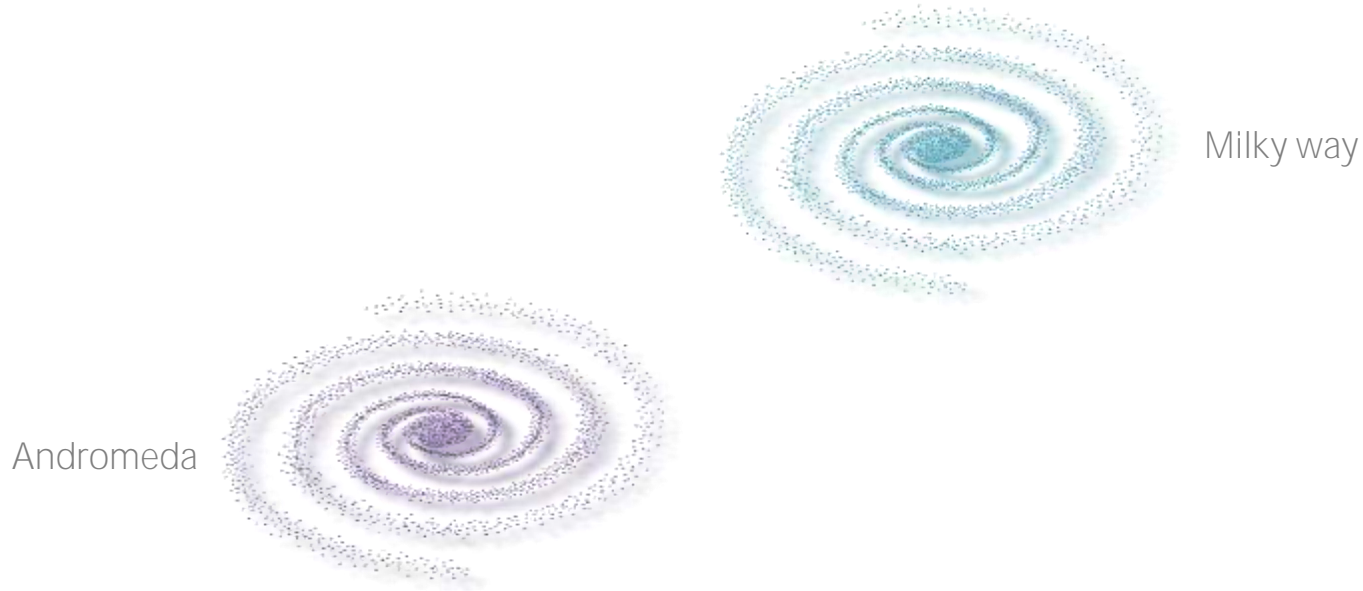


Figure. Two galaxies and their gravitational influences



# High-level Galaxies Simulation Modeling

Overview

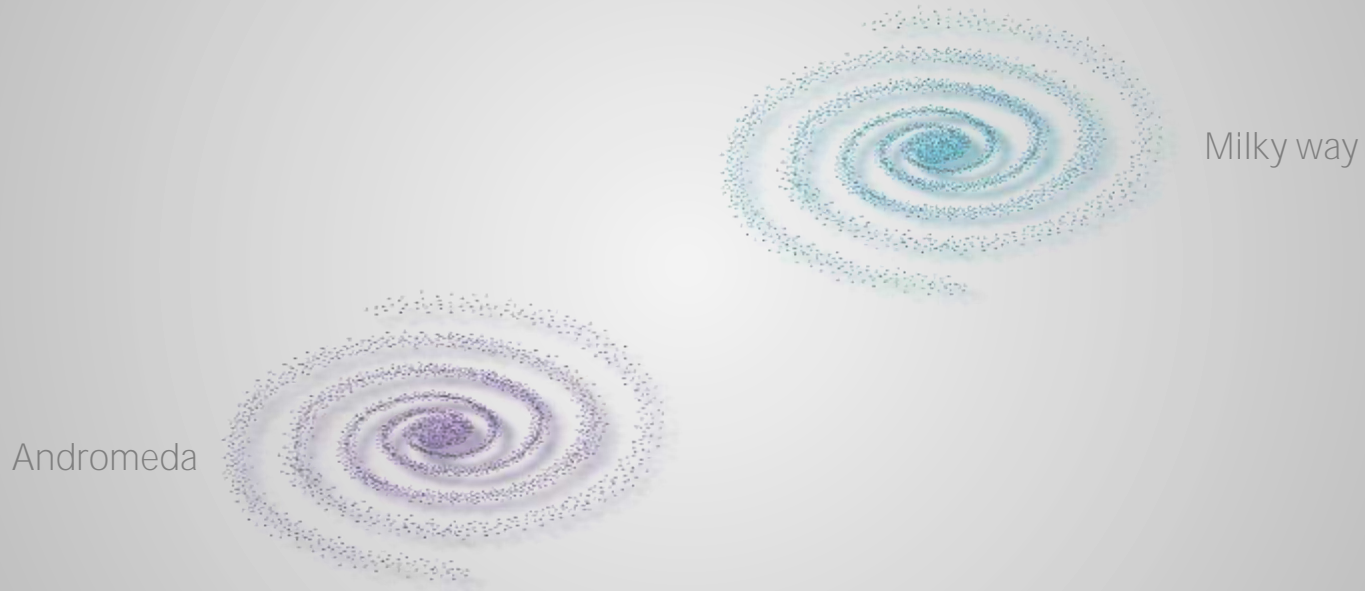


Figure. Two galaxies and their gravitational influences

# High-level Galaxies Simulation Modeling

Overview

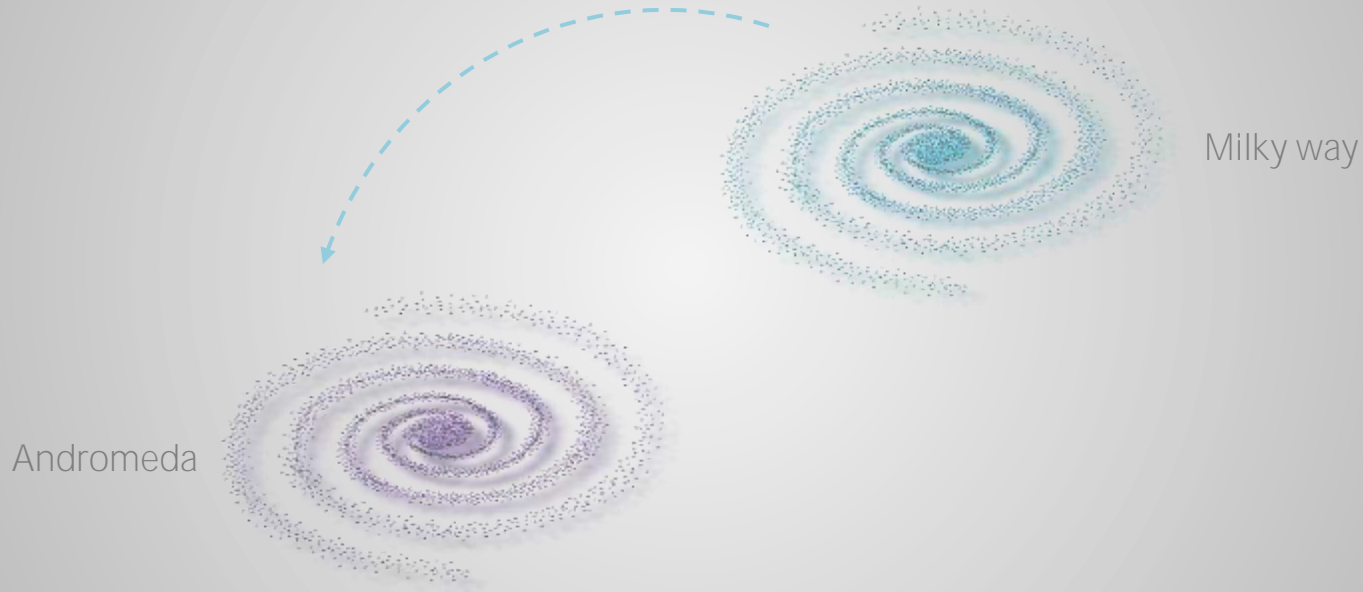


Figure. Two galaxies and their gravitational influences

# High-level Galaxies Simulation Modeling

Overview

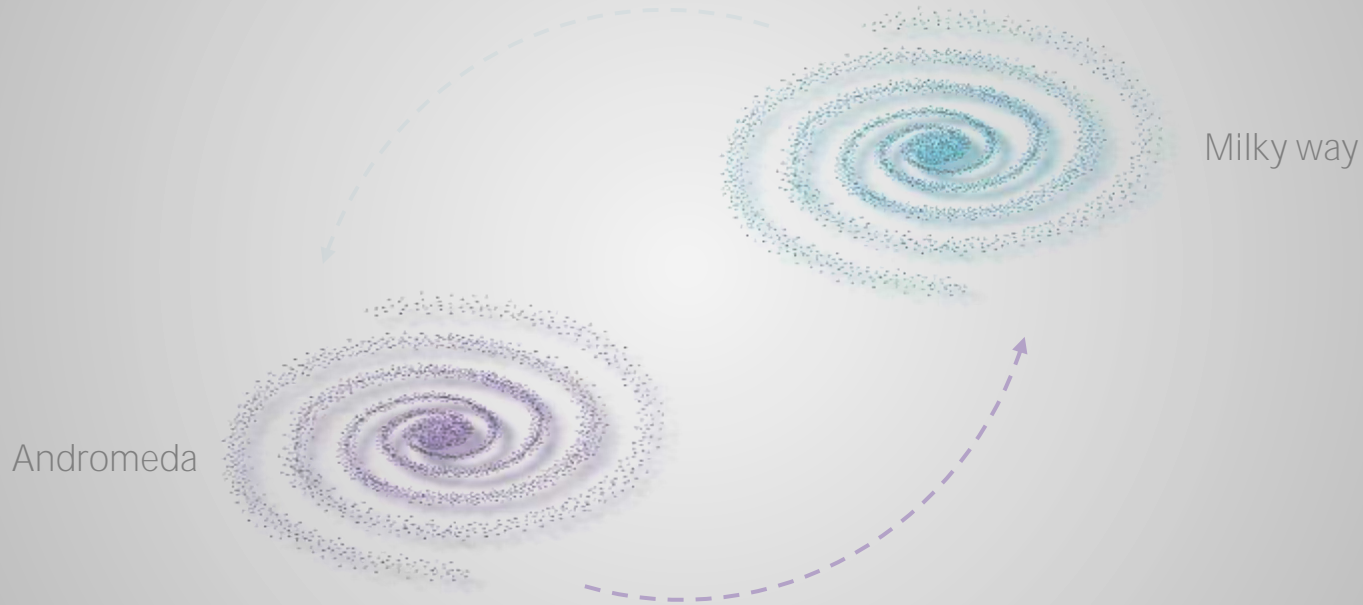


Figure. Two galaxies and their gravitational influences

# High-level Galaxies Simulation Modeling

*Spiral galaxies morphology*

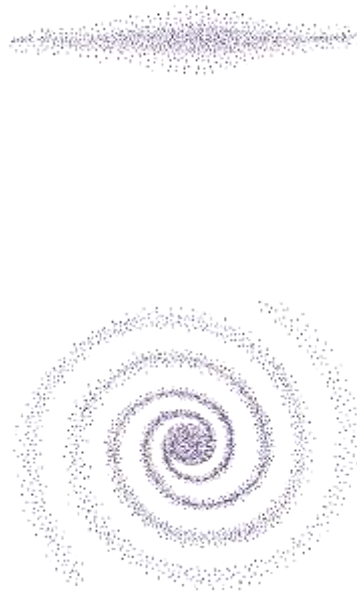


Figure. Face and side views of a disk galaxy

# High-level Galaxies Simulation Modeling

Spiral galaxies morphology

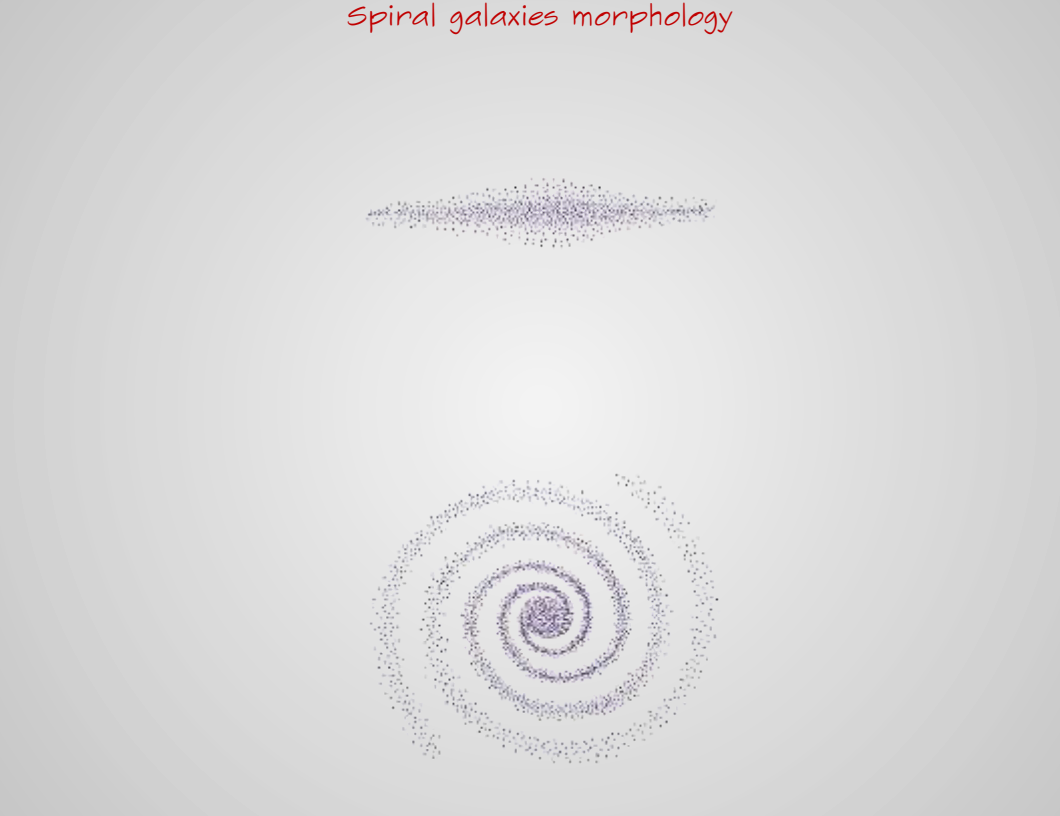


Figure. Face and side views of a disk galaxy

# High-level Galaxies Simulation Modeling

*Spiral galaxies morphology*

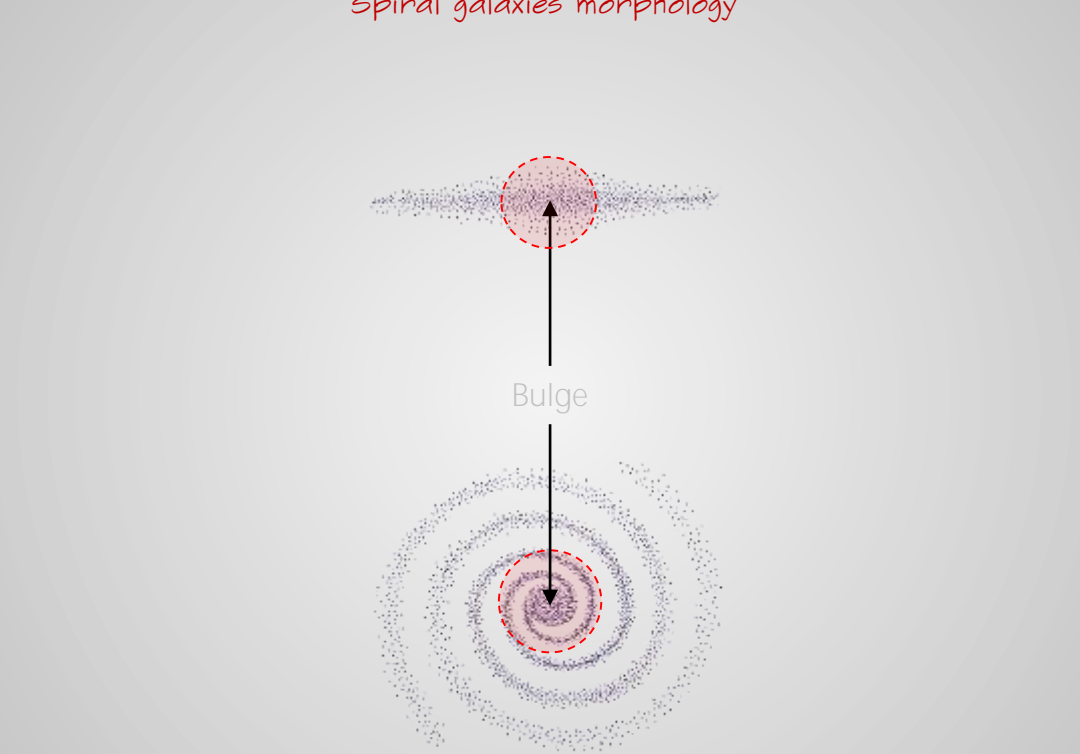


Figure. Face and side views of a disk galaxy

# High-level Galaxies Simulation Modeling

*Spiral galaxies morphology*

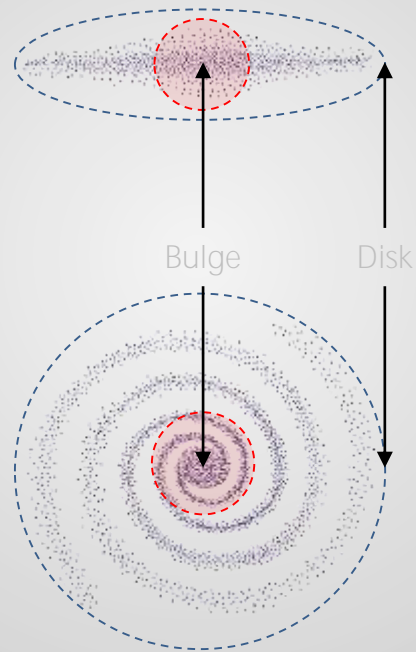


