

Assignment 1—Group:

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Abstract

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1 Introduction

Non-negative matrix factorization (NMF) is a matrix decomposition technique that approximates a data matrix with non-negative entries with the multiplication of two non-negative matrices

$$V \approx WH.$$

In this approximation, matrix W is the basis and matrix H is the weight matrix corresponding to the new dictionary W . As NMF only is applicable to an extensive range of domain. [Lee and Seung \[1999\]](#) suggest that NMF is useful for image processing problems including facial recognition.

Matrices W and H are often referred as the basis images and weights. This is because the observed image V is approximated by a linear combination of W with positive coefficients H . Due to the additive nature of the algorithm, the dictionary W are often parts of images. This property also distinguishes NMF from other traditional image processing methods such as principal components analysis (PCA). [Guillamet and Vitrià \[2002\]](#) demonstrate that NMF performs better in image classification problems in comparison with PCA.

Moreover, NMF is also applicable to text mining such as semantic analysis. Generally, NMF is useful to discover semantic features of an article by counting the frequency of each word, and then approximating the document from a subset of a large array of features [[Lee and Seung, 1999](#)].

In practice, face images could be easily corrupted during data collection by noise with large magnitude. Corruption may result from lighting environment, facial expression or facial details. An NMF algorithm that is robust to large noise is desired for real-world application. Therefore, the objective of this project is to analyse the robustness of NMF algorithms on corrupted datasets. We implement two NMF algorithms designed by [Lee and Seung \[2001\]](#) on real face image datasets, ORL dataset and Extended YaleB dataset. We add artificial noises to the face images are contaminated.

The rest of the report is organized as follows. We describe noisy design and two NMF algorithms including Euclidean Distance and Kullback-Leibler Divergence (KLD) in Section 2. Section 3 shows experiment setup and empirical results which

demonstrate the robustness of the two NMF algorithms. The conclusions and future work are discussed in Section 5.

2 Related work

Researchers proposed many NMF algorithms. As the objective function of NMF is non-convex, for which the traditional gradient decent method could be very sensitive to step sizes and converge slowly, [Lee and Seung \[2001\]](#) first proposes to algorithms which minimises Euclidean distance or Kullback-Leibler divergence (KLD) between the original matrix and its approximation. Although this algorithm is easy to implement and have reasonable convergent rate [[Lee and Seung, 2001](#)], it may fail on seriously corrupted dataset which violates its assumption of Gaussian noise or Poisson noise, respectively [[Guan et al., 2017](#)]. Moreover, [Yang et al. \[2011\]](#) indicate that KLD is sensitive to the initial value of matrix factors and requires many iterations to retrieve from wrong initial values. To improve the robustness of NMF, many methods have been proposed. [Lam \[2008\]](#) proposes L_1 -norm based NMF to model noisy data by a Laplace distribution which is less sensitive to outliers. However, as L_1 -norm is not differentiable at zero, the optimization procedure is computationally expensive. [Kong et al. \[2011\]](#) proposed an NMF algorithm using L_{21} -norm loss function which is robust to outliers. The updating rules used in L_{21} -norm NMF, however, converge slowly because of a continual use of the power method [[Guan et al., 2017](#)].

Apart from different loss functions, several optimization methods were proposed to improve the performance of NMF. After [Lee and Seung \[2001\]](#) proposed multiplicative update rules MUR, [Lin \[2007\]](#) proposed a modified MUR which guaranteed the convergence to a stationary point. This modified MUR, however, did not improve the convergence rate of traditional MUR [[Guan et al., 2012](#)]. Moreover, as MUR is not able to shrink all entries in matrix factors to zero, [Berry et al. \[2007\]](#) proposed a projected nonnegative least square (PNLS) method to overcome this problem. Although, in each nonnegative least square (NLS) sub-problem, the least squares solution is directly projected to the negative quadratic, PNLS does not guarantee convergence [[Guan et al., 2012](#)]. In contrast to these gradient-based optimization methods, [Kim and Park \[2008\]](#) presented an active set method AS which divides variables into two sets, free set and active set and update free set in each iteration by solving unconstrained equation. Even though AS has good convergence rate, it assumes strictly convexity in each NLS sub-problem [[Kim and Park, 2008](#)]. These assumptions are easily violated in real applications.

