

# Ladder operators and quasinormal modes in Banados-Teitelboim-Zanelli black holes

Based on 2205.15610 (T.K. and Masashi Kimura)

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# Outline

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1. Introduction
2. Review: Mass ladder operators
3. Review: Quasinormal modes
4. Quasinormal modes in Banados-Teitelboim-Zanell spacetimes
5. Mass ladder operators in Banados-Teitelboim-Zanell spacetimes
6. Shift of quasinormal mode frequencies

# Black hole perturbations

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- Test fields in BH spacetimes play an important role in the understanding of phenomena in the strong-gravity regime

e.g., gravitational waves from binary black holes,  
time evolution of ultralight bosons around black holes,  
linear stabilities of black hole spacetimes,  
relaxation phenomena within AdS/CFT, ...

- In many problems, master equations take a form of Schrodinger eq:

$$\left[ \frac{d^2}{dx^2} + \omega^2 - V(x) \right] \phi = 0$$

- Mathematical tools in quantum mechanics can be useful in BH perturbation theory  
e.g., ladder operators

# Ladder operators

- In quantum mechanics,  
ladder operators allow to relate the different energy eigenvalues
- Ladder operators in curved spacetimes: **Mass ladder operators** [Cardoso et al, 2017]  
[Cardoso et al, 2018]

$$[\square - \mu^2] \Phi = 0 \quad \rightarrow \quad [\square - (\mu^2 + \delta\mu^2)] D\Phi = 0$$

Mass ladder operator  $D$  maps a Klein-Gordon field onto another Klein-Gordon field  
This is constructed from spacetime conformal symmetry

Question:

Does a mass ladder operator keep physics determined by boundary conditions?

This work: Application of mass ladder operators to black hole physics

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# Review: Mass Ladder operators from spacetime conformal symmetry

- Material: spacetime conformal symmetry

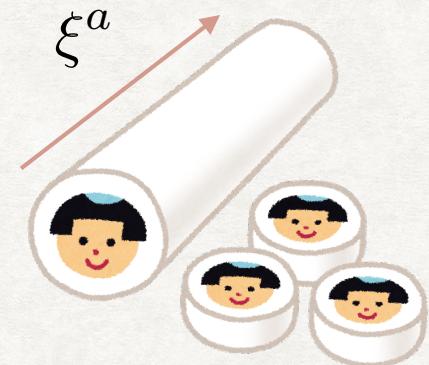
[Wald's textbook]

*Definition: spacetime symmetry*

Spacetime  $(\mathcal{M}, g_{ab})$  possesses symmetry

iff  $g_{ab}$  admits an isometry defined by  $\varphi_t : \mathcal{M} \rightarrow \mathcal{M}$  such that  $\varphi_t^* g_{ab} = g_{ab}$

Isometry group is generated by  $x^a \rightarrow \bar{x}^a - \xi^a$   
along a **Killing vector field** that satisfies  $\mathcal{L}_\xi g_{ab} = 0$

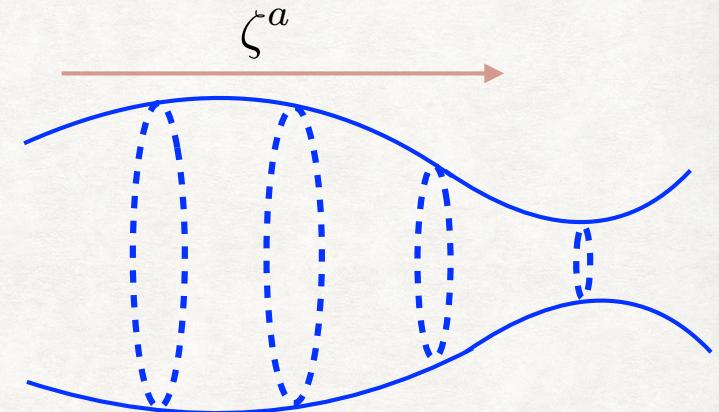


Kintaro-ame

# Review: Mass Ladder operators from spacetime conformal symmetry

- What is spacetime conformal symmetry?

Generalization of spacetime symmetry



[Wald's textbook]

*Definition: spacetime conformal symmetry*

Spacetime  $(\mathcal{M}, g_{ab})$  possesses conformal symmetry

iff  $g_{ab}$  admits a conformal isometry defined by  $\varphi_t : \mathcal{M} \rightarrow \mathcal{M}$

such that  $\varphi_t^* g_{ab} = \exp(2Q) g_{ab}$ , where  $Q$  is a function on  $\mathcal{M}$

Conformal isometry group is generated by  $x^a \rightarrow \bar{x}^a - \zeta^a$   
along a **conformal Killing vector field** that satisfies  $\mathcal{L}_\zeta g_{ab} = 2Q g_{ab}$

- \* Conformal Killing vector field is  
a Killing vector field of  $\tilde{g}_{ab} = \Omega^2 g_{ab}$  with  $Q = -2\zeta^a \nabla_a \ln \Omega$

# Construction of mass Ladder operators

[Cardoso et al, 2017]

[Cardoso et al, 2018]

- “Closed” condition for conformal Killing vector fields:  $\nabla_a \zeta_b = \nabla_b \zeta_a$
- Assumption:  $R^a_b \zeta^b = \chi(n - 1)\zeta^a$ ,  $\chi \in \mathbb{R}$   
e.g., n-dimensional (anti-) de Sitter spacetimes with  $\chi = \frac{\Lambda}{n - 1}$
- **Mass ladder operator**:  $D_k := \mathcal{L}_\zeta - \frac{k}{n} (\nabla_\sigma \zeta^\sigma)$ ,  $k \in \mathbb{R}$
- Commutation relation holds:  $[\square, D_k] = \chi(2k + n - 2) D_k + 2Q(\square + \chi k(k + n - 1))$

$$\Rightarrow D_{k-2} [\square - \mu^2] \Phi = [\square - (\mu^2 + \delta\mu^2)] D_k \Phi, \quad \begin{aligned} \mu^2 &= -\chi k(k + n - 1), \\ \delta\mu^2 &= \chi(2k + n - 2) \end{aligned}$$

If  $\Phi$  is a Klein-Gordon field with  $\mu^2$ ,  
 $D_k \Phi$  is also a Klein-Gordon field but with  $\mu^2 + \delta\mu^2$

Note: Mass ladder operator is an onto map

# Mass Ladder operators

[Cardoso et al, 2017]

[Cardoso et al, 2018]

- Mass ladder operator connects Klein-Gordon fields with different mass squared:

$$[\square - \mu^2] \Phi = 0 \longrightarrow [\square - (\mu^2 + \delta\mu^2)] D_k \Phi = 0$$

$\mu^2 = -\chi k(k+n-1),$

$\delta\mu^2 = \chi(2k+n-2)$

- $k$  is required to be real, so  leads to inequalities:

$$\mu^2 \geq \frac{\chi}{4} (n-1)^2 \quad (\text{for } \chi < 0), \quad \mu^2 \leq \frac{\chi}{4} (n-1)^2 \quad (\text{for } \chi > 0)$$

Note: In AdS case ( $\chi < 0$ ) , the lower bound coincides with the BF bound

- When parametrizing  $\mu^2 = -\chi\nu(\nu+n-1)$  ( $\nu \geq -1$ ),

leads to two solutions,  $k_+ = -n+1-\nu$ ,  $k_- = \nu$

$$[\square + \chi\nu(\nu+n-1)] \Phi = 0$$

$\xrightarrow{\hspace{100pt}}$

$[\square + \chi\tilde{\nu}(\tilde{\nu}+n-1)] D_{k_+} \Phi = 0$

$\tilde{\nu} = \nu + 1$  mass raising (lowering) for  $\chi < 0$  ( $> 0$ )