Figure. Face and side views of a disk galaxy

# High-level Galaxies Simulation Modeling

*Spiral galaxies morphology*

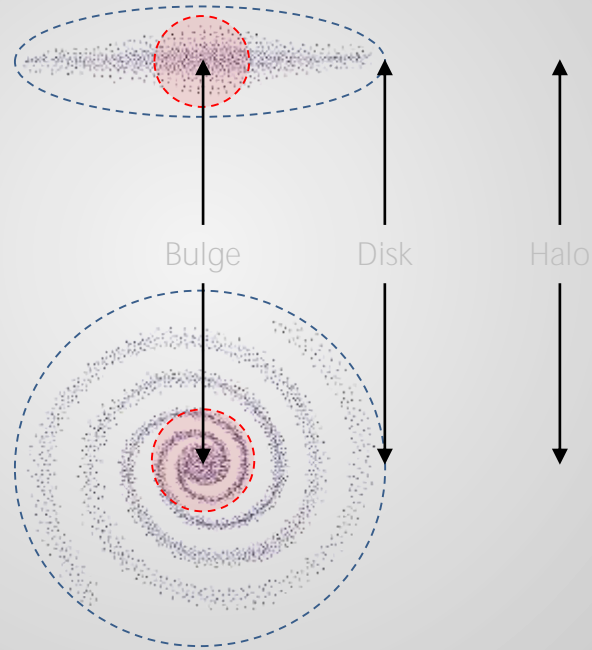


Figure. Face and side views of a disk galaxy



# Particle Modeling

# Particle Modeling

*Characterization and coordinate system*

Each celestial body is a particle characterized by its **position**, **velocity** and **mass**. To represent a particle in a three-dimensional (3D) space, a Cartesian coordinate system is conventionally used

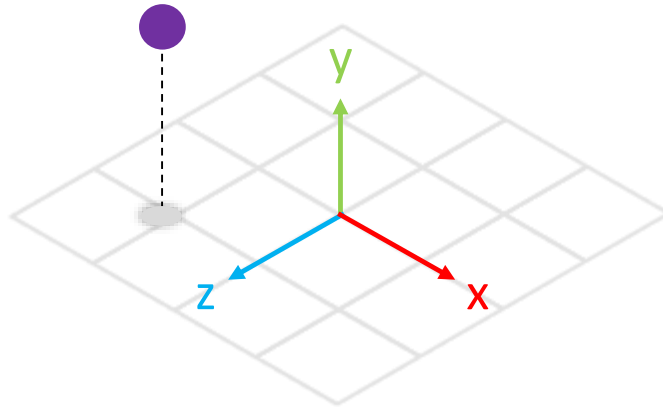


Figure. A three-dimensional space using a cartesian coordinate system, AKA « right-handed » system

# Particle Modeling

## Motion

The motion of a particle is done in a Cartesian coordinates system

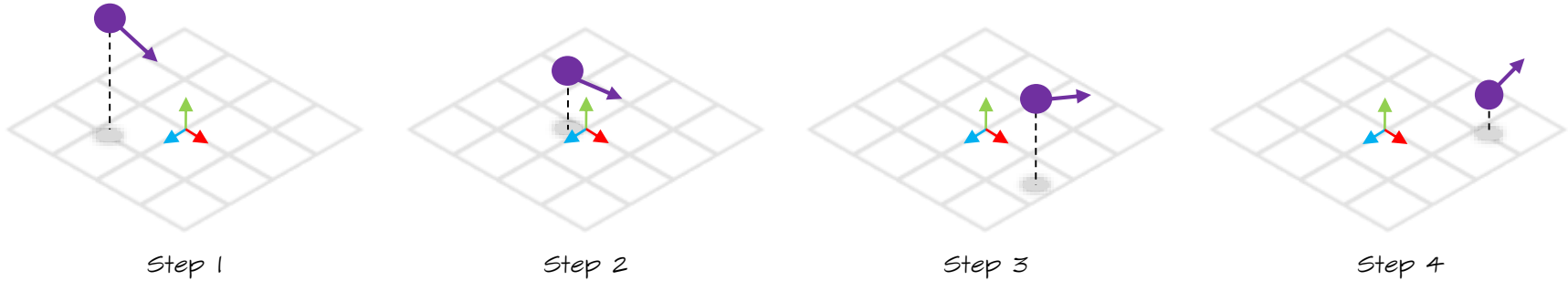


Figure. A particle is moving in a Cartesian system

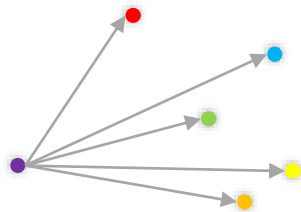
This motion is not only influenced by its own mass and velocity but also by positions and masses of other particles in the given “bodies” set  $S$

The position of a particle at time  $t+1$  is then computed by summing every particle contributions in set  $S$ . It is a “brute force” method.

