

# Exploring Exoplanet Properties

By: Daniel Zurawski, Victor Karkour and Jake Kamen



# Introduction & Motivation

## Intro:

- Our exoplanet system: GJ 436 b
- Explore how we measure exoplanet mass, radius, density
- Using Radial Velocity and transit data
- Compare data with M-R relation from Chen & Kipping (2016)
- Explore abundance of types of exoplanets!

## Some Questions We Look to Answer:

- How can Radial Velocity measurements lead to a deeper understanding of the abundance of planet types in the Cosmos?
- How do our measurements of mass and radius compare to other exoplanets of similar mass and radii?

# Methods

## List of used Methods:

- NEA for data collection
- Google Colab for Python implementation
- Uncertainty for Relative Flux Calculation
- Lomb-Scargle Periodogram
- RadVel python Package
- Matplotlib for simple python plotting

```
17 string sInput;  
18 int iLength, iN;  
19 double dblTemp;  
20 bool again = true;  
21  
22 while (again) {  
23     iN = -1;  
24     again = false;  
25     getline(cin, sInput);  
26     system("cls");  
27     stringstream(sInput) >> dblTemp;  
28     iLength = sInput.length();  
29     if (iLength < 4) {  
30         again = true;  
31         continue;  
32     } else if (sInput[iLength - 3] != '.') {  
33         again = true;  
34         continue;  
35     } while (++iN < iLength) {  
36         if (isdigit(sInput[iN])) {  
37             continue;  
38         } else if (iN == (iLength - 3)) {  
39             continue;  
40         }  
41     }  
42 }
```

# Methods Continued...

Equations (shown on right):

- K-value equation
- Planet Mass equation
- Kepler's 3rd Law
- Transit depth equation

$$K = \frac{M_p}{M_\star + M_p} \frac{na \sin(i)}{\sqrt{1 - e^2}}$$

$$M_p = \frac{KM_\star \sqrt{1 - e^2}}{\frac{2\pi}{T} \left( \frac{GM_\star T^2}{4\pi^2} \right)^{\frac{1}{3}}}$$

$$a = \left( \frac{GM_\star T^2}{4\pi^2} \right)^{\frac{1}{3}}$$

$$\delta = \left( \frac{R_p}{R_\star} \right)^2$$

# Results

## Mass:

Approximate mass = 20.003 Earth Masses

## Using Uncertainties:

Mass range is  $\pm 1.99963 M_{\text{earth}}$

## Radius:

Radius = 4.10156 Earth Radii

## Using Uncertainties:

Radius range is  $\pm 0.24674$  Earth Radii

## Density:

Density = 1.59297 g/cm<sup>3</sup>

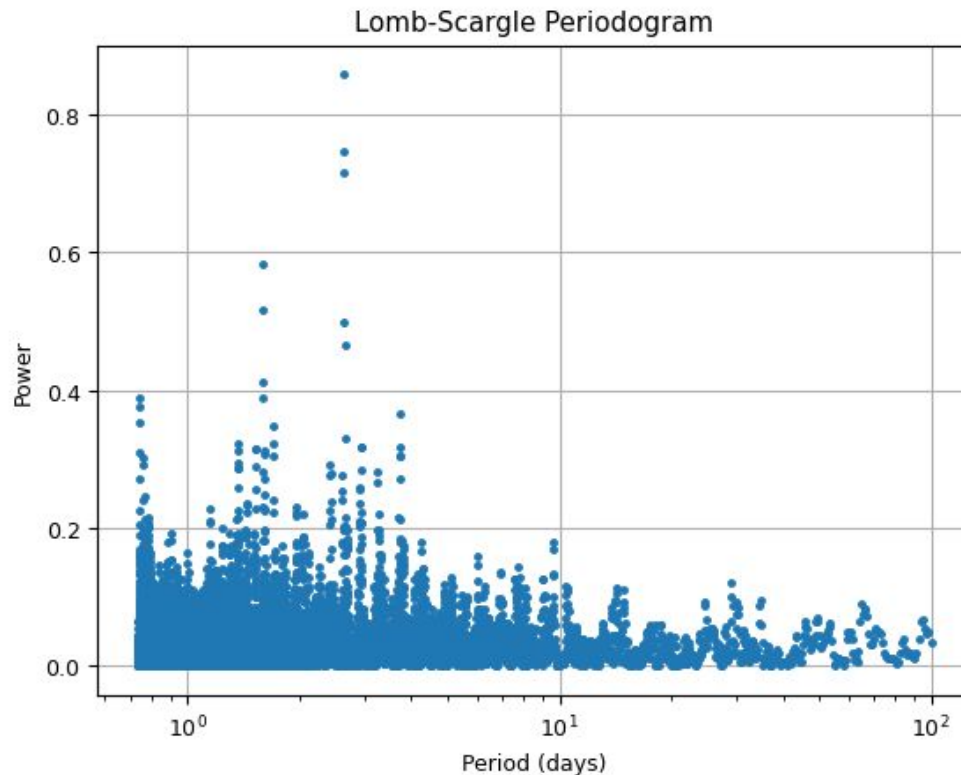
Uncertainty =  $\pm 0.18585$  g/cm<sup>3</sup>