$\xrightarrow{\hspace{100pt}}$

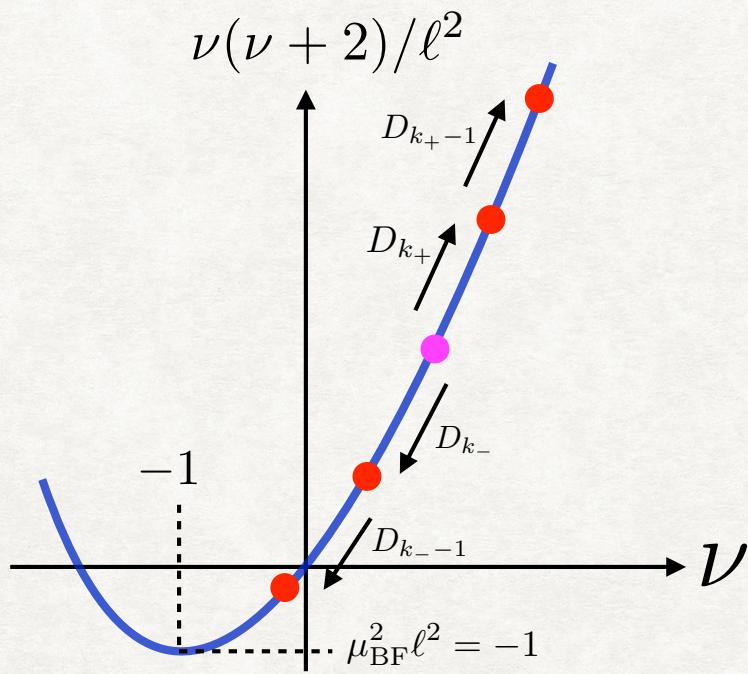
$[\square + \chi\tilde{\nu}(\tilde{\nu}+n-1)] D_{k_-} \Phi = 0$

$\tilde{\nu} = \nu - 1$  mass lowering (raising) for  $\chi < 0$  ( $> 0$ )

# Mass shift

Example:  $[\square - \mu^2] \Phi = 0$  on AdS<sub>3</sub> with length scale  $\ell := \sqrt{-1/\Lambda}$

Mass parametrization:  $\mu^2 = \nu(\nu + 2)/\ell^2$  ( $\nu \geq -1$ )



Mass ladder operators  $D_{k_\pm}$  make mass squared raise or lower

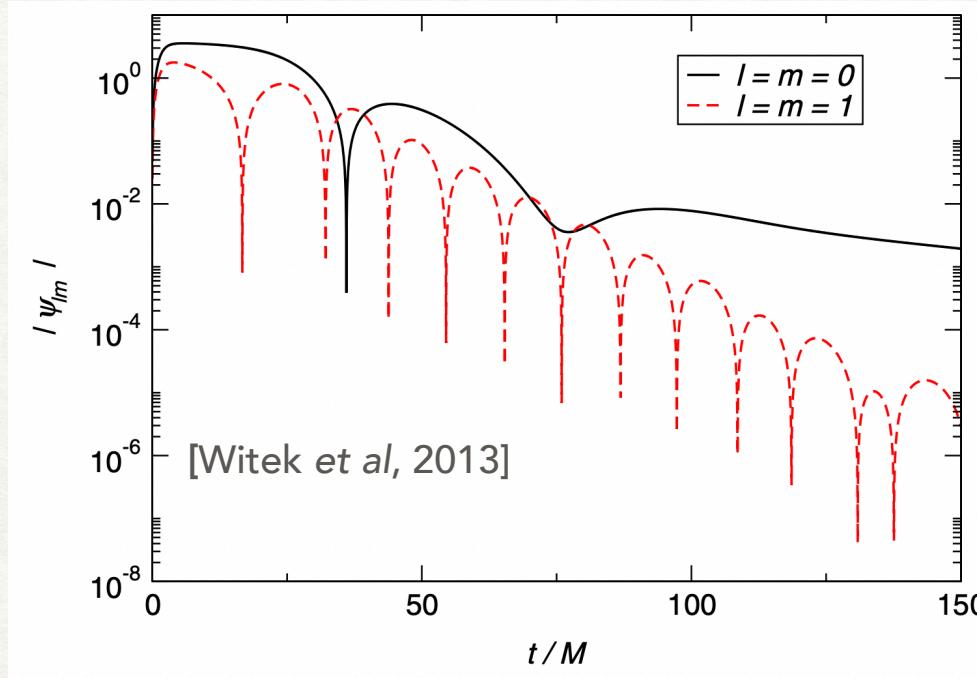
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# Quasinormal modes

- Quasinormal modes (QNMs) describe characteristic dynamics of test fields on BH spacetimes as a linear response



- Many applications:
  - modeling ringdown gravitational waveforms, [Giesler et al, 2019]
  - analysis of relaxation phenomena within AdS/CFT, [Horowitz and Hubeny, 2000]
  - linear mode stability of BH spacetimes,... [Regge and Wheeler, 1957]

# Brief review: QNMs in asymptotically flat BH spacetimes

- Equation for linear perturbations on asymptotically flat background:

$$\left[ \frac{d^2}{dx^2} + \omega^2 - V(x) \right] \phi(x; \omega) = 0$$

e.g., spin-s field perturbation on Schwarzschild backgrounds

$$\begin{aligned} \left[ \frac{d^2}{dx^2} + \omega^2 - V^{(s)} \right] \Phi^{(s)} &= 0, \quad \frac{dr}{dx} = 1 - \frac{r_H}{r} \\ V^{(s)} &= \left(1 - \frac{r_H}{r}\right) \left( \frac{\ell(\ell+1)}{r^2} + \frac{(1-s^2) r_H}{r^3} \right), \end{aligned}$$

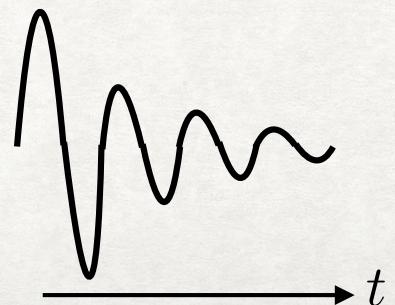
- Appropriate boundary conditions at the horizon  $x = -\infty$  and infinity  $x = \infty$ ,

$$\lim_{x \rightarrow -\infty} \phi = e^{-i\omega x} \quad (\text{purely ingoing})$$

$$\lim_{x \rightarrow +\infty} \phi = e^{i\omega x} \quad (\text{purely outgoing})$$

define QNMs and (a discrete set of) QNM frequencies

- QNM frequencies are complex,  $\phi \propto e^{-i\text{Re}[\omega]t} e^{\text{Im}[\omega]t}$



# Brief review: QNMs as poles of Green's function

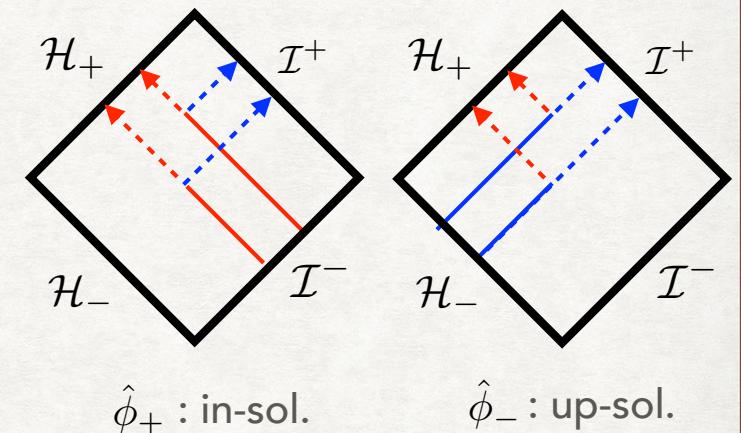
- Equation for linear perturbations in time domain:  $[\partial_x^2 - \partial_t^2 - V(x)] \phi(t, x) = 0$

Laplace transform leads to  $\left[ \frac{d^2}{dx^2} + \omega^2 - V(x) \right] \hat{\phi}(x; \omega) = [i\omega\phi - \partial_t\phi] \Big|_{t=0}$

- Two homogeneous sols. such that

$$\hat{\phi}_+ \simeq \begin{cases} e^{-i\omega t}, & x \rightarrow -\infty, \\ a_1(\omega) e^{-i\omega t} + a_2(\omega) e^{+i\omega t}, & x \rightarrow +\infty \end{cases}$$

$$\hat{\phi}_- \simeq \begin{cases} b_1(\omega) e^{-i\omega t} + b_2(\omega) e^{+i\omega t}, & x \rightarrow -\infty \\ e^{+i\omega t}, & x \rightarrow +\infty, \end{cases}$$



