

# COMP498G/691G COMPUTER VISION

## LECTURE 17 PHOTOMETRIC STEREO



# Administrative

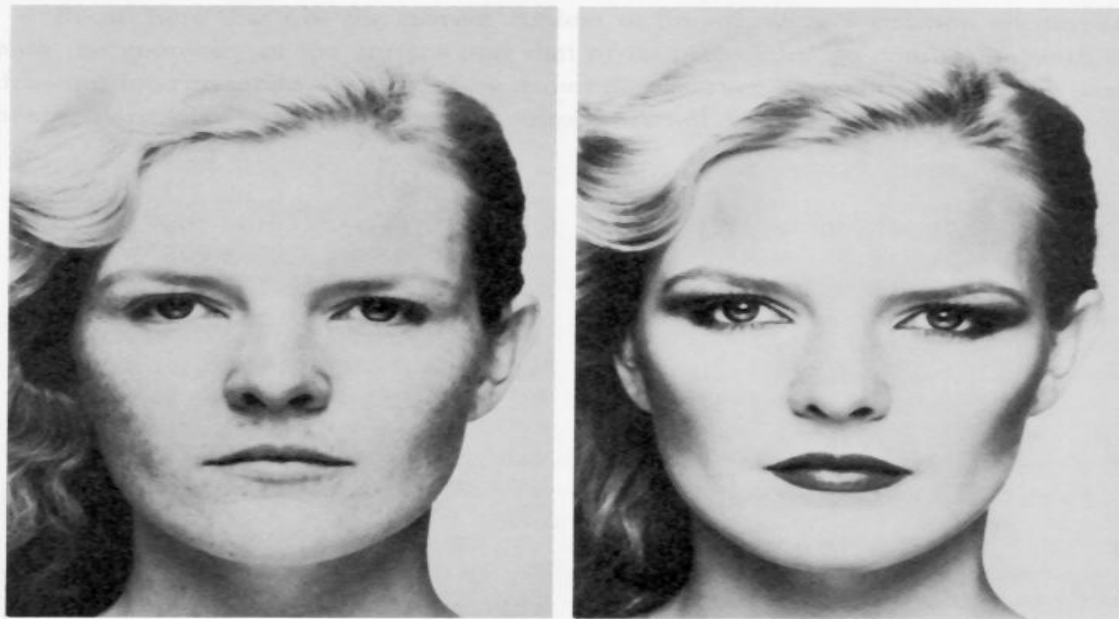
- Assignment #3 due
- Project assigned
  - Hard deadline: 11 April 2017
  - Demo in class/tutorial
- Tonight's tutorial
  - Quiz #1 demo

# Today's Lecture

- Photometric Stereo
  - Project
- Slides acknowledgment: Lana Lazebnik, Fei-Fei Li, Rob Fergus, Antonio Torralba, and Jean Ponce
- Questions

# Photometric Stereo

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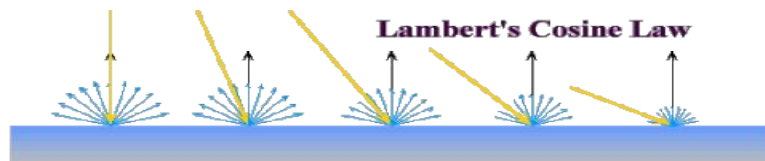
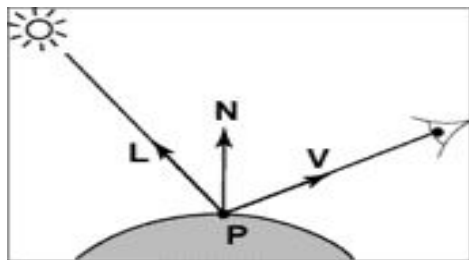


Merle Norman Cosmetics, Los Angeles

## Readings

- Optional: Woodham's original photometric stereo paper
  - <http://www.cs.ubc.ca/~woodham/papers/Woodham80c.pdf>