3 Methods

Some carefully designed NMF are robust to various noises. These robust algorithms aim to significantly reduce the amount of noise while preserving the edges without blurring the image [Barbu, 2013]. Figure 4.1.4 shows three kinds of noises we designed, including Gaussian noise, Poisson noise, and Salt & Pepper noise.

3.1 NMF and Gaussian noise

Gaussian noise is a noise with a probability density function being normal with mean zero. Lee and Seung [2001] propose the first NMF with the objective function between image V and its NMF factorisation W and H being

$$\|V - WH\| = \sum_{ij} [V_{ij} - (WH)_{ij}]^2. \quad (1)$$

To minimise this object function of least square, Lee and Seung [2001] prove the convergence of the multiplication update rule

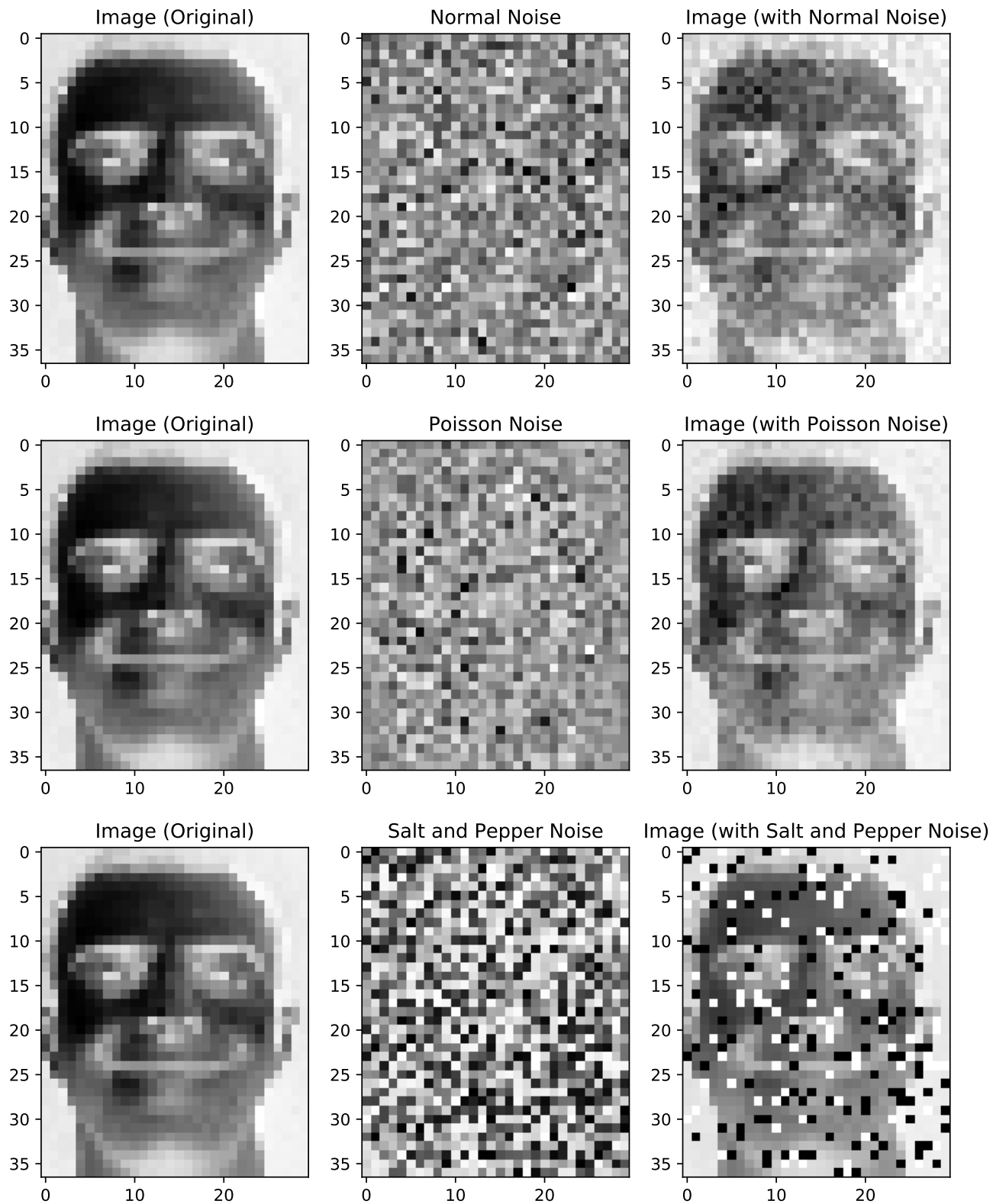
$$H_{jk} \leftarrow H_{jk} \frac{(W^T V)_{jk}}{(W^T WH)_{jk}} \text{ and } W_{ij} \leftarrow W_{ij} \frac{(VH)_{ij}}{(WHH^T)_{ij}}. \quad (2)$$