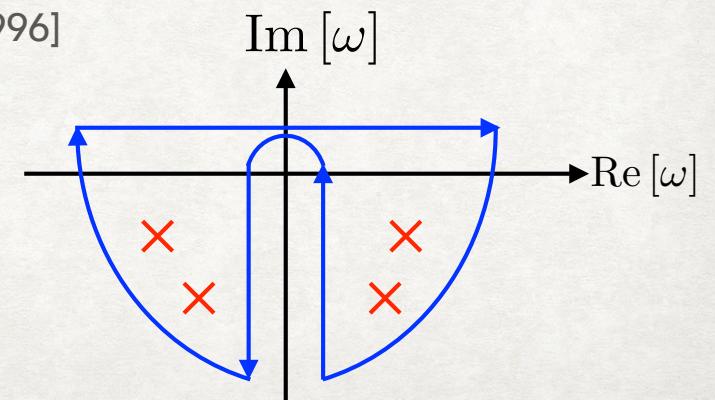
- Green's function in frequency domain: [Leaver, 1988]

$$\hat{G}(x, x'; \omega) = \frac{\hat{\phi}_+(x; \omega) \hat{\phi}_-(x'; \omega)}{W(\omega)}$$

[Anderson, 1996]

where  $W$  is a Wronskian of  $\hat{\phi}_+$  and  $\hat{\phi}_-$

QNM frequencies are determined by  $W = 0$

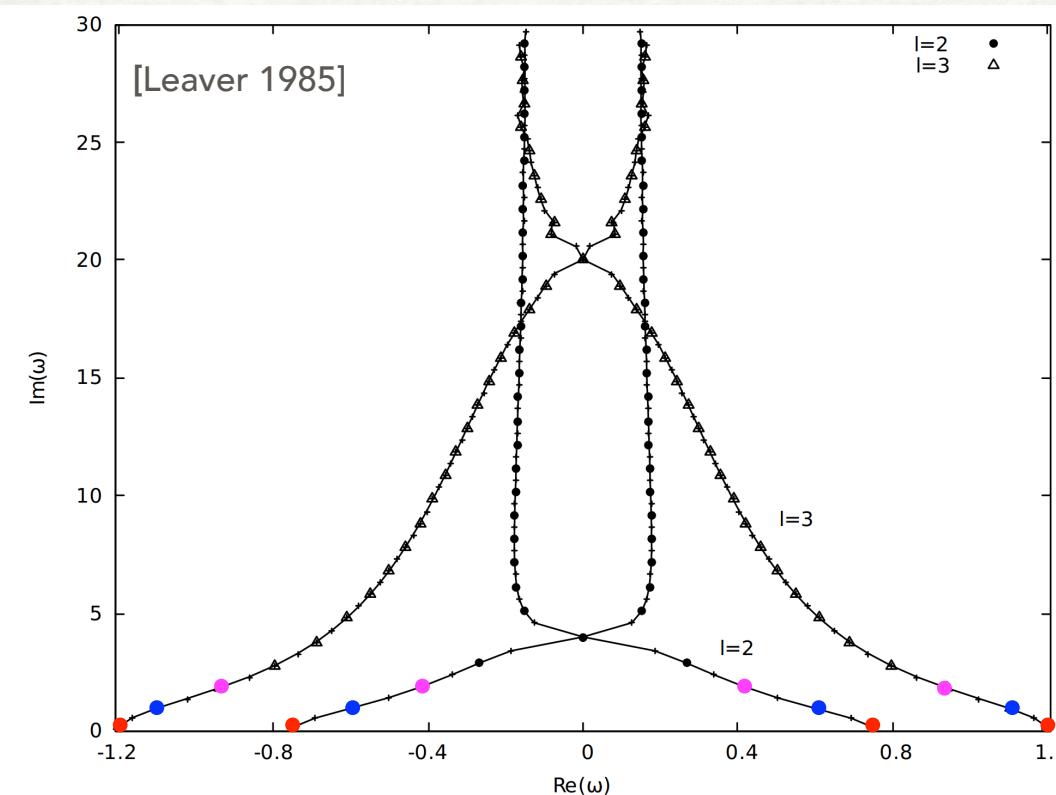


# Brief review: Overtones

- QNM takes a form:  $\Phi_{\text{QNM}} = \sum_{\ell=2}^{\infty} \sum_{m=-\ell}^{\ell} \sum_{n=0}^{\infty} \phi_{\ell mn}(x) e^{-i\omega_{\ell mn} t} Y_{\ell m}(\theta, \varphi)$

For each mode with  $\ell$ ,  
there exists a discrete set of modes labeled by  $n (= 0, 1, 2, \dots)$ : **overtones**

- Index of overtones is defined in the order from the smallest value of  $|\text{Im}[\omega]|$   
 $n = 0$  : fundamental mode,  $n = 1$  : 1st overtones,  $n = 2$  : 2nd overtones, ...



# QNMs in AdS BH spacetimes

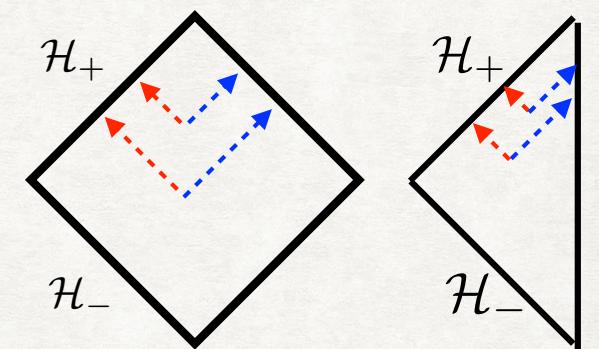
- QNMs in AdS BH spacetimes can be defined in the same manner [Berti et al, 2009]
- Variety of boundary condition at infinity exists due to the asymptotic structure

e.g.,  $\lim_{x \rightarrow \infty} \phi = A + \frac{B}{x^3}$  for massless scalar field

$A = 0$  : Dirichlet condition

$B = 0$  : Neumann condition

$B = \kappa A, \kappa \in \mathcal{R}$  : Robin condition



Asym. flat

Asym. AdS



Rich structure of QNM dynamics appears in AdS (as will be seen later)

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# Question and our work

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Mass ladder operator works in curved spacetimes with conformal symmetry:

$$[\square - \mu^2] \Phi = 0 \quad \rightarrow \quad [\square - (\mu^2 + \delta\mu^2)] D\Phi = 0$$

Question:

Does a mass ladder operator keep physics determined by boundary conditions?

We study QNMs of a massive Klein-Gordon field  
in Banados-Teitelboim-Zanelli black hole spacetimes

Why BTZ?: BTZ spacetime is the simplest system,  
in which QNMs and mass ladder operators can be exactly derived

# Static Banados-Teitelboim-Zanelli spacetime

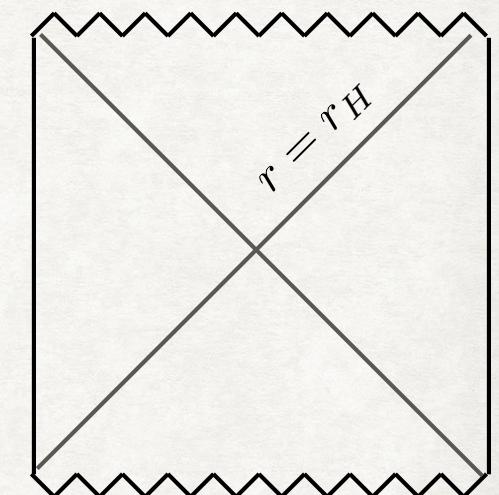
- BTZ geometry describes asymptotically AdS black hole spacetimes in 3 dim.
- Line element in  $(t, r, \varphi)$  coordinates: [Banados, Teitelboim, and Zanelli, 1992]

$$ds^2 = -N^2(r) dt^2 + \frac{1}{N^2(r)} dr^2 + r^2 d\varphi^2, \quad N^2(r) = \frac{r^2 - r_H^2}{\ell^2}$$

$$-\infty < t < \infty, \quad r_H < r < \infty, \quad 0 \leq \varphi < 2\pi$$

- Horizon is located at  $r = r_H$  such that  $N^2(r_H) = 0$

- locally isometric to  $\text{AdS}_3$



\*BTZ BH can rotate but for simplicity we consider the static version

# Massive Klein-Gordon fields

- Massive Klein-Gordon field:  $\left[ \nabla_\mu \nabla^\mu - \frac{\nu(\nu+2)}{\ell^2} \right] \Phi = 0$   
 $\nu \geq -1$