# Diffuse reflection



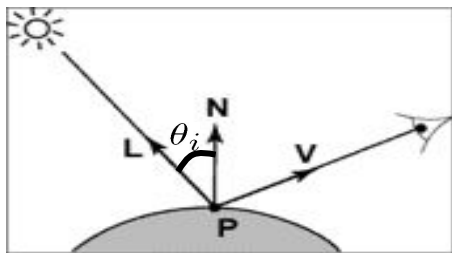
$$R_e = k_d \mathbf{N} \cdot \mathbf{L} R_i$$

image intensity of P  $\longrightarrow I = k_d \mathbf{N} \cdot \mathbf{L}$

## Simplifying assumptions

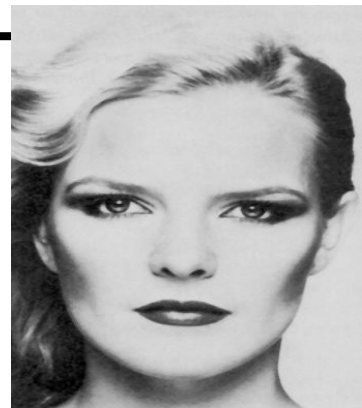
- $I = R_e$ : camera response function  $f$  is the identity function:
- can always achieve this in practice by solving for  $f$  and applying  $f^{-1}$  to each pixel in the image
- $R_i = 1$ : light source intensity is 1
- can achieve this by dividing each pixel in the image by  $R_i$

# Shape from shading



Suppose  $k_d = 1$

$$\begin{aligned} I &= k_d \mathbf{N} \cdot \mathbf{L} \\ &= \mathbf{N} \cdot \mathbf{L} \\ &= \cos \theta_i \end{aligned}$$

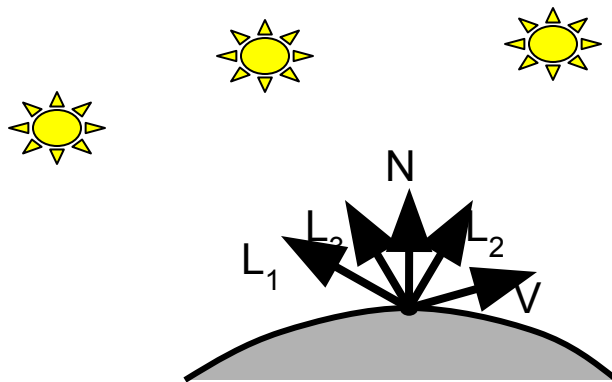


You can directly measure angle between normal and light source

- Not quite enough information to compute surface shape
- But can be if you add some additional info, for example
- assume a few of the normals are known (e.g., along silhouette)
- constraints on neighboring normals—"integrability"
- smoothness
- Hard to get it to work well in practice
- plus, how many real objects have constant albedo?

# Photometric stereo

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$$I_1 = k_d \mathbf{N} \cdot \mathbf{L}_1$$

$$I_2 = k_d \mathbf{N} \cdot \mathbf{L}_2$$

$$I_3 = k_d \mathbf{N} \cdot \mathbf{L}_3$$

Can write this as a matrix equation:

# Solving the equations

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$$\underbrace{\begin{bmatrix} I_1 & \dots & I_n \end{bmatrix}}_{\substack{\mathbf{I} \\ 1 \times 3}} = \underbrace{k_d}_{\substack{\mathbf{G} \\ 1 \times 3}} \mathbf{N}^T \underbrace{\begin{bmatrix} \mathbf{L}_1 & \dots & \mathbf{L}_n \end{bmatrix}}_{\substack{\mathcal{L} \\ 3 \times 3}}$$

$$\mathbf{G} = \mathbf{I} \mathbf{L}^{-1}$$

$$k_d = \|\mathbf{G}\|$$

$$\mathbf{N} = \frac{1}{k_d} \mathbf{G}$$



# More than three lights

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Get better results by using more lights

$$\begin{bmatrix} I_1 & \dots & I_n \end{bmatrix} = k_d \mathbf{N}^T \begin{bmatrix} \mathbf{L}_1 & \dots & \mathbf{L}_n \end{bmatrix}$$

Least squares solution:

$$\begin{aligned} \mathbf{I} &= \mathbf{G}\mathbf{L} \\ \mathbf{I}\mathbf{L}^T &= \mathbf{G}\mathbf{L}\mathbf{L}^T \\ \mathbf{G} &= (\mathbf{I}\mathbf{L}^T)(\mathbf{L}\mathbf{L}^T)^{-1} \end{aligned}$$

Solve for  $\mathbf{N}$ ,  $k_d$  as before

What's the size of  $\mathbf{L}\mathbf{L}^T$ ?



