

HYBRID METHODS FOR SIMULATION
OF MUON IONIZATION COOLING CHANNELS

BY
JOSIAH D. KUNZ

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Approved _____
Advisor

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LIST OF SYMBOLS

Symbol	Definition
$g(u)$	Angular distribution function
θ_0	Angular distribution Gaussian width
s	Arc length coordinate
Z	Atomic charge
N	Atomic density
N_A	Avagadro's Number
(p_x, p_y, p_z)	Beamline momentum coordinate system
(x, y, z)	Beamline position coordinate system
u_0	Characteristic cosine variable
Q_i	Charge number of particle i
π	Circle constant
$*$	Complex conjugate
K	Constant (context dependent)
$u = \cos \theta$	Cosine variable (context dependent; see also 'Dummy variable')
ζ	COSY amplitude of scattering tail
b_c	COSY offset of scattering tail
Σ	Cross section
ρ	Density
δ	Density correction parameter (straggling; see also 'Dirac function')
δ	Dirac function (see also 'density correction parameter')

γ^α	Dirac gamma matrix
f	Distribution function (general; context dependent)
h	Distribution function (general; context dependent)
F	Distribution function antiderivative (general; context dependent)
H	Distribution function antiderivative (general; context dependent)
u	Dummy variable (context dependent; see also ‘Cosine variable’)
z_{ch}	Electric charge
N_{el}	Electron density
e^-, e^+	Electron or antielectron
m_e	Electron mass
$\nu_e, \bar{\nu}_e$	Electron neutrino or electron antineutrino
r_e	Electron radius
ϵ	Emittance (see also ‘energy loss fluctuation’)
E	Energy (general; context dependent)
$g(E)$	Energy loss distribution function
$\Delta E = \epsilon$	Energy loss fluctuation (see also ‘emittance’)
$\langle dE/dx \rangle$	Energy loss per unit length, mean (Bethe Bloch)
C_{Euler}	Euler’s constant (≈ 0.577)
P	Four-momentum
e	Fundamental charge (such that $z_{ch} = eQ$)
a, b_r, d_r, p_g, q_g	GEANT4 angular parameters

z_g	Geometric path length
\dagger	Hermitian conjugate (see also ‘transpose conjugate’)
h_i	Highland corrections ($i = 1, 2$)
b	Impact parameter
I	Ionization energy (mean)
λ	Landau parameter
ϕ	Laplace transformed function
P_k	Legendre polynomials
L	Length (of absorber or step size)
γ	Loretnz factor
m	Mass (general; context dependent)
\mathcal{M}	Matrix element (scattering) (see also ‘Scattering amplitude’)
\mathcal{A}, \mathcal{B}	Matrix, general
T_{max}	Maximum transferrable kinetic energy
μ	Mean (context dependent; see also ‘Muon’)
η	Minkowski metric
M	Moments of a function
p	Momentum (total)
μ	Muon (context dependent; see also ‘Mean’)
m_μ	Muon mass
ν_μ	Muon neutrino
C, C_i	Normalization constant (see also ‘Shell correction parameter’)

f_i	Oscillator strengths
σ^j	Pauli matrices
h	Planck's constant
\hbar	Planck's constant divided by 2π
X_0	Radiation length
θ	Scattered angle
\mathcal{M}	Scattering amplitude (context dependent; see also 'transfer map')
$g(u)$	Scattering function
C	Shell correction parameter (for straggling; see also 'Normalization constant')
Ω	Solid angle
c	Speed of light in a vacuum
U	Spinor
σ	Standard deviation
$g(E)$	Straggling function
t	Time (variable)
ℓ	Time-of-flight in units of length (COSY coordinate)
λ_k	Transport free mean paths, k^{th} value
\dagger	Transpose conjugate (see also 'Hermitian conjugate')
T	Transpose of a matrix
T	Transverse coordinate
t_z	True path length

I_j	Unit matrix of rank j
\vec{v}	Unit vector
λ_v	Vavilov parameter
κ	Vavilov limit parameter
v	Velocity
ψ	Wavefunction, time-independent component
w	Weight for a distribution function (context dependent)

ABSTRACT

COSY Infinity is an arbitrary-order beam dynamics simulation and analysis code. It can determine high-order transfer maps of combinations of particle optical elements of arbitrary field configurations. For precision modeling, design, and optimization of next-generation muon beam facilities, its features make it a very attractive code. New features are being developed for inclusion in COSY to follow the distribution of charged particles through matter. To study in detail some of the properties of muons passing through material, the transfer map approach alone is not sufficient. The interplay of beam optics and atomic processes must be studied by a hybrid transfer map–Monte Carlo approach in which transfer map methods describe the average behavior of the particles in the accelerator channel including energy loss, and Monte Carlo methods are used to provide small corrections to the predictions of the transfer map accounting for the stochastic nature of scattering and straggling of particles. The advantage of the new approach is that it is very efficient in that the vast majority of the dynamics is represented by fast application of the high-order transfer map of an entire element and accumulated stochastic effects as well as possible particle decay. The gains in speed are expected to simplify the optimization of muon cooling channels which are usually very computationally demanding due to the need to repeatedly run large numbers of particles through large numbers of configurations. Progress on the development of the required algorithms is reported.

CHAPTER 1

INTRODUCTION

1.1 Muon-based Accelerators

Muons (μ) were first discovered experimentally in 1947 by Powell *et al.* [1] who were looking for the Yukawa meson. It is now known that muons in fact fit into a fundamental particle group called leptons, and fit into the standard model as shown in Figure 1.

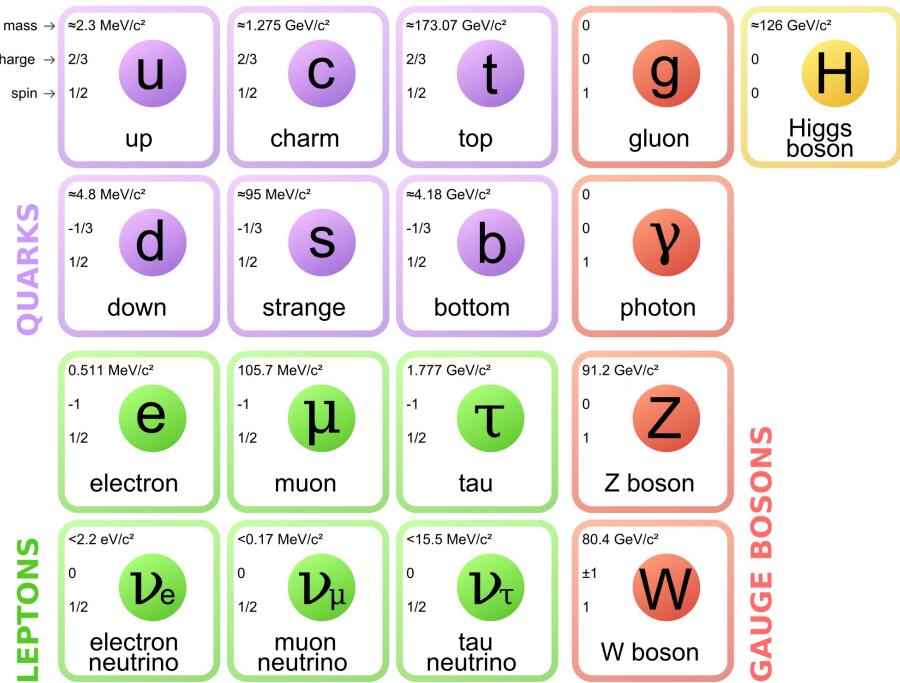


Figure 1.1: The current model of particle physics. Image courtesy of [2].

Similar to the electron (e), the muon carries a fundamental charge of ± 1 , a spin of $1/2$, and observes the electromagnetic and weak forces. Moreover, the muon also has a corresponding neutrino: the muon neutrino (ν_μ). However, the muon (mass = $105.7 \text{ MeV}/c^2$) is about 200 times heavier than the electron (mass = $0.511 \text{ MeV}/c^2$).

Indeed, sometimes it is useful to think of a muon simply as a heavy electron, but the mass implies several unique characteristics. One of these are the instability of muons, since the following process is not forbidden by mass, charge, or lepton number conservation:

$$\mu \rightarrow e + \bar{\nu}_e + \nu_\mu.$$

This is quite interesting, as it means muons are a double-edged sword. On the one hand, their fundamentality means that the muon collisions will be clean. That is to say, muon colliders have a great advantage over, e.g., proton colliders since protons are composed of three quarks. Each quark may have a different flavor or energy level and hence adds more variables to the analysis. Furthermore, each quark is bound and so gluon interactions must also be considered. These quarks and gluons may undergo hadronization when interacting with one another, creating a plethora of possible hadrons. Hadronization is not fully understood, and so these sprays of hadrons are typically lumped together as a single "jet". While there are many working models for jet analysis, none are exact. Conversely, any data from muon interactions will have relatively little noise and will not produce jets.

Clean collisions can also be achieved with linear electron colliders. Yet unlike the electron, the muon does not emit a large amount of synchrotron radiation as it is accelerated. This is because the power irradiated off a particle due to synchrotron radiation is $\propto 1/m^4$ [3]. Therefore, the relative power loss due to synchrotron radiation for electrons and muons is $P_e/P_\mu \propto m_\mu^4/m_e^4 \approx 1.8E9$. For this reason, it is possible to have a circular muon accelerator. Furthermore, due to the small mass of a muon compared to a proton (~ 105 MeV/ c^2 vs ~ 938 MeV/ c^2), muons will be easier to accelerate. This means that a muon facility will be much smaller than its proton counterpart.

However, there is one problem for a muon accelerator. A rest frame lifetime of $2 \mu\text{s}$ requires the muons to be accelerated quickly before they decay. This is a challenge for circular colliders which require a high-intensity beam.

Overall, it appears that there are several advantages and one key disadvantage: the $2 \mu\text{s}$ mean rest frame lifetime of the muon. However, this is only a disadvantage for muon colliders. Another application which turns the moderately short lifetime into an advantage is a neutrino factory. This is a facility which is dedicated to the output of a neutrino beam. Muons have two primary advantages over fission reactor neutrino sources. The first is that muons decay into exactly two flavors of neutrino: electron and muon. Therefore, the initial composition of the neutrino beam would be roughly known. This is important since neutrinos can change their flavors over time. Secondly, since neutrinos have no electric charge, they cannot be manipulated via electromagnetic focusing methods. However, the beam of muons can be focused into a high-intensity beam, and the intensity of the neutrino beam will reflect this.

1.2 COSY Infinity

COSY Infinity is a beamline simulations tool used in the design, analysis, and optimization of particle accelerators [4]. COSY uses the transfer map approach, which evaluates the overall effect of a system on a beam of particles using differential algebra. This involves phase space vectors being expanded into multivariate Taylor polynomials up to arbitrary order [5]. An example of a phase space vector can be seen in Equation 1.1, where the coordinates are transverse positions (x, y), time-of-flight in units of length (l), transverse angles w.r.t. the reference particle (a, b), and kinetic energy deviations w.r.t. the reference particle (δ). The 0 subscript in the definitions denotes the reference particle properties. However, any arbitrary “coordinates” such as spin would also work in principle.

$$\mathbf{Z} = \begin{pmatrix} x \\ y \\ l = k(t - t_0) \\ a = p_x/p_0 \\ b = p_y/p_0 \\ \delta = (E - E_0)/E_0 \end{pmatrix} \quad (1.1)$$

In beam physics, phase space vectors \mathbf{Z} are subject to physical processes. This subjugation can usually be represented by a differential equation. For example, a particle in an electric field is subject to the force law [3]

$$\mathbf{F} = Q(\mathbf{E} + \mathbf{v} \times \mathbf{B}),$$

or using Maxwell's equations

$$\frac{d}{dt}\mathbf{p} = Q(\mathbf{E} + \frac{1}{m}\mathbf{p} \times \mathbf{B}).$$

This is only the equation for the momentum-dependent components of the aforementioned \mathbf{Z} , a and b . However, the explicit dependence is easily worked out.

Fortunately, any phase space vector \mathbf{Z} in an arbitrarily-ordered ordinary differential equation (ODE) can be rewritten as a first-order ODE [5]. For an order n ODE, a first-order ODE is constructed by introducing $n - 1$ new variables. This is to say that

$$\frac{d^n}{ds^n}\mathbf{Z} = \mathbf{f}\left(\frac{d^0}{ds^0}\mathbf{Z}, \dots, \frac{d^{n-1}}{ds^{n-1}}\mathbf{Z}\right)$$

can be rewritten as

$$\frac{d}{ds} \begin{pmatrix} \mathbf{Z} \\ \mathbf{Z}_1 \\ \vdots \\ \mathbf{Z}_{n-1} \end{pmatrix} = \begin{pmatrix} \mathbf{Z}_1 \\ \mathbf{Z}_2 \\ \vdots \\ \mathbf{f}(\mathbf{Z}, \dots, \mathbf{Z}_{n-1}) \end{pmatrix}.$$

Here, \mathbf{f} represents the real physics processes. For the Maxwell's equations example, the equation is first-order already with the momentum component of \mathbf{f} as

$$\mathbf{f}(\mathbf{p}) = Q(\mathbf{E} + \frac{1}{m}\mathbf{p} \times \mathbf{B}).$$

Furthermore, the independent variable in principle does not matter. COSY does not use time, but rather arc length s (see Figure 1.2)).

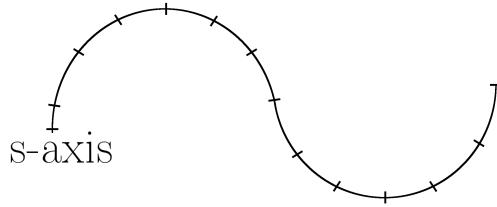


Figure 1.2: The reference orbit. Figure courtesy of [6].

Now that there is a system of equations, the transfer map \mathcal{M} will uniquely predict the evolution of \mathbf{Z} . This is because most beam elements follow differential equations which yield unique solutions dependent on initial conditions (such as Maxwell's equations). Mathematically, this relationship is $\mathbf{Z}(s) = \mathcal{M}(s_0, s) * \mathbf{Z}(s_0)$, with the $*$ representing the composition of functions. An example of this relationship can be seen in Figure 1.3. The initial phase space occupied by the beam of particles is at the coordinate s_0 . Physically, there exists some deterministic beamline element. This element can be represented by the map \mathcal{M} , which creates a bijection for the

phase space vectors $\mathbf{Z}(s_0)$ and $\mathbf{Z}(s)$ between the initial coordinate s_0 and an arbitrary final coordinate s .

Entire lattices may also be represented by a single transfer map. This is done by dividing the lattice into its base components or elements. A transfer map for each corresponding element may then be produced. For two example elements between coordinates s_0 and s_2 (see Figure 1.4), the composition of two maps yeilds another map: $\mathcal{M}(s_1, s_2) \times \mathcal{M}(s_0, s_1) = \mathcal{M}(s_0, s_2)$. Therefore, it is possible to “cut out” the middle part s_1 . In this way, the transfer maps from small components build up into a single transfer map for the whole system. Computationally this is advantageous because once calculated, it is much faster to apply a single transfer map to a distribution of particles than to simulate that same distribution through many meters of individual lattice elements.

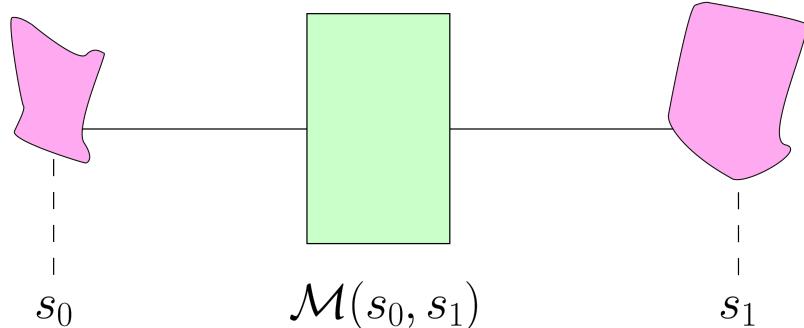


Figure 1.3: Example of some map \mathcal{M} creating a bijection from s_0 to s_1 .

While a transfer map is technically not a nonlinear matrix, it is sometimes helpful to think of them as one in the same. Along with the evaluation of particles through a lattice, COSY also has a plethora of analysis and optomization tools, including (but not limited to) lattice aberration and correction tools, support for

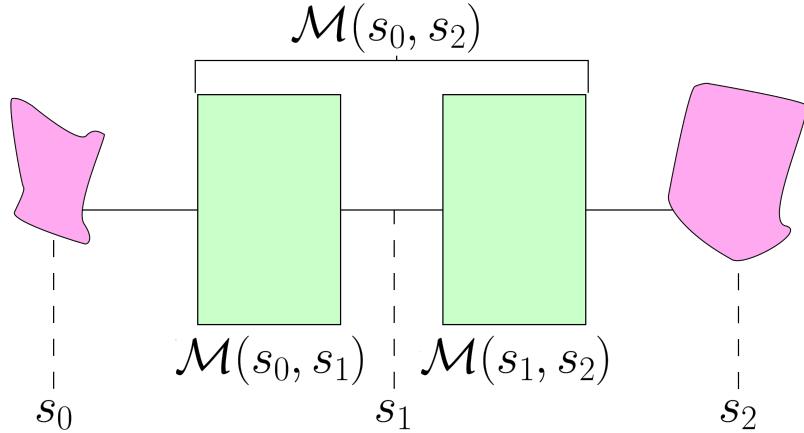


Figure 1.4: Example of two maps $\mathcal{M}(s_0, s_1)$ and $\mathcal{M}(s_1, s_2)$. These two maps may be combined together to reduce to a single map, $\mathcal{M}(s_0, s_2)$.

Twiss parameters, support for tunes and nonlinear tune shifts, built-in optimizers (for lattice design), and spin tracking.

Valid elements are any beamline tools which are deterministic. Elements used in this study are magnetic multipoles (dipoles, quadrupoles, etc.), solenoidal coils, RF cavities, and drifts. Currently supported elements in COSY include but are not limited to: various magnetic and electric multipoles (with fringe effects), homogeneous and inhomogeneous bending elements, Wien filters, wigglers and undulators, cavities, cylindrical electromagnetic lenses, general particle optical elements, and deterministic polynomial absorbers of arbitrary order, with the last element being of particular interest.

1.2.1 Emittance. Often, it is advantageous to represent the volume of phase space occupied by the beam by the emittance, ϵ . It is sometimes beneficial to normalize the emittance by the relativistic factors, γ and β . This is useful because when a beam accelerates the transverse angles decrease. This results in a reduction of ϵ_x (where x represents the general transverse direction) even though nothing was done to the

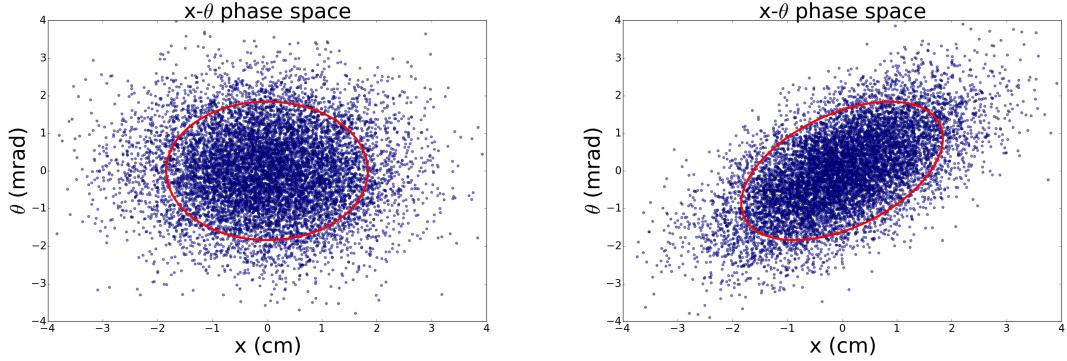


Figure 1.5: x - θ phase space examples. Left: example beam with no x - θ correlation ($\langle x\theta \rangle = 0$). Right: example beam with significant x - θ correlation ($\langle x\theta \rangle = 0.8$).

transverse phase space directly. However, ϵ_x^N stays constant when a beam accelerates. For the relevant transverse emittance,

$$\epsilon_x^N = \beta\gamma\epsilon_x = \beta\gamma\sqrt{\langle x^2 \rangle \langle \theta^2 \rangle - \langle x\theta \rangle^2}. \quad (1.2)$$

Here, azimuthal symmetry is assumed, and so x represents the general transverse position, and $\theta = \sin^{-1} p_x/p$ is the projection of the divergence angle of the particle trajectory onto the x - z plane, where z is the longitudinal Cartesian coordinate.

The full derivation of emittance can be found in Appendix A. It is assumed that the beam is nearly Gaussian in both x and θ . Qualitatively, if there exists no cross-dependence in x and θ , then the distribution should be a non-tilted ellipse. Then the volume of this bivariate Gaussian is proportional to the geometric mean, $\sqrt{\langle x^2 \rangle \langle \theta^2 \rangle}$. However, if there is some cross-dependence then the result is a tilted ellipse. For a sample comparison of tilted vs. non-tilted ellipses, observe Figure 1.5. Here, the emittance is outlined in red.

Now it is possible to derive the effect of cooling absorbers on emittance as it pertains to muon ionization cooling. Closely following [7], Eqn. 1.2 is differentiated

by z :

$$\frac{d\epsilon_x^N}{dz} = \epsilon_x \frac{d(\beta\gamma)}{dz} + \beta\gamma \frac{d\epsilon_x}{dz}, \quad (1.3)$$

It will soon be shown that the first term represents ‘cooling’ and the second term represents ‘heating’ in the transverse plane. Starting with the heating term,

$$\frac{d\epsilon_x^N}{dz}(\text{heat}) = \beta\gamma \frac{d\epsilon_x}{dz}.$$

It can be assumed that the cooling is taking place near the beam waist (that is, the part of the trajectory where the beam spread is smallest). For this case, it is a reasonable approximation that the transverse phase space has little $x - \theta$ cross-dependence, and so the $\langle x\theta \rangle$ term may be neglected. Furthermore, with strong focusing it is possible to neglect the rate of position growth of the beam. This is cited in [7] provided the conditions that $\sigma_{xo}^2 \gg \theta_c^2 L / 2\omega^2$ and $\sigma_{xo}^2 \gg \theta_c^2 / 4\omega^3$, where L is the absorber length, ω is the focusing strength parameter, σ_{xo} is the original beam size incident upon the absorber, and θ_c is the coefficient of the RMS scattering angle given in [8]. Using these two approximations, the only derivative in the heating term is that of $\langle \theta^2 \rangle$:

$$\frac{d\epsilon_x^N}{dz}(\text{heat}) \approx \beta\gamma \frac{d}{dz} \sqrt{\langle x^2 \rangle \langle \theta^2 \rangle} \approx \frac{\beta\gamma}{2\epsilon_x} \langle x^2 \rangle \frac{d}{dz} \langle \theta^2 \rangle.$$

From betatron focusing theory, $\langle x^2 \rangle$ may be rewritten as $\beta_\perp \epsilon_x$. Moreover, using [8] $\langle \theta^2 \rangle$ can be written as $h(z)/\beta^2 E$. Here, $h(z)$ is the Highland z -dependence, the form of which many theories disagree. However, the first order term of $h(z)$ is generally not model-dependent, and so this becomes

$$\langle \theta^2 \rangle \approx \frac{E_s}{\beta^2 E} \sqrt{\frac{z}{X_0}},$$

where E_s is some characteristic energy (≈ 14 MeV) and X_0 is the radiation length of the given material. Then

$$\frac{d\epsilon_x^N}{dz}(\text{heat}) \approx \beta\gamma \frac{\beta_\perp}{2} \frac{d}{dz} \left(\frac{E_s}{\beta^2 E} \sqrt{\frac{z}{X_0}} \right),$$

$$\frac{d\epsilon_x^N}{dz}(\text{heat}) \approx \frac{\beta_\perp}{2} \frac{E_s^2}{\beta^3 Emc^2} \frac{1}{X_0}. \quad (1.4)$$

Now it is clearer why this is referred to as the heating term. Since the derivative of emittance is positive, this part represents emittance growth. Moreover, to reduce this growth it is necessary to have highly energetic particles through a material with a large radiation length (typically $X_0 \propto 1/Z$ (nuclear charge)).

Observe the cooling term of Eqn. 1.3:

$$\frac{d\epsilon_x^N}{dz}(\text{cool}) = \epsilon_x \frac{d(\beta\gamma)}{dz} = \epsilon_x \cdot (\beta \frac{d\gamma}{dz} + \gamma \frac{d\beta}{dz}).$$

First,

$$\frac{d\beta}{dz} = \frac{d}{dz}(1 - \gamma^{-2})^{\frac{1}{2}} = \frac{1}{\beta\gamma^3} \frac{d\gamma}{dz},$$

and then,

$$\frac{d\gamma}{dz} = \frac{\gamma}{E} \frac{dE}{dz},$$

with E as the total energy of the muon beam. Finally,

$$\frac{d\epsilon_x^N}{dz}(\text{cool}) = \epsilon_x \cdot (\beta \frac{d\gamma}{dz} + \gamma \frac{1}{\beta\gamma^3} \frac{d\gamma}{dz}) = \epsilon_x \cdot \frac{d\gamma}{dz} \beta \left(1 + \frac{1}{\beta^2\gamma^2}\right) \frac{d\epsilon_x^N}{dz}(\text{cool}) = \epsilon_x \frac{dE}{dz} \frac{\gamma}{E\beta}.$$

Normalizing the emittance by first multiplying and then dividing by $\beta\gamma$ results in

$$\frac{d\epsilon_x^N}{dz}(\text{cool}) = \frac{1}{\beta} \frac{dE}{dz} \frac{\epsilon_x^N}{E}$$

However, there are two corrections which should be observed. First, even for a perfect monoenergetic beam energy loss is stochastic, not deterministic. Two muons with identical initial conditions will likely lose different amounts of energy as they pass through the same medium. Since emittance is a collective effect, is it more appropriate to talk about an average energy loss, which turns $\frac{dE}{dz}$ into $\langle \frac{dE}{dz} \rangle$. The second observation is with regard to sign convention. Typically, one refers to energy loss as a positive quantity (e.g. the beam lost 10 MeV from point A to point B), even though

it is clear that $\frac{dE}{dz}$ is negative. For this reason, the energy loss term must again be changed from $\langle \frac{dE}{dz} \rangle$ to $-\left| \langle \frac{dE}{dz} \rangle \right|$. This leaves the full cooling term as

$$\frac{d\epsilon_x^N}{dz}(cool) = -\frac{1}{\beta} \left| \left\langle \frac{dE}{dz} \right\rangle \right| \frac{\epsilon_x^N}{E}. \quad (1.5)$$

Similar to the heating term, it is now understandable why this is referred to as the cooling term. The rate of change of the normalized transverse emittance is negative, and so the transverse phase space shrinks. However, at this point Eqn. 1.5 does not explain precisely how the rate changes, and so is purely conceptual. The shape of the cooling term is determined by the average energy loss, $\left| \langle \frac{dE}{dz} \rangle \right|$. However, usually it is not $\left| \langle \frac{dE}{dz} \rangle \right|$ that is measured, but rather the stopping power $S(E) = -\frac{1}{\rho} \langle \frac{dE}{dx} \rangle$, typically in units of MeV cm²/g. A good example of a stopping power curve can be seen in Figure 1.6 [9]. For this example, there are several effects noted in Figure 1.6, and one may be tempted to think that based on 1.5 the most cooling should occur in the Anderson-Ziegler regime at $\beta\gamma \sim 0.01$ (since there is a $1/E$ dependence in Eqn. 1.5, the high energy part of the curve does no good). Indeed, for $\rho = 8.92$ g/cm³ it is easy to see that at $\beta\gamma = 0.01$ then $\frac{d\epsilon_x^N}{dz}(cool) \approx -900 \text{ cm}^{-1} \cdot \epsilon_x^N$.

However, the correct strategy is just the opposite—one should endeavor to build a cooling cell with the minimum ionization point in mind (in the Bethe regime, where $\beta\gamma \sim 2$). At the peak ($\beta\gamma \sim 0.01$), the muon beam slows down quite a bit through a single cooling cell and one is at risk of losing their beam to decay. Moreover, Figure 1.6 only depicts the stopping power, not the actual energy loss that individual particles experience. The individual energy loss is a stochastic effect, and the distribution of energy losses broadens with decreasing beam energy. This means that for lower energies, there will be more particles which stop completely. Further, with a higher beam energy the stochastic energy loss profile is more sharply peaked, and so there are fewer decays and more particles which are in the RF stability region. (Remember, the longitudinal emittance is observed by the RF cavities, as seen in

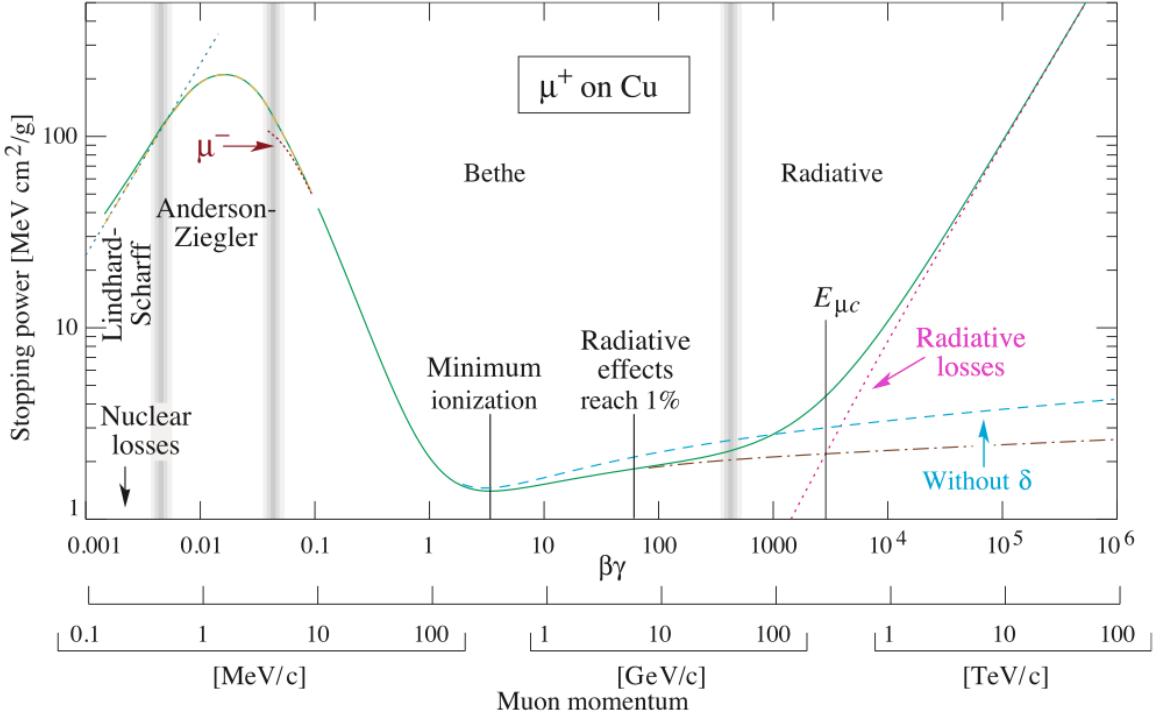


Figure 1.6: Stopping power curve for antimuons on copper, courtesy of [9].

Figure 1.10. The longitudinal emittance increases with energy loss according to [7].) Finally, Eqn. 1.4 suggests that higher energies are better, since they reduce the heating term. For these reasons, it is apparent that a muon cooling cell should be made of low Z materials such as liquid hydrogen or lithium hydride and operate anywhere between 100 MeV/ c to 400 MeV/ c .

When particles impinge upon a fixed target, they lose some of their energy. For massive particles like muons, there are a number of major effects that contribute. However, it is important to note that in the muon cooling regime only ionization contributes significantly to the energy loss. To see this, take for an example the relatively heavy element of iron (iron is chosen because it is a fairly relatable element to those who are not experts in this field). In Table 1.1, the energy loss contributions of these four effects can be seen. However, the non-ionization effects only start to

Muons in Iron (Fe)

T MeV	p MeV/ c	Ionization	Brems	Pair prod MeV cm ² /g	Photonucl	Total	CSDA range g/cm ²
10	47.04	5.494			5.494	5.494	1.025
14	56.16	4.321			4.321	4.321	1.854
20	68.02	3.399			3.399	3.399	3.437
30	85.09	2.654			2.654	2.654	6.812
40	100.3	2.274			2.274	2.274	10.91
80	152.7	1.717			1.717	1.717	31.78
100	176.4	1.616			1.616	1.616	43.82
140	221.8	1.516			1.516	1.516	69.50
200	286.8	1.463			1.463	1.463	109.9
274	364.2	1.451			1.451	<i>Min. ionization</i>	
300	391.7	1.452			1.453	1.453	178.7
400	494.5	1.467			1.467	1.467	247.2
800	899.5	1.548			1.548	1.548	512.4
1000	1101	1.581	0.001		1.582	1.582	640.2
1400	1502	1.635	0.001		0.001	1.637	888.5

Table 1.1: Energy loss table for muons in iron. Table courtesy of [10].

contribute at an initial beam kinetic energy of 1.40 GeV—far outside of the muon cooling regime.

This is a good example because all of the cooling materials currently being considered have less muon stopping power than iron. This means that iron represents the maximum contribution of non-ionization energy loss per initial beam energy; that is, the non-ionization effects begin to emerge at much higher energies for lighter materials such as liquid hydrogen. Note that Figure 1.6 may also be useful when discussing this subject, as the minimum ionization regime is explicitly characterized.

Here, *ionization* refers to both true ionization (the production of δ rays) and excitation (the promotion of an inner electron to a higher shell). To reiterate the previous section, the term ‘straggling’ refers to the fluctuation about some mean energy loss. That is to say, the amount of energy that a particle will lose is intrinsically random—there exists some distribution with some average value, and when the particle of interest transverses a length of matter the amount of energy that this particle loses is selected from this distribution. This intrinsic randomness can be attributed to quantum-like behaviors (e.g. the energy loss cross section which comes from the wave nature of the particle in question).

1.3 Introduction to Matter-Dominated Lattices

1.3.1 Introduction to Stochastic Effects. This section introduces stochastic effects, or effects that are intrinsically random. These effects are in contrast to deterministic effects, which are not random and hence can be predicted exactly.

This work is concerned with the interactions between a muon beam and some stationary target called the ‘absorber’—typically a cylinder or wedge of a low-z material such as liquid hydrogen or lithium hydride. The simplest model of a beam interacting with some stationary target can be found in several textbooks [11, 12], but perhaps the most helpful model comes from [13] in Figure 1.7.

In this figure b is referred to as the impact parameter and is measured with respect to the particle’s initial trajectory. For a beam of noninteracting particles and a perfectly stationary target, classical mechanics suggest that this is a purely deterministic problem (see, e.g., ‘hard-sphere scattering’ in [12]); the particle with a smaller impact parameter b will deflect more than its neighbor with a slightly larger impact parameter of $b + db$. This model is entirely based on initial conditions, and if this were reality it would be relatively easy to implement these effects into the

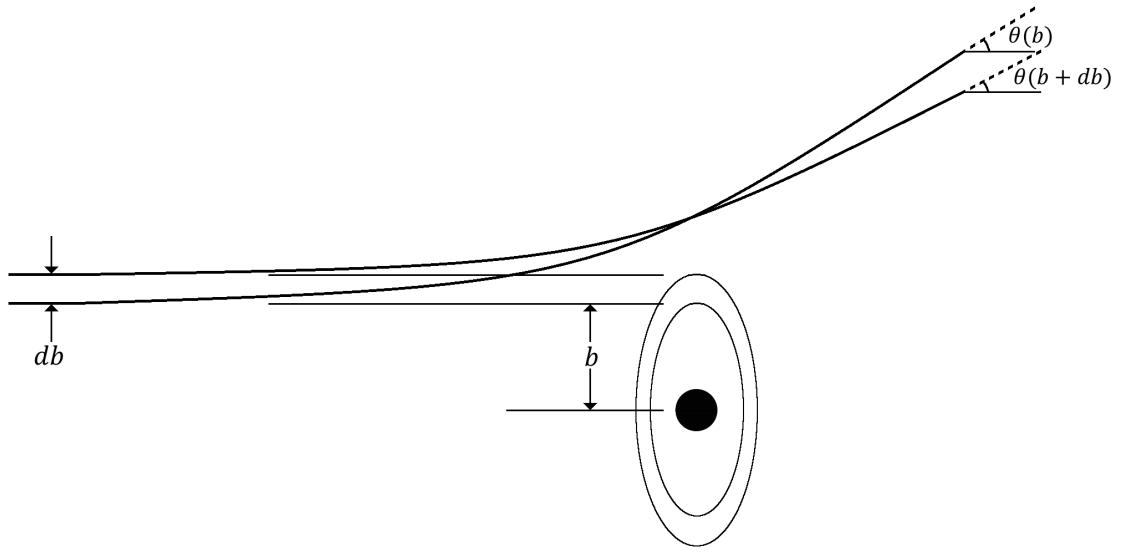


Figure 1.7: Classical muon–target interaction model courtesy of [13].

map methods of COSY Infinity. However, reality does not follow the model of Figure 1.7. A more accurate model is illustrated by Figure 1.8. Here the target is still approximated as fixed, but the incident particle is now approximated as a travelling plane wave and the scattered particle is approximated as a spherical wave (at least locally). It will be shown later that this model predicts

1. a spectrum of energy loss $\omega(\epsilon)$ which yields the probability of losing an amount of energy ϵ (see Section 2.1), and
2. a scattering amplitude which gives the shape of the probability distribution of scattering in a given direction θ (see Section 2.2).

Hence quantum theory suggests that even if two identical particles have identical initial conditions their final conditions will not be the same.

1.3.2 Muon Ionization Cooling. In Section 1.1, the primary disadvantage to the considered muon-based accelerator was the mean muon rest frame lifetime (2

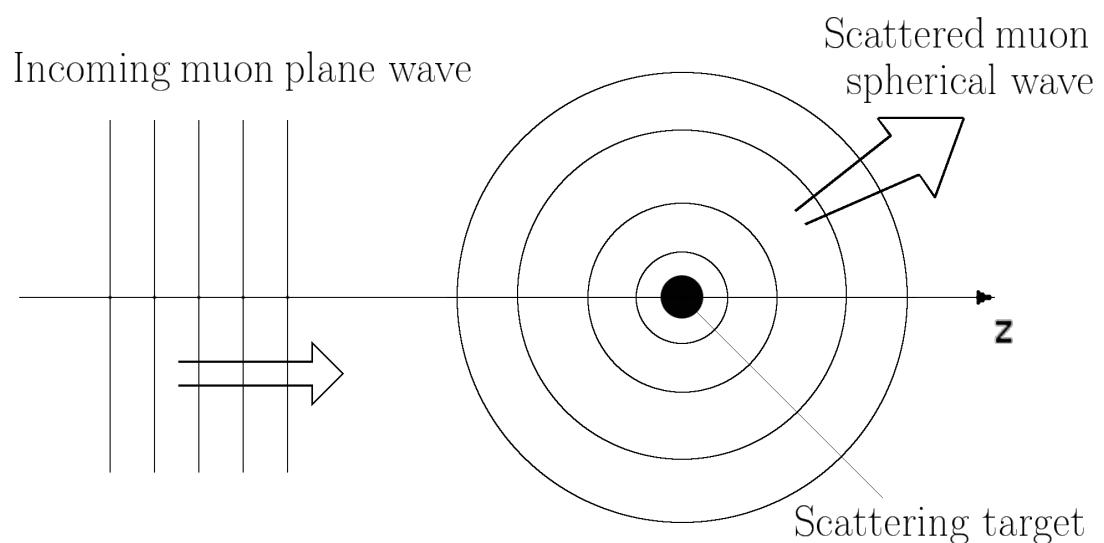


Figure 1.8: Quantum muon–target interaction model. The incoming particle wavefunction is represented as a plane wave and is scattered locally as a spherical wave. Image courtesy of [12].

μs). This is expected to be far too short of a timespan to be useful in a traditional accelerator scheme. For the moment, observe the two proposed schematics in Figure 1.9 [14]. The section labelled ‘Proton Driver’ produces protons, which will eventually yield muons. This is essential since muons do not naturally occur in great quantities at a convenient extraction point. To do this, the protons strike some large target (which must be optimized to produce the highest yield), resulting in a spray of protons, muons, electrons, and pions (even shorter-lived particles consisting of a quark-antiquark pair). To further optimize this process, it is advantageous to let the pions decay into muons via $\pi^+ \rightarrow \mu^+ + \nu_\mu$. The resulting ensemble of muons is then split up into bunches. The next section is where all of the interesting physics lies, and is the focus of this thesis. For now, it will be simply addressed as the ‘cooling channel’ and delved into later. Cooling the beam simply means to reducing the beam’s phase space. Reducing the momenta of the beam will result in a reduced physical size. The purpose of cooling is to increase the density of the resulting beam. This lets the beam fit into smaller magnets, which are less expensive than larger magnets. However, this increase in density also increases the luminosity of the beam in the long run. For a collider, this means more collisions; for a neutrino factory, this means more neutrino counts in the detectors (since neutrinos are neutral, the ν beam would disperse too early). After this critical step, the beam accelerates to an appropriate energy. For a neutrino factory, this is 5 GeV, and the muon and antimuon beams are allowed to decay in a storage ring. For a muon collider, the center-of-momentum energy is ~ 126 GeV for a Higgs factory or up to 10 TeV for high energy studies.

The cooling channel is the crux of the entire operation for either purpose. Traditional cooling methods (e.g. FODO, electron cooling, etc.) are expected to take far too long to allow a significant portion of the muon beam to survive. However, ionization cooling is a much faster method. Ionization cooling requires the beam to deposit its energy in matter, thus ionizing the matter.

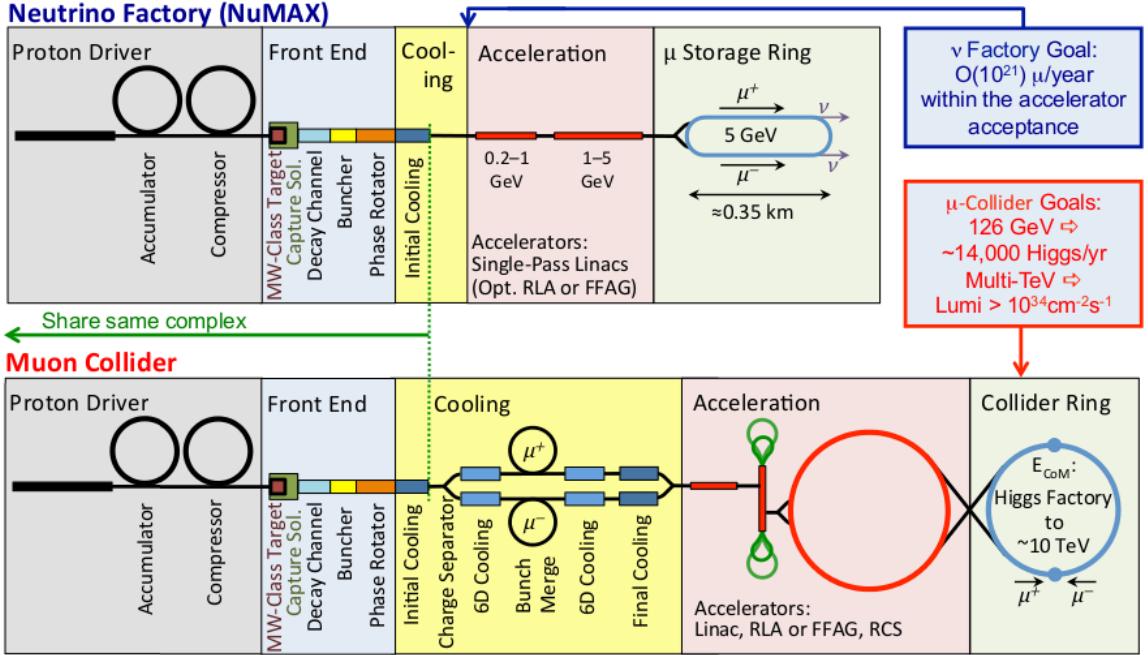


Figure 1.9: Proposed muon accelerator shematics. Image courtesy of [14].

While the idea of ionization cooling has been around since at least 1956 [15,16], it did not appear to be a viable option until roughly 1970 [17]. It was believed that the multiple scattering effect would mask any cooling benefits. Scattering is a quantum effect which deflects two objects when they are in close proximity at high energies, and hence is intrinsically random. This means that while the energy deposition reduced the beam's phase space, the beam would also grow in the transverse direction due to this multiple scattering.

Now, this seems to be quite the opposite of what is intended; ionization cooling potentially produces a slow, fat beam, and a good muon beam should be narrow and fast so that the Lorentz boost keeps the muons from decaying. Fortunately, there exists a schematic that solves both of these problems simultaneously. Physically, it is represented in Figure 1.10 and it is vectorally depicted in Figure 1.11.

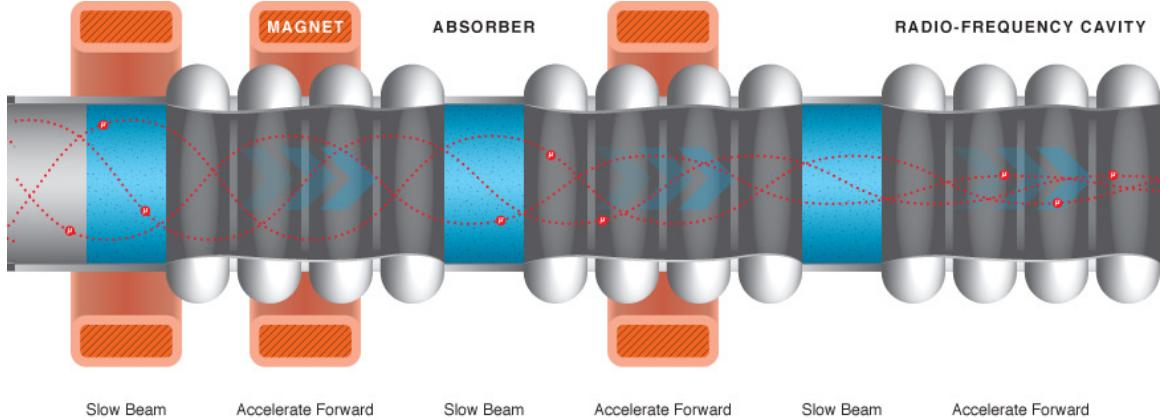


Figure 1.10: Cartoon of a cooling channel, courtesy of [14].

