

Experiment: Motion of a Particle under a Central Force

$$F = -\frac{K}{r^3} \left(1 - \frac{\alpha}{r}\right) \mathbf{r}$$

Aim

To study the motion of a particle under the central force

$$F = -\frac{K}{r^3} \left(1 - \frac{\alpha}{r}\right) \mathbf{r},$$

and to show that the resulting orbit is a **precessing ellipse**, depending on the constants K and α .

Apparatus Required

Computer or microprocessor system capable of numerical computation (Python with NumPy and Matplotlib recommended).

You don't need to write Theory in the Record

Theory-This is for your understanding. You can just skim it. No need to write

Consider a particle of mass m moving under a central potential $V(r)$ which depends only on the distance r from a fixed center.

Potential Energy and Force Components

The given central force

$$\mathbf{F}(r) = -\frac{K}{r^3} \left(1 - \frac{\alpha}{r}\right) \mathbf{r}$$

corresponds to the potential energy function

$$V(r) = -\frac{K}{r} + \frac{K\alpha}{2r^2}.$$

Definitions of Constants:

- m : Mass of the particle (kg)
- K : Strength of the inverse-square (attractive) central force

- α : Small positive constant introducing deviation from the pure inverse-square law
- r : Radial distance from the center (m)
- \mathbf{r} : Position vector of the particle

Since V depends only on r , the force is purely radial:

$$F_\theta = 0, \quad F_r = -\frac{dV}{dr} = -\frac{K}{r^2} + \frac{K\alpha}{r^3}.$$

Equation of Motion in Polar Coordinates

The radial and angular equations of motion are

$$m(\ddot{r} - r\dot{\theta}^2) = F_r(r), \quad m(r\ddot{\theta} + 2\dot{r}\dot{\theta}) = 0.$$

The second equation implies conservation of angular momentum:

$$\boxed{L = mr^2\dot{\theta} = \text{constant.}}$$

Here L is the angular momentum of the particle about the center of force.

Substituting for $\dot{\theta}$ in the radial equation gives

$$\ddot{r} = \frac{L^2}{m^2 r^3} - \frac{K}{mr^2} + \frac{K\alpha}{mr^3}.$$

Reduction using $u = 1/r$

Let $u = 1/r$. Then

$$\frac{d^2 u}{d\theta^2} + u = -\frac{m}{L^2 u^2} F_r\left(\frac{1}{u}\right).$$

Substituting $F_r = -K(1 - \alpha u)u^2$ gives

$$\frac{d^2 u}{d\theta^2} + \underbrace{\left(1 + \frac{mK\alpha}{L^2}\right)}_{\beta^2} u = \frac{mK}{L^2},$$

where

$$\beta = \sqrt{1 + \frac{mK\alpha}{L^2}}$$

is the *precession parameter*.

Solution of the Orbit Equation

The general solution of the above differential equation is

$$u(\theta) = \frac{mK/L^2}{1 + \frac{mK\alpha}{L^2}} [1 + e \cos(\beta(\theta - \theta_0))] ,$$

where

- e : Eccentricity-like parameter that determines the shape of the orbit,
- θ_0 : Angular position of the periapsis (initial phase).

Hence the trajectory is

$$\boxed{r(\theta) = \frac{p_{\text{eff}}}{1 + e \cos(\beta(\theta - \theta_0))}} , \quad p_{\text{eff}} = \frac{L^2 + mK\alpha}{mK} .$$

When $\alpha = 0$, $\beta = 1$ and the orbit is a closed Keplerian ellipse. When $\alpha \neq 0$, $\beta \neq 1$ and the orbit precesses.

Precession of the Orbit

The radial distance repeats when the angle advances by $2\pi/\beta$, so the apsidal precession per revolution is

$$\boxed{\Delta\varpi = 2\pi \left(1 - \frac{1}{\sqrt{1 + \frac{mK\alpha}{L^2}}} \right)} .$$

Thus, the orbit precesses in the direction of motion (prograde precession) for positive α .