# Color images

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## The case of RGB images

- get three sets of equations, one per color channel:

$$\mathbf{I}_R = k_{dR} \mathbf{N}^T \mathcal{L} \quad \text{--- call this } \mathbf{J}$$

$$\mathbf{I}_G = k_{dG} \mathbf{N}^T \mathcal{L}$$

$$\mathbf{I}_B = k_{dB} \mathbf{N}^T \mathcal{L}$$

- Simple solution: first solve for  $\mathbf{N}$  using one channel
- Then substitute known  $\mathbf{N}$  into above equations to get  $k^d$  s:

$$\mathbf{I}_R = k_{dR} \mathbf{J}$$

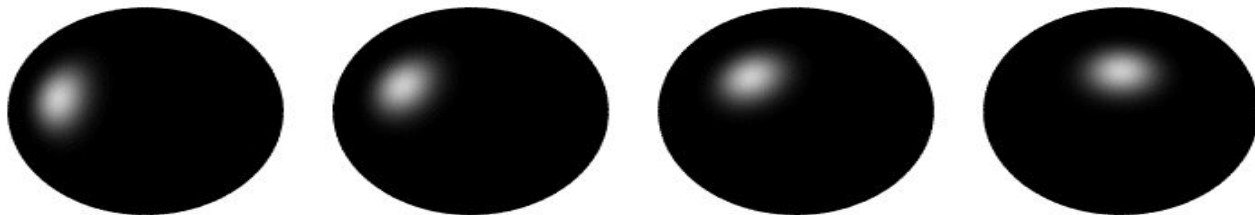
$$\mathbf{J} \cdot \mathbf{I}_R = k_{dR} \mathbf{J} \cdot \mathbf{J}$$

$$k_{dR} = \frac{\mathbf{J} \cdot \mathbf{I}_R}{\mathbf{J} \cdot \mathbf{J}}$$

# Computing light source directions

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Trick: place a chrome sphere in the scene

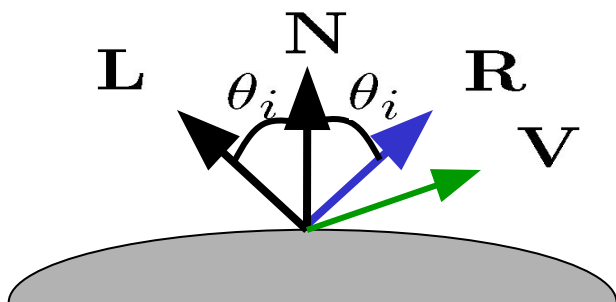


- the location of the highlight tells you where the light source is

# Recall the rule for specular reflection

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For a perfect mirror, light is reflected about  $\mathbf{N}$



$$R_e = \begin{cases} R_i & \text{if } \mathbf{V} = \mathbf{R} \\ 0 & \text{otherwise} \end{cases}$$

We see a highlight when  $\mathbf{V} = \mathbf{R}$

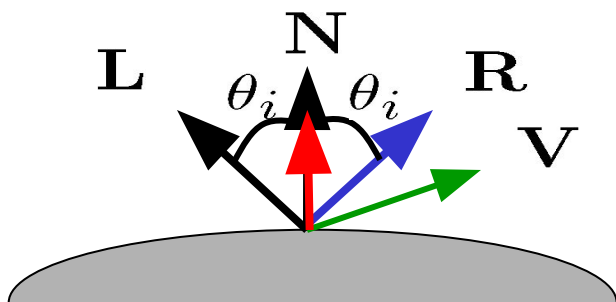
- then  $\mathbf{L}$  is given as follows:

$$\mathbf{L} = 2(\mathbf{N} \cdot \mathbf{R})\mathbf{N} - \mathbf{R}$$

# Recall the rule for specular reflection

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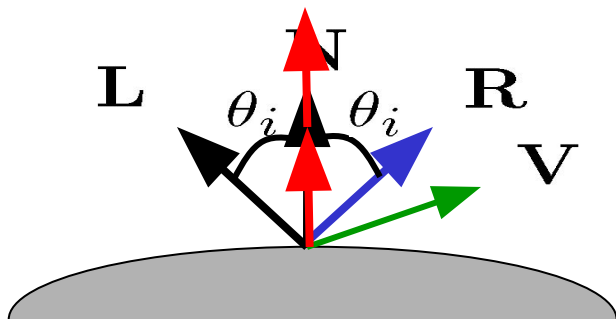
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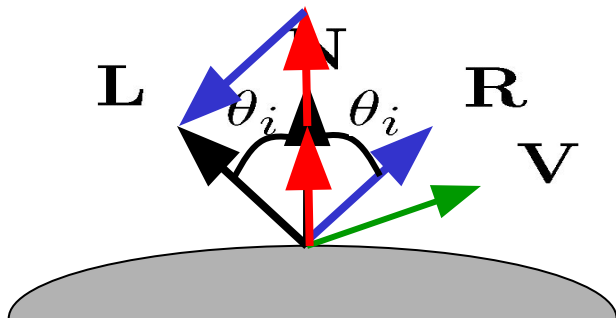
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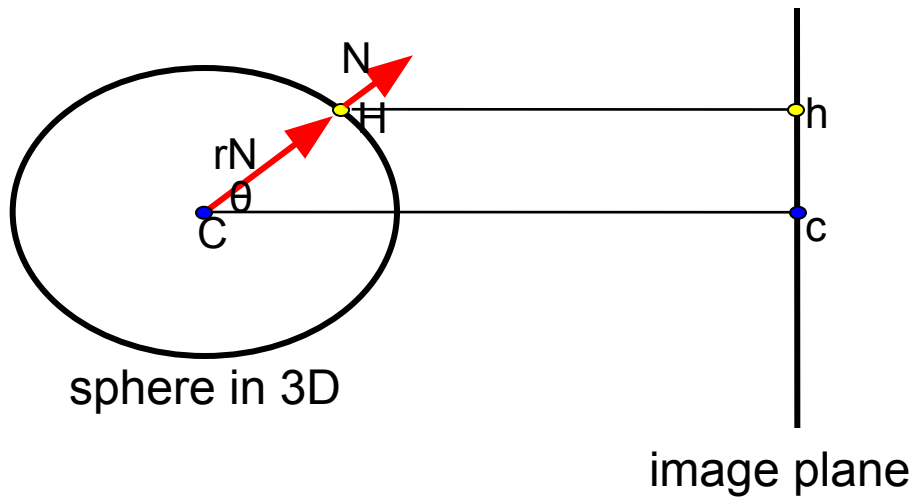
We see a highlight when  $\mathbf{V} = \mathbf{R}$

- then  $\mathbf{L}$  is given as follows:

$$\mathbf{L} = 2(\mathbf{N} \cdot \mathbf{R})\mathbf{N} - \mathbf{R}$$

# Computing the light source direction

Chrome sphere that has a highlight at position  $h$  in the image



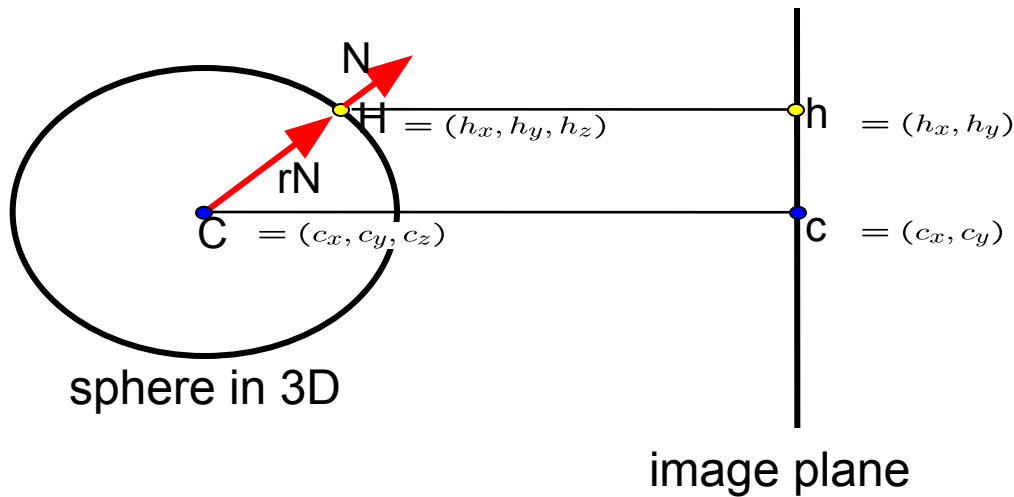
Can compute  $\theta$  (and hence  $N$ ) from this figure