The key is to use a cell which contains both the ionizing material and a radio frequency cavity. It is clear why this solves the latter problem: after the beam loses energy, it gains that much energy again, and so the Lorentz boost stays constant on average. However, to understand why this solves the former problem it is necessary to observe Figure 1.11. In Figure 1.11, the longitudinal momentum (p_l) is plotted against the transverse momentum (p_t).

1. Beam deposits energy in material, reducing the momentum in all directions.
2. Multiple scattering effects are observed. The transverse phase space increases (or ‘heats up’).
3. The beam is re-accelerated by the RF cavity, increasing the longitudinal momentum only. The result is a reduction in transverse momentum (i.e. blue vector compared to black vector).

1.4 Looking Forward

In short, this work endeavors to take these and other basic principles to successfully implement muon ionization cooling simulation tools into COSY Infinity. The

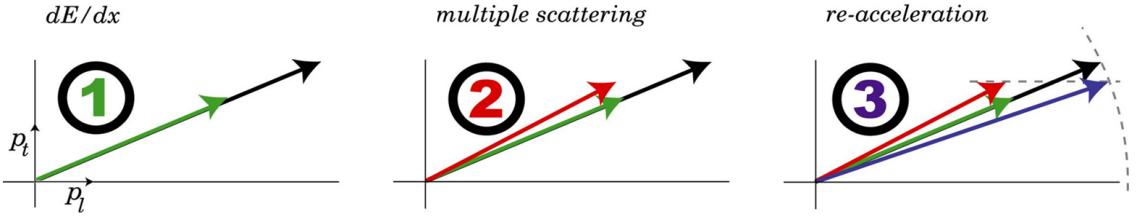


Figure 1.11: Vector diagram illustrating the principle of ionization cooling in Figure 1.10.

routines which will be implemented are derived from first or second principles as will be seen from other codes in Chapter 2 and novelly in Chapter 3. Furthermore, these routines will augmented by fitting the results to experimental data and benchmarked against external sources. The results of benchmarking and validation will be discussed in Chapter . Here, it will be seen that the new routines in COSY produce robust, accurate, and fast simulations.

CHAPTER 2

STOCHASTIC PROCESSES IN OTHER CODES

In this chapter, the two major stochastic processes of energy straggling and multiple scattering are detailed. Because there are so many models for these two key processes, two case studies will be shown as a frame for their derivations. Both of these codes use particle-by-particle propagation and do not use transfer maps.

The first case study is ICOOL [18] and will cover the classic models. Much of the classic theory is also used in the new COSY routines. Consequently, the sections concerning ICOOL serve as a foundation for the derivations in Chapter 3. ICOOL was created at Brookhaven National Laboratory for the benefit of the members of the Neutrino Factory and Muon Collider Collaboration, who specifically study ionization cooling problems.

The second case study is G4Beamline [19], which is based on GEANT4 [20]. GEANT4 implements a more novel approach. Many of the routines use empirically-tuned parameters. Consequently, G4Beamline is accurate for a variety of particles over a wide range of energies. It should be noted that muons, which have little experimental data available, are usually grouped with protons as heavy charged particles. Therefore, for lack of experimental data muons use the same routines as protons.

2.1 Energy Straggling in ICOOL

ICOOL [18] employs four levels of straggling models, three of which were used in these muon simulations studies. Discussed below, these are

1. Gaussian (Bohr),

2. Landau distribution, and
3. Vavilov distribution (with appropriate limits).

The fourth model is “restricted energy fluctuations from continuous processes with energy below DCUTx” and was not used for this study.

2.1.1 Gaussian (Bohr) Straggling Model. The first model uses a Guassian function to model the energy loss distribution. Recall that Gaussian distributions are defined by two parameters: the standard deviation σ and the mean μ .

According to [20], the first of these is given by:

$$\sigma^2 = 2\pi e^2 N_A T_c \frac{Z\rho L}{A} \frac{1 - \beta^2/2}{\beta}, \quad (2.1)$$

where e is the fundamental electric charge, N_A is Avagadro’s number, T_c is the cut kinetic energy of δ -electrons, Z is the nucleic charge, ρ is the density of the material, L is the length of the material, A is the nucleic mass, and $\beta = v/c$ is the relativistic speed.

The mean comes from the Bethe-Bloch equation [21], which will be derived here and yeilds Eqn. 2.5. The derivation is the so-called classical derivation, which takes first principles and mechanically derives an expression. The classically-derived intermediate equation is then subject to modern corrections, which account for the approximations made or phenomena neglected during the classical derivation.

Classical Derivation

The process begins with a particle of charge e moving through a material. It is assumed that the mass of the incoming particle is much greater than the mass of the electron since there will be no scattering in this derivation. Moreover, the electron must be at rest and fixed into place (which is nonphysical) or the collision time of

the particle and electron must be very small compared to the electron orbital time. The longitudinal (z) axis is aligned with the particle velocity such that $v_x = v_y = 0$. The momentum transfer Δp from the incoming particle to the orbiting electron is sought, and the first observation is that the change in the longitudinal momentum of the particle is zero (i.e. $\Delta p_z = 0$). This is true since the particle feels an average force “forward” just as much as it feels a force “backwards” due to symmetry. Since this model bars scattering, the change in momentum must be interpreted as energy imparted upon the electron. Then following Figure 2.1,

$$\Delta p_x = \int_{-\infty}^{\infty} F_x dt = \int_{-\infty}^{\infty} F_x \frac{dz}{v} = \int_{-\infty}^{\infty} F \cos \theta \frac{dz}{v},$$

where F is the electrostatic force (e^2/r^2) between the particle and the stationary electron, r is the distance between the particle and the electron, and v is the particle’s velocity. Observe from Figure 2.1 that $\cos \theta = b/r$, where b is the impact parameter (the closest distance between the particle and the electron). Then

$$\begin{aligned} \Delta p_x &= \int_{-\infty}^{\infty} \frac{e^2}{r^2} \frac{b}{r} \frac{dz}{v} = \int_{-\infty}^{\infty} \frac{e^2}{z^2 + b^2} \frac{b}{\sqrt{z^2 + b^2}} \frac{dz}{v} \\ &= \frac{e^2 b}{v} \int_{-\infty}^{\infty} \frac{1}{(z^2 + b^2)^{3/2}} dz, \end{aligned}$$

By substitution of the following

$$z = b \tan \theta,$$

$$dz = b \sec^2 \theta d\theta,$$

$$z = \infty \rightarrow \theta = \frac{\pi}{2},$$

$$z = -\infty \rightarrow \theta = \frac{-\pi}{2}.$$

the integral becomes

$$\begin{aligned} \Delta p_x &= \frac{e^2 b}{v} \int_{-\pi/2}^{\pi/2} \frac{b \sec^2 \theta d\theta}{(b^2(\tan^2 \theta + 1))^{3/2}} = \frac{e^2 b}{v} \int_{-\pi/2}^{\pi/2} \frac{d\theta}{b^2 |\sec \theta|} \\ &= \frac{2e^2}{vb}. \end{aligned}$$

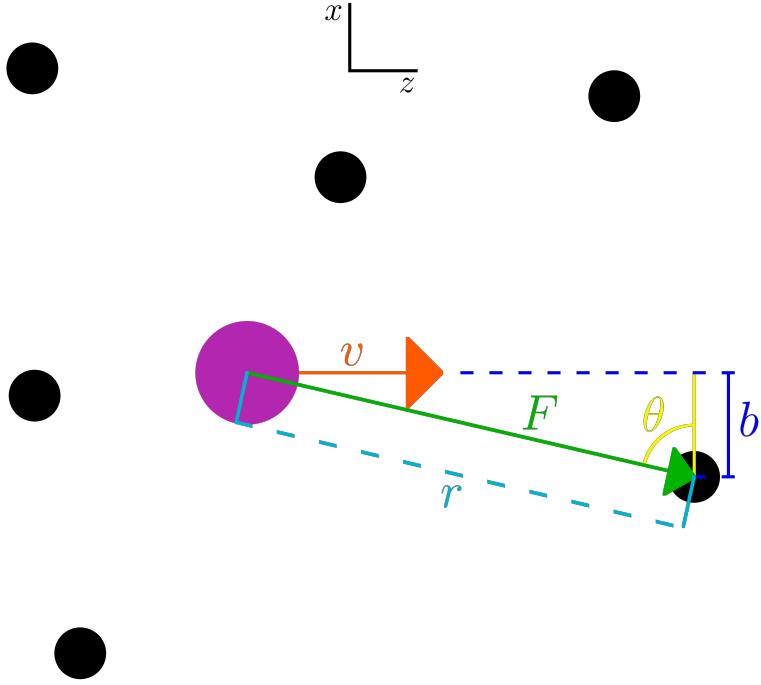


Figure 2.1: Classical model of particle through matter.

The non-relativistic approach yeilds the energy loss for the interaction between a muon and a single electron:

$$\epsilon = \frac{\Delta p^2}{2m_e} = \frac{2e^4}{v^2 b^2 m_e} \quad (2.2)$$

Here, it is useful to note that if the interaction of the muon with the nucleus was desired, ϵ for a single interaction would be proportional to Z^2/m_{nuc} and would be added on to Eqn. 2.2. However, since $1/m_e \gg 1/m_{nuc}$ the second term will be disregarded.

Observe from Figure 2.1 that only a single electron has been considered for Eqn. 2.2. For multiple electrons, the average electron density is $N_{el} = N_A \cdot Z\rho/A$, where N_A is Avagadro's number, Z is the nucleic charge, ρ is the density of the material, and A is the nucleic mass. From this expression the average energy loss may be obtained. The cylindrical symmetry of the system is exploited by integrating the

single electron interaction in cylindrical coordinates, with θ as the cylindrical angle, b (the impact parameter) as radius, and z as length. The bounds are $\theta \in [0, 2\pi]$, $z \in [0, L]$ (where L is the total length of material traversed), and $b \in [b_{min}, b_{max}]$ (which will be discussed shortly). Then

$$\begin{aligned}\langle \epsilon \rangle &= \int_{b_{min}}^{b_{max}} \int_0^L \int_0^{2\pi} \frac{2e^4}{v^2 b^2 m_e} N_{el} d\theta dz b db \\ &= \frac{4\pi N_A e^4 Z \rho L}{v^2 m_e} \frac{A}{A} \int_{b_{min}}^{b_{max}} \frac{db}{b}.\end{aligned}\quad (2.3)$$

It is known that b_{min} comes from the maximum amount of energy (T_{max}) which a muon may impart upon an electron. This quantity is derived later and can be seen in Eqn. 2.8, but is only symbolic here. Similarly, the minimum transferrable energy will be taken as the mean ionization energy, I , which is usually an empirical value. Any energy below this value will not transfer, since this is the minimum amount of energy to free the electron from its orbit. Then

$$\begin{aligned}\epsilon_{max} &= \frac{2e^4}{v^2 b_{min}^2 m_e} = T_{max} \\ b_{min} &= \frac{e^2}{v} \sqrt{\frac{2}{m_e T_{max}}} \\ \epsilon_{min} &= \frac{2e^4}{v^2 b_{max}^2 m_e} = T_{min} = I \\ b_{max} &= \frac{e^2}{v} \sqrt{\frac{2}{m_e I}}.\end{aligned}$$

Once the bounds have been acquired, the integral in Eqn. 2.3 is solvable.

$$\begin{aligned}\langle \epsilon \rangle &= \frac{4\pi N_A e^4 Z \rho L}{v^2 m_e} \frac{A}{A} \int_{b_{min}}^{b_{max}} \frac{db}{b}, \\ \langle \epsilon \rangle &= \frac{4\pi N_A e^4 Z \rho L}{v^2 m_e} \frac{A}{A} \ln \frac{b_{max}}{b_{min}}, \\ \langle \epsilon \rangle &= \frac{2\pi N_A e^4 Z \rho L}{v^2 m_e} \ln \frac{T_{max}}{I}.\end{aligned}\quad (2.4)$$

There are a number of corrections to the Bethe-Bloch equation. The first is a correction under the logarithm due to the relativistic flattening of the incoming particle's electric field. This factor is $2m_e\beta^2\gamma^2/I$. Another correction is known as the density correction δ and arises due to polarization of the material. This is important for high energies since it can limit the range of the flattened electric field in the material. The last correction is the shell correction C and accounts for the fact that the electron is not at rest (or the interaction time is not much faster than the electron orbit). This is important for low energies. These corrections may or may not have a large impact on these medium energy studies, since the logarithmic and density corrections arise at high energies and the shell correction occurs at low energies. Finally, in its more familiar form, the constants in Eqn. 2.4 are condensed into a single constant $K = \frac{2\pi e^4 N_A}{m_e} \approx 15.4 \text{ (MeV} \cdot \text{cm}^3\text{)}/(\text{m} \cdot \text{g})$. Since the model assumes natural units (i.e. $c = 1$), v^2 is usually replaced with β^2 . Moreover, to avoid confusion $\langle\epsilon\rangle$ is given a negative sign in order to emphasize energy *loss*.

The ultimate result is the Bethe-Bloch equation for the average energy lost by a particle traversing some medium with correction factors:

$$\langle\epsilon\rangle = -\frac{K}{\beta^2} \frac{Z\rho L}{A} \left(\ln \frac{2m_e\beta^2\gamma^2 T_{max}}{I^2} - 2\beta^2 - \delta - 2\frac{C}{Z} \right). \quad (2.5)$$

According to the Particle Data Group [9], this equation is generally accurate for intermediate Z materials (up to a few % when compared to experimental data) in the energy regime of $0.1 \lesssim \beta\gamma \lesssim 1000$. The lower limit of this equation is reached when the incoming particle velocity is comparable to the atomic electron velocities and the upper limit is attained due to radiative effects, and both limits exhibit some Z dependence. Figure 1.6 depicts this region with muons on copper. Clearly, the region typically used for muon ionization cooling ($100\text{MeV}/c < p_\mu < 1000\text{MeV}/c$) falls in the middle of the Bethe region.

Derivation of T_{max}

Although T_{max} is used symbolically in many equations, here it will be derived explicitly from a relativistic point of view. This derivation yields Eqn. 2.8 and works in natural units such that $c = 1$. Conservation of energy for a muon incident upon an electron at rest requires that

$$\begin{aligned} E_\mu + m_e &= E_{\mu,f} + E_e, && \text{or} \\ \sqrt{p_\mu^2 + m_\mu^2} + m_e &= \sqrt{p_{\mu,f}^2 + m_\mu^2} + T_e + m_e, && \text{or} \\ p_{\mu,f}^2 &= p_\mu^2 + T_e^2 - 2T_e \sqrt{p_\mu^2 + m_\mu^2} \end{aligned} \quad (2.6)$$

with

$$\begin{aligned} E_e &= T_e + m_e = \sqrt{p_e^2 + m_e^2}, && \text{or} \\ p_e^2 &= (T_e + m_e)^2 - m_e^2 \end{aligned} \quad (2.7)$$

where T_e is the final kinetic energy of the electron, p_μ is the initial muon momentum, m_μ is the mass of the muon, and $p_{\mu,f}$ is the final muon momentum.

Conservation of momentum requires that

$$\begin{aligned} \vec{p}_\mu &= \vec{p}_{\mu,f} + \vec{p}_e, && \text{or} \\ p_{\mu,f}^2 &= p_\mu^2 + p_e^2 - 2p_\mu p_e \cos \theta. \end{aligned}$$

Using Eqn. 2.7 for p_e on the right-hand side this becomes

$$p_{\mu,f}^2 = p_\mu^2 + (T_e + m_e)^2 - m_e^2 - 2p_\mu \cos \theta \sqrt{(T_e + m_e)^2 - m_e^2}.$$

Now substitution of Eqn. 2.6 for $p_{\mu,f}^2$ on the left-hand side yields

$$p_\mu^2 + T_e^2 - 2T_e \sqrt{p_\mu^2 + m_\mu^2} = p_\mu^2 + (T_e + m_e)^2 - m_e^2 - 2p_\mu \cos \theta \sqrt{(T_e + m_e)^2 - m_e^2}.$$

For maximum energy (i.e. to attain $T_e = T_{max}$), $\cos \theta = 1$, which is representative of a head-on collision. Then

$$\begin{aligned} T_{max}^2 - 2T_{max}\sqrt{p_\mu^2 + m_\mu^2} &= T_{max}^2 + 2T_{max}m_e - 2p_\mu\sqrt{T_{max}^2 + 2T_{max}m_e} \\ - 2T_{max}(\sqrt{p_\mu^2 + m_\mu^2} + m_e) &= -2p_\mu\sqrt{T_{max}^2 + 2T_{max}m_e} \\ \sqrt{p_\mu^2 + m_\mu^2} + m_e &= p_\mu\sqrt{1 + \frac{2m_e}{T_{max}}}. \end{aligned}$$

Substituting the initial muon energy $E_\mu = \sqrt{p_\mu^2 + m_\mu^2} = \gamma m_\mu$ and the initial muon momentum $p_\mu^2 = E_\mu^2 - m_\mu^2 = m_\mu^2(\gamma^2 - 1) = m_\mu^2\gamma^2\beta^2$ results in

$$\begin{aligned} \gamma m_\mu + m_e &= m_\mu \gamma \beta \sqrt{1 + \frac{2m_e}{T_{max}}} \\ (\gamma m_\mu + m_e)^2 &= m_\mu^2 \gamma^2 \beta^2 \left(1 + \frac{2m_e}{T_{max}}\right). \end{aligned}$$

Finally, solving for the maximum transferrable energy from an incident muon to an electron at rest is

$$T_{max} = \frac{2m_e\beta^2\gamma^2}{1 + 2\gamma\frac{m_e}{m_\mu} + (\frac{m_e}{m_\mu})^2}. \quad (2.8)$$

2.1.2 Landau Straggling Model.

The second ICOOL straggling model selects an energy loss from a Landau distribution [22]. This derivation is particularly important since it is the same model that was implemented into COSY in this work. Landau begins the derivation by stipulating that this theory assumes fast particles (“so that the usual ionisation theory may be applied”, here taken as particles whose energy is in the Bethe regime in Figure 1.6). Moreover, the thickness of the absorber should be small enough such that the energy loss is small compared to the initial energy. Landau defines the weight function $w(E, u)$ as the probability per unit length of an energy loss u given the instantaneous total energy E . For the aforementioned constraints, the weight function may be written simply as $w(u)$ (since E is now a constant instead of a variable). Another

way of describing this constraint is that the total energy of the particle is roughly constant while traversing the medium.

Let $f(L, \epsilon)$ be desired distribution function for the energy loss. This means that the particle will lose an amount of energy between ϵ and $\epsilon + d\epsilon$ while traversing an absorber of length L . Then on one hand

$$\text{change in } f \text{ per unit length} = \frac{\partial f(L, \epsilon)}{\partial L}.$$

On the other hand, the change in f may also be expressed as the difference of two functions: one at a length L with possible energy losses between ϵ and $\epsilon + d\epsilon$ and another at length L and with possible energy losses between $\epsilon - u$ and $\epsilon + d\epsilon - u$, with u accounting for the infinitesimal change in length. Then

$$\text{change in } f \text{ per unit length} = \left[\int_0^\infty w(u) f(L, \epsilon - u) du \right] - f(L, \epsilon).$$

Often referred to as the integral transport equation, the two previous definitions of the change in f per unit length may be combined as

$$\frac{\partial f(L, \epsilon)}{\partial L} = \left[\int_0^\infty w(u) f(L, \epsilon - u) du \right] - f(L, \epsilon). \quad (2.9)$$

Since L and ϵ are independent and implicit variables, this allows for a Laplace transformation. Take the transformed function with respect to ϵ as

$$\phi(p, L) = \int_0^\infty e^{-p\epsilon} f(\epsilon) d\epsilon.$$

(Note that p is simply a dummy variable and not the momentum.) Then the inverse transformation gives

$$f(L, \epsilon) = \frac{1}{2\pi i} \int_{K-i\infty}^{K+i\infty} \phi(p, L) e^{p\epsilon} dp. \quad (2.10)$$

Observe here that $K > 0$ and so the integral is just to the right of the imaginary axis. This is the desired distribution function, and so will be useful later. However,

without the form of ϕ it is not useful. The goal then will be to find a closed form of ϕ .

Multiplying both sides of Eqn. 2.9 by $e^{-p\epsilon}$ and integrating with respect to $d\epsilon$ yields

$$\int_0^\infty \frac{\partial f}{\partial L} e^{-p\epsilon} d\epsilon = \int_0^\infty \left[\int_0^\infty w(u) f(L, \epsilon - u) du \right] e^{-p\epsilon} d\epsilon - \int_0^\infty f(L, \epsilon) e^{-p\epsilon} d\epsilon.$$

On the left side, the operations of partial derivative and integration are commutable, and are therefore switched. On the right side, the first term has commutable integrations and so the order is switched. For the second term, note that $w(u)$ is normalized, so adding the integral of $w(u)$ over all u changes nothing. Dropping the implicit L for now,

$$\begin{aligned} \frac{\partial}{\partial L} \int_0^\infty e^{-p\epsilon} f(\epsilon) d\epsilon &= \int_0^\infty \left[\int_0^\infty e^{-p\epsilon} f(\epsilon - u) d\epsilon \right] w(u) du - \\ &\quad \int_0^\infty \left[\int_0^\infty e^{-p\epsilon} f(\epsilon) d\epsilon \right] w(u) du. \end{aligned}$$

Parts of the left side and the second term on the right side may be substituted for ϕ directly, while the first term on the right side should be shifted by $-u$, resulting in

$$\frac{\partial}{\partial L} \phi(p, L) = \int_0^\infty \left[\int_{-u}^\infty e^{-p(\epsilon+u)} f(\epsilon) d\epsilon \right] w(u) du - \int_0^\infty \phi(p, L) w(u) du.$$

Now recall that $f(\epsilon)$ is the desired function for energy loss. Therefore, $f(\epsilon < 0) = 0$ (i.e. the particle cannot gain energy while traversing a medium), and so

$$\begin{aligned} \frac{\partial}{\partial L} \phi(p, L) &= \int_0^\infty e^{-pu} \left[\int_0^\infty e^{-p\epsilon} f(\epsilon) d\epsilon \right] w(u) du - \phi(p, L) \int_0^\infty w(u) du \\ &= \phi(p, L) \int_0^\infty w(u) (e^{-pu} - 1) du. \end{aligned}$$

This differential equation is a first-order, undriven normal linear ODE (ordinary differential equation) [23]. Let prime ($'$) denote a partial derivative with respect

to L . Then the strategy to solve this ODE is to define a function $h(L)$ as

$$h(L) = - \int_0^\infty w(u)(e^{-pu} - 1) du$$

and its antiderivative as

$$H(L) = \int h(L)dL = L \int_0^\infty w(u)(1 - e^{-pu}) du.$$

Then

$$\phi' + \phi \cdot h(L) = 0. \quad (2.11)$$

Since $(e^{H(L)})' = e^{H(L)} \cdot H'(L) = e^{H(L)} \cdot h(L)$, it is useful to multiply both sides of the ODE in Eqn. 2.11 by $e^{H(L)}$, resulting in

$$\begin{aligned} \phi' \cdot e^{H(L)} + \phi \cdot h(L) \cdot e^{H(L)} &= 0, \\ (\phi \cdot e^{H(L)})' &= 0. \end{aligned}$$

Integrating both sides yields

$$\begin{aligned} \phi \cdot e^{H(L)} &= K_1, \\ \phi &= K_1 e^{-H(L)}, \\ \phi(p, L) &= K_1 \exp \left[-L \int_0^\infty w(u)(1 - e^{-pu}) du \right]. \end{aligned}$$

This differential equation is now solvable provided that there are initial conditions. The first observation is that for $L = 0$, the only possible energy loss is zero. Mathematically, this means that $f(0, \epsilon) = \delta(\epsilon)$; that is, the probability of energy loss is 100% for an energy loss of zero and 0% for all other energy losses. Then the boundary condition on ϕ is

$$\begin{aligned} \phi(p, 0) &= \int_0^\infty \delta(\epsilon) e^{-p\epsilon} d\epsilon \\ \phi(p, 0) &= e^{p \cdot 0} \\ \phi(p, 0) &= 1 = K_1 \exp[0]. \end{aligned}$$

Then

$$\phi(p, L) = \exp \left[-L \int_0^\infty w(u)(1 - e^{-pu}) du \right].$$

Now using Eqn. 2.10, the energy loss distribution function f in terms of $w(u)$ is

$$f(L, \epsilon) = \frac{1}{2\pi i} \int_{K-i\infty}^{K+i\infty} \exp \left[p\epsilon - L \int_0^\infty w(u)(1 - e^{-pu}) du \right] dp. \quad (2.12)$$

In principle, this is the general solution to the energy loss profile for a particle traversing some medium. In practice, the only thing which is inhibiting implementation is an algorithm to generate a number from this distribution (which will be discussed in Chapter 3) and the function $w(u)$. Once $w(u)$ is obtained, the integral may be found using numeric integration, provided that it is not computationally expensive.

Livingston and Bethe [24] derived the form of $w(u)$ in 1937, and it is

$$w(u) = \frac{\xi}{L} \frac{1}{u^2},$$

where

$$\xi = \frac{2\pi r_e^2 m_e N_A Z \rho L}{\beta^2 A} \quad (2.13)$$

and r_e is the classical electron radius, m_e is the electron mass, N_A is Avagadro's number, Z is the nucleic charge, ρ is the density of the material, L is the length of the material, A is the nucleic mass, and $\beta = v/c$ is the relativistic speed. Now the form of $w(u)(1 - e^{-pu})$ may now be seen in Figure 2.2. It can be seen that the most important values of p are those where $pu \ll 1$ and $1 \ll pu$ since values of $pu \approx 1$ will tend to vanish.

Rather than zero, let u_{min} be the minimum possible energy loss. Now consider a certain energy u_1 such that $u_{min} \ll u_1$ and $pu_1 \ll 1$. Then the integral in the exponent of Eqn. 2.12 has the form

$$L \int_0^\infty w(u)(1 - e^{-pu}) du = \xi \left(\int_0^{u_1} \frac{1 - e^{-pu}}{u^2} du + \int_{u_1}^\infty \frac{1 - e^{-pu}}{u^2} du \right).$$

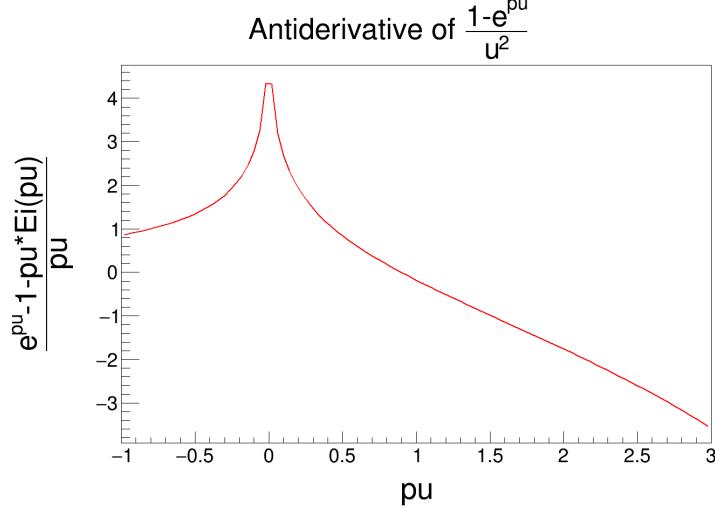


Figure 2.2: Plot of $\int w(u)(1 - e^{-pu}) du$ for various pu . On the vertical axis is the antiderivative of $w(u)(1 - e^{-pu})$, with $Ei(x)$ representing the exponential integral.

Since the first integral is over the small values, it is possible to write $1 - e^{pu} \approx 1 - (1 - pu) = pu$. Furthermore, the minimum energy loss is not 0 but rather u_{min} . Then

$$\begin{aligned} \frac{L}{\xi} \int_0^\infty w(u)(1 - e^{-pu}) du &= \int_{u_{min}}^{u_1} \frac{p u}{u^2} du + \int_{u_1}^\infty \frac{1 - e^{-pu}}{u^2} du \\ &= p \int_{u_{min}}^{u_1} \frac{1}{u} du + \int_{u_1}^\infty \frac{1 - e^{-pu}}{u^2} du \\ &= p \ln \frac{u_1}{u_{min}} + \int_{u_1}^\infty \frac{1 - e^{-pu}}{u^2} du. \end{aligned}$$

For the second term, integration by parts and evaluation at the boundaries gives

$$\frac{L}{\xi} \int_0^\infty w(u)(1 - e^{-pu}) du = p \ln \frac{u_1}{u_{min}} + \frac{1 - e^{-pu_1}}{u_1} + p \int_{u_1}^\infty \frac{e^{-pu}}{u} du.$$

Recalling that u_1 was chosen such that $pu_1 \ll 1$, the exponential in the second term may be approximated as $e^{-pu_1} \approx 1 - pu_1$. Then

$$\frac{L}{\xi} \int_0^\infty w(u)(1 - e^{-pu}) du = p \ln \frac{u_1}{u_{min}} + p + p \int_{u_1}^\infty \frac{e^{-pu}}{u} du. \quad (2.14)$$

The final integral may be evaluated first by letting $K_2 = pu$:

$$\begin{aligned}
\frac{L}{\xi} \int_{u_1}^{\infty} \frac{e^{-pu}}{u} du &= \int_{pu_1}^{\infty} \frac{e^{-K_2}}{K_2} dK_2 \\
&= \int_{pu_1}^{\infty} \frac{e^{-K_2}}{K_2} dK_2 + \int_{pu_1}^1 \frac{dK_2}{K_2} - \int_{pu_1}^1 \frac{dK_2}{K_2} \\
&= \int_{pu_1}^1 \frac{e^{-K_2}}{K_2} dK_2 + \int_1^{\infty} \frac{e^{-K_2}}{K_2} dK_2 + \int_{pu_1}^1 \frac{dK_2}{K_2} - \int_{pu_1}^1 \frac{dK_2}{K_2} \\
&= \int_{pu_1}^1 \left(\frac{e^{-K_2}}{K_2} - \frac{1}{K_2} \right) dK_2 + \int_1^{\infty} \frac{e^{-K_2}}{K_2} dK_2 + \int_{pu_1}^1 \frac{dK_2}{K_2} \\
&= \int_{pu_1}^1 \frac{dK_2}{K_2} + \left[\int_{pu_1}^1 \frac{e^{-K_2} - 1}{K_2} dK_2 + \int_1^{\infty} \frac{e^{-K_2}}{K_2} dK_2 \right].
\end{aligned}$$

The first term is easily evaluated, and the second term in brackets is approximately¹ the negative of Euler's constant, $-C_{Euler}$. Putting this back into Eqn. 2.14 yields

$$\begin{aligned}
\frac{L}{\xi} \int_0^{\infty} w(u)(1 - e^{-pu}) du &= p \ln \frac{u_1}{u_{min}} + p(1 - C_{Euler} - \ln pu_1) \\
&= p(1 - C_{Euler} - \ln pu_{min}).
\end{aligned}$$

Now Eqn. 2.12 may be rewritten as

$$\begin{aligned}
f(L, \epsilon) &= \frac{1}{2\pi i} \int_{K-i\infty}^{K+i\infty} \exp \left[p\epsilon - L \int_0^{\infty} w(u)(1 - e^{-pu}) du \right] dp \\
&= \frac{1}{2\pi i} \int_{K-i\infty}^{K+i\infty} \exp \left[p\epsilon - p\xi(1 - C_{Euler} - \ln pu_{min}) \right] dp.
\end{aligned}$$

If a unitless variable of integration is desired, then $p \rightarrow p/\xi$, and

$$f(\xi, \epsilon) = \frac{1}{2\pi i \xi} \int_{K-i\infty}^{K+i\infty} \exp \left[p \frac{\epsilon}{\xi} - p(1 - C_{Euler} - \ln \frac{u_{min}}{\xi}) + p \ln p \right] dp,$$

or simply

$$f(\lambda) = \frac{1}{2\pi i \xi} \int_{K-i\infty}^{K+i\infty} \exp \left[p \ln p + \lambda p \right] dp, \quad (2.15)$$

where evidently

$$\lambda = \frac{\epsilon}{\xi} - 1 + C_{Euler} + \ln \frac{\xi}{u_{min}}.$$

¹This expression would be exactly $-C_{Euler}$ if $pu_1 = 0$.

However, typically the Landau function is evaluated for the desired *fluctuation about the mean energy loss*, not the energy loss itself. Because of this, λ must be shifted by the mean, $\langle \epsilon \rangle$. Furthermore, a relativistic correction of $-\beta^2$ is added and $u_{min} \rightarrow T_{max}$ (from Eqn. 2.8), resulting in the final form of the Landau parameter:

$$\lambda \equiv \frac{\epsilon - \langle \epsilon \rangle}{\xi} - (1 - C_{Euler}) - \beta^2 - \ln(\xi/T_{max}). \quad (2.16)$$

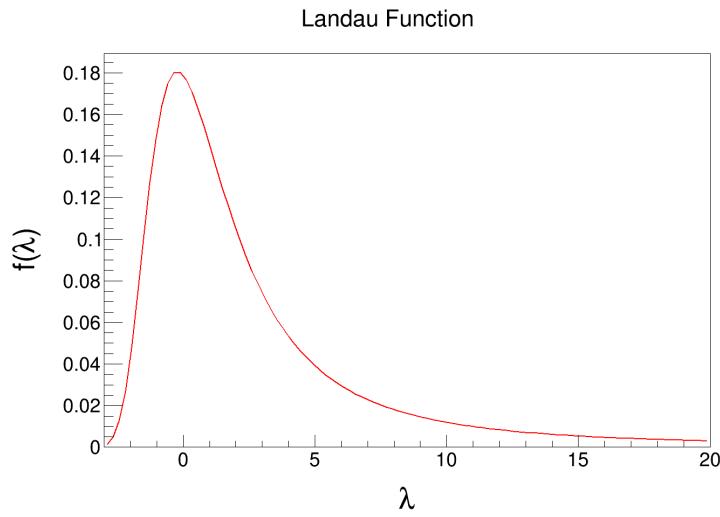


Figure 2.3: Example of the Landau function.

In summary, the ICOOL Landau model uses the mean energy loss from the Bethe-Bloch equation (Eqn. 2.5). The routine then adds noise based on the Landau function (see, e.g., Figure 2.3), which has a mode of zero. This value is then accepted as the total energy loss of the particle.

2.1.3 Vavilov Straggling Model.

The Vavilov model [25] is similar to the Landau model, except that the thickness of the absorber is not required to be small enough such that the energy loss is small compared to the initial energy. The derivation is also similar, except that Vavilov introduced a limit on the maximum transferable energy due to a single collision

(recall that in Landau theory the upper limit was ∞). The result is

$$f(\lambda_v, \kappa, \beta^2) = \frac{1}{\xi} \phi_v(\lambda_v, \kappa, \beta^2), \quad (2.17)$$

where

$$\begin{aligned} \phi_v(\lambda_v, \kappa, \beta^2) &= \frac{1}{2\pi i} \int_{K+i\infty}^{K-i\infty} \phi(p, \kappa, \beta^2) e^{\lambda p}, \\ \phi(p, \kappa, \beta^2) &= \exp[\kappa(1 + \beta^2 \gamma)] \exp[\psi(p, \kappa, \beta^2)], \\ \psi(p) &= p \ln \kappa + (p + \beta^2 \kappa)[\ln(p/\kappa) + E_1(p/\kappa)] - \kappa e^{-p/\kappa}, \\ E_1(p) &= \int_{\infty}^z \frac{e^{-u}}{u} du \quad (\text{the exponential integral}), \\ \lambda_v &= \kappa \left(\frac{\epsilon - \langle \epsilon \rangle}{\xi} - (1 - C_{Euler}) - \beta^2 \right) = \kappa(\lambda + \ln \kappa), \text{ and} \\ \kappa &= \xi/T_{max}. \end{aligned}$$

While this function in theory is universal for the needs of this study, clearly such a distribution function requires numeric methods, and as such is more costly than the Landau method. Fortunately, the Vavilov distribution converges to either the Landau distribution (for $\kappa \rightarrow 0$) or a Gaussian distribution (for $\kappa \rightarrow \infty$) at its extrema. The cutoffs for these extrema in practice are

$$\phi_v = \begin{cases} \text{Landau via Eqn. 2.15} & \kappa \leq 0.01 \\ \text{Vavilov via Eqn. 2.17} & 0.01 < \kappa < 10 \\ \text{Gaussian via } Gaus(\langle \epsilon \rangle, \sigma) \text{ in Eqns. 2.5 and 2.1} & 10 \leq \kappa \end{cases} .$$

2.2 Multiple Scattering in ICOOL

ICOOL version 3.30 [18] boasts a total of seven models of multiple scattering:

1. Gaussian (σ determined by Rossi-Greisen model),

2. Gaussian (σ determined by Highland model),
3. Gaussian (σ determined by Lynch-Dahl model),
4. Bethe version of Molière distribution with Rutherford limit,
5. Rutherford,
6. Fano with Rutherford limit, and
7. Tollestrup with Rutherford limit.

The scattering model used in this work is Fano with Rutherford limit, which is also the default scattering model in ICOOL. The following is a derivation of the Rutherford model accompanied by a brief discussion of the Fano model. Unless otherwise stated, the derivation of the Rutherford model closely follows [12].

The Rutherford model is used in part in four out of the seven models because it is so robust. The derivation can be done using the Coulomb potential, classical mechanics, the Born approximation, and quantum field theory, with the quantum mechanics Born approximation being the weapon of choice here. Recall that the quantum model used in this work is represented by Figure 2.4. Here there is an incoming plane wave (mathematically represented by e^{ikz}) and a spherical wave (represented by e^{ikr}/r). r and z are coordinates, i is the imaginary unit, and k is the wave number, classically related to the energy as $k = \sqrt{2mE}/\hbar$. Then basic quantum mechanics suggests that the solution to the Schrödinger equation has the form

$$\psi(r, \theta) \approx A \left(e^{ikz} + f(\theta) \frac{e^{ikr}}{r} \right), \quad (2.18)$$

where A is the total amplitude and $f(\theta)$ is the scattering amplitude. The probability of the particle scattering in a particular direction is given by the amplitude squared,

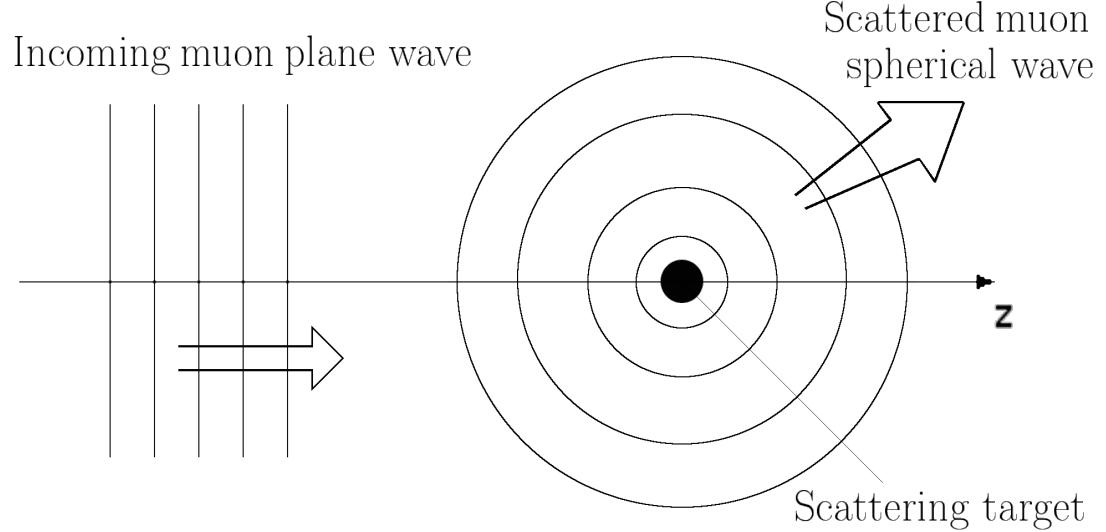


Figure 2.4: Quantum scattering model.

$|f(\theta)|^2$, and is the object of this derivation. Now, ψ solves the differential (time-independent) Schrödinger equation, which usually has the form

$$-\frac{\hbar^2}{2m}\nabla^2\psi + V\psi = E\psi,$$

where V is the system potential and E is the energy of the wavefunction. This can be expressed alternatively by assigning $Q \equiv V\psi \cdot 2m/\hbar^2$ and recalling that $k = \sqrt{2mE}/\hbar$.

Then

$$(\nabla^2 + k^2)\psi = Q. \quad (2.19)$$

The strategy is to solve Eqn. 2.19 for ψ . A comparison of the solutions for ψ from Eqns. 2.18 and 2.19 will yield the scattering amplitude $f(\theta)$. However, if Eqn. 2.19 is solved for ψ then it suggests that ψ will be in integral form since Eqn. 2.19 is a differential equation. Therefore, some special techniques will be required. Note that it does not matter at this point that Q is a function of ψ , since the task will be to compare the integrated solution of Eqn. 2.19 with that of Eqn. 2.18. Moving

forward, observe that Eqn. 2.19 may again be rewritten as

$$(\nabla^2 + k^2)\psi(\vec{r}) = Q(\vec{r}) = \int \delta^3(\vec{r} - \vec{r}_0)Q(\vec{r}_0)d^3\vec{r}_0. \quad (2.20)$$

Now it is natural to guess that there exists some function $G(\vec{r})$ such that

$$\psi(\vec{r}) = \int G(\vec{r} - \vec{r}_0)Q(\vec{r}_0)d^3\vec{r}_0, \quad (2.21)$$

in which case

$$(\nabla^2 + k^2)\psi(\vec{r}) = (\nabla^2 + k^2) \int G(\vec{r} - \vec{r}_0)Q(\vec{r}_0)d^3\vec{r}_0,$$

or

$$(\nabla^2 + k^2)\psi(\vec{r}) = \int [(\nabla^2 + k^2)G(\vec{r} - \vec{r}_0)]Q(\vec{r}_0)d^3\vec{r}_0. \quad (2.22)$$

Combining Eqns. 2.20 and 2.22, it can be seen that

$$\int \delta^3(\vec{r} - \vec{r}_0)Q(\vec{r}_0)d^3\vec{r}_0 = \int [(\nabla^2 + k^2)G(\vec{r} - \vec{r}_0)]Q(\vec{r}_0)d^3\vec{r}_0.$$

Consequently, one comes to the conclusion that

$$(\nabla^2 + k^2)G(\vec{r}) = \delta^3(\vec{r}). \quad (2.23)$$

If Eqn. 2.23 seems familiar, it is because this is the Helmholtz equation with a delta function source. $G(\vec{r})$ is Green's function for the Helmholtz equation, and in this case it is the response to the delta function source. Now, if one accepts that there exists a well-known particular solution for Eqn. 2.23, then one could skip forward to the solution in Eqn. 2.26; however, it will also be derived subsequently.

The usual strategy in solving systems like Eqn. 2.23 is to Fourier transform both the Green's function and the delta function. Then creating the dummy variable \vec{s} , the transform yields

$$G(\vec{r}) = \frac{1}{(2\pi)^{\frac{3}{2}}} \int e^{i\vec{s}\cdot\vec{r}} g(\vec{s})d^3\vec{s},$$

and

$$\delta^3(\vec{r}) = \frac{1}{(2\pi)^3} \int e^{i\vec{s}\cdot\vec{r}} d^3\vec{s}.$$

Then from Eqn. 2.23,

$$\frac{1}{(2\pi)^{\frac{3}{2}}} \int (k^2 e^{i\vec{s}\cdot\vec{r}} - s^2 e^{i\vec{s}\cdot\vec{r}}) g(\vec{s}) d^3\vec{s} = \frac{1}{(2\pi)^3} \int e^{i\vec{s}\cdot\vec{r}} d^3\vec{s},$$

it is clear that $g(\vec{s}) = 1/(2\pi)^{3/2}(k^2 - s^2)$. Now all that is left is to find $G(\vec{r})$ from its transformation:

$$G(\vec{r}) = \frac{1}{(2\pi)^3} \int e^{i\vec{s}\cdot\vec{r}} \frac{d^3\vec{s}}{k^2 - s^2}$$

$$G(\vec{r}) = \frac{1}{(2\pi)^3} \int_0^\infty \frac{s}{k^2 - s^2} \left(\int_0^\pi e^{isr \cos \theta} \sin \theta d\theta \right) ds \int_0^{2\pi} d\phi.$$

The integration over ϕ is trivial, and the integration over θ can be done via u substitution, with $u = \cos \theta$. The last integral is over s , and is

$$G(\vec{r}) = \frac{1}{2\pi^2 r} \int_0^\infty \frac{s \sin sr}{k^2 - s^2} ds = \frac{1}{4\pi^2 r} \int_{-\infty}^\infty \frac{s \sin sr}{k^2 - s^2} ds.$$

This integral is not simple to solve, but it does have two poles at $s = k$ and $s = -k$, which implies the technique of choice should be to use Cauchy's integral formula for simple poles²:

$$\oint \frac{f(z)}{z - z_0} dz = 2\pi i f(z_0), \quad (2.24)$$

where the integral is done over some path in the complex plane and z_0 is the pole of interest which lies in the enclosed path (note: the integral is zero if there exist no poles in the enclosed path). It follows then that $f(z)$ is not simply any function, but a necessarily complex function which is closed. For this reason, the (strictly real) Green's function integral should be split up into two (strictly complex) functions, as depicted by Figure 2.5. This can be done by expanding the $\sin sr$ term and factoring $k^2 - s^2$:

$$G(\vec{r}) = \frac{i}{8\pi^2 r} \left[\int_{-\infty}^\infty \frac{se^{isr}}{(s - k)(s + k)} ds - \int_{-\infty}^\infty \frac{se^{-isr}}{(s - k)(s + k)} ds \right]. \quad (2.25)$$

²This touches an area of mathematics known as the calculus of residues, and is based on the Laurent series.