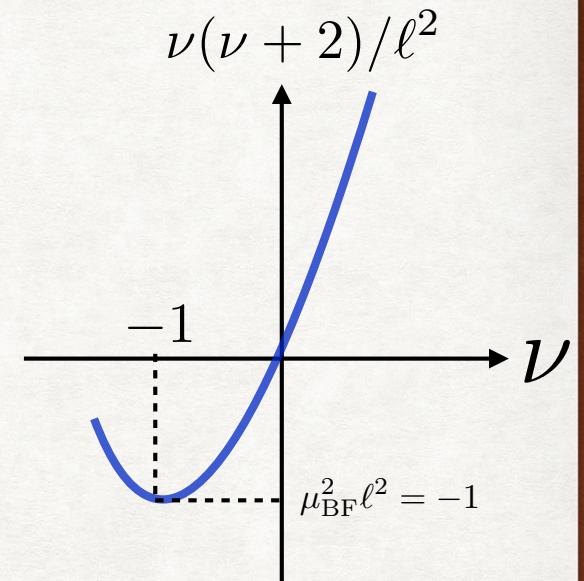
- Expanding the field:  $\Phi = \sum_m \phi_m(r) e^{-i\omega_m t} e^{im\varphi}$

$$\phi'' + \left( \frac{1}{r} + \frac{(N^2)'}{N^2} \right) \phi' + \frac{1}{N^2} \left( \frac{\omega^2}{N^2} - \frac{m^2}{r^2} - \frac{\nu(\nu+2)}{\ell^2} \right) \phi = 0$$

- Appropriate B.C. selects a discrete set of eigenvalues, i.e., QNM frequencies

At the horizon: Ingoing-wave condition

$$\phi = A \left( 1 - \frac{r_H^2}{r^2} \right)^{-\frac{i\omega\ell^2}{2r_H}} \left( \frac{r_H}{r} \right)^{\nu+2} {}_2F_1(a, b; c; 1 - r_H^2/r^2)$$



$$a = \frac{\nu+2}{2} - i \frac{\ell}{2r_H} (\omega\ell - m)$$

$$b = \frac{\nu+2}{2} - i \frac{\ell}{2r_H} (\omega\ell + m)$$

$$c = 1 - i \frac{\omega\ell^2}{r_H}$$

# QNMs in BTZ spacetimes

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- B.C. at infinity: Dirichlet B.C. for  $\nu > -1$  ( $\mu^2 > \mu_{\text{BF}}^2$ )

$$\phi = A_{\text{I}}(\omega) \left( \frac{r_H}{r} \right)^{-\nu} [1 + \dots] + A_{\text{II}}(\omega) \left( \frac{r_H}{r} \right)^{\nu+2} [1 + \dots]$$

$$a = \frac{\nu + 2}{2} - i \frac{\ell}{2r_H} (\omega\ell - m)$$

$$b = \frac{\nu + 2}{2} - i \frac{\ell}{2r_H} (\omega\ell + m)$$

$$c = 1 - i \frac{\omega\ell^2}{r_H}$$

$A_{\text{I}}(\omega) \left( := A \frac{\Gamma(c)\Gamma(a+b-c)}{\Gamma(a)\Gamma(b)} \right)$  vanishes

if  $a = -n$  or  $b = -n$  for  $n = 0, 1, 2, \dots$  due to  $1/\Gamma(-n) = 0$

- QNMs frequencies:

$$\boxed{\omega_D = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} (2n + 2 + \nu)}$$

[Cardoso and Lemos, 2001]

$$\phi = A \left( 1 - \frac{r_H^2}{r^2} \right)^{-i \frac{\omega_D \ell^2}{2r_H}} \left( \frac{r_H}{r} \right)^{2+\nu} \sum_{k=0}^n \frac{(a)_k (b)_k}{k! (c)_k} \left( 1 - \frac{r_H^2}{r^2} \right)^k \quad \text{where } (\xi)_k \equiv \Gamma(z+k)/\Gamma(z)$$

Imaginary parts are negative, indicating linear mode stability

# QNMs in BTZ spacetimes: Other boundary condition

- B.C. at infinity: Neumann B.C. for  $\nu > -1$  ( $\mu^2 > \mu_{\text{BF}}^2$ )

$$\phi = A_I(\omega) \left( \frac{r_H}{r} \right)^{-\nu} [1 + \dots] + \cancel{A_{II}(\omega)} \left( \frac{r_H}{r} \right)^{\nu+2} [1 + \dots]$$

→ 
$$\boxed{\omega_N = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} (2n - \nu)}$$

Imaginary part can be nonnegative if  $\nu \geq 0$  ( $\mu^2 \geq 0$ ),  
indicating linear mode instability  
due to the presence of non-normalizable mode

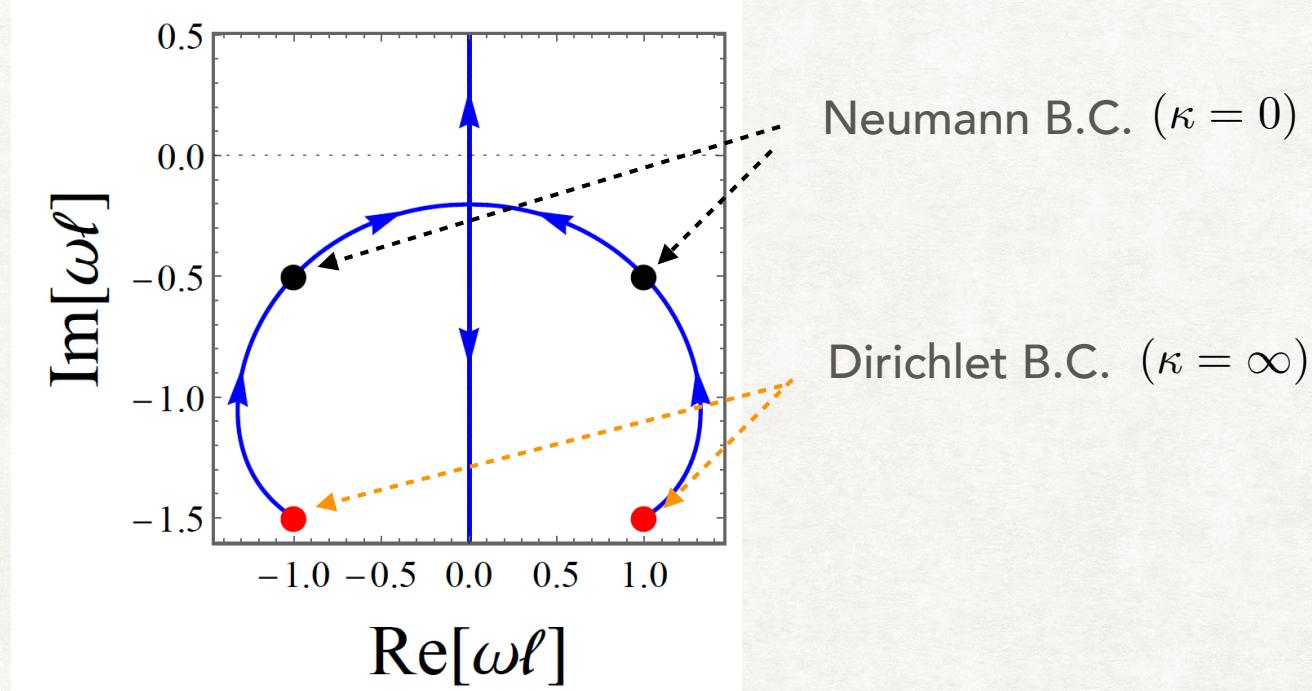
- B.C. at infinity: Robin B.C. for  $-1 < \nu < 0$  ( $\mu_{\text{BF}}^2 < \mu^2 < 0$ )

$$\boxed{A_{II}/A_I = \kappa \quad (\kappa \in \mathbb{R})}$$

including the Dirichlet  $\kappa = \infty$  ( $A_I = 0$ ) and Neumann B.C.  $\kappa = 0$  ( $A_{II} = 0$ )

