Continuum effects for the shell-model calculation near the drip line oxygen isotopes

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JPS annual meeting, September 23rd 2007, Hokkaido

- From stable line to drip line
- Continuum effect
- Calculation and Results
 - The first excited state of ²⁴O
 - The first excited state of ²³O
 - The ground state of ²⁵O
 - Density
- Summary
- Compliment
 - The ground state of ²⁶O(preliminary)
 - Convergence

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Near the β -stable line is well described with shell-model in which s.p. states are treated as bound state and the wave functions are described as the eigenfunction of H.O potential.

From β -stable to exotic Near the drip line, excitation strength distribution is explained by the continuum coupling in which not only the resonant state but also the whole continuum states are mixed through the residual interaction.

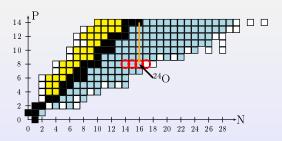


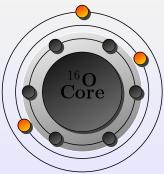
Fig 2.1: Nuclear chart in small mass region Univ. of Tokvo

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The drip line of oxygen isotopes

- Many peaks are observed near the particle emission threshold.
- New magic numbers N = 16, 14 seem to be confirmed. (Stanoiu, et al., PRC 69 2004)
- $0d_{3/2}$ orbit of neutron is lying at almost unbound level.



Binding energy of oxygen isotopes relative to ¹⁶O by shell-model calculation with the lowest configuration which reproduce the full calculation

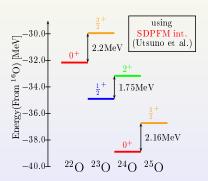


Fig 3.1: The lowest configuration

We reproduce the experimentally measured one neutron separation energy of ²³O with full shell-model calculation, modifying

$$\delta < s1/2, d5/2 | V | s1/2, d5/2 >_{\text{mono pole}}^{T=1}$$

= -0.03MeV(< 5%modification).

And then reproduce the levels of ²³⁻²⁵O relative to ²²O in terms of the filling configuration, modifying both s.p.e and 2-body matrix elements

Generating the continuum $d_{3/2}$ states

$$\mathcal{H}_0 = T + U_{WS} + V_{wall}$$
 $R = 1.09A^{1/3}, \text{diff} = 0.67,$ $V_{ls} = -0.44V_0$

 V_0 is determined so that W.S. potential satisfies the relation bellow.

$$< 0d_{3/2}|T+U_{WS}|0d_{3/2}> = 2.22$$
MeV

Note: $0d_{3/2}$ is H.O. wave function

The condition that the place of wall L should at least meet

t_r =time scale corresponding the excitation energy

 $E(\sim 0.1 - 1 \text{MeV})$ t_c =classical time scale of which outgoing wave comes back to the origin after reflected at the wall.

$$t_r < t_c \Rightarrow L > O(10^1)[\text{fm}]$$

For the calculation done later, the spectrum obtained should be adequately dense.

$$\Rightarrow$$
 We set $L = 1000$ [fm]

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Diagonalization of the hamiltonian with residual interaction

 J^+ states of ²⁴O can be constructed with these bases

$$\left|iJ^{+}\right\rangle = \left|\left[1s_{1/2} \otimes id_{3/2}^{c}\right]; J^{+}M >$$
 $(i = 1, \cdots, n_{\text{max}}, \text{ corresponding to } 0 - 20\text{MeV})$

then we diagonalize the hamiltonian

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$$\mathcal{H} = \sum_i \epsilon^i a_i^\dagger a_i + \frac{1}{4} \sum_{ijkl} \bar{v}_{res}^{ij,kl} a_i^\dagger a_j^\dagger a_l a_k$$
 Type1° $V_{res}(r_i,r_j) = g_d \delta(r_{ij})$, and Type2° $V_{res}(r_i,r_j) = g_1(1+a_1\sigma_1\cdot\sigma_2) \exp(-r_{ij}^2/\mu_1^2) + g_2(1+a_2\sigma_1\cdot\sigma_2) \exp(-r_{ij}^2/\mu_2^2)$ where $\mu_1 = 1.415$, $\mu_2 = 0.7$