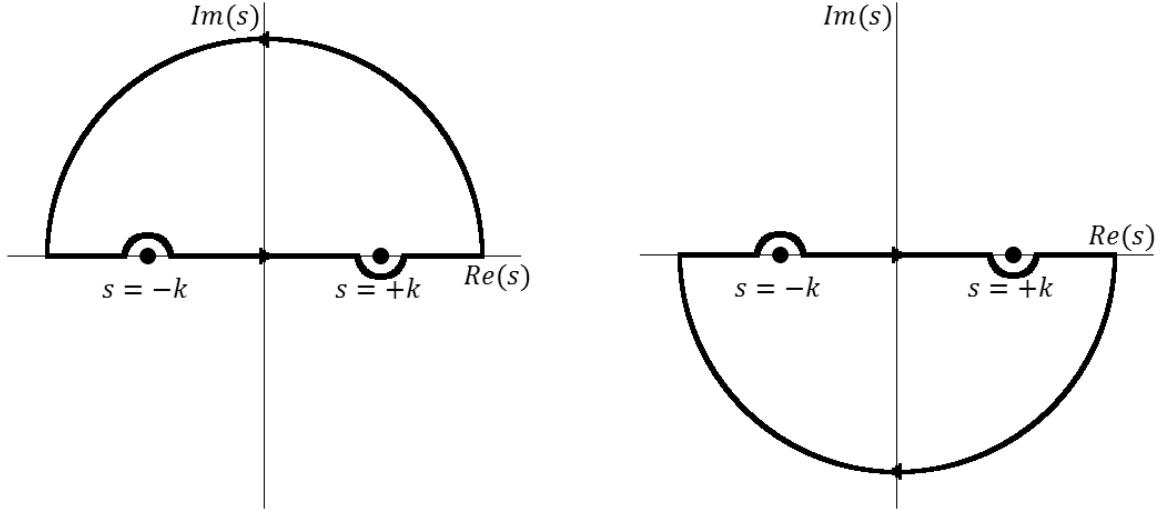


Figure 2.5: Two parts of Green's function with their poles at $s = \pm k$, altered to be closed by a semicircle at $|s| = \pm \infty$.

Observe that in Eqn. 2.25 the path integrals at $|s| = \pm \infty$ have been left out. This is because they do not contribute to the integral, since the first integrand corresponds to the left side of Figure 2.5 and goes like e^{isr} , hence going to zero at large positive imaginary numbers. Similarly, the second integrand goes like e^{-isr} and goes to zero at large negative imaginary numbers.

Combining Eqns. 2.24 and 2.25, it can be seen that

$$G(\vec{r}) = \frac{i}{8\pi^2 r} [(i\pi e^{ikr}) - (-i\pi e^{ikr})] = -\frac{e^{ikr}}{4\pi r}.$$

This is a particular solution to the inhomogeneous Helmholtz equation. To get a general solution, the solution to the homogeneous Helmholtz equation must be added:

$$G(\vec{r}) = G_0(\vec{r}) - \frac{e^{ikr}}{4\pi r}; \quad (2.26)$$

that is, $G_0(\vec{r})$ solves $(\nabla^2 + k^2)G_0(\vec{r}) = 0$.

Using Eqn. 2.21 in conjunction with Eqn. 2.26 gives rise to the *integrated* time-independent Schrödinger equation:

$$\psi(\vec{r}) = \psi_0(\vec{r}) - \frac{m}{2\pi\hbar^2} \int \frac{e^{ik|\vec{r}-\vec{r}_0|}}{|\vec{r}-\vec{r}_0|} V(\vec{r}_0) \psi(\vec{r}_0) d^3\vec{r}_0. \quad (2.27)$$

That is, nothing has been said about the potential and so this equation is quite general. Eqn. 2.27 gives the recursive form for ψ . Let $g = -me^{ik|\vec{r}-\vec{r}_0|}/2\pi\hbar^2|\vec{r}-\vec{r}_0|$. Then Eqn. 2.27 says that

$$\psi = \psi_0 + \int gV\psi = \psi_0 + \int gV(\psi_0 + \int gV\psi).$$

Applying this recursion several times gives the Born series:

$$\psi = \psi_0 + \int gV\psi_0 + \int \int gVgV\psi_0 + \int \int \int gVgVgV\psi_0 + \dots$$

The zeroth term is exact if there is no scattering whatsoever (i.e. $V = 0$) and the first term is accurate if the scattering potential is “weak” (that is, if V is small enough for the second term to be neglected). For the purposes of this derivation, first order will be a sufficient approximation for ψ . Then

$$\psi = \psi_0 + \frac{m}{2\pi\hbar^2} \int \frac{e^{ik|\vec{r}-\vec{r}_0|}}{|\vec{r}-\vec{r}_0|} V(\vec{r}_0) \psi_0(\vec{r}_0) d^3\vec{r}_0$$

Recall that the strategy was to compare this solution of the Helmholtz equation with the wavefunction from the quantum scattering model (Eqn. 2.18) in order to find the scattering amplitude, $f(\theta)$. Doing so yields

$$f(\theta) = -\frac{r}{e^{ikr}} \frac{m}{2\pi\hbar^2 A} \int \frac{e^{ik|\vec{r}-\vec{r}_0|}}{|\vec{r}-\vec{r}_0|} V(\vec{r}_0) \psi_0(\vec{r}_0) d^3\vec{r}_0.$$

The Coulomb potential goes as $1/r^2$, and as such $V(\vec{r}_0)$ is localized about $\vec{r}_0 = 0$. Since particles are observed far away from their scattering centers, it is advantageous to make use of the fact that $\vec{r} \gg \vec{r}_0$. However, caution must be taken,

since one must evaluate the exponential and non-exponential terms separately, for they are of separate orders. For the non-exponential terms, this is simple, since

$$\frac{r}{|\vec{r} - \vec{r}_0|} \approx 1.$$

The exponential terms look like

$$e^{ik(|\vec{r} - \vec{r}_0| - r)},$$

and so it is useful to expand the absolute value as

$$|\vec{r} - \vec{r}_0|^2 = r^2 + r_0^2 - 2\vec{r} \cdot \vec{r}_0 \approx r^2 \left(1 - 2\frac{\vec{r} \cdot \vec{r}_0}{r^2}\right),$$

or simply

$$|\vec{r} - \vec{r}_0| \approx r - \hat{r} \cdot \vec{r}_0.$$

This leaves

$$f(\theta) = \frac{m}{2\pi\hbar^2} \int e^{-ik\hat{r} \cdot \vec{r}_0} V(\vec{r}_0) e^{ik\hat{z} \cdot \vec{r}_0} d^3 \vec{r}_0, \quad (2.28)$$

where $\psi_0(\vec{r}_0) = Ae^{ik\hat{z} \cdot \vec{r}_0}$. Now define $\vec{\kappa} \equiv k(\hat{z} - \hat{r})$ such that $\kappa = 2k \sin \theta/2$. The exponential term becomes

$$e^{i\vec{\kappa} \cdot \vec{r}_0} = e^{i\kappa r_0 \cos \theta_0}.$$

Furthermore, the form of the potential V is known. Since the scattering for the low Z target happens at low temperatures, it is possible to use the Fermi-Thomas approximation for screening [26]. This modifies the Maxwell equation

$$\nabla^2 V(r) = -\frac{\rho}{\epsilon_0}$$

to

$$(\nabla^2 - b^2)V(r) = -\frac{Q}{\epsilon_0}\delta(r),$$

where ρ is the charge density, b is some screening constant, ϵ_0 is the permittivity of free space, and Q is the charge of the potential (in this case, the atomic number Z). The solution is

$$V(r) = \frac{Q}{4\pi\epsilon_0} \frac{e^{-br}}{r},$$

Then the quantum scattering equation (Eqn. 2.28) becomes

$$f(\theta) \propto \int e^{i\kappa r_0 \cos(\theta_0)} e^{-br_0} r_0 \sin(\theta_0) dr_0 d\theta_0 d\phi_0,$$

where the amplitude terms have been left out since $|f(\theta)|^2$ is normalized anyway. Here it is seen that there does exist some θ dependence in $f(\theta)$ (by virtue of κ). The integral over ϕ_0 is trivial, and the integral over θ_0 can be done via u substitution, with $u = \cos \theta_0$. This leaves

$$f(\theta) \propto \frac{1}{\kappa} \int_0^\infty e^{-br_0} \sin(\kappa r_0) dr_0,$$

which has the solution

$$f(\theta) \propto \frac{1}{b^2 + \kappa^2}.$$

For classical Rutherford scattering, $b \rightarrow 0$. Recalling that $\kappa = 2k \sin \theta / 2$ yields the final Rutherford scattering distribution:

$$\begin{aligned} |f(\theta)|^2 &\propto \frac{1}{\sin^4 \frac{\theta}{2}} = \frac{1}{(\sin^2 \frac{\theta}{2})^2} = \frac{1}{\left(\frac{1}{2} [1 - \cos^2(2 \cdot \frac{\theta}{2})]\right)^2}, \\ |f(\theta)|^2 &\propto \frac{1}{(1 - \cos^2 \theta)^2}. \end{aligned} \tag{2.29}$$

Only the tail of Eqn. 2.29 should be used since it blows up at $\theta = 0$. This may seem like a contradiction, since the Born series was truncated at the weak potential term (the first nontrivial term), and hence the incoming muons should not scatter much at all! In fact, the muons are not scattering much at all for each individual atom they encounter. Over the entire absorber, however, some muons will have a net scattering angle which is small (and therefore should not be approximated by the Rutherford tail) whereas others will have a large net scattering angle (and are well represented by Eqn. 2.29). Many weak potentials can still induce a relatively large scattering angle.

2.3 Energy Straggling in G4Beamline

G4Beamline uses the in-house straggling model described by GEANT4 in [20]. GEANT4 has determined that a “thick absorber” occurs when the following conditions are met:

$$\begin{aligned} |\delta E| &> \kappa T_c \\ T_{max} &\leq 2T_c, \end{aligned} \tag{2.30}$$

where δE is the energy loss, κ is the Vavilov limit parameter, T_c is the kinetic energy cut of δ -electrons, and T_{max} is the maximum transferrable kinetic energy between the incoming particle and target electron. These symbols can also be found in the Definition of Terms at the beginning of this document.

For thick absorbers, the energy loss is sampled according to a simple Gaussian distribution with average equal to the Bethe-Bloch energy loss (Eqn. 2.5) and a standard deviation according to Bohr’s variance (Eqn. 2.1).

If these conditions are not met, a “thin absorber” algorithm is called. Here, atoms are assumed to have only two energy levels with binding energies E_1 and E_2 . The interacting muon can then lose energy via excitation, yielding an energy loss of E_1 or E_2 , or lose energy via δ ray production, yielding an energy loss according to $g(E) \propto 1/E^2$, or (more likely) some combination or weighted average of the two. $g(E)$ may then be normalized:

$$\int_{E_0}^{T_{up}} g(E)dE = 1 \rightarrow g(E) = \frac{E_0 T_{up}}{T_{up} - E_0} \frac{1}{E^2}, \tag{2.31}$$

where E_0 is the ionization energy of the atom in question and T_{up} is some kinetic energy cutoff (either the production threshold for delta rays or the maximum transferrable energy, whichever is smaller).

The probability for obtaining any one of these energy losses is given by the macroscopic cross section, Σ_i , where $i = 1, 2, 3$. For excitation ($i = 1, 2$), the cross

section has a form similar to the deterministic Bethe-Bloch equation (Eqn. 2.5):

$$\Sigma_i = C \frac{f_i}{E_i} \frac{\ln(2m_e c^2 (\beta\gamma)^2 / E_i)}{\ln(2m_e c^2 (\beta\gamma)^2 / I)} (1 - r). \quad (2.32)$$

Here, C and $r = 0.55$ are model parameters with r describing the relative contribution of excitation to ionization, f_i are the relative oscillator strengths of the energy levels of E_i , I is the average ionization energy, and the other symbols have their usual meaning. For continuous energy loss, the cross section is given by

$$\Sigma_3 = C \frac{T_{up} - E_0}{T_{up} E_0 \ln(T_{up}/E_0)} r. \quad (2.33)$$

The oscillator strengths are relative to one another and hence should satisfy

$$f_1 + f_2 = 1. \quad (2.34)$$

The next constraint comes from [27] and states that all the energy levels should be weighted and add logarithmically to the total ionization energy:

$$f_1 \ln E_1 + f_2 \ln E_2 = \ln I. \quad (2.35)$$

Moreover, f_1 and f_2 can be thought of as representing the relative number of loosely and tightly bound electrons. Then using the first constraint, it is easy to see that the absolute number of loosely bound electrons are $Z \cdot f_1$ and the absolute number of tightly bound electrons are $Z \cdot f_2$ (since $Z \cdot f_1 + Z \cdot f_2 = Z$). For modeling purposes, GEANT4 has placed empirical initial conditions on these parameters:

$$f_2 = 0 \quad \text{for } Z = 1, \quad (2.36)$$

$$f_2 = 2/Z \quad \text{for } Z \geq 2, \quad (2.37)$$

$$E_2 = 10 \text{ eV } Z^2, \quad (2.38)$$

$$E_0 = 10 \text{ eV}. \quad (2.39)$$

From these, f_1 and E_1 can be found from Eqns. 2.34 and 2.35 given a particular muon through a particular material (from which Z , I , β , and the like are taken).

Finally, an energy loss for a thin absorber can be sampled. For the contribution due to excitation, two numbers n_1 and n_2 are sampled randomly from a Poisson distribution. These numbers represent the relative contributions of the energy levels of E_1 and E_2 , respectively:

$$\epsilon_{exc} = n_1 E_1 + n_2 E_2.$$

The contribution due to ionization can be found by inverting the cumulative distribution function of $g(E)$ (see Eqn. 2.31):

$$G(E) = \int_{E_0}^E g(E)dE \rightarrow E = \frac{E_0}{1 - u \frac{T_{up} - E_0}{T_{up}}},$$

where u is uniformly randomly selected from $[0, 1]$. However, this treatment so far has only been executed for a single ionization. For an absorber of length L , the number of ionizations n_3 is again sampled from a Poisson distribution. Then the total energy loss for thin absorbers is

$$\epsilon = n_1 E_1 + n_2 E_2 + \sum_{j=1}^{n_3} \frac{E_0}{1 - u_j \frac{T_{up} - E_0}{T_{up}}}.$$
 (2.40)

It should be noted that [20] does make a brief mention of a width correction algorithm. This algorithm allegedly decreases the dependence of the results on kinetic energy cuts and step sizes and works for any thickness of material. However, the section in the manual is less than a page long and purely conceptual without any mathematics or data on which to elaborate. The width correction algorithm is relevant to this work in order to make a fair comparison and understand why this work and G4Beamline disagree in some places.

Finally, the GEANT4 straggling routine can be summarized as such:

1. Determine if absorber is “thick” via Eqn. 2.30.
2. If absorber is thick, use Gaussian distribution with mean $\mu = \langle dE/dx \rangle$ (from Eqn. 2.5) and standard deviation from Eqn. 2.1.

3. If absorber is “thin”, use Eqn. 2.40.

- Select n_1 , n_2 , and n_3 from a Poisson distribution.
- Select u_j from a uniform distribution on $[0, 1]$ (where $j = 1 \dots n_3$).
- Find E_2 , f_2 , and E_0 from Eqn. 2.36.
- Find E_1 from Eqns. 2.34 and 2.35.

4. Apply width correction algorithm.

2.4 Multiple Scattering in G4Beamline

The G4Beamline scattering model in [19] uses the GEANT4 Urbán model [20], and again parameterizes according to experimental data and Lewis theory [28]. For this section, the scattering distribution is $g(u)$, where $u = \cos \theta$ and θ is the scattering angle.

Based on the models available, it can be inferred that the function responsible for the sampling of the angular distribution, $g(u)$, is based off both the Goudsmit-Saunderson treatment of scattering [29] and Rutherford scattering. One fundamental result from Goudsmit and Saunderson is that for small angles, the scattering distribution is Gaussian. Recall that Rutherford scattering was derived in Section 2.2 and resulted in Eqn. 2.29. It should be noted that the shape of $g(u)$ was chosen empirically, and is

$$g(u) = q_g [p_g g_1(u) + (1 - p_g) g_2(u)] + (1 - q_g) g_3(u), \quad (2.41)$$

where $0 \leq p_g, q_g \leq 1$ and

$$\begin{aligned} g_1(u) &= C_1 e^{-a(1-u)} & -1 \leq u_0 \leq u \leq 1 \\ g_2(u) &= C_2 \frac{1}{(b_r - u)^{d_r}} & -1 \leq u \leq u_0 \leq 1 \\ g_3(u) &= C_3 & -1 \leq u \leq 1 \end{aligned}$$

are normalized over $[-1, 1]$, where C_i are normalization constants and a , b_r , d_r , and u_0 are empirical parameters. All of these parameters will be discussed in this section, and can be found in Table 2.1.

Observe that for small angles, $g(u)$ is nearly Gaussian since $\exp(1 - u) = \exp(1 - \cos \theta) \approx \exp(\theta^2/2)$. For large angles, $g(u)$ resembles the Rutherford dependence of Eqn. 2.29 for $b_r \approx 1$ and d_r close to 2. Moreover, at small q_g the shape of $g(u)$ is nearly constant since the constant function $g_3(u)$ dominates.

While Eqn. 2.41 is the main point of this section, it would be incomplete without detailing the nine parameters mentioned. The C_i (with $i = 1, 2, 3$) are normalization constants. a , u_0 , and d_r are chosen based off theoretical and experimental data and are a bit more interesting. Finally, p_g , q_g , and b_r can be found using constraints.

Finding a

Observe from Eqn. 2.41 that a only appears in $g_1(u)$. Further observe from Eqn. 2.41 that $g_1(u)$ is only valid for u close to 1. Since $u = \cos \theta$ (where θ is the scattering angle), $g_1(u)$ must be the part of $g(u)$ that determines small angle scattering.

It has already been noted that $g_1(u)$ is approximately Gaussian for u close to 1. Highland [8] provided an estimate for the width of the (approximate) Gaussian scattering distribution. Let the width of this Gaussian distribution be θ_0 . Lynch and Dahl [30] gave corrections to Highland's form of θ_0 in 1991. If θ_0 truly is the width of the Gaussian scattering distribution, then it follows that

$$g_1(u) \underset{\sim}{\propto} \exp\left(-\frac{\theta^2}{2\theta_0^2}\right) \approx \exp\left(\frac{1}{2} \cdot \frac{1 - \cos \theta}{1 - \cos \theta_0}\right),$$

G4Beamline Scattering Distribution Parameters

Parameter	Physical Meaning	Found Via	Eqn.
$C_1, C_2,$ and C_3	Normalization constants	Normalization	N/A
a	Related to Gaussian-like σ	Relating the Gaussian-like behavior of $g_1(u)$ to Highland-like theory [8]	2.42
u_0	The boundary between the Gaussian-like $g_1(u)$ and the Rutherford-like $g_2(u)$	Emperical parameterization	2.44
d_r	The Rutherford-like exponent in $g_2(u)$	Emperical parameterization	2.45
p_g	Relative contribution of the Gaussian-like $g_1(u)$ to the Rutherford-like $g_2(u)$	Demanding continuity	2.48
b_r	Relative u offset of the Rutherford-like $g_2(u)$	Demanding smoothness	2.49
q_g	The relative contribution of the varying functions $g_1(u)$ and $g_2(u)$ to the constant function $g_3(u)$	Demanding that g gives the same mean value as Lewis theory	2.50

Table 2.1: The nine parameters of the scattering distribution used by G4Beamline (see Eqn. 2.41).

where $\underset{\sim}{\propto}$ means “approximately proportional to”. Recall from Eqn. 2.41 that

$$g_1(u) \underset{\sim}{\propto} \exp(-a(1-u)).$$

Then it is reasonable to choose a as

$$a = \frac{0.5}{1 - \cos \theta_0} \quad (2.42)$$

so that

$$\exp(-a(1-u)) = \exp\left(\frac{1}{2} \cdot \frac{1 - \cos \theta}{1 - \cos \theta_0}\right).$$

For heavy charged particles (such as muons), the model for θ_0 has been modified even more from Lynch and Dahl by GEANT4 [20]. Its form is now chosen by GEANT4 [20] as

$$\theta_0 = \frac{13.6 \text{ MeV}}{\beta cp} z_{ch} \sqrt{\frac{t}{X_0} \left[1 + 0.105 \ln\left(\frac{t}{X_0}\right) + 0.0035 \left(\ln\left(\frac{t}{X_0}\right) \right)^2 \right]} \left(1 - \frac{0.24}{Z(Z+1)} \right). \quad (2.43)$$

Finding u_0

Observe from Eqn. 2.41 that the parameter u_0 is the boundary between the Gaussian-like $g_1(u)$ and the Rutherford-like $g_2(u)$. GEANT4 [20] has chosen u_0 as

$$u_0 = 1 - \frac{3}{a}. \quad (2.44)$$

It is assumed that this parameter has been chosen as such based on empirical results, but no formal explanation is given. While GEANT4 [20] does not elaborate on this parameterization, validation in Section 4.2 will show that this choice is reasonable.

Finding d_r

Observe from Eqn. 2.41 that the parameter d_r the Rutherford-like exponent. Note that a classical Rutherford distribution would have $d_r = 2$. For heavy particles like muons, the parameter d_r has been chosen by GEANT4 [20] as

$$d_r = 2.40 - 0.027 Z^{\frac{2}{3}}. \quad (2.45)$$

It is assumed that this parameter has been chosen as such based on empirical results, but no formal explanation is given. While GEANT4 [20] does not elaborate on this parameterization, validation in Section 4.2 will show that this choice is reasonable.

Finding p_g

Now p_g will be found using constraints on the scattering distribution $g(u)$.

Observe from Eqn. 2.41 that the parameter p_g is the relative contribution of the Gaussian-like $g_1(u)$ to the Rutherford-like $g_2(u)$. It is reasonable to demand that $g(u)$ be continuous and smooth on $[-1, 1]$. Taking continuity and smoothness at $u = u_0$ yields the following constraints:

$$p_g g_1(u_0) = (1 - p_g) g_2(u_0), \quad (2.46)$$

$$a \cdot p_g g_1(u_0) = (1 - p_g) g_2(u_0) \cdot \frac{d_r}{b_r - u_0}. \quad (2.47)$$

From Eqn. 2.46, it is easy to see that

$$p_g = \frac{g_2(u_0)}{g_1(u_0) + g_2(u_0)}. \quad (2.48)$$

Finding b_r

Observe from Eqn. 2.41 that the parameter b_r is the relative u offset of the Rutherford-like distribution. For a classical Rutherford distribution, $b_r = 1$. From Eqn. 2.47,

$$a = \frac{d_r}{b_r - u_0}.$$

Rearranging this yields

$$b_r = \frac{a}{d_r} + u_0. \quad (2.49)$$

Finding q_g

Lastly, q_g is found by knowing that $g(u)$ must give the same mean value as Lewis theory. Observe from Eqn. 2.41 that the parameter q_g is the relative con-

tribution of the varying functions $g_1(u)$ and $g_2(u)$ to the constant function $g_3(u)$. GEANT4 [20] shows that

$$q_g = \frac{(1 - \frac{\lambda_{10} - \lambda_{11}}{\lambda_{10}})^{\frac{t}{\lambda_{10} - \lambda_{11}}}}{p_g \langle u \rangle_1 + (1 - p_g) \langle u \rangle_2}, \quad (2.50)$$

where λ_{10} is the value of the first transport free mean path at the beginning of the step, λ_{11} is this value at the end of the step, t is the true path length, and $\langle u \rangle_i$ is the mean value of u computed from the distribution $g_i(u)$.

CHAPTER 3

STOCHASTIC PROCESSES IN COSY INFINITY

3.1 Energy Straggling in COSY

For energy loss, COSY uses the Bethe-Bloch equation (Eqn. 2.5) to find the mean energy loss. This can be done in the transfer map paradigm since average energy loss is a deterministic effect. However, this work is concerned with simulating realistic stochastic fluctuations about the average energy loss. For this reason, this work has implemented Landau theory [22] to describe the straggling distribution. Landau theory is discussed in detail in Section 2.1.2.

The average of the Landau distribution is infinity, which is clearly nonphysical. Given the universal Landau parameter (Eqn. 2.16),

$$\lambda = \frac{\epsilon - \langle \epsilon \rangle}{\xi} - (1 - C_{Euler}) - \beta^2 - \ln(\xi/T_{max}),$$

fluctuations about the mean energy loss $\langle \epsilon - \langle \epsilon \rangle \rangle$ will also be divergent given enough samples. The possibility of divergence results in a sensitivity to step size since smaller steps will effectively produce a large sample size. In order to combat this, an artificial cutoff is given to λ such that the average Landau ϵ is equal to the average Bethe-Bloch energy loss $\langle \epsilon \rangle$. That is, it is required that

$$\begin{aligned} \langle \lambda \rangle &= \left\langle \frac{\epsilon - \langle \epsilon \rangle}{\xi} \right\rangle - (1 - C_{Euler}) - \beta^2 - \ln(\xi/T_{max}) \\ &= -(1 - C_{Euler}) - \beta^2 - \ln(\xi/T_{max}). \end{aligned}$$

If this cutoff is λ_{max} , then $\langle \lambda \rangle$ can be calculated by the definition of an average over the range $[0, \lambda_{max}]$:

$$\langle \lambda \rangle = \frac{\int_0^{\lambda_{max}} \lambda * f(\lambda) d\lambda}{\int_0^{\lambda_{max}} f(\lambda) d\lambda}.$$

The task then is to numerically find λ_{max} such that it satisfies

$$\frac{\int_0^{\lambda_{max}} \lambda * f(\lambda) d\lambda}{\int_0^{\lambda_{max}} f(\lambda) d\lambda} = -(1 - C_{Euler}) - \beta^2 - \ln(\xi/T_{max}).$$

GEANT version 3.21 [31] suggests the following form for λ_{max} :

$$\lambda_{max} = 0.60715 + 1.1934 \langle \lambda \rangle + (0.67794 + 0.052382 \langle \lambda \rangle) \exp[0.94753 + 0.74442 \langle \lambda \rangle)] \quad (3.1)$$

(note that GEANT4 does not have a section on the Landau cutoff since the Urbán model is used instead). However, a plot of required cutoffs (λ_{max}) vs. desired means ($\langle \lambda \rangle$) was produced independently in this work with muon ionization cooling parameters in mind. The form suggested by GEANT3 (see, e.g., Eqn. 3.1) was used to fit the plot in Figure 3.1. The determined form of the function is

$$\lambda_{max} = 0.517891 + 1.17765 \langle \lambda \rangle + (0.476074 + 0.00880733 \langle \lambda \rangle) \exp[1.15467 + 0.984008 \langle \lambda \rangle]. \quad (3.2)$$

Based on Eqn. 2.16, it is possible to find ϵ_{max} :

$$\epsilon_{max} = \xi[\lambda_{max} + (1 - C_{Euler}) + \beta^2 + \ln(\xi/T_{max})] + \langle \epsilon \rangle.$$

Therefore, during the energy loss sampling, if any energy loss ϵ is selected which is greater than ϵ_{max} it is thrown out and the sampling is performed again. However, if the result has been thrown out 100 times, the particle is assumed to have lost too much energy and is considered lost.

3.2 Multiple Scattering in COSY Infinity

Similar to ICOOL's fifth method of scattering, the Rutherford model (see Sec. 2.2), COSY utilizes a piecewise distribution function which is Gaussian at small angles (as Goudsmit and Saunderson suggested [29]) and Rutherford-like at large angles. This Rutherford-like tail is derived at length in Appendix C, with a review of

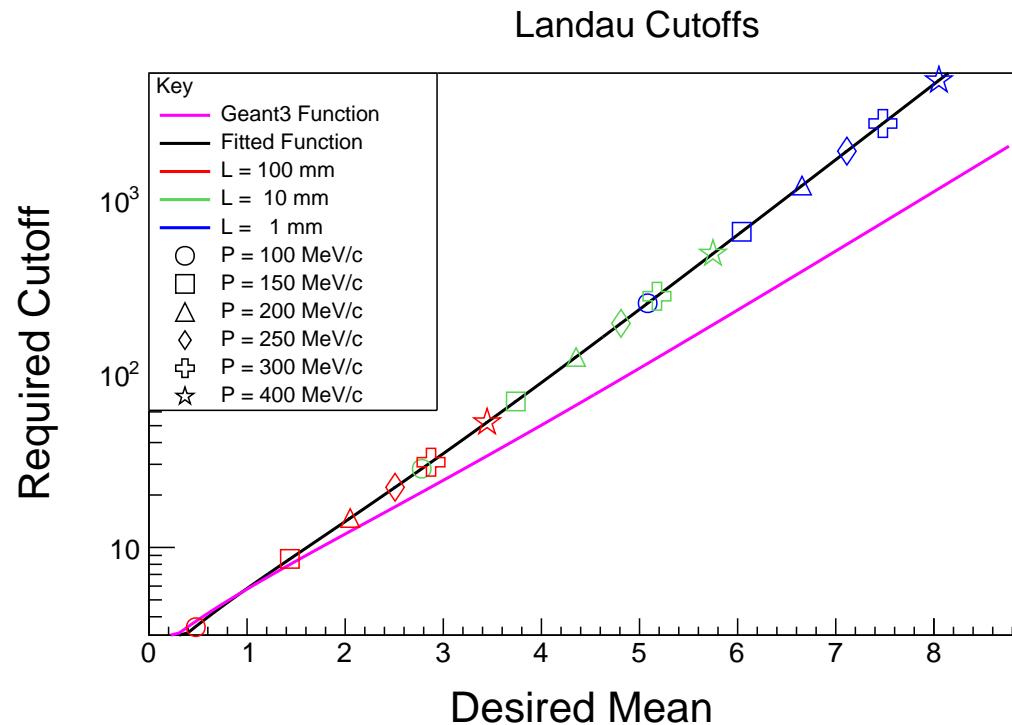


Figure 3.1: λ_{max} vs. $\langle \lambda \rangle$ over a variety of liquid hydrogen absorber lengths and initial beam momenta. The pink line is the form given by GEANT3 (see Eqn. 3.1) and the black line is the fitted curve (see Eqn. 3.2).

the relevant particle physics symbols and methods in Appendix B. The result is the (differential) Mott scattering cross section:

$$\frac{d\sigma}{d\Omega} \propto \frac{1 + \frac{(\beta\gamma)^2}{2}(1 + \cos\theta)}{(1 - \cos\theta)^2}, \quad (3.3)$$

where $d\sigma/d\Omega$ is the differential scattering cross section, β is the relativistic velocity, $\gamma = 1/\sqrt{1 - \beta^2}$, and θ is the scattering angle. Observe that for the non-relativistic limit $\beta \rightarrow 0$ the cross section does indeed approach a Rutherford distribution (Eqn.2.29). The practical implementation of this cross section into the probability distribution function is then discussed in Sec. 3.2.1.

3.2.1 Implementation. Now that the forms of the Gaussian and Rutherford-like scattering cross sections have been obtained, implementation of these cross sections will be discussed. In COSY, when a particle passes through matter, the change in angle of this particle is selected from a probability distribution. For $u = \cos\theta$, this distribution should be Gaussian-like at small angles [29] and follow the Mott cross section at large angles. Based on a Gaussian-like cross section for small angles and the cross section in Eqn. 3.3, the distribution has been chosen as

$$g(u) = \begin{cases} e^{-a(1-u)} & u_0, \leq u \\ \zeta \frac{1 + \frac{1}{2}(\beta\gamma)^2(1+u-b_c)}{(1-u+b_c)^2} & u \leq u_0 \end{cases}, \quad (3.4)$$

where u_0 is the cutoff between the Gaussian-like distribution and the Mott distribution, ζ is the amplitude of the Mott distribution, and b_c is the relative u shift of the Mott distribution. The parameter a is an empirical parameter, based off Highland theory [8], and can be found in Eqn. 2.42. Eqn. 2.42 is reproduced here for convenience:

$$a = \frac{0.5}{1 - \cos\theta_0}.$$

where θ_0 has the form [8]

$$\theta_0 = \frac{E_s}{p\beta} \sqrt{\frac{L}{X_0}}.$$

However, this definition has been extended in this work to include empirical corrections based on GEANT4's [20] treatment of θ_0 (see Eqn. 2.43):

$$\theta_0 = \frac{13.6 \text{ eV}}{\beta p} \sqrt{\frac{L}{X_0} \left[1 + h_1 \ln \frac{L}{X_0} + h_2 \left(\ln \frac{L}{X_0} \right)^2 \right]}. \quad (3.5)$$

The Highland correction terms have been chosen novelly in this work as $h_1 = 0.12$ and $h_2 = 0.006$. This was done by fitting the curve given by Eqn. 3.4 to match the MuScat results [32], the results of which can be found in Section ???. It is important to note that these correction terms are tunable for future data of muons through higher Z material.

u_0 is the point at which the Gaussian term meets the Mott tail. This has been chosen empirically as

$$u_0 = 1 - \frac{4.5}{a}. \quad (3.6)$$

This parameter was fitted alongside the Highland correction terms to match the experimental results in [32].

ζ and b_c are the angular scattering distribution's amplitude and offset for the tail. These are found by demanding continuity and smoothness at u_0 :

$$e^{-a(1-u_0)} = \zeta \frac{1 + \frac{1}{2}(\beta\gamma)^2(1+u_0-b_c)}{(1-u_0+b_c)^2},$$

$$ae^{-a(1-u_0)} = \zeta \frac{1 + \frac{1}{2}(\beta\gamma)^2(1+u_0-b_c)}{(1-u_0+b_c)^2} \left(\frac{2}{1-u_0+b_c} + \frac{(\beta\gamma)^2}{2 + (\beta\gamma)^2(1+u_0-b_c)} \right).$$

Then

$$a = \frac{2}{1-u_0+b_c} + \frac{(\beta\gamma)^2}{2 + (\beta\gamma)^2(1+u_0-b_c)}.$$

Solving this for $(u_0 - b_c)$ yields a quadratic with the solution

$$b_c = u_0 + \frac{A_2 + \sqrt{A_2^2 - 4A_1A_3}}{2A_1}, \quad (3.7)$$

with

$$\begin{aligned} A_1 &= -a(\beta\gamma)^2, \\ A_2 &= -(\beta\gamma)^2 - 2a, \\ A_3 &= (\beta\gamma)^2(a - 3) + 2a - 4. \end{aligned}$$

The constraint for continuity of $g(u_0)$ can now be solved to find ζ :

$$\zeta = \frac{e^{-a(1-u_0)}(1-u_0+b_c)^2}{1+\frac{1}{2}(\beta\gamma)^2(1+u_0-b)}. \quad (3.8)$$

Now that the distribution function has a concrete form, it is implemented by inverting the cumulative distribution function (CDF). Let $G(u)$ be the integral of $g(u)$. Then the variable G will be uniformly sampled over the region $[0, G_{max}]$. If $G \geq G(u_0) \equiv G_0$ then the Gaussian part of the distribution is used to generate u (i.e. $G(u \geq u_0)$). Otherwise, the tail of the distribution is used. Figure 3.2 shows $G(u)$ for $L = 100$ mm, $p = 200$ MeV/c, and a radiation length of $X_0 = 8.66$ m (to simulate a liquid hydrogen target).

However, since this is a piecewise function, the CDF will be inverted in pieces. The tail of the CDF will be found first:

$$G(u \leq u_0) = \int_{-1}^u \zeta \frac{1 + \frac{1}{2}(\beta\gamma)^2(1+u'-b_c)}{(1-u'+b_c)^2} du'.$$

This integral may be solved by substituting $v = 1 - u' + b_c$ and simply splitting the numerator into separate parts:

$$\begin{aligned} G(u \leq u_0) &= -\zeta \left(1 + \frac{1}{2}(\beta\gamma)^2(1-b_c)\right) \int_{2+b_c}^{1-u+b_c} v^{-2} dv - \\ &\quad \zeta \frac{(\beta\gamma)^2}{2} \int_{2+b_c}^{1-u+b_c} (v^{-2} - v^{-1} + b_c v^{-2}) dv \\ G(u \leq u_0) &= \zeta \left(1 + (\beta\gamma)^2\right) \left(\frac{1}{1-u+b_c} - \frac{1}{2+b_c}\right) + \zeta \frac{(\beta\gamma)^2}{2} \ln \left(\frac{1-u+b_c}{2+b_c}\right) \quad (3.9) \end{aligned}$$

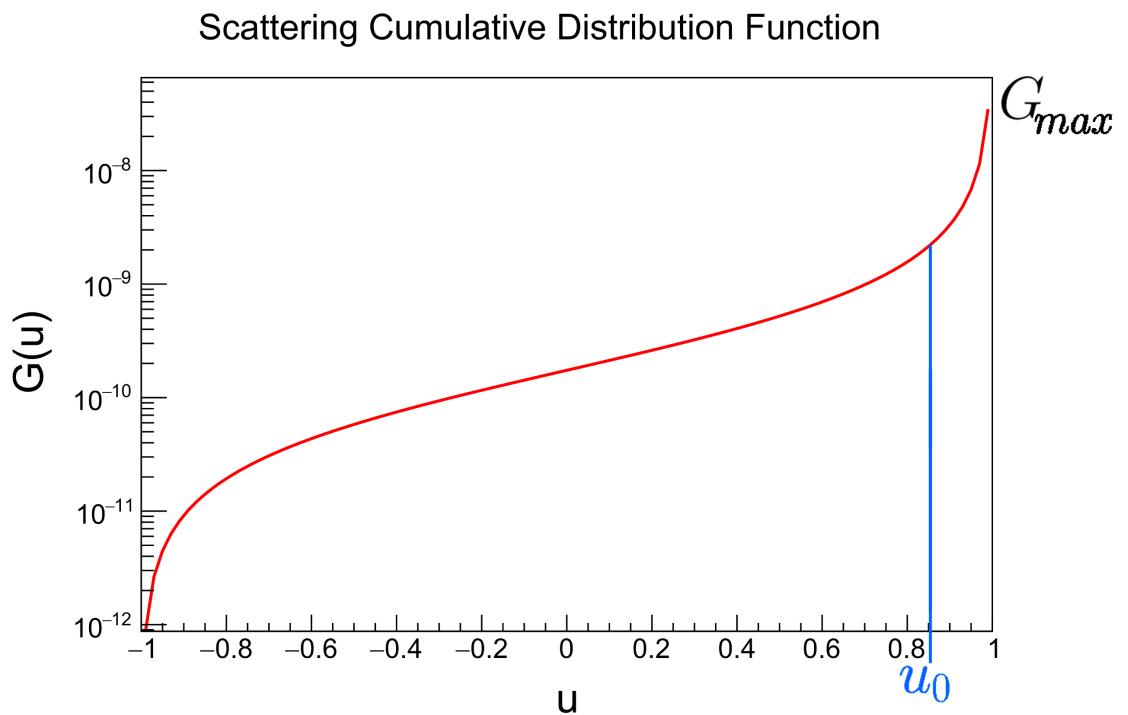


Figure 3.2: Example of the cumulative angular distribution function for muons with momenta of 200 MeV/c through 100 mm of liquid hydrogen. Note that the y axis is log scaled due to the very sharp peak. Furthermore, note that u_0 is greatly exaggerated, since its actual value for these parameters is 0.99987.

The goal is to solve Eqn. 3.9 for $u(G)$. However, using direct inversion is extremely difficult and involves special functions. Therefore, it is more prudent to generate u via bisection method (see Figure 3.3). In this method, the true G is sampled uniformly on the range $[0, G_{max}]$, where $G_{max} \equiv G(1)$. Explicitly, G_{max} is

$$\begin{aligned} G_{max} \equiv G(1) &= \int_{-1}^{u_0} \zeta \frac{1 + \frac{1}{2}(\beta\gamma)^2(1+u-b_c)}{(1-u+b_c)^2} du + \int_{u_0}^1 e^{-a(1-u)} du, \\ &= G_0 + \frac{1}{a} - \frac{e^{-a(1-u_0)}}{a}. \end{aligned}$$

If $G < G_0$, then the tail is sampled. A trial u called \bar{u} (as in ‘average’) is selected from some range which is known to contain the actual u . The range is described as $[u_{min}, u_{max}]$, and $\bar{u} = (u_{min} + u_{max})/2$.

Initially, the range is chosen as $u_{min} = -1$ and $u_{max} = u_0$ (since that is the largest range on which $G(u \leq u_0)$ is valid and hence u is guaranteed to be in this range). $\bar{G} \equiv G(\bar{u})$ is found using Eqn. 3.9, and then the routine is subject to the following conditions:

- | | |
|---|---|
| If $\bar{G} \in [G - \delta G, G + \delta G]$ | then $u = \bar{u}$, return value. |
| If $\bar{G} < G - \delta G$ | then $u_{min} = \bar{u}$, rerun with new u_{min} . |
| If $\bar{G} > G + \delta G$ | then $u_{max} = \bar{u}$, rerun with new u_{max} . |

δG is a precision, and has been chosen for this work as $\delta G = 10^{-8}$. However, it is conceded that in the future δG should be a percentage of G_{max} rather than an absolute number.

For the peak, $u_0 \leq u$ and so the CDF becomes

$$G(u_0 \leq u) = \int_{-1}^{u_0} g(u) du + \int_{u_0}^u e^{-a(1-u)} du.$$

The first term is simply G_0 , the cumulative distribution function at u_0 . The second term is easily integratable and yields

$$G(u_0 \leq u) = G_0 + \frac{e^{-a(1-u)} - e^{-a(1-u_0)}}{a}. \quad (3.10)$$

Scattering Cumulative Distribution Function

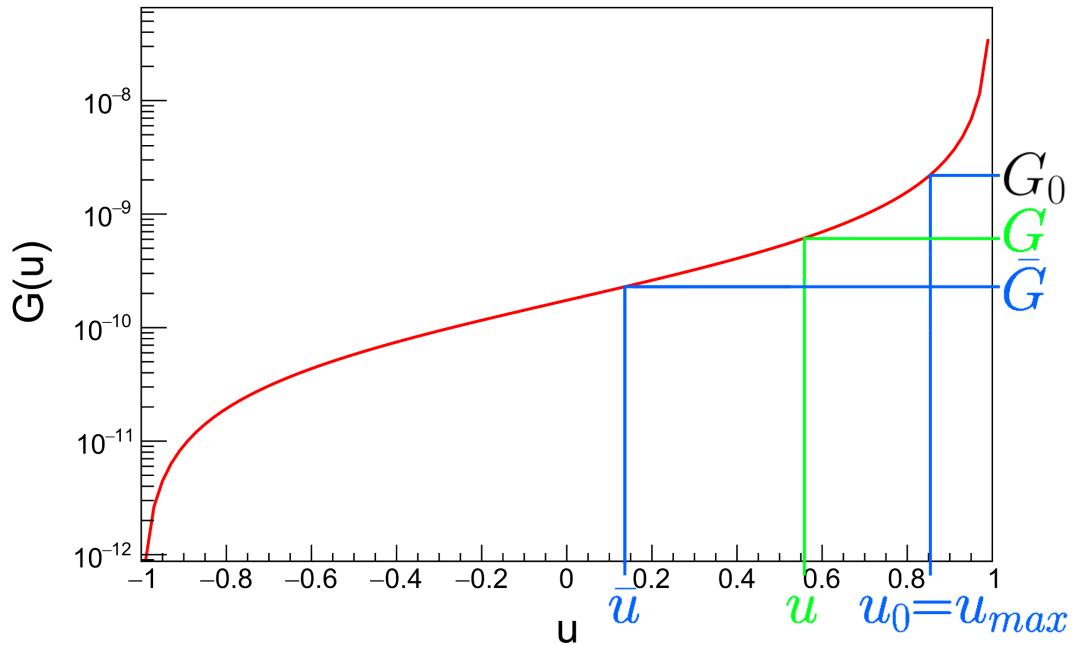


Figure 3.3: Example of the first iteration of the algorithm to obtain the true u (in green). The true G is chosen uniformly from $G \in [0, G_{max}]$. If $G < G_0$, then the tail is sampled via bisection method. In this case, since $\bar{G} < G$, \bar{u} is the new u_{min} and \bar{u} is calculated again.

The inversion of this function is quite straightforward, and is

$$u(G_0 \leq G) = 1 + \frac{1}{a} \ln [a(G - G_0) + e^{-a(1-u_0)}]. \quad (3.11)$$

Therefore, if $G \in [0, G_{max}]$ is greater than or equal to G_0 , then it is simply inserted into Eqn. 3.11 and the true u is obtained.

It is a subtle yet important point to note that the Mott cross section (Eqn. 3.3), upon which the probability distribution function $g(u)$ in Eqn. 3.4 is based, assumes an on-axis straight line trajectory (that is, $x = y = p_x = p_y = 0$). The treatment of this subtlety will be discussed here.

The implemented routine for this work which uses the scattering distribution $g(u)$ is called SCATDIST. This routine takes two arguments: θ_0 (the Highland-like critical scattering angle from Eqn. 3.5) and p , the total momentum. θ_0 includes not only the dependence on material parameters (L, X_0) but also dependence energy terms ($1/\beta p$). p is the momentum *after* the straggling routine has been called. SCATDIST returns not the scattered angle, but the new z momentum $p_z = pu = p \cos \theta$.

