1st Place Solution for ICDAR 2021 Competition on Mathematical Formula Detection

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- ▶ Introduction to Mathematical Formula Detection
- ▶ Major challenges
- ► Key points of our solution²
- Experiment

²Yuxiang Zhong, et al.: 1st Place Solution for ICDAR 2021 Competition on Mathematical Formula Detection. CoRR abs/2107.05534 (2021)

MFD Task

where S(4) is the action of the field . Using standard (non-rigorous) methods of quantum field theory a number of new and unexpected mathematical results have been been derived from topological models, results which in many cases have then been fully proved by more standard mathematical methods, but which would probably not have been discovered without the insights gained from the quantum field theory. (An early appearance of topological invariants in the quantum field theoretic situation is due to Belavin. Polyakov, Schwarz and Tyupin [1]. A more recent example of the powerful application of topological quantum field theory in mathematics may be found in [2], while fuller accounts of earlier work in this field may be found in the books of Nash [3] and Schwarz [4].) Most functional integrals such as (1), and related expressions with operator insertions, have not at present been properly defined. However, since these integrals have such astonishing mathematical power, it seems that an attempt to define these objects rigorously should be more than worth while. In this talk we show how this may be done for the simplest topological model, the topological particle, and describe briefly some recent work by Hrabak [5] which might lead to progress in the canonical quantization of topological field theories.

Some rigorous results on path integrals (that is, functional integrals in quantum mechanics) are known. The basic classical result (which is described by Simon in [6]) for a particle of unit mass moving in one dimension with Hamiltonian

$$H = \frac{1}{2}p^2 + V(x)$$
 (2)
gives the action of the imaginary time evolution operator $\overline{\exp(-H!)}$ on a
wave function $\overline{\mathbb{P}(x)}$ by the formula

 $\exp(-Ht)\psi(x) = \int d\mu \exp(-\int V((x(s))ds))\psi(x(t))$

where $\frac{W}{W}$ denotes Wiener measure starting from and $\frac{W}{W}$ are corresponding Brownian paths; the potential $\frac{W}{W}$ must sarisky certain analytic conditions. The curved space analogue of this result for a Riemannian mariotid blas been structured to the starting of the starting the starting of the starting terms of the starting terms of the scalar Laplacian looks identical to (3), but with $\frac{W}{W}$ a process depending on metric and connection rather than simply flat space Brownian motion. Tangent space geometry plays an essential part in the theory. The present analyse has threst excluded these are started by the weight of all targets theory and the late of their excluded the simple of the depth of the starting three starting from the starting three starting from the starting terms and the starting from the starting fr

class => {0: embedded. 1: isolated} _____ # x rel y rel width height class 22.18 53.76 14.58 3.22 22.18 62.21 47.55 3.86 15.77 22.94 4.01 1.46 48.45 15.82 1.24 1.37 69.18 58.15 8.64 1.51 29.79 59.86 4.08 1.51 67.24 3.39 1.46 67.29 1.37 2.28 57.98 67.68 1.11 0.68 69.04 1.66 1.03 13.27 1.51 74.07 75.54 74.17 1.38 1.07 60.68 75.83 3.39 1.46

Ground Truth.

Isolated and Embedded Formulas.

manifolds to give Brownian motion on supermanifolds in a suitable form for

Samples

where \square is the original ranking weeter bundle over \square . Scond, we must specify a line bundled on the surface \square which we will glue together with \square 01 to produce \square 0. A priori, \square 0 is arbitrary, but, as we will see, it is actually subject to strong constraints. To find these constraints, we first use the Fourier-Mukai rankformation to construct a vector bundle \square 01 from the spectral data $\prod_{i=1}^{\infty} \square$ 1 that is

$$(\pi^*z, \ell) \longrightarrow V_1$$
 (5

where $[\underline{\Gamma}]$ is a rank $[\underline{\Pi}]$ vector bundle over the curve $[\underline{\underline{\Gamma}} = \underline{\sigma} - \underline{\sigma}^*] = \underline{\Lambda}$ Now, the fact that, by construction, the base component of the bulk five-brane class associated with $[\underline{\Pi}]$ must be equal to $[\underline{W}_{\underline{\sigma}} = \underline{\sigma}^*]$, implies that $[\underline{E} = \underline{\sigma}^*]$ for some line bundle $[\underline{\Pi}]$ on the curve $[\underline{\Pi}]$ in the base. A simple Fourier-Mukai calculation shows that $[\underline{\Gamma}]$ is just