Here, $()_{ij}/()_{ij}$ denotes elementwise division of the two matrix. Liu and Tao [2016] shows this NMF algorithm minimises Gaussian.

3.2 KLNMF and Poisson noise

Poisson noise or shot noise is a type of electronic noise that occurs when the finite number of particles that carry energy, such as electrons in an electronic circuit or photons in an optical device, is small enough to give rise to detectable statistical fluctuations in a measurement. Lee and Seung [2001] suggest that KLNMF is a algorithm that minimising the Kullback-Leibler divergence

$$\begin{aligned} D(V||WH) &= \sum_{ij} \left(V_{ij} \log \frac{V_{ij}}{(WH)_{ij}} - V_{ij} + (WH)_{ij} \right) \\ &= \sum_{ij} \left(-V_{ij} \log (WH)_{ij} + (WH)_{ij} + C(V_{ij}) \right). \end{aligned} \quad (3)$$



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Figure 1: The original images (left) are corrupted by Gaussian Noise (top), Poisson Noise (middle), and Salt & Pepper Noise (bottom). The corrupted images are shown on the right.

where $C(V_{ij}) = V_{ij} \log V_{ij} - V_{ij}$. $C(V_{ij})$ is a function of the observed image matrix V only. Lee and Seung [2001] also suggest a multiplication update rule to find as the optimisation procedure of KLNMF

$$H_{jk} \leftarrow H_{jk} \frac{\sum_i W_{ij} V_{ik} / (WH)_{jk}}{\sum_{i'} W_{i'j}} \text{ and } W_{ij} \leftarrow W_{ij} \frac{\sum_k H_{jk} V_{ik} / (WH)_{jk}}{\sum_{k'} H_{ik'}}. \quad (4)$$

As this original image matrix V is observed, minimising this Kullback-Leibler divergence (3) is equivalent to minimising

$$\sum_{ij} \left(-V_{ij} \log (WH)_{ij} + (WH)_{ij} + C(V_{ij}) \right).$$

, for arbitrary bounded function $C(V_{ij})$. Taking exponential of the negative of this score function, the problem transforms to maximising the following likelihood function

$$L(WH|V) = \prod_{ij} \left((WH)_{ij}^{V_{ij}} e^{-(WH)_{ij}} + C(V_{ij}) \right).$$

Choosing constant $C(V_{ij})$ to be $-\log V_{ij}!$ gives

$$L(WH|V) = \prod_{ij} \left(\frac{(WH)_{ij}^{V_{ij}} e^{-(WH)_{ij}}}{V_{ij}!} \right).$$

Hence, the probability density function of each element of the original matrix V is Poisson

$$P(V_{ij}) = \frac{(WH)_{ij}^{V_{ij}} e^{-(WH)_{ij}}}{V_{ij}!}$$

is a sufficient condition to yield this likelihood. Hence KLNMF is most suitable for images with Poisson noise.