This new z momentum is in the rotated frame, i.e. the frame in which $p_x = p_y = 0$ (see Figure 3.4), and so is called $p_{z,R}$. The transverse momentum in the rotated frame, $p_{T,R}$, can be found via $p_{T,R} = \sqrt{p^2 - p_{z,R}^2}$, but this is the total transverse angle ($\theta_o + \theta_s$), which includes the original angle (θ_o), not simply the transverse angle which was gained via scattering (θ_s). Rotation back into the lab frame is given by

$$\begin{pmatrix} p_z \\ p_T \end{pmatrix} = \begin{pmatrix} \cos \theta_o & -\sin \theta_o \\ \sin \theta_o & \cos \theta_o \end{pmatrix} \begin{pmatrix} p_{z,R} \\ p_{T,R} \end{pmatrix}.$$

Note that at this point one must not uniformly distribute p_T into p_x and p_y . This is because, for example, if p_x was positive before scattering it should have a strong probability of being positive after scattering. If one were to distribute p_T

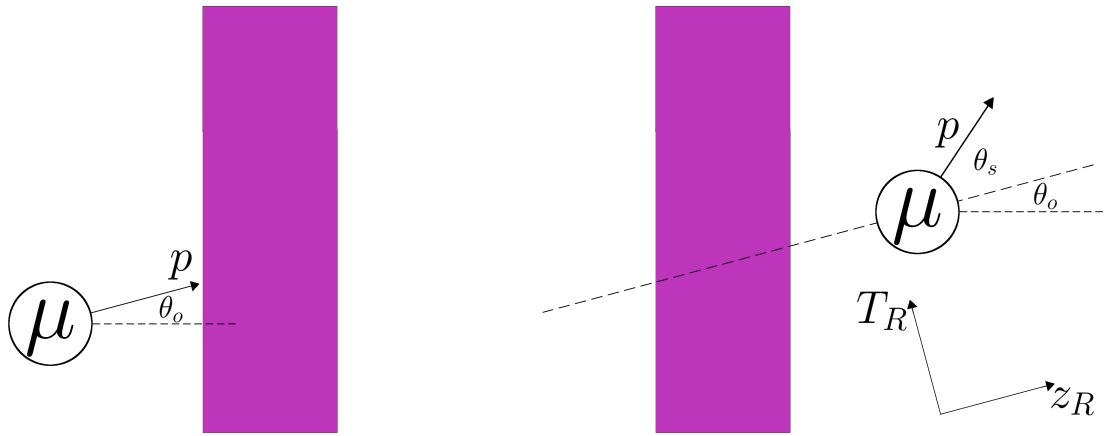


Figure 3.4: Example of a muon entering an absorber (purple) with some nonzero initial angle θ_o . The muon then scatters an angle θ_s with respect to its initial momentum \vec{p} . The scattering distribution $g(u)$ assumes a straight, on-axis particle ($x = y = p_x = p_y = 0$), and so is in the rotated frame, represented by T_R, z_R .

uniformly then p_x would have a 50/50 chance of being negative. For this reason, only the transverse momentum gained via scattering should be uniformly distributed into p_x and p_y .

Let the original momenta (after straggling but before scattering) be denoted with the subscript o in the same fashion that θ_o is the angle before scattering. Then the amount of transverse momentum which was gained via scattering is $P_T - P_{T,o}$. This new amount of transverse momentum must be added to the original $p_{x,o}$ and $p_{y,o}$ uniformly due to cylindrical symmetry. Consequently, let ϕ be an angle chosen from $[0, 2\pi]$. Then the final p_x and p_y are

$$p_x = p_{x,o} + (P_T - P_{T,o}) \cos \phi$$

$$p_y = p_{y,o} + (P_T - P_{T,o}) \sin \phi.$$

In summation, the angular distribution used by COSY Infinity is based on a piecewise function which is Gaussian for small angles [29] and has a Mott tail for large angles. This distribution is represented by $g(u)$ in Eqn. 3.4, where $u \equiv \cos \theta$. $g(u)$ has four parameters: a , an empirical parameter that is based on Highland theory [8] and is dependent on some critical angle θ_0 , defined in Eqn. 3.5; u_0 , the empirical cutoff angle that distinguishes which angles are Gaussian and which are not, found by Eqn. 3.6; b_c , a parameter derivable from smoothness of g at u_0 , which represents the offset of the Mott tail, found in Eqn. 3.7; and ζ , a parameter derivable from continuity of g at u_0 which represents the amplitude of the Mott tail, found in Eqn. 3.8. From $g(u)$, its antiderivative $G(u)$ may be found and a particular G may be picked from the range $[0, G_{max}]$. If $G < G_0 \equiv G(u_0)$, then u comes from the Mott tail and a bisection method is used to find u . If $G_0 \leq G$ then u comes from the Gaussian peak and $G(u)$ may be inverted to find $u(G)$. This scattered angle must then be rotated into the lab frame and the additional transverse momentum must be uniformly distributed into p_x and p_y .

3.3 Transverse Displacement in COSY Infinity

Due to multiple scattering events, when a particle traverses matter a direct correlation between the particle transverse position and scattered angle is not always clear. Two identical particles with identical initial conditions may end up with identical scattered angles but different transverse positions (see Figure 3.5). This is because these two particles may take different paths through the absorber. While both of these paths may lead to a similar final angle with respect to the z axis, the positions will likely be different due to their trajectories.

For this reason, transverse empirical corrections have been implemented in COSY. Due to a lack of empirical data at the time, COSY was match ‘empirically’ with G4Beamline across several initial momenta and absorber lengths. The result

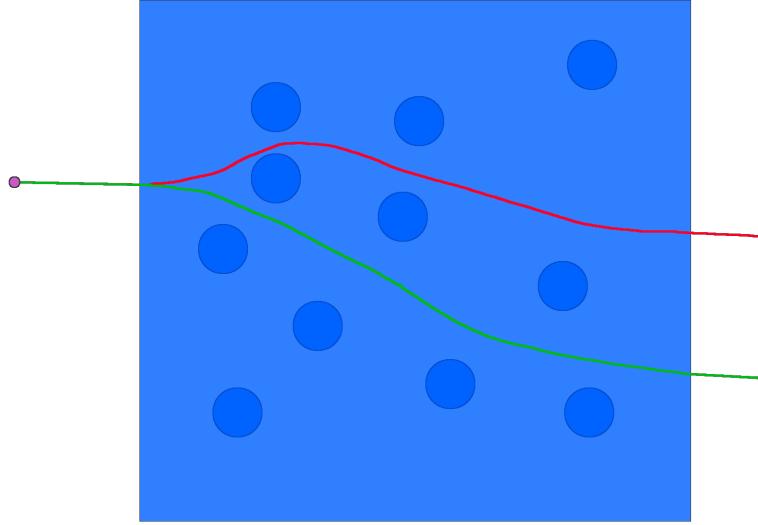


Figure 3.5: Two examples of true paths which a particle might take when traversing a medium. Note that both the red path and green path have the same final scattered angle but different transverse positions.

was

$$x = x_o + x_D + \text{Gaus}(\theta_{diff} * L/2, \theta_c/(2\sqrt{3})), \quad (3.12)$$

where x_o is the original x position, $x_D = L*P_{x,o}/P_{z,o}$ is the deterministic gain in x , and $\text{Gaus}(\mu, \sigma)$ is a randomly selected number from a Gaussian distribution with mean μ and standard deviation σ . The forms of μ and σ were selected based on a combination of the Particle Data Group [9] and Fernow and Gallardo [33]. Note that the average $\mu = \theta_{diff} * L/2$ represents the transverse displacement that would have occurred if all of the angular scattering had happened at the point $L/2$. $\theta_{diff} = \theta_{final} - \theta_o$ is the amount of deflection which occurred due to scattering and $\theta_c = 13.6 \text{ eV}/\beta p \cdot \sqrt{1/X_0}$ is the coefficient from Highland theory [8].

The fitting of these parameters will be discussed next. As previously mentioned, the data for the fits was generated by G4Beamline [19]. Here, a particular

example of a simulation of 10^6 particles through 1 cm of liquid hydrogen will be shown. The initial beam distribution was a pencil beam of momentum 200 MeV/ c . A plot of the histogram of (x, p_x) phase space can be seen in Figure 3.6. It can be seen in Figure 3.7 that the cross section of a given transverse momentum results in a nearly-Gaussian x histogram. It can also be inferred by the figure that the mean and standard deviation of the Gaussian fit varies depending on which p_x is chosen. For example, higher values of p_x appear to have both a larger mean and standard deviation. While this trend is conceptual, it will aid in the understanding of the form Eqn. 3.12.

The phase space portrait in Figure 3.8 shows that the distribution is only locally Gaussian. However, fitting in the range of $p = 100\text{--}400$ MeV/ c , $L = 1\text{--}100$ mm has shown that this is a good approximation for the majority of particles (see Section 4.1).

An example of the success of this implementation can be seen in Figure 3.9, where 10^6 muons of momentum 250 MeV/ c were simulated through 100 mm of liquid hydrogen. The simulation was carried out in both COSY and ICOOL [18] with excellent agreement.

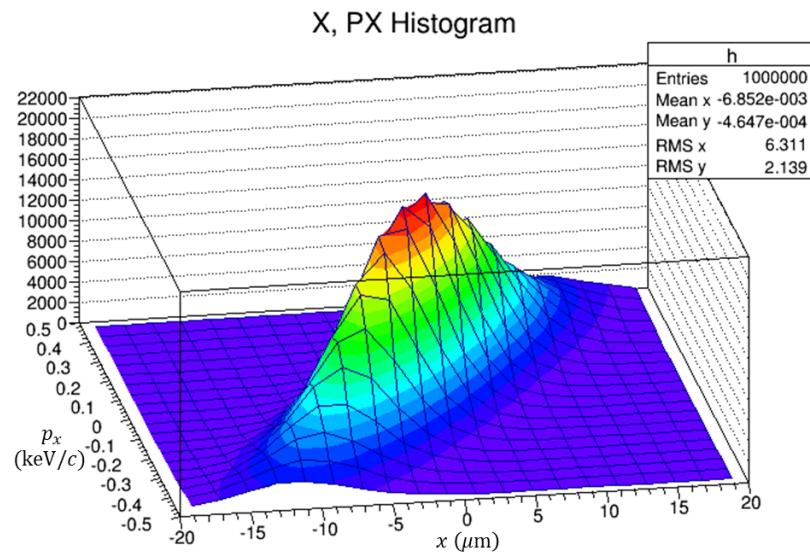


Figure 3.6: 2D histogram of (x, p_x) phase space for 10^6 muons of momentum 200 MeV/ c through 1 cm of liquid hydrogen.

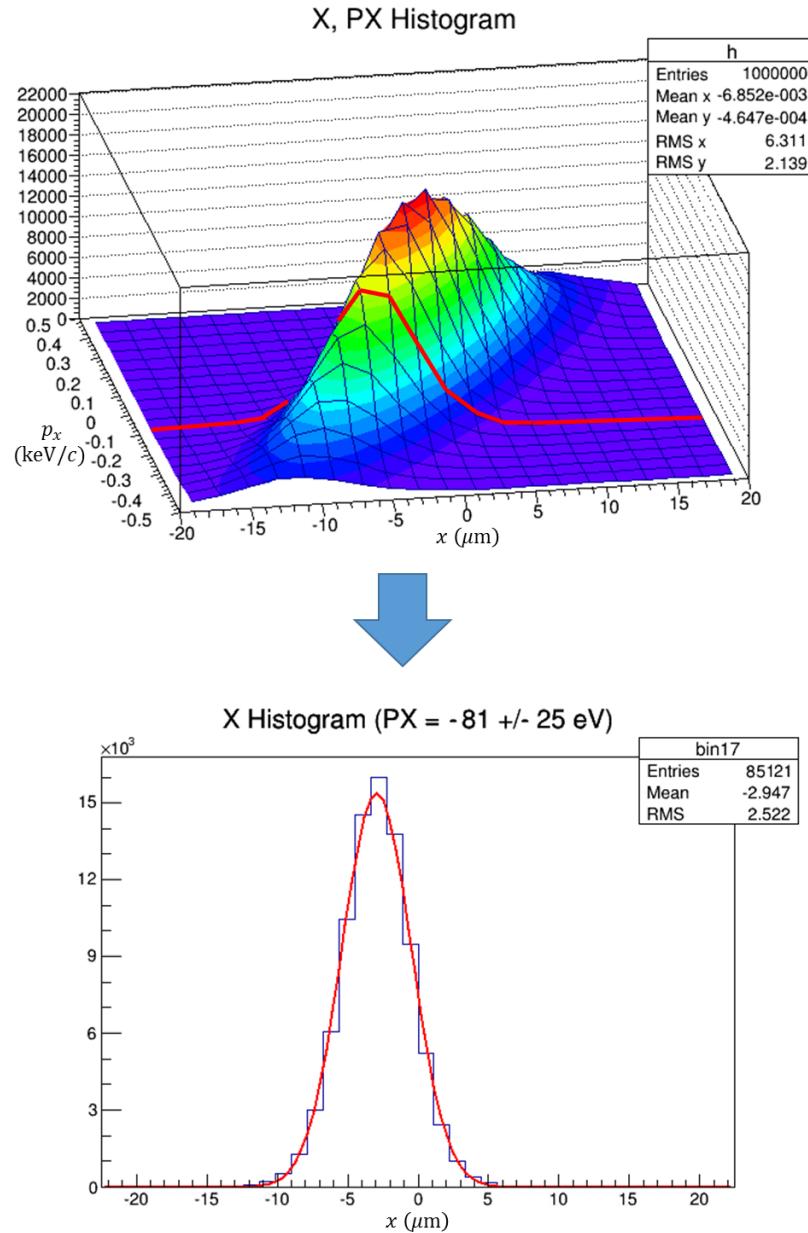


Figure 3.7: Cross section of Figure 3.6 at $p_x = (-0.081 \pm 0.025)$ keV/c. The resulting x histogram fit (in red) is nearly Gaussian.

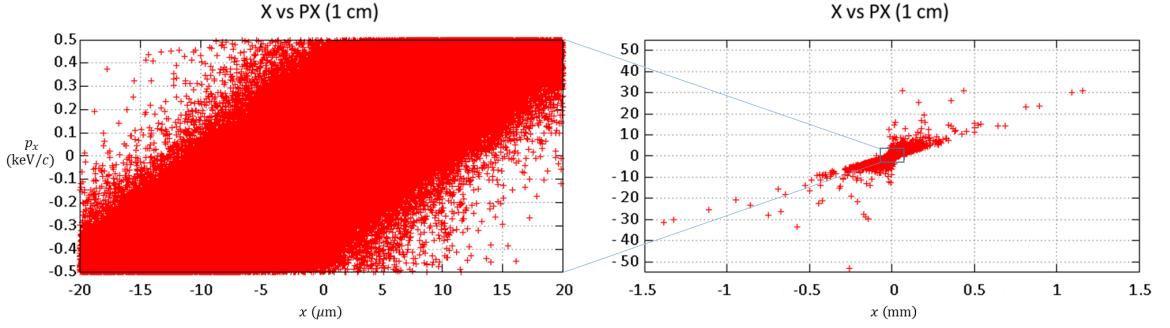


Figure 3.8: Phase space portrait of Figure 3.6. Observe that the bulk in Figure 3.6 is represented on the left. Further observe that the extrema of this distribution do not follow a Gaussian curve.

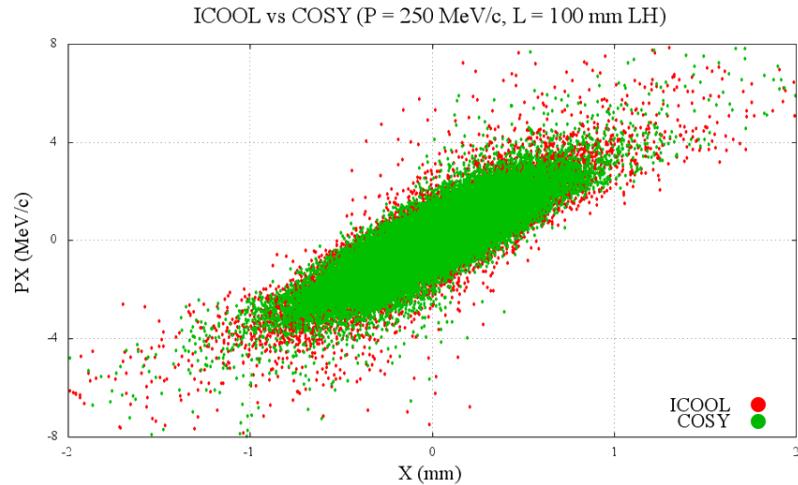


Figure 3.9: Sample simulation results for the implementation of the transverse coordinate algorithms discussed in this section.

3.4 Temporal Displacement in COSY Infinity

For the time-of-flight offset, both the deterministic and stochastic processes are handled in the same routine. To first order, the particle decelerates at some constant (or average) rate through an absorber of length L . If a is the constant acceleration

then

$$v_f = v_o + a\Delta t,$$

or

$$a = \frac{v_f - v_o}{\Delta t}.$$

At the same time, $v_f^2 = v_o^2 + 2aL$, and so

$$a = \frac{v_f^2 - v_o^2}{2L}.$$

Then

$$\Delta t = \frac{(v_f - v_o)2L}{v_f^2 - v_o^2}.$$

Given $\beta = p/E$ and $v = \beta c$, then

$$\Delta t = \frac{\left(\frac{p_f}{E_f} - \frac{p_o}{E_o}\right)2L}{\left(\frac{p_f^2}{E_f^2} - \frac{p_o^2}{E_o^2}\right)c}. \quad (3.13)$$

However, COSY does not have a time variable, but rather a variable ℓ that is described as the time-of-flight in units of length. In COSY [4], this is defined as

$$\ell = \frac{-(t - t_0)v_0\gamma}{1 + \gamma},$$

where the subscript 0 signifies the reference particle. Let the time before a step be denoted t_1 and the time after a step denoted t_2 . Then to find ℓ_2 given ℓ_1 and Δt from Eqn. 3.13, observe that

$$\begin{aligned} \ell_1 &= \frac{(t_{01} - t_1)v_{01}\gamma_1}{1 + \gamma_1} = (t_{01} - t_1)A_1 \\ \ell_2 &= \frac{(t_{02} - t_2)v_{02}\gamma_2}{1 + \gamma_2} = (t_{02} - t_2)A_2, \end{aligned} \quad (3.14)$$

where

$$A_n \equiv v_{0n}\gamma_n/(1 + \gamma_n) \quad \text{for } n = 1, 2. \quad (3.15)$$

Then

$$\begin{aligned}
\ell_2 - \ell_1 &= (t_{02} - t_2)A_2 - (t_{01} - t_1)A_1 \\
&= [(t_{02} - t_{01}) + t_{01}] - [(t_2 - t_1) + t_1]A_2 - (t_{01} - t_1)A_1 \\
&= (\Delta t_0 - \Delta t)A_2 + (t_{01} - t_1)A_2 - (t_{01} - t_1)A_2 \\
&= (\Delta t_0 - \Delta t)A_2 + (t_{01} - t_1)(A_2 - A_1).
\end{aligned}$$

Eqn. 3.14 says that $t_{01} - t_1 = \ell_1/A_1$. Moving ℓ_1 to the right hand side,

$$\begin{aligned}
\ell_2 &= (\Delta t_0 - \Delta t)A_2 + \frac{\ell_1}{A_1}(A_2 - A_1) + \ell_1 \\
&= (\Delta t_0 - \Delta t)A_2 + \ell_1\left(\frac{A_2 - A_1}{A_1} + \frac{A_1}{A_1}\right) \\
&= (\Delta t_0 - \Delta t)A_2 + \ell_1\frac{A_2}{A_1}.
\end{aligned}$$

Substituting for A_n via Eqn. 3.15 yields the final result

$$\ell_2 = \frac{(\Delta t_0 - \Delta t)v_{02}\gamma_2}{1 + \gamma_2} + \ell_1\frac{v_{02}\gamma_2(1 + \gamma_1)}{v_{01}\gamma_1(1 + \gamma_2)}. \quad (3.16)$$

Using Eqn. 3.16, Δt from Eqn. 3.13 can be directly input into the new COSY variable for time-of-flight in units of length. Figure 3.10 shows the simulation results of a beam of 10^6 muons of momentum 172 MeV/c through 109 mm of liquid hydrogen, which were the parameters of the MuScat experiment [32]. The COSY results are shown alongside ICOOL [18]. Note that the agreement is quite good. ICOOL displays a thicker bulk, but it is approximately 1 fs in width—about 0.2% of the average time-of-flight.

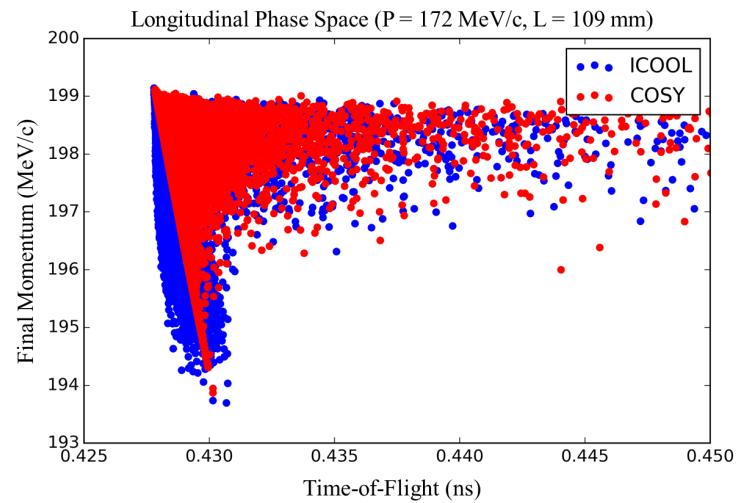


Figure 3.10: Sample simulation results for the implementation of the temporal displacement algorithm discussed in this section.

CHAPTER 4

RESULTS

In this chapter, comparison of the new COSY routines will be benchmarked against both ICOOL and G4Beamline for pencil beams. Recall that a pencil beam is a beam with an RMS of zero for all transverse coordinates. Pencil beams are used in order to simulate a plethora of possible paths and energy losses for a particular initial condition. Both validation against past experiments and predictions of current muon ionization cooling channel efforts will be shown.

4.1 Benchmark Against Other Codes

This section briefly discusses the comparison of the new COSY routines with those of two other codes, ICOOL [18] and G4Beamline [19].

Please refer to Appendix F for Figures F.1–F.12 (12 figures, each with three subplots) mentioned here. The benchmarking was done over both the typical momentum range of cooling channels (100, 200, 300, and 400 MeV/ c) and various absorber lengths (1, 10, and 100 mm). These simulations were performed using a pencil beam of 5×10^4 muons with the aforementioned momenta through liquid hydrogen. Note that the transverse position and transverse momentum histograms each have 10^5 entries. This is because the histograms included both x and y components since the system was cylindrically symmetric. Further note that the absolute value of the coordinate was plotted. For ICOOL and COSY, the step size was the entire absorber length. For G4Beamline, the default step size was used except in the 1 mm absorber case. In this case, the step size was limited to 0.1 mm. This was done because of the heavy dependence of the transverse position on step size for short absorbers.

For the 1 mm figures (Figures F.1, F.4, F.7, and F.10) there is much disagreement between the codes' transverse momentum distributions. Since the transverse momentum and transverse position coordinates are related, there is also disagreement among the transverse position distributions. This is likely because the default ICOOL scattering model uses a Fano peak with a Rutherford tail whereas both COSY and G4Beamline use a Gaussian peak. It can be seen by, e.g., Figure F.1 that COSY, like G4Beamline, uses the Gaussian model for the peak but, like ICOOL, then switches to a Rutherford tail.

For the 10 mm figures (Figures F.2, F.5, F.8, and F.11), both x and p_x distributions agree well. For COSY, the tail of the distribution falls off slightly faster when the number of particles becomes less than 100.

For the final momentum plots, ICOOL and COSY agree quite well. This is not surprising since both codes use Landau theory to describe the energy loss distribution. G4Beamline occasionally disagrees. An example of this can be seen in Figure F.4. However, given the precision of the horizontal axis this disagreement is much smaller than it initially appears.

4.2 Validation

The new COSY routines were also compared to the Muon Scattering Experiment [32], often referred to as MuScat. This experiment measured the scattering of a beam of collimated muons through seven materials. To emulate this, pencil beams with momentum $P = 172$ MeV/c were created in COSY, G4Beamline, and ICOOL. The pencil beam consisted of 5×10^6 particles and was propagated through 109 mm of liquid hydrogen, 159 mm of liquid hydrogen, and 3.73 mm of beryllium. Liquid hydrogen was chosen to represent muons through large absorbers of low Z materials. Beryllium, on the other hand, was chosen to represent muons through much thinner

and denser media. COSY and ICOOL took a single step through the absorbers while G4BL took a step size of 1 mm. The data points were normalized to the total probability, was calculated via Simpson's rule. The probability per radian was then found by dividing the probability of the data point by the scattered angle at that point. The results are shown below.

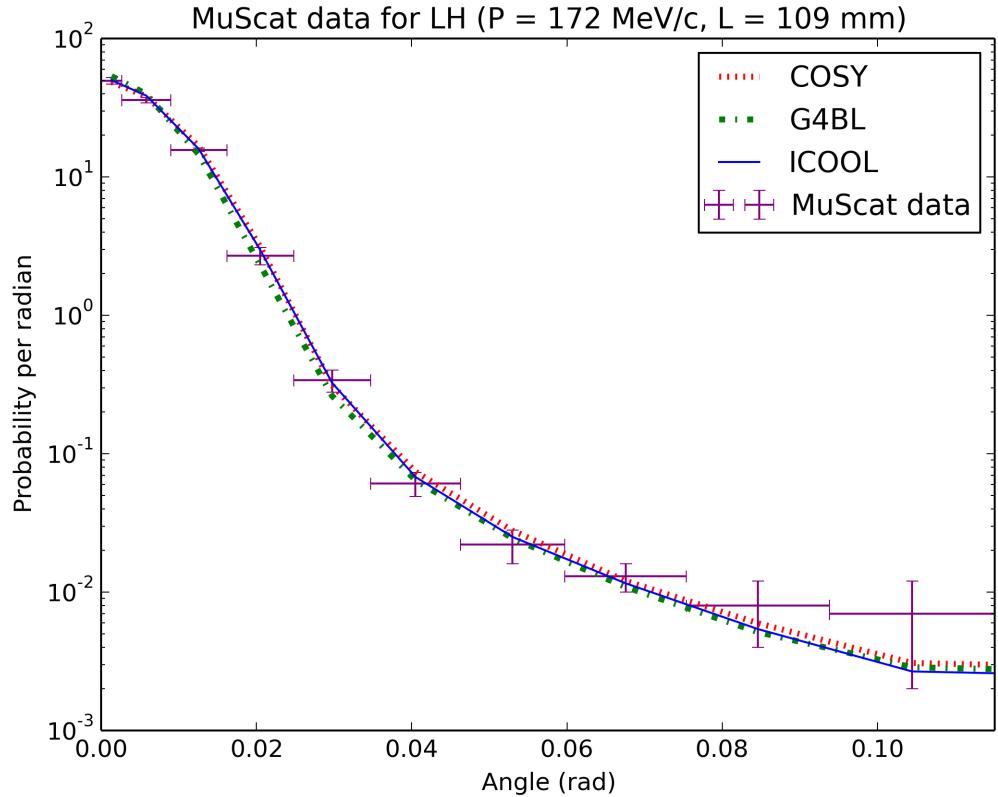


Figure 4.1: MuScat results for 109 mm of liquid hydrogen compared against COSY (red), G4BL (green), and ICOOL (blue).

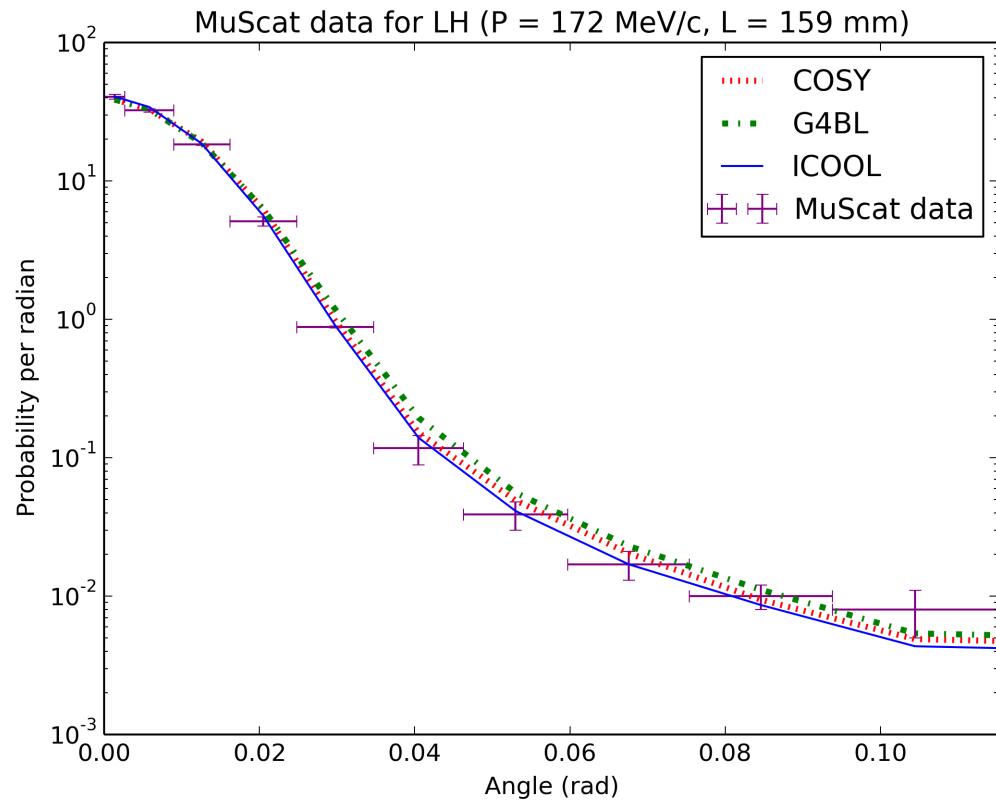


Figure 4.2: MuScat results for 159 mm of liquid hydrogen compared against COSY (red), G4BL (green), and ICOOL (blue).

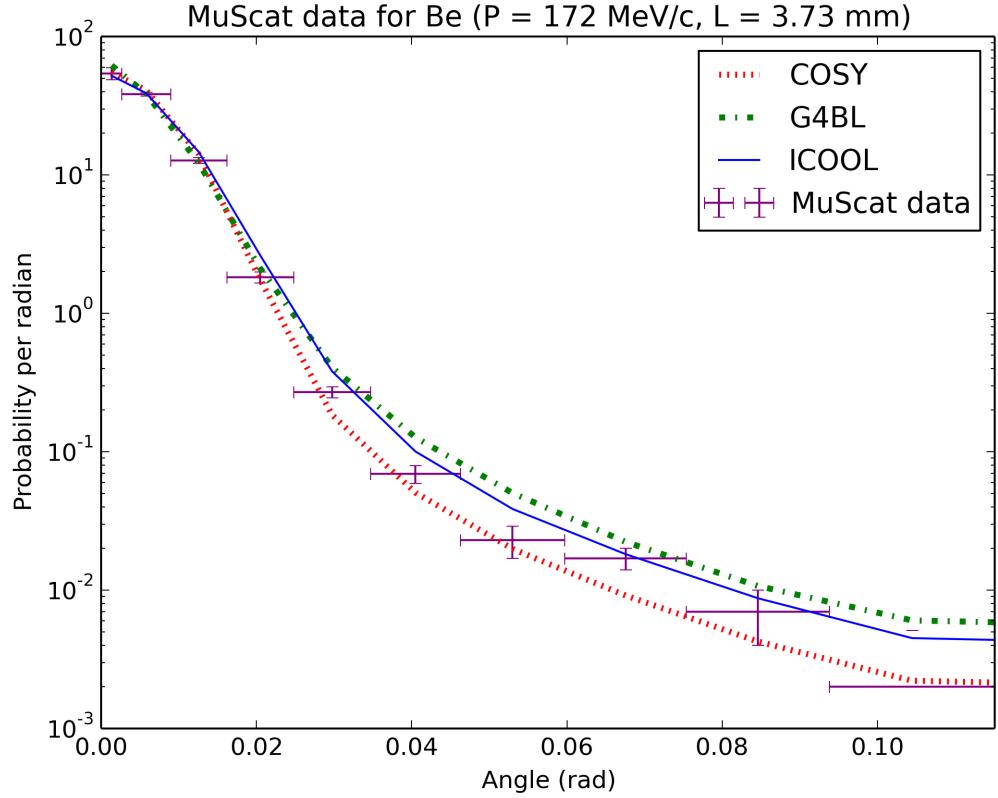


Figure 4.3: MuScat results for 3.73 mm of beryllium compared against COSY (red), G4BL (green), and ICOOL (blue).

In the liquid hydrogen cases, COSY appears to match both G4Beamline and ICOOL very well, as well as the MuScat data points. ICOOL appears to have a sharper peak than either COSY or G4Beamline, which can be more easily seen in the 159 mm liquid hydrogen case than in the 109 mm case. In the case of beryllium, COSY matches the MuScat points slightly better than G4Beamline and ICOOL, particularly for the two data points between 0.04 and 0.06 radians.

4.3 The Muon Ionization Cooling Experiment

This section introduces the Muon Ionization Cooling Experiment (MICE, [34]), a practical application for the new absorber routines in COSY. The results of the

MICE simulations in both G4Beamline and COSY will be examined, showing good agreement for liquid hydrogen.

4.3.1 Introduction to MICE. The Muon Ionization Cooling Experiment (MICE [34]) is an experiment currently being developed at the Rutherford Appleton Laboratory in Oxfordshire, U.K. Its goal is to show a proof-of-principle demonstration of muon ionization cooling. While there are six steps, the MICE Step IV configuration is explored in this work. The Step IV cell includes 12 magnetic coils positioned symmetrically around a liquid hydrogen absorber, shown in Figure 4.4.

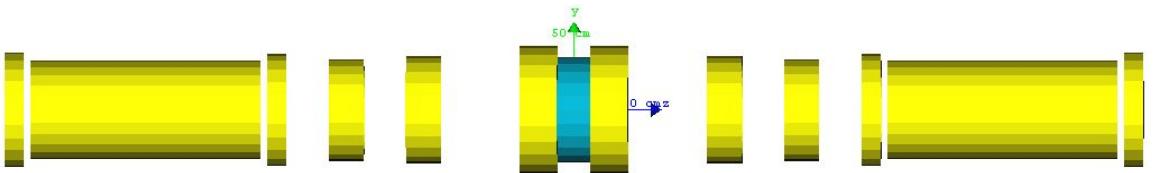


Figure 4.4: MICE Step IV cell. Magnetic coils are shown in yellow and the absorber is shown in grey. The green and blue axes are the y and z axes, here drawn to scale as 50 cm each. Image rendered via G4Beamline [19].

4.3.2 Results of the MICE Simulation. 1,000,000 particles were simulated through the cell in Figure 4.4. The coil parameters may be found in Table 4.1. The absorber was a 350 mm cylindrical block of liquid hydrogen centered at $z = 0$. The decay process was disabled in all simulation codes. The beam started at -2.45105 m and ended at 2.450 m. The Gaussian parameters for the initial distribution of particles can be found in Table 4.2.

For COSY, it was found that a 5th order simulation was sufficient. Through the coil-only portion of the simulation, 32 steps were taken on each side of the absorber (or roughly a step size of 76 mm both upstream and downstream). The particles were tracked through the momentary transfer map after each step and then the transfer map was cleared. It was noted that a single map was not sufficient even at the 21st

MICE Step IV Coil Parameters

Name	z position mm	Length mm	Inner radius mm	Outer radius mm	Current density A/mm ²
End2	∓ 3200.28	110.642	258	325.783	± 126
Center	∓ 2450.275	1314.3	258	280.125	± 148
End1	∓ 1700.29	110.642	258	318.905	± 133
Match2	∓ 1300.29	199.492	258	288.925	± 132
Match1	∓ 860.645	201.268	258	304.165	± 133
Focus	∓ 202.2	213.3	267.6	361.9	± 104

Table 4.1: MICE Step IV coil parameters corresponding to Figure 4.4.

MICE Step IV Initial Distribution Parameters

Parameter	Mean	Standard deviation
x (mm)	0	32
y (mm)	0	32
z (mm)	0	0
p_x (MeV/ c)	0	20
p_y (MeV/ c)	0	20
p_z (MeV/ c)	200	30

Table 4.2: MICE Step IV initial distribution Gaussian parameters.

order. This is likely due to the large phase space volume of the initial beam and the complexity of the magnetic field.

Compounding the map without propagating the beam also gave poor results. When one takes the composition of two n^{th} order transfer maps, a transfer map of order $n \times n$ is the result. For example, the first step in the MICE simulation would yield a 5th order transfer map. Taking the second step would give a new transfer map of order $5 \times 5 = 25$. However, since COSY is operating in the 5th order mode, the new transfer map would not be 25th order, but rather it would be truncated to a 5th order map. For this reason, the particles were propagated through the momentary transfer map after each step in the simulation.

When the beam entered the region of the absorber, COSY switched to a step size of 1 mm in order to simulate the superposition of the magnetic coils and the absorber.

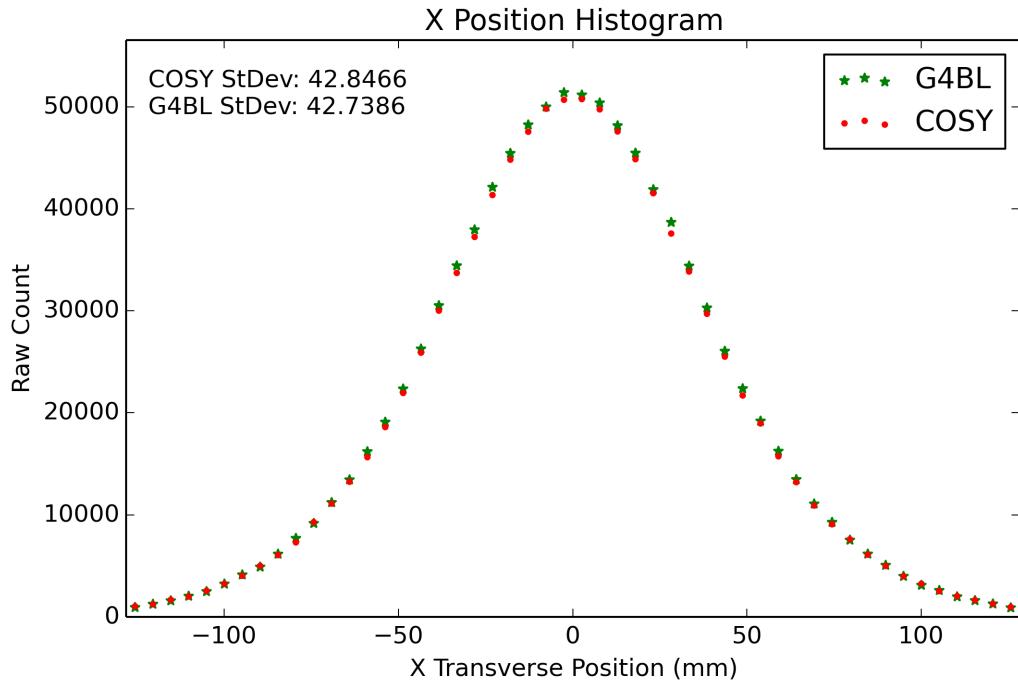


Figure 4.5: MICE Step IV x position results.

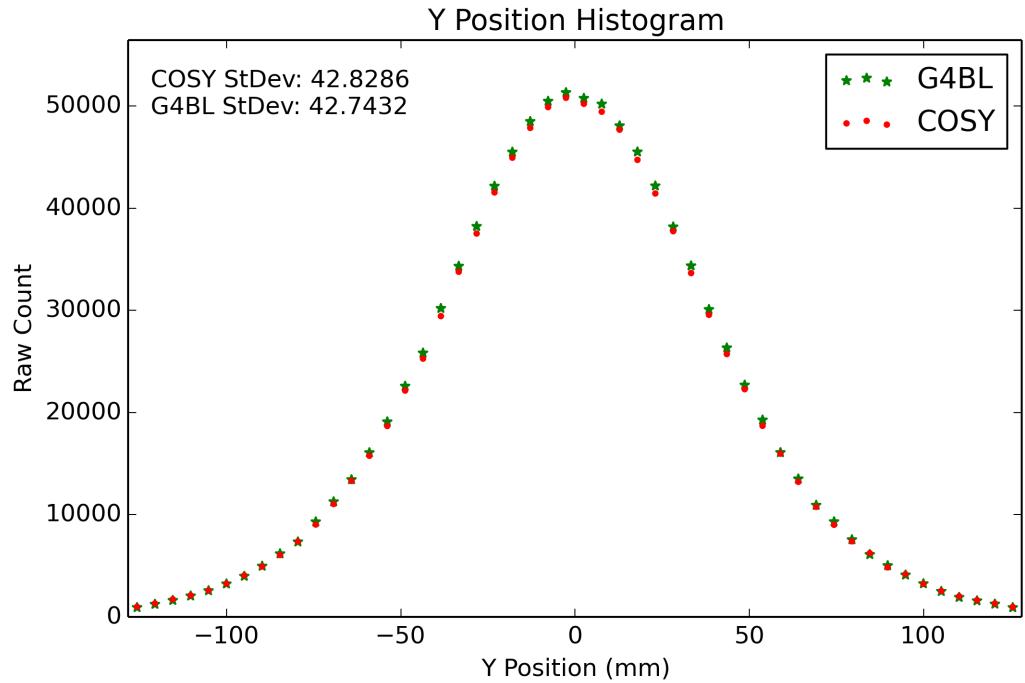


Figure 4.6: MICE Step IV y position results.

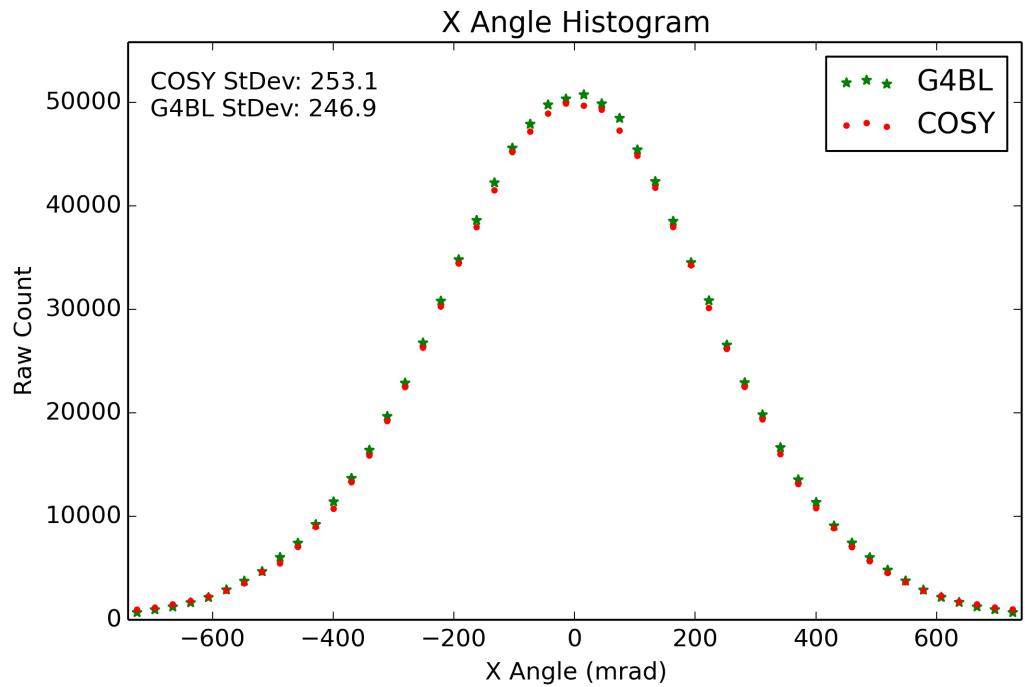


Figure 4.7: MICE Step IV x angle results.

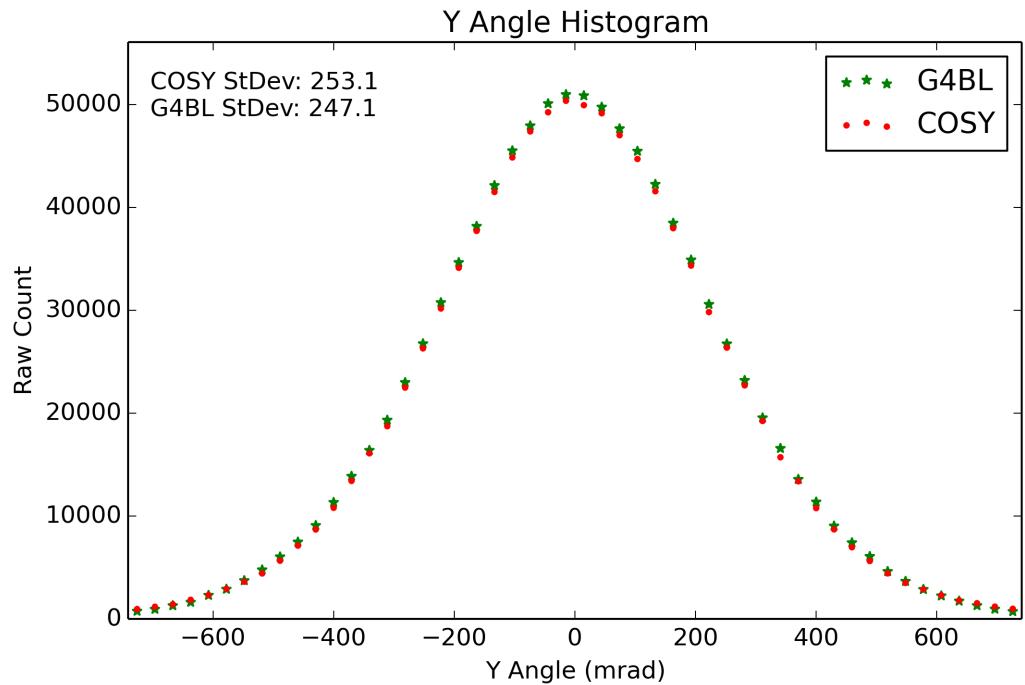


Figure 4.8: MICE Step IV y angle results.

