

# **Electromagnetics and Differential Forms**

Metric and  
Hodge star operator

# In today's lecture we talk about:

- **Riemannian manifold**
- Metric volume form
- Hodge star operator
- Constitutive relations

# Riemannian manifold

The pair  $(M, g)$  is called **Riemannian manifold**, where  $M$  is a differentiable manifold and

$$g : \mathcal{X}(M) \times \mathcal{X}(M) \rightarrow C^\infty(M)$$

$$\forall v, w \in \mathcal{X}(M)$$

the **metric tensor** field.

- Locality  $g(v, w)|_P = g(\tilde{v}, \tilde{w})|_P$   $\left\{ \begin{array}{l} \tilde{v}, \tilde{w} \in \mathcal{X}(M) : \\ \tilde{v}_P = v_P, \tilde{w}_P = w_P \end{array} \right.$   

$\tilde{v}_P = v_P$

$\tilde{w}_P = w_P$

smooth continuations
- Symmetry  $g(v, w) = g(w, v)$
- Bilinearity  $g(v + fv', w) = g(v, w) + fg(v', w) \quad \begin{array}{l} \forall v' \in \mathcal{X}(M) \\ \forall f \in C^\infty(M) \end{array}$
- Positive definiteness  $g(v, v) > 0, \quad v \neq 0$

# Metric tensor field

The metric tensor field  $g$  is symmetric bilinear, not to be confused with a 2-form  $\omega \in \mathcal{F}^2(M)$ , which is antisymmetric bilinear.

In a chart with coordinates  $(x^1, \dots, x^n)$  we obtain the **metric coefficients**

$$g_{ij} = g \left( \frac{\partial}{\partial x^i}, \frac{\partial}{\partial x^j} \right) \in \mathcal{C}^\infty(M) \quad \begin{array}{l} i, j = 1, \dots, n \\ n = \dim M \end{array}$$

- Orthonormal basis fields  $g_{ij} = \delta_{ij}$
- In general, it is not possible to choose coordinates such that the basis fields are orthogonal throughout. Obstruction: **curvature**  
Example: breakdown of spherical coordinates in the poles

# Euclidean space $\mathbb{R}^n$

In Euclidean space  $(\mathbb{R}^n, \cdot)$  we let  $g(v, w) := v \cdot w$ .

Since Euclidean space is flat, there exist coordinate systems such that the basis fields are orthogonal,  $g_{ij} = h_i \delta_{ij}$ , where  $0 < h_i \in C^\infty(\mathbb{R}^3)$  are the **scaling functions**.

Coordinate system (in 3 dimensions)	Coordinates $(x^1, x^2, x^3)$	Scaling functions $(h_1, h_2, h_3)$
Cartesian coordinates	$(x, y, z)$	$(1, 1, 1)$
Cylindrical coordinates	$(r, \varphi, z)$	$(1, r, 1)$
Spherical coordinates	$(r, \theta, \varphi)$	$(1, r, r \sin \theta)$

# Musical isomorphisms

The metric tensor field induces the **musical isomorphisms**

$$\flat : \mathcal{X}(M) \rightarrow \mathcal{F}^1(M)$$
$$\sharp = \flat^{-1} : \mathcal{F}^1(M) \rightarrow \mathcal{X}(M)$$

where  $\flat(v)(w) := g(v, w)$ ,  $v \in \mathcal{X}(M)$ ,  $\forall w \in \mathcal{X}(M)$

- In Euclidean space we have  $\flat(v) = {}^1v$
- Isomorphisms  $\flat$  and  $\sharp$  amount to **lowering** and **raising** of indices:

$$v = \underbrace{v^i}_{\text{contravariant components}} \frac{\partial}{\partial x^i} \quad \hookrightarrow \quad \flat(v) = \underbrace{g_{ij} v^i}_{= v_j}_{\text{covariant components}} dx^j$$

Checking this is  
an easy exercise.