$$V_z = i_*(L \otimes K_B) \qquad (5.6)$$

where

is the embedding

iv: 3 → X of the curve

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curve $\overline{\mathbb{Q}}$ in $\overline{\mathbb{W}}$ where their data overlap. This is done by specifying a surjection

$$\xi: V|_1 \longrightarrow V_*$$

That such a surjection exists and is unique will become clear shortly. Given this relation, one can define a "bundle" \square on \square via the exact sequence³ $0 \to \widetilde{V} \to V \to V_* \to 0 \qquad (5.8)$

It can be shown that because

Hence, to complete our construction, we will have to show that $\widehat{\mathbb{M}}$ can be "smoothed out" to ${}^{2}Aa$ technical aside, we note that the vector bundle" giver the curved can be foundly extended over the Calchia' Yea threshold $\widehat{\mathbb{M}}$ as an object that vanishes everywhere orisis the curved $\widehat{\mathbb{M}}$ is distinct to $\widehat{\mathbb{M}}$ out $\widehat{\mathbb{M}}$ this extended object will also be denoted $\widehat{\mathbb{M}}_{\mathbb{M}}^{m}$ by well let context distant which of these notions is to be used. This remark conducts at $\widehat{\mathbb{M}}$ of the brundles discussed the curve form $\widehat{\mathbb{M}}$ is the conduct of the cond

The bundle \(\frac{1}{2} \) defined in this way bears a special name. It is a called a Hecke transform of \(\mathbb{N} \) and the pair \(\mathbb{K} \), \(\mathbb{V}_L \) is called the center of the Hecke transform.

"The notion of stability here is similar to that used for vector bundles. The differential prometric country required of a stable shall be a Hermitian Negality differential on the vector bundles of which is smooth outside of the curve [T. A] met has a delta function behavior along the curve [B. other words, a stable towards for shall [B] be night-tain-prometry interaction of a summarized on the curve [B. This justifies the terminology "small instanton phase transition" that we use below to describe the two term process of first controlling 0 and 000 and the deferminal [B] as mountly works bundle [B] and [B]

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the Hitchin spectral curve \mathbb{Q} , then the kernel of the map $[z-\zeta]$ from $H^1(K_{\sigma})$ to $H^1(K_{\sigma})$ is nontrivial. But the cohomology exact sequence implies an isomorphism

$$\text{Ker}(z - \zeta) \cong H_{\mathcal{D}}^{0}(\mathbb{S}^{1}, E|_{z=\zeta}).$$
 (3)

Therefore $[H_1](S_1, E_{1:m_1}]$ is nontrivial as well. This means that the bolomout of Π along the circle 1 = 3 has $\frac{1}{2} \log 2\pi n$ as one of its eigenvalues. Thus the point $\{I_1, \exp(2\pi n)\} \in \mathbb{C} \times \mathbb{C}^n$ belongs to the monopole spectral curve \mathbb{S} . Moreover, the fibers of the spectral line bundles on Ω and S are given by $[K_{\Pi}^{-1} = 1]$ and $[H_{\Pi}^{-1}(S_1, E_{1:m_1})]$ respectively. Thus we also get an isomorphism of the learning M.

5.4 Exactness Of The Cohomology Sequence

Consider again the complex of sheaves

$$0 \rightarrow E \xrightarrow{(z-\zeta)} E \xrightarrow{rest.} E|_{z=\zeta} \rightarrow 0,$$
 (3)

where \overline{prst} is the restriction to $[\underline{z} = \zeta]$. Since $[D_p, p = 0, 1]$ commutes with $[\underline{z} - \zeta]$ and \overline{prst} , this complex is included in a double complex $[D^{prot}]$.