3.3 Gaussian and Poisson are asymptotic equivalent

We design an Gaussian noise and a Poisson noise with different magnitude. Poisson distribution with parameter λ (integer) is equivalent to the sum of λ Poisson distributions with parameter 1 [Walck, 1996, p. 45]. Hence for λ large, Central Limit Theorem implies that Poisson distribution with parameter λ is well approximated by $N(\lambda, \lambda)$. When applying Poisson noise to an image, we do not degree of freedom to choose any parameter. The variance is the magnitude of the pixels. To compare the robustness of KLNMF with NMF with different noise, we choose the



Figure 2: Compare a Gaussian noise $N(0, 40)$ with Poisson noise $\text{Poi}(40) - 40$. They two distributions are asymptotically equivalent and have similar density functions.

variance of Gaussian noise to be the different from the magnitude of the pixel, that is, $N(0, \text{Var}) \neq N(0, V) \approx \text{Poi}(V) - V$. Figure 2 visualises the similarity of Poisson distribution and Normal distribution with parameter $V = 40$. To overcome this issue, the Gaussian noise we use should be larger or smaller variance in comparison with the mean of the images. We choose the variance of the noise to be 80.

3.4 Salt & Pepper noise

Apart from Gaussian and Poisson noises, we also test our two algorithms on the commonly seen **Salt & Pepper noise**. The noise presents itself by having dark pixels in bright regions and bright pixels in dark regions [Sampat et al., 2005].

3.5 Multiple initial estimates avoid local minima

As discussed in the section of related work. The problem of nonnegative matrix factorization is not a convex problem. Hence the results update rules (2) and (4) coverage to may be local minima instead of global minima. To address this

issue, we implement several (i.e. n) initial estimates for each matrix factorization problem. We use the factorized matrices W and H corresponding to the least residual (1) and (3), for NMF and KLNMF, algorithms respectively, as the final result of factorization. This design of multiple starting point improves the stability of algorithm, but it requires more computational power. To improve the computational speed, we make the number n equal to the number of cores of the CPU. We assign each of the n initial estimates randomly with an uniform distribution. Then these each of the n initial estimates are assigned to a different core of the CPU. This boost the CPU utilisation to 100% instantly and improved the computational speed by 70% on the ORL data. The following part of our code implements this idea of parallel computing.

Algorithm 1: Centring image data

```
1 args = zip(repeat(V,ncpu), repeat(r,ncpu),
            repeat(niter[name2],ncpu), repeat(min_error[name2],ncpu))
2 result = pool.starmap(algo, args)
```

where algo is the NMF algorithm and niter is the number of iterations.

3.6 KLNMF requires more iterations

KLNMF converges slower than NMF. We made the plot of the logarithm of residual (1) and (3) with respect to the logarithm of the number of iterations. We measured the slope of the log-log plot for both algorithms and found that the slope of the NMF residual plot is 2.5 times as that of the KLNMF plot (Figure 3). This estimates that the rate of convergence of NMF is 2.5 faster than KLNMF. As a result, we set the number of iterations as 500 and 1200 for NMF and KLNMF algorithms, respectively (i.e. roughly 2.5 more iterations).

3.7 Evaluation metrics and their confidence intervals

The assignment instruction asks us to compare the performance of NMF and KLNMF by using evaluations metrics including Relative Reconstruction Errors (RRE), Average Accuracy (AA), Normalized Mutual Information (NMI). The instruction states the formulae of these metrics. However, to systematically compare the metrics, we construct an 95% confidence interval for any metric (e.g. RRE) by Studentized bootstrapping method

$$(RRE - t_{97.5}^* \cdot \hat{se}_{RRE}, RRE - t_{2.5}^* \cdot \hat{se}_{RRE}), \quad (5)$$