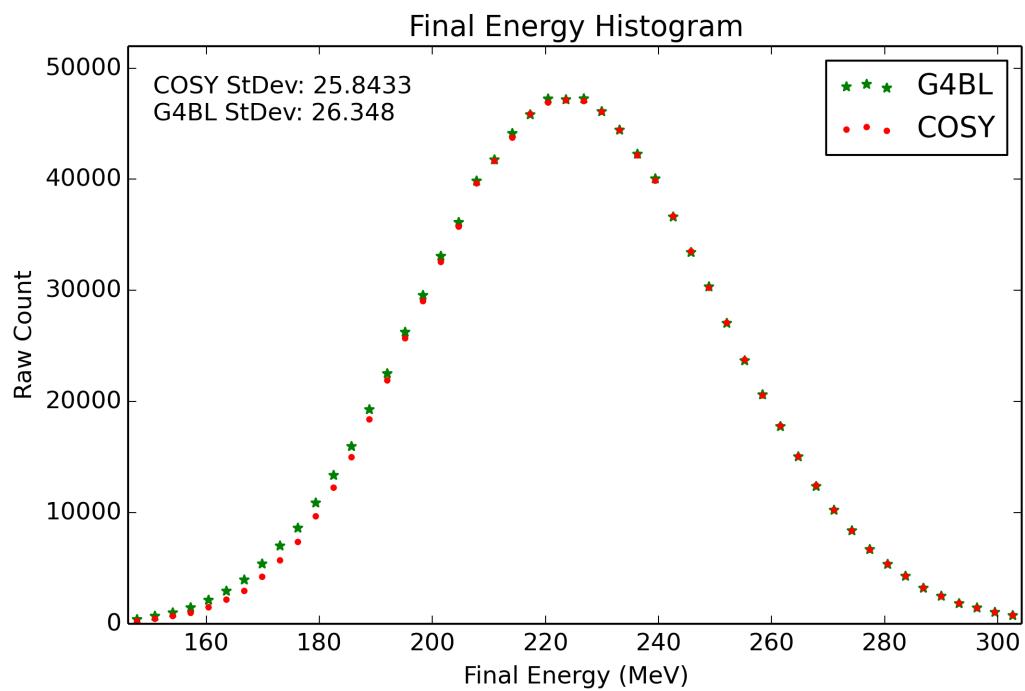


Figure 4.9: MICE Step IV final energy results.

Run times for the MICE Step IV simulation

Number of particles:	10^6	10^5	10^4	10^3	10^2
G4Beamline:	154 min 23 s	9 min 1 s	56 s	6 s	2 s
COSY:	25 min 58 s	51 s	10 s	6 s	6 s

Table 4.3: Run Times for the MICE Step IV Simulation. Note that the G4Beamline initialization time was not added to the run time values.

The runtimes of COSY and G4Beamline are listed in Table 4.3. To reiterate, COSY was run at 5th order with 32 steps before the absorber, 350 steps inside the absorber, and 32 steps after the absorber. G4Beamline was left to its default step size. Further note that the initialization time for G4Beamline to create the field maps was 33 seconds. However, since G4Beamline only has to create the field map once the initialization time is not added to the run times in Table 4.3.

4.4 Code Implementation

This section discusses in detail the organization and internal structure of the code. For reference, a reproduction of the code itself may be found in Appendix G. First, the user input will be discussed. Next, the `cosy.fox` level of the code will be examined. Finally, the structure of the FORTRAN code will briefly be considered.

Figure 4.10 will be referenced during these discussions.

In addition to the user's input file, COSY also uses a `cosy.fox` file. This file contains a plethora of global variables, functions, and routines which the user can access. However, `cosy.fox` can also be used to hide most of the complicated machinery from the user. For example, for this study the routines in `cosy.fox` transform the

COSY coordinates (x, a, y, b, ℓ, d) into absolute coordinates (x, p_x, y, p_y, t, E) ³, and then relays the information to the FORTRAN code. Both the user file and the `cosy.fox` file are written in COSYScript, the programming language of COSY.

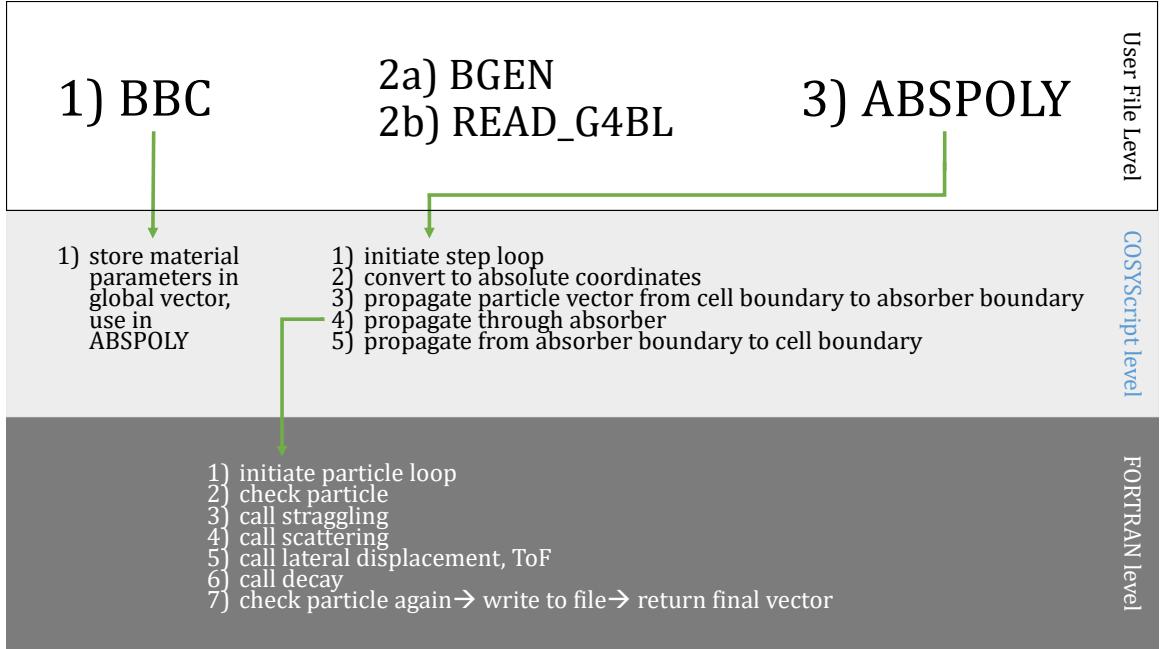


Figure 4.10: A flowchart for the structure of the new COSY routines implemented in this work.

4.4.1 User Input. As with the deterministic absorber routine previously present in COSY, `WA`, the user must first call the procedure

`BBC <Z> <A> <ρ> <I> <δ> <C> ;`

This stores the material parameters in a global array. The arguments are the nuclear charge Z , the atomic mass A , the density of the material ρ , the ionization energy I , the density correction δ , and the shell correction C . Next, the user must either generate a distribution of particles or read a distribution of particles from a file. Using

`BGEN <n> <V> <μx> <σx> <μpx> <σpx> <μy> <σy> <μpy> <σpy> <μt> <σt>`

³Recall from Eqn. 1.1 that x and y are the transverse coordinates; $a = p_x/p_0$; $b = p_y/p_0$; $\ell = -(t - t_0)v_0\gamma/(1 + \gamma)$; $\delta = (K - K_0)/K_0$; E is the total energy; K is the kinetic energy; and the subscript 0 denote the coordinate of the reference particle.

$\langle\mu_{p_z}\rangle \langle\sigma_{p_z}\rangle$;

the user can generate a Gaussian beam of n particles into a 2D vector V . Alternatively, the user may use

`READ_G4BL <file> <n> <V>` ;

to read a G4Beamline-formatted file of n particles and store it into a vector V . Finally, the user can call

`ABSPOLY <S1> <S2> <n> <L> <A> <V> <X0> <Sn> <Lc> <O>` ;

where S_1 is an n^{th} order polynomial describing the entrance surface, S_2 is an n^{th} order polynomial describing the exit surface, L is the on-axis length of the absorber, A is the aperture, V is the 2D input and output particle vector, X_0 is the radiation length of the material, S_n is the number of steps inside the absorber, L_c is the length of the absorber cell, and O is the output save number (e.g. “12” to save the results in `fort.12`). A depiction of the parameters can be found in Figure 4.11.

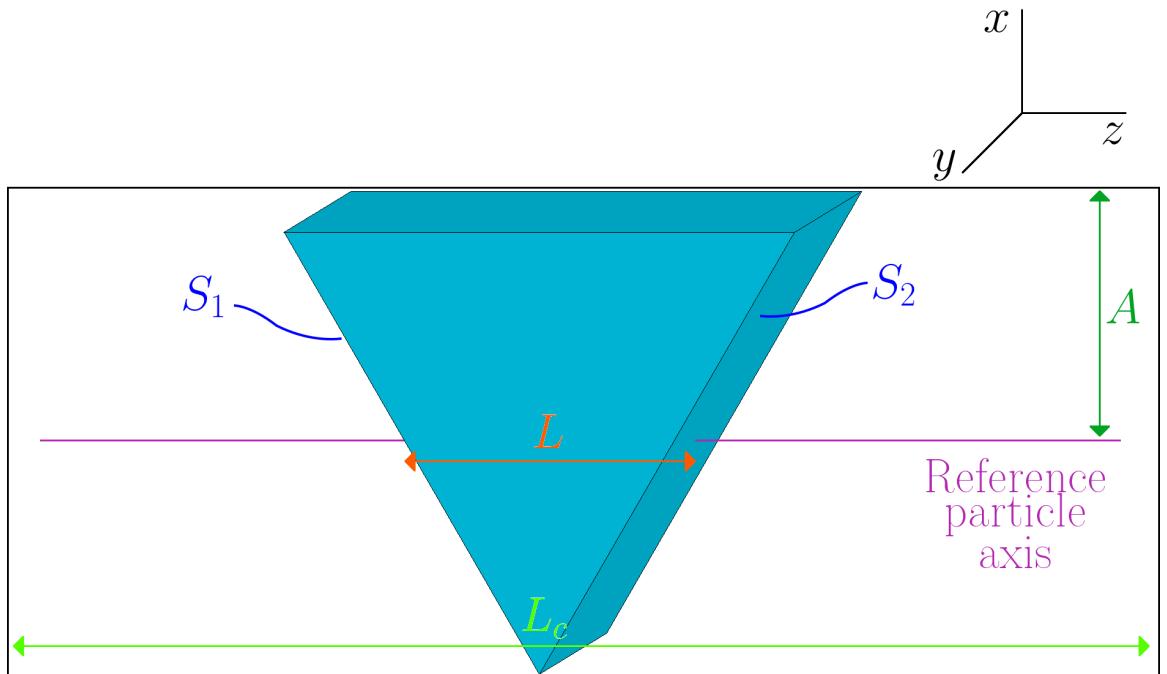


Figure 4.11: Cartoon example of some of the `ABSPOLY` parameters.

4.4.2 COSYScript Level.

As previously mentioned, the **ABSPOLY** routine is defined in the external file **cosy.fox**. Here, the step loop over S_n is initialized. Next, the coordinates of the input 2D vector V are converted from coordinates relative to the reference particle to absolute coordinates. The particle vector is then propagated from the cell boundary (whose width is defined by L_c) to the absorber (whose on-axis width is defined by L). If $L_c < L$ then no propagation occurs. When the particles are at the absorber boundary, the 2D particle vector V is passed on to the FORTRAN level of the code, where propagation through the absorber takes place. The routine linking the COSYScript and FORTRAN levels is **STOABS** and will be discussed in Section 4.4.3. After the 2D particle vector V is returned, V is propagated to the end of the cell and returned to the user.

4.4.3 FORTRAN Level. Except for the propagation of particles through the absorber, the processes found in Section 4.4.2 are much faster in COSYScript than in FORTRAN. This is because processes like coordinate conversion, propagation through a vacuum, etc. are deterministic and can therefore be handled by transfer maps.

The routine

STOABS $\langle V \rangle \langle m \rangle \langle L \rangle \langle MP \rangle \langle X_0 \rangle \langle O \rangle \langle A \rangle \langle n \rangle$;

takes input parameters from the COSYScript level and returns the input 2D particle vector V . Here, m is the particle mass (taken from the reference particle), L is an n -dimensional array containing the path lengths for each particle, MP is an array containing the material parameters (taken from the global array *BETHEBLOCHC* set up by the routine **BBC**), X_0 is the radiation length of the material, O is the output save number, A is the aperture, and n is the number of particles.

Once the FORTRAN level has this information, a loop over n particles is started. Each particle is checked at the beginning of each loop for the following flags: particle stopped, particle hit aperture, particle missed absorber, drift. If either the

particle was stopped ($E \leq m$) or the particle hit the aperture ($x^2 + y^2 >= A^2$) then the particle is terminated. If the particle missed the absorber completely then the drift flag is turned on. If the drift flag is on then the particle is simply propagated without the stochastic processes called.

Provided that the particle has not been flagged, the particle is subject to energy straggling and then multiple scattering. After these two routines, the lateral displacement and time-of-flight corrections are called. It should be noted that the order matters for energy straggling and multiple scattering. The order does not matter for lateral displacement and time-of-flight provided that both straggling and scattering have already occurred. Finally, decay can be called. Note that for this work, decay has been turned off completely, and so this step is skipped.

Finally, the particle is checked again. If the output save number O is not equal to zero then the particles are written to a file. Lastly, the 2D particle vector V is returned to the COSYScript level.

4.5 Summary

In summary, new simulation tools for muon ionization cooling have been added to COSY Infinity for particle-by-particle propagation. The energy straggling, multiple scattering, transverse position, and time-of-flight models were developed from first principles. The algorithms implemented have been modified via empirical fit to MuScat [32] while keeping reasonable agreement with G4Beamline [19]. Fitted with this new software, COSY has simulated one of the current muon ionization cooling efforts, MICE Step IV [34], yielding excellent agreement with G4Beamline. The code developed in this work is accurate, fast, and user-friendly.

APPENDIX A
DERIVATION OF TRANSVERSE EMITTANCE

The following is a derivation of Eqn. 1.2. First, it is assumed that the beam is an on-axis Gaussian beam. Explicitly, this is to say that the distribution of transverse coordinates are Gaussian and that $(\langle x \rangle, \langle \theta \rangle) = (0, 0)$. Recall that emittance is a measure of phase space volume. To measure this volume, it will be necessary to use a bivariate Gaussian distribution.

In general, a multivariate Gaussian distribution has the form

$$f_n(X) = \frac{1}{(2\pi)^{n/2}|\mathcal{A}|^{1/2}} \exp\left(-\frac{1}{2}(X - \mu)^T \mathcal{A}^{-1}(X - \mu)\right),$$

where $X = (x_1, \dots, x_n)$ is a column vector of independent variables, $\mathcal{A} \in S_{++}^n$ is the covariance matrix, and $\mu = (\mu_1, \dots, \mu_n)$ is a column vector of average values (which are all zero for this derivation) [35]. S_{++}^n is the space of symmetric positive definite $n \times n$ matrices. The covariance matrix \mathcal{A} is a matrix that describes the covariance between each variable. That is, \mathcal{A} describes how much two given variables change with each other. =

As an example, consider the multivariate Gaussian distribution of order 1. The only variable is x and its scale of change with itself is $\langle x \cdot x \rangle = \sigma^2$. Then

$$X = x \quad \mathcal{A} = \sigma^2$$

and

$$\begin{aligned} f_1(X) &= \frac{1}{(2\pi)^{1/2}\sigma} \exp\left(-\frac{1}{2}x \frac{1}{\sigma^2} x\right) \\ &= \frac{1}{\sqrt{2\pi}\sigma} \exp\left(-\frac{x^2}{2\sigma^2}\right). \end{aligned}$$

For transverse emittance, there are two variables, x and θ . They change with themselves as their respective σ^2 and with one another as $\langle x\theta \rangle$, and so

$$X = (x, \theta) \quad \mathcal{A} = \begin{pmatrix} \sigma_x^2 & \langle x\theta \rangle \\ \langle x\theta \rangle & \sigma_\theta^2 \end{pmatrix}.$$

The distribution is

$$f_2(X) = \frac{1}{2\pi\sqrt{\sigma_x^2\sigma_\theta^2 - \langle x\theta \rangle^2}} \exp\left(-\frac{x^2\sigma_\theta^2 - 2x\theta\langle x\theta \rangle + \theta^2\sigma_x^2}{2(\sigma_x^2\sigma_\theta^2 - \langle x\theta \rangle^2)}\right).$$

Since the boundary of this function is undefined, it is conventional to use a metric called the Mahalanobis distance [36]. This is a measure of how far some point is away from the center of a multivariate distribution. For multivariate Gaussian distributions, it is the negative of the exponential argument, or generally

$$\text{distance} = \frac{1}{2}(X - \mu)^T \mathcal{A}^{-1}(X - \mu).$$

For this particular case, the boundary of interest is all x and θ combinations which are unitary distance from the central peak. Then

$$1 = \frac{x^2\sigma_\theta^2 - 2x\theta\langle x\theta \rangle + \theta^2\sigma_x^2}{2(\sigma_x^2\sigma_\theta^2 - \langle x\theta \rangle^2)}.$$

Now it is apparent that this equation represents an ellipse in $x - \theta$ phase space:

$$1 = Ax^2 + 2Bx\theta + C\theta^2, \quad (\text{A.1})$$

with

$$A = \sigma_\theta^2/D$$

$$B = -\langle x\theta \rangle / D$$

$$C = \sigma_x^2/D$$

$$D = 2(\sigma_x^2\sigma_\theta^2 - \langle x\theta \rangle^2).$$

Indeed, an ellipse is precisely what one would expect. Observe from Figure A.1 that the projection of equal heights from a bivariate Gaussian onto the $x - \theta$ plane is an ellipse. This ellipse can get bigger or smaller depending on the Mahalanobis distance chosen, with the extrema being 0 and ∞ .

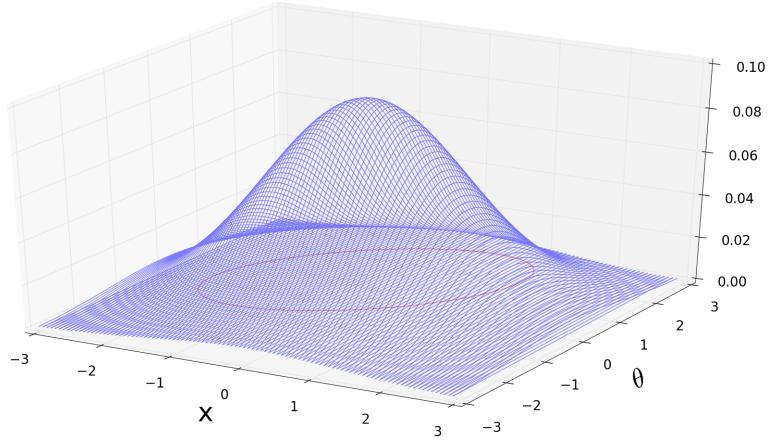


Figure A.1: Sample scatterplot of a bivariate Gaussian (blue). The ellipse (red) is a projection of points of equal heights onto the $x - \theta$ plane.

The area of this ellipse is the emittance (the volume of phase space) that is desired. For an untilted ellipse, this area is equal to π times the length of the major axis (a) times the length of the minor axis (b), or $\text{Area} = \pi ab$. The goal, then, is to apply a transformation of coordinates to the (potentially) tilted ellipse. This transformation to a new coordinate system (x', θ') is a simple rotation such that the major and minor axes of the ellipse are coincident with the x' and θ' axes. Then it will be much easier to find the major and minor axis lengths.

Note that Eqn. A.1 can also be represented by

$$1 = \begin{pmatrix} x & \theta \end{pmatrix} \mathcal{B} \begin{pmatrix} x \\ \theta \end{pmatrix}$$

where

$$\mathcal{B} = \begin{pmatrix} A & B \\ B & C \end{pmatrix}.$$

This is a useful representation, since the “normal” form (with the ellipse aligned to the x' and θ' axes) is desired:

$$1 = A'x'^2 + C'\theta'^2, \quad (\text{A.2})$$

or equivalently

$$1 = \begin{pmatrix} x' & \theta' \end{pmatrix} \mathcal{B}' \begin{pmatrix} x' \\ \theta' \end{pmatrix}$$

with

$$\mathcal{B}' = \begin{pmatrix} A' & 0 \\ 0 & C' \end{pmatrix}.$$

The goal, then, is to diagonalize \mathcal{B} . The diagonalization may be done via finding the eigenvalues and eigenvectors or by introducing some angle that relates x' and θ' to x and θ . Regardless, the result is

$$\begin{aligned} A' &= \frac{1}{2} \left(A + C + \sqrt{A^2 - 2AC + C^2 + 4B^2} \right) \\ C' &= \frac{1}{2} \left(A + C - \sqrt{A^2 - 2AC + C^2 + 4B^2} \right). \end{aligned}$$

In accordance with Eqn. A.2, the major and minor axis lengths are given by $1/2\sqrt{A'}$ and $1/2\sqrt{C'}^4$. Then the area of the ellipse in question is

$$\text{Area} = \pi \frac{1}{2\sqrt{A'}} \frac{1}{2\sqrt{C'}} = \pi \sqrt{\sigma_x^2 \sigma_\theta^2 - \langle x\theta \rangle^2}.$$

For the emittance, the π term is usually absorbed into the units (e.g. millimeter pi radians), and so the emittance is simply

$$\epsilon = \sqrt{\sigma_x^2 \sigma_\theta^2 - \langle x\theta \rangle^2}. \quad (\text{A.3})$$

⁴It may help to recall that the normal form equation for an ellipse may also be represented by $\frac{x^2}{a^2} + \frac{y^2}{b^2} = 1$, where a and b are the major or minor axis half-lengths.

APPENDIX B
A BRIEF REVIEW OF PARTICLE PHYSICS

Since this dissertation concerns only muons interacting with electrons, this brief review will only cover particles with spin 1/2. This review follows [1]. Particles of four-momentum $P = (E, \vec{p}) = (E, p_x, p_y, p_z)$ must satisfy the relativistic energy-momentum relation:

$$P^\alpha P_\alpha - m^2 = 0, \quad (\text{B.1})$$

since

$$\begin{aligned} P^\alpha P_\alpha - m^2 &= E^2 - p^2 - m^2 \\ &= E^2 - (p^2 + m^2) \\ &= E^2 - E^2 = 0. \end{aligned}$$

For particles at rest (i.e. $\vec{p} = 0$), this can be written as:

$$P^\alpha P_\alpha - m^2 = (p^0 + m)(p^0 - m) = 0,$$

and so clearly the energy $p^0 = E$ has to be either the rest mass or the negative of the rest mass (for antiparticles). However, for particles which are not at rest, if this same form is desired then it should look something like

$$P^\alpha P_\alpha - m^2 = (\gamma^\beta P_\beta + m)(\gamma^\delta P_\delta - m).$$

One may solve this equation for the coefficients $\gamma^\beta = (\gamma^0, \gamma^1, \gamma^2, \gamma^3)$, but no scalars solve this complex system. Dirac proposed that the various γ s were actually matrices, not scalars. This approach has the solution:

$$\gamma^0 = \begin{pmatrix} I_2 & 0 \\ 0 & -I_2 \end{pmatrix} \quad \gamma^j = \begin{pmatrix} 0 & \sigma^j \\ -\sigma^j & 0 \end{pmatrix},$$

where σ^j are the Pauli matrices. This leads to the Dirac equation for particles (for antiparticles, simply switch the sign of the mass):

$$\gamma^\alpha P_\alpha - m = 0. \quad (\text{B.2})$$

These particles can be represented by the wavefunctions for free particles,

$$\psi(x) = C e^{-(i/\hbar)p \cdot x} U^{(s)}(P),$$

where C is some amplitude, U is the spinor, and s is the spin state of the particle (in this case $s = 1, 2$). Antiparticle spinors are represented by V (unused in this work). Spinor adjoints are defined by

$$\bar{U} = U^\dagger \gamma^0 = U^{*T} \gamma^0, \quad (\text{B.3})$$

where $*$ represents the complex conjugate and T the transpose. These spinors satisfy the Diracequation:

$$(\gamma^\alpha P_\alpha - m)U = 0,$$

they are orthogonal:

$$\bar{U}^{(1)} U^{(2)} = 0,$$

and they are normalized:

$$\bar{U} U = 2m,$$

and they are complete:

$$\sum_{s=1}^2 U^{(s)} \bar{U}^{(s)} = (\gamma^\alpha P_\alpha + m). \quad (\text{B.4})$$

These spinors are four-component column vectors, and spinors describing particles (U) and antiparticles (V) for spin 1/2 particles could be represented as, for instance

$$U^{(1)} = \begin{pmatrix} 1 \\ 0 \\ 0 \\ 0 \end{pmatrix} \quad U^{(2)} = \begin{pmatrix} 0 \\ 1 \\ 0 \\ 0 \end{pmatrix}$$

$$V^{(1)} = \begin{pmatrix} 0 \\ 0 \\ 1 \\ 0 \end{pmatrix} \quad V^{(2)} = \begin{pmatrix} 0 \\ 0 \\ 0 \\ 1 \end{pmatrix},$$

but are usually more complicated if unpolarized.

APPENDIX C
DERIVATION OF IMPLEMENTED SCATTERING CROSS SECTION

Electron-muon scattering is a textbook scattering problem following Feynman rules [1]. Similar to both Sec. 2.1 and Sec. 2.2, either the collision time of the muon and electron is assumed to be very small compared to the electron orbital time or the electron is initially at rest. The longitudinal (z) axis is aligned with the muon velocity such that $v_x = v_y = 0$. The reference frame is the lab frame. Please refer to the information in Appendix B for a review on symbols and methods.

To find the (differential) scattering cross section $d\sigma/d\Omega$, it is necessary to have both the scattering amplitude \mathcal{M} (sometimes also called the ‘matrix element’, although this is a little ambiguous) and the available phase space over which to integrate. However, [1] notes that the integration over the available phase space is a constant, and so only the scattering amplitude is of interest.

Recall that the desired cross section is the Rutherford-like tail of a Gaussian–Rutherford-like piecewise function. The multiplicative constants will be ignored since only the forms of the individual Gaussian and Rutherford-like functions are desired. Then

$$\frac{d\sigma}{d\Omega} \propto |\mathcal{M}|^2.$$

To find the scattering amplitude \mathcal{M} , observe that Figure C.1 represents a muon-electron interaction via exchange of a virtual photon from the electron vertex α to the muon vertex β . The fermion flow is the path which goes along the directions of the arrows. For example, the first fermion flow which is observed is the muon flow: a muon enters from the left, there is an interaction at the propagator vertex β , and a muon exits from the right. Each muon contributes its spinor. The outgoing muon and incoming muon are adjoints of one another (i.e. if the spinor of P_2 is $U(P_2)$ then the spinor of P_4 is $\bar{U}(P_4)$). The propagator vertex contributes a factor of $-ieQ_\mu\gamma^\beta$. Since matrices compound right-to-left, it is necessary to start at the end of the fermion flow

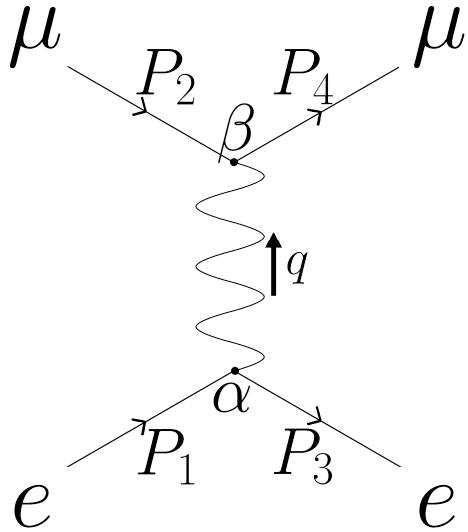


Figure C.1: Feynman diagram for electron-muon scattering. P_1 and P_2 are the incoming four-momenta for the electron and muon and P_3 and P_4 are the outgoing four-momenta. q represents the four-momentum carried by the virtual photon.

and work backwards. Figure C.2 helps to show that the scattering amplitude is

$$i\mathcal{M} = \bar{U}_4(-ieQ_\mu\gamma^\beta)U_2 \cdot \frac{i}{q^2}(-\eta_{\alpha\beta} + \frac{q_\alpha q_\beta}{q^2}) \cdot \bar{U}_3(-ieQ_e\gamma^\alpha)U_1, \quad (\text{C.1})$$

where U_i is the spinor for the i^{th} particle, q is the momentum carried by the virtual photon, γ^i is the γ matrix corresponding to vertex i in Figure C.1, α, β denote the vertices in Figure C.1, and $\eta_{i,j}$ is the $(i,j)^{th}$ component of the Minkowski metric described in Section E.4.

Ignoring the multiplicative constants, Eqn. C.1 can be rearranged to

$$\mathcal{M} \propto \bar{U}_4\gamma^\beta U_2 * \frac{1}{q^2}(-\eta_{\alpha\beta} + \frac{q_\alpha q_\beta}{q^2}) * \bar{U}_3\gamma^\alpha U_1.$$

Note that the second term in the parenthesis (i.e. $q_\alpha q_\beta/q^2$) vanishes. To see this, observe that when the expression is multiplied out the second term $q_\alpha q_\beta/q^2$ becomes proportional to $\bar{U}_4\gamma^\beta U_2 q_\beta$. By conservation of P_β at vertex β , it can be seen that

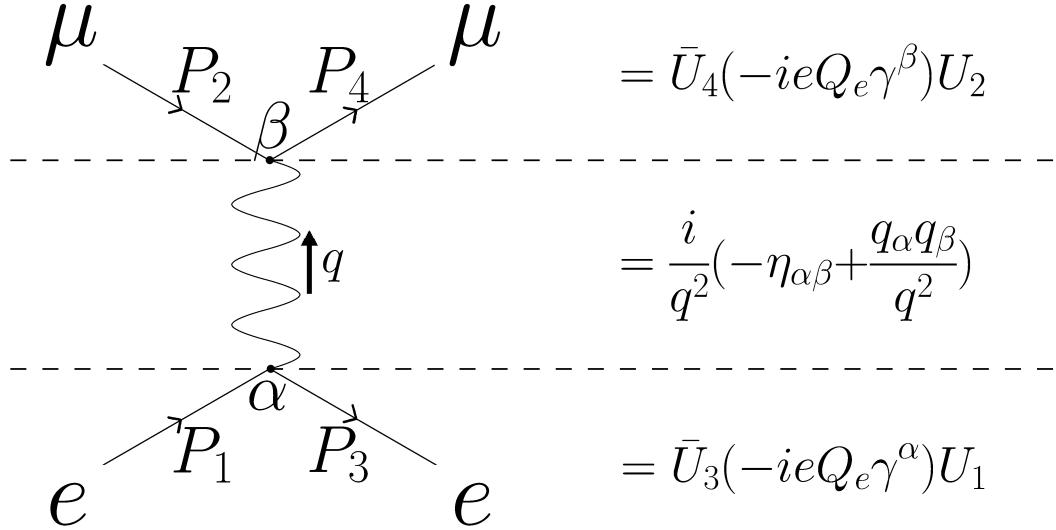


Figure C.2: Mathematical interpretation the three branches of the Feynman diagram in Figure C.1. The top part is the muon flow, the middle is the virtual photon, and the bottom is the electron flow.

$P_{2\beta} + q_\beta = P_{4\beta}$, or more usefully $q_\beta = P_{4\beta} - P_{2\beta}$. Then

$$\begin{aligned} \bar{U}_4\gamma^\beta U_2 q_\beta &= \bar{U}_4(\gamma^\beta(P_{4\beta} - P_{2\beta}))U_2 \\ &= \bar{U}_4(\gamma^\beta P_{4\beta} - \gamma^\beta P_{2\beta})U_2. \end{aligned}$$

Eqn. B.2 states that $\gamma^\alpha P_\alpha - m = 0$, and so

$$\bar{U}_4\gamma^\beta U_2 q_\beta = \bar{U}_4(m_\mu - m_\mu)U_2 = 0.$$

For this reason, the second term in parenthesis vanishes entirely, leaving

$$\mathcal{M} \propto \bar{U}_4\gamma^\beta U_2 * \frac{\eta_{\alpha\beta}}{q^2} * \bar{U}_3\gamma^\alpha U_1.$$

Propagating the $\eta_{\alpha\beta}$ through,

$$\mathcal{M} \propto \bar{U}_4\gamma^\beta U_2 * \frac{1}{q^2} * \bar{U}_3\gamma_\beta U_1.$$

The final cross section is proportional to $|\mathcal{M}|^2$, and so

$$|\mathcal{M}|^2 \propto \frac{1}{q^4} * \bar{U}_4\gamma^\beta U_2 [\bar{U}_4\gamma^\delta U_2]^* * \bar{U}_3\gamma_\beta U_1 [\bar{U}_3\gamma_\delta U_1]^*.$$

Since $\bar{U}_4\gamma^\beta U_2$ is the same as $\bar{U}_3\gamma_\beta U_1$ except for notation, from here the quantity $\bar{U}_4\gamma^\beta U_2[\bar{U}_4\gamma^\delta U_2]^*$ will be reduced and the same treatment will be applied to $\bar{U}_3\gamma_\beta U_1[\bar{U}_3\gamma_\delta U_1]^*$.

Since $\bar{U}_4\gamma^\delta U_2 \in \mathbb{C}$ in Dirac space,

$$\begin{aligned} [\bar{U}_4\gamma^\delta U_2]^* &= [\bar{U}_4\gamma^\delta U_2]^\dagger \\ &= U_2^\dagger \gamma^{\delta\dagger} \bar{U}_4^\dagger. \end{aligned}$$

Observe that since γ^δ has the form

$$\gamma^\delta = \begin{pmatrix} A & B \\ -B & -A \end{pmatrix}$$

then

$$\begin{aligned} \gamma^0 \gamma^\delta \gamma^0 &= \begin{pmatrix} I_2 & 0 \\ 0 & I_2 \end{pmatrix} \begin{pmatrix} A & B \\ -B & -A \end{pmatrix} \begin{pmatrix} I_2 & 0 \\ 0 & I_2 \end{pmatrix} \\ &= \begin{pmatrix} A & -B \\ B & -A \end{pmatrix} \\ &= \gamma^{\delta\dagger}, \end{aligned}$$

and so

$$[\bar{U}_4\gamma^\delta U_2]^* = U_2^\dagger [\gamma^0 \gamma^\delta \gamma^0] \bar{U}_4^\dagger.$$

Recall the definition of the spinor adjoint in Eqn. B.3:

$$\bar{U} = U^\dagger \gamma^0.$$

Then

$$\begin{aligned} \bar{U}_4^\dagger &= (U_4^\dagger \gamma^0)^\dagger \\ &= \gamma^{0\dagger} U_4^{\dagger\dagger} \\ &= \gamma^0 U_4, \end{aligned}$$

which results in

$$[\bar{U}_4 \gamma^\delta U_2]^* = U_2^\dagger [\gamma^0 \gamma^\delta \gamma^0] [\gamma^0 U_4],$$

or more suggestively,

$$\begin{aligned} [\bar{U}_4 \gamma^\delta U_2]^* &= [U_2^\dagger \gamma^0] \gamma^\delta [\gamma^0 \gamma^0] U_4 \\ &= \bar{U}_2 \gamma^\delta U_4. \end{aligned}$$

This results in

$$|\mathcal{M}|^2 \propto \frac{1}{q^4} * \bar{U}_4 \gamma^\beta U_2 \bar{U}_2 \gamma^\delta U_4 * \bar{U}_3 \gamma_\beta U_1 \bar{U}_1 \gamma_\delta U_3. \quad (\text{C.2})$$

In principle, if the spinors of the muons were known (e.g. the muons were part of a polarized beam) this could be evaluated directly. However, in general the spinors are not known. As an approximation, it is necessary to average the spinors over their spins. Again, starting with the muons and extrapolating the treatment to the electrons,

$$\begin{aligned} \langle \bar{U}_4 \gamma^\beta U_2 \bar{U}_2 \gamma^\delta U_4 \rangle_{\text{spin } 2} &= \frac{1}{2} \sum_{s_2=1}^2 \bar{U}_4 \gamma^\beta U_2^{(s_2)} \bar{U}_2^{(s_2)} \gamma^\delta U_4 \\ &= \frac{1}{2} \bar{U}_4 \gamma^\beta \left[\sum_{s_2=1}^2 U_2^{(s_2)} \bar{U}_2^{(s_2)} \right] \gamma^\delta U_4. \end{aligned}$$

Recall the completeness property of spinors (Eqn. B.4):

$$\sum_s U^{(s)} \bar{U}^{(s)} = \gamma^\alpha P_\alpha + m.$$

Then

$$\langle \bar{U}_4 \gamma^\beta U_2 \bar{U}_2 \gamma^\delta U_4 \rangle_{\text{spin } 2} = \frac{1}{2} \bar{U}_4 \gamma^\beta (\gamma^\epsilon P_{2\epsilon} + m_\mu) \gamma^\delta U_4.$$

Averaging similarly over the spin for particle #4 yields

$$\langle \bar{U}_4 \gamma^\beta U_2 \bar{U}_2 \gamma^\delta U_4 \rangle_{\text{spin } 2, \text{ spin } 4} = \frac{1}{2} \frac{1}{2} \sum_{s_4=1}^2 \bar{U}_4^{(s_4)} [\gamma^\beta (\gamma^\epsilon P_{2\epsilon} + m_\mu) \gamma^\delta] U_4^{(s_4)}.$$

Now, spinors do not generally commute, but their components (which are scalars) do.

Observing that \bar{U} is a 1x4 row vector, the bracketed term is a 4x4 matrix, and U is a 4x1 column vector, it is possible to write out the matrix multiplication as an explicit double sum over the indices i, j :

$$\begin{aligned}\langle \bar{U}_4 \gamma^\beta U_2 \bar{U}_2 \gamma^\delta U_4 \rangle_{s_2, s_4} &= \frac{1}{4} \sum_{s_4=1}^2 \sum_{i,j=1}^4 \bar{U}_{4i}^{(s_4)} [\gamma^\beta (\gamma^\epsilon P_{2\epsilon} + m_\mu) \gamma^\delta]_{ij} U_{4j}^{(s_4)} \\ &= \frac{1}{4} \sum_{i,j=1}^4 \left([\gamma^\beta (\gamma^\epsilon P_{2\epsilon} + m_\mu) \gamma^\delta]_{ij} \left[\sum_{s_4=1}^2 \bar{U}_4^{(s_4)} U_4^{(s_4)} \right]_{ij} \right)\end{aligned}$$

Spinor completeness (Eqn. B.4) requires a summation over $U\bar{U}$, not $\bar{U}U$. However, $\bar{U}U$ is still useful. Since $A^T + B^T = [A + B]^T$,

$$\begin{aligned}\sum \bar{U}U &= \left[\sum (\bar{U}U)^T \right]^T \\ &= \left[\sum U^T \bar{U}^T \right]^T.\end{aligned}$$

Again explicitly writing the matrix multiplication,

$$\begin{aligned}\sum \bar{U}U &= \left[\sum \left(\sum_{k,l}^4 (U_k)^T (\bar{U}_l)^T \right) \right]^T \\ &= \left[\sum_{k,l}^4 \left(\sum (U_k)^T (\bar{U}_l)^T \right) \right]^T \\ &= \left[\sum_{k,l}^4 \left(\sum U_l \bar{U}_k \right) \right]^T \\ &= \left[\sum_{k,l}^4 \left(\sum U \bar{U} \right)_{l,k} \right]^T.\end{aligned}$$

Now it is possible to use spinor completeness (Eqn. B.4):

$$\begin{aligned}\sum \bar{U}U &= \left[\sum_{k,l} (\gamma^\alpha P_{alp}ha + m)_{l,k} \right] \\ &= (\gamma^\alpha P_\alpha + m)^T.\end{aligned}$$

Conclusively, if the i - j^{th} component of $\sum \bar{U}U$ is desired, then the indices of Eqn. B.4 must be exchanged:

$$\begin{aligned}\left[\sum \bar{U}U \right]_{i,j} &= [(\gamma^\alpha P_\alpha + m)^T]_{i,j} \\ &= [(\gamma^\alpha P_\alpha + m)_{i,j}]^T \\ &= [\gamma^\alpha P_\alpha + m]_{j,i}.\end{aligned}$$

This results in

$$\langle \bar{U}_4 \gamma^\beta U_2 \bar{U}_2 \gamma^\delta U_4 \rangle_{s2, s4} = \frac{1}{4} \sum_{i,j=1}^4 \left([\gamma^\beta (\gamma^\epsilon P_{2\epsilon} + m_\mu) \gamma^\delta]_{ij} [\gamma^\kappa P_{4\kappa} + m_\mu]_{j,i} \right).$$

Upon the concatenation of the subscripts i, j and j, i , it can be seen that

$$\langle \bar{U}_4 \gamma^\beta U_2 \bar{U}_2 \gamma^\delta U_4 \rangle_{s2, s4} = \frac{1}{4} \sum_{i=1}^4 \left([\gamma^\beta (\gamma^\epsilon P_{2\epsilon} + m_\mu) \gamma^\delta] (\gamma^\kappa P_{4\kappa} + m_\mu) \right)_{ii}.$$

This is now the definition of a trace ($\sum_i M_{ii} \equiv \text{Tr}(M)$). Then

$$\langle \bar{U}_4 \gamma^\beta U_2 \bar{U}_2 \gamma^\delta U_4 \rangle_{s2, s4} = \frac{1}{4} \text{Tr}[\gamma^\beta (\gamma^\epsilon P_{2\epsilon} + m_\mu) \gamma^\delta (\gamma^\kappa P_{4\kappa} + m_\mu)].$$

Using the addition property of traces, it is clear that there are four terms to evaluate:

- i) $\text{Tr}(\gamma^\beta \gamma^\epsilon P_{2\epsilon} \gamma^\delta \gamma^\kappa P_{4\kappa})$
- ii) $\text{Tr}(\gamma^\beta \gamma^\epsilon P_{2\epsilon} \gamma^\delta m_\mu)$
- iii) $\text{Tr}(\gamma^\beta m_\mu \gamma^\delta \gamma^\kappa P_{4\kappa})$
- iv) $\text{Tr}(\gamma^\beta m_\mu \gamma^\delta m_\mu)$.

The derivations for the solutions to these traces can be found in Appendix E. The first term is the trace of four γ matrices and can be solved by using Eqn. E.9:

$$\begin{aligned}\text{Tr}(\gamma^\beta \gamma^\epsilon P_{2\epsilon} \gamma^\delta \gamma^\kappa P_{4\kappa}) &= P_{2\epsilon} P_{4\kappa} \text{Tr}(\gamma^\beta \gamma^\epsilon \gamma^\delta \gamma^\kappa) \\ &= 4P_{2\epsilon} P_{4\kappa} (\eta^{\beta\epsilon} \eta^{\delta\kappa} - \eta^{\beta\delta} \eta^{\epsilon\kappa} + \eta^{\beta\kappa} \eta^{\epsilon\delta}) \\ &= 4(P_2^\beta P_4^\delta - P_2 \cdot P_4 \eta^{\beta\delta} + P_2^\delta P_4^\beta).\end{aligned}$$

The second term is the trace of an odd number of γ matrices, and is equal to zero via Eqn. E.6:

$$\begin{aligned}\text{Tr}(\gamma^\beta \gamma^\epsilon P_{2\epsilon} \gamma^\delta m_\mu) &= P_{2\epsilon} m_\mu \text{Tr}(\gamma^\beta \gamma^\epsilon \gamma^\delta) \\ &= 0.\end{aligned}$$

The third term is also a trace of an odd number of γ matrices:

$$\begin{aligned}\text{Tr}(\gamma^\beta m_\mu \gamma^\delta \gamma^\kappa P_{4\kappa}) &= m_\mu P_{4\kappa} \text{Tr}(\gamma^\beta \gamma^\delta \gamma^\kappa) \\ &= 0.\end{aligned}$$

The final term is the trace of two γ matrices, and results in the Minkowski metric via Eqn. E.8:

$$\begin{aligned}\text{Tr}(\gamma^\beta m_\mu \gamma^\delta m_\mu) &= m_\mu^2 \text{Tr}(\gamma^\beta \gamma^\delta) \\ &= 4m_\mu^2 \eta^{\beta\delta}.\end{aligned}$$

Putting it all together results in

$$\langle \bar{U}_4 \gamma^\beta U_2 \bar{U}_2 \gamma^\delta U_4 \rangle_{s2, s4} = P_2^\beta P_4^\delta - P_2 \cdot P_4 \eta^{\beta\delta} + P_2^\delta P_4^\beta + m_\mu^2 \eta^{\beta\delta}.$$

The electron portion is mathematically the same. Symbolically, contravariant indices become covariant (e.g. γ^β becomes γ_β), the mass is now the electron mass (i.e. m_μ becomes m_e), and the even subscripts become odd (i.e. P_2 , P_4 become P_1 , P_3). Explicitly, Eqn. C.2 becomes

$$\begin{aligned}\langle |\mathcal{M}|^2 \rangle &\propto \frac{1}{q^4} * \langle \bar{U}_4 \gamma^\beta U_2 \bar{U}_2 \gamma^\delta U_4 \rangle_{s2, s4} * \langle \bar{U}_3 \gamma_\beta U_1 \bar{U}_1 \gamma_\delta U_4 \rangle_{s1, s3} \\ &\propto \frac{1}{q^4} * (P_2^\beta P_4^\delta - P_2 \cdot P_4 \eta^{\beta\delta} + P_2^\delta P_4^\beta + m_\mu^2 \eta^{\beta\delta}) * \\ &\quad (P_{1\beta} P_{3\delta} - P_1 \cdot P_3 \eta_{\beta\delta} + P_{1\delta} P_{3\beta} + m_e^2 \eta_{\beta\delta}).\end{aligned}$$

Explicitly, the algebra is

$$\begin{aligned} \langle |\mathcal{M}|^2 \rangle \propto \frac{1}{q^4} * [& (P_1 \cdot P_2)(P_3 \cdot P_4) - (P_1 \cdot P_3)(P_2 \cdot P_4) + (P_1 \cdot P_4)(P_2 \cdot P_3) + \\ & m_e^2(P_2 \cdot P_4) - (P_1 \cdot P_3)(P_2 \cdot P_4) + 4(P_1 \cdot P_3)(P_2 \cdot P_4) - \\ & (P_1 \cdot P_3)(P_2 \cdot P_4) - 4m_e^2(P_2 \cdot P_4) + (P_1 \cdot P_4)(P_2 \cdot P_3) - \\ & (P_1 \cdot P_3)(P_2 \cdot P_4) + (P_1 \cdot P_2)(P_3 \cdot P_4) + m_e^2(P_2 \cdot P_4) + \\ & m_\mu^2(P_1 \cdot P_3) - 4m_\mu^2(P_1 \cdot P_3) + m_\mu^2(P_1 \cdot P_3) + 4m_\mu^2 m_e^2], \end{aligned}$$

with $\eta^{\alpha\beta}\eta_{\alpha\beta} = 4$. Gathering the like terms, this reduces to

$$\begin{aligned} \langle |\mathcal{M}|^2 \rangle \propto \frac{1}{q^4} [& 2(P_1 \cdot P_4)(P_2 \cdot P_3) + 2(P_1 \cdot P_2)(P_3 \cdot P_4) - \\ & 2m_e^2(P_2 \cdot P_4) - 2m_\mu^2(P_1 \cdot P_3) + 4m_\mu^2 m_e^2]. \end{aligned}$$

Up until here, this is a quite general expression for two particles interacting via virtual photon exchange. For COSY, straggling and scattering are mutually exclusive processes: first the straggling routine is called and then the scattering routine is called. Therefore, for this model it must be assumed that there is no energy exchange between the muon and electron. Consequently, since the electron is bound it will remain fixed and the muon will scatter. This can be seen diagrammatically in Figure C.3.