$$E \otimes \Lambda^{0.2} \xrightarrow{G \leftarrow G} E \otimes \Lambda^{0.2} \xrightarrow{resd.} 0$$

$$D_{\uparrow} \xrightarrow{D_{\uparrow}} D_{\uparrow} \xrightarrow{D_{\uparrow}} D_{\uparrow}$$

$$E \otimes \Lambda^{0.1} \xrightarrow{G \leftarrow G} E \otimes \Lambda^{0.1} \xrightarrow{rresd.} E \otimes \Lambda^{0.1} |_{z=c}$$

$$D_{\uparrow} \xrightarrow{D_{\uparrow}} \xrightarrow{D_{\uparrow}} E \otimes \Lambda^{0.0} \xrightarrow{rrest.} E \otimes \Lambda^{0.0} |_{z=c}$$

Computing the cohomology of the rows, we obtain the first term of the "vertical" spectral sequence:

On the second level the "vertical" spectral sequence degenerates to zero: $E_{\nu\nu}^{pov} = 0$

Major Challenges

► Large scale span

Large variation of the ratio between height and width

► Rich character set and mathematical expressions

Our Solution •00000

Generalized Focal Loss (GFL)

Adaptive Training Sampling Strategy (ATSS)

Feature Pyramid Network (FPN)

Comparison between anchor-based and anchor-free methods

Our Solution

Pre-set anchors vs no anchors

Pre-set scale and ratio vs any points on object for positive sampling

Regression of the box residual vs direct regression

Drawbacks of anchor-based method on MFD task

It is hard to preset effective anchors to cover large span of scales and aspect ratios.

Our Solution 000000

Even with ATSS, small objects cannot be assigned positive anchors because anchor candidate will be filtered out by IoU.

GFL

Quality Focal Loss (QFL)

Distribution Focal Loss (DFL)

Random sampling vs ATSS sampling.

which after substitution in Eq. (12) leads to

$$a^2 - D\beta^2 = \frac{d_b^2 \aleph^2 R_0^4}{4}$$
 (14)

In these coordinates the dilaton, or equivalently the string coupling squared, is $\overline{y_s^2} = e^{2\phi} = M^{-1}\eta^2 R^{0}q$ and the Kalb-Ramond field strength is $H = 8R^{0}de R J$.

Although there is a world of possibilities in the above class of solutions, perhaps the most interesting case is the one where [5,-4], since then the uncompactified part of spacetime is just Minkowski space: Taking [7,-4], since then the uncompactified part of spacetime is just Minkowski space: Taking [7,-4], since then the uncompactified part of spacetime is cosmological time, __ reads

$$ds^2 = d\tau^2 - d\tilde{x}_{LD} - 2R_0^2 B(\tau)^{-1} \eta_{cd} dy^a dy^b,$$

 $v^{2b} = \left(\sqrt{2}R_0\right)^{d_0} \tau^{-1} B(\tau)^{-d_0/2},$
 $H = 8R_0^2 d_0^{1/2} \tau^{-1} B(\tau)^{-2} d\tau \wedge \sqrt{f},$ (15)
 $B(\tau) = \tau^{2/\sqrt{d_0}} + \tau^{-2/\sqrt{d_0}}.$ (16)

As one can see, this is a completely regular solution, modulo the usual gravitational singularities, which amonthly interpolates between two Kasare-like regions [6]. From the lower dimensional point of view, the Ansatz considered above corresponds to a solution of dilaton-gravity coupled to moduli [16], where the breathing mode, Ω , and Ω are the scalar fields parameterizing an BLI(2,E)(I) [1] cost model. When Ω = Ω , the above solutions can be obtained from the solutions given in [2] by applying an BLL(2,E)(I) transformation on the moduli.

2 RR case

In much the same way as in the foregoing subsection, we can use the RR two form in type IIA, to trigger compactification. In this case the equations of motion and the Bianchi identity imply that

$$E_{(2)} = \aleph \mathcal{J} = \frac{1}{2} \mathcal{R} \mathcal{J}_{km} dy^m \wedge dy^n$$
. (

Applying the same steps as in the foregoing paragraph, one finds

$$\begin{aligned} 0 &= (\log R)^{n} + \frac{(d_{n} - \log n)}{2} M \eta^{D} R^{d_{n} - 1} \\ 0 &= (\log \eta^{n} + 2 \frac{d_{n}^{2}}{2} M \eta^{D} R^{d_{n} - 1} - \lambda \eta^{-2} M^{2}) \\ 0 &= ((\log M)^{n} - \frac{d_{n}^{2}}{2} M \eta^{D} R^{d_{n} - 1} - D \lambda \eta^{-2} M^{2}) \\ 0 &= ((\log M)^{n} - D ((\log \eta))^{2} - d_{n}^{2} ((\log R))^{n}) - \frac{d_{n}^{2}}{2} M \eta^{D} R^{d_{n} - 1} - D M^{2} \eta^{-2}) \end{aligned}$$

