General Formulas

$$\begin{array}{l} e^{i\theta} = \cos\theta + i\sin\theta \\ -\frac{\hbar^2}{2m} \frac{\delta\Psi(x,t)}{\delta x^2} + V(x)\Psi(x,t) = i\hbar\frac{\delta\Psi(x,t)}{\delta t} \\ \sin\alpha + \sin\beta = 2\cos\frac{\alpha-\beta}{2}\sin\frac{\alpha+\beta}{2} \end{array}$$

Light

 $\mathcal{E}_0 = WaveAmplitude$ n=IndexOfRefractionk=Wavenumbera = Single Slit Width $h{=}Planck'sConstant$ $\omega{=}Wavelength$ T=PeriodK=KineticEnergy $\nu = Ordinary Frequency$ W = WorkFunction $\phi = PhaseShift$ p=momentum $c{=}SpeedOfLightVacuum$ $\mathcal{E} = \mathcal{E}_0 \cos(kx - \omega t) \ _{(1.1)}$

 $k = \frac{2\pi}{\lambda}_{(1.2)}$ $\omega = \frac{2\pi}{T} = 2\pi\nu_{(1.3)}$ $\nu = 1/T$ $\mathcal{E} = \mathcal{E}_0 \cos(kx - \omega t + \phi) \ _{(1.6)}$ $e^{i\theta} = \cos\theta + i\sin\theta_{(1.7)}$ $\omega = kc_{(1.11)}$ $\omega\nu = c_{(1.12)}$ $\frac{\delta^2 \mathcal{E}}{\delta x^2} - \frac{n^2}{c} \frac{\delta^2 \mathcal{E}}{\delta t^2} = 0_{(1.13)}$ $\lambda \nu = \frac{c}{n}_{(1.14)}$ $a\sin\theta = n\lambda_{(minima)}$ Single Slit $E = h\nu_{(1.18)}$ $K = h\nu - W_{(1.19)}$ $h\nu_0 = hc/\lambda_0 = W_{(1.20)}$ $p = \frac{h}{\lambda}_{~(1.21)}$ $\lambda' - \lambda = \frac{h}{mc}(1 - \cos\theta)_{~(1.28)} \text{ Compton}$

The First Principle of Quantum Mechanics The probability of an event = z^*z (1.32) The probability of detecting a particle is equal to z^*z , where z is called the probability amplitude and z^* is its conjugate

The Second Principle of Quantum Mechanics To determine the probability amplitude for a process that can be viewed as taking place in a series of steps we multiply the probability amplitudes for each of these steps. Examples of this are propagation of a photon from a light source or a beam splitter, transmission at the beam splitter, and propagation to a photo detector $z = z_a z_b \cdots (1.38)$

The Third Principle of Quantum Mechanics If there are multiple ways that an event can occur we add the amplitudes for each of these ways.

$$\begin{split} z &= z_1 + z_2 + \cdots_{(1.47)} \\ \phi &= kx \\ z &= x + iy = r\cos\phi + ir\sin\phi = re^{i\phi} \\ z^* &= x - iy = r\cos\phi - ir\sin\phi = re^{-i\phi} \end{split}$$

Wave Mechanics

*In this section we assume a free particle, V(x)=0j=ProbabilityCurrent $\Delta x = Uncertainty$ $\langle x \rangle = Expectation Value$

 $\lambda = \frac{h}{p}~_{(2.1)}$ de Broglie wavelength $d\sin\theta = n\lambda~_{(2.3)~(maxima)}$ Double Slit $x_{n+1} - x_n = \frac{L\lambda}{d} (2.4)$