In this sense, Robin B.C. is more general and admits rich structure

# Fundamental modes in Robin condition



As  $\kappa$  decreases, the trajectories approach the imaginary axis, eventually intersect, and split into two parts

There exist growing modes,  
indicating linear mode instability due to the boundary condition

[Ishibashi and Wald, 2004]

[TK and Harada, 2021]

# QNMs in BTZ spacetimes: BF bound case

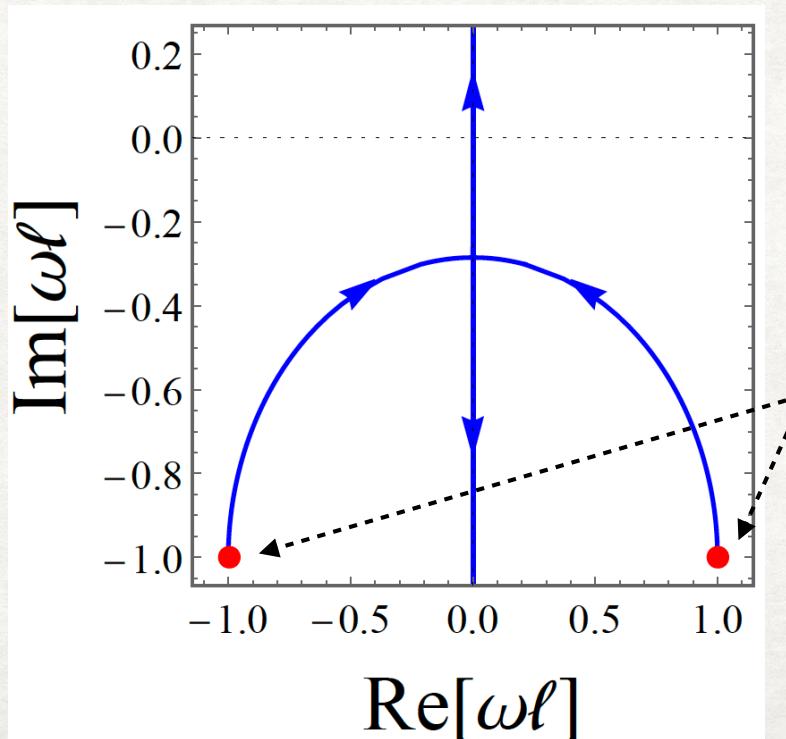
- B.C. at infinity: Dirichlet-Neumann B.C. for  $\nu = -1$  ( $\mu^2 = \mu_{\text{BF}}^2$ ) [Ishibashi and Wald, 2004]

$$\phi(r) = A_{\text{I,BF}} \frac{r_H}{r} + \cancel{A_{\text{II,BF}}} \frac{r_H}{r} \ln \left( \frac{r_H}{r} \right) + \mathcal{O}(1/r^3),$$

→  $\omega_{\text{DN}} = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} (2n + 1)$

- B.C. at infinity: Robin B.C. for  $\nu = -1$  ( $\mu^2 = \mu_{\text{BF}}^2$ )

$$A_{\text{II,BF}}/A_{\text{I,BF}} = 1/\kappa_{\text{BF}} \quad (\kappa_{\text{BF}} \in \mathbb{R})$$



Dirichlet - Neumann B.C. ( $\kappa_{\text{BF}} \rightarrow -\infty$ )

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# Mass ladder operators in BTZ spacetimes

- Mass ladder operators:  $D_{i,k} := \mathcal{L}_{\zeta_i} - \frac{k}{3} \nabla_\mu \zeta_i^\mu, \quad k \in \mathbb{R}$
- Four independent CCKVs ( $i = 0, 1, 2, 3$ ) exist in the BTZ spacetime, thus:

$$D_{0,k} = e^{\frac{r_H}{\ell^2}t} \left( \frac{1}{\sqrt{r^2 - r_H^2}} \partial_t - \frac{r\sqrt{r^2 - r_H^2}}{\ell^2 r_H} \partial_r + k \frac{\sqrt{r^2 - r_H^2}}{\ell^2 r_H} \right),$$

$$D_{1,k} = e^{-\frac{r_H}{\ell^2}t} \left( \frac{1}{\sqrt{r^2 - r_H^2}} \partial_t + \frac{r\sqrt{r^2 - r_H^2}}{\ell^2 r_H} \partial_r - k \frac{\sqrt{r^2 - r_H^2}}{\ell^2 r_H} \right),$$

$$D_{2,k} = e^{\frac{r_H}{\ell}\varphi} \left( \frac{r^2 - r_H^2}{\ell r_H} \partial_r + \frac{1}{r} \partial_\varphi - k \frac{r}{\ell r_H} \right),$$

$$D_{3,k} = e^{-\frac{r_H}{\ell}\varphi} \left( -\frac{r^2 - r_H^2}{\ell r_H} \partial_r + \frac{1}{r} \partial_\varphi + k \frac{r}{\ell r_H} \right).$$

Acting on  $\Phi \propto e^{-i\omega t} e^{im\varphi}$ , the factors  $e^{\pm \frac{r_H^2}{\ell^2}t}$  shift  $\omega \rightarrow \omega \pm ir_H/\ell^2$

while  $e^{\pm \frac{r_H}{\ell}\varphi}$  break the periodicity to  $\varphi$



We mainly focus on  $D_{0,k}, D_{1,k}$  ( $D_{2,k}, D_{3,k}$  still work by multiple action)

# Mass ladder operators in BTZ spacetimes

- Commutation relation:  $[\nabla_\mu \nabla^\mu, D_{i,k}] = -\frac{2k+1}{\ell^2} D_{i,k} + \frac{2}{3} (\nabla_\mu \zeta_i^\mu) \left[ \nabla_\mu \nabla^\mu - \frac{k(k+2)}{\ell^2} \right],$

When choosing  $k = k_+ := -2 - \nu$

$$D_{i,k_+-2} \left[ \nabla_\mu \nabla^\mu - \frac{\nu(\nu+2)}{\ell^2} \right] \Phi = \left[ \nabla_\mu \nabla^\mu - \frac{(\nu+1)(\nu+3)}{\ell^2} \right] D_{i,k_+} \Phi.$$

When choosing  $k = k_- := \nu$

$$D_{i,k_--2} \left[ \nabla_\mu \nabla^\mu - \frac{\nu(\nu+2)}{\ell^2} \right] \Phi = \left[ \nabla_\mu \nabla^\mu - \frac{(\nu-1)(\nu+1)}{\ell^2} \right] D_{i,k_-} \Phi.$$



For the massive Klein-Gordon field  $\Phi$  with  $\nu(\nu+2)/\ell^2$

$D_{i,k_\pm} \Phi$  is also that with  $\tilde{\nu}(\tilde{\nu}+2)/\ell^2$  ( $\tilde{\nu} = \nu \pm 1$ )

# Mass ladder operators and QNM boundary conditions

- QNM with Dirichlet B.C.:  $\Phi = A \left(1 - \frac{r_H^2}{r^2}\right)^{-i\frac{\ell^2}{2r_H}\omega_D} \left(\frac{r_H}{r}\right)^{2+\nu} {}_2F_1(a, b; c; 1 - r_H^2/r^2) e^{-i\omega_D t + im\varphi},$

$$\omega_D = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} (2n + 2 + \nu)$$

$$\Phi|_{r \approx r_H} = 2^{-i\frac{\ell^2}{2r_H}\omega_D} A \left(\frac{r - r_H}{r_H}\right)^{-i\frac{\ell^2}{2r_H}\omega_D} [1 + \mathcal{O}(r - r_H)] e^{-i\omega_D t + im\varphi}. \quad (\text{Ingoing wave})$$

- Acting  $D_{0,k_\pm}, D_{1,k_\pm}$  on the QNMs, at  $r \rightarrow r_H$

$$D_{0,k_\pm} \Phi = c_{0,k_\pm} \left(\frac{r - r_H}{r_H}\right)^{-i\frac{\ell^2}{2r_H}(\omega + i\frac{r_H}{\ell^2})} [1 + \mathcal{O}(r - r_H)] e^{-i(\omega + i\frac{r_H}{\ell^2})t + im\varphi},$$

$$D_{1,k_\pm} \Phi = c_{1,k_\pm} \left(\frac{r - r_H}{r_H}\right)^{-i\frac{\ell^2}{2r_H}(\omega - i\frac{r_H}{\ell^2})} [1 + \mathcal{O}(r - r_H)] e^{-i(\omega - i\frac{r_H}{\ell^2})t + im\varphi}.$$