Now it is clear that $P_1 = P_3 = (m_e, 0, 0, 0)$. Furthermore, since the total energy of the muon is conserved $E_2 = E_4 = E_\mu$ and $P_2 = (E_\mu, \vec{p}_2)$ and $P_4 = (E_\mu, \vec{p}_4)$ and so

$$\begin{aligned} P_1 \cdot P_2 &= E_\mu m_e & P_1 \cdot P_3 &= m_e^2 \\ P_1 \cdot P_4 &= E_\mu m_e & P_2 \cdot P_3 &= E_\mu m_e \\ P_2 \cdot P_4 &= E_\mu^2 - \vec{p}_2 \cdot \vec{p}_4 & &= E_\mu^2 - p_\mu^2 \cos \theta \\ P_3 \cdot P_4 &= E_\mu m_e. & & \end{aligned}$$

Then

$$\langle |\mathcal{M}|^2 \rangle \propto \frac{1}{q^4} [2(E_\mu m_e)(E_\mu m_e) + 2(E_\mu m_e)(E_\mu m_e) - 2m_e^2(E_\mu^2 - p_\mu^2 \cos \theta) - 2m_\mu^2 m_e^2 + 4m_\mu^2 m_e^2].$$

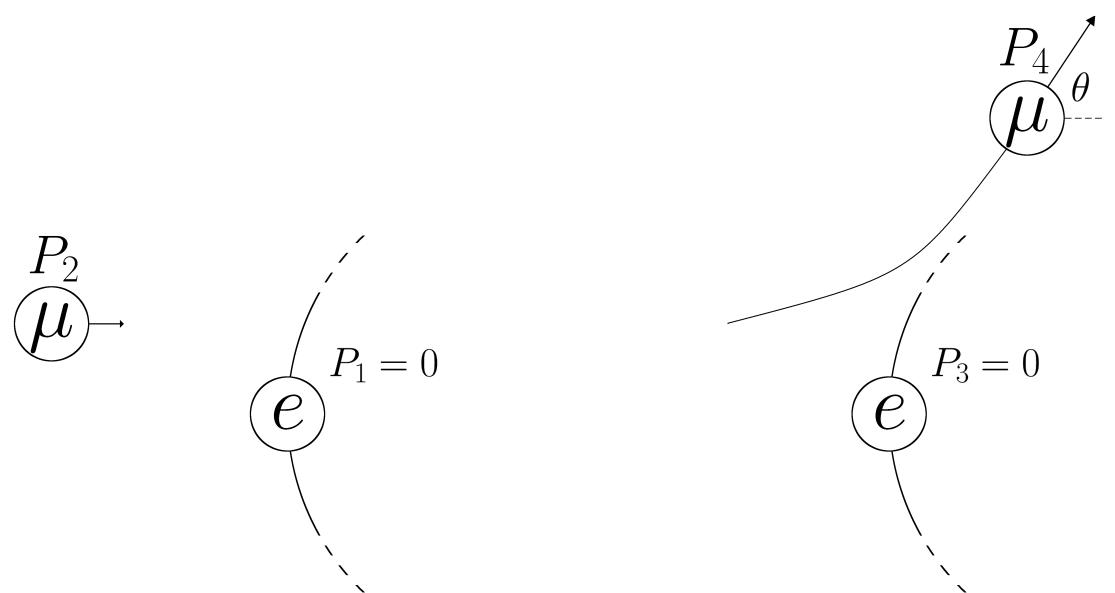


Figure C.3: COSY treatment of muon-electron scattering. Since the routines are called separately, the model assumes no straggling while scattering.

Factoring out the term $2m_e^2$,

$$\begin{aligned}\langle |\mathcal{M}|^2 \rangle &\propto \frac{1}{q^4} (E_\mu^2 + E_\mu^2 - E_\mu^2 + p_\mu^2 \cos \theta - m_\mu^2 + 2m_\mu^2) \\ &\propto \frac{1}{q^4} (E_\mu^2 + p_\mu \cos \theta + m_\mu^2) \\ &\propto \frac{1}{q^4} (p_\mu + m_\mu + p_\mu \cos \theta + m_\mu^2) \\ &\propto \frac{1}{q^4} (2m_\mu^2 + p_\mu(1 + \cos \theta)).\end{aligned}$$

Factoring out $2m_\mu^2$ and observing that $p/m = \beta\gamma$ yields

$$\langle |\mathcal{M}|^2 \rangle \propto \frac{1 + \frac{(\beta\gamma)^2}{2}(1 + \cos \theta)}{q^4}.$$

Again using conservation of P_β at vertex β , it can be seen that

$$P_{2\beta} + q_\beta = P_{4\beta} \quad \rightarrow \quad q_\beta = P_{4\beta} - P_{2\beta}$$

. Then

$$\begin{aligned}q^2 &= (P_{4\beta} - P_{2\beta})^2 \\ &= ((E_4, \vec{p}_4) - (E_2, \vec{p}_2))^2 \\ &= ((0, \vec{p}_4 - \vec{p}_2))^2 \\ &= -(p_4^2 + p_2^2 - p_4 p_2 \cos \theta) \\ &= -2p_\mu(1 - \cos \theta).\end{aligned}$$

Squaring q^2 and throwing out the constant p_μ ,

$$\langle |\mathcal{M}|^2 \rangle \propto \frac{1 + \frac{(\beta\gamma)^2}{2}(1 + \cos \theta)}{(1 - \cos \theta)^2}.$$

Finally, the Mott cross section is obtained:

$$\frac{d\sigma}{d\Omega} \propto \frac{1 + \frac{(\beta\gamma)^2}{2}(1 + \cos \theta)}{(1 - \cos \theta)^2}. \quad (\text{C.3})$$

Observe that for a non-relativistic beam of muons, $\beta\gamma \rightarrow 0$ and this reduces to the Rutherford cross section (Eqn. 2.29).

APPENDIX D
EXPLICIT FORMS OF THE DIRAC GAMMA MATRICES AND PAULI
MATRICES

From the conventions in [1] it is observed that the Pauli matrices are defined as:

$$\sigma^1 = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} \quad \sigma^2 = \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix} \quad \sigma^3 = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}.$$

Then the Dirac matrices follow as:

$$\gamma^0 = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix} \quad \gamma^1 = \begin{pmatrix} 0 & 0 & 0 & 1 \\ 0 & 0 & 1 & 0 \\ 0 & -1 & 0 & 0 \\ -1 & 0 & 0 & 0 \end{pmatrix} \quad (D.1)$$

$$\gamma^2 = \begin{pmatrix} 0 & 0 & 0 & -i \\ 0 & 0 & i & 0 \\ 0 & i & 0 & 0 \\ -i & 0 & 0 & 0 \end{pmatrix} \quad \gamma^3 = \begin{pmatrix} 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & -1 \\ -1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \end{pmatrix}.$$

APPENDIX E

PROOFS OF USEFUL DIRAC GAMMA MATRIX TRACE IDENTITIES

E.1 Proof of $(\gamma^5)^2 = I_4$

Let $\gamma^5 \equiv i\gamma^0\gamma^1\gamma^2\gamma^3$. Then explicitly carrying out the multiplication yields

$$\begin{aligned} \gamma^5 &= i \begin{pmatrix} 0 & 0 & 0 & 1 \\ 0 & 0 & 1 & 0 \\ 0 & 1 & 0 & 0 \\ 1 & 0 & 0 & 0 \end{pmatrix} \cdot -i \begin{pmatrix} 0 & 1 & 0 & 0 \\ 1 & 0 & 0 & 0 \\ 0 & 0 & 0 & 1 \\ 0 & 0 & 1 & 0 \end{pmatrix} \\ \gamma^5 &= \begin{pmatrix} 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \\ 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \end{pmatrix}. \end{aligned} \quad (\text{E.1})$$

Then

$$(\gamma^5)^2 = I_4. \quad (\text{E.2})$$

E.2 Proof of $\gamma^5\gamma^\alpha = -\gamma^\alpha\gamma^5$

Let the form of γ^α be represented as

$$\begin{pmatrix} A & B \\ -B & -A \end{pmatrix}$$

(i.e. for $\alpha = 0$, $A = I_2$ and $B = 0$; otherwise $A = 0$ and $B = \sigma^\alpha$). Then using the explicit form of γ^5 in Eqn. E.1:

$$\begin{pmatrix} 0 & I_2 \\ I_2 & 0 \end{pmatrix} \begin{pmatrix} A & B \\ -B & -A \end{pmatrix} + \begin{pmatrix} A & B \\ -B & -A \end{pmatrix} \begin{pmatrix} 0 & I_2 \\ I_2 & 0 \end{pmatrix} = \begin{pmatrix} 0 & 0 \\ 0 & 0 \end{pmatrix}.$$

Therefore,

$$\gamma^5 \gamma^\alpha = -\gamma^\alpha \gamma^5. \quad (\text{E.3})$$

E.3 Proof of $\text{Tr}(\text{odd number of } \gamma \text{ matrices}) = 0$

Let there be the trace of an arbitrary odd number of γ matrices

$$\text{Tr}(\gamma^{\alpha_1} \gamma^{\alpha_2} \dots \gamma^{\alpha_n}),$$

such that n is odd. Now insert into the beginning of the trace $\gamma^5 \gamma^5$,

$$\text{Tr}(\gamma^5 \gamma^5 \gamma^{\alpha_1} \gamma^{\alpha_2} \dots \gamma^{\alpha_n}),$$

and there are two cases. The first makes use of $(\gamma^5)^2 = I_4$ (Eqn. E.2), and results in

$$\begin{aligned} \text{Tr}(\gamma^5 \gamma^5 \gamma^{\alpha_1} \gamma^{\alpha_2} \dots \gamma^{\alpha_n}) &= \text{Tr}(I_4 \gamma^{\alpha_1} \gamma^{\alpha_2} \dots \gamma^{\alpha_n}) \\ \text{Tr}(\gamma^5 \gamma^5 \gamma^{\alpha_1} \gamma^{\alpha_2} \dots \gamma^{\alpha_n}) &= \text{Tr}(\gamma^{\alpha_1} \gamma^{\alpha_2} \dots \gamma^{\alpha_n}) \end{aligned} \quad (\text{E.4})$$

The next case uses the cyclic property of traces, which is $\text{Tr}(ABCD) = \text{Tr}(BCDA) = \text{Tr}(CDAB) = \dots$. Then

$$\text{Tr}(\gamma^5 \gamma^5 \gamma^{\alpha_1} \gamma^{\alpha_2} \dots \gamma^{\alpha_n}) = \text{Tr}(\gamma^5 \gamma^{\alpha_1} \gamma^{\alpha_2} \dots \gamma^{\alpha_n} \gamma^5).$$

Now using $\gamma^5 \gamma^\alpha = -\gamma^\alpha \gamma^5$ (Eqn. E.3):

$$\text{Tr}(\gamma^5 \gamma^5 \gamma^{\alpha_1} \gamma^{\alpha_2} \dots \gamma^{\alpha_n}) = \text{Tr}(\gamma^5 \gamma^{\alpha_1} \gamma^{\alpha_2} \dots \gamma^5 \gamma^{\alpha_n})(-1)^1.$$

Applying this n times yields

$$\text{Tr}(\gamma^5 \gamma^5 \gamma^{\alpha_1} \gamma^{\alpha_2} \dots \gamma^{\alpha_n}) = \text{Tr}(\gamma^5 \gamma^5 \gamma^{\alpha_1} \gamma^{\alpha_2} \dots \gamma^{\alpha_n})(-1)^n.$$

Since n is negative, $(-1)^n = -1$. Again using $(\gamma^5)^2 = I_4$ (Eqn. E.2), this yields

$$\text{Tr}(\gamma^5 \gamma^5 \gamma^{\alpha_1} \gamma^{\alpha_2} \dots \gamma^{\alpha_n}) = -\text{Tr}(\gamma^{\alpha_1} \gamma^{\alpha_2} \dots \gamma^{\alpha_n}). \quad (\text{E.5})$$

Combining Eqns. E.4 and E.5:

$$\begin{aligned}\text{Tr}(\gamma^5 \gamma^5 \gamma^{\alpha_1} \gamma^{\alpha_2} \dots \gamma^{\alpha_n}) &= \text{Tr}(\gamma^{\alpha_1} \gamma^{\alpha_2} \dots \gamma^{\alpha_n}) \\ \text{Tr}(\gamma^5 \gamma^5 \gamma^{\alpha_1} \gamma^{\alpha_2} \dots \gamma^{\alpha_n}) &= -\text{Tr}(\gamma^{\alpha_1} \gamma^{\alpha_2} \dots \gamma^{\alpha_n}).\end{aligned}$$

Since this trace is both positive and negative the only conclusion is that it must be zero:

$$\text{Tr}(\gamma^{\alpha_1} \gamma^{\alpha_2} \dots \gamma^{\alpha_n}) = 0 \quad \text{for odd } n. \quad (\text{E.6})$$

E.4 Proof of $\text{Tr}(\gamma^\alpha \gamma^\beta) = 4\eta^{\alpha\beta}$

$\eta^{\alpha\beta}$ is an element of the Minkowski metric, which is defined here as

$$\eta = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix}.$$

Unlike γ^α , $\eta^{\alpha\beta}$ is simply a number. For example, if $\alpha \neq \beta$ then $\eta^{\alpha\beta} = 0$, if $\alpha = \beta = 0$ then $\eta^{\alpha\beta} = 1$, and so on. While it can be shown explicitly, it will be treated as fact that the defining algebra is

$$\gamma^\alpha \gamma^\beta + \gamma^\beta \gamma^\alpha = 2\eta^{\alpha\beta} I_4, \quad (\text{E.7})$$

which is simply to say that

$$\begin{aligned}(\gamma^0)^2 &= I_4 \\ (\gamma^{1,2, \text{ or } 3})^2 &= -I_4 \\ \gamma^\alpha \gamma^\beta &= 0 \quad \text{for } \alpha \neq \beta.\end{aligned}$$

Now, concerning the problem at hand

$$\text{Tr}(\gamma^\alpha \gamma^\beta) = \frac{1}{2}(\text{Tr}(\gamma^\alpha \gamma^\beta) + \text{Tr}(\gamma^\beta \gamma^\alpha)).$$

Using first the cyclic property of traces, the addition of two traces, and the defining algebra (Eqn. E.7)

$$\begin{aligned}\text{Tr}(\gamma^\alpha \gamma^\beta) &= \frac{1}{2}(\text{Tr}(\gamma^\alpha \gamma^\beta) + \text{Tr}(\gamma^\beta \gamma^\alpha)) \\ &= \frac{1}{2}\text{Tr}(\gamma^\alpha \gamma^\beta + \gamma^\beta \gamma^\alpha) \\ &= \frac{1}{2}\text{Tr}(2\eta^{\alpha\beta} I_4).\end{aligned}$$

Finally,

$$\text{Tr}(\gamma^\alpha \gamma^\beta) = 4\eta^{\alpha\beta}. \quad (\text{E.8})$$

E.5 Proof of $\text{Tr}(\gamma^\alpha \gamma^\beta \gamma^\delta \gamma^\epsilon) = 4(\eta^{\alpha\beta} \eta^{\delta\epsilon} - \eta^{\alpha\delta} \eta^{\beta\epsilon} + \eta^{\alpha\epsilon} \eta^{\beta\delta})$

Using the defining algebra (Eqn. E.7), commuting α and β yields that

$$\begin{aligned}\text{Tr}(\gamma^\alpha \gamma^\beta \gamma^\delta \gamma^\epsilon) &= \text{Tr}((2\eta^{\alpha\beta} - \gamma^\beta \gamma^\alpha) \gamma^\delta \gamma^\epsilon) \\ &= \text{Tr}(2\eta^{\alpha\beta} \gamma^\delta \gamma^\epsilon - \gamma^\beta \gamma^\alpha \gamma^\delta \gamma^\epsilon).\end{aligned}$$

Using the same method first for α and δ and then α and ϵ ,

$$\begin{aligned}\text{Tr}(\gamma^\alpha \gamma^\beta \gamma^\delta \gamma^\epsilon) &= \text{Tr}(2\eta^{\alpha\beta} \gamma^\delta \gamma^\epsilon - \gamma^\beta (2\eta^{\alpha\delta} - \gamma^\delta \gamma^\alpha) \gamma^\epsilon) \\ &= \text{Tr}(2\eta^{\alpha\beta} \gamma^\delta \gamma^\epsilon - 2\eta^{\alpha\delta} \gamma^\beta \gamma^\epsilon + \gamma^\beta \gamma^\delta \gamma^\alpha \gamma^\epsilon) \\ &= \text{Tr}(2\eta^{\alpha\beta} \gamma^\delta \gamma^\epsilon - 2\eta^{\alpha\delta} \gamma^\beta \gamma^\epsilon + \gamma^\beta \gamma^\delta (2\eta^{\alpha\epsilon} - \gamma^\beta \gamma^\delta \gamma^\epsilon \gamma^\alpha)) \\ &= \text{Tr}(2\eta^{\alpha\beta} \gamma^\delta \gamma^\epsilon - 2\eta^{\alpha\delta} \gamma^\beta \gamma^\epsilon + 2\eta^{\alpha\epsilon} \gamma^\beta \gamma^\delta - \gamma^\beta \gamma^\delta \gamma^\epsilon \gamma^\alpha).\end{aligned}$$

Recalling the addition properties of traces and observing Eqn. E.8,

$$\begin{aligned}\text{Tr}(\gamma^\alpha \gamma^\beta \gamma^\delta \gamma^\epsilon) &= 2\eta^{\alpha\beta} \text{Tr}(\gamma^\delta \gamma^\epsilon) - 2\eta^{\alpha\delta} \text{Tr}(\gamma^\beta \gamma^\epsilon) + 2\eta^{\alpha\epsilon} \text{Tr}(\gamma^\beta \gamma^\delta) - \text{Tr}(\gamma^\beta \gamma^\delta \gamma^\epsilon \gamma^\alpha) \\ &= 8\eta^{\alpha\beta} \eta^{\delta\epsilon} - 8\eta^{\alpha\delta} \eta^{\beta\epsilon} + 8\eta^{\alpha\epsilon} \eta^{\beta\delta} - \text{Tr}(\gamma^\beta \gamma^\delta \gamma^\epsilon \gamma^\alpha).\end{aligned}$$

Now recalling the cyclic permutative properties of traces, it can be seen that $\text{Tr}(\gamma^\beta \gamma^\delta \gamma^\epsilon \gamma^\alpha) = \text{Tr}(\gamma^\alpha \gamma^\beta \gamma^\delta \gamma^\epsilon)$. Then it is possible to move the last term on the right side to the left side, yielding the desired outcome

$$\text{Tr}(\gamma^\alpha \gamma^\beta \gamma^\delta \gamma^\epsilon) = 4\eta^{\alpha\beta}\eta^{\delta\epsilon} - 4\eta^{\alpha\delta}\eta^{\beta\epsilon} + 4\eta^{\alpha\epsilon}\eta^{\beta\delta}. \quad (\text{E.9})$$

APPENDIX F
BENCHMARK FIGURES

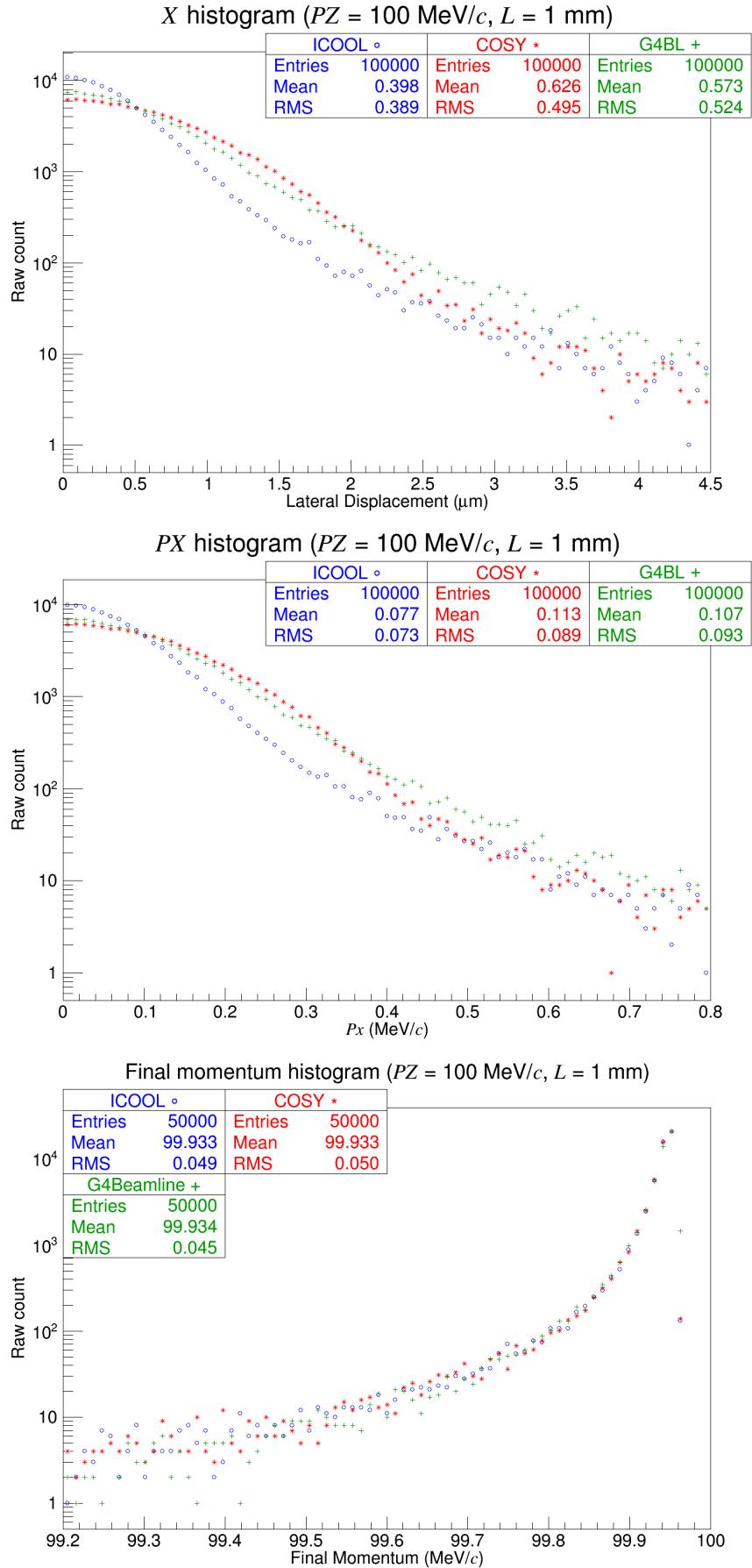


Figure F.1: Muons of momentum $100 \text{ MeV}/c$ through 1 mm liquid hydrogen.

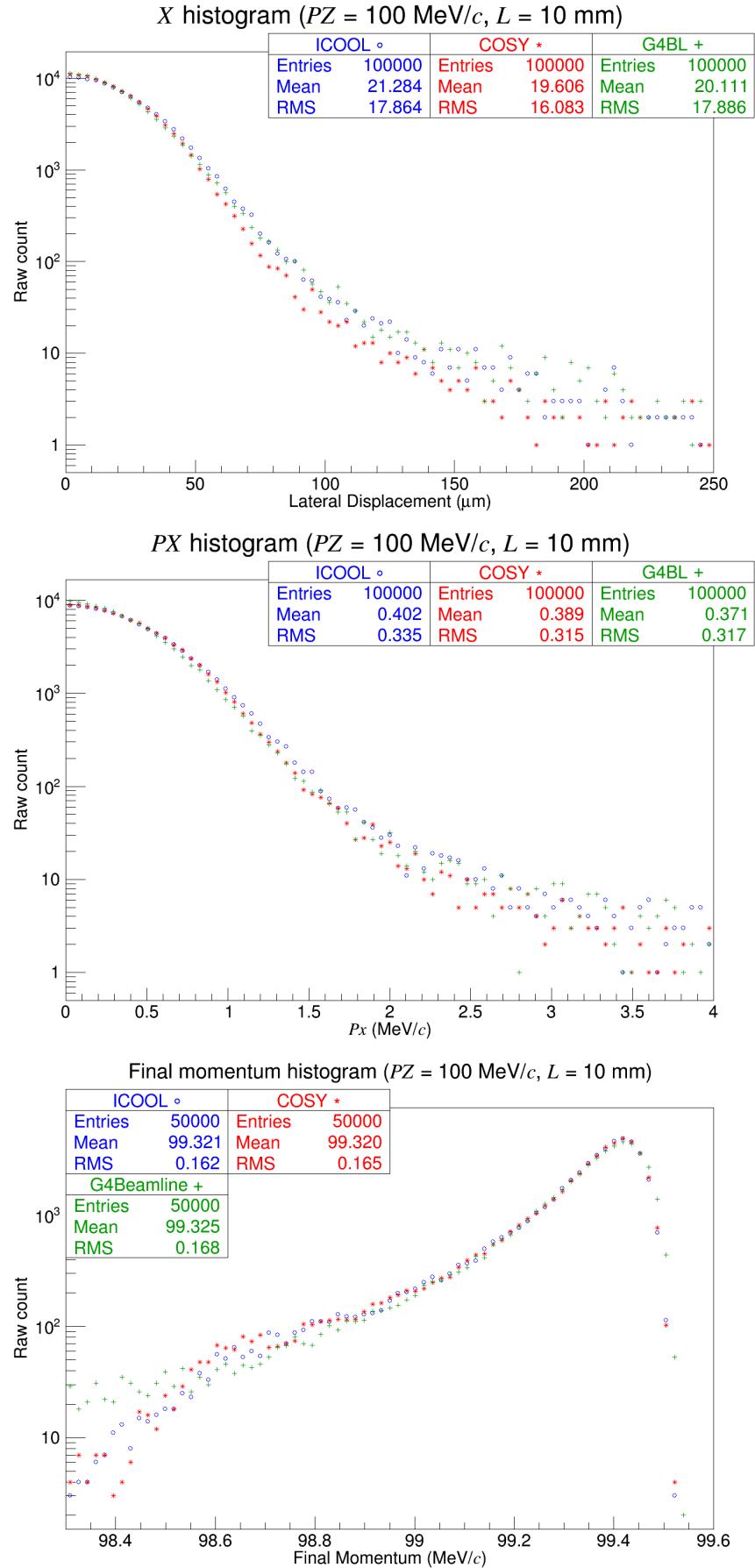


Figure F.2: Muons of momentum $100 \text{ MeV}/c$ through 10 mm liquid hydrogen.

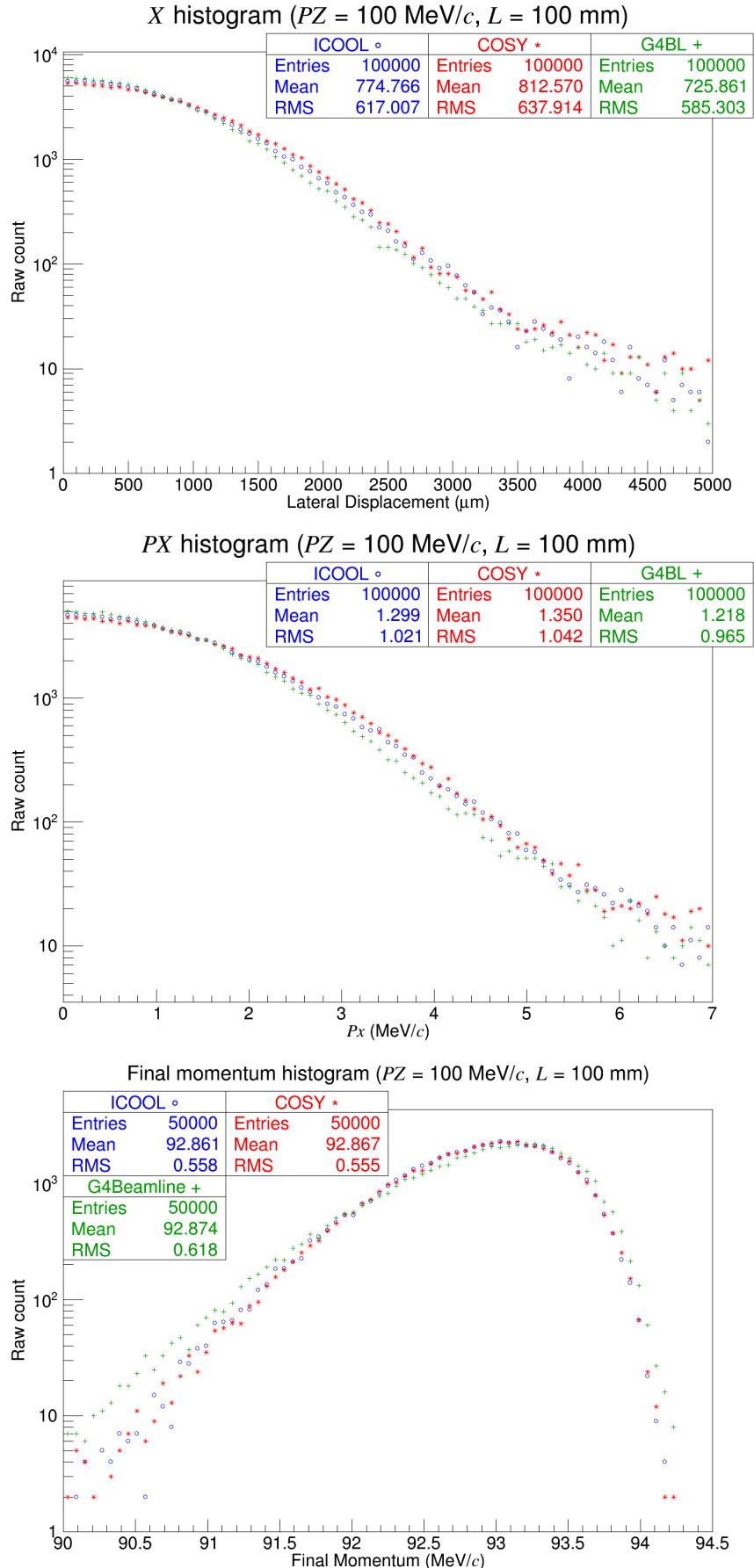


Figure F.3: Muons of momentum $100 \text{ MeV}/c$ through 100 mm liquid hydrogen.

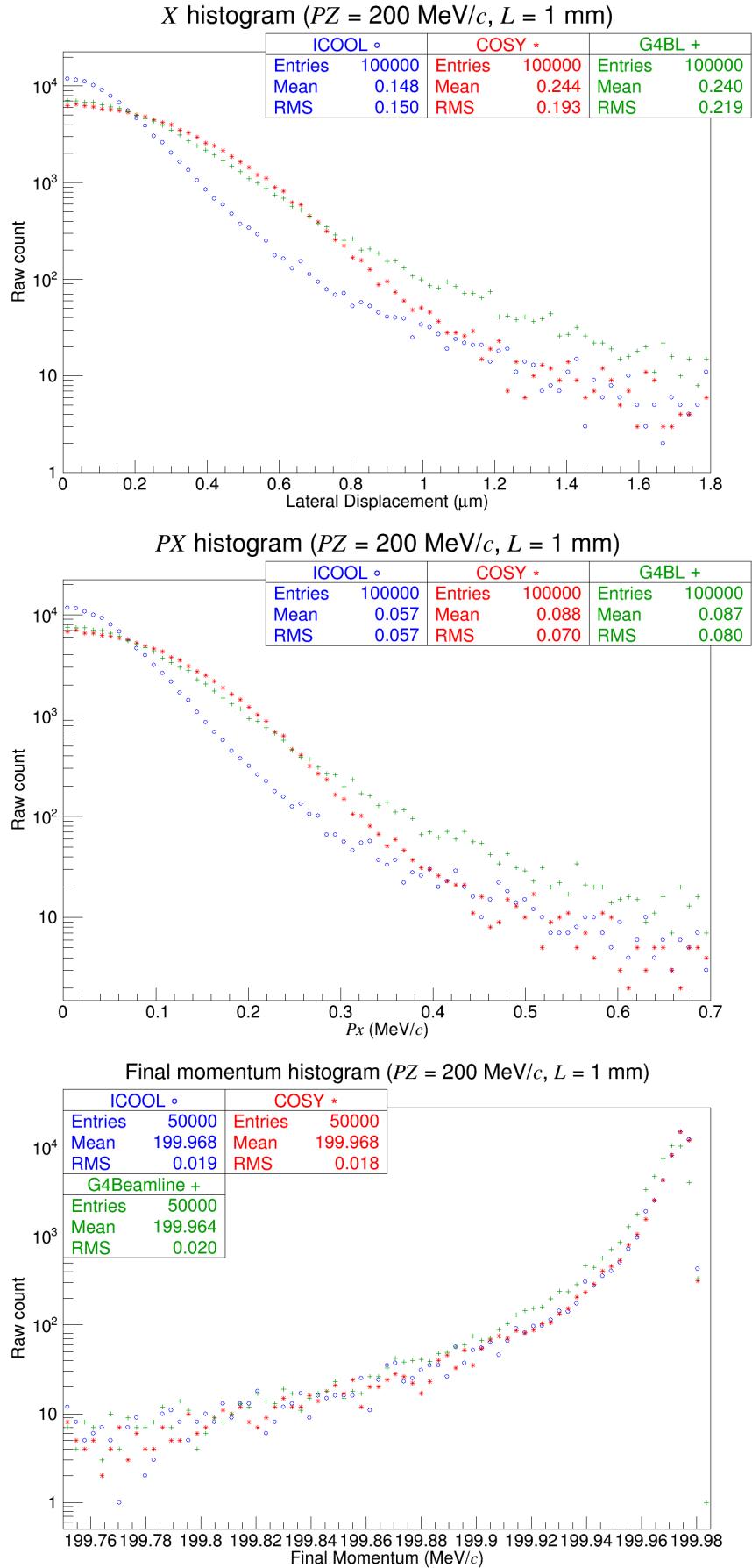


Figure F.4: Muons of momentum $200 \text{ MeV}/c$ through 1 mm liquid hydrogen.

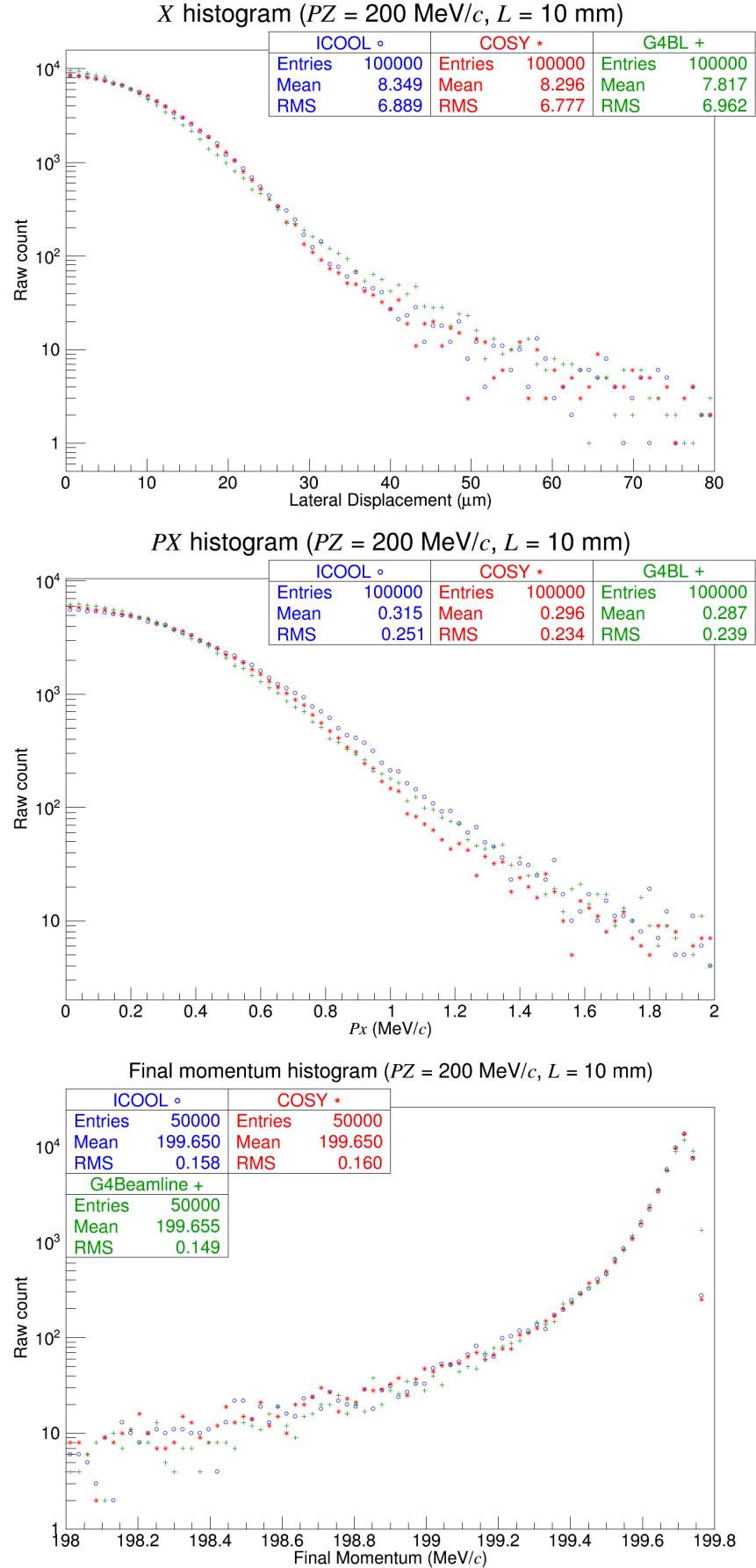


Figure F.5: Muons of momentum $200 \text{ MeV}/c$ through 10 mm liquid hydrogen.

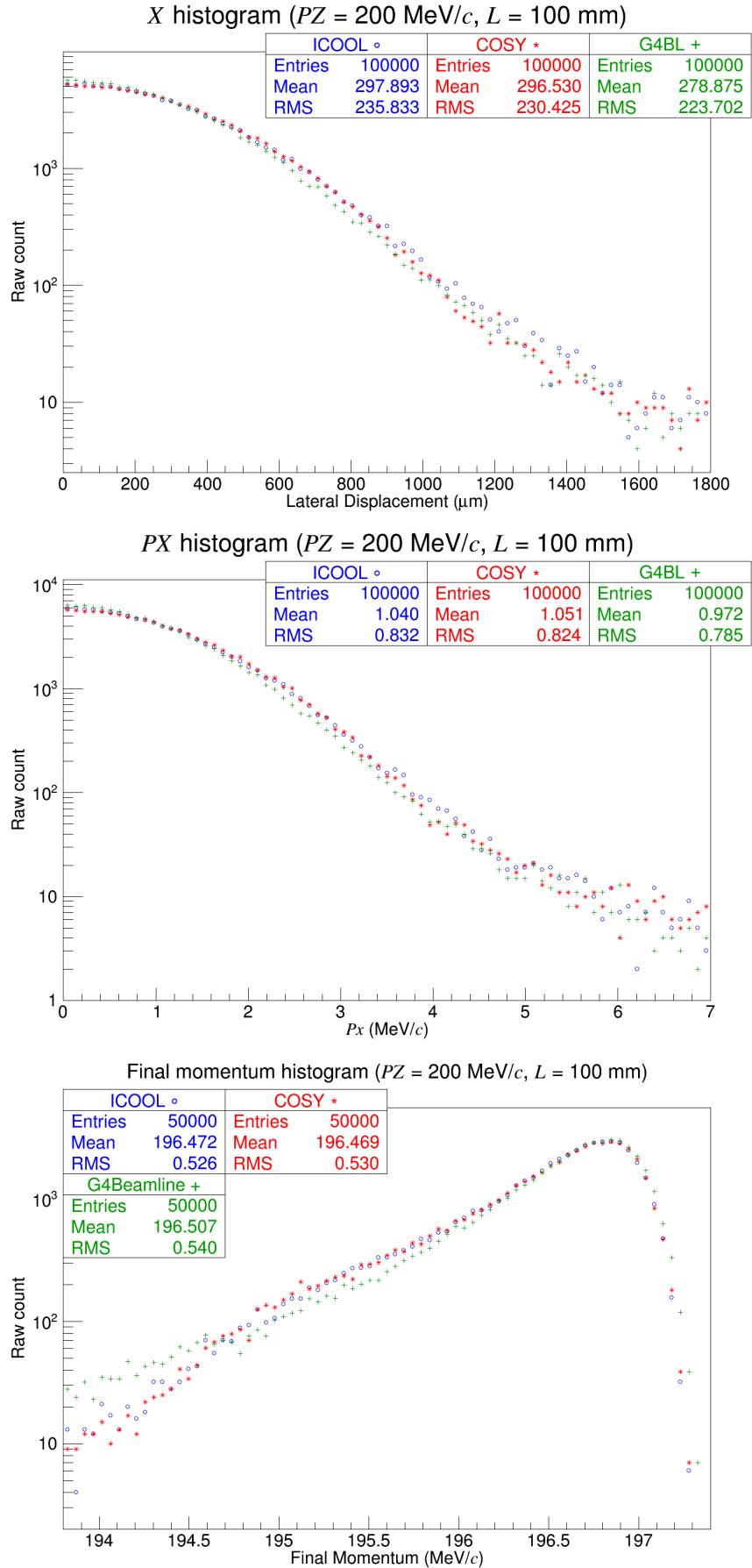


Figure F.6: Muons of momentum $200 \text{ MeV}/c$ through 100 mm liquid hydrogen.

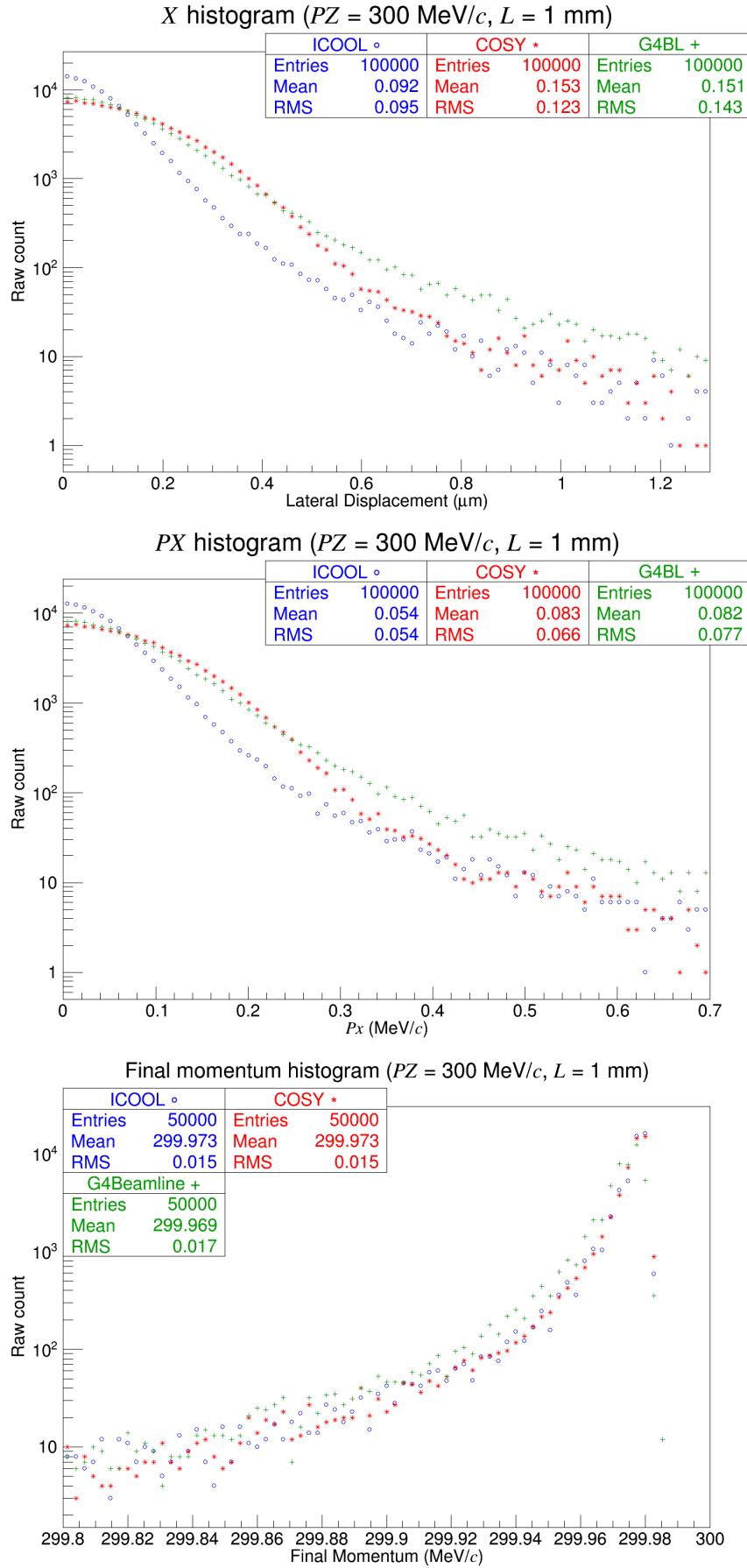


Figure F.7: Muons of momentum $300 \text{ MeV}/c$ through 1 mm liquid hydrogen.