 $2d\sin\theta = n\lambda_{(2.5)}$ Bragg relation $\begin{aligned} & 2d\sin\theta - \hbar\lambda \ (2.5) \ \text{Bragg relation} \\ & - \frac{\hbar^2}{2m} \frac{\delta\Psi(x,t)}{\delta x^2} + V(x)\Psi(x,t) = i\hbar \frac{\delta\Psi(x,t)}{\delta t} \ (2.6) \\ & - \frac{\hbar^2}{2m} \frac{\delta\Psi(x,t)}{\delta x^2} = i\hbar \frac{\delta\Psi(x,t)}{\delta t} \ (2.7) \\ & \frac{\delta^2 \mathcal{E}}{\delta x^2} = \frac{n^2}{c} \frac{\delta^2 \mathcal{E}}{\delta t^2} \ (2.8) \\ & E = h\nu - \frac{h}{2\pi} 2\pi\nu = \hbar\omega \ (2.9) \\ & p = \frac{h}{\lambda} = \frac{h\pi}{2\pi} \frac{2\pi}{\lambda} = \hbar k \ (2.10) \\ & \hbar \psi = \hbar k c \ (2.11) \end{aligned}$ $\hbar\omega = \hbar k c_{(2.11)}$ $E = pc_{(2.12)}$ $\hbar\omega = \frac{\hbar^2 k^2}{2m} (2.15)$ $p = \frac{\hbar}{\lambda} = \hbar k (2.16)$ $E = h\nu = \hbar\omega_{(2.17)}$ $E = \frac{p^2}{2m}$ (2.18) $|\Psi(x,t)|^2 dx = the \ probability \ of finding \ the$ particle between x and x+dx at the time t if a measurement of the particle's position is carried out $|\Psi(x,t)|^2$ probability density $\int_{-\infty}^{\infty} |\Psi(x,t)|^2 dx = 1_{(2.19)}$ $\frac{\delta|\Psi|^2}{\delta t} = \frac{\Psi^* \Psi}{\delta t} = \Psi^* \frac{\delta \Psi}{\delta t} + \Psi \frac{\delta \Psi^*}{\delta t} (2.20)$ $j_x(x,t) = \frac{\hbar}{2mi} (\Psi^* \frac{\delta \Psi}{\delta x} - \Psi \frac{\delta \Psi^*}{\delta x}) (2.24)$ $\frac{d}{dt} \int_{-\infty}^{\infty} |\Psi(x,t)|^2 dx = -j_x(x,t)|_{-\infty}^{\infty} = 0$ $\Psi(x,t) = \int_{-\infty}^{\infty} A(k)e^{i(kx-\omega t)}dk$ (2.29) $\Delta x \Delta k \geq \frac{1}{2} (2.30)$ $\Delta x \Delta p_x \geq \frac{\hbar}{2} (2.31) \text{ Heisenberg}$ $v_{ph} = \frac{\omega}{k} = \frac{2\pi\nu}{(2\pi/\lambda)} = \lambda\nu (2.33)$ The phase velocity is the speed at which a point on the wave, such as a crest, moves. $v_{ph} = \frac{\omega}{k} = \frac{\hbar\omega}{\hbar k} = \frac{E}{p} = \frac{mv^2/2}{mv} \frac{v}{2}$ (2.34) $v_q = \frac{d\omega}{dk} (2.36)$ The group velocity is the speed of a localized packet of waves that has been generated by superposing many waves together $\Psi(x,t) = \int_{-\infty}^{\infty} A(k)e^{i(kx-\omega t)}dk$ (2.37) $\omega \cong \omega_0 + v_g(\tilde{k} - k_0) \tag{2.39}$ Dispersion relation is the relationship between ω and k $\langle x \rangle = \int_{-\infty}^{\infty} x |\Psi(x,t)|^2 dx$ (2.53) The average values $\langle x \rangle$ are referred to as the expectation values $\langle x^2 \rangle = \int_{-\infty}^{\infty} x^2 |\Psi(x,t)|^2 dx$ (2.55) $(\Delta x)^2 = \langle x^2 \rangle - \langle x \rangle^2 \ _{(2.56)}$ Δx , the standard deviation, is also called the uncertainty $(\Delta p_x)^2 = \langle p_x^2 \rangle - \langle p_x \rangle^2$ (2.57) $\frac{\frac{d\langle x \rangle}{dt}}{dt} = \frac{\langle p_x \rangle}{m} (2.58)$ $\langle p_x \rangle = \int_{-\infty}^{\infty} \Psi^* \frac{\hbar}{i} \frac{\delta \Psi}{\delta x} dx$ (2.63) $\frac{d\langle p_x \rangle}{dt} = \langle -\frac{\delta V}{\delta x} \rangle_{(2.64)}$ $\langle p_x \rangle = \int_{-\infty}^{\infty} \Psi \frac{\hbar}{i} \frac{\delta \Psi}{\delta x} dx$ $\Delta p_x = (\langle p_x^2 \rangle - \langle p_x \rangle^2)^{1/2}$ Schrodinger equation for a free particle: $\Psi(x,t) = Ae^{i(kx-\omega t)}$ Where: $p = \hbar k = h/\lambda$ $E = \hbar\omega - h\nu$ $\Psi(x,t) = \int_{-\infty}^{\infty} A(k)e^{i(kx - \omega t)}dk$ Wave move with group velocity: $v_q = d\omega/dk = \hbar k/m = p/m$