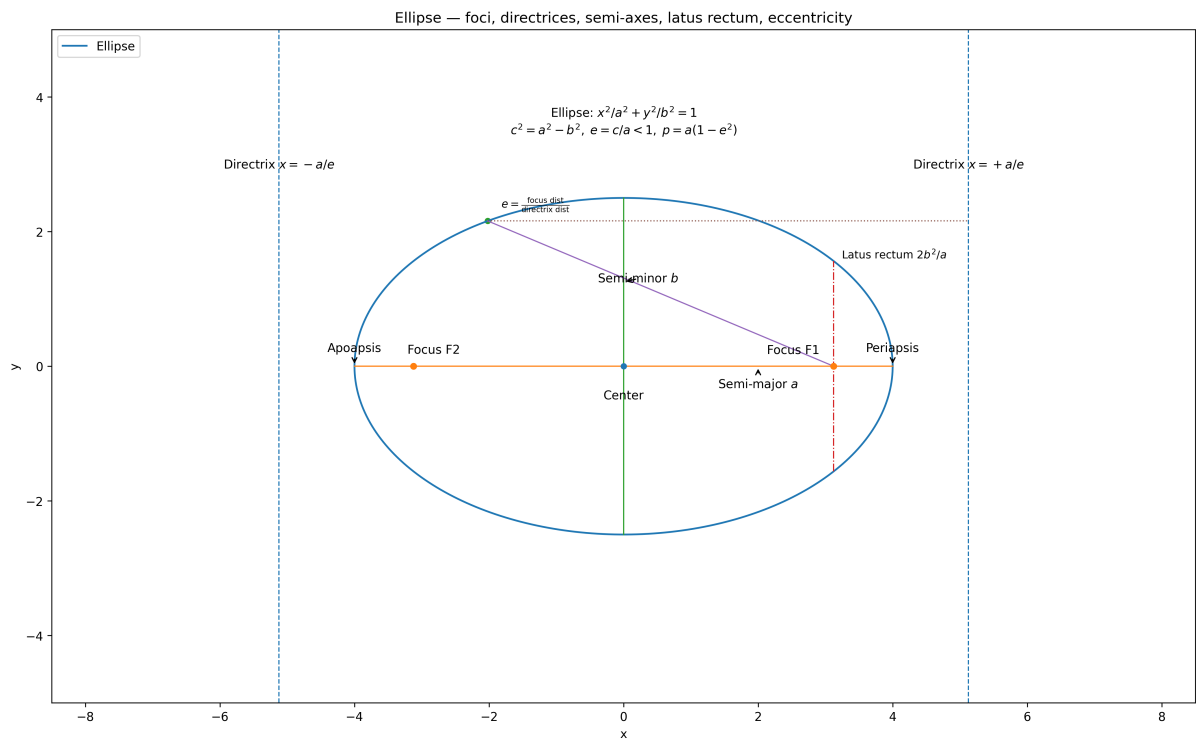


Figure 1: Caption

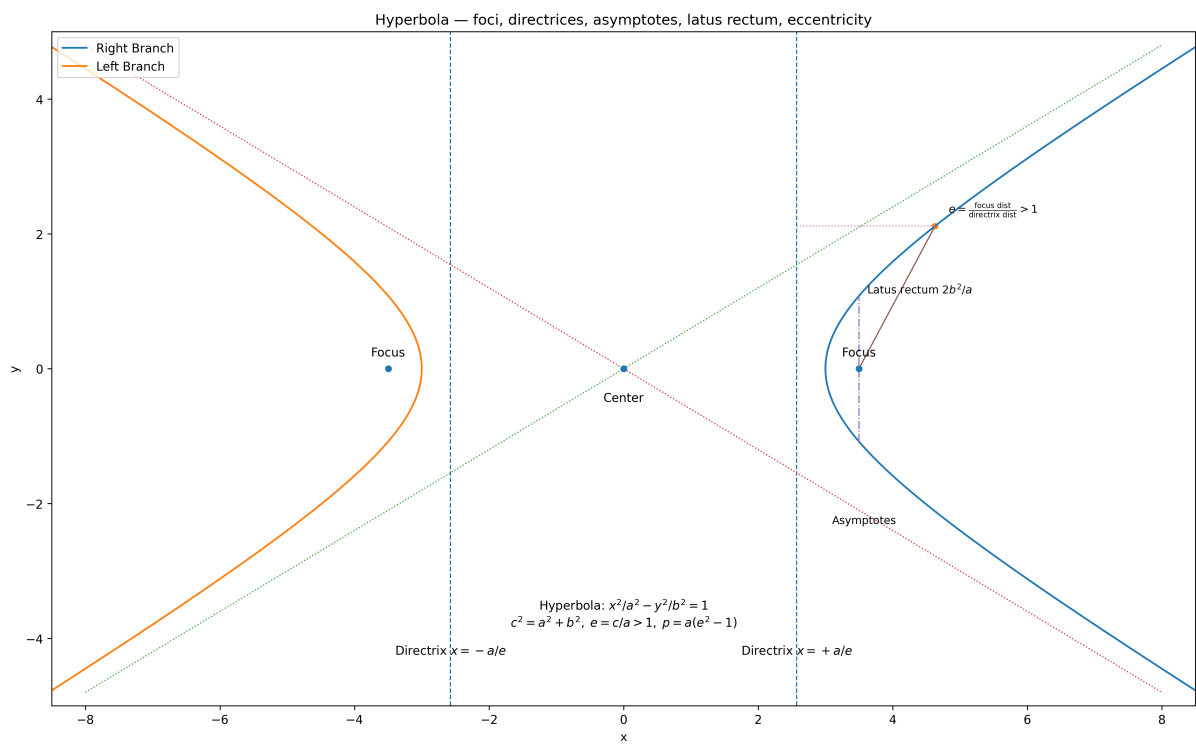


Figure 2: Caption

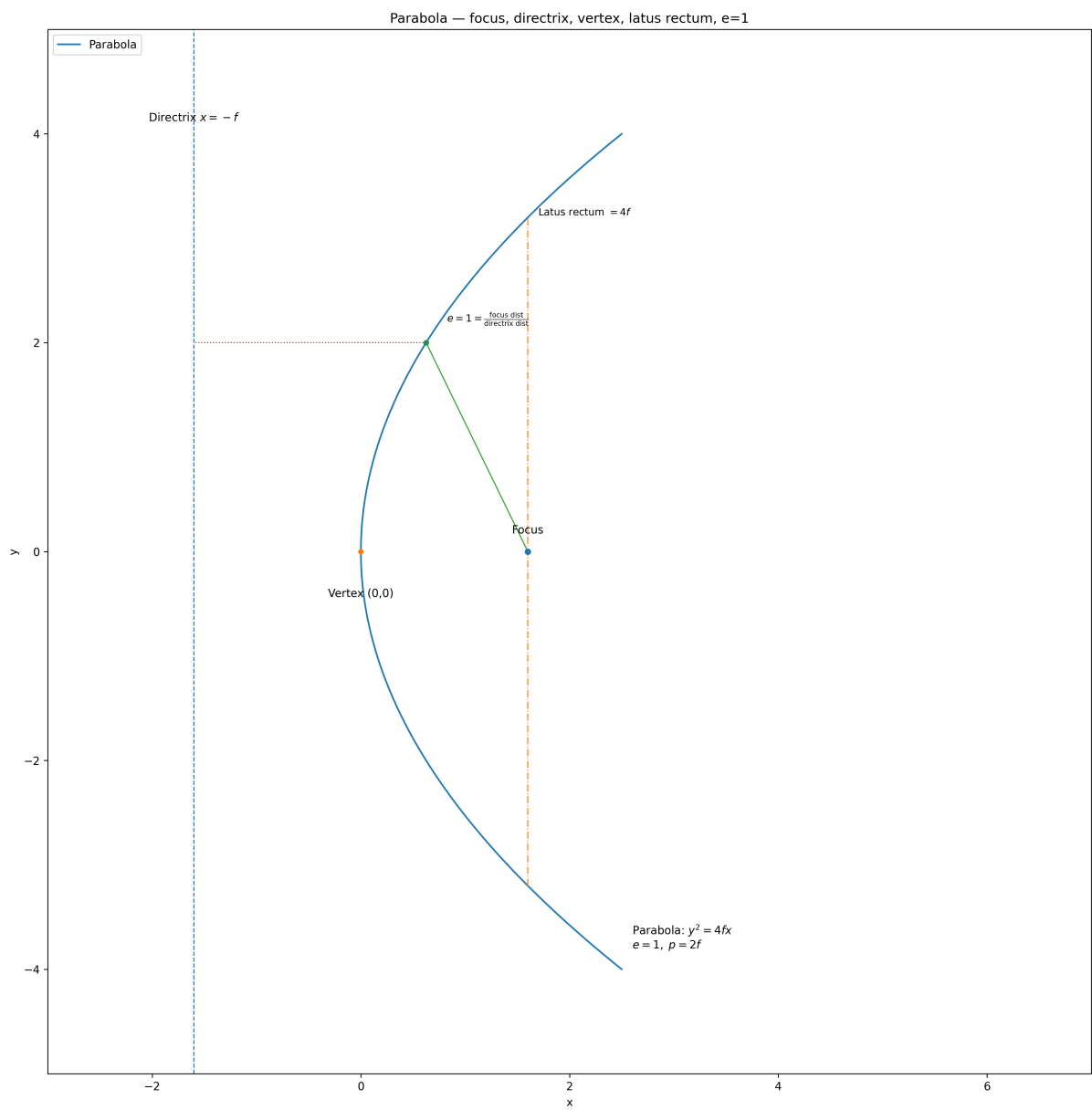


Figure 3: Caption

Procedure

1. Set the values of m, K, L, α and the eccentricity e .

2. Using the formula

$$r(\theta) = \frac{p_{\text{eff}}}{1 + e \cos(\beta(\theta - \theta_0))},$$

compute r for a range of θ values (e.g. 0 to 12π).