- Successfully found our exoplanets mass from the radial velocity data and radius from transit data
- Calculated uncertainties of each

# Results Continued...

## Lomb-Scargle Periodogram (Graph 1)

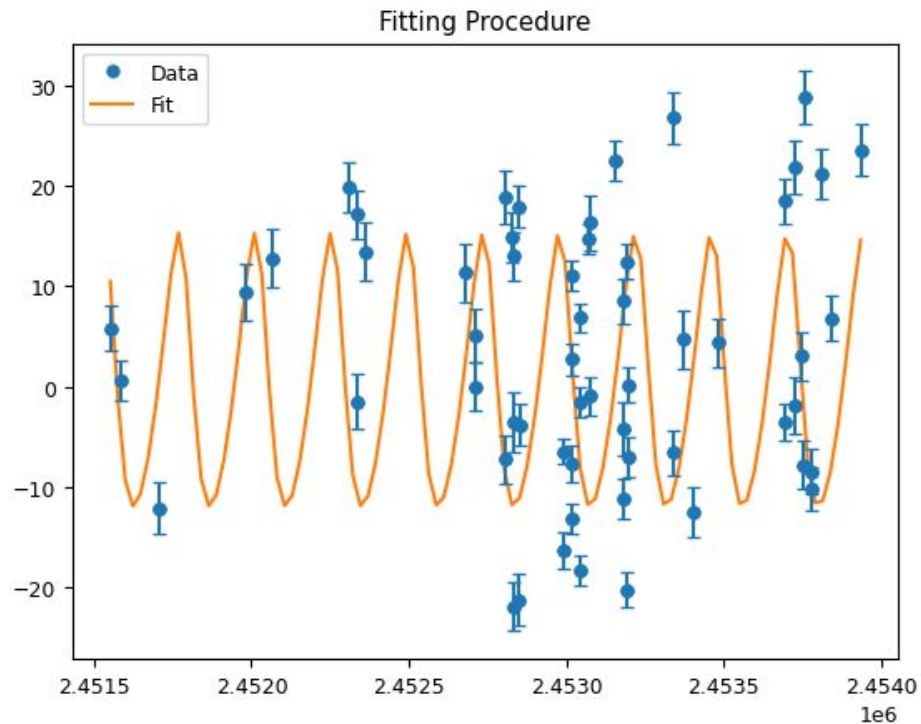
- Graph that determines period through unevenly spaced data
- More data at smaller periods
- Large spikes between 1-10
- Largest spike at our period
- future graphs (graph 2)
- Leads to calculations for mass