 $\omega = \hbar k^2/2m$

The Time-Independent Schrödinger Equation

*In this section we assume V(x) is independent of t $\delta_{nm} = Kronecker Delta$ $\psi_a = Eigenfunction$ a=EigenvalueT=TransmissionCoef. $\Psi(x,t) = \psi(x)f(t) (3.2)$ $\frac{\delta^2 \Psi(x,t)}{\delta x^2} = f(t) \frac{d^2 \psi(x)}{dx^2} \quad (3.3)$ $\frac{\delta \Psi(x,t)}{\delta t} = \psi(x) \frac{df(t)}{dt} \quad (3.4)$ $\frac{df(t)}{dt} = \frac{-iE}{\hbar} f(t) \quad (3.8)$ $-\frac{\hbar^2}{2m} \frac{\delta \psi(x)}{\delta x^2} + V(x) \psi(x) = E\psi(x) \quad (3.9)$ $f(t) = f(0) e^{-iEt/\hbar} \quad (3.10)$ $f(t) = f(0)e^{-i\omega t}$ (3.11) $E = \hbar \omega_{(3.12)}$ $\Psi(x,t) = \psi(x)e^{-iEt/\hbar}$ (3.13) $|\Psi(x,t)|dd = |\psi(x)|^2$ (3.14)

$$V(x) = \begin{cases} 0, & 0 < x < L. \\ \infty, & \text{elsewhere.} \end{cases}$$

 $\begin{array}{l} -\frac{\hbar^2}{2m}\frac{\delta\psi}{\delta x^2} = E\psi_{~(3.16)}~0 < x < L \\ k^2 = \frac{2mE}{\hbar^2}~_{(3.17)} \\ \psi(x) = A\sin kx + B\cos kx~_{(3.21)}~0 < x < L \end{array}$ $k_n = \frac{n\pi}{L} (3.26)$ $E_n = \frac{\hbar k_n^2}{2m} = \frac{n^2 \hbar^2 \pi^2}{2mL^2} (3.27)$ $\psi(x) = A_n \sin \frac{n\pi x}{L} (3.28) \quad 0 < x < L$

$$\psi_n(x) = \begin{cases} \sqrt{\frac{2}{L}} \sin \frac{n\pi x}{L}, & 0 < x < L. \\ 0, & \text{elsewhere.} \end{cases}$$

 $\Psi(x) = c_1 \psi_1(x) + c_2 \psi_2(x)$ $c_1(t) = \frac{1}{\sqrt{2}} e^{-iE_1 t/\hbar}$ (3.39) $\Psi = \sum_{n=1}^{\infty} c_n \psi_n(x)$ (3.40)

$$\delta_{nm} = \begin{cases} 1, & m = n. \\ 0, & m \neq n. \end{cases}$$

 ψ_n 's are orthanormal if they satisfy: $\int_{-\infty}^{\infty} \psi_x^*(x) \psi_n(x) dx = \delta_{nm \ (3.49)}$ $|c_n|^2 = P_{n-(3.59)}$