Looking at the above expressions, one sees that they simplify enormously when one considers the case $\overline{g}_k = \overline{x}_k$. In that case the Kähler breathing mode decouples completely and one has $\overline{R} = R_0 e^{\alpha}$. Equating also the powers of \overline{M} and \overline{g} in the equations, i.e. putting $\overline{M} = \overline{g}^{\alpha - 2}$, one necessarily has to impose

$$\lambda = \frac{\aleph^2}{4}(D+3)$$
 (22)

which after substitution in Eq. (12) leads to

$$-D\beta^2 = \frac{d_k \aleph^2 R_0^4}{\epsilon}$$
. (14)

In these coordinates the dilaton, or equivalently the string coupling squared, is $\overline{\mu_i^2} = M^{-1} H^{-1} K^{-1}$ and the Kalb-Ramond field strength is $H = N H dt \wedge J$.

Although there is a world of possibilities in the above class of solutions, perhaps the most interesting case is the one where [7,], since then the uncompactified part of spacetime is just Minkowski space: Taking [7,] for convenience, one finds that the solution in, string, cosmological time, a reads

$$ds^2 = dr^2 - d\vec{x}_{(D)} - 2R_0^2B(\tau)^{-1}h_{mn}dy^mdy^n$$
,
 $e^{2\phi} = \left(\sqrt{2}R_0\right)^{d_0}\tau^{-1}B(t)^{-d_0/2}$,
 $H = 8R_0^2f_0^{-1/2}\tau^{-1}B(t)^{-2}d\tau \wedge \mathcal{J}$, (15)
 $B(\tau) = \tau^{2/\sqrt{d_0}} + \tau^{-2/\sqrt{d_0}}$. (16)

As one can see, this is a ompletely regular solution, modulo the usual gravitational singularities, which sumes of the term of the regular gravitational singularities, which summer between two Kasarev delwers two Kasarev from the lower to the control point of view, the Ausarz considered above corresponds to a solution of district moduli [10], where the breathing mode, a solution of district from the control point of the present the control point of the present the control point of the co

2 RR case

In much the same way as in the foregoing subsection, we can use the RR two form in type IIA, to trigger compactification. In this case the equations of motion and the Bianchi identity imply that

$$F_{(2)} = \aleph \mathcal{J} = \frac{1}{2} \& J_{mn} dy^m \wedge dy^n$$
. (17)

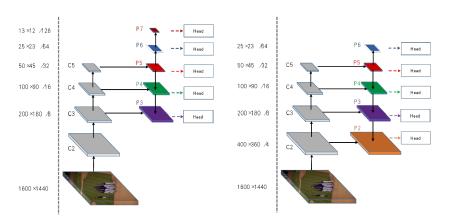
Applying the same steps as in the foregoing paragraph, one finds

Looking at the above expressions, one sees that they simplify enormously when one considers the case $\overline{B_{10}} = \overline{B_{1}}$. In that case the Kähler breathing mode decouples completely and one has $\overline{B_{10}} = \overline{B_{10}} = \overline{B_{10}}$. Equating also the powers of $\overline{B_{10}}$ and $\overline{B_{10}}$ in the equations, i.e. putting $\overline{M_{10}} = \overline{B_{10}} = \overline{B_{10}} = \overline{B_{10}}$. So ne necessarily has to impose

$$\lambda = \frac{\aleph^2}{4!} (D + 3).$$

(22)

FPN



Our Solution 000000

FPN 3-7

FPN 2-6

- Double Training Epoch
- Large Crop Size
- Random Flip
- ResNeSt
- Synchronized Batch Normalization
- Deformable Convolution Network
- Larger Batch Size
- Ranger
- RegMax
- **FPN** Selection
- Weight Box Fusion