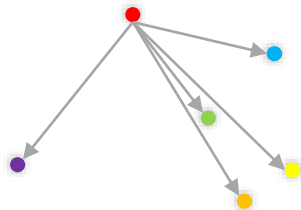
Although very accurate, this simulation model exhibits a time complexity of  $\mathcal{O}(n^2)$

# Particle Modeling

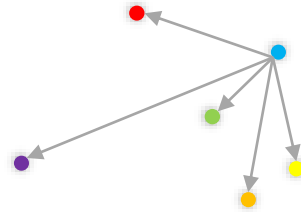
*Contribution and complexity*



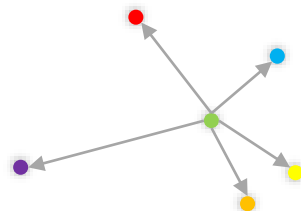
(a)



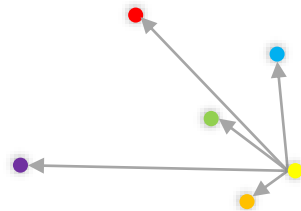
(b)



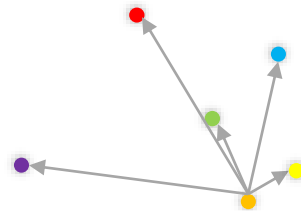
(c)



(d)



(e)



(f)

Figure. Five particles in purple, red, blue, green, yellow and orange

# Particle Modeling

*Contribution and complexity*

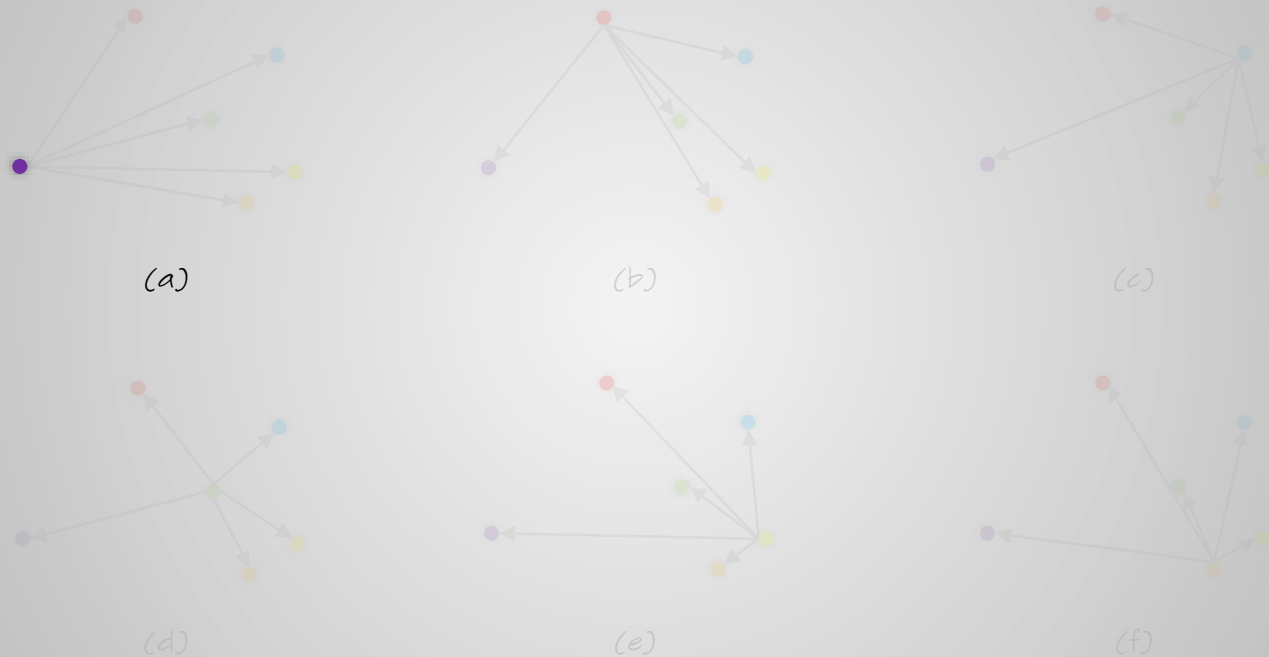


Figure. Five particles in purple, red, blue, green, yellow and orange

# Particle Modeling

*Contribution and complexity*

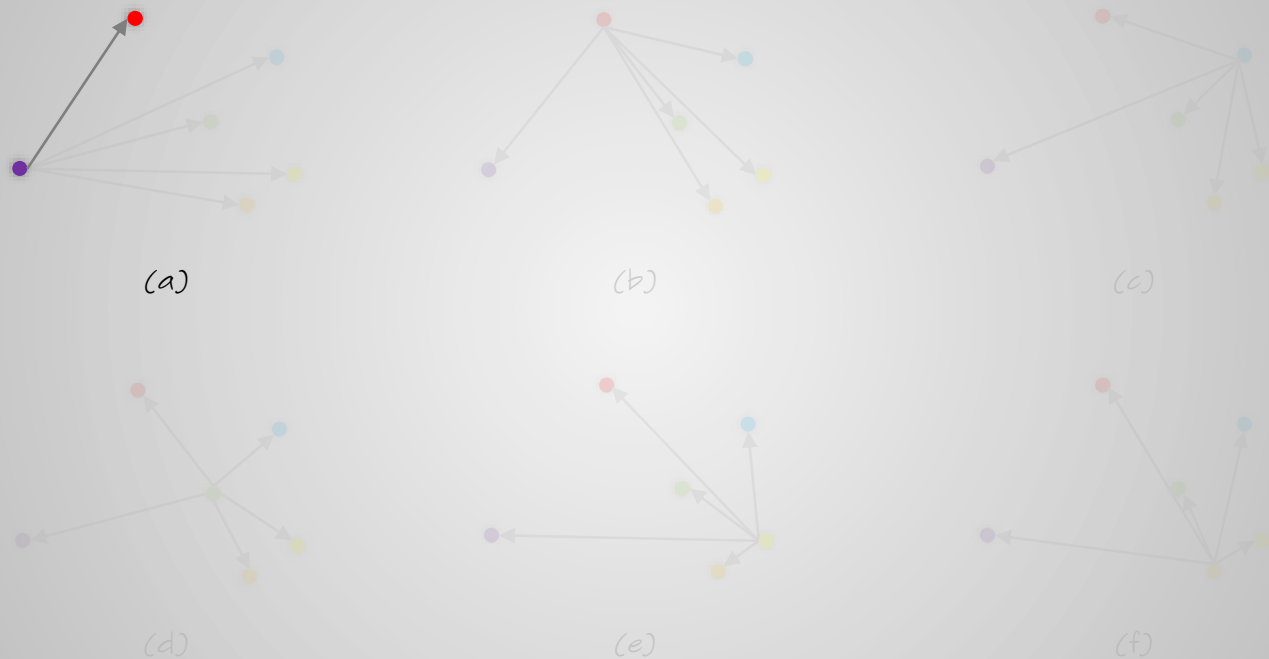


Figure. Five particles in purple, red, blue, green, yellow and orange

# Particle Modeling

*Contribution and complexity*

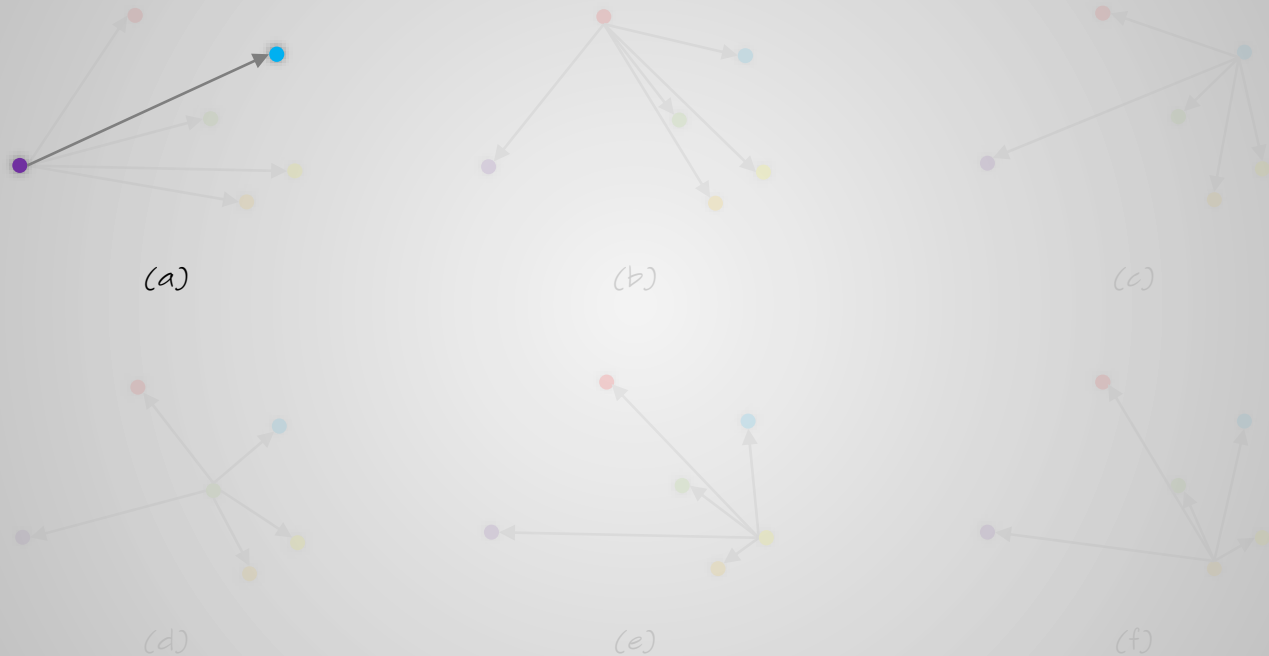


Figure. Five particles in purple, red, blue, green, yellow and orange

# Particle Modeling

*Contribution and complexity*

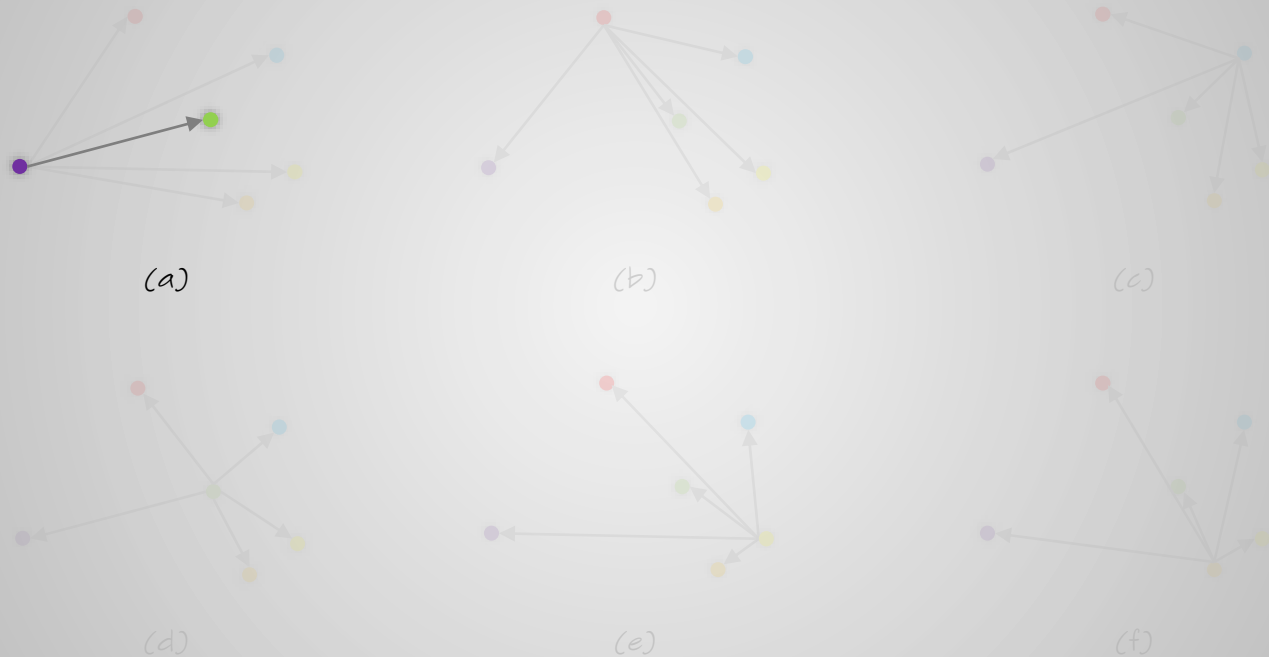


Figure. Five particles in purple, red, blue, green, yellow and orange



# Particle Modeling

*Contribution and complexity*



(a)



(b)



(c)



(d)



(e)



(f)