The above is the probability of obtaining E_n if a measurement of the energy of a particle with wave function Ψ is carried

out
$$\langle E \rangle = \sum_{n=1}^{\infty} |c_n|^2 E_{n (3.61)}$$

$$c_n = \int_{-\infty}^{\infty} \psi^*(x) \psi(x) dx$$

$$H \psi = E \psi$$

Where H is the energy operator:

$$H = \frac{(p_{xop})^2}{2m} + V(x)$$

$$A_{op}\psi_a = a\psi_a \ (3.63)$$

$$x_{op} = x \ (3.64)$$

$$p_{xop} = \frac{\hbar}{i} \frac{\delta}{\delta x} \ (3.65)$$

$$E_{op} = \frac{(p_{xop})^2}{2m} + V(x_{op}) \ (3.71)$$

$$H \equiv E_{op} = -\frac{\hbar^2}{2m} \frac{\delta^2}{\delta x^2} + V(x) \ (3.72)$$

$$\langle E \rangle = \int_{-\infty}^{\infty} \Psi^* H \Psi dx \ (3.81)$$

One-Dimensional Potentials

Energy eigenfunctions are oscillatory in

regions where E>V and exponential in V>E.

$$V(x) = \begin{cases} 0, & |x| < a/2. \\ V_0, & |x| > a/2. \end{cases}$$

$$k = \frac{\sqrt{2mE}}{\hbar}$$

$$\psi(x) = Ae^{ikx} + Be^{-ikx} |x| < a/2$$

$$\kappa = \frac{\sqrt{2m(V_0 - E)}}{\hbar} > 0$$

Principles of Quantum Mechanics

Constants

$$\hbar = 6.582 \times 10^{-16}$$

$$\epsilon_0 = 8.854 \times 10^{-12}$$

$$e = 1.602 \times 10^{-19}$$

$$k_B = 8.617 \times 10^{-5}$$

 $\Psi^*\Psi dx$ is the probability of finding the particle between x and x + dx

$$\int_{-\infty}^{\infty} \Psi^* \Psi dx = 1_{(5.97)}$$

A Hermitian operator satisfies:

$$\int_{-\infty}^{\infty} \phi^* A_{op} \psi dx = \int_{-\infty}^{\infty} (A_{op} \phi)^* \psi dx$$
 (5.98)

$$A_{op}\psi_a = a\psi_{a\ (5.99)}$$

Orthonormal wave functions satisfy:

$$\Psi = \sum_{a} c_a \psi_{a (5.100)}$$

Probability of obtaining a:

$$|c_a|^2 = \left| \int_{-\infty}^{\infty} \psi_a^* \Psi dx \right|^2$$
 (5.101)

Average or expectation value:

$$\langle A \rangle = \sum_{a} |c_a|^2 a = \int_{-\infty}^{\infty} \Psi^* A_{op} \Psi dx$$
 (5.102)
Commutator:

$$[A_{op}, B_{op}] = A_{op}B_{op} - B_{op}A_{op}$$
(5.103)

If:
$$[A_{op}, B_{op}] = iC_{op}$$
 (5.104)

Then:
$$\Delta A \Delta B \ge \frac{\left|\langle C \rangle\right|}{2}$$
 (5.105)

$$\Delta x \Delta p_x \ge \frac{\hbar}{2} (5.106)$$

$$[x_{op}, p_{x_{op}}] = i\hbar_{(5.107)}$$

$$H\Psi(x,t) = i\hbar \frac{\delta \Psi(x,t)}{\delta t}$$
(5.108)

Hamiltonian, the energy operator:
$$H=-\frac{\hbar^2}{2m}\frac{\delta^2}{\delta x^2}+V(x)_{(5,109)}$$

Expectation values $\frac{d\langle A \rangle}{dt} =: \frac{i}{\hbar} \int_{-\infty}^{\infty} \Psi^* [H, A_{op}] \Psi dx + \int_{-\infty}^{\infty} \Psi^* \frac{\delta A_{op}}{\delta t} \Psi dx$ If the Hamiltonian commutes withe the operator corresponding to the observable A and $\delta A_{op}/\delta t = 0$, then $\langle A \rangle$ is

independent of time.

$$\Delta E \frac{\Delta A}{\left|\frac{d < A >}{dt}\right|} \ge \frac{\hbar}{2} \tag{5.111}$$

 $\Delta E \Delta t \geq \frac{\hbar}{2} (5.112)$

Parity operator:

$$\Pi \psi(x) = \psi(-x)$$
 (5.4)

If there exists a single eigenfunction with eigenvalue a then its nondegenerate.