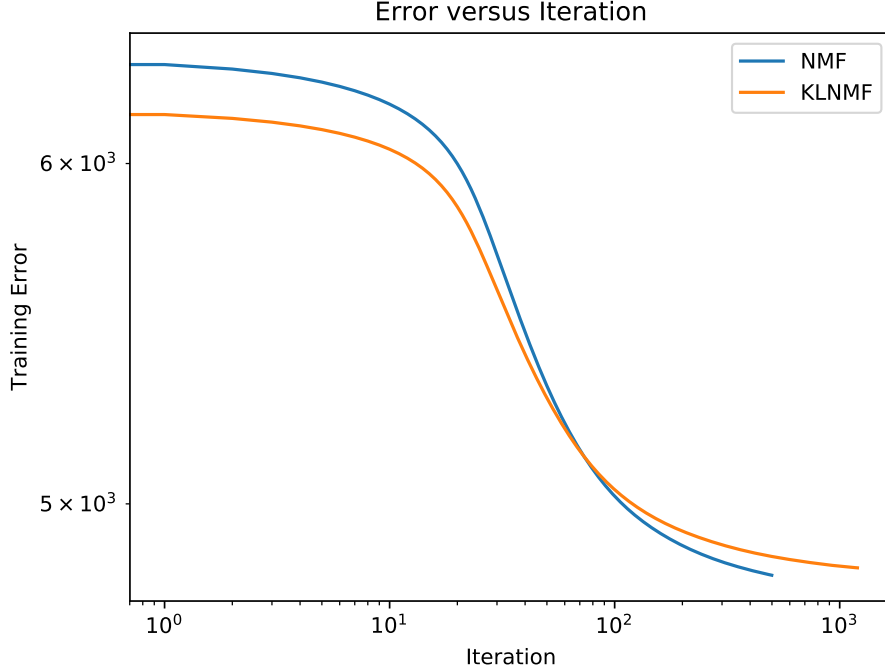


Figure 3: Residual of objective function (1) and (3) versus the number of iterations. NMF converges more than twice faster than KLNMF.

where $t_{2.5}^*$ is the 2.5 percentile of the bootstrapped Student's t-distribution from our sample space with 80 simulations and \hat{s}_{RRE} is the sample standard error. We run 80 simulations so that the confidence interval we construct is precise.

Boostrapping does require the sample space follows specific distributions. We apply this nonparametric method to construct confidence interval here because we do not have the knowledge about the specific distributions of these three metrics.

3.8 Compare the robustness of algorithms

We implement Kolmogorov-Smirnov test to test the hypothesis that the algorithms NMF and KLNMF have different robustness. Again Kolmogorov-Smirnov test is distribution free, so we do not need to know the distributions of the evaluation metrics.

Let RRE_i denote the RRE generated from our NMF algorithm by the i th simulation. Define the empirical distribution of a sample set generated by algorithm α , perhaps

the 80 RRE results, as

$$\hat{F}_\alpha(x) = \frac{1}{n} \sum_{i=1}^{80} 1_{\text{RRE}_i \leq x}. \quad (6)$$

The test statistic is the supremum among the differences of the empirical distribution generated using definition (6) [Walck, 1996]

$$D = \sup_x |F_{\alpha_2}(x) - F_{\alpha_1}(x)|. \quad (7)$$

We compare the test statistics D with the critical value of 0.215, which corresponds to 80 samples and a 95% of confidence level. We reject the null hypotheses that the two algorithms produce similar RRE (or other evaluation metrics) with 95% confidence level if the test statistics $D > 0.215$. The technique is extended to compare AA and NMI.

4 Experiments

4.1 Noise

We implement three kinds of noises including Gaussian noise, Poisson noise and Salt & Pepper noise.

4.1.1 Gaussian Noise

We design the Gaussian noise by normal distribution with 0 mean and 80 standard deviation (Algorithm 2). The ORL dataset has a global pixel mean of 40 and the CroppedYale data set has that of 70. Hence the designed Gaussian noise corrupt the images significantly. We choose the standard deviation to be 80 so that our Gaussian noise are less likely to coincident with the designed Poisson noise.