Figure. Five particles in purple, red, blue, green, yellow and orange

# Particle Modeling

*Contribution and complexity*

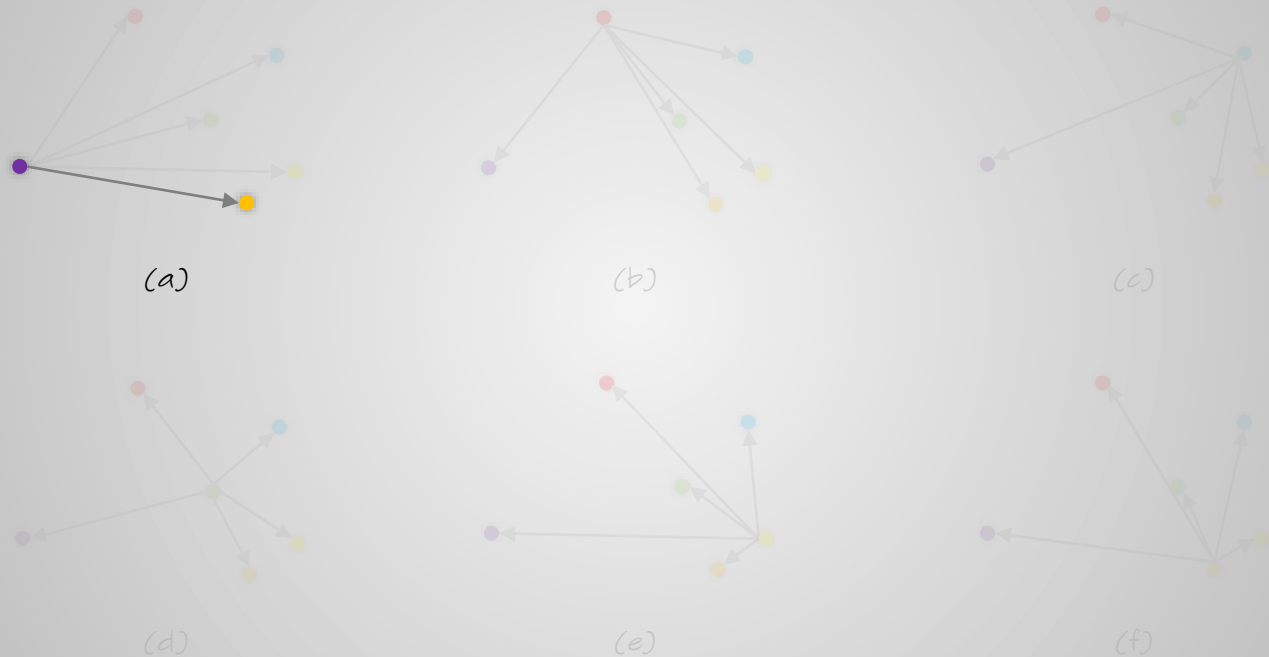


Figure. Five particles in purple, red, blue, green, yellow and orange

# Particle Modeling

## Damping factor

The damping factor  $\zeta$  influences a particle velocity over time.

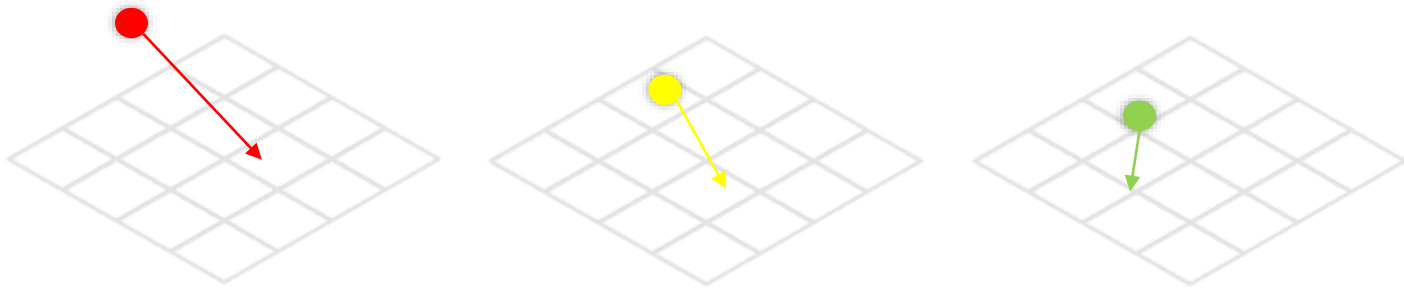


Figure. Step by step influence of the damping factor  $\zeta$  on one particle

A  $\zeta$  factor  $< 1.0$  decreases a particle velocity while a  $\zeta$  factor  $> 1.0$  increases a particle velocity.

# Particle Modeling

*Simulation algorithm*

```
1  FOR each step DO
2    FOR each particle DO
3      particle acceleration  $\leftarrow$  0
4      FOR each neighbor particle DO
5        particle acceleration  $\leftarrow$  Add Acceleration ( particle, neighbor particle )
6      END
7    END
8    FOR each particle DO
9      Update Particle Positions ( particle, particle acceleration )
10   END
11   Do Anything Useful ( )
12 END
```

*Algorithm. Pseudo-code simulation algorithm*

# Particle Modeling

## Parameters

• $n$	number of particles	scalar
• $\vec{p}_i, \vec{p}_j$	position vectors	vector3
• $\vec{\Delta}_{ij}$	slope vector between particles i and j	vector3
• $d_{ij}$	distance between particles i and j	scalar
• $\vec{a}_i$	acceleration vector of particle i	vector3
• $m_i, m_j$	particles masses	scalar
• $M$	mass factor	scalar
• $\zeta$	particles damping factor	scalar

# Particle Modeling

Finding a particle acceleration

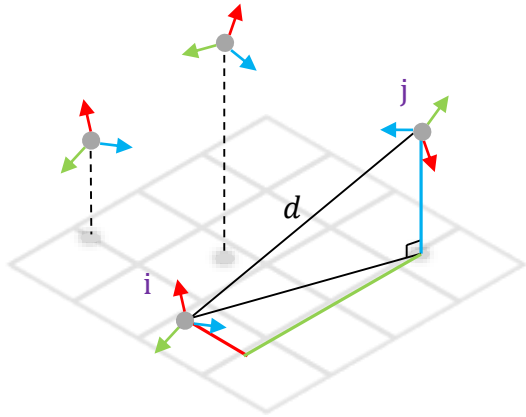


Figure. Two particles  $i, j$   
and their Euclidean  
distance  $d$

- $\vec{p}_i = (x_i, y_i, z_i)$
- $\vec{p}_j = (x_j, y_j, z_j)$
- $\vec{\Delta}_{ij} = \vec{p}_j - \vec{p}_i$        $\vec{\Delta}_{ij} = \begin{pmatrix} x_j - x_i \\ y_j - y_i \\ z_j - z_i \end{pmatrix}$
- $d_{ij} = \sqrt{(x_j - x_i)^2 + (y_j - y_i)^2 + (z_j - z_i)^2}$
- $\vec{a}_i = \sum_{j \neq i, j=0}^{n-1} \vec{\Delta}_{ij} \times M \times \zeta \times \frac{1}{d_{ij}^3} \times m_j$