Now just reflect  $V$  about  $N$  to obtain  $L$



# Computing the light source direction

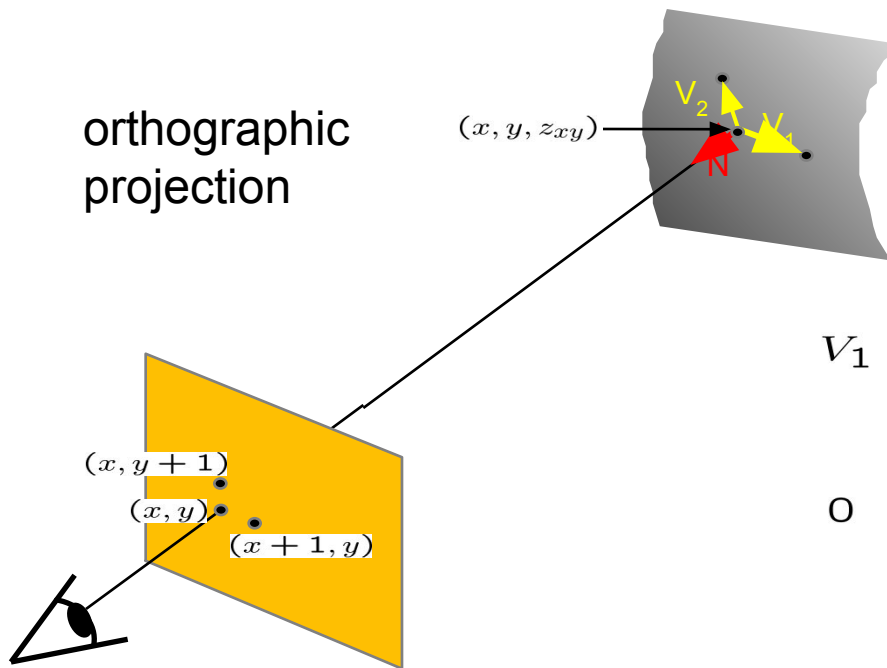
Chrome sphere that has a highlight at position  $h$  in the image



Can compute  $N$  by studying this figure

- Hints:
  - use this equation:  $\|H - C\| = r$
  - can measure  $c$ ,  $h$ , and  $r$  in the image
  - can choose  $c_z = 0$

# Depth from normals



orthographic  
projection

$(x, y, z_{xy})$

$$\begin{aligned} V_1 &= (x + 1, y, z_{x+1,y}) - (x, y, z_{xy}) \\ &= (1, 0, z_{x+1,y} - z_{xy}) \end{aligned}$$

$$\begin{aligned} 0 &= N \cdot V_1 \\ &= (n_x, n_y, n_z) \cdot (1, 0, z_{x+1,y} - z_{xy}) \\ &= n_x + n_z(z_{x+1,y} - z_{xy}) \end{aligned}$$

Get a similar equation for  $V_2$

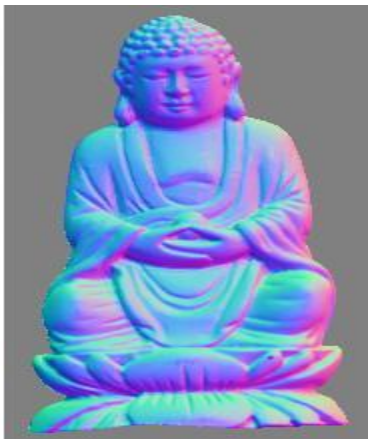
- Each normal gives us two linear constraints on  $z$
- compute  $z$  values by solving a matrix equation

# Results...

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Input  
(1 of 12)



Normals



Normals



Shaded  
rendering



Textured  
rendering

# Results...

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from Athos Georghiades

<http://cvc.yale.edu/people/Athos.html>

# Limitations

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## Big problems

- doesn't work for shiny things, semi-translucent things
- shadows, inter-reflections