Algorithm 2: Gaussian Noise Design

```

1 def normal(subVhat):
2     """Design a Gaussian noise."""
3     V_noise = np.random.normal(0, 80, subVhat.shape) #*
4               np.sqrt(subVhat)
5     V = subVhat + V_noise
6     V[V < 0] = 0
7     return V, V_noise

```

4.1.2 Poisson Noise

The Poisson noise is not additive and has no hyperparameters to be set. We draw the corrupted image from Poisson distribution and then the Poisson noise is calculated from the difference between the corrupted image and the original image, as discussed in Section 3 and demonstrated in algorithm 3.


Algorithm 3: Poisson Noise Design

```

1 def possion(subVhat):
2     """Design a Possion noise."""
3     V = np.random.poisson(subVhat)
4     V_noise = V-subVhat
5     return V, V_noise

```

4.1.3 Salt and Pepper Noise

Salt and Pepper noises are added by drawing random integers from discrete uniform distribution of the interval $[0, 255)$  ⁴

Algorithm 4: Salt and Pepper Noise Design

```

1 """Design a salt and pepper noise where make some pixel value
   zeros."""
2 # obtain one Image
3 image = subVhat[:, 0]
4 V_noise = np.random.randint(low=0, high=255,
   size=subVhat.shape, dtype=int)
5 V = subVhat.copy()
6 V[V_noise <= 20] = 0
7 V[V_noise >= 230] = 255
8 return V, V_noise

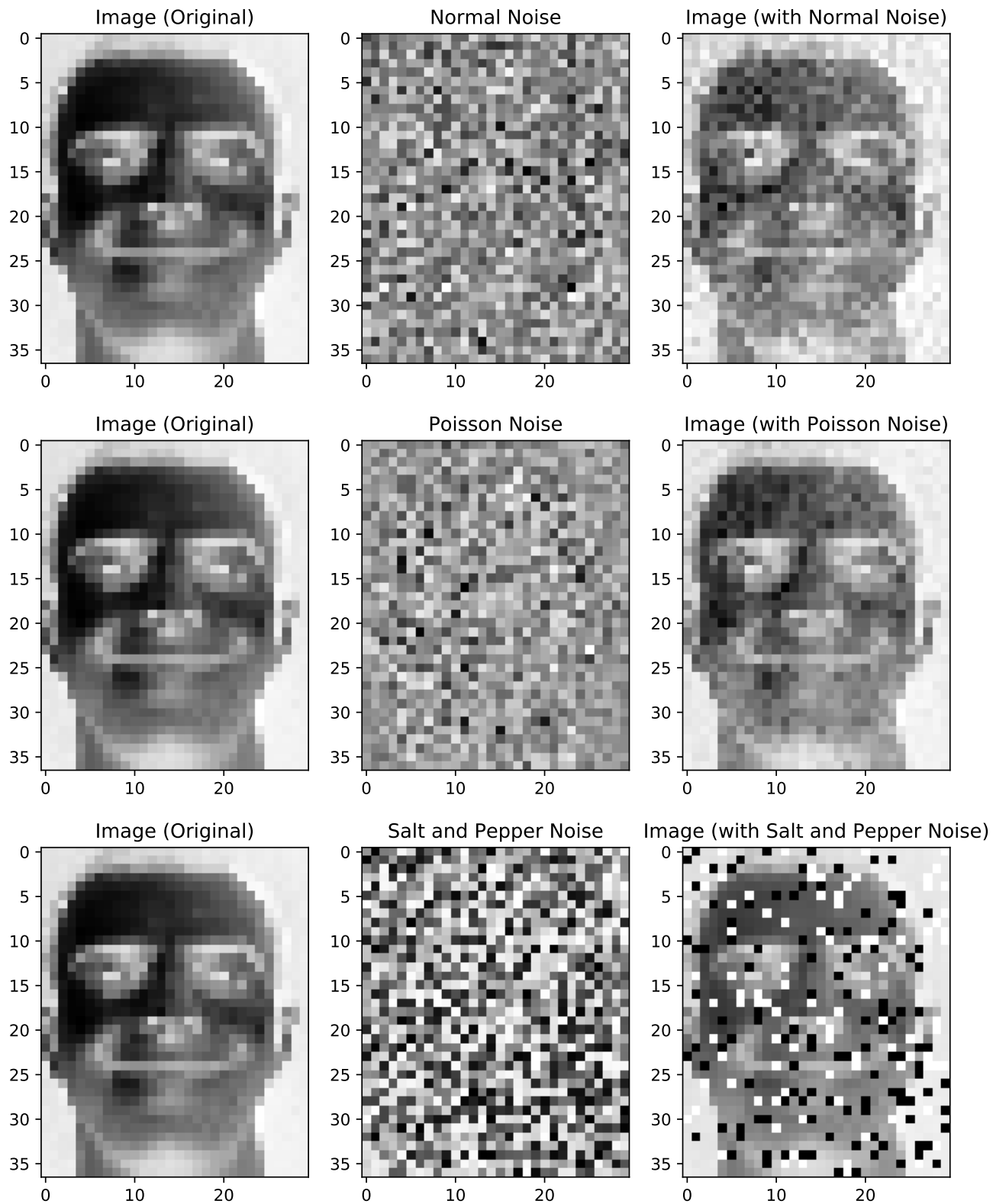
```