# Results Continued...

## RadVel (Graph 2)

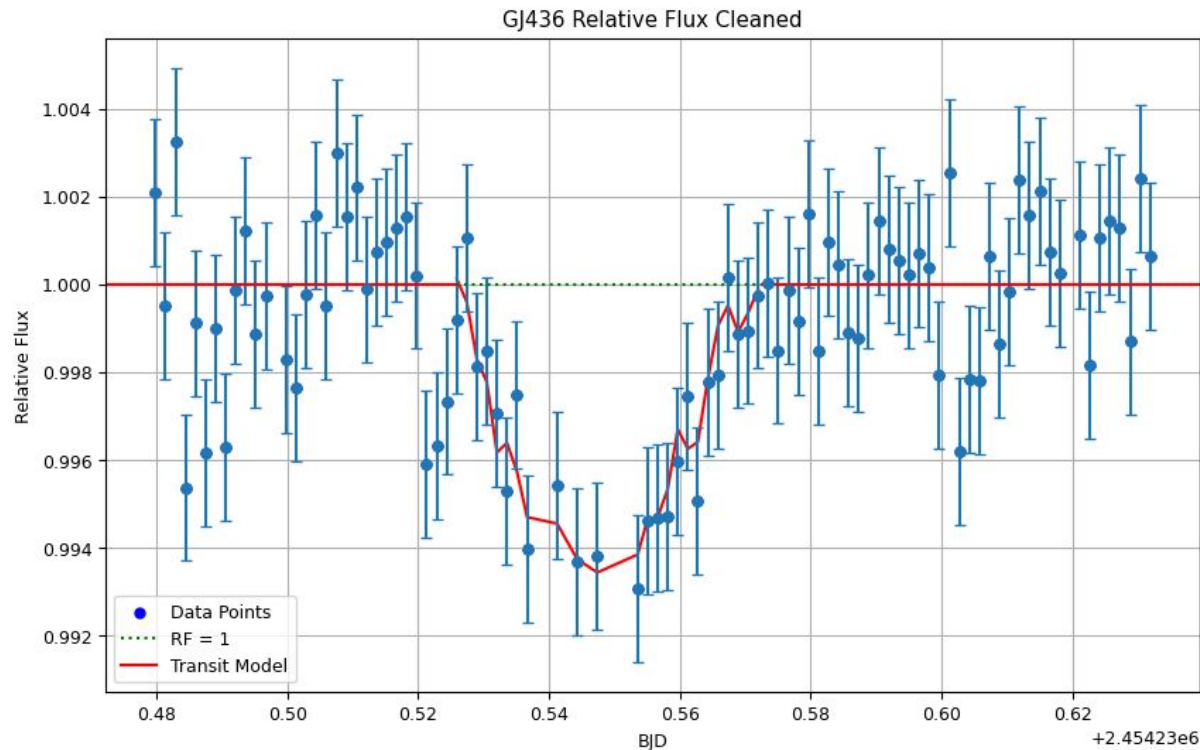
- “Fitting Procedure”
- Clearly sinusoidal
- Notice error bars
- Led to us finding our Mass value



# Results Continued...

## Transit Light Curve (Graph 3)

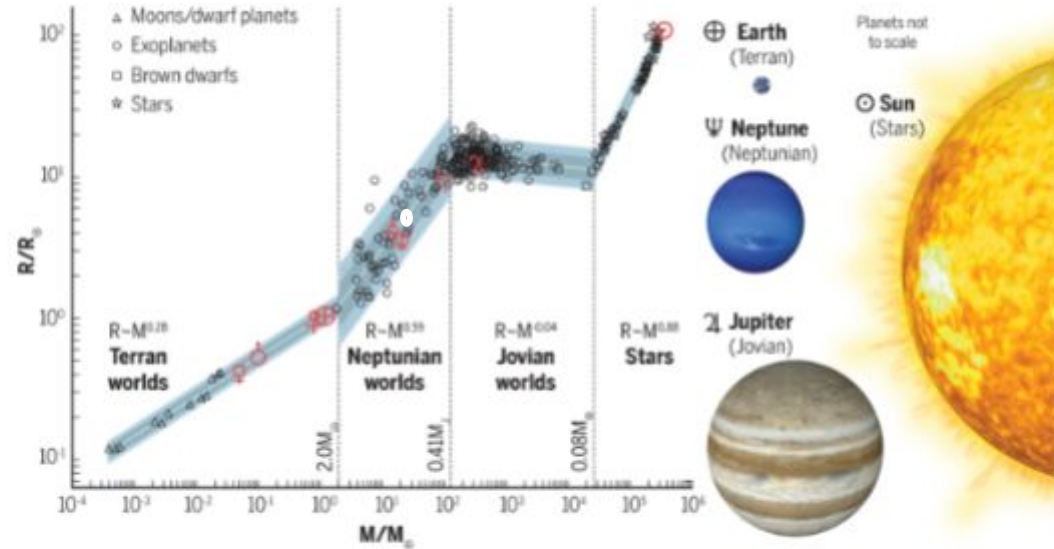
- Notice initial clear dip in middle of graph
- Error Bars from standard deviation
- This graph is “cleaned” (by  $2\sigma$ )
- Red line (transit model)
- Clearly see transit depth
- Leads to radius calculation





# Chen and Kipping (2016)

- Mass-radius relation from Chen and Kipping (2016)
- Our measurement of mass  
= 20.003 Earth masses
- Our measurement of radius  
= 0.3659 Jupiter Radii  $\sim 4$  earth radii
- Our measurements would put our planet in the Neptunian worlds section of the M-R relation (White circle)



# Conclusion

We were able to successfully calculate the radius, mass, and density of our planet using RV and transit data.