## Smaller problems

- camera and lights have to be distant
- calibration requirements
  - measure light source directions, intensities
  - camera response function

# Trick for handling shadows

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Weight each equation by the pixel brightness:

$$I_i(I_i) = I_i[k_d \mathbf{N} \cdot \mathbf{L}_i]$$

Gives weighted least-squares matrix equation:

$$\begin{bmatrix} I_1^2 & \dots & I_n^2 \end{bmatrix} = k_d \mathbf{N}^T \begin{bmatrix} I_1 \mathbf{L}_1 & \dots & I_n \mathbf{L}_n \end{bmatrix}$$

Solve for  $\mathbf{N}$ ,  $k_d$  as before

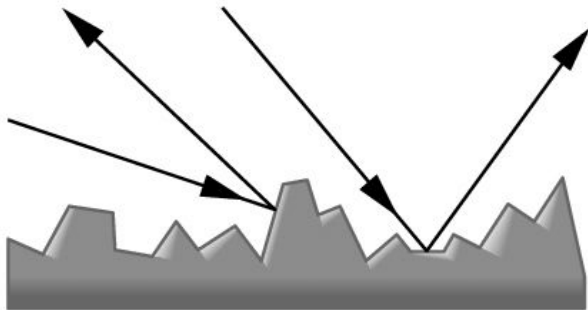
A large, solid blue abstract shape that spans across the middle of the slide. It has a jagged, angular appearance, starting from the left edge, dipping down, and then rising towards the right edge.

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# Diffuse Reflection

- Diffuse reflector scatters light equally to all directions
- Called *Lambertian* surface
- Diffuse reflection coefficient  $k_d$ ,  $0 \leq k_d \leq 1$

- What happens if the lighting direction is changed?
- What happens if the viewer's direction is changed?

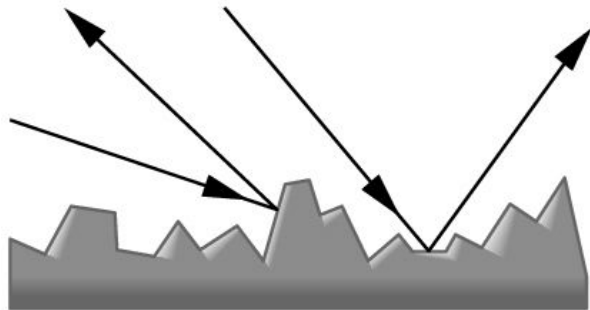




# Diffuse Reflection

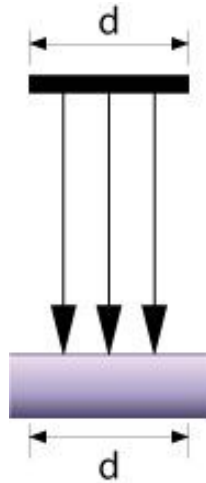
- Diffuse reflector scatters light equally to all directions
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- Diffuse reflection coefficient  $k_d$ ,  $0 \leq k_d \leq 1$
- Only the angle of incoming light is important

- What happens if the lighting direction is changed?
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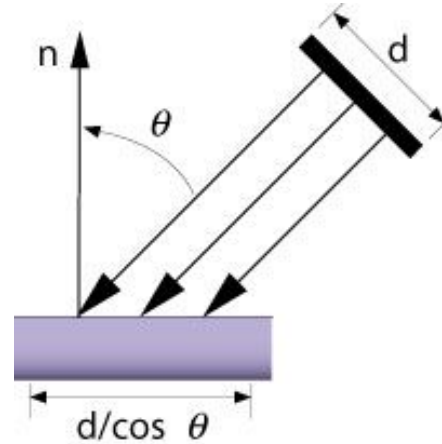


# Lambert's Law

Intensity depends on angle of incoming light



(a)



(b)

# Diffuse Light Intensity Depends on Angle of Incoming Light

## Recall

$l$  = unit vector to light

$n$  = unit surface normal

$\theta$  = angle to normal

- $\cos \theta = l \cdot n$

- $I_d = k_d L_d (l \cdot n)$

- With attenuation:

$q$  = distance to light

$$I_d = \frac{k_d L_d}{a + bq + cq^2} (l \cdot n)$$

$L_d$  = diffuse component of light