## 1 Imaginary Wilson coefficients and $R_K$ observables

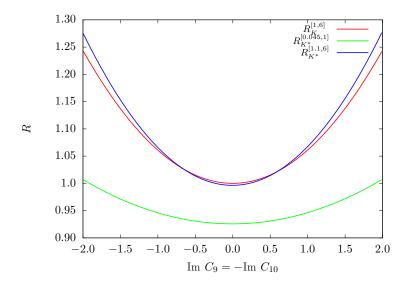


Figure 1: Values of  $R_K$  and  $R_{K^*}$  whith imaginary Wilson coefficients.

Figure ?? shows the values of the ratios  $R_K$  and  $R_{K^*}$  in their respective  $q^2$  ranges, when both Wilson coefficients  $C_9$  and  $C_{10}$  are imaginary, and also we assume  $C_9 = -C_{10}$ . In all cases the minimum value for the ratio is attained at the SM point  $C_9 = -C_{10} = 0$ . The addition of non-zero imaginary Wilson coefficients results in larger values of  $R_K$  and  $R_{K^*}$ , at odds with the experimental values  $R_{K^{(*)}}^{\exp} < R_{K^{(*)}}^{SM}$ .

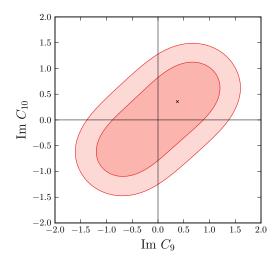


Figure 2: Allowed regions for imaginary Wilson coefficients.

In Figure ?? we plot the allowed regions for imaginary values of  $C_9$  and  $C_{10}$  when fitting to measurements of a series of  $b \to s\mu^+\mu^-$  observables. In the fit we have included the ratios  $R_K$  and  $R_{K^*}$ , the angular observables  $P_4'$  and  $P_5'$  and the branching ratios  $\mathrm{BR}(B_s \to \mu^+\mu^-)$  and  $\mathrm{BR}(B^0 \to \mu^+\mu^-)$ . The fit was carried by the software flavio, wich computes the  $\chi^2$  function with each  $(C_9, C_{10})$  pair. The global minimum of the  $\chi^2$  was found at  $C_9 = 0.31\,i$ ,  $C_{10} = 0.29\,i$ . The pull of the SM, defined as the probability that the SM scenario can describe the best fit assuming that  $\Delta\chi^2 = \chi^2_{\mathrm{SM}} - \chi^2_{\mathrm{min}}$  follows a  $\chi^2$  distribution with 5 degrees of freedom, is of just  $2.1 \times 10^{-5}\sigma$ . That is, the SM result and the best fit are indistinguishable from an statistical point of view. The shaded regions in the plot are  $1\sigma$  (darker pink) and  $2\sigma$  (lighter pink) away from the best fit.

## 2 Z' fit

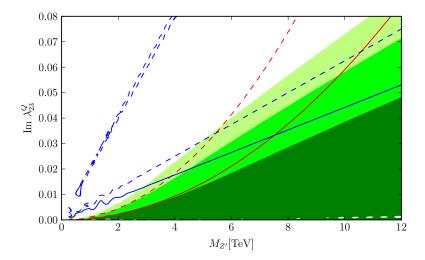


Figure 3: Bounds on Z' parameter space imposed by  $b \to s\mu^+\mu^-$  decays and  $B_s$  mixing.

Figure ?? shows the bounds on the Z' mass  $M_{Z'}$  and the imaginary coupling coupling  $\lambda_{23}^Q$  (setting  $\lambda_{22}^L=1$ ) imposed by  $b\to s\mu^+\mu^-$  decays and  $B_s$  mixing. Blue lines correspond to the fit to  $B_s$ -mixing observables  $\Delta M_s$  and  $A_{\rm CP}^{\rm mix}$  (solid line:  $\Delta\chi^2=1$ , dashed lines  $\Delta\chi^2=4$ ), red lines to  $b\to s\mu^+\mu^-$  as in Figure ?? (solid line:  $\Delta\chi^2=1$ , dashed lines  $\Delta\chi^2=4$ ), and green regions to the combined fit (dark green:  $\Delta\chi^2=1$ , medium green  $\Delta\chi^2=4$ , light green:  $\Delta\chi^2=9$ ). For the  $B_s$ -mixing fit, the SM offers the best fit, while the best fits for the  $b\to s\mu^+\mu^-$  and global fits is  $(M_{Z'}=11.0~{\rm TeV},~\lambda_{23}^Q=0.001i)$ , although the improvement with respect to the SM is negligible. In the low mass region  $(M_{Z'}\lesssim 7~{\rm TeV})$  the global fit is dominated by  $b\to s\mu^+\mu^-$  observables, and  $B_s$ -mixing becomes more important at larger masses. The allowed regions correspond to 'small' Wilson coefficients (Im $C_{10}<0.7,~C_{bs}^{LL}>-3\times10^{-4}$ ), except for a narrow region in the  $B_s$  fit with  $C_{bs}^{LL}\sim-2.7\times10^{-3}$ .

## 3 Leptoquark fit

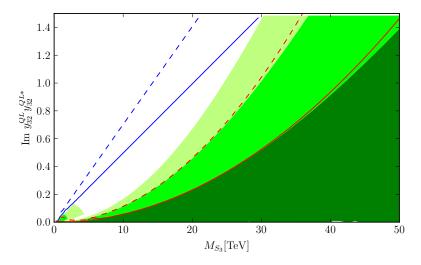


Figure 4: Bounds on  $S_3$  leptoquark parameter space imposed by  $b \to s\mu^+\mu^-$  decays and  $B_s$  mixing.

Figure ?? shows the bounds on the  $S_3$  leptoquark mass  $M_{S_3}$  and the imaginary coupling coupling  $y_{23}^{QL}y_{23}^{QL*}$  imposed by  $b \to s\mu^+\mu^-$  decays and  $B_s$  mixing. The observables used in the respective fits are the same as in Figure ??. The SM scenario was preferred in all the fits. In contrast to Z' models,  $b \to s\mu^+\mu^-$  observables remain the main contributors to the bounds at masses as large as  $M_{S_3} \sim 50$  TeV.