$$x_{op} = x_{(3.64)}$$

$$p_{x_{op}} = \frac{\hbar}{i} \frac{\delta}{\delta x}$$
 (3.66)

Quantum Mechanics in Three Dimensions

Cubic Box:

Cubic Box:
$$E_{n_x,n_y,n_z} = \frac{(n_x^2 + n_y^2 + n_z^2)\hbar^2\pi^2}{2mL^2} \text{ (6.104)}$$
 Hydrogenic Atom:

$$E_n = -\frac{mZ^2e^4}{(4\pi\epsilon_0)^22\hbar^2n^2} = -\frac{(1.36eV)Z^2}{n^2}$$
(6.105)

$$\psi(x) = Ce^{\kappa x} + De^{-\kappa x} \ |x| > a/2$$

$$\psi(x) = \begin{cases} Ce^{\kappa x}, & x \le -a/2. \\ 2A\cos kx, & -a/2 \le x \le a/2. \\ Ce^{-\kappa x}, & x \ge a/2. \end{cases}$$

$$V(x) = \begin{cases} 0, & x < 0. \\ V_0, & x > 0. \end{cases}$$

Allowed values of l

$$l = 0, 1, 2, ..., n - 1$$

$$\nabla^2 = \frac{\delta^2}{\delta x^2} + \frac{\delta^2}{\delta y^2} + \frac{\delta^2}{\delta z^2}$$
$$\nabla^2 =$$

$$| \nabla^2 |$$

 $\frac{1}{r^2}\frac{\delta}{\delta r}(r^2\frac{\delta}{\delta r}) + \frac{1}{r^2\sin\theta}\frac{\delta}{\delta\theta}(\sin\theta\frac{\delta}{\delta\theta}) + \frac{1}{r^2\sin^2\theta}\frac{\delta^2}{\delta\phi^2}$ Angular momentum operator:

$$L_{op}^2 = L_{x_{op}}^2 + L_{y_{op}}^2 + L_{z_o}^2$$

$$\begin{split} L_{op}^2 &= L_{x_{op}}^2 + L_{y_{op}}^2 + L_{z_{op}}^2 \\ L_{op}^2 Y_{l,m_l}(\theta,\phi) &= l(l+1)\hbar^2 Y_{l,m_l}(\theta,\phi) \ _{(6.106)} \\ L_{op}^2 &= -\hbar^2 \big[\frac{1}{\sin\theta} \frac{\delta}{\delta\theta} \big(\sin\theta \frac{\delta}{\delta\theta}\big) + \frac{1}{\sin^2\theta} \frac{\delta^2}{\delta\phi^2} \big] \\ L &= r \times p \end{split}$$

$$L = r \times p$$

 $L_x = yp_z - zp_y$

The eigenfunctions Y_{l,m_l} , the spherical harmonics, also satisfy:

 $L_{op}Y_{l,m_l}(\theta,\phi) = m_l \hbar Y_{l,m_l}(\theta,\phi)$ (6.107)

The energies (6.105) are independent of m_l because of the rotational symetry of the potential energy $-Ze^2/4\pi\epsilon_0 r$

The maximum value for L_z is $l\hbar$ which is always less than the total angular

momentum
$$\sqrt{l(l+1)}\hbar$$

$$\begin{bmatrix} L_{x_{op}}, L_{y_{op}} \end{bmatrix} = i\hbar L_{z_{op}}$$

$$\begin{bmatrix} L_{yop}, L_{zop} \end{bmatrix} = i\hbar L_{xop}$$
$$\begin{bmatrix} L_{zop}, L_{xop} \end{bmatrix} = i\hbar L_{yop}$$
(6.108)

$$\Delta L_x \Delta_y \ge \frac{\hbar}{2} |\langle L_z \rangle|_{(6.109)}$$

$$L_{z_{op}} = \frac{\hbar}{i} \frac{\delta}{\delta \phi}$$

Spin angular momentum S:

 $\chi \pm$ are two dimensional column vectors:

$$S_{op}^2 \chi \pm = s(s+1)\hbar^2 \chi \pm s = 1/2$$
 (6.110)

$$S_{z_{op}}\chi \pm = \pm \frac{\hbar}{2}\chi \pm _{(6.111)}$$

The magnetic moment:

$$\mu = 2.00232 \left(\frac{-e}{2m}\right) S_{(6.112)}$$

The spherical harmonics with

$$l=0, l=1, \underbrace{an\underline{dl}}=2$$

$$Y_{0,0}(\theta,\phi) = \sqrt{\frac{1}{4\pi}}$$

$$Y_{1,\pm 1}(\theta,\phi) = \sqrt{\frac{3}{8\pi}} \sin \theta e^{\pm i\theta}$$

$$Y_{1,0}(\theta,\phi) = \sqrt{\frac{3}{4\pi}}\cos\theta$$

$$Y_{2,\pm 2}(\theta,\phi) = \sqrt{\frac{15}{32\pi}} \sin^2 \theta e^{\pm 2i\theta}$$

$$Y_{2,\pm 1}(\theta,\phi) = \mp \sqrt{\frac{15}{8\pi}} \sin \theta \cos \theta e^{\pm i\theta}$$

$$Y_{2,0}(\theta,\phi) = \sqrt{\frac{5}{16\pi}} (3\cos^2\theta - 1)$$

These spherical harmonics are the eigenfunctions.

For a three dimensional box from 0 to Lthe energy eigenfunctions (ψ_{n_x,n_y,n_z}) are:

$$\left(\frac{2}{L}\right)^{3/2} \sin \frac{n_x \pi x}{L} \sin \frac{n_y \pi y}{L} \sin \frac{n_z \pi z}{L}$$
The Energy Eigenvalues are:

$$E_{n_x,n_y,n_z} = \frac{(n_x^2 + n_y^2 + n_z^2)\hbar^2 \pi^2}{2mL^2}$$

$$\begin{split} k &= \frac{\sqrt{2mE}}{\hbar} \\ k_0 &= \sqrt{k^2 - \frac{2mV_0}{\hbar^2}} \\ \psi(x) &= \begin{cases} Ae^{ikx} + Be^{-ikx}, & x < 0. \\ Ce^{ik_0x}, & x > 0. \end{cases} \\ j_x &= \begin{cases} \frac{\hbar k}{m} (|A|^2 - |B|^2), & x < 0. \\ \frac{\hbar k_0}{m} |C|^2, & x > 0. \end{cases} \end{split}$$

$$T \cong \left(\frac{4\kappa k}{k^2 + \kappa^2}\right)^2 e^{-2\kappa a}$$

Identical Particles

Bosons are symmetric under particle exchange, Fermions are anti-symmetric. Bosons have integral intrinsic spin,

Fermions have half integral intrinsic spin. Valid Boson states:

$$\Psi_S(1,2) = \psi_{\alpha}(1)\psi_{\alpha}(2)$$
 (7.104)

$$\Psi_S(1,2) = \frac{1}{\sqrt{2}} [\psi_{\alpha}(1)\psi_{\beta}(2) + \psi_{\beta}(1)\psi_{\alpha}(2)]$$

Bosons are more likely to be in the same state than are distinguishable particles.

The average number of identical Bosons in a state with energy E in thermal equilibrium at temperature T is given by: $n(E) = \frac{1}{e^{(E-\mu)/k_b T} - 1}$ (7.106)

Where μ is the chemical potential.

Valid Fermions:

$$\Psi_A(1,2) = \frac{1}{\sqrt{2}} \left[\psi_\alpha(1)\psi_\beta(2) - \psi_\beta(1)\psi_\alpha(2) \right]$$

The average number of identical Fermions in a state with energy E in thermal equilibrium at temperature T is given by:

 $n(E) = \frac{1}{e^{(E-E_F)/k_b T} + 1}$ (7.108)