Implementation

# Targeted hardware

*NVIDIA Jetson TK1 development board*

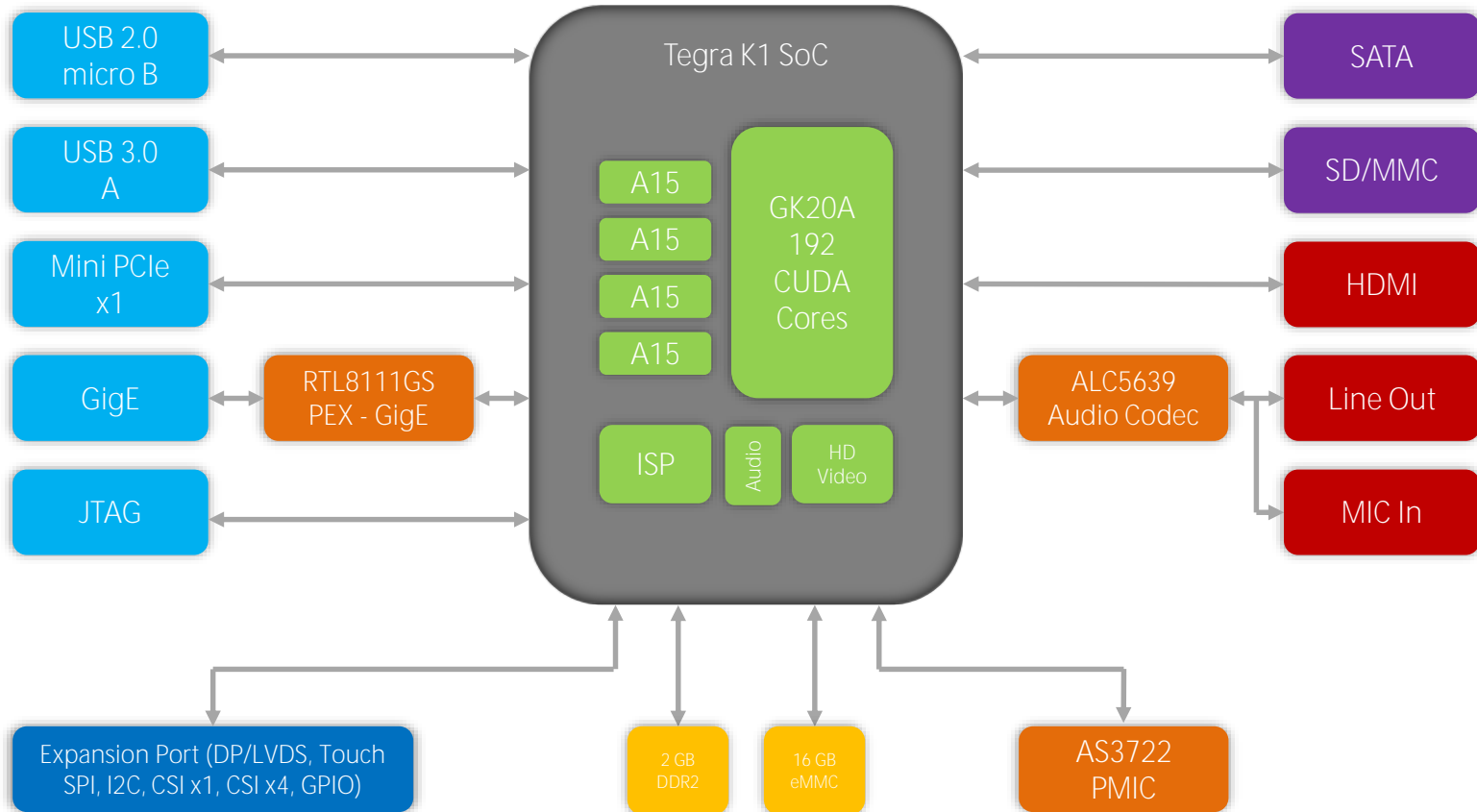


<http://www.nvidia.com/object/tegra-k1-processor.html>

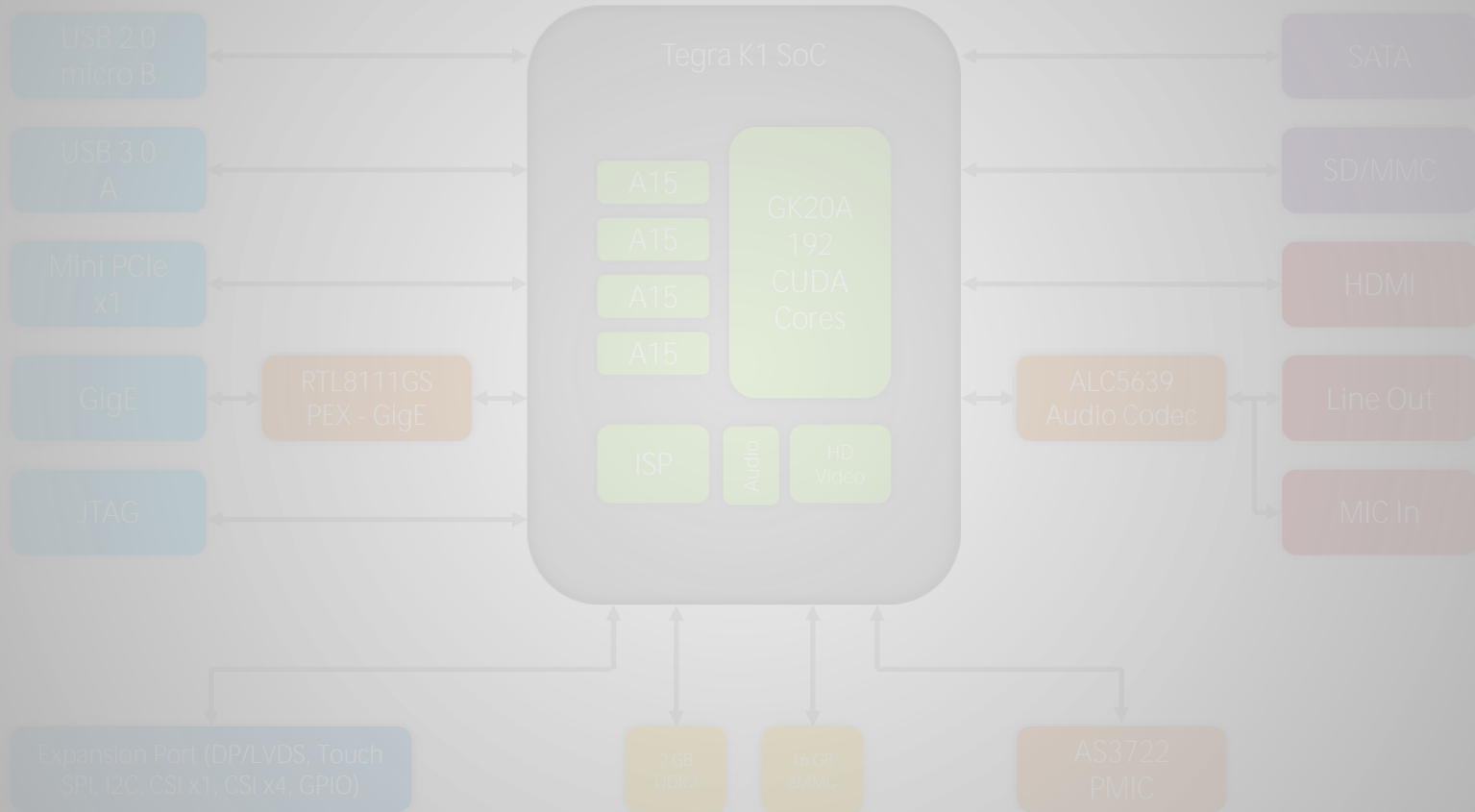
<http://www.hardware.fr/focus/imprimer/94/>