Diagonalization of the hamiltonian with residual interaction

The parameters g_d , g_1 , g_2 , a_1 , a_2 are determined so as to meet the relations(a_2 is given by hand, and can be different)

$$ar{v}_{res}^{sd,sd,J=2,T=1} = -0.4165 \text{MeV}, \quad ar{v}_{res}^{sd,sd,J=1,T=1} = 0.6246 \text{MeV},$$
 and $ar{v}_{res}^{ss,ss,J=0,T=1} = -1.1967 \text{MeV}.$ $\Rightarrow g_1 = -21.0, g_2 = 440 \text{MeV}, \qquad a_1 = -5.0, a_2 = 0.2$ $g_d = -301 \text{MeV} \cdot \text{fm}$

Where, $1s_{1/2}$, $0d_{3/2}$ is H.O. wave function. The parameters are determined so that they can reproduce the shell-model matrix elements of SDPFM.

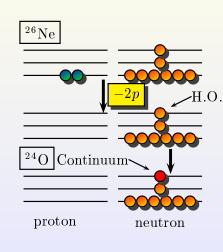
$$\left|J_{k}^{+}\right\rangle = \sum_{j=1}^{n_{\max}} C_{(k)}^{j} \left|\left[1s_{1/2} \otimes j \frac{d_{3/2}^{c}}{3/2}; J^{+}\right]\right\rangle \qquad (k = 1, \cdots, n_{\max})$$

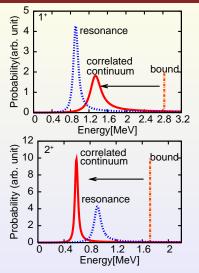
The neutron emission probability in ²⁴O

We assume the low-lying states in ²⁴O is obtained through the reaction 9 Be(26 Ne, 24 O)X. After that, a neutron is emitted if the low-lying states are unbound. In this case, the emission probability is considered to be proportional to the overlap between these states bellow

Prob.
$$\propto \left|\left\langle 1s_{1/2}0d_{3/2}; J^+ \middle| J_k^+ \right\rangle\right|^2$$

$$\propto \left|\sum_i C_{(k)}^j \left\langle 0d_{3/2} \middle| jd_{3/2}^c \right\rangle\right|^2$$





Experimental data (C.Hoffman, private communication) is now preliminary, but roughly

$$E(2^+) \sim 670 \mathrm{keV}$$

 $E(1^+) \sim 1.35 \mathrm{MeV}$

Our result

$$E(2^+) \sim 600 \text{keV}$$

 $E(1^+) \sim 1.35 \text{MeV}$

Fig 4.1: Emission probability in ²⁴O*

The neutron emission probability in ²³O

- The spin parity of the ground state is 1/2+ (Sauvan et al,PLB,2000)
- The excited states(resonance), measured by using the reaction ²²O(d,p)²³O, is
 4.0MeV(3/2+),5.3MeV(5/2+) (Elekes et al, PRL98,102502,2007)

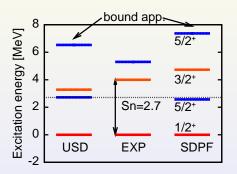


Fig 4.2: Excited states of ²³O. Both ends are the shell-model calculation with two deferent interactions.

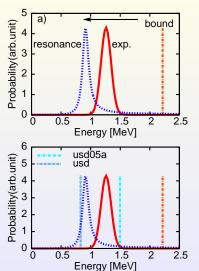


Fig 4.3: overlap value for ²³O Figure

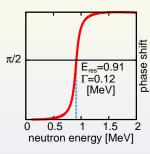
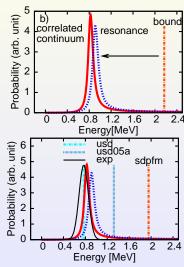


Fig 4.4: one particle phase shift for the W.S. potential given by the ²²O core

'resonance' corresponds to the one particle resonant state.

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Neutron emission probability²⁵O



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Assuming ⁹Be(²⁶F,²⁵O)X Experimental data

$$E_{\rm decay} = 770(20) {\rm keV}$$
 $\Gamma \sim 140 {\rm keV}$

Our result

$$E_{d3/2} = 820 \text{keV}$$

 $\Gamma \sim 65 \text{keV}$

Note:USD seems to be the best, but it has another problem. Using this interaction, ²⁶O is bound for the two neutron separation.