# Translation isomorphisms – coordinate-free [\[1\]](#)

RECALL THIS

In  $(\mathbb{R}^3, \cdot, \times)$  we may identify scalar/vector fields with differential forms.

Let  $M \subset \mathbb{R}^3$  be an open set in  $\mathbb{R}^3$  and  $u, v, w \in \mathcal{X}(M)$

$$\begin{aligned} {}^1. & : \mathcal{X}(M) \rightarrow \mathcal{F}^1(M) \\ & a \mapsto {}^1a \end{aligned}$$

$${}^1a(v) := a \cdot v$$

$$\begin{aligned} {}^2. & : \mathcal{X}(M) \rightarrow \mathcal{F}^2(M) \\ & b \mapsto {}^2b \end{aligned}$$

$${}^2b(u, v) := b \cdot (u \times v)$$

$$\begin{aligned} {}^3. & : \mathcal{C}^\infty(M) \rightarrow \mathcal{F}^3(M) \\ & c \mapsto {}^3c \end{aligned}$$

$${}^3c(u, v, w) := c \, u \cdot (v \times w)$$

# A note on the translation isomorphisms

! no summation convention !

Consider Euclidean space  $(\mathbb{R}^3, \cdot, \times)$  equipped with orthogonal coordinates  $(x^1, x^2, x^3)$  such that  $\left(e_i = \frac{1}{h_i} \frac{\partial}{\partial x^i}\right)_{i=1}^3$  is a direct **orthonormal basis** and  $(\varepsilon^j = h_j dx^j)_{j=1}^3$  its dual basis.

Then for the **translation isomorphisms** it holds that

$${}^1e_i = \varepsilon^i = h_i dx^i \quad i = 1, 2, 3$$

$${}^2e_i = \varepsilon^{jk} = h_j dx^j \wedge h_k dx^k \quad ijk = 123, 231, 312$$

$${}^31 = \varepsilon^{123} = h_1 dx^1 \wedge h_2 dx^2 \wedge h_3 dx^3$$

Example: cylindrical coordinates  ${}^1a = a_r dr + a_\varphi r d\varphi + a_z dz$

Physical components  
wrt  $(e_r, e_\varphi, e_z)$



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- Metric tensor
- **Metric volume form**
- Hodge star operator
- Constitutive relations

# Gram determinant

Disclaimer: In the sequel,  $v$  can be either understood as tangent vector,  $v \in T_P M$ , or as vector field  $v \in \mathcal{X}(M)$ ,  $n = \dim M$ .

In the latter case, the constructions are defined pointwise.

Consider a  $p$ -tuple  $(v_1, \dots, v_p)$ ,  $1 \leq p \leq n$ .

Define the **Gram determinant**

$$G = G(v_1, \dots, v_p) := \det(g(v_i, v_k)), \quad i, k = 1, \dots, p$$

Properties:

1.  $G \geq 0$
  2.  $G = 0$  iff  $(v_1, \dots, v_p)$  are linearly dependent
- } Cauchy inequality for  $p = 2$

# Volume of a parallelotope (“Parallelfach”)

Consider the parallelotope spanned by the vectors  $(v_1, \dots, v_p)$ .

Define its **volume** by

$$\text{Vol}(v_1, \dots, v_p) := \sqrt{G}$$

} Insensitive  
to orientation

Examples in Euclidean space  $(\mathbb{R}^3, \cdot, \times)$ :

1. Single vector:  $\text{Vol}(v) = |v|$
2. Parallelogramm:  $\text{Vol}(v_1, v_2) = |v_1 \times v_2|$
3. Parallelepipiped (“Spat”):  $\text{Vol}(v_1, v_2, v_3) = |v_1 \cdot (v_2 \times v_3)|$

# Orientable manifolds and volume forms

Let  $M$  be a  $n$ -dimensional manifold.

A  $n$ -form  $\omega \in \mathcal{F}^n(M)$  on  $M$  is called a **volume form** if  $\omega_P \neq 0 \quad \forall P \in M$  .