Ablation Study

DTE	LCS	Flip	NeSt	SyBN	DCN	LBS	Ranger		FPN (2-6)	WBF	F1-score Embedded	F1-score Isolated	F1-score Total
											89.88 p:91.34 r:88.47	85.92 p:89.77 r:82.39	89.17 p:91.02 r:87.39
√											91.36 p:92.80 r:89.96	86.29 p:90.93 r:82.10	90.45 p:92.26 r:88.71
\checkmark	\checkmark										92.95 p:94.43 r:91.52	87.84 p:92.04 r:84.01	92.03 p:93.58 r:90.53
\checkmark	\checkmark	\checkmark									93.58 p:94.79 r:92.41	88.38 p:92.29 r:84.78	92.66 p:94.35 r:91.03
\checkmark	\checkmark	\checkmark	\checkmark								93.49 p:93.82 r:93.16	91.33 p:93.56 r:89.20	93.12 p:93.71 r:92.54
\checkmark	\checkmark	\checkmark	\checkmark	\checkmark							93.81 p:94.00 r:93.63	91.60 p:93.82 r:89.48	93.42 p:93.97 r:92.89
\checkmark	\checkmark	\checkmark	\checkmark	\checkmark	\checkmark						94.33 p:95.62 r:93.07	95.25 p:95.62 r:94.88	94.49 p:95.62 r:93.39
\checkmark					94.58 p:95.16 r:94.00	95.60 p:95.97 r:95.23	94.76 p:95.31 r:94.22						
\checkmark				95.22 p:95.70 r:94.74	95.81 p:96.12 r:95.51	95.33 p:95.79 r:94.87							
\checkmark			95.01 p:95.86 r:94.16	97.28 p:97.18 r:97.38	95.41 p:96.10 r:94.73								
\checkmark		95.67 p:96.34 r:95.00	97.67 p:97.46 r:97.88	96.03 p:96.54 r:95.53									
\checkmark	96.01 p:95.79 r:96.23	98.14 p:97.32 r:98.98	96.33 p:96.81 r:95.85										

Table 1: Experimental results of baseline, DTE (double training epoch), LCS (larger crop size), Flip (training and testing time flip), SyBN (SyncBN), DCN (deformable convolution network), LBS (larger batch size), Ranger optimizer, Reg 24 (a hyperparameter regmax), FPN (2-6) (FPN selection), WBF (weighted box fusion).

- CutMix data augmentation
- Soft NMS vs original NMS
- loU loss or GloU loss vs smooth L1 loss
- BFP or PAFPN instead of original FPN
- Swish or Mish vs ReLU