Where \tilde{E}_F is the Fermi energy. Photons are Bosons, thus the distribution of electromagnetic energy in a cavity:

$$\rho(\nu)d\nu = \frac{8\pi h\nu^3}{c^3(e^{h\nu/k_bT}-1)} (7.109)$$

Total energy per unit area per unit time:

$$\sigma T^4_{(7.67)}$$

$$\sigma = \hat{5}.67 \times 10^{-8} \,_{(7.68)}$$

$$\lambda_{max}T = 2.9 \times 10^{-3} (7.69)$$

$$E_F = \frac{1}{2}mV_F^2$$

$$E_F = \frac{\hbar^2}{2m} \left(\frac{3\pi^2 N}{V}\right)^{2/3}$$

$$E_{total} = \frac{3}{5} N E_F$$

Degeneracy Pressure:

$$P = -\frac{dE_{total}}{dV} = \frac{2}{5} (3\pi^2)^{2/3} \frac{\hbar^2}{2m} (\frac{N}{V})^{5/3}$$

Compressibility:

$$K = (-V \frac{dP}{dV})^{-1}$$

Ways of choosing N_1 from group of N: $\frac{N!}{N_1!(N-N_1)!}$

Fermi temperature: $E_F = k_B T_F$ $I = I_0(e^{e\varphi/k_bT - 1})$

e in exponent is charge on electron.

Solid State Physics

Bloch ansatz:

$$\psi(x+a) = e^{i\theta}\psi(x) (8.4)$$

$$\sqrt{2mE} - k$$

 $\frac{\sqrt{2mE}}{\hbar} = k_{(8.10)}$ $\cos \theta = \cos ka + \frac{\alpha \sin ka}{2ka}_{(8.26)}$ W

Contact Potential: $(W_b - W_a)/e$ For semiconductors: $E_F = \frac{E_g}{2}$ (8.33)

$$n(E) = \frac{1}{e^{E_g/2k_bT} + 1}$$
 (8.34)

Nuclear Physics

 $m_{nucleus} = Zm_p + (A - Z)m_n - B.E./c^2$ Where $B.E. \times (1, 0, -1)$ is:

$$a_1 A - a_2 A^{2/3} - a_3 \frac{Z^2}{A^{1/3}} - a_4 \frac{(z - \frac{A}{2})^2}{A} + \frac{a_5}{\sqrt{A}}$$

$$N(t) = N(0)e^{-Rt} = N(0)e^{-t/\tau}$$
(9.79)

Radius of a nucleus composed of A nucleons:

$$R = r_0 A^{1/3}$$

$$r_0 = 1.2 fm = 12 \times 10^{-15}$$

The lifetime of an unstable nucleus:

$$\tau = 1/R$$

$$t_{1/2} = \tau \ln 2$$

$$a_{1}^{'} = 15.75 MeV$$

$$a_2 = 17.8 MeV$$

$$a_3 = 0.711 MeV$$

$$a_4 = 94.8 MeV$$

$$a_5 = 11.2 MeV$$

Relativity

$$c = 3 \times 10^8 m/s$$

Conservation of Kinetic Energy is not a thing.

Galilean relativity:

$$t' = t$$

$$x' = x - \beta t$$

Coordinate time is Δt between two events recorded by a pair of clocks synchronized at rest in a given inertial reference frame (one clock present at each event). This is basically the difference on the time axis.

$$\Delta t = t_A - t_B$$

The spacetime interval Δs can never be larger than the coordinate time Δt .

Coordinate time Δt

 $t_1 - t_2$ as measured by clocks in any inertial frame.

The clocks for the two events could be different.

Many possible values for this (many inertial frames).

Proper time $\Delta \tau$

Measured by a single clock present at both

This clock could be in a non-inertial frame.

Many possible different values for this (many possible clock paths).

Spacetime Interval Δs

 $t_1 - t_2$ as measured by one clock present at both events and attached to an inertial

Unique for any pair of events.

$$\Delta t > \Delta s > \Delta \tau$$

$$\Delta \tau = \int \sqrt{1 - v^2} dt$$

$$\Delta s^2 = \Delta t^2 - \Delta d^2$$

Spacetime interval is the same in any inertial frame, the others are different. If speed is constant: $\Delta \tau = \sqrt{1 - v^2} \Delta t$ Lorentz Transformation Equations $t' = \gamma(t - \beta x)$

$$x' = \gamma(x - \beta t)$$

Lorentz Contraction

$$L = \sqrt{1 - \beta^2} L_R$$

$$\Delta s^2 \geq 0$$
 Time-like.

 $\Delta s^2 = 0$ Light-like. $\Delta s^2 < 0$ Space-like.