A manifold  $M$  is called **orientable** if there exists a volume form on  $M$ .

Remark:

- Not all manifolds are orientable. Example: Möbius strip
- The notion of integration can be extended to non-orientable manifolds. See for example [\[7\]](#).
- If there exists a metric on the manifold, a special volume form is distinct by assigning a unit-volume the value  $\pm 1$ . ( **metric volume form** )

# Metric volume form: definition

Start from an orientable Riemannian manifold  $(M, g)$  and orient it, by picking a volume form  $\omega \in \mathcal{F}^n(M)$ .

Assertion: There exists a unique **metric volume form**  $\omega_M \in [\omega]$  that replies in each point to a direct basis  $(v_1, \dots, v_n)$  with the volume of the associated parallelotope,

$$\omega_M(v_1, \dots, v_n) = +\text{Vol}(v_1, \dots, v_n)$$

Hence we write for an **oriented Riemannian manifold**

$$(M, g, \omega_M)$$

Equivalently, one may ask that  $\omega_M$  assigns each direct orthonormal basis the value +1

# Metric volume form in local coordinates

Consider an oriented Riemannian manifold  $(M, g, \omega_M)$  and a chart  $\phi : U \rightarrow \mathbb{R}^n$ , with local coordinates  $(x^1, \dots, x^n)$ , such that  $\omega_M \sim dx^1 \wedge \dots \wedge dx^n$ . Hence  $\exists \lambda > 0 \in C^\infty(U)$  such that

$$\omega_M = \lambda dx^1 \wedge \dots \wedge dx^n.$$

Assertion: It can be shown that

$$\omega_M = \sqrt{\det(g_{ij})} dx^1 \wedge \dots \wedge dx^n$$

**Invariant volume element**  
in vector analysis

In orthogonal coordinates:

$$\omega_M = h_1 dx^1 \wedge \dots \wedge h_n dx^n$$

Example:  
spherical coordinates  
 $d\tau = r^2 \sin \theta \, dr \, d\theta \, d\varphi$

# Metric volume form in local coordinates: proof

Proof of the assertions.

$$\begin{aligned}\text{Consider } \text{Vol}(v_1, \dots, v_n) &= \sqrt{\det(g(v_k, v_\ell))} \\ &= \sqrt{\det(g_{ij} v_k^i v_\ell^j)} \\ &= \sqrt{\det(g_{ij}) |\det(v_k^i)|} \\ &= \sqrt{\det(g_{ij}) |\det(dx^i(v_k))|} \\ &= \underbrace{\sqrt{\det(g_{ij})}}_{:= \lambda} dx^1 \wedge \dots \wedge dx^n(v_1, \dots, v_n) \\ &= \omega_M(v_1, \dots, v_n) > 0\end{aligned}$$

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- Constitutive relations



# Hodge star operator: basic idea

Consider an oriented Riemannian manifold  $(M, g, \omega_M)$ . The spaces of  $p$ -covectors and  $(n - p)$ -covectors share the same dimension,

$$\dim \bigwedge^p T_P^* M = \binom{n}{p} = \binom{n}{n-p} = \dim \bigwedge^{n-p} T_P^* M$$

↪ The spaces  $\mathcal{F}^p(M)$  and  $\mathcal{F}^{n-p}(M)$  are isomorphic, but not canonically isomorphic.