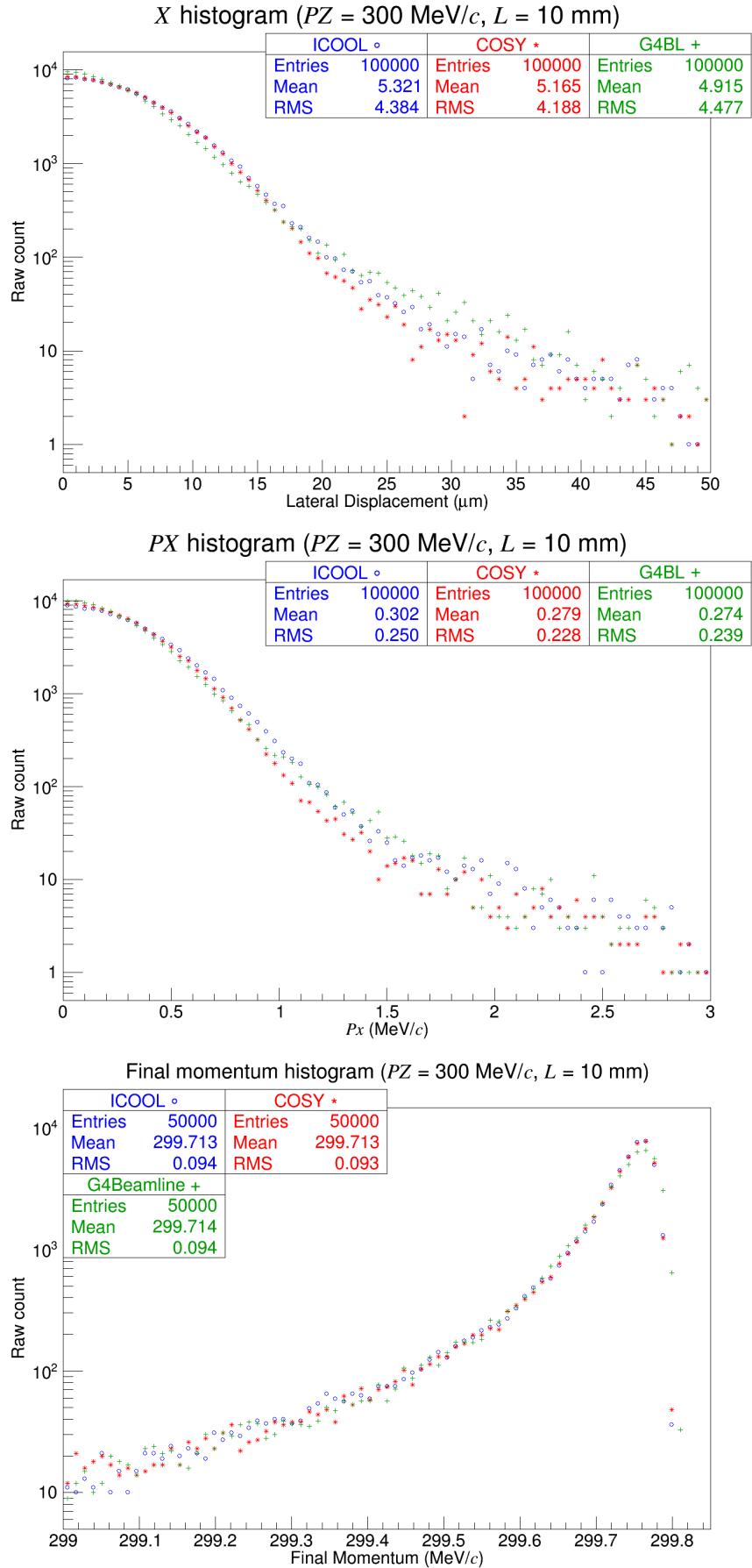


Figure F.8: Muons of momentum $300 \text{ MeV}/c$ through 10 mm liquid hydrogen.

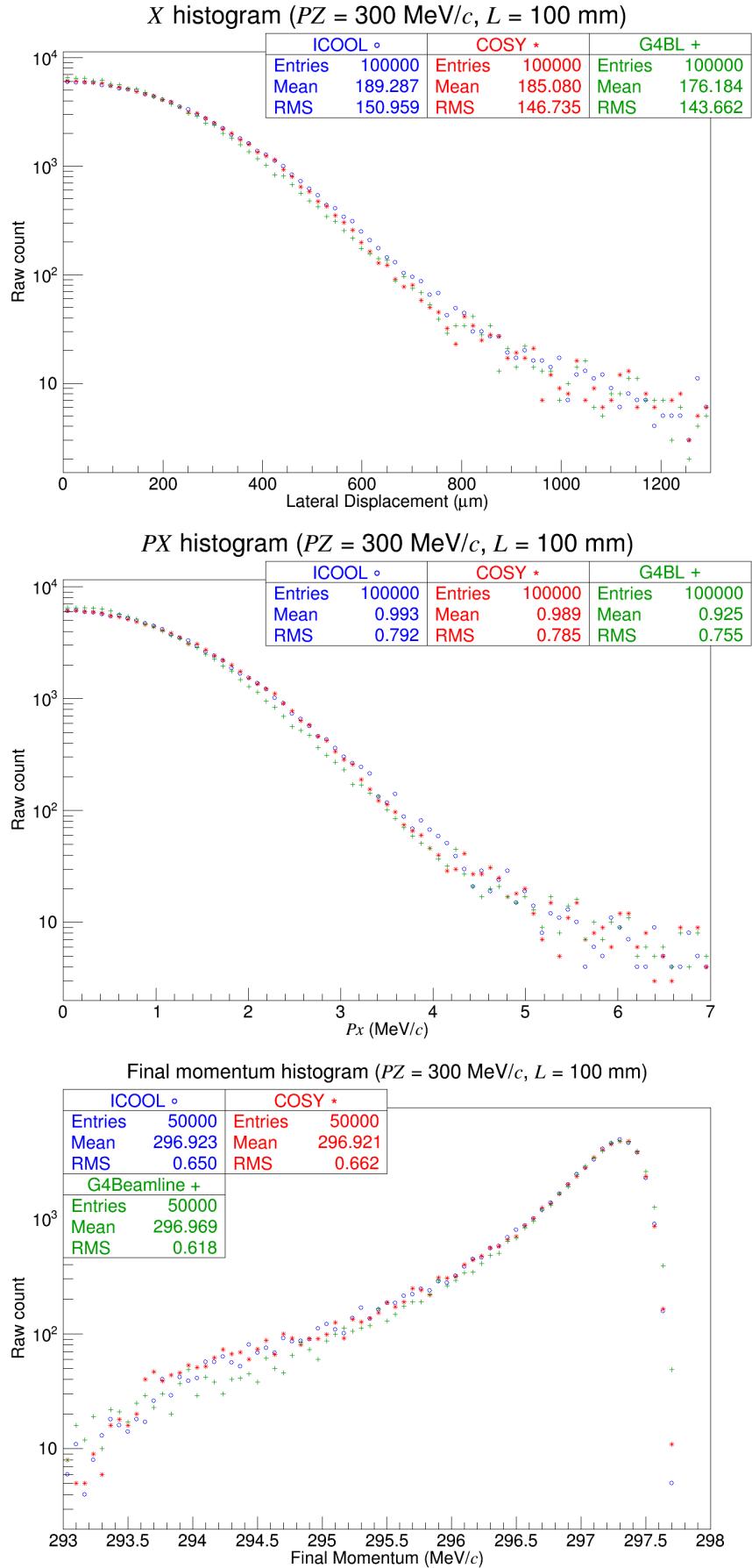


Figure F.9: Muons of momentum $300 \text{ MeV}/c$ through 100 mm liquid hydrogen.

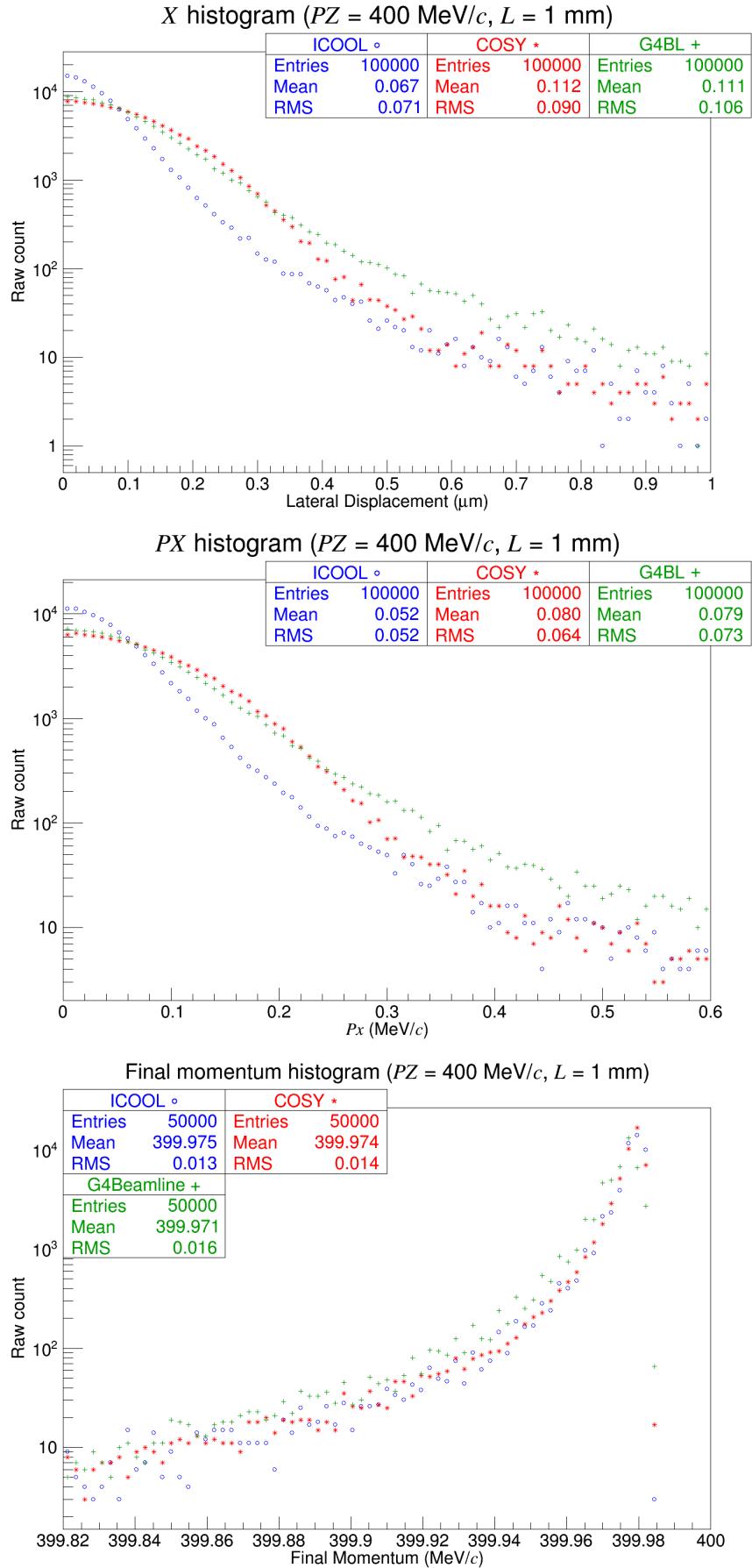


Figure F.10: Muons of momentum $400 \text{ MeV}/c$ through 1 mm liquid hydrogen.

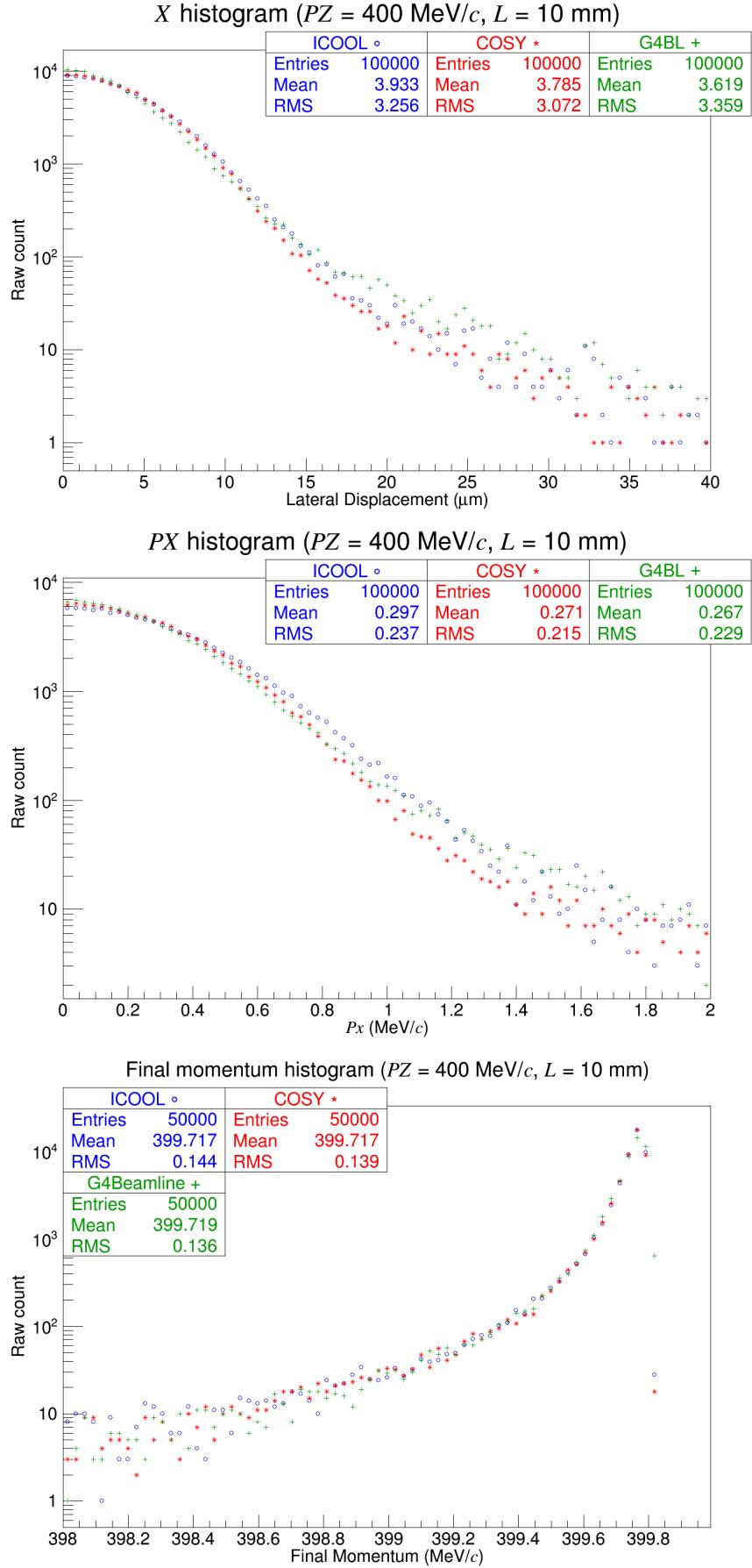


Figure F.11: Muons of momentum $400 \text{ MeV}/c$ through 10 mm liquid hydrogen.

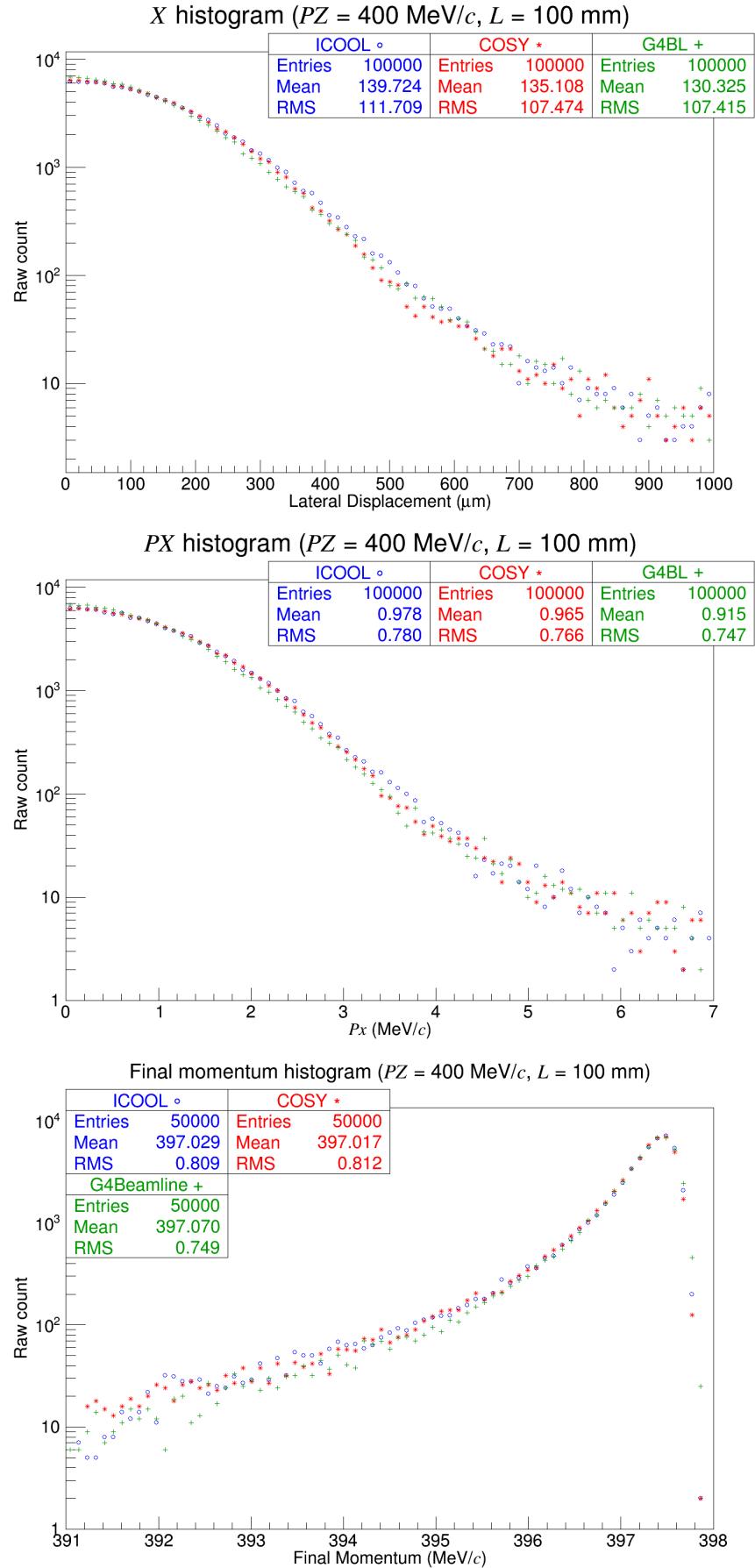


Figure F.12: Muons of momentum $400 \text{ MeV}/c$ through 100 mm liquid hydrogen.

APPENDIX G
REPRODUCTION OF IMPLEMENTED CODE

The following is a verbatim reproduction of the code that was implemented into COSY Infinity. This is broken up into two levels: COSYScript (inside the file *cosy.fox*) and FORTRAN (inside the file *absfox.f*). It should be noted that this appendix contains only code that was added to COSY version 9.1. As such, this section does not present a stand-alone code.

G.1 COSYScript

```

FUNCTION IN1DVE SIZE VAL ;
{INITIALIZES 1D VECTOR WITH ZEROS. VECTOR VE MUST BE EMPTY}

VARIABLE TMPX SIZE ;VARIABLE SIZETMP 1 ;
VARIABLE I 1 ; VARIABLE COUNT 1 ;
TMPX :=VAL ;
{FIRST LOOP}
SIZETMP :=SIZE ;
LOOP I 0 (INT(LOG(SIZETMP)/LOG(2))-1) ;
TMPX:=TMPX&TMPX ;
ENDLOOP;
IN1DVE :=TMPX ;
TMPX := 0 ;
COUNT :=COUNT+2^(INT(LOG(SIZETMP)/LOG(2))) ;
SIZETMP :=SIZETMP-2^(INT(LOG(SIZETMP)/LOG(2))) ;
WHILE (COUNT<SIZE) ;
LOOP I 0 (INT(LOG(SIZETMP)/LOG(2))-1) ;
TMPX:=TMPX&TMPX ;
ENDLOOP;
IN1DVE :=IN1DVE&TMPX ;
TMPX := 0 ;

```

```

COUNT := COUNT + 2^(INT(LOG(SIZETMP)/LOG(2))) ;
SIZETMP := SIZETMP - 2^(INT(LOG(SIZETMP)/LOG(2))) ;
ENDWHILE ;
ENDFUNCTION ;

PROCEDURE BGEN NPART V X SX PX SPX Y SY PY SPY T ST PZ SPZ ;
{PROCEDURE TO QUICKY GENERATE A GAUSSIAN BEAM.

NPART = NUMBER OF PARTICLES
V = 2D VECTOR TO PUT PARTICLES IN
X = AVERAGE X POSITION
SX = SIGMA X
ETC.

}

VARIABLE I 1 ; VARIABLE GAM NPART 1 ; VARIABLE PV 1 12 ;
{}

{INITIALIZE VECTOR V IS NECESSARY}

IF LENGTH(V(1))#NPART ;
LOOP I 1 6 ;
V(I) := IN1DVE(NPART,0);
ENDLOOP ;
ENDIF ;
{}

PV(1) := X ; PV(2) := SX ; PV(3) := PX ; PV(4) := SPX ;
PV(5) := Y ; PV(6) := SY ; PV(7) := PY ; PV(8) := SPY ;
PV(9) := T ; PV(10) := ST; PV(11) := PZ; PV(12) := SPZ;
GAUVEC PV 6 V NPART ;

```

```

{CONVERT BACK TO COSY COORDINATES}

V(6) :=SQRT(V(2)^2+V(4)^2+V(6)^2+(CONS(M0)*AMUMEV)^2) ;{NOW V(6) IS E}
V(6) :=V(6)-CONS(M0)*AMUMEV ;{NOW V(6) IS KE}
GAM(1) :=V(6)/(CONS(M0)*AMUMEV)+1 ;
V(2) :=V(2)/CONS(P0) ;
V(4) :=V(4)/CONS(P0) ;
V(6) :=(V(6)-CONS(E0))/CONS(E0) ;{NOW V(6) IS D}
V(5) :=-V(5)*CONS(V0)*GAM(1)/(1+GAM(1)) ;{ASSUMES TREF=0}
ENDPROCEDURE ;

PROCEDURE READ_G4BL FILE NPART V ;
{PROCEDURE TO READ PARTICLES FROM A G4BL-FORMATTED FILE AND PUT IT
INTO A COSY VECTOR V. INPUTS ARE:
FILE = FILE NUMBER TO READ (E.G. THE 12IN 'FORT.12')
NPART = NUMBER OF PARTICLES
V = VECTOR TO READ TO
}
VARIABLE I 1 ; VARIABLE GAM NPART ;VARIABLE MASS 1 ;
{INITIALIZE VECTOR IF NECESSARY}
IF LENGTH(V(1))#NPART ;
LOOP I 1 6 ;
V(I) :=IN1DVE(NPART,0);
ENDLOOP ;
ENDIF ;
{READ FILE AT FORTRAN LEVEL}
RG4BL FILE NPART V ;
{OUTPUT IS IN (X,PX,Y,PY,T,E) SO NEED TO TRANSLATE TO COSY COORDS}

```

```

MASS:=CONS(M0)*AMUMEV ;
GAM:=V(6)/MASS ;
V(2):=V(2)/CONS(P0) ;
V(4):=V(4)/CONS(P0) ;
V(5):=-V(5)*CONS(V0)*GAM/(1+GAM) ;
V(6):=(V(6)-MASS)/CONS(E0)-1 ;
ENDPROCEDURE ;

PROCEDURE READ_ICOOL FILE NPART V ;
{PROCEDURE TO READ PARTICLES FROM AN ICOOL-FORMATTED FILE AND PUT IT
INTO A COSY VECTOR V. INPUTS ARE:
FILE    = FILE NUMBER TO READ (E.G. THE 12IN 'FORT.12')
NPART   = NUMBER OF PARTICLES
V       = VECTOR TO READ TO
}
VARIABLE I 1 ; VARIABLE GAM NPART ;VARIABLE MASS 1 ;
{INITIALIZE VECTOR IF NECESSARY}
IF LENGTH(V(1))#NPART ;
LOOP I 1 6 ;
V(I):=IN1DVE(NPART,0);
ENDLOOP ;
ENDIF ;
{READ FILE AT FORTRAN LEVEL}
RICOOL FILE NPART V ;
{OUTPUT IS IN (X,PX,Y,PY,T,E) SO NEED TO TRANSLATE TO COSY COORDS}
MASS:=CONS(M0)*AMUMEV ;
GAM:=V(6)/MASS ;

```

```
V(2):=V(2)/CONS(P0) ;
```

```
V(4):=V(4)/CONS(P0) ;
```

```
V(5):=-V(5)*CONS(V0)*GAM/(1+GAM) ;
```

```
V(6):=(V(6)-MASS)/CONS(E0)-1 ;
```

```
ENDPROCEDURE ;
```

PROCEDURE WRITE_ICOOL Z REG NPART V UI APE ;

{PROCEDURE TO TAKE COSY ARRAY AND WRITE IT INTO ICOOL STYLE.

Z = Z-POSITION

REG = REGION NUMBER

NPART = NUMBER OF PARTICLES

V = VECTOR TO WRITE

UI = UI TO WRITE (E.G. UI=12 WRITES TO FORT.12)

APE = APERATURE (<=0 TO TURN OFF)

VECTOR SHOULD BE NORMAL COSY VECTOR:

V = (X,A,Y,B,L,D)

}

VARIABLE GAM1 LENGTH(V(1)) 1 ; VARIABLE L1 LENGTH(V(1)) 1 ;

VARIABLE V1 LENGTH(V(1)) 6 ;

{}

{NEED VECTOR "V1" =(X,PX,Y,PY,TOF,EI)}

V1(1) :=V(1) ;

V1(2) :=V(2)*CONS(P0) ;

V1(3) :=V(3) ;

```

V1(4) :=V(4)*CONS(P0) ;
V1(6) :=CONS(E0)*(V(6)+1)+CONS(M0)*AMUMEV ;
L1(1) :=V(5) ;
GAM1(1) :=V1(6)/(CONS(M0)*AMUMEV) ;
{TOF ASSUMING REF T =0}
V1(5) :=-L1(1)*(1+GAM1(1))/(CONS(V0)*GAM1(1)) ;
WICOOL Z REG NPART V1 UI APE ;
ENDPROCEDURE ;

PROCEDURE CMSTPT CURD R1 R2 L ZST S1 S2 AX AY SS ;
{TILTED CMSTP WITH THE USUAL ARGUMENTS +TILTS IN DEGREES +STEP SIZE.
CURD = CURRENT DENSITY
R1 = INNER RADIUS
R2 = OUTER RADIUS
L = LENGTH OF MAGNET
ZST = STARTING POSITION OF MAGNET
S1 = LEFT SLICE Z-POSITION
S2 = RIGHT SLICE Z-POSITION
AX = X TILT IN DEGREES
AY = Y TILT IN DEGREES
SS = STEPSIZE (CAN BE ZERO) }

VARIABLE I 1 ; VARIABLE CURDA 1 1 ;VARIABLE R1A 1 1 ;
VARIABLE R2A 1 1 ;VARIABLE LA 1 1 ;VARIABLE ZSTA 1 1 ;
VARIABLE XOFF 1 ; VARIABLE YOFF 1 ; VARIABLE ZOFF 1 ;
VARIABLE START 1 ;

```

{CORRECTION FOR X,Y (INPUT SHOULD BE SAME AS G4BL TILTS; MUST
TRANSLATE TO COSY. E.G. IF G4BL HAS A -3X TILT, THIS IS +3Y TILT.}

I :=AY;

AY:=-AX;

AX:=I;

CURDA(1):=CURD; R1A(1):=R1; R2A(1):=R2; LA(1):=L; ZSTA(1):=ZST;

DL ZST-S1+L/2 ;{GET TO CENTER OF MAGNET}

TA AX AY ;

DL -(ZST-S1+L/2) ;{BACK TO CELL'S START}

IF SS#0 ;

LOOP I S1 S2-SS SS ;

CMSTP 1 CURDA R1A R2A LA ZSTA I I+SS;

ENDLOOP ;

ELSEIF LO(1) ;

CMSTP 1 CURDA R1A R2A LA ZSTA S1 S2 ;

ENDIF ;

DL ZST-S2+L/2 ;{BACK TO MAGNET'S CENTER}

TA -AX -AY ;{TILT BACK IN LAB FRAME}

DL -(ZST-S2+L/2) ;{TO CELL'S EXIT}

ENDPROCEDURE ;

FUNCTION CEILING X ;

VARIABLE RX 1 ;

```

RX := NINT(X) ;
IF (RX<X) ;RX :=RX+1 ;ENDIF ;{CASE FOR NINT ROUNDED DOWN}
CEILING :=RX ;
ENDFUNCTION ;

PROCEDURE ABSPOLY S1 S2 N LEN APERTURE V RADLEN SPLIT LCELL SAV ;
{-----}

VARIABLE SS1 1000 N+1 N+1 ;VARIABLE SS2 1000 N+1 N+1 ;
VARIABLE LF1 NM1 1 ; VARIABLE PATH LENGTH(V(1)) 1 ;
VARIABLE V1 LENGTH(V(1)) 6 ; VARIABLE MASS 1 ; VARIABLE SUBLEN 1 ;
VARIABLE NPART 1 ; VARIABLE J 1 ; VARIABLE V01 1 ; VARIABLE V02 1 ;
VARIABLE I 1 ; VARIABLE K 1 ; VARIABLE LF NM1 ;VARIABLE SUBCELL 1 ;
VARIABLE ZEROES 1 6 ;VARIABLE L1 LENGTH(V(1)) 1 ; VARIABLE DT0 1 ;
VARIABLE GAM1 LENGTH(V(1)) 1 ; VARIABLE GAM2 LENGTH(V(1)) 1 ;
VARIABLE SAVNUM 1 ;

{-----}

{INITIALIZE}

IF (NINT(SPLIT)#SPLIT) ;
  WRITE 6 '###ERROR IN ABSPOLY###' ;
  WRITE 6 ' >ITERATION NUMBER MUST BE INTEGER!<' ;
  QUIT 0 ;
ELSEIF ((SPLIT<0)+(SPLIT=0)) ;
  WRITE 6 '###ERROR IN ABSPOLY###' ;
  WRITE 6 ' >ITERATION NUMBER MUST BE POSITIVE!<' ;
  QUIT 0 ;
ENDIF ;

```

```

MASS :=CONS(M0)*AMUMEV ;
NPART :=LENGTH(V(1)) ;
V01 :=CONS(V0) ;
LOOP I 1 6 ;ZEROES(I) := 0 ;ENDLOOP ;
DT0 := 0 ;

{-----}
{NEED VECTOR "V1" =(X,PX,Y,PY,TOF,EI)}

V1(1) :=V(1) ;
V1(2) :=V(2)*CONS(P0) ;
V1(3) :=V(3) ;
V1(4) :=V(4)*CONS(P0) ;
V1(6) :=CONS(E0)*(V(6)+1)+CONS(M0)*AMUMEV ;
L1(1) :=V(5) ;
GAM1(1) :=V1(6)/(CONS(M0)*AMUMEV) ;
{TOF ASSUMING REF T =0}
V1(5) :=-L1(1)*(1+GAM1(1))/(CONS(V0)*GAM1(1)) ;

{-----}
{PROPAGATE PARTICLES FROM CELL BOUNDARY TO WEDGE BOUNDARY}

IF (LCELL-LEN)>0 ;
LOOP I 1 N+1 ;
LOOP J 1 N+1 ;
SS1(I,J) := 0 ;
SS2(I,J) :=-1*S1(I,J) ;

```

```

        ENDLOOP ;

ENDLOOP ;

WL SS1 SS2 N 0 (LCELL-LEN)/2 LF ; {GET PATHS TRAVELED}

LF1(1) :=LF ; {REFERENCE TIME}

POLVAL 1 LF1 1 V 6 PATH 1 ;

V1(1) :=V1(1)+PATH(1)*V1(2)/SQRT(V1(6)^2-MASS^2) ;

V1(3) :=V1(3)+PATH(1)*V1(4)/SQRT(V1(6)^2-MASS^2) ;

V1(5) :=V1(5)+PATH(1)/(299792458*SQRT(V1(6)^2-MASS^2)/V1(6)) ;

DT0 :=DT0+(LCELL-LEN)/(2*V01) ; {REFERENCE TIME}

ENDIF ;

{-----}

{PROPAGATE THROUGH ABSORBER}

SUBLEN :=LEN/SPLIT ;

SUBCELL :=LCELL/SPLIT ;

LOOP K 1 SPLIT ;

LOOP I 1 N+1 ;

    LOOP J 1 N+1 ;

        SS1(I,J) :=(1-(K-1)/SPLIT)*S1(I,J)-(K-1)/SPLIT*S2(I,J) ;

        SS2(I,J) :=-(1-K/SPLIT)*S1(I,J)+K/SPLIT*S2(I,J) ;

    ENDLOOP ;

ENDLOOP ; {END I,J SUBABSORBER LOOPS}

WL SS1 SS2 N 0 SUBLEN LF ; {FIND DISTANCE TRAVELED}

LF1(1) :=LF ;

POLVAL 1 LF1 1 V 6 PATH 1 ; {PATH =DETERMINISTIC LENGTH}

SAVNUM := 0 ;

IF ((LCELL-LEN)=0)*(K=SPLIT) ; {ONLY SAVE ON LAST SPLIT}

```

```

SAVNUM :=SAV ;
ENDIF ;

STOABS V1 MASS PATH(1) BETHEBLOCHC RADLEN V1 SAVNUM APERTURE NPART ;
UM ;

WA SS1 SS2 N SUBLIN APERTURE ; {UPDATE REFERENCE PARTICLE}
UM ;

V02 :=CONS(V0) ; {NEW REFERENCE VELOCITY}
DT0 :=DT0+(V02-V01)*2*SUBLIN/(V02^2-V01^2) ;{NEW REFERENCE TIME}
ENDLOOP ; {END K LOOP}

{-----}
{PROPAGATE PARTICLES FROM WEDGE BOUNDARY TO CELL BOUNDARY}

IF (LCELL-LEN)>0 ;
LOOP I 1 N+1 ;
LOOP J 1 N+1 ;
SS1(I,J) :=-1*S2(I,J) ;
SS2(I,J) := 0 ;
ENDLOOP ;
ENDLOOP ;
WL SS1 SS2 N 0 (LCELL-LEN)/2 LF ;
LF1(1) :=LF ;
POLVAL 1 LF1 1 V 6 PATH 1 ;
STOABS V1 MASS PATH(1) ZEROES 0 V1 SAV APERTURE NPART ;
DT0 :=DT0+(LCELL-LEN)/(2*V02) ;
ENDIF ;

```

```
{-----}

{XFORM BACK TO COSY COORDINATES}

V(1) := V1(1) ;
V(2) := V1(2)/CONS(P0) ;
V(3) := V1(3) ;
V(4) := V1(4)/CONS(P0) ;
V(6) := (V1(6)-CONS(M0)*AMUMEV)/CONS(E0)-1 ;
GAM2(1) := V1(6)/(CONS(M0)*AMUMEV) ;
V(5) := (DT0-V1(5))*V02*GAM2(1)/(1+GAM2(1)) ;
V(5) := V(5)+L1(1)*V02*GAM2(1)*(1+GAM1(1))/(V01*GAM1(1)*(1+GAM2(1))) ;
ENDPROCEDURE ;
SAVE 'COSY' ;
```

G.2 FORTRAN

```
SUBROUTINE STOABS(IV1,IMA,IL,IBB,IRA,IV2,ISA,IAP,INP)
!      CALCULATES CONSTANTS FIRST WITHOUT INITIALIZING VECTORS
!
!      FOR EXPLANATION OF TERMS AND DERIVATIONS OF THEORETICAL
!      MODELS USED, PLEASE SEE JOSIAH KUNZ DISSERTATION 2016.
IMPLICIT DOUBLE PRECISION (A-H,O-Z)
DOUBLE PRECISION MASS
DOUBLE PRECISION BETHEBLOCHC(6)
INTEGER SAV,NPART
!-----MEMORY MANAGEMENT -----
PARAMETER(LMEM=1000000000,LVAR=100000000,LDIM=10000)
INTEGER NTYP(LVAR),NBEG(LVAR),NEND(LVAR),NMAX(LVAR),
```

```

*           NC(LMEM) ,NDIM(LDIM)

DOUBLE PRECISION CC(LMEM)

COMMON NTYP,NBEG,NEND,NMAX, CC,NC, NDIM, IDIM, IVAR, IMEM
COMMON /TYID/ NRE,NST,NLO,NCM,NVE,NIN,NIV,NDA,NCD,NTM,NGR

!

IF(NTYP(IMA).NE.NRE) CALL FOXNTY(IMA)
IF(NTYP(IAP).NE.NRE) CALL FOXNTY(IAP)
IF(NTYP(IRA).NE.NRE) CALL FOXNTY(IRA)
IF(NTYP(INP).NE.NRE) CALL FOXNTY(INP)

!

MASS = CC(NBEG(IMA))
RADLEN = CC(NBEG(IRA))
SAV = NINT(CC(NBEG(ISA)))
APE = CC(NBEG(IAP))
NPART = CC(NBEG(INP))

DO 5 I=0,6
    BETHEBLOCHC(I) = CC(NBEG(IBB+I))
5 CONTINUE

CALL STORUN(IV1,MASS,IL,BETHEBLOCHC,RADLEN,IV2,SAV,APE,NPART)
RETURN
END

! ****
SUBROUTINE STORUN(IV1,MASS,IL,BETHEBLOCHC,RADLEN,IV2,SAV,APE,
& NPART)
! TAKES PARTICLE VECTOR V1 = (X,PX,Y,PY,TOF,EI)
! THROUGH ABSORBER OF LENGTH L. THE ABSORBER HAS PROPERTIES
! DESCRIBED BY THE BETHEBLOCH ARRAY AND A RADIATION LENGTH

```

```

!      OF RADLEN. OUTPUT IS THE VECTOR V2.

IMPLICIT DOUBLE PRECISION (A-H,O-Z)

DOUBLE PRECISION IP,LO,LAMBAR,LAMMAX,LATDIS,LO,MASS,MASSE,
&                  MAXSTEP,MPL,NA,L

DOUBLE PRECISION BETHEBLOCHC(6),MP(6),TH0A(6),DA(6), BAR(6)
DOUBLE PRECISION, DIMENSION(:), ALLOCATABLE :: LV !LV(NPART)
DOUBLE PRECISION, DIMENSION(:, :), ALLOCATABLE :: V1 !V1(NPART,6)

INTEGER FLAG,I,J,LDK,LCNT,LIT,NPART,SAV

!-----MEMORY MANAGEMENT -----
PARAMETER(LMEM=1000000000, LVAR=100000000, LDIM=10000)

INTEGER NTYP(LVAR),NBEG(LVAR),NEND(LVAR),NMAX(LVAR),
*          NC(LMEM),NDIM(LDIM)
DOUBLE PRECISION CC(LMEM)

COMMON NTYP,NBEG,NEND,NMAX, CC,NC, NDIM, IDIM, IVAR, IMEM
COMMON /TYID/ NRE,NST,NLO,NCM,NVE,NIN,NIV,NDA,NCD,NTM,NGR

!
INTEGER IAllocateStatus,IAllocateStatus2
ALLOCATE(LV(NPART+10),Stat=IAllocateStatus)
ALLOCATE(V1(NPART+10,7),Stat=IAllocateStatus2)

IF (IAllocateStatus /= 0) STOP "*** Not enough memory ***"
IF (IAllocateStatus2 /= 0) STOP "*** Not enough memory ***"

MP(1) = BETHEBLOCHC(1)           ! Z NUMBER
MP(2) = BETHEBLOCHC(2)           ! ATOMIC MASS IN g/mol

```

```

MP(3) = BETHEBLOCHC(3)*1E6 ! DENSITY IN g/m^3
MP(4) = BETHEBLOCHC(4)      ! IONIZATION POTENTIAL IN MeV
MP(5) = BETHEBLOCHC(5)      ! DENSITY CORRECTION PARAMETER
MP(6) = BETHEBLOCHC(6)      ! SHELL CORRECTION PARAMETER

!
! GET INDIVIDUAL PATH LENGTHS AND (X, PX, ...)
!
*****  

DO 7 I=1,NPART
LV(I) = CC(NBEG(IL)+I-1)
DO 6 J=1,6
V1(I,J) = CC(NBEG(IV1+J)+I-1)
6    CONTINUE
7    CONTINUE
!
DO 18 I=1,NPART ! START PARTICLE LOOP
X0    = V1(I,1)
PX0   = V1(I,2)
Y0    = V1(I,3)
PY0   = V1(I,4)
TOFO  = V1(I,5)
EIO   = V1(I,6)
LO    = LV(I)
P0    = SQRT(EIO**2-MASS**2)
PZ0   = SQRT(P0**2-PX0**2-PY0**2)
!
ORIGINAL ANGLE TO ROTATE BETWEEN FRAMES
THO  = SIGN(ACOS(PZ0/P0),X0+Y0)
CTHO = PZ/P

```

```

STHO = SQRT(1-CTHO**2)

!
! NOTE THAT ORIGINAL COMPONENTS ARE DENOTED WITH AN '0' (E.G. XO,
! AS IN 'ORIGINAL', NOT ZERO, 0). COMPONENTS ON THIS PARTICULAR
! STEP OF THE SIMULATION WILL HAVE NO SUBSCRIPT (E.G. X).

X = XO
PX = PXO
Y = YO
PY = PYO
TOF = TOFO
EI = EIO
L = LO
P = PO
PZ = PZO

!
! DEFINE PHYSICAL CONSTANTS
!
*****  

C = 299792458.D0           ! SPEED OF LIGHT IN m/s
EULER = 0.5772156649015329d0
BETA = SQRT(1-(MASS/EI)**2)
GAM = 1/SQRT(1-BETA**2)
MASSE = 0.510998928d0       ! MASS OF ELECTRON IN MeV/c**2
NA = 6.02214129d23         ! AVAGADRO
PLANCK = 1.239842d-12       ! hc IN MeV*m
RELECTRON = 2.8179403267d-15 ! CLASSICAL ELECTRON RADIUS IN m
AO = 0.52917721092d-10     ! BOHR IN m
PI = 3.141592653589793d0

```

```

!
DZ = 0.D0          ! PATHLENGTH CORRECTION PLACEHOLDER

!
!      CHECK FLAGS
!
!*****  

10    CONTINUE  

FLAG = 0  

IF (EI.LE.MASS .OR. EIO.NE.EIO .OR. LO.NE.LO) THEN  

  FLAG = 1  !'PARTICLE STOPPED' FLAG  

  GOTO 16  !TERMINATE  

ELSEIF (SQRT(X**2+Y**2).GE.APE) THEN  

  FLAG = 2  !'PARTICLE HIT APERTURE' FLAG  

  GOTO 16  

ELSEIF (L.LE.0) THEN  

  FLAG = 3  !PARTICLE MISSED ABSORBER COMPLETELY  

  GOTO 16  

ELSEIF ((MP(1).EQ.0).OR.(MP(2).EQ.0).OR.(MP(3).EQ.0).OR.  

&           (MP(4).EQ.0)) THEN  

  FLAG = 5  !'DRIFT' FLAG (DEALS WITH Z=0)  

  GOTO 16  

ENDIF  

!
!      L = 2*L/(1+CTH) !TRUE PATHLENGTH CORRECTION PLACEHOLDER
!
!
!      STRAGGLING
!
!*****
```

```

15    CONTINUE

DE  = STRAG(MASS,EI,L,MP)

IF ((DE.GE.EI-MASS).OR.(EI.EQ.MASS)) THEN

FLAG = 1                      !PARTICLE STOPPED

GOTO 16

ENDIF

EI  = EI-DE                    !NEW TOTAL ENERGY

P   = SQRT(EI**2-MASS**2) !NEW TOTAL MOMENTUM

PZ  = P*PZO/PO                !ANGLE DID NOT CHANGE

PX  = P*PXO/PO

PY  = P*PYO/PO

!

! SCATTER

!

*****  

IF (RADLEN.LE.0) RADLEN = 100*716.4/ !CRUDE APPROXIMATION IF
& (RHO*1E-6*Z*(Z+1)*LOG(287/SQRT(Z))) !DO NOT KNOW RAD LEN
!  

XD  = L*PXO/PZO                !X DETERMINISTIC

YD  = L*PYO/PZO

!  

! 1ST ORDER HIGHLAND CORRECTION

HC1 = 0.12d0

! 2ND ORDER HIGHLAND CORRECTION

HC2 = 0.006d0

! CRITICAL ANGLE (LIKE GAUS SIGMA)

TH0  = 13.6/(BETA*P)*SQRT(L/RADLEN*
&           (1+HC1*LOG(L/RADLEN)+HC2*  