Mass ladder operators keep the ingoing-wave condition

# Mass ladder operators and QNM boundary conditions

$$\Phi|_{r \simeq \infty} = A_{\text{II}} \left( \frac{r_H}{r} \right)^{2+\nu} [1 + \mathcal{O}(1/r^2)] e^{-i\omega t + im\varphi} \quad (\text{Dirichlet B.C.})$$

- Acting  $D_{0,k_+}, D_{1,k_+}$  on the QNMs, at  $r \rightarrow \infty$

$$D_{0,k_+} \Phi = c_{0,k_+}^{(\text{D})} \left( \frac{r_H}{r} \right)^{3+\nu} [1 + \mathcal{O}(1/r^2)] e^{-i(\omega + i \frac{r_H}{\ell^2})t + im\varphi},$$
$$D_{1,k_+} \Phi = c_{1,k_+}^{(\text{D})} \left( \frac{r_H}{r} \right)^{3+\nu} [1 + \mathcal{O}(1/r^2)] e^{-i(\omega - i \frac{r_H}{\ell^2})t + im\varphi}.$$

Asymptotic behaviors of the Klein-Gordon field with  $\tilde{\nu} (\tilde{\nu} + 2) / \ell^2$  ( $\tilde{\nu} = \nu + 1$ )

Mass ladder operators  $D_{0,k_+}, D_{1,k_+}$  keep the Dirichlet B.C.

# Mass ladder operators and QNM boundary conditions

$$\Phi|_{r \simeq \infty} = A_{\text{II}} \left( \frac{r_H}{r} \right)^{2+\nu} [1 + \mathcal{O}(1/r^2)] e^{-i\omega t + im\varphi} \quad (\text{Dirichlet B.C.})$$

- Acting  $D_{0,k_-}, D_{1,k_-}$  on the QNMs, at  $r \rightarrow \infty$

$$D_{0,k_-} \Phi = c_{0,k_-}^{(\text{D})} \left( \frac{r_H}{r} \right)^{1+\nu} [1 + \mathcal{O}(1/r^2)] e^{-i(\omega + i \frac{r_H}{\ell^2})t + im\varphi},$$

$$D_{1,k_-} \Phi = c_{1,k_-}^{(\text{D})} \left( \frac{r_H}{r} \right)^{1+\nu} [1 + \mathcal{O}(1/r^2)] e^{-i(\omega - i \frac{r_H}{\ell^2})t + im\varphi}.$$

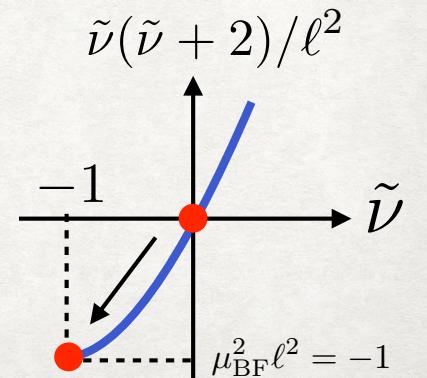
- For  $\nu > 0$  ( $\mu^2 > 0$ ) :

Asymptotic behaviors of the Klein-Gordon field with  $\tilde{\nu}(\tilde{\nu} + 2)/\ell^2$  ( $\tilde{\nu} = \nu - 1$ )

Mass ladder operators  $D_{0,k_-}, D_{1,k_-}$  keep the Dirichlet B.C.

- For  $\nu = 0$  ( $\mu^2 = 0$ ):  $\tilde{\nu} = -1$  corresponds to  $\mu_{\text{BF}}^2 \ell^2 = -1$

Dirichlet B.C. changes to the Dirichlet-Neumann B.C.



# Mass ladder operators and QNM boundary conditions

$$D_{0,k_-} \Phi = c_{0,k_-}^{(\text{D})} \left( \frac{r_H}{r} \right)^{1+\nu} [1 + \mathcal{O}(1/r^2)] e^{-i(\omega + i \frac{r_H}{\ell^2})t + im\varphi},$$

$$D_{1,k_-} \Phi = c_{1,k_-}^{(\text{D})} \left( \frac{r_H}{r} \right)^{1+\nu} [1 + \mathcal{O}(1/r^2)] e^{-i(\omega - i \frac{r_H}{\ell^2})t + im\varphi}.$$

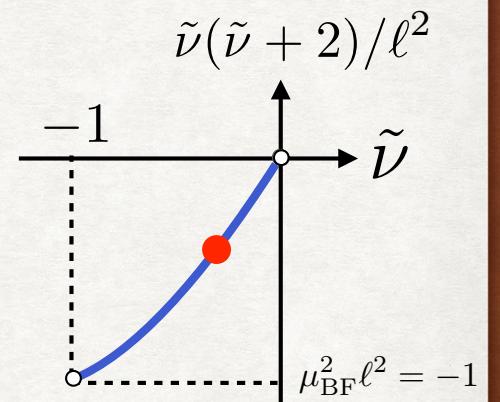
For  $-1 < \nu < 0$  ( $\mu_{\text{BF}}^2 < \mu^2 < 0$ ) :

The above corresponds to  $A_{\text{II}}(\omega) = 0$  (Neumann B.C.)  
of the asymptotic behavior

$$\phi = A_{\text{I}}(\omega) \left( \frac{r_H}{r} \right)^{-\tilde{\nu}} [1 + \dots] + \cancel{A_{\text{II}}(\omega) \left( \frac{r_H}{r} \right)^{\tilde{\nu}+2} [1 + \dots]}$$

with  $\tilde{\nu} = |\nu| - 1$   
 $(-1 < \tilde{\nu} < 0)$

Dirichlet B.C. changes to Neumann B.C.



# Brief summary: Changes of QNM boundary conditions

- At the horizon:

Mass ladder operators  $D_{0,k\pm}, D_{1,k\pm}$  keep the ingoing-wave condition

- At infinity:

Mass ladder operators  $D_{0,k_+}, D_{1,k_+}$

Dirichlet B.C. ————— Dirichlet B.C.

Mass ladder operators  $D_{0,k_-}, D_{1,k_-}$

For  $\nu > 0$  ( $\mu^2 > 0$ )

Dirichlet B.C. ————— Dirichlet B.C.

For  $\nu = 0$  ( $\mu^2 = 0$ )

Dirichlet B.C. ————— Dirichlet-Neumann B.C.

For  $-1 < \nu < 0$  ( $\mu_{\text{BF}}^2 < \mu^2 < 0$ )

Dirichlet B.C. ————— Neumann B.C.

# Brief summary: Neumann case

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- At infinity:

Mass ladder operators  $D_{0,k_+}, D_{1,k_+}$

Neumann B.C.  $\longrightarrow$  Neumann B.C.

Mass ladder operators  $D_{0,k_-}, D_{1,k_-}$

For  $\nu > 0$  ( $\mu^2 > 0$ )

Neumann B.C.  $\longrightarrow$  Neumann B.C.

For  $\nu = 0$  ( $\mu^2 = 0$ )

Neumann B.C.  $\longrightarrow$  Dirichlet-Neumann B.C.

For  $-1 < \nu < 0$  ( $\mu_{\text{BF}}^2 < \mu^2 < 0$ )

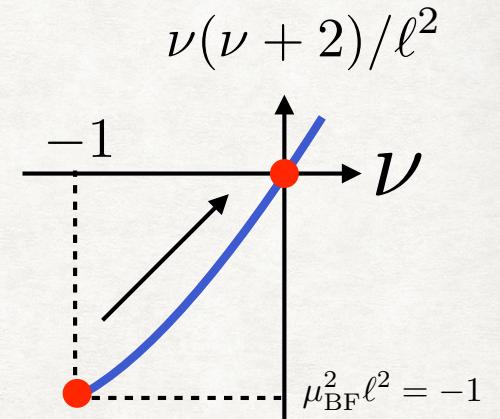
Neumann B.C.  $\longrightarrow$  Dirichlet B.C.