On an oriented Riemannian manifold, a specific isomorphism is defined by the **Hodge star operator**:

$$\star : \mathcal{F}^p(M) \xrightarrow{\sim} \mathcal{F}^{n-p}(M) : \omega \mapsto \star \omega$$

# Hodge star operator: definition

The Hodge star operator is defined pointwise as follows, where all test vectors  $\xi \in T_P M$ . This extends to differential forms.

$$\star \omega_P(\xi_1, \dots, \xi_q) := \omega_P(\xi_1^\perp, \dots, \xi_p^\perp)$$

$$p + q = n$$

$$n = \dim M$$

Given  $q$ -tupel  
of linearly  
independent  
test vectors

$p$ -tupel  
constructed  
from test  
vectors

↪ The definition boils down to defining the map  $(\xi_1, \dots, \xi_q) \mapsto (\xi_1^\perp, \dots, \xi_p^\perp)$

# Hodge star operator: definition (cont'd)

The map  $(\xi_1, \dots, \xi_q) \mapsto (\xi_1^\perp, \dots, \xi_p^\perp)$  is characterized by three conditions:

1.  $(\xi_1^\perp, \dots, \xi_p^\perp)$  span the orthogonal complement of  $(\xi_1, \dots, \xi_q)$ :

$$\text{span}(\xi_1^\perp, \dots, \xi_p^\perp) = \text{span}(\xi_1, \dots, \xi_q)^\perp \quad \text{i.e.} \quad g(\xi_j^\perp, \xi_i) = 0$$

2.  $(\xi_1^\perp, \dots, \xi_p^\perp), (\xi_1, \dots, \xi_q)$  span the same volume:  $j = 1, \dots, q$   
 $i = 1, \dots, p$

$$\text{Vol}(\xi_1^\perp, \dots, \xi_p^\perp) = \text{Vol}(\xi_1, \dots, \xi_q)$$

3.  $(\xi_1^\perp, \dots, \xi_p^\perp), (\xi_1, \dots, \xi_q)$  form a direct basis of  $T_P M$ :

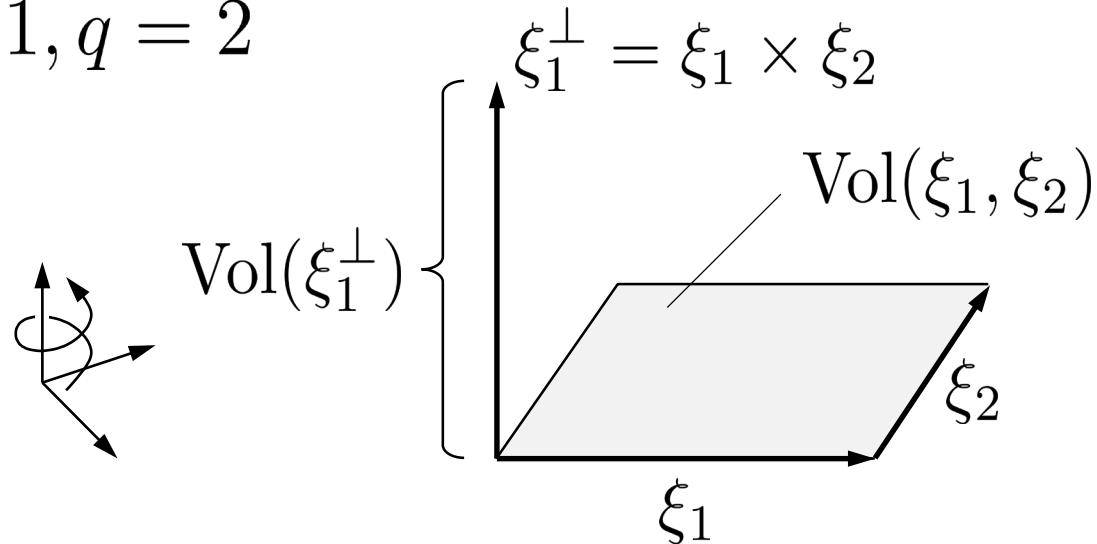
$$\omega_M(\xi_1^\perp, \dots, \xi_p^\perp, \xi_1, \dots, \xi_q) > 0$$

# Hodge star operator: definition (cont'd)

Conditions 1. – 3. fix  $(\xi_1^\perp, \dots, \xi_p^\perp)$  up to a transformation from the special linear group<sup>1</sup>  $\curvearrowright \omega_P(\xi_1^\perp, \dots, \xi_p^\perp)$  uniquely defined