Calculation and Results

Density

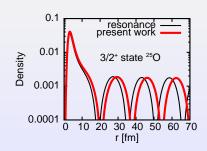
Density of the neutrons is

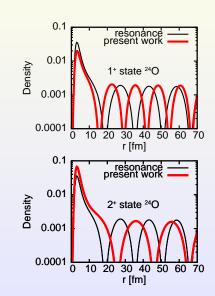
$$\rho_k(r) = \langle \Psi_k | \hat{\rho}(r) | \Psi_k \rangle$$

$$= n_s |\phi_{1s}(r)|^2$$

$$+ \sum_{ij} c^{i*}_{(k)} c^j_{(k)} \phi^*_{id^c}(r) \phi_{jd^c}(r)$$

$$= \rho_s(r) + \rho_{d,k}(r)$$





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Result and discussion

- The peak denoted by 'resonance' is corresponding to the one-particle resonance of W.S. potential(~ 0.9[MeV]).
- The coupling to the continuum states lowers the energy of the 1st excited 2⁺ state of ²⁴O by about 1MeV, but 2⁺ state is still unbound. 1⁺ state is pushed up from the resonance.
- In the case of ²⁵O, the ground state is lowered by because of the coupling between bound and continuum states, but is still unbound, and it is consistent with the recent experiment.
- With the residual interaction, wave function changes from that
 of the resonance in and near the nuclear surface.

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The ground state of ²⁶O(preliminary)

We diagonalize the hamiltonian in M = 0(J = 0, 2) space. Wave function is written as

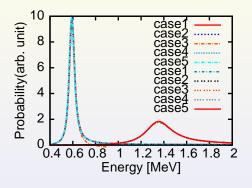
$$\begin{split} |\Psi_k; M &= 0 > = \sum_{i,j} c^{ij}_{(k)} \left| s_{1/2,1/2} s_{1/2,-1/2} i d^c_{3/2,3/2} j d^c_{3/2,-3/2} \right\rangle \\ &= \sum_{i,j} d^{ij}_{(k)} \left| s_{1/2,1/2} s_{1/2,-1/2} i d^c_{3/2,1/2} j d^c_{3/2,-1/2} \right\rangle \\ &=: \sum_{i,j} c^{ij}_{(k)} |\Psi^{ij}_1 > + d^{ij}_{(k)} |\Psi^{ij}_2 > \end{split}$$

$$H(M=0) \doteq \left(\begin{array}{ccc} A + B^{+} + C^{+} & B^{-} + D \\ B^{-} + D & A + B^{+} + C^{-} \end{array} \right)$$

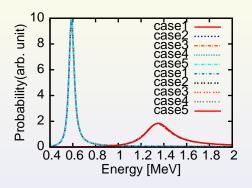
└─Compliment └─The ground state of ²⁶O(preliminary)

Where

$$\begin{split} A(ij,i'j') &= E(^{24}O) + (\epsilon_{id} + \epsilon_{jd})\delta_{ii'}\delta_{jj'} \\ &+ \frac{1}{N_{ij}N_{i'j'}} \left(\frac{5}{2} < sid|V|si'd >_{J=2} + \frac{3}{2} < sid|V|si'd >_{J=1}\right)\delta_{jj'} \\ &+ \frac{1}{N_{ij}N_{i'j'}} \left(\frac{5}{2} < sjd|V|sj'd >_{J=2} + \frac{3}{2} < sjd|V|sj'd >_{J=1}\right)\delta_{ii'} \\ B^{\pm}(ij,i'j') &= \frac{1}{4N_{ij}N_{i'j'}} \left(< ij|V|i'j' >_{J=2} \pm < ij|V|i'j' >_{J=0} \right) \\ C^{\pm}(iji'j') &= \frac{1}{20N_{ij}N_{i'j'}} \left(\left\{ \begin{array}{c} 9\\1 \end{array} \right\} < ij|V|i'j' >_{J=1} \\ &+ \left\{ \begin{array}{c} 1\\9 \end{array} \right\} < ij|V|i'j' >_{J=3} \right) (1 - \delta_{ij})(1 - \delta_{i'j'}) \\ D(ij,i'j') &= \frac{3}{20N_{ij}N_{i'j'}} \left(< ij|V|i'j' >_{J=3} - < ij|V|i'j' >_{J=1} \right) \end{split}$$



(d)
$$A = 24$$
, $nc = 300$



(e)
$$A = 24$$
, $nc = 300$

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