# Brief summary: Dirichlet-Neumann, Robin cases

- Dirichlet-Neumann case  $\nu = -1$  ( $\mu^2 = \mu_{\text{BF}}^2$ ):

Mass ladder operators  $D_{0,k_+}, D_{1,k_+}$  ( $k_+ = k_-$ )

Dirichlet-Neumann B.C.  $\longrightarrow$  Dirichlet B.C.



- Robin boundary condition is kept  
but the resulting boundary condition parameter is complex

# Outline

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1. Introduction
2. Review: Mass ladder operators
3. Review: Quasinormal modes
4. Quasinormal modes in Banados-Teitelboim-Zanell spacetimes
5. Mass ladder operators in Banados-Teitelboim-Zanell spacetimes
6. Shift of quasinormal mode frequencies

# QNM frequency shift

- Frequency shift from expressions of mass ladder operators

$$D_{0,k} = e^{\frac{r_H}{\ell^2}t} \left( \frac{1}{\sqrt{r^2 - r_H^2}} \partial_t - \frac{r\sqrt{r^2 - r_H^2}}{\ell^2 r_H} \partial_r + k \frac{\sqrt{r^2 - r_H^2}}{\ell^2 r_H} \right),$$

$$D_{1,k} = e^{-\frac{r_H}{\ell^2}t} \left( \frac{1}{\sqrt{r^2 - r_H^2}} \partial_t + \frac{r\sqrt{r^2 - r_H^2}}{\ell^2 r_H} \partial_r - k \frac{\sqrt{r^2 - r_H^2}}{\ell^2 r_H} \right),$$

The factors  $e^{\pm \frac{r_H^2}{\ell^2}t}$  suggest  $\omega \rightarrow \omega \pm ir_H/\ell^2$

- Example: Shift by  $D_{0,k_+}$

Original QNM frequency with Dirichlet B.C.:  $\omega_D = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} (2n + 2 + \nu)$

Mass parameter shift:  $\nu \rightarrow \tilde{\nu} = \nu + 1$

QNM frequency shift:  $\omega_D \rightarrow \tilde{\omega}_D = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} [2 \underbrace{(n - 1)}_{n \rightarrow n - 1} + 2 + \tilde{\nu}]$

Note: no “negative overtones” are generated,  $D_{0,k_+}$  [fundamental mode] = 0

# QNM frequency shift

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Operators	$D_{0,k_+}$	$D_{0,k_-}$	$D_{1,k_+}$	$D_{1,k_-}$
Frequencies	$\omega_D(\nu + 1, n - 1)$	$\omega_D(\nu - 1, n) \ (\nu > 0)$ $\omega_{DN}(n) \ (\nu = 0)$ $\omega_N( \nu  - 1, n) \ (-1 < \nu < 0)$	$\omega_D(\nu + 1, n)$	$\omega_D(\nu - 1, n + 1) \ (\nu > 0)$ $\omega_{DN}(n + 1) \ (\nu = 0)$ $\omega_N( \nu  - 1, n + 1) \ (-1 < \nu < 0)$

where  $\omega_D(\nu, n) = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} (2n + 2 + \nu),$  (Dirichlet B.C.)

$\omega_N(\nu, n) = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} (2n - \nu),$  (Neumann B.C.)

$\omega_{DN}(n) = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} (2n + 1),$  (Dirichlet-Neumann B.C.)

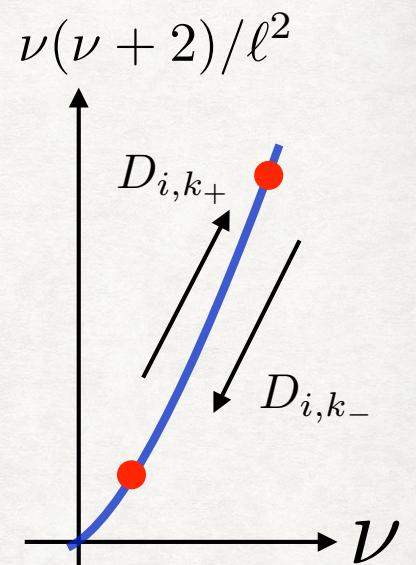
Mass ladder operators change not only mass squared but also indices of overtones

# Overtone shift by multiple actions

- Multiple actions  $D_{0,k_+ - 1} D_{0,k_-}$  or  $D_{1,k_+ - 1} D_{1,k_-}$  keep mass squared but QNM frequencies are shifted:

$$\tilde{\omega} = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} [2(n-1) + 2 + \nu] \quad \text{for } D_{0,k_+ - 1} D_{0,k_-}$$

$$\tilde{\omega} = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} [2(n+1) + 2 + \nu] \quad \text{for } D_{1,k_+ - 1} D_{1,k_-}$$



Note: no “negative overtones” are generated from the fundamental mode

All overtones can be generated from the fundamental mode

# Regular solutions generated by multiple actions of $D_{2,k\pm}$ , $D_{3,k\pm}$

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$$D_{2,k} = e^{\frac{r_H}{\ell}\varphi} \left( \frac{r^2 - r_H^2}{\ell r_H} \partial_r + \frac{1}{r} \partial_\varphi - k \frac{r}{\ell r_H} \right),$$

$$D_{3,k} = e^{-\frac{r_H}{\ell}\varphi} \left( -\frac{r^2 - r_H^2}{\ell r_H} \partial_r + \frac{1}{r} \partial_\varphi + k \frac{r}{\ell r_H} \right).$$

Factors  $e^{\pm \frac{r_H}{\ell}\varphi}$  break the periodicity to  $\varphi$ ;  
thus, the single action of  $D_{2,k\pm}$ ,  $D_{3,k\pm}$  fails to generate a regular solution

- Multiple actions can remove those singular factors,

e.g.,  $D_{2,k_+ - 1} D_{3,k_+}$

$$\tilde{\omega} = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} [2(n-1) + 2 + (\nu + 2)] \quad \text{for Dirichlet B.C.}$$

$$\tilde{\omega} = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} [2(n+1) - (\nu + 2)] \quad \text{for Neumann B.C.}$$

e.g.,  $D_{2,k_- - 1} D_{3,k_+}$

QNM frequency does not change

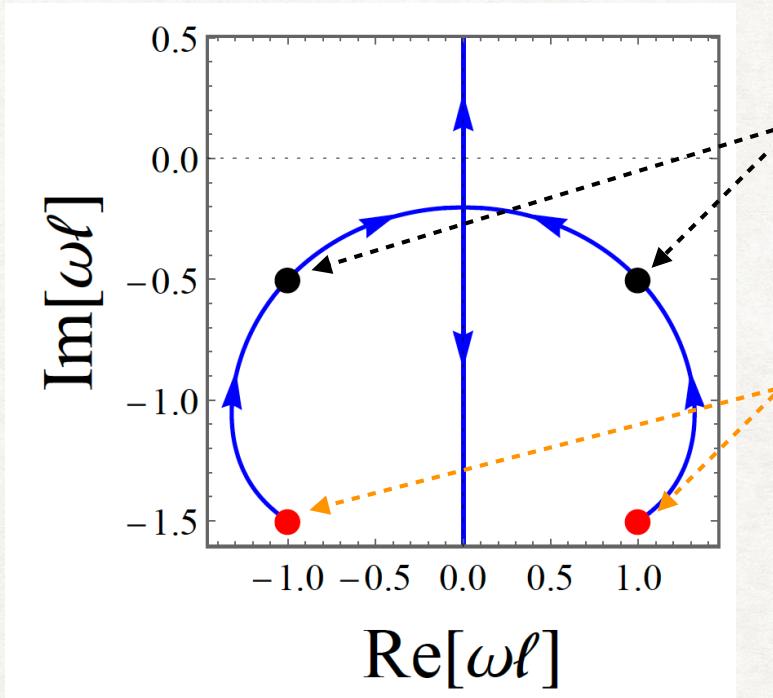
In fact,  $D_{2,k_- - 1} D_{3,k_+} = \mathcal{L}_\xi$  is a symmetry operator

# Other boundary condition: Robin case

- Robin B.C. ( $-1 < \nu < 0$ ):  $A_{\text{II}}/A_{\text{I}} = \kappa$  ( $\kappa \in \mathbb{R}$ ),