Example in Euclidean space  $(\mathbb{R}^3, \cdot, \times) \curvearrowright \omega_{\mathbb{R}^3} = {}^3 1$

$$p = 1, q = 2$$



1.  $\xi_1^\perp$  perpendicular to  $\xi_1, \xi_2$
2. Length of arrow  $\xi_1^\perp$  equals area of parallelogram  $\xi_1, \xi_2$
3.  $\xi_1^\perp, \xi_1, \xi_2$  right-handed system

<sup>1</sup> A linear transformation whose matrix has determinant +1

# Hodge star operator: properties

For proofs and background information see [1, 5.8], [2, 12.3], [4, 4.2]

- Well-definedness.  $\star \omega_P$  does define a multivector (not obvious<sup>1</sup>)

- $C^\infty(M)$  - linearity

$$\star(\omega + f\eta) = \star\omega + f\star\eta$$

$$\forall \omega, \eta \in \mathcal{F}^p(M)$$

$$\forall f \in C^\infty(M)$$

- Self-inverse property

$$\star\star\omega = (-1)^{pq}\omega$$

in odd dimensions

$$\curvearrowright \star^{-1} = \star$$

- Symmetry

$$\omega \wedge \star\eta = \eta \wedge \star\omega$$

plus positive definiteness

$$\omega \wedge \star\omega \sim \omega_M, \quad \omega \neq 0$$

<sup>1</sup> See material on the learning platform about an equivalent reformulation of the definition.

# Hodge star operator: properties (cont'd)

In coordinates:

$$\omega \cdot \eta = \omega_I \eta_J \det(g^{IJ})$$

- Pointwise inner product. By symmetry and positive definiteness we define

$$\begin{aligned} \mathcal{F}^p(M) \times \mathcal{F}^p(M) &\rightarrow C^\infty(M) : (\omega, \eta) \mapsto \omega \cdot \eta : \\ (\omega \cdot \eta) \omega_M &= \omega \wedge \star \eta \end{aligned}$$

Note also:

$$\star 1 = \omega_M$$

$$\star \omega_M = 1$$

- Hodge and contraction

$$\mathbf{i}_v \star \omega = \star(\omega \wedge \flat(v))$$

$$\forall \omega \in \mathcal{F}^p(M)$$

$$\forall v \in \mathcal{X}(M)$$

- Scaling of metric tensor  $(M, g, \omega_M) \mapsto (M, \lambda g, \lambda^{n/2} \omega_M)$

$$\star \omega \mapsto \lambda^{n/2-p} \star \omega$$

$$0 < \lambda \in C^\infty(M)$$

Interesting case:  $n$  even,  $p = n/2$

# Hodge star operator: coordinate expression

Consider orthogonal coordinates  $(x^1, x^2, x^3)$  in Euclidean space  $(\mathbb{R}^3, \cdot, \times)$ , with metric coefficients  $(g_{ij}) = \text{diag}(h_1^2, h_2^2, h_3^2)$ .

Then it holds

$$\star : \begin{cases} 1 \mapsto h_1 dx^1 \wedge h_2 dx^2 \wedge h_3 dx^3 \mapsto 1 \\ h_i dx^i \mapsto h_j dx^j \wedge h_k dx^k \mapsto h_i dx^i \end{cases} \quad \begin{matrix} ijk = 123, \\ 231, 312 \end{matrix}$$

Proof: easy consequence of

$$\star 1 = \omega_{\mathbb{R}^3}, \quad \star \star = 1,$$

$$\mathbf{i}_{\left(\frac{1}{h_i} \frac{\partial}{\partial x^i}\right)} \star 1 = \star \left(1 \wedge \flat\left(\frac{1}{h_i} \frac{\partial}{\partial x^i}\right)\right) = \star h_i dx^i$$