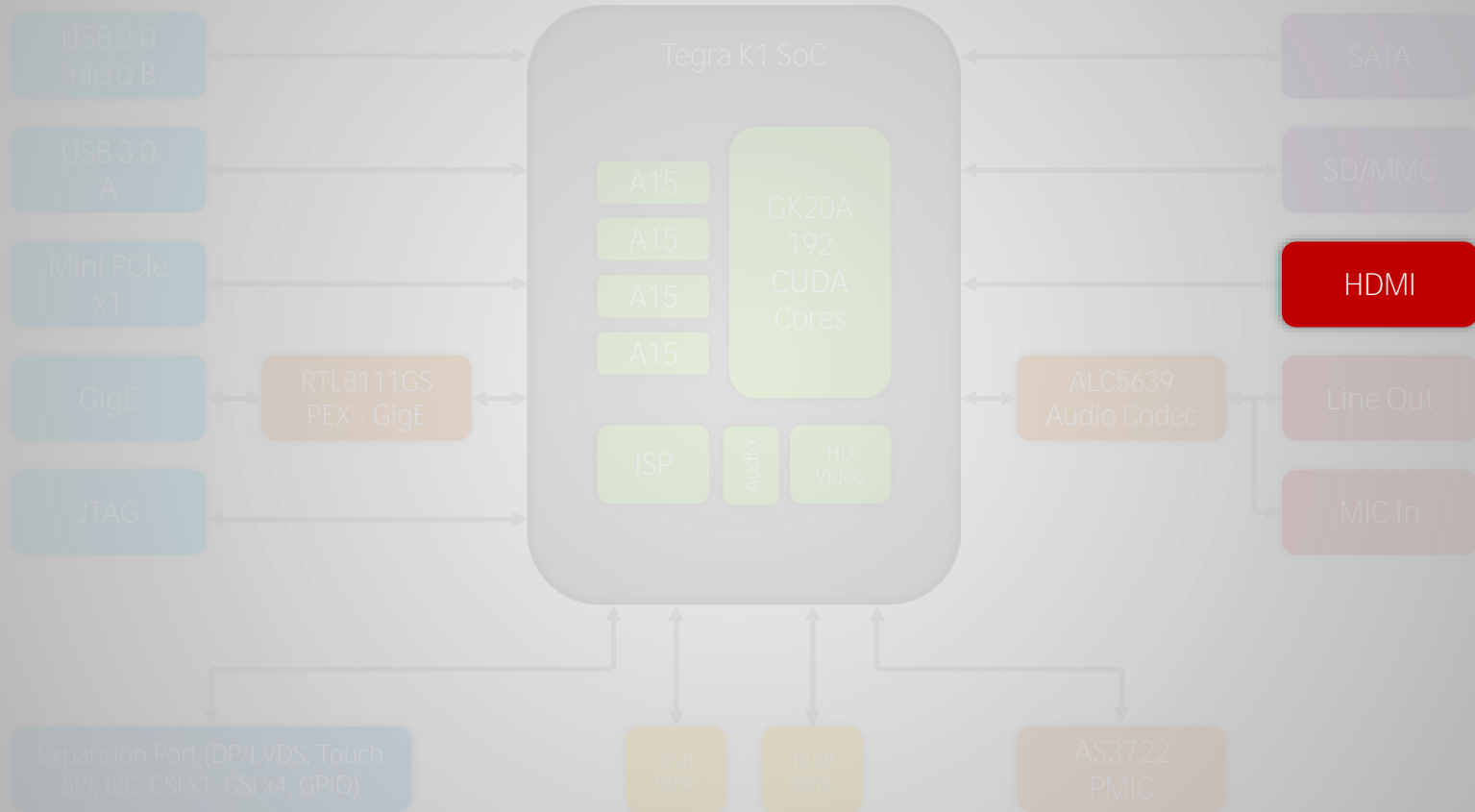
# NVIDIA Jetson TK1 development board



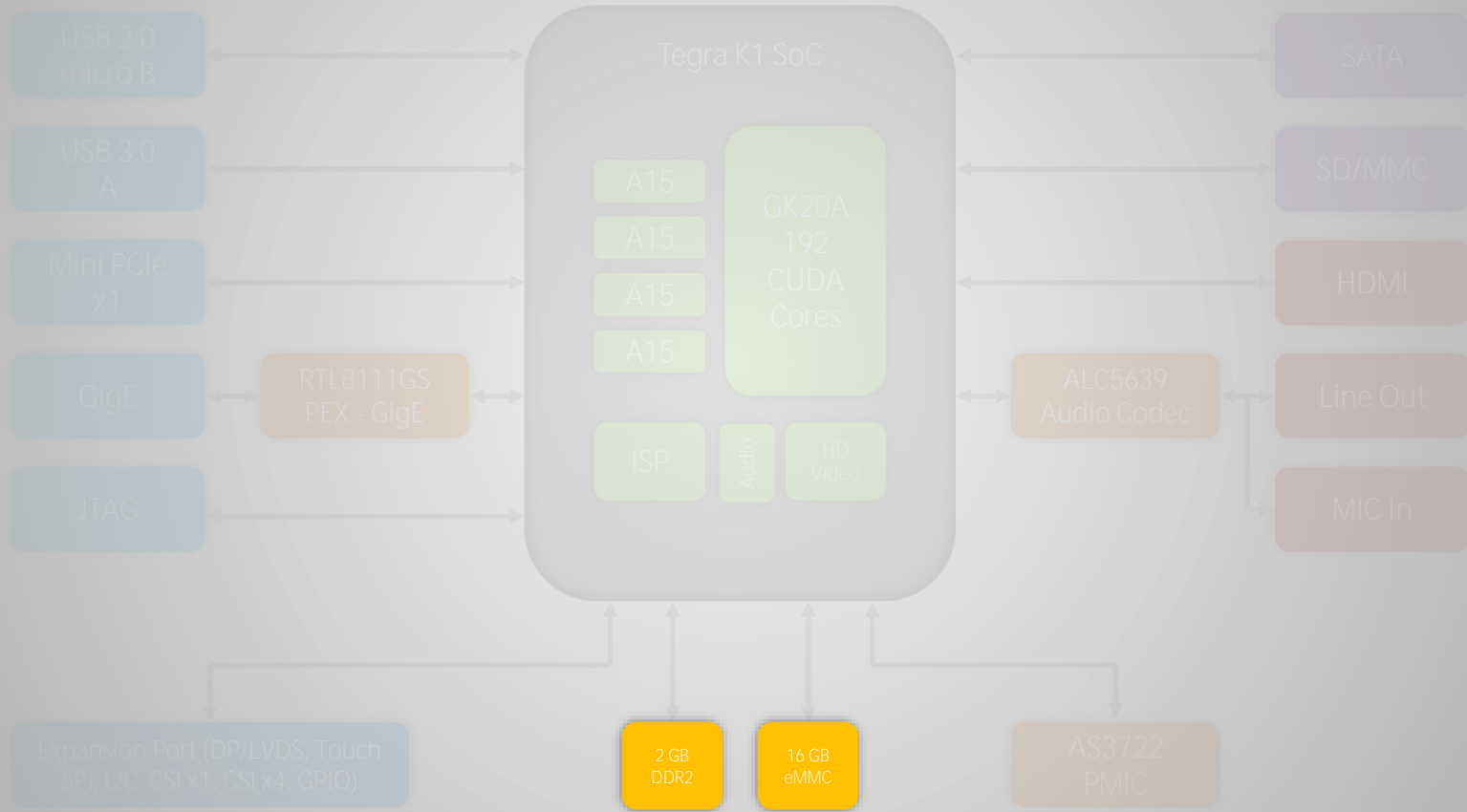
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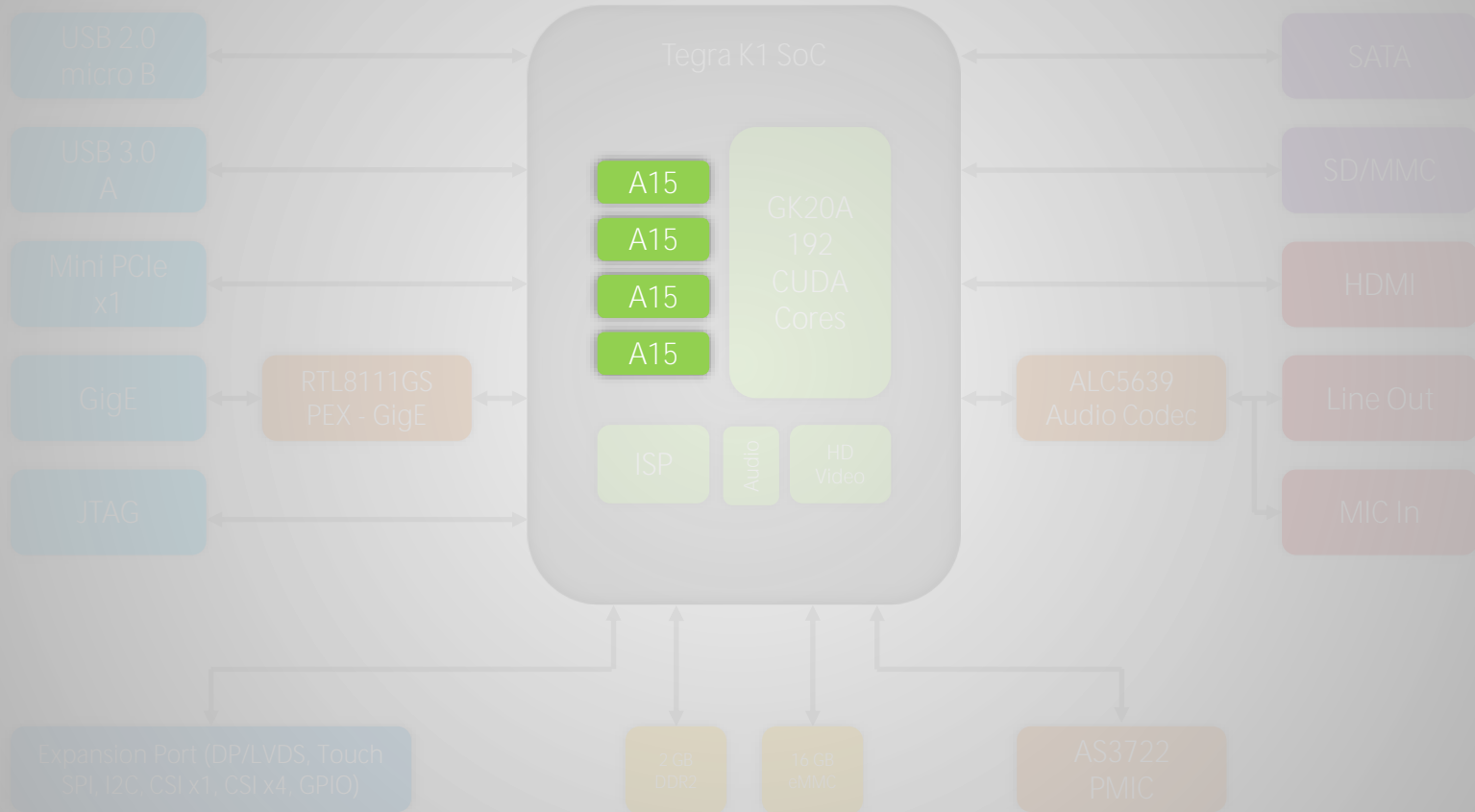
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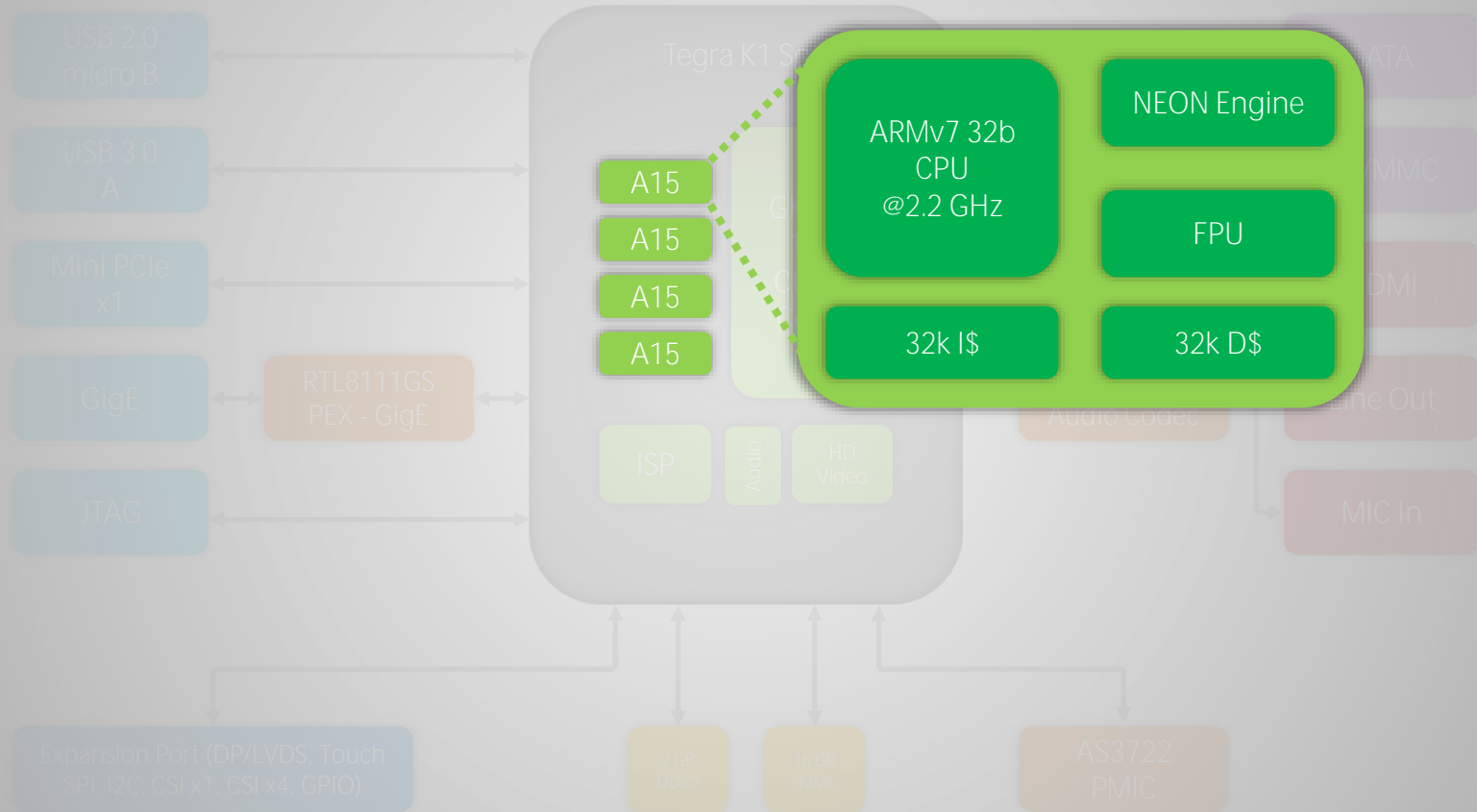