[Ishibashi and Wald, 2004]

$$\phi \simeq A_{\text{I}}(\omega) \left( \frac{r_H}{r} \right)^{-\nu} + A_{\text{II}}(\omega) \left( \frac{r_H}{r} \right)^{\nu+2}$$



Neumann B.C. ( $\kappa = 0$ )

$$\omega_N(0) = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} [-\nu]$$

Dirichlet B.C. ( $\kappa = \infty$ )

$$\omega_D(0) = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} [2 + \nu]$$

- Acting the mass ladder operators,  $\omega \rightarrow \tilde{\omega} = \omega \pm ir_H/\ell^2$

$$\kappa \rightarrow \tilde{\kappa} \text{ (complex value)}$$

At least, for  $\kappa$  so that  $\omega$  is purely imaginary,  
 $\tilde{\kappa}$  is real, and  $\tilde{\omega}$  is also purely imaginary

# Summary

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We have studied QNMs of massive Klein-Gordon fields in static BTZ spacetimes in terms of mass ladder operators

- Ladder operator is a useful tool to understand mathematical properties of QNMs
- Mass ladder operator change not only mass squared but also QNM frequencies
- In particular, an index of overtones are shifted
- All overtones can be generated from fundamental modes by their multiple actions

# Other boundary condition: Neumann case

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Operators	$D_{0,k_+}$	$D_{0,k_-}$	$D_{1,k_+}$	$D_{1,k_-}$
Frequencies	$\omega_N(\nu + 1, n)$	$\omega_{DN}(n - 1) \ (\nu = 0)$ $\omega_D( \nu  - 1, n - 1) \ (-1 < \nu < 0)$	$\omega_N(\nu + 1, n + 1)$	$\omega_N(\nu - 1, n) \ (\nu > 0)$ $\omega_{DN}(n) \ (\nu = 0)$ $\omega_D( \nu  - 1, n) \ (-1 < \nu < 0)$

where  $\omega_D(\nu, n) = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} (2n + 2 + \nu), \quad (\text{Dirichlet B.C.})$

$\omega_N(\nu, n) = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} (2n - \nu), \quad (\text{Neumann B.C.})$

$\omega_{DN}(n) = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} (2n + 1), \quad (\text{Dirichlet-Neumann B.C.})$

# BF bound case

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Operators	$D_{0,-1}$	$D_{1,-1}$	$D_0^{\text{BF}}$
Frequencies	$\omega_D(0, n-1)$	$\omega_D(0, n)$	$\omega_{DN}(0) \rightarrow \omega_N(0, 0)$

$$\omega_D(\nu, n) = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} (2n + 2 + \nu), \quad (\text{Dirichlet B.C.})$$

$$\omega_N(\nu, n) = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} (2n - \nu), \quad (\text{Neumann B.C.})$$

$$\omega_{DN}(n) = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2} (2n + 1), \quad (\text{Dirichlet-Neumann B.C.})$$

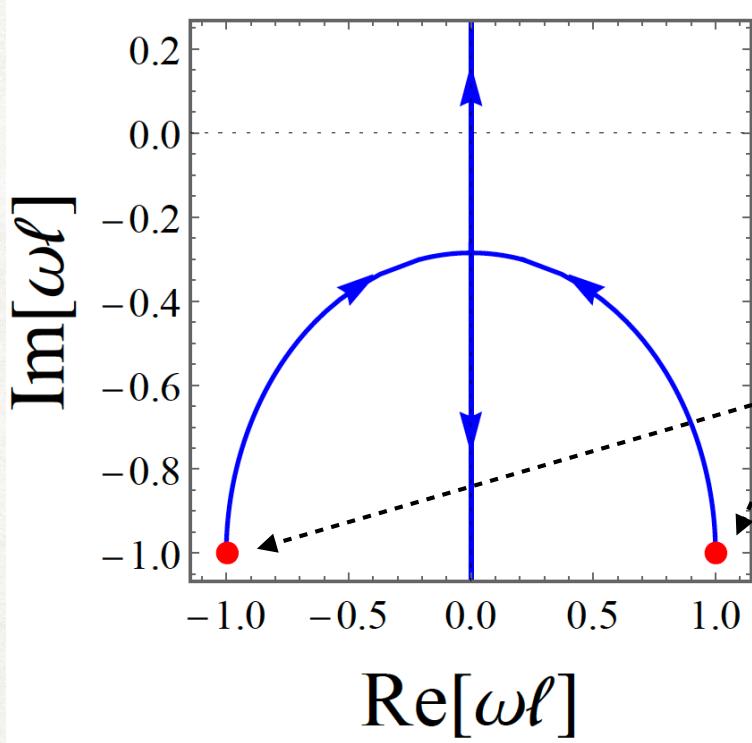
- We find new mass ladder operator only for the fundamental mode in BF bound:

$$D_0^{\text{BF}} \Phi_0 := (\nabla_\mu \zeta_0^\mu) \Phi_0 = e^{r_H t / \ell^2} \frac{\sqrt{r^2 - r_H^2}}{\ell^2 r_H} \Phi_0,$$

# Other boundary condition: Robin case with BF bound

- Robin B.C. ( $\nu = -1$ ):  $A_{\text{II,BF}}/A_{\text{I,BF}} = 1/\kappa_{\text{BF}}$  ( $\kappa_{\text{BF}} \in \mathbb{R}$ ), [Ishibashi and Wald, 2004]

$$\phi(r) = A_{\text{I,BF}} \frac{r_H}{r} + A_{\text{II,BF}} \frac{r_H}{r} \ln \left( \frac{r_H}{r} \right) + \mathcal{O}(1/r^3),$$



Dirichlet - Neumann B.C. ( $\kappa_{\text{BF}} \rightarrow -\infty$ )

$$\omega_{\text{DN}}(0) = \pm \frac{m}{\ell} - i \frac{r_H}{\ell^2}$$

- Acting the mass ladder operators,  $\omega \rightarrow \tilde{\omega} = \omega \pm ir_H/\ell^2$   
 $\kappa_{\text{BF}} \rightarrow \tilde{\kappa}_{\text{BF}}$  (complex value)

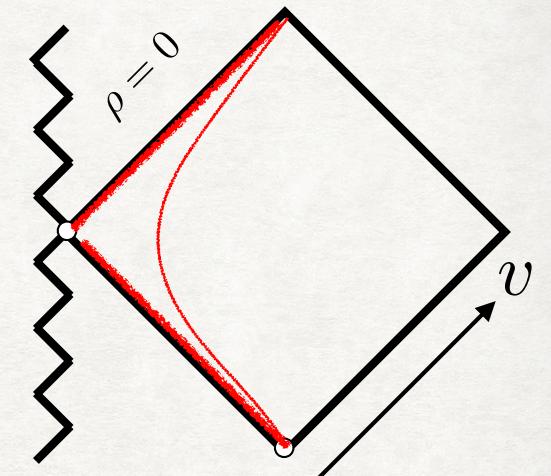
At least, for  $\kappa_{\text{BF}}$  so that  $\omega$  is purely imaginary,  
 $\tilde{\kappa}_{\text{BF}}$  is real, and  $\tilde{\omega}$  is also purely imaginary

# Application: scalars in near-horizon geometry

- We have derived conserved quantities along black hole horizons [TK and Kimura 2021]  
[TK and Kimura 2022] by exploiting mass ladder operators near the horizon

- Vicinity of black holes with zero Hawking temperature is highly-symmetric geometry called near-horizon geometry:

$$ds^2 = \underbrace{-\lambda_0 \rho^2 dv^2}_{\text{AdS}_2} + \underbrace{2dvd\rho + \gamma_0 d\Omega_{n-2}^2}_{S^{n-2}}$$



- Reduction of scalars on near-horizon geometry to that on AdS2:

$$\square \Phi = 0 \longrightarrow \left[ \square_{\text{AdS}_2} - \frac{\ell(\ell + n - 3)}{\gamma_0} \right] \phi_\ell = 0$$

$$\Phi = \sum_{\ell=0}^{\infty} \sum_{m=-\ell}^{\ell} \phi_\ell(v, \rho) Y_{\ell m}(\theta, \varphi)$$

Mass ladder operators connect different multipole modes

- We obtain conservation laws along the horizon:  $\partial_v [\partial_\rho D_1 D_2 \cdots D_\ell \phi_\ell]|_{\rho=0} = 0$