```

```

&      (LOG(L/RADLEN))**2))

!

! PZ ROTATED, NOT IN LAB FRAME YET

PZR = SCATDIST(TH0,P)

PT = SQRT(P**2-PZ**2) !TRANS P AFTER STRAG

IF (PT.NE.PT) PT = 0 !IN CASE PT = NaN

PTR = SQRT(P**2-PZR**2) !PT ROTATED FRAME

PZL = PZR*COS(TH0)-PTR*SIN(TH0) !PZ IN LAB FRAME

PTL = SQRT(P**2-PZL**2) !PT IN LAB FRAME

!

CALL RANDOM_NUMBER(R1)

R1 = R1*2*PI

PX = PX+(PTL-PT)*SIGN(COS(R1),R1-PI)

PY = PY+(PTL-PT)*SIN(R1)

!

TRANVERSE DISPLACEMENT

*****  

THETAC = 13.6/(BETA*P)*SQRT(1/RADLEN)

AMOUNT OF DEFLECTION IN X

ANGDIF = ASIN(PX/P)-ASIN(PX0/PO)

SHIFT+DETERMINISTIC+FLUCTUATION

X = X+XD+LATDIS(ANGDIF,L,THETAC)

ANGDIF = ASIN(PY/P)-ASIN(PY0/PO)

Y = Y+YD+LATDIS(ANGDIF,L,THETAC)

TOF = TOF+DTOF(EIO,P,MASS,L+DZ)

CHECK FLAGS AGAIN

IF (FLAG.EQ.0) THEN

```

```

    IF (EI.LE.MASS) THEN
!
!      'PARTICLE STOPPED' FLAG
!
    FLAG = 1

    ELSEIF (SQRT(X**2+Y**2).GE.APE) THEN
!
!      'PARTICLE HIT APERTURE' FLAG
!
    FLAG = 2

    ENDIF
!
    ENDIF

16   CONTINUE

!
!      'PARTICLE MISSED ABSORBER' OR 'DRIFT' FLAG
!
    IF ((FLAG.EQ.3).OR.(FLAG.EQ.5)) THEN
!
        EI = EIO
!
        X = X0+LO*PX0/PO
!
        Y = Y0+LO*PY0/PO
!
        TOF = TOFO+LO/(C*SQRT(EI**2-MASS**2)/EI)
!
    STOPPED, APERTURE, OR DECAY
!
    ELSEIF ((FLAG.EQ.1).OR.(FLAG.EQ.2).OR.(FLAG.EQ.4)) THEN
!
        X = 0.D0
!
        PX = 0.D0
!
        Y = 0.D0
!
        PY = 0.D0
!
        TOF = 0.D0
!
        EI = 0.D0
!
    ENDIF
!
    CC(NBEG(IV1+1)+I-1) = X
!
    CC(NBEG(IV1+2)+I-1) = PX
!
    CC(NBEG(IV1+3)+I-1) = Y

```

```

CC(NBEG(IV1+4)+I-1) = PY
CC(NBEG(IV1+5)+I-1) = TOF
CC(NBEG(IV1+6)+I-1) = EI
IF (SAV.GT.0) THEN
  WRITE(SAV,20) X,PX,Y,PY,TOF,EI,FLAG
ENDIF
20 FORMAT(6E22.14,I5)
21 FORMAT(1A3)
18 CONTINUE
RETURN
END
!
! ****
FUNCTION STRAG(MASS,EI0,L0,MP)
!
! DETERMINES ENERGY LOSS OF A PARTICLE TRAVELLING THROUGH
! ABSORBER WITH ALPHALAN AS THE APPROXIMATE MEAN ENERGY LOSS,
! XI AS A SCALING PARAMETER, DMAX AS THE MAXIMUM ENERGY LOSS
! (FOR THE LANDAU DISTRIBUTION), EMAX AS THE MAXIMUM ENERGY
! LOSS PER INTERACTION, AND BETA*C AS THE INCIDENT SPEED OF
! THE PARTICLE.
!
IMPLICIT DOUBLE PRECISION (A-H,O-Z)
DOUBLE PRECISION MP(6)
DOUBLE PRECISION PAR(6)
DOUBLE PRECISION IP,L,LAMBAR,LAMBARI,LAMMAX,LAMMAXI,LANDAU,
&           LANDAU2,L0,LLI,MASS,MASSE,NA
INTEGER J,K
!
!
```

```

!      DEFINE PHYSICAL CONSTANTS

! ****

EULER = 0.5772156649015329d0

BETA = SQRT(1-(MASS/EI)**2)

GAM = 1/SQRT(1-BETA**2)

MASSE = 0.510998928d0      !MASS OF ELECTRON IN MeV/c**2

NA = 6.02214129d23         !AVAGADRO

PLANCK = 1.239842d-12       !hc IN MeV*m

RELECTRON = 2.8179403267d-15 !CLASSICAL ELECTRON RADIUS IN m

A0 = 0.52917721092d-10     !BOHR IN m

PI = 3.141592653589793d0

!      LANDAU LAMBDA CUTOFF PARAMETERS

PAR = (/ 0.517891D0, 1.17765D0, 0.476074D0, 0.00880733D0,
&           1.15467D0 , 0.984008D0 /)

!

!      DEFINE SPECIFIC MATERIAL CONSTANTS FOR ENERGY LOSS

! ****

Z   = MP(1)

A   = MP(2)

RHO = MP(3)

IP  = MP(4)

DCP = MP(5)

SCP = MP(6)

!

!      CALL FLAGS FOR CALCULATING DENSITY, SHELL PARAMETERS

! ****

DF = 0.D0

```

```

SF = 0.D0

IF (DCP.EQ.0) DF = 1 !USER DID NOT SUPPLY CORRECTION PARAMETER
IF (SCP.EQ.0) SF = 1 !SO WE WILL APPROXIMATE
!
! CAN ONLY CALCULATE ONE STEP AT A TIME, SINCE E IS DYNAMIC
! (CANNOT SIMPLY SPLIT ABSORBER INTO EVEN PIECES).
! THAT IS, IF E IS TOO HIGH THE STEP MAY NOT BE A 'LANDAU' STEP
! ANYMORE, INVALIDATING THE THEORY.
!
CUL = 0.D0 !CUMULATIVE LENGTH
DE = 0.D0 !ENERGY LOSS
!
DO 25 WHILE (CUL.LT.L0) !LOOP THROUGH LANDAU-EXCLUSIVE STEPS
EF = E10-DE
GAM = EF/MASS
BETA = SQRT(1-1/GAM**2)
!
! DISABLE CRUDE APPROXIMATION FOR NOW
! IF (DF.NE.0) DCP = GET_DCP(BETA*GAM,MP)
! IF (SF.NE.0) SCP = GET_SCP(BETA*GAM,MP)
EMAX = 2*MASSE*(BETA*GAM)**2/(1+2*GAM*MASSE/MASS+
& (MASSE/MASS)**2)
J = 0
!
! FIND L WHICH WILL PRODUCE A LANDAU ENERGY PROFILE
!
*****DO 5 WHILE ((XI/EMAX>0.01D0).OR.(J<1))
J = J+1
L = L0/J

```

```

XI = 2*PI*RELECTRON**2*MASSE*NA*Z*RHO*L/(A*BETA**2)

IF (J>10000) THEN !PARTICLE IS GOING TOO SLOW, FAIL IT

STRAG = EI0+1

RETURN

ENDIF

5 CONTINUE

!

CUL = CUL+L

!

IF (CUL.GT.L0) THEN !ONLY USED ON THE FINAL STEP

L = L0-(CUL-L)

XI = 2*PI*RELECTRON**2*MASSE*NA*Z*RHO*L/(A*BETA**2)

ENDIF

IF (L.LE.0) GOTO 25 !PRECISION ERROR

!

! NOW PROPAGATE PARTICLE THROUGH L

!

*****  

EBAR = XI*(LOG(2*MASSE*(BETA*GAM/IP)**2*EMAX)-2*BETA**2-DCP-
& SCP)

ALPHALAN = EBAR+XI*(BETA**2+LOG(XI/EMAX)+1-EULER)

LAMBAR = -(1-EULER)-BETA**2-LOG(XI/EMAX)

LAMMAX = PAR(1)+PAR(2)*LAMBAR+(PAR(3)+PAR(4)*LAMBAR)*
& EXP(PAR(5)+PAR(6)*LAMBAR)

DEMAX = XI*(LAMMAX+1-EULER+BETA**2+LOG(XI/EMAX))+EBAR

DE = DE+LANDAU(ALPHALAN,XI,DEMAX)

25 CONTINUE

STRAG = DE

```

```

RETURN

END

! ****
REAL FUNCTION LANDAU(MEAN,SIGMA,DEMAX)
! Generate a random number following a Landau distribution
! with Landau parameters alphalan = mean, betalan= xi = sigma.
!
! This is a copy from the source file TRandom from SLAC.
! Originally converted by Rene Brun from CERNLIB routine
! ranlan(G110).

IMPLICIT DOUBLE PRECISION (A-H,O-Z)

DOUBLE PRECISION MEAN

INTEGER I,J,seed

DOUBLE PRECISION F(982)

F = (/
& 0.        , 0        , 0        ,0        ,0        ,-2.244733 ,
& -2.204365,-2.168163,-2.135219,-2.104898,-2.076740,-2.050397 ,
& -2.025605,-2.002150,-1.979866,-1.958612,-1.938275,-1.918760 ,
& -1.899984,-1.881879,-1.864385,-1.847451,-1.831030,-1.815083 ,
& -1.799574,-1.784473,-1.769751,-1.755383,-1.741346,-1.727620 ,
& -1.714187,-1.701029,-1.688130,-1.675477,-1.663057,-1.650858 ,
& -1.638868,-1.627078,-1.615477,-1.604058,-1.592811,-1.581729 ,
& -1.570806,-1.560034,-1.549407,-1.538919,-1.528565,-1.518339 ,
& -1.508237,-1.498254,-1.488386,-1.478628,-1.468976,-1.459428 ,
& -1.449979,-1.440626,-1.431365,-1.422195,-1.413111,-1.404112 ,
& -1.395194,-1.386356,-1.377594,-1.368906,-1.360291,-1.351746 ,
& -1.343269,-1.334859,-1.326512,-1.318229,-1.310006,-1.301843 ,

```

& -1.293737, -1.285688, -1.277693, -1.269752, -1.261863, -1.254024,
 & -1.246235, -1.238494, -1.230800, -1.223153, -1.215550, -1.207990,
 & -1.200474, -1.192999, -1.185566, -1.178172, -1.170817, -1.163500,
 & -1.156220, -1.148977, -1.141770, -1.134598, -1.127459, -1.120354,
 & -1.113282, -1.106242, -1.099233, -1.092255,
 & -1.085306, -1.078388, -1.071498, -1.064636, -1.057802, -1.050996,
 & -1.044215, -1.037461, -1.030733, -1.024029, -1.017350, -1.010695,
 & -1.004064, -.997456, -.990871, -.984308, -.977767, -.971247,
 & -.964749, -.958271, -.951813, -.945375, -.938957, -.932558,
 & -.926178, -.919816, -.913472, -.907146, -.900838, -.894547,
 & -.888272, -.882014, -.875773, -.869547, -.863337, -.857142,
 & -.850963, -.844798, -.838648, -.832512, -.826390, -.820282,
 & -.814187, -.808106, -.802038, -.795982, -.789940, -.783909,
 & -.777891, -.771884, -.765889, -.759906, -.753934, -.747973,
 & -.742023, -.736084, -.730155, -.724237, -.718328, -.712429,
 & -.706541, -.700661, -.694791, -.688931, -.683079, -.677236,
 & -.671402, -.665576, -.659759, -.653950, -.648149, -.642356,
 & -.636570, -.630793, -.625022, -.619259, -.613503, -.607754,
 & -.602012, -.596276, -.590548, -.584825, -.579109, -.573399,
 & -.567695, -.561997, -.556305, -.550618, -.544937, -.539262,
 & -.533592, -.527926, -.522266, -.516611, -.510961, -.505315,
 & -.499674, -.494037, -.488405, -.482777,
 & -.477153, -.471533, -.465917, -.460305, -.454697, -.449092,
 & -.443491, -.437893, -.432299, -.426707, -.421119, -.415534,
 & -.409951, -.404372, -.398795, -.393221, -.387649, -.382080,
 & -.376513, -.370949, -.365387, -.359826, -.354268, -.348712,
 & -.343157, -.337604, -.332053, -.326503, -.320955, -.315408,

& - .309863, - .304318, - .298775, - .293233, - .287692, - .282152,
 & - .276613, - .271074, - .265536, - .259999, - .254462, - .248926,
 & - .243389, - .237854, - .232318, - .226783, - .221247, - .215712,
 & - .210176, - .204641, - .199105, - .193568, - .188032, - .182495,
 & - .176957, - .171419, - .165880, - .160341, - .154800, - .149259,
 & - .143717, - .138173, - .132629, - .127083, - .121537, - .115989,
 & - .110439, - .104889, - .099336, - .093782, - .088227, - .082670,
 & - .077111, - .071550, - .065987, - .060423, - .054856, - .049288,
 & - .043717, - .038144, - .032569, - .026991, - .021411, - .015828,
 & - .010243, - .004656, .000934, .006527, .012123, .017722,
 & .023323, .028928, .034535, .040146, .045759, .051376,
 & .056997, .062620, .068247, .073877,
 & .079511, .085149, .090790, .096435, .102083, .107736,
 & .113392, .119052, .124716, .130385, .136057, .141734,
 & .147414, .153100, .158789, .164483, .170181, .175884,
 & .181592, .187304, .193021, .198743, .204469, .210201,
 & .215937, .221678, .227425, .233177, .238933, .244696,
 & .250463, .256236, .262014, .267798, .273587, .279382,
 & .285183, .290989, .296801, .302619, .308443, .314273,
 & .320109, .325951, .331799, .337654, .343515, .349382,
 & .355255, .361135, .367022, .372915, .378815, .384721,
 & .390634, .396554, .402481, .408415, .414356, .420304,
 & .426260, .432222, .438192, .444169, .450153, .456145,
 & .462144, .468151, .474166, .480188, .486218, .492256,
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& .608554, .614762, .620980, .627207, .633444, .639689,
 & .645945, .652210, .658484, .664768,
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 & .709036, .715400, .721775, .728160, .734556, .740963,
 & .747379, .753807, .760246, .766695, .773155, .779627,
 & .786109, .792603, .799107, .805624, .812151, .818690,
 & .825241, .831803, .838377, .844962, .851560, .858170,
 & .864791, .871425, .878071, .884729, .891399, .898082,
 & .904778, .911486, .918206, .924940, .931686, .938446,
 & .945218, .952003, .958802, .965614, .972439, .979278,
 & .986130, .992996, .999875, 1.006769, 1.013676, 1.020597,
 & 1.027533, 1.034482, 1.041446, 1.048424, 1.055417, 1.062424,
 & 1.069446, 1.076482, 1.083534, 1.090600, 1.097681, 1.104778,
 & 1.111889, 1.119016, 1.126159, 1.133316, 1.140490, 1.147679,
 & 1.154884, 1.162105, 1.169342, 1.176595, 1.183864, 1.191149,
 & 1.198451, 1.205770, 1.213105, 1.220457, 1.227826, 1.235211,
 & 1.242614, 1.250034, 1.257471, 1.264926, 1.272398, 1.279888,
 & 1.287395, 1.294921, 1.302464, 1.310026, 1.317605, 1.325203,
 & 1.332819, 1.340454, 1.348108, 1.355780,
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 & 1.554068, 1.562275, 1.570503, 1.578754, 1.587028, 1.595325,
 & 1.603644, 1.611987, 1.620353, 1.628743, 1.637156, 1.645593,
 & 1.654053, 1.662538, 1.671047, 1.679581, 1.688139, 1.696721,
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& 1.757506, 1.766293, 1.775106, 1.783945, 1.792810, 1.801703,
& 1.810623, 1.819569, 1.828543, 1.837545, 1.846574, 1.855631,
& 1.864717, 1.873830, 1.882972, 1.892143, 1.901343, 1.910572,
& 1.919830, 1.929117, 1.938434, 1.947781, 1.957158, 1.966566,
& 1.976004, 1.985473, 1.994972, 2.004503, 2.014065, 2.023659,
& 2.033285, 2.042943, 2.052633, 2.062355, 2.072110, 2.081899,
& 2.091720, 2.101575, 2.111464, 2.121386, 2.131343, 2.141334,
& 2.151360, 2.161421, 2.171517, 2.181648, 2.191815, 2.202018,
& 2.212257, 2.222533, 2.232845, 2.243195,
& 2.253582, 2.264006, 2.274468, 2.284968, 2.295507, 2.306084,
& 2.316701, 2.327356, 2.338051, 2.348786, 2.359562, 2.370377,
& 2.381234, 2.392131, 2.403070, 2.414051, 2.425073, 2.436138,
& 2.447246, 2.458397, 2.469591, 2.480828, 2.492110, 2.503436,
& 2.514807, 2.526222, 2.537684, 2.549190, 2.560743, 2.572343,
& 2.583989, 2.595682, 2.607423, 2.619212, 2.631050, 2.642936,
& 2.654871, 2.666855, 2.678890, 2.690975, 2.703110, 2.715297,
& 2.727535, 2.739825, 2.752168, 2.764563, 2.777012, 2.789514,
& 2.802070, 2.814681, 2.827347, 2.840069, 2.852846, 2.865680,
& 2.878570, 2.891518, 2.904524, 2.917588, 2.930712, 2.943894,
& 2.957136, 2.970439, 2.983802, 2.997227, 3.010714, 3.024263,
& 3.037875, 3.051551, 3.065290, 3.079095, 3.092965, 3.106900,
& 3.120902, 3.134971, 3.149107, 3.163312, 3.177585, 3.191928,
& 3.206340, 3.220824, 3.235378, 3.250005, 3.264704, 3.279477,
& 3.294323, 3.309244, 3.324240, 3.339312, 3.354461, 3.369687,
& 3.384992, 3.400375, 3.415838, 3.431381, 3.447005, 3.462711,
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& 3.542497, 3.558711, 3.575012, 3.591402, 3.607881, 3.624450,

& 3.641111, 3.657863, 3.674708, 3.691646, 3.708680, 3.725809,
& 3.743034, 3.760357, 3.777779, 3.795300, 3.812921, 3.830645,
& 3.848470, 3.866400, 3.884434, 3.902574, 3.920821, 3.939176,
& 3.957640, 3.976215, 3.994901, 4.013699, 4.032612, 4.051639,
& 4.070783, 4.090045, 4.109425, 4.128925, 4.148547, 4.168292,
& 4.188160, 4.208154, 4.228275, 4.248524, 4.268903, 4.289413,
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& 4.436781, 4.458395, 4.480154, 4.502060, 4.524114, 4.546319,
& 4.568676, 4.591187, 4.613854, 4.636678, 4.659662, 4.682807,
& 4.706116, 4.729590, 4.753231, 4.777041, 4.801024, 4.825179,
& 4.849511, 4.874020, 4.898710, 4.923582, 4.948639, 4.973883,
& 4.999316, 5.024942, 5.050761, 5.076778, 5.102993, 5.129411,
& 5.156034, 5.182864, 5.209903, 5.237156, 5.264625, 5.292312,
& 5.320220, 5.348354, 5.376714, 5.405306, 5.434131, 5.463193,
& 5.492496, 5.522042, 5.551836, 5.581880, 5.612178, 5.642734,
& 5.673552, 5.704634, 5.735986, 5.767610,
& 5.799512, 5.831694, 5.864161, 5.896918, 5.929968, 5.963316,
& 5.996967, 6.030925, 6.065194, 6.099780, 6.134687, 6.169921,
& 6.205486, 6.241387, 6.277630, 6.314220, 6.351163, 6.388465,
& 6.426130, 6.464166, 6.502578, 6.541371, 6.580553, 6.620130,
& 6.660109, 6.700495, 6.741297, 6.782520, 6.824173, 6.866262,
& 6.908795, 6.951780, 6.995225, 7.039137, 7.083525, 7.128398,
& 7.173764, 7.219632, 7.266011, 7.312910, 7.360339, 7.408308,
& 7.456827, 7.505905, 7.555554, 7.605785, 7.656608, 7.708035,
& 7.760077, 7.812747, 7.866057, 7.920019, 7.974647, 8.029953,
& 8.085952, 8.142657, 8.200083, 8.258245, 8.317158, 8.376837,
& 8.437300, 8.498562, 8.560641, 8.623554, 8.687319, 8.751955,

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& 8.817481, 8.883916, 8.951282, 9.019600, 9.088889, 9.159174,
& 9.230477, 9.302822, 9.376233, 9.450735, 9.526355, 9.603118,
& 9.681054, 9.760191, 9.840558, 9.922186, 10.005107, 10.089353,
& 10.174959, 10.261958, 10.350389, 10.440287, 10.531693, 10.624646,
& 10.719188, 10.815362, 10.913214, 11.012789, 11.114137, 11.217307,
& 11.322352, 11.429325, 11.538283, 11.649285,
& 11.762390, 11.877664, 11.995170, 12.114979, 12.237161, 12.361791,
& 12.488946, 12.618708, 12.751161, 12.886394, 13.024498, 13.165570,
& 13.309711, 13.457026, 13.607625, 13.761625, 13.919145, 14.080314,
& 14.245263, 14.414134, 14.587072, 14.764233, 14.945778, 15.131877,
& 15.322712, 15.518470, 15.719353, 15.925570, 16.137345, 16.354912,
& 16.578520, 16.808433, 17.044929, 17.288305, 17.538873, 17.796967,
& 18.062943, 18.337176, 18.620068, 18.912049, 19.213574, 19.525133,
& 19.847249, 20.180480, 20.525429, 20.882738, 21.253102, 21.637266,
& 22.036036, 22.450278, 22.880933, 23.329017, 23.795634, 24.281981,
& 24.789364, 25.319207, 25.873062, 26.452634, 27.059789, 27.696581,
& 28.365274, 29.068370, 29.808638, 30.589157, 31.413354, 32.285060,
& 33.208568, 34.188705, 35.230920, 36.341388, 37.527131, 38.796172,
& 40.157721, 41.622399, 43.202525, 44.912465, 46.769077, 48.792279,
& 51.005773, 53.437996, 56.123356, 59.103894 /)

```

J = 0

30 CONTINUE

IF (SIGMA.LE.0) THEN

WRITE(6,*) "Error in LANDAU: sigma cannot be less than zero!"

GOTO 35

ENDIF

CALL RANDOM_NUMBER(X)

```

!
U = 1000*X
I = NINT(U)
U = U-I
IF ((I.GE.70.D0).AND.(I.LT.800.D0)) THEN
  RANLAN = F(I) + U*(F(I+1)-F(I))
ELSEIF ((I.GE.7.D0).AND.(I.LE. 980.D0)) THEN
  RANLAN = F(I) + U*(F(I+1)-F(I)-0.25*(1-U)*(F(I+2)-F(I+1)-F(I)
&           +F(I-1)));
ELSEIF (I.LT.7.D0) THEN
  V = LOG(X)
  U = 1/V
  RANLAN = ((0.99858950+(3.45213058E1+1.70854528E1*U)*U)/
&           (1           +(3.41760202E1+4.01244582 *U)*U))*(
&           (-LOG(-0.91893853-V)-1));
ELSE
  U = 1-X
  V = U*U
  IF (X.LE.0.999D0) THEN
    RANLAN = (1.00060006+2.63991156E2*U+4.37320068E3*V)/
&           ((1           +2.57368075E2*U+3.41448018E3*V)*U)
  ELSE
    RANLAN = (1.00001538+6.07514119E3*U+7.34266409E5*V)/
&           ((1           +6.06511919E3*U+6.94021044E5*V)*U);
  ENDIF
ENDIF
LANDAU = MEAN+SIGMA*RANLAN

```

```

J = J+1

!REPICK AT MOST 100 TIMES

IF ((ABS(LANDAU).GT.ABS(DEMAX)).AND.(J.LE.100)) GOTO 30

35  CONTINUE

RETURN

END

! ****
FUNCTION GAURAN(AVG,SIG)

! SIMPLE GAUSSIAN GENERATOR

IMPLICIT DOUBLE PRECISION (A-H,O-Z)

IF (SIG.EQ.0) THEN

GAURAN = AVG

RETURN

ENDIF

45  CONTINUE

CALL RANDOM_NUMBER(R1)

IF (R1.EQ.0) GOTO 45

CALL RANDOM_NUMBER(R2)

R2 = R2*6.283185

GAURAN = AVG+SIG*SIN(R2)*SQRT(-2*LOG(R1))

RETURN

END

! ****

SUBROUTINE GAUVEC(IPV,IPN,IV,IN)

! SIMPLE GAUSSIAN GENERATOR. OUTPUT IS ARRAY X.

IMPLICIT DOUBLE PRECISION (A-H,O-Z)

INTEGER I,N

```

```

! -----MEMORY MANAGEMENT -----
PARAMETER(LMEM=1000000000, LVAR=100000000, LDIM=10000)

INTEGER NTYP(LVAR), NBEG(LVAR), NEND(LVAR), NMAX(LVAR),
*      NC(LMEM), NDIM(LDIM)

DOUBLE PRECISION CC(LMEM)

COMMON NTYP,NBEG,NEND,NMAX, CC,NC, NDIM, IDIM, IVAR, IMEM
COMMON /TYID/ NRE,NST,NLO,NCM,NVE,NIN,NIV,NDA,NCD,NTM,NGR

!
!     CONSISTENCY CHECKS AND PREPARATION
!
! ****
IF(NTYP(IN) .NE. NRE) CALL FOXNTY(IN)
IF(NTYP(IPN). NE. NRE) CALL FOXNTY(IPN)

!
!     GET PARTICLE VARS
!
! ****
N    = NINT(CC(NBEG(IN)))
PN   = NINT(CC(NBEG(IPN)))
DO 7 I=1,N
    DO 6 J=1,PN
        ! V(I,J)=GAURAN(PV(J),PV(J+1))
        AVG=CC(NBEG(IPV+2*J-1))
        SIG=CC(NBEG(IPV+2*J))
        CC(NBEG(IV+J)+I-1)=GAURAN(AVG,SIG)
6    CONTINUE
7    CONTINUE

!

```

```

RETURN

END

! ****
SUBROUTINE GAUS(IAV,ISI,IX,IN)
! NOTE: DOES NOT WORK
! SIMPLE GAUSSIAN GENERATOR. OUTPUT IS ARRAY X.
IMPLICIT DOUBLE PRECISION (A-H,O-Z)
INTEGER I,N
!-----MEMORY MANAGEMENT -----
PARAMETER(LMEM=1000000000,LVAR=100000000,LDIM=10000)
INTEGER NTYP(LVAR),NBEG(LVAR),NEND(LVAR),NMAX(LVAR),
* NC(LMEM),NDIM(LDIM)
DOUBLE PRECISION CC(LMEM)
COMMON NTYP,NBEG,NEND,NMAX, CC,NC, NDIM, IDIM, IVAR, IMEM
COMMON /TYID/ NRE,NST,NLO,NCM,NVE,NIN,NIV,NDA,NCD,NTM,NGR
!
! CONSISTENCY CHECKS AND PREPARATION
! ****
IF(NTYP(IAV).NE.NRE) CALL FOXNTY(IAV)
IF(NTYP(ISI).NE.NRE) CALL FOXNTY(ISI)
IF(NTYP(IN) .NE.NRE) CALL FOXNTY(IN)

!
! GET PARTICLE VARS
! ****
AVG = CC(NBEG(IAV))
SIG = CC(NBEG(ISI))
N = NINT(CC(NBEG(IN)))

```

```

!
DO 47 I=1,N
IF (SIG.EQ.0) THEN
X=AVG
ELSE
46    CONTINUE
CALL RANDOM_NUMBER(R1)
IF (R1.EQ.0) GOTO 46
CALL RANDOM_NUMBER(R2)
R2 = R2*6.283185
X = AVG+SIG*SIN(R2)*SQRT(-2*LOG(R1))
ENDIF
CC(NBEG(IX+1)+I-1) = X
47    CONTINUE
RETURN
END
!
*****FUNCTION SCATDIST(TH0,P)
IMPLICIT DOUBLE PRECISION (A-H,O-Z)
DOUBLE PRECISION MASS
INTEGER I,J,N
A = 0.5/(1-COS(TH0))
MASS = 105.6583715
E = SQRT(P**2+MASS**2)
GAM = E/MASS
BETA = SQRT(1-1/GAM**2)
BG2 = (BETA*GAM)**2

```

```

XI1 = 4.5d0 !EMPERICAL CUTOFF BETWEEN DISTRIBUTIONS
D = 2.d0      !POWER IN DENOMINATOR (I.E. ~ 1/U^D)
U0 = 1-XI1/A
A1 = -A*BG2 !SEE THESIS
A2 = BG2*(1-D)-2*A
A3 = BG2*(A-D-1)+2*(A-D)
B = U0+(A2+SQRT(A2**2-4*A1*A3))/(2*A1)
ZETA = EXP(-A*(1-U0))*(1-U0+B)**D/(1+0.5*BG2*(1+U0-B))
G0 = ZETA*((1+BG2)*(1/(1-U0+B)-1/(2+B))+BG2/2*LOG((1-U0+B)/
& (2+B)))
GMAX = G0+(1-EXP(-A*(1-U0)))/A
CALL RANDOM_NUMBER(G)
G = G*GMAX
!
! USE GAUSSIAN -GENERATION VIA INVERSION
IF (G.GT.G0) THEN
  U = 1+LOG(A*(G-G0)+EXP(-A*(1-U0)))/A
!
! USE MOTT -GENERATION VIA BISECTION METHOD
ELSE
  ACC = 1.D-8           !+/- BINWIDTH OF g (ACCURACY)
  UMAX = U0
  UMIN = -1
80  CONTINUE
  U = (UMAX+UMIN)/2
!
! 'G' UPPER-CASE' GENERATOR-CDF NOT PDF
  TEST = GUC_TAIL(U,ZETA,B,BG2)
  IF (TEST.GT.G*(1+ACC)) UMAX = U

```

```

    IF (TEST.LT.G*(1-ACC)) UMIN = U
    IF (NOT((TEST.GE.G*(1-ACC)).AND.(TEST.LE.G*(1+ACC)))) GOTO 80
  ENDIF
85  CONTINUE
!
SCATDIST = P*U
RETURN
END
! ****
FUNCTION GLC_TAIL(U,ZETA,B,BG2)
! RETURNS THE TAIL OF THE PDF OF THE SCATTERING FUNCTION
! GIVEN THE SCALE, SHIFT, AND ENERGY.
IMPLICIT DOUBLE PRECISION (A-H,O-Z)
GLC_TAIL = ZETA*(1+0.5*BG2+(1+U-B))/((1-U+B)**2)
RETURN
END
! ****
FUNCTION GUC_TAIL(U,ZETA,B,BG2)
! RETURNS THE TAIL OF THE CDF OF THE SCATTERING FUNCTION
! GIVEN THE SCALE, SHIFT, AND ENERGY.
IMPLICIT DOUBLE PRECISION (A-H,O-Z)
GUC_TAIL = ZETA*((1+BG2)*(1/(1-U+B)-1/(2+B))+BG2/2*LOG
& ((1-U+B)/(2+B)))
RETURN
END
! ****
REAL FUNCTION LATDIS(THETA,Z,THC)

```

```

!      RETURNS TRANSVERSE COORDINATE. BASED ON FERNOW AND
!      GALLARDO, "Muon transverse ionization cooling:
!      Stochastic approach", PHYS. REV. E 52 1039 (1995).
!      THETA IS FLUCTUATION (I.E. THETA-THETA0), Z IS ABS
!      LENGTH, AND THC IS CRITICAL ANGLE COEFFICIENT.

IMPLICIT DOUBLE PRECISION (A-H,O-Z)

DOUBLE PRECISION MU

MU = THETA*Z/1.8

SIG = THETAC/(2*SQRT(3.))

LATDIS = GAURAN(MU,SIG)

RETURN

END

! ****

FUNCTION DTOF(EI,P2,MASS,L)

!      RETURNS THE TEMPORAL DISPLACEMENT CORRECTION ASSUMING
!      AN INITIAL STRAIGHT TRAJECTORY (I.E. PZ = P).

IMPLICIT DOUBLE PRECISION (A-H,O-Z)

DOUBLE PRECISION L,MASS

C = 299792458.D0           ! m/s

P1 = SQRT(EI**2-MASS**2)

DTOF = (P2/SQRT(P2**2+MASS**2)-P1/SQRT(P1**2+MASS**2))

&      *2*L/(C*(P2**2/(P2**2+MASS**2)-P1**2/(P1**2
&      +MASS**2)))

RETURN

END

! ****

INTEGER FUNCTION DECAY(GAM1,GAM2,TOF)

```

```

!      RETURNS 4 IF PARTICLE DECAY OCCURED OR ZERO OTHERWISE.

IMPLICIT DOUBLE PRECISION (A-H,O-Z)

TAU = 2.197034D-6          !MEAN LIFETIME IN [s]

GAM = (GAM1+GAM2)/2        !AVG GAMMA; PROBLEM FOR LARGE STEPS

T    = TOF/GAM

CALL RANDOM_NUMBER(R1)

TEST = EXP(-T/TAU)

DECAY = 0

IF (R1.GT.TEST) DECAY = 4

RETURN

END

! ****

FUNCTION GET_DCP(ETA,MP)

IMPLICIT DOUBLE PRECISION(A-H,O-Z)

DOUBLE PRECISION KDE, MP(6)

! ADOPT DENSITY EFFECT PARAMETER FROM PDG:

! pdg.web.cern.ch/pdg/2013/AtomicNuclearProperties/adndt.pdf

! I HAVE YET TO IMPLEMENT THE 'CONDUCTORS' PORTION

HWP = 28.816*SQRT(MP(3)/(1.E6)*MP(1)/MP(2))*1.E-6 ! PLAMSA ENERGY

X   = LOG10(ETA)

C   = 2*LOG(MP(4)/HWP)+1

X0 = 0.2D0

KDE = 3.D0      !K FOR THE DENSITY EFFECT

! DETERMINE X1, X0 FOR SOLIDS, LIQUIDS:

IF (MP(4).GE.0.0001D0) THEN

X1 = 3.D0

IF (C.GE.5.215D0) X0 = 0.326*C-1.5

```

```
ELSE  
X1 = 2.D0  
IF (C.GE.3.681D0) X0 = 0.326*C-1.0  
ENDIF  
! DETERMINE X1, X0 FOR GASSES:  
! YET TO IMPLEMENT STATE DISTINCTION FOR COSY!  
! *****  
! X1 = 4.D0 !DEFAULT  
! IF (C.GE.13.804D0) THEN  
! X0 = 0.326*C-1.5  
! X1 = 5.D0  
! ELSEIF (C.GE.12.25) THEN  
! X0 = 2.D0  
! X1 = 5.D0  
! ELSEIF (C.GE.11.50) THEN  
! X0 = 2.D0  
! ELSEIF (C.GE.11.00) THEN  
! X0 = 1.9D0  
! ELSEIF (C.GE.10.50)  
! X0 = 1.8D0  
! ELSEIF (C.GE.10.00)  
! X0 = 1.7D0  
! ELSE  
! X0 = 1.6D0  
! ENDIF  
! ADE = (C-2*LOG(10.)*X0)/((X1-X0)**3)  
!
```

```

!      "TABULATED DATA" FOR HYDROGEN IS AT BUBBLE-CHAMBER CONDITIONS.

!      THEREFORE A CORRECTION MUST BE MADE:

R = 1

IF (MP(1)/MP(2).EQ.1.D0/1.008D0) R = MP(3)/(0.060*1.D6)

C = C-LOG(R)

X0 = X0-0.5*LOG10(R)

X1 = X1-0.5*LOG10(R)

!

IF (X.GE.X1) THEN

DCP = 2*LOG(10.)*X-C

ELSEIF (X.GT.X0) THEN

DCP = 2*LOG(10.)*X-C+ADE*(X1-X)**KDE

ELSE

DCP = 0.D0

ENDIF

GET_DCP = DCP

RETURN

END

! ****

FUNCTION GET_SCP(ETA, MP)

IMPLICIT DOUBLE PRECISION(A-H,O-Z)

DOUBLE PRECISION BAR(6), MP(6)

! ADOPT SHELL CORRECTION APPROX BY BARKAS.

! THIS APPROXIMATION IS OUT OF DATE, AND IS THEREFORE QUITE ROUGH.

! SEE PDG FOR MORE INFO:

! pdg.web.cern.ch/pdg/2013/AtomicNuclearProperties/adndt.pdf

IF (ETA>0.13D0) THEN !ONLY GOOD FOR ETA>0.13 (MUON: KE = 0.89)

```

```

BAR = (/ 0.422377, 0.0304043, -0.00038106, 3.858019,
&           -0.1667989, 0.00157955 /)

SCP = (BAR(1)/ETA**2+BAR(2)/ETA**4+BAR(3)/ETA**6)*MP(4)**2*1.E6
&           +(BAR(4)/ETA**2+BAR(5)/ETA**4+BAR(6)/ETA**6)*MP(4)**3*1.E9

ENDIF

GET_SCP = SCP

RETURN

END

! *****
REAL FUNCTION MASSBYID(ID)
INTEGER ID
! RETURNS PARTICLE MASS IN MeV BASED ON PDG ID.
! CURRENTLY ONLY SUPPORTS MUONS (13).
IF (ABS(ID).EQ.11) MASSBYID=0.5109989461
IF (ABS(ID).EQ.13) MASSBYID=105.6583715
RETURN
END

! *****
REAL FUNCTION MASSBYIDICOOL(ID)
INTEGER ID
! RETURNS PARTICLE MASS IN MeV BASED ON PDG ID.
! CURRENTLY ONLY SUPPORTS MUONS (2).
IF (ABS(ID).EQ.1) MASSBYIDICOOL=0.5109989461
IF (ABS(ID).EQ.2) MASSBYIDICOOL=105.6583715
RETURN
END

! *****

```

```

SUBROUTINE RG4BL(IFILE, INPART, IV)
IMPLICIT DOUBLE PRECISION(A-H,O-Z)
INTEGER I
DOUBLE PRECISION MASSBYID

!      READS PARTICLES FROM A G4BL-FORMATTED FILE. INPUTS ARE:
!
!      IFILE - THE NUMBER OF THE FILE (E.G. 12 IN 'FORT.12')
!
!      INPART - THE NUMBER OF PARTICLES TO READ
!
!      IV - THE OUTPUT VECTOR
!
!
!      OUTPUT VECTOR IS OF THE FORM
!
!      V = (X,PX,Y,PY,T,E) IN m, MeV, AND s

!-----MEMORY MANAGEMENT -----
PARAMETER(LMEM=1000000000,LVAR=100000000,LDIM=10000)
INTEGER NTYP(LVAR),NBEG(LVAR),NEND(LVAR),NMAX(LVAR),
*          NC(LMEM),NDIM(LDIM)
DOUBLE PRECISION CC(LMEM)

COMMON NTYP,NBEG,NEND,NMAX, CC,NC, NDIM, IDIM, IVAR, IMEM
COMMON /TYID/ NRE,NST,NLO,NCM,NVE,NIN,NIV,NDA,NCD,NTM,NGR

!
!      CONSISTENCY CHECKS AND PREPARATION
!
! ****
IF(NTYP(IFILE).NE.NRE) CALL FOXNTY(IFILE)
IF(NTYP(INPART).NE.NRE) CALL FOXNTY(INPART)
!
!
!      GET VARS
!
! *****
FNAME = NINT(CC(NBEG(IFILE)))

```

```

NPART = NINT(CC(NBEG(INPART)))

!
! GET RID OF FIRST THREE LINES
!
! ****
READ(FNAME,*)
READ(FNAME,*)
READ(FNAME,*)

!
! READ THE REST OF THE FILE
!
! ****
DO 5 I=1,NPART
  READ(FNAME,* ,END=10) X,Y,Z,PX,PY,PZ,T,PDGID,EID,TID,PID,W
  E=SQRT(PX**2+PY**2+PZ**2+MASSBYID(NINT(PDGID))**2)
  CC(NBEG(IV+1)+I-1)=X/1000
  CC(NBEG(IV+2)+I-1)=PX
  CC(NBEG(IV+3)+I-1)=Y/1000
  CC(NBEG(IV+4)+I-1)=PY
  CC(NBEG(IV+5)+I-1)=T
  CC(NBEG(IV+6)+I-1)=E
5   CONTINUE
10  CONTINUE
END
!
! ****
SUBROUTINE RICOOL(IFILE, INPART, IV)
IMPLICIT DOUBLE PRECISION(A-H,O-Z)
INTEGER I
DOUBLE PRECISION MASSBYID

```

```

!      READS PARTICLES FROM AN ICOOL-FORMATTED FILE. INPUTS ARE:
!
!      IFILE - THE NUMBER OF THE FILE (E.G. 12 IN 'FORT.12')
!
!      INPART - THE NUMBER OF PARTICLES TO READ
!
!      IV      - THE OUTPUT VECTOR
!
!
!      OUTPUT VECTOR IS OF THE FORM
!
!      V = (X,PX,Y,PY,T,E) IN m, MeV, AND s
!
!-----MEMORY MANAGEMENT -----
!
PARAMETER(LMEM=1000000000,LVAR=100000000,LDIM=10000)
!
INTEGER NTYP(LVAR),NBEG(LVAR),NEND(LVAR),NMAX(LVAR),
*
*           NC(LMEM),NDIM(LDIM)
!
DOUBLE PRECISION CC(LMEM)
!
COMMON NTYP,NBEG,NEND,NMAX, CC,NC, NDIM, IDIM, IVAR, IMEM
COMMON /TYID/ NRE,NST,NLO,NCM,NVE,NIN,NIV,NDA,NCD,NTM,NGR
!
!
!      CONSISTENCY CHECKS AND PREPARATION
!
!      ****
!
IF(NTYP(IFILE).NE.NRE) CALL FOXNTY(IFILE)
!
IF(NTYP(INPART).NE.NRE) CALL FOXNTY(INPART)
!
!
!      GET VARS
!
!      ****
!
FNAME = NINT(CC(NBEG(IFILE)))
NPART = NINT(CC(NBEG(INPART)))
!
!
!      READ THE FILE
!
!      ****

```

```

DO 5 I=1,NPART

READ(FNAME,*,END=10) PID,PNUM,PDGID,FLAG,T,W,X,Y,Z,PX,PY,PZ,
& SPINX,SPINY,SPINZ

E=SQRT(PX**2+PY**2+PZ**2+(MASSBYIDICOOL(NINT(PDGID))/1000)
& **2)*1000

CC(NBEG(IV+1)+I-1)=X
CC(NBEG(IV+2)+I-1)=PX*1000
CC(NBEG(IV+3)+I-1)=Y
CC(NBEG(IV+4)+I-1)=PY*1000
CC(NBEG(IV+5)+I-1)=T
CC(NBEG(IV+6)+I-1)=E

5   CONTINUE

10  CONTINUE

END

!
! ****
!
SUBROUTINE WICOOL(IZ,IREG,INPART,IV,IUI,IAPE)
!
! WRITES PARTICLES TO UI IN ICOOL-STYLE. INPUTS ARE:
!
! IZ - Z LOCATION
!
! IREG - REGION NUMBER
!
! INPART - NUMBER OF PARTICLES
!
! IV - VECTOR TO WRITE
!
! IUI - UI TO WRITE TO
!
! IAPE - APERATURE
!
!
! VECTOR IV MUST BE OF THE FORM
!
! V = (X,PX,Y,PY,T,E) IN m, MeV, AND s

```

```

!
IMPLICIT DOUBLE PRECISION(A-H,O-Z)

DOUBLE PRECISION MASS

DOUBLE PRECISION, DIMENSION(:,:),
ALLOCATABLE :: V ! V(NPART,6)

INTEGER IALLOCATESTATUS,I,J

!-----MEMORY MANAGEMENT -----
PARAMETER(LMEM=1000000000,LVAR=100000000,LDIM=10000)

INTEGER NTYP(LVAR),NBEG(LVAR),NEND(LVAR),NMAX(LVAR),
* NC(LMEM),NDIM(LDIM)

DOUBLE PRECISION CC(LMEM)

COMMON NTYP,NBEG,NEND,NMAX, CC,NC, NDIM, IDIM, IVAR, IMEM
COMMON /TYID/ NRE,NST,NLO,NCM,NVE,NIN,NIV,NDA,NCD,NTM,NGR

!
CONSISTENCY CHECKS AND PREPARATION

*****
IF(NTYP(IZ).NE.NRE) CALL FOXNTY(IAV)
IF(NTYP(IREG).NE.NRE) CALL FOXNTY(ISI)
IF(NTYP(INPART).NE.NRE) CALL FOXNTY(IN)
IF(NTYP(IUI).NE.NRE) CALL FOXNTY(IUI)
IF(NTYP(IAPE).NE.NRE) CALL FOXNTY(IAPE)

!
GET VARS
*****
Z = CC(NBEG(IZ))
REG = NINT(CC(NBEG(ISI)))
NPART = NINT(CC(NBEG(INPART)))
UI = NINT(CC(NBEG(IUI)))

```

```

APE = CC(NBEG(IAPE))

MASS = 105.6583715

ALLOCATE(V(NPART+10,7),Stat=IALLOCATESTATUS)

IF (IALLOCATESTATUS /= 0) STOP "*** Not enough memory ***"

!

DO 7 I=1,NPART

X =CC(NBEG(IV+1)+I-1)

PX=CC(NBEG(IV+2)+I-1)

Y =CC(NBEG(IV+3)+I-1)

PY=CC(NBEG(IV+4)+I-1)

T =CC(NBEG(IV+5)+I-1)

E =CC(NBEG(IV+6)+I-1)

P =SQRT(E**2-MASS**2)

PZ=SQRT(P**2-PX**2-PY**2)

! MAKE SURE PARTICLE ISN'T LOST

TOL=1E10

IF (APE.LE.0) APE=TOL

IF ((X.GE.TOL).OR.(Y.GE.TOL).OR.(T.GE.TOL).OR.(PZ.GE.TOL).OR.

&      (SQRT(X**2+Y**2).GE.APE)) THEN

T =0

X =0

Y =0

Z =-1

PX=0

PY=0

PZ=-1

ENDIF

```

```
      WRITE(U1,20) I,1,2,0,1,T,X,Y,Z,PX,PY,PZ,0,0,0,1,0,0,0,0,0,0,0  
7      CONTINUE  
20      FORMAT(I,4I2,7E22.14,11I2)  
      END
```

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