Group ID	Type	F1 score (Ts10 + Ts11)	F1 score Task dependent (Ts11)	F1 score Task independent (Ts10)
PAPCIC	E	94.79 (94.89 , 94.69)	95.11 (95.11, 95.11)	94.64 (94.79 , 94.50)
	I	98.76 (98.25 , 99.28)	98.70 (98.34, 99.07)	98.79 (98.21 , 99.37)
	S	95.47 (95.47 , 95.47)	95.68 (95.62, 95.73)	95.37 (95.40 , 95.35)
Lenovo	E	94.29 (95.36 , 93.25)	93.98 (95.41 , 92.60)	94.44 (95.34 , 93.56)
	I	98.19 (98.26 , 98.12)	97.85 (98.04 , 97.67)	98.33 (98.35 , 98.31)
	S	94.96 (95.86 , 94.08)	94.60 (95.84 , 93.40)	95.13 (95.87 , 94.39)
DLVCLab	E	93.79 (94.54 , 93.05)	93.88 (94.70 , 93.07)	93.75 (94.46 , 93.04)
	I	98.54 (98.19 , 98.89)	98.61 (98.33 , 98.88)	98.51 (98.13 , 98.90)
	S	94.60 (95.17 , 94.04)	94.64 (95.29 , 93.99)	94.59 (95.12 , 94.07)
TYAI	E	93.39 (94.43, 92.38)	93.94 (95.05 , 92.86)	93.13 (94.13 , 92.15)
	I	98.55 (98.19, 98.92)	98.42 (98.15 , 98.70)	98.61 (98.21 , 99.02)
	S	94.28 (95.08, 93.49)	94.66 (95.55 , 93.79)	94.1 (94.86 , 93.35)
SPDBLab	E	92.80 (93.25 , 92.36)	92.14 (92.83 , 91.46)	93.12 (93.45 , 92.79)
	I	98.06 (98.06 , 98.06)	97.76 (97.85 , 97.67)	98.19 (98.15 , 98.23)
	S	93.70 (94.08 , 93.33)	93.03 (93.63 , 92.44)	94.01 (94.28 , 93.74)
YoudaoAI	E	92.73 (93.57 , 91.91)	92.71 (93.95 , 91.51)	92.74 (93.39 , 92.10)
	I	98.34 (97.66 , 99.03)	98.38 (97.97 , 98.79)	98.32 (97.52 , 99.13)
	S	93.70 (94.28 , 93.12)	93.63 (94.61 , 92.67)	93.74 (94.14 , 93.34)
PKUF-MFD	E	91.94 (92.18 , 91.70)	92.32 (92.93 , 91.72)	91.76 (91.82, 91.69)
	I	96.56 (96.88 , 96.24)	96.87 (97.28 , 96.46)	96.43 (96.72, 96.15)
	S	92.72 (92.98 , 92.47)	93.04 (93.62 , 92.47)	92.57 (92.68, 92.47)
HW-L	E	90.53 (91.55 , 89.53)	90.57 (91.82 , 89.35)	90.51 (91.42, 89.61)
	I	98.94 (98.81 , 99.06)	98.61 (98.33 , 98.88)	99.08 (99.02, 99.13)
	S	91.97 (92.81 , 91.15)	91.86 (92.88 , 90.86)	92.02 (92.77, 91.28)
Komachi	E	90.39 (90.92 , 89.86)	89.69 (90.43 , 88.97)	90.72 (91.16 , 90.28)
	I	98.57 (98.27 , 98.87)	98.6 (98.69 , 98.51)	98.55 (98.09 , 99.02)
	S	91.79 (92.19 , 91.39)	91.11 (91.75 , 90.48)	92.10 (92.39 , 91.81)
AIG	E	89.71 (90.18 , 89.25)	89.19 (89.96 , 88.44)	89.95 (90.28 , 89.63)
	I	95.95 (99.00 , 93.09)	96.07 (99.01 , 93.30)	95.90 (99.00 , 93.00)
	S	90.75 (91.61 , 89.90)	90.26 (91.34 , 89.21)	90.97 (91.74 , 90.22)
PKUSG	E	89.10 (90.26 , 87.97)	88.59 (90.23 , 87.00)	89.34 (90.27 , 88.42)
	I	97.96 (97.85 , 98.06)	97.94 (98.31 , 97.58)	97.96 (97.66 , 98.27)
	S	90.62 (91.58 , 89.68)	90.09 (91.54 , 88.68)	90.87 (91.60 , 90.15)
TAL	E	87.87 (88.56 , 87.20)	88.51 (89.12 , 87.92)	87.57 (88.29 , 86.85)
	I	96.85 (96.31 , 97.40)	97.09 (96.33 , 97.86)	96.75 (96.30 , 97.21)
	S	89.42 (89.91 , 88.93)	89.89 (90.29 , 89.49)	89.19 (89.73 , 88.67)
UIT	E	86.04 (85.60 , 86.49)	85.64 (85.62 , 85.65)	86.23 (85.58 , 86.89)
	I	97.05 (95.12 , 99.06)	98.11 (97.08 , 99.16)	96.60 (94.31 , 99.02)
	S	87.94 (87.26 , 88.63)	87.63 (87.47 , 87.80)	88.08 (87.16 , 89.02)
AV-DFKI	E	85.35 (87.37 , 83.41)	84.75 (86.96 , 82.66)	85.63 (87.57 , 83.77)
	I	97.48 (97.12 , 97.84)	97.40 (97.30 , 97.49)	97.52 (97.04 , 97.99)
	S	87.45 (89.10 , 85.87)	86.80 (88.67 , 85.01)	87.76 (89.30 , 86.27)
VH	E	84.25 (83.49 , 85.02)	84.39 (83.76 , 85.02)	84.18 (83.37 , 85.01)
	I	98.59 (98.51 , 98.67)	98.51 (98.42 , 98.60)	98.62 (98.55 , 98.70)
	S	86.67 (86.01 , 87.34)	86.61 (86.06 , 87.18)	86.70 (85.99 , 87.41)