3. Plot $r(\theta)$ as a parametric curve in Cartesian coordinates:

$$x = r \cos \theta, \quad y = r \sin \theta.$$

4. Observe the effect of varying α and K on the orbit.

Observations

Typical results obtained from simulation:

| α | Nature of Orbit | Remarks |
|----------|-----------------------------|--------------------------------|
| 0.0 | Closed ellipse | No precession (Keplerian) |
| 0.02 | Slightly precessing ellipse | Small prograde shift per orbit |
| 0.05 | Rosette-like orbit | Significant precession visible |

Result

The motion of a particle under the central force

$$F = -\frac{K}{r^3} \left(1 - \frac{\alpha}{r}\right) \mathbf{r}$$

is a **precessing conic section**. The amount of precession per revolution is given by

$$\Delta\varpi = 2\pi \left(1 - \frac{1}{\sqrt{1 + \frac{mK\alpha}{L^2}}} \right).$$

The precession increases with K and α and decreases with larger angular momentum L .

Precautions

1. Use small step sizes in the computation for higher accuracy.
2. Ensure that the initial conditions correspond to a bound orbit ($e < 1$).
3. Avoid very large α values to prevent numerical divergence.
4. Check that angular momentum $L = mr^2\dot{\theta}$ remains constant throughout the computation.

Python Code for Conic Sections

Python Simulation Code: Central Force Motion

In the code avoid every that is in green. That is for your understanding

```
1  #
2  #Central Force Motion:  $F = -(K/r^3) * (1 - \alpha/r) * r_{\text{vec}}$ 
3  #-----
4  #This program visualizes the motion of a particle under the
5  #given central force. Instead of integrating with time, we
6  #use the analytic form of the orbit:
7
8  #  $r(\theta) = p_{\text{eff}} / [1 + e * \cos(\beta * (\theta - \theta_0))]$ 
9
10 #where:
11 #  $p_{\text{eff}} = (L^2 + m*K*\alpha) / (m*K)$ 
12 #  $\beta = \sqrt{1 + (m*K*\alpha)/(L^2)}$ 
13 #  $e = \text{eccentricity-like parameter}$ 
14 #  $\theta_0 = \text{phase angle (set to 0 here)}$ 
15
16 #This gives a precessing conic (rosette-like orbit). When
17 # $\alpha = 0$ , the orbit reduces to a perfect Keplerian ellipse.
18 #
19 #####
20
21 # -----
22 # Import required packages
23 # -----
24 import numpy as np
25 import matplotlib.pyplot as plt
26
27 # -----
28 # Define constants and parameters
29 # -----
30 m = 1.0          # mass of the particle (arbitrary units)
31 K = 1.0          # central force constant (attractive)
32 alpha = 0.03     # small positive number --> precession occurs
33 L = 1.0          # angular momentum (assumed constant)
34 e = 0.25         # eccentricity-like parameter ( $0 < e < 1$ )
35 theta0 = 0.0     # initial angular phase ( $r$  is max when  $\theta =$ 
36                   # 0)
37
38 # -----
39 # Derived quantities from the theory
40 # -----
41 beta = np.sqrt(1 + (m * K * alpha) / (L ** 2))
42 p_eff = (L ** 2 + m * K * alpha) / (m * K)
43
44 # -----
```

```

44 # Compute r(theta) for a range of theta values
45 # -----
46 theta = np.linspace(0, 12 * np.pi, 2000)
47 r = p_eff / (1 + e * np.cos(beta * (theta - theta0)))
48
49 x = r * np.cos(theta)
50 y = r * np.sin(theta)
51
52 # -----
53 # Plot in the xy-plane (spatial trajectory)
54 # -----
55 plt.figure(figsize=(7, 6))
56 plt.plot(x, y, linewidth=1.2)
57 plt.scatter([0], [0], color='black', s=20, label='Center of
    Force')
58 plt.axis('equal')
59 plt.xlabel('x')
60 plt.ylabel('y')
61 plt.title('Orbit under
    $F=-(K/r^3)(1-\\alpha/r)\\mathbf{r}$')
62 plt.grid(True)
63 plt.legend()
64 plt.show()
65
66 # -----
67 # Plot r(theta) parametric curve (r vs theta)
68 # -----
69 plt.figure(figsize=(7, 4.5))
70 plt.plot(theta, r, linewidth=1.2)
71 plt.xlabel(r'$\theta$ (radians)')
72 plt.ylabel(r'$r(\theta)$')
73 plt.title('Parametric plot $r(\theta)$ showing precession')
74 plt.grid(True)
75 plt.show()
76
77 # -----
78 # Print theoretical precession per orbit
79 # -----
80 Delta_varpi = 2 * np.pi * (1 - 1 / beta)
81
82 print(f'beta (precession parameter): {beta:.6f}')
83 print('Predicted apsidal advance per revolution (radians):
    {Delta_varpi:.6f}')
84 print(f'In degrees: {np.degrees(Delta_varpi):.3f}')

```

Listing 1: Simulation of a particle under the central force $F = -(K/r^3)(1 - \alpha/r)\mathbf{r}$ showing precessing orbit.