# Hodge star operator and translation isomorphisms

In Euclidean space  $(\mathbb{R}^3, \cdot, \times)$  it holds that

$$\star^0 f = {}^3 f$$

$$\star^1 a = {}^2 a$$

$$\star^2 b = {}^1 b$$

$$\star^3 g = {}^0 g$$

$$\forall f, g \in C^\infty(M)$$

$$\forall a, b \in \mathcal{X}(M)$$

$$M \subset \mathbb{R}^3 \text{ open}$$

Proof: From the definitions of Hodge star and the translation isomorphisms it follows that

$$\begin{aligned} (\star^0 f)(u, v, w) &= {}^0 f u \cdot (v \times w) = f u \cdot (v \times w) = {}^3 f(u, v, w) & u, v, w \\ (\star^1 a)(u, v) &= {}^1 a(u \times v) = a \cdot (u \times v) = {}^2 a(u, v) & \in \mathcal{X}(M) \end{aligned}$$



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# Constitutive relations for simple media

Consider Euclidean space. **Simple media** are linear and local

↪ tensorial constitutive relations. Tensors shall be time-independent, symmetric positive definite<sup>1</sup>, no electric-magnetic coupling<sup>2</sup>.

If – in addition – the media are homogeneous and isotropic they can be characterized by constant parameters.

Symbol	Parameter	Physical dimension
$\varepsilon$	permittivity	$\text{U}^{-1} \text{I T L}^{-1}$
$\mu$	permeability	$\text{U I}^{-1} \text{T L}^{-1}$
$\sigma$	conductivity	$\text{U}^{-1} \text{I L}^{-1}$

<sup>1</sup> Positive semidefinite in case of conductivity   <sup>2</sup> Excludes in particular moving media and bi-isotropic media [[Link](#)]

# Constitutive relations for simple media (cont'd)

Under these assumptions, the constitutive relations read

Ohm's law

$$\left. \begin{aligned} D &= \varepsilon E \\ B &= \mu H \\ J &= \sigma E \end{aligned} \right\} \iff \left\{ \begin{aligned} \mathcal{D} &= \varepsilon \star \mathcal{E} \\ \mathcal{B} &= \mu \star \mathcal{H} \\ \mathcal{J} &= \sigma \star \mathcal{E} \end{aligned} \right.$$

This is a simple consequence of the translation isomorphisms and Hodge star

Note: We defined the metric tensor and Hodge star dimensionless.

In physics, the metric tensor acquires the **physical dimension**

$\text{pd}(g) = \text{L}^2$ . From the scaling law it follows on  $p$ -forms in  $n$  dimensions

$$\text{pd}(\star) = \text{L}^{n-2p}$$

$$\mathcal{D} = \varepsilon \star \mathcal{E}$$

```
graph TD; D["D"] --- ε["ε"]; D --- star["⋆"]; D --- E["E"]; ε --- IT["IT"]; star --- UL["U⁻¹ITL⁻¹"]; E --- L["L"]; UL --- U["U"]; L --- U;
```

# Constitutive relations for simple media (cont'd)

The material properties can be absorbed in the metric tensor with the help of the scaling law, by choosing  $\lambda$  successively as  $\varepsilon^2, \mu^2, \sigma^2$ .

For 1-forms  $\omega$  in 3 dimensions this yields  $\star_\lambda \omega = \lambda \star \omega$ , hence:

$$\mathcal{D} = \star_\varepsilon \mathcal{E}$$

$$\mathcal{B} = \star_\mu \mathcal{H}$$

$$\mathcal{J} = \star_\sigma \mathcal{E}$$

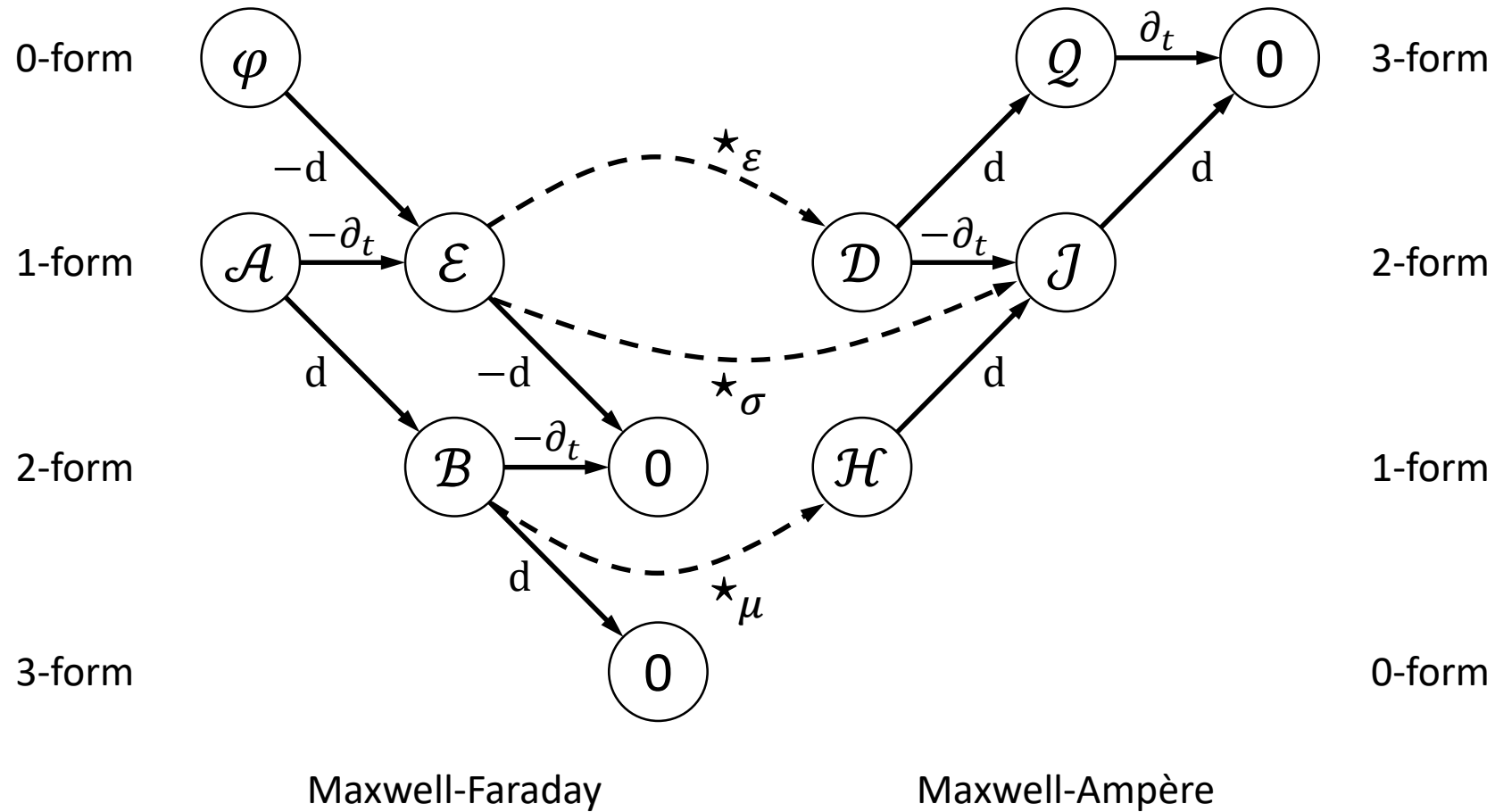
$$\mathcal{D} = \star_\varepsilon \mathcal{E}$$

I T    U<sup>-1</sup> I T    U

! Length is irrelevant !

Note: It can be shown that the requirements of homogeneity and isotropy can be dropped. Simple media can always be represented by adapted metrics and Hodge star operators [15].

# Tonti-diagram in 3D [16, §10.8][17, Table III]



# Literature

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