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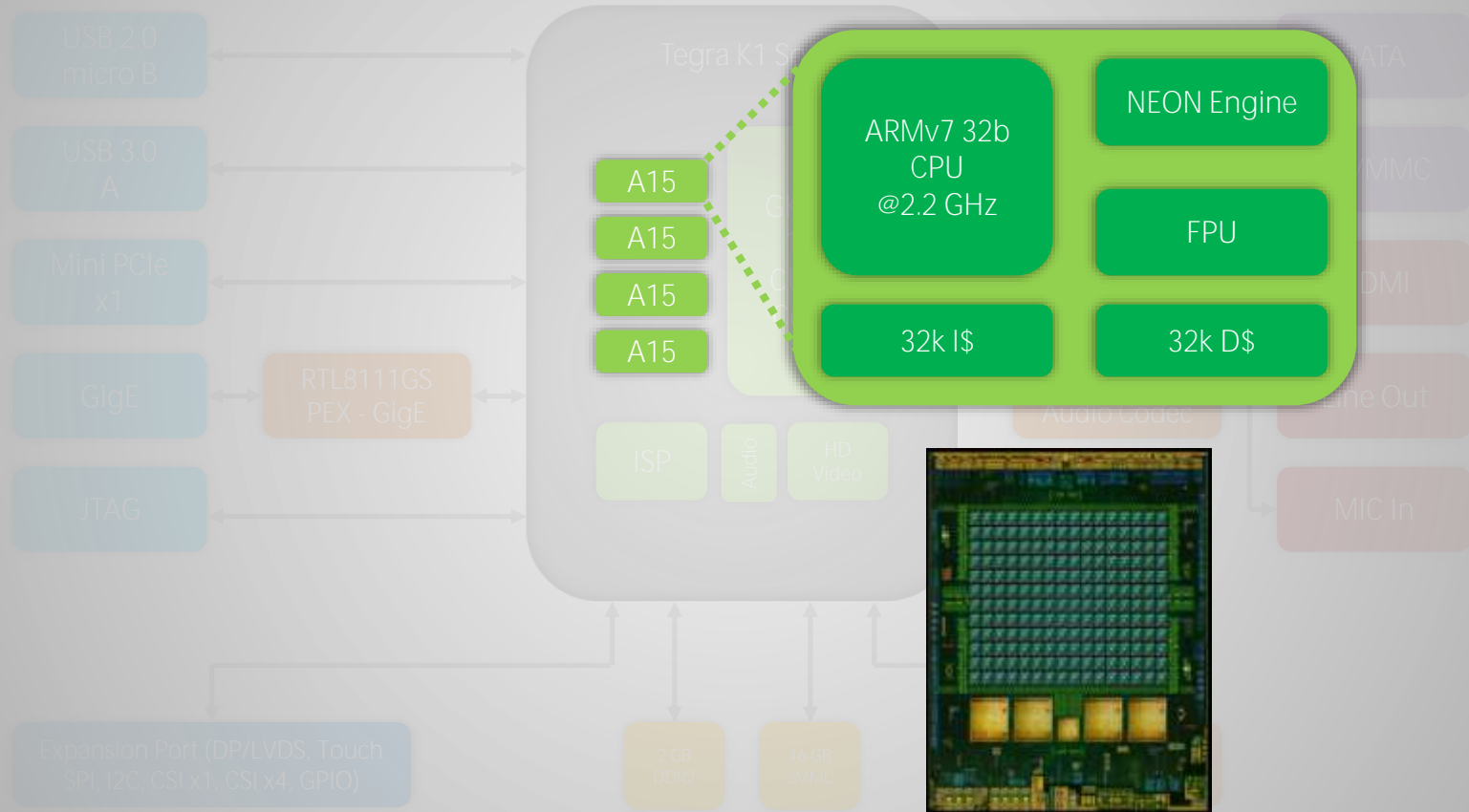
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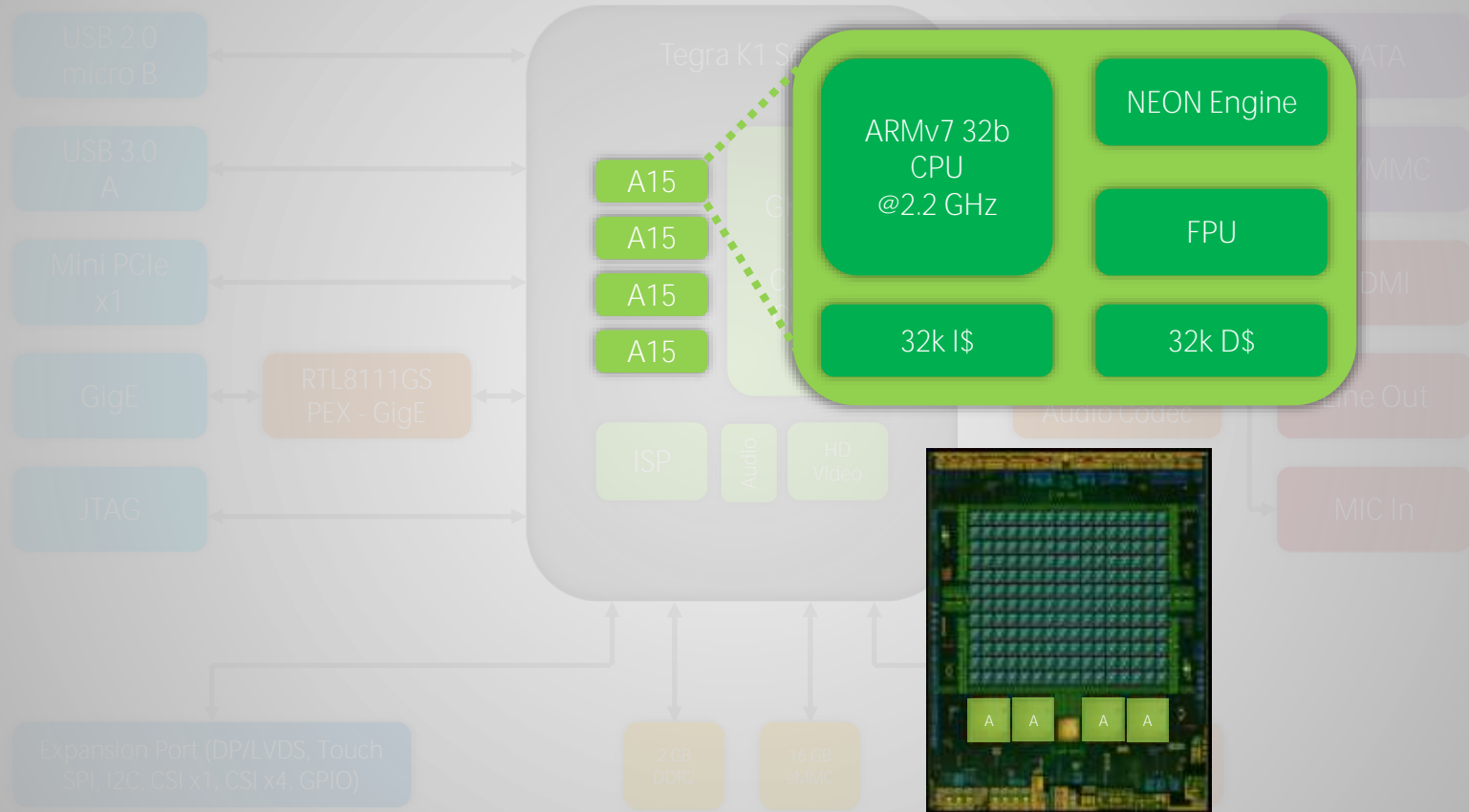
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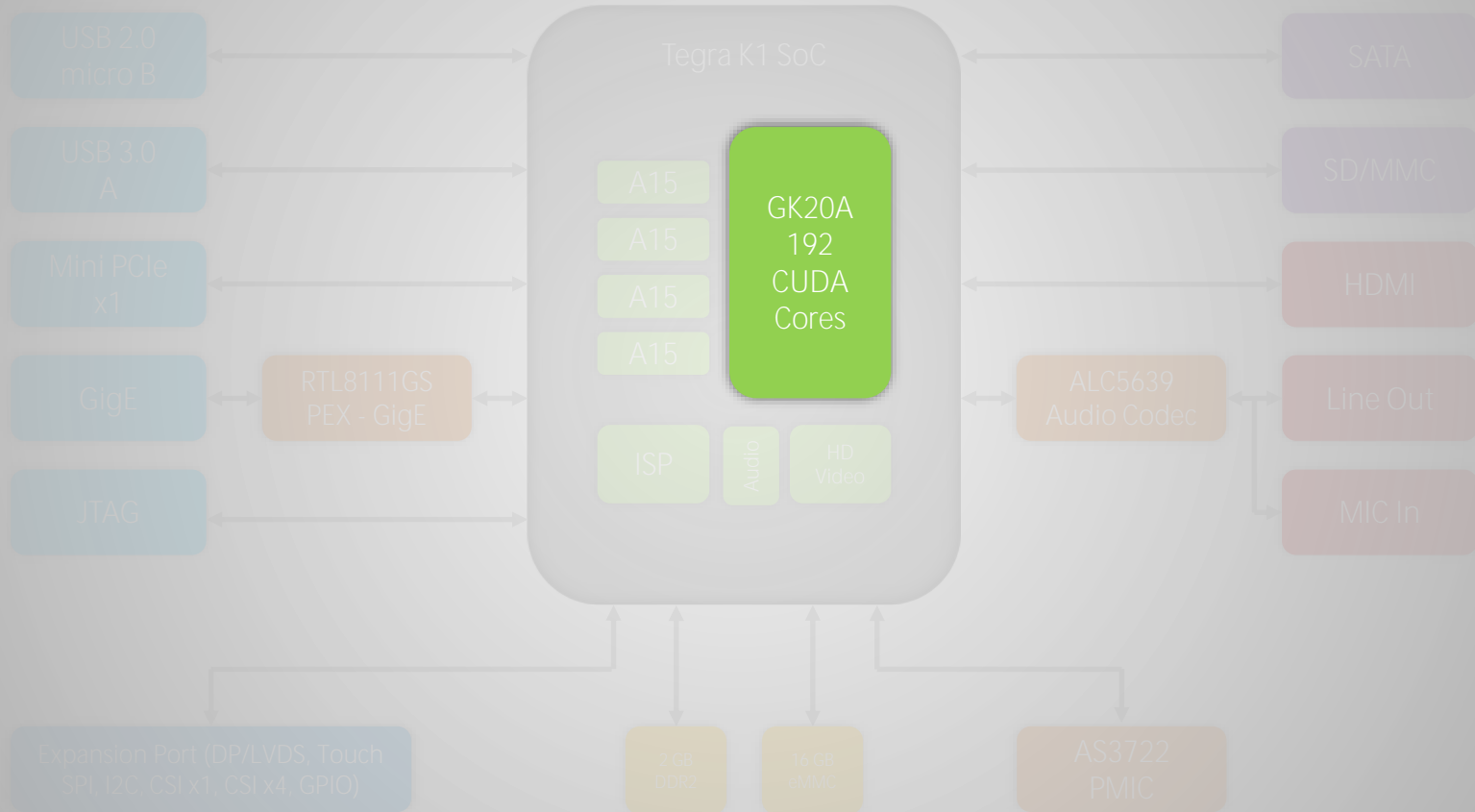