If casually connected the order matters, otherwise the order can be switched in different reference frames.

Einstein Velocity Transformation

$$v_x' = \frac{v_x - \beta}{1 - \beta v_x}$$

$$v_y' = \frac{v_y \sqrt{1-\beta}}{1-\beta v}$$

$$v_z' = \frac{v_z \sqrt{1-\beta^2}}{1-\beta v}$$

 $\begin{aligned} v_y' &= \frac{v_y\sqrt{1-\beta^2}}{1-\beta v_x} \\ v_z' &= \frac{v_z\sqrt{1-\beta^2}}{1-\beta v_x} \\ \text{For inverse transformation swap primes} \end{aligned}$ with unprime and change β to $-\beta$.

$$\vec{p}_{tot} = \sum m_i \begin{bmatrix} v_{ix} \\ v_{iy} \\ v_{iz} \end{bmatrix}$$

Four-momentum:

$$P_{tot} = \sum \frac{m_i}{\sqrt{1 - v_i^2}} \begin{bmatrix} 1\\ v_{ix}\\ v_{iy}\\ v_{iz} \end{bmatrix}$$

At low speeds:

At low speeds:
$$P = \sum \frac{m}{\sqrt{1-v^2}} \begin{bmatrix} 1 \\ v_x \\ v_y \\ v_z \end{bmatrix} \approx \begin{bmatrix} m + \frac{1}{2}mv^2 \\ mv_x \\ mv_y \\ mv_z \end{bmatrix}$$
 Relativistic K: $k = \frac{m}{\sqrt{1-v^2}} - m$

Relativistic K: $k = \frac{m}{\sqrt{1-v^2}}$ Binomial approximation:

$$(1+x)^{\alpha} \approx 1 + \alpha x$$

$$\begin{array}{ccc}
f & 10^{-15} \\
p & 10^{-12}
\end{array}$$

$$n = 10^{-9}$$

$$\mu = 10^{-6}$$

$$m = 10^{-3}$$

$$M 10^6$$

Name Solution:

$$\Psi = \frac{e^{i\pi}}{2}\psi_1 + \frac{\sqrt{3}}{2}\psi_2$$

$$Z_{ABO} = \frac{1}{12} e^{i\pi} \frac{1}{12} e^{i\kappa k_1} = \frac{1}{2} e^{i\kappa k_2}$$

Probability of a photon arriving at PM, ...
$$P_{ano} = (Z_{AB,O}^{*} + Z_{ACO}^{*})(Z_{ABO} + Z_{ACO})$$

$$=\frac{1}{4}\left(2-e^{-ik(\hat{x}_1-\hat{x}_2)}-e^{-ik(\hat{x}_1-\hat{x}_2)}\right)$$

$$(6ets) \qquad \frac{1}{e^{(mx)}} = \frac{1}{2} \left[1 - \cos\left(k(\lambda_1, \lambda_2)\right) \right]$$

 $E_{E} = \frac{1}{2} M_{0} V_{E}^{2}$ $E_{E} = \frac{\pi^{2}}{2} \left(\frac{3 \pi^{2} N}{V} \right)^{2/3}$

DAR FE is known can relentate VE

$$\frac{N}{V} = \frac{0.975}{1 \text{ cm}^3} \times \frac{1 \text{ m/s}}{225} \times \left(\frac{190 \text{ cm}}{1 \text{ m}}\right)^3 \times \frac{(.923 \times 10^{23} \text{ parkish})}{1 \text{ m/k}}$$

$$= 2.54 \times 10^{27} \text{ gleeters}$$

5.) N-type semiconductors lare diged to have more electrons in the conduction band This means Ex is higher than in pure Silican

P-type summenductors are depend to have more holes in the valence band. This means Ex is lower than in pure silven.

When a place of P-type is joined to a prea of N-type, a contect potential Sorms. The P-type and up at higher potential energy

> Formed biasing pushes the plentiful N-type conduction elections over the contact potential. Current Flows Freely.

Reverse bissing pushes whatever Prtype conduction aluctions there are found the Nortyke side. Aren't many p-tyke conduction electrons. Small amount of current (affectively none) Flows