4.2 Experiments Results

The results of noise design are shown below:

5 Conclusion

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Figure 4: The original images (left) are corrupted by Gaussian Noise (top), Poisson Noise (middle), and Salt & Pepper Noise (bottom). The corrupted images are shown on the right.

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A Appendix

Algorithm 5: The Levenberg–Marquardt algorithm iteratively finds optimal macro-scale Robin boundary conditions.

```

1 %A=imread('C:\Users\chenc\OneDrive - UNSW\machine
   learning\assignment
   1\data\CroppedYaleB\yaleB01\yaleB01_P00A+000E+00.pgm');
2 red=3;
3 k=40;
4
5 imagefiles = dir('data/ORL/*/*.pgm');
6 imagefiles2=struct2cell(imagefiles);
7 imagefiles=imagefiles((~endsWith(imagefiles2(1,:), 'Ambient.pgm'))');
8 imagefiles2=struct2cell(imagefiles);
9 A=imread(strcat(imagefiles(1).folder, '\', imagefiles(1).name));
10 if size(A,1)==112
11     A=A(1:111,1:90);
12 end
13 A_list=zeros(size(A,1)/red, size(A,2)/red, red);
14 nfiles = length(imagefiles); % Number of files found
15 matrix_image=zeros(prod(size(A))/red^2, nfiles);
16 temp=struct2cell(imagefiles);
17 names=temp(1,:);
18 for ii=1:nfiles
19     currentfilename = imagefiles(ii).name;
20     currentfilename=strcat(imagefiles(ii).folder, '\', currentfilename);
21     currentimage = imread(currentfilename);
22     if abs(size(A,1)-112)<=1
23         currentimage=currentimage(1:111,1:90);
24     end
25
26     for i=1:red
27         A_list(:, :, i)=currentimage(i:red:end, i:red:end);
28     end
29     A2=uint8(mean(A_list, 3));
30     matrix_image(:, ii) = A2(:);
31 end
32
33 [w h]=NeNMF(matrix_image, k);
34 idx = kmeans(h', k)
35 Y_pred=zeros(size(matrix_image, 2), 1)
36 names=str2mat(string(imagefiles2(2,:)))

```

```

37 names=str2num(names(:,end-1:end))
38 for ii=unique(idx)'
39     ind= (idx==ii);
40     Y_pred(ind)=mode(names(ind,:));
41 end
42 nmi(Y_pred, names)

```

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