# NVIDIA Jetson TK1 development board

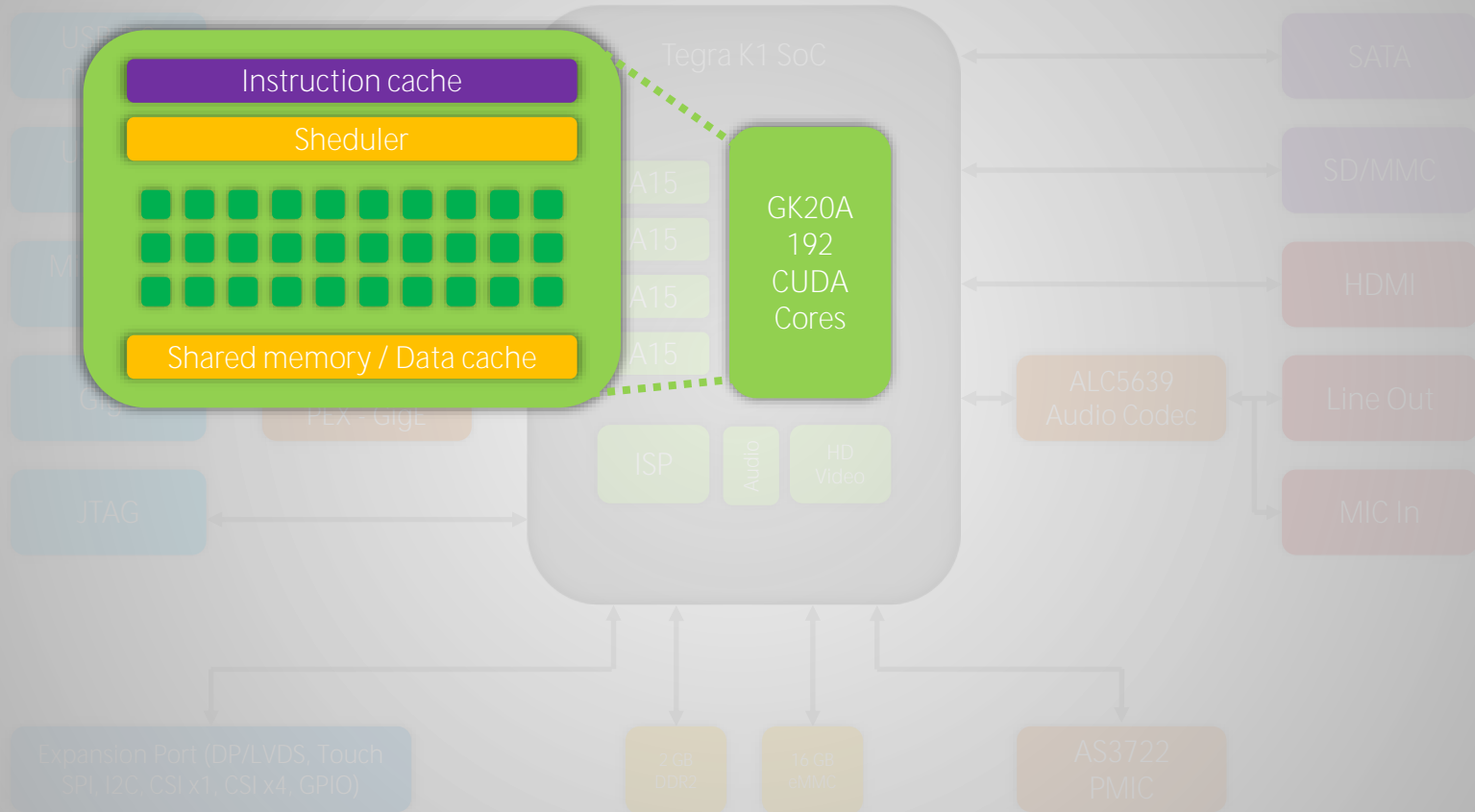




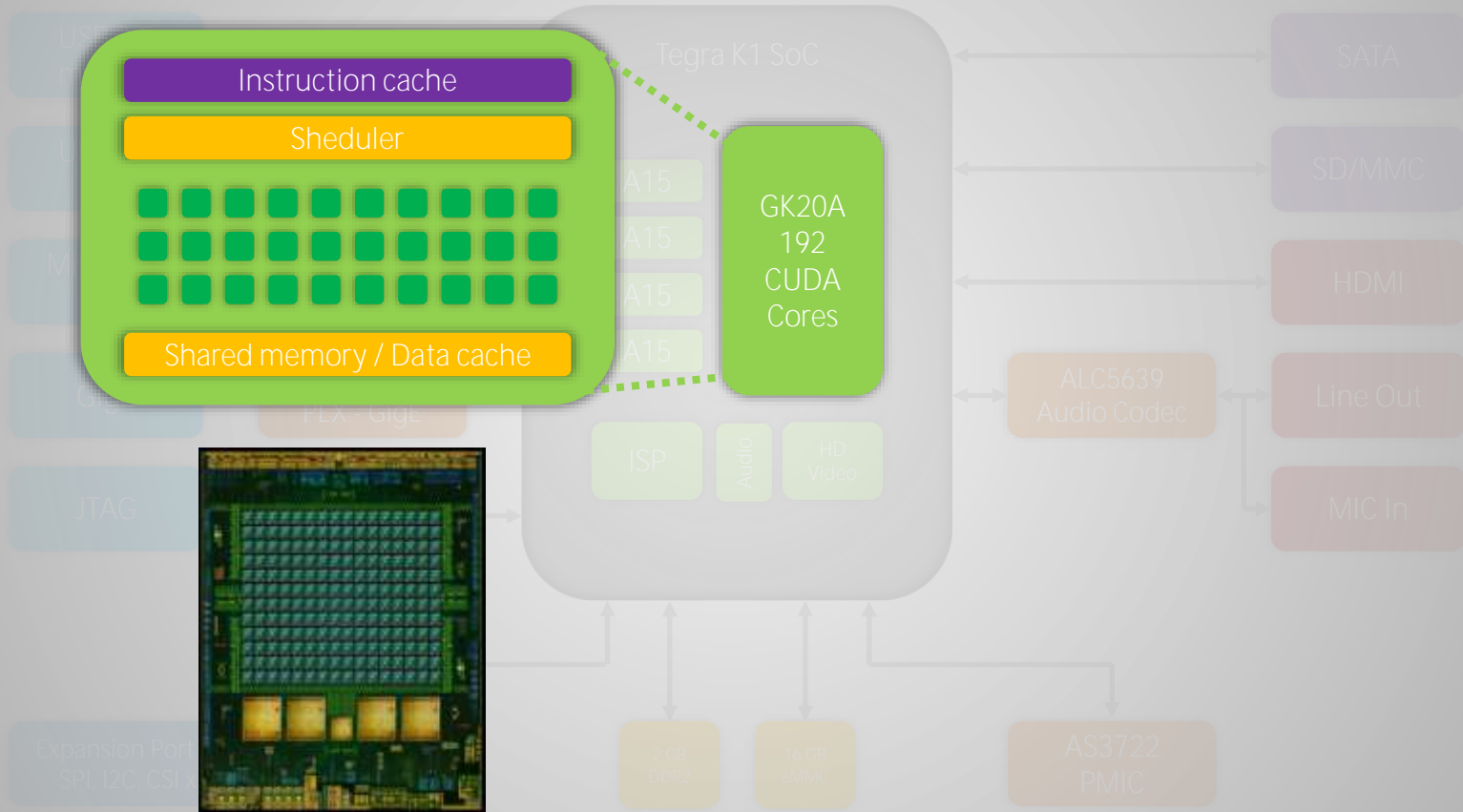
# NVIDIA Jetson TK1 development board



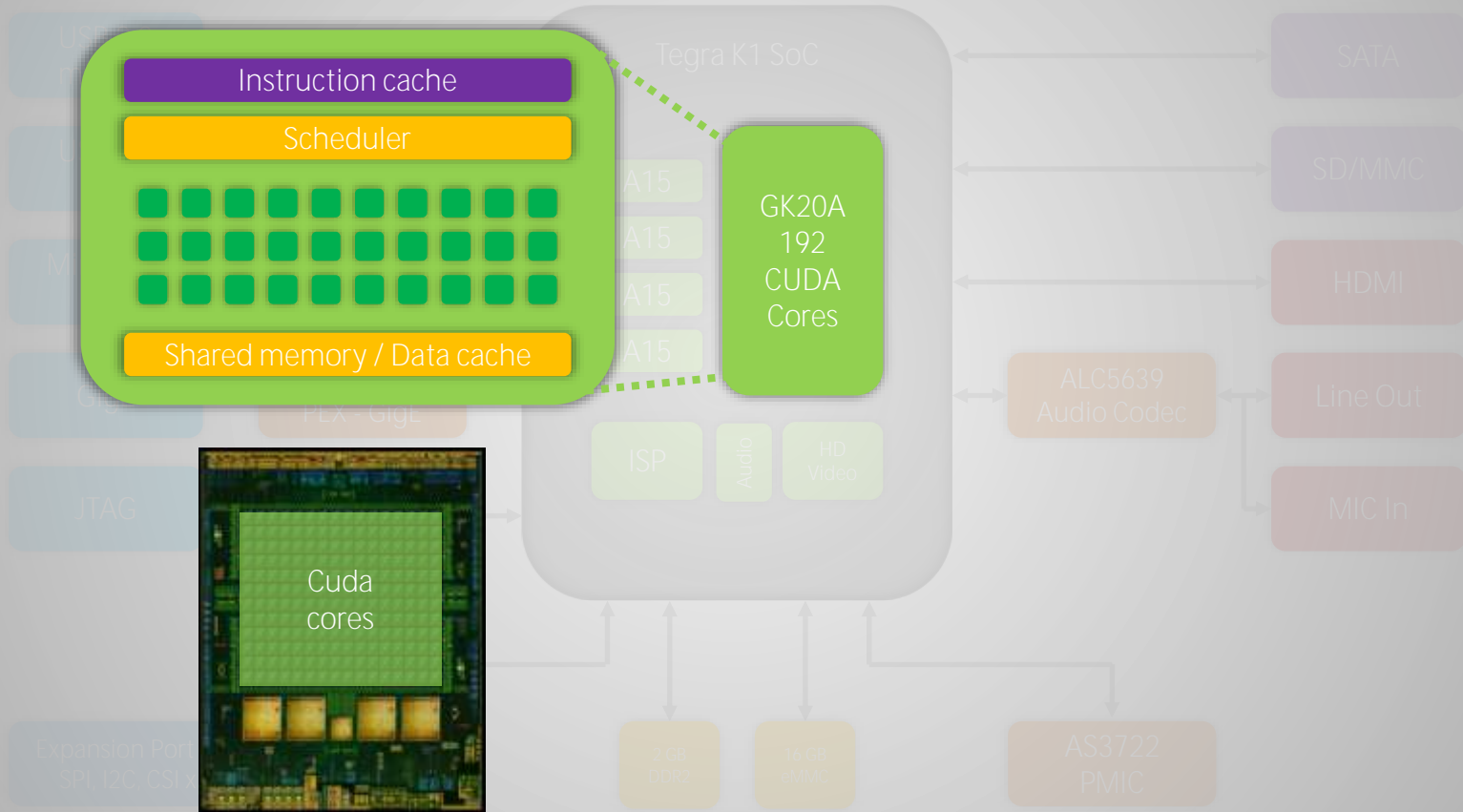
# NVIDIA Jetson TK1 development board



# NVIDIA Jetson TK1 development board



# NVIDIA Jetson TK1 development board



# NVIDIA Jetson TK1 development board

*operating system Ubuntu 14.04 LTS*



<https://developer.nvidia.com/linux-tegra-rel-19>

# NVIDIA Jetson TK1 development board

*Useful compatible libraries*

The Open graphics library  
GPU.



n API for rendering 2D and 3D vector graphics using a

The OpenGL extension wrangler library GLEW  
using extensions.



ables OpenGL latest graphics capabilities



or Simple Directmedia Layer is also a library which provides low access to windows,  
keyboard and mouse.

The Open Multi-Processing API



CPUs parallel architectures to unleash relevant  
application speed-up. To achieve further acceleration, NVIDIA Compute Unified Device Architecture  
makes uses of NV GPUs.



# NVIDIA Jetson TK1 development board

## Useful links

- ❑ <https://www.opengl.org/>
- ❑ <http://glew.sourceforge.net/>
- ❑ <https://www.libsdl.org/>
- ❑ <http://openmp.org/wp/>
- ❑ <http://www.nvidia.fr/object/cuda-parallel-computing-fr.html>
- ❑ <https://computing.llnl.gov/tutorials/pthreads/>

A close-up of a bearded man (Leonidas) with a determined, intense expression, shouting or yelling. He is wearing a dark, rugged garment. In the background, a woman in a white dress stands in a classical setting with columns.

“Madness ?!...”

Leonidas



Where to start ?

Where to start ?

*One small step for man...*



Boot your board and log in (SE or SEE)  
Please take care of it !

# Where to start ?

*Inside the ./Documents/examples directory*



01\_helloworld



02\_sdl



03\_opengl



04\_pthread



05\_openmp

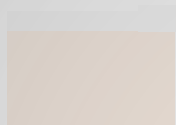


06\_cuda

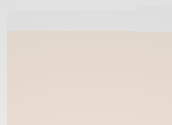
A set of 6 basic examples to help you getting started

# Where to start ?

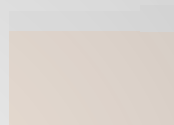
*Inside the ./Documents/examples directory*



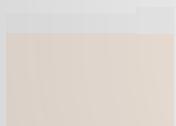
01\_helloworld



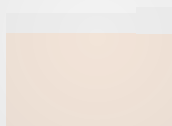
02\_sdl



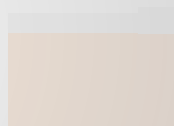
03\_opengl



04\_pthread



05\_openmp



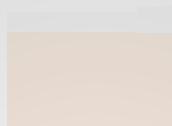
06\_cuda

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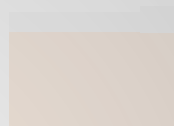
*Inside the ./Documents/examples directory*



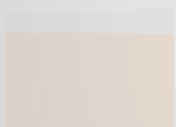
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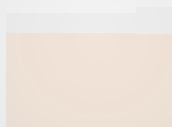
02\_sdl



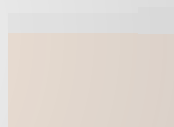
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04\_pthread



05\_openmp

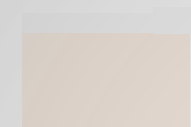


06\_cuda

```
> $. cd 01_helloworld  
> $. make all  
> $. ./bin/helloworld
```

# Where to start ?

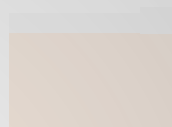
*Inside the ./Documents/examples directory*



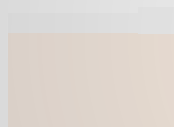
01\_helloworld



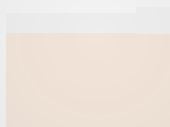
02\_sdl



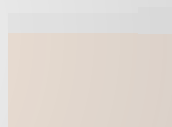
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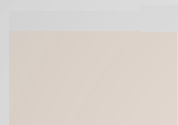


06\_cuda

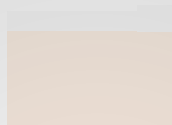
```
> $. cd 02_sdl
> $. make all
> $. ./bin/sdl
```

# Where to start ?

*Inside the ./Documents/examples directory*



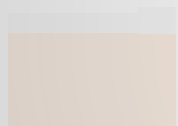
01\_helloworld



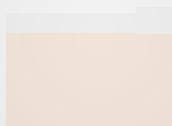
02\_sdl



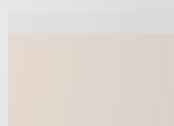
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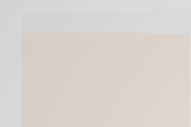


06\_cuda

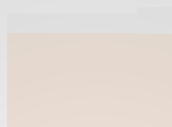
```
> $: cd 03_opengl  
> $: make all  
> $: ./bin/opengl
```

# Where to start ?

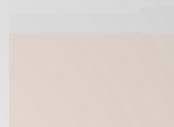
*Inside the ./Documents/examples directory*



01\_helloworld



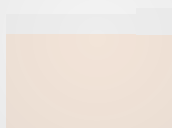
02\_sdl



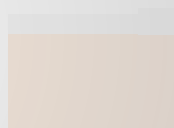
03\_opengl



04\_pthread



05\_openmp



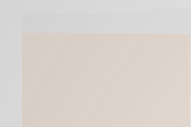
06\_cuda

```
> $: cd 04_pthread  
> $: make all  
> $: ./bin/pthread
```

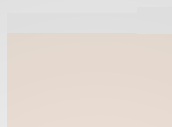


# Where to start ?

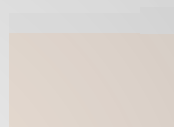
*Inside the ./Documents/examples directory*



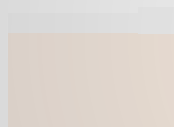
01\_helloworld



02\_sdl



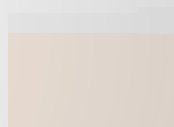
03\_opengl



04\_pthread



05\_openmp

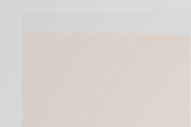


06\_cuda

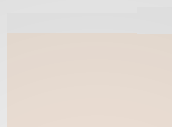
```
> $: cd 05_openmp
> $: make all
> $: ./bin/openmp
```

# Where to start ?

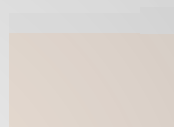
*Inside the ./Documents/examples directory*



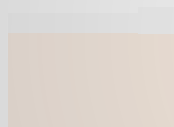
01\_helloworld



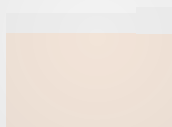
02\_sdl



03\_opengl



04\_pthread



05\_openmp



06\_cuda

```
> $: cd 06_cuda  
> $: make all  
> $: ./bin/cuda
```

# Where to start ?

*Milky way and andromeda galaxies data set*

Go to this page :

<http://bima.astro.umd.edu/nemo/archive/#dubinski>

and download John **Dubinski's** archive :

[dubinski.tab.gz](http://dubinski.tab.gz)



John Dubinski

# Where to start ?

John Dubinski's file format

Num. of particles	Total particles	Galaxy	Element
16384	16384	Milky way	disk
16384	32768	Andromeda	disk
8192	40960	Milky way	bulge
8192	49152	Andromeda	bulge
16384	65536	Milky way	halo
16384	81920	Andromeda	halo

The text file describes each particle line by line. Each line is a string that follows this format, where space is used as delimiter :

Mass PositionX PositionY PositionZ VelocityX VelocityY VelocityZ

Video time



“Just one more thing...”

Columbo



Just one more thing...

Q & A

Q: We need to open a new window and we want to use my mouse and my keyboard to play with. How do I do that ?

A: Use SDL, it's not the only available library to do it, but at least it's a good one





Just one more thing...

Q & A

Q: How do we visualize our particles in a 3D space ?

A: OpenGL is the key. There is a basic example for manipulating 3D in your ./Documents/examples directory



Just one more thing...

Q & A

Q: OK, We're able to draw points in 3D, now what ?

A : See if everything is correct.  
Download John's Dubinsky data set  
and see what's happening. If it looks  
good, then you can start playing with  
maths :)



Just one more thing...

Q & A

Q: Our simulation is working but is slow as a Romero's zombie, is it normal ?

A : Yes, it is. But you can use OpenMP or/and CUDA to speed-up things



Just one more thing...

Q & A

Q: What if we succeed in this project ?

A : Well, then you could say ...





*"We came, we saw, we kicked  
its ass!"*

*Dr. Peter Venkman*