

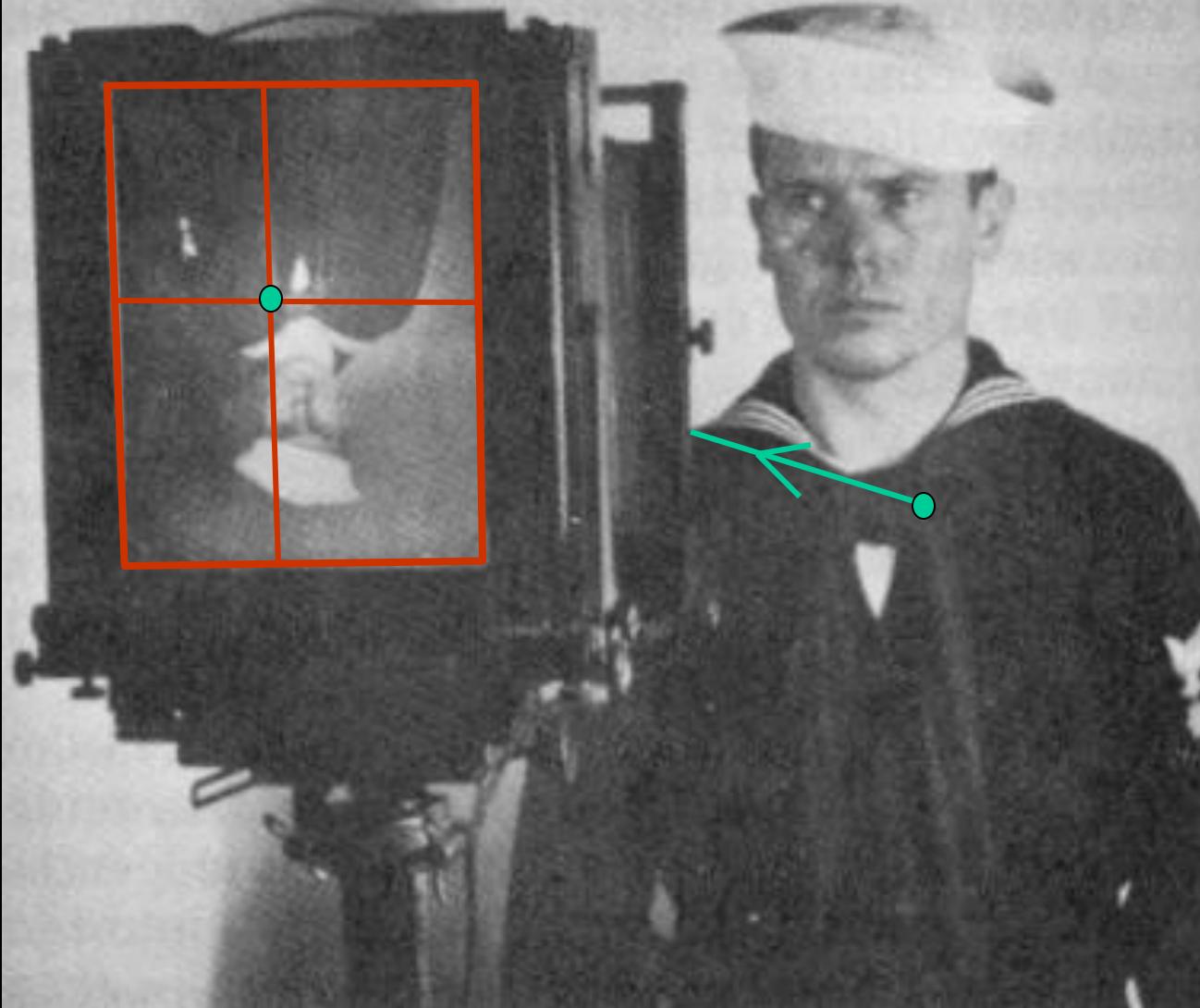
# Elements of Camera geometry and Image processing

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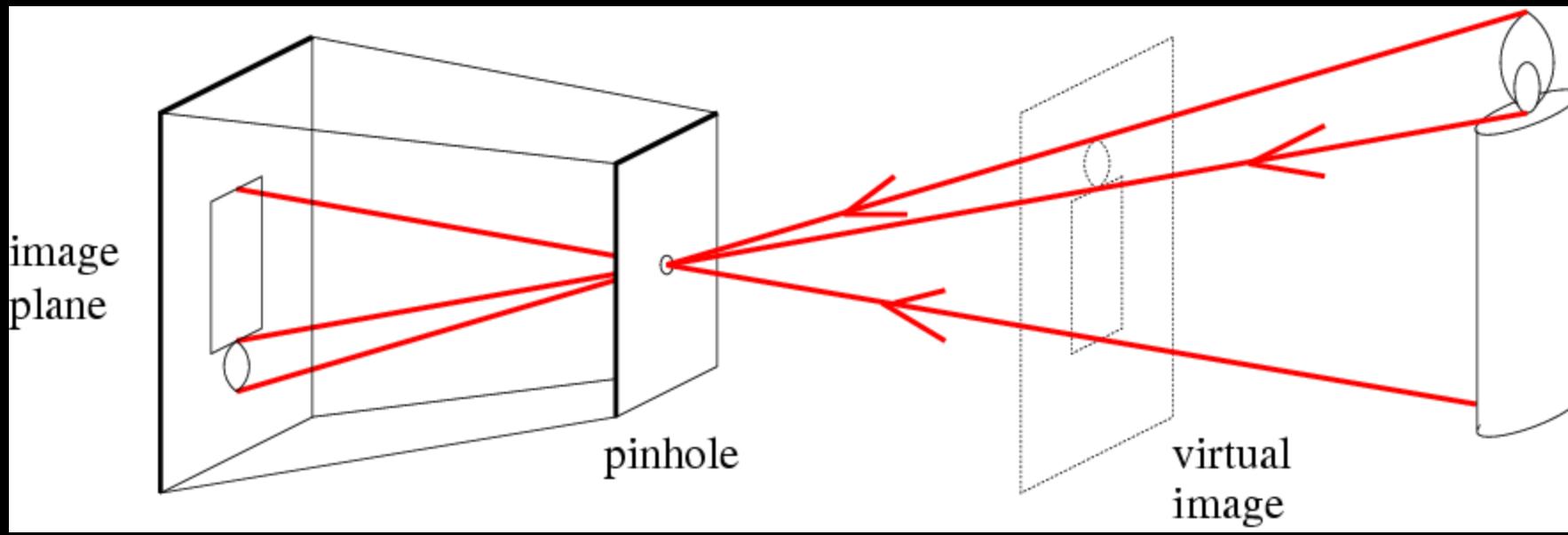
# Camera geometry and calibration

- Pinhole perspective projection
- Orthographic and weak-perspective models
- Non-standard models
- A detour through sensing country
- Intrinsic and extrinsic parameters
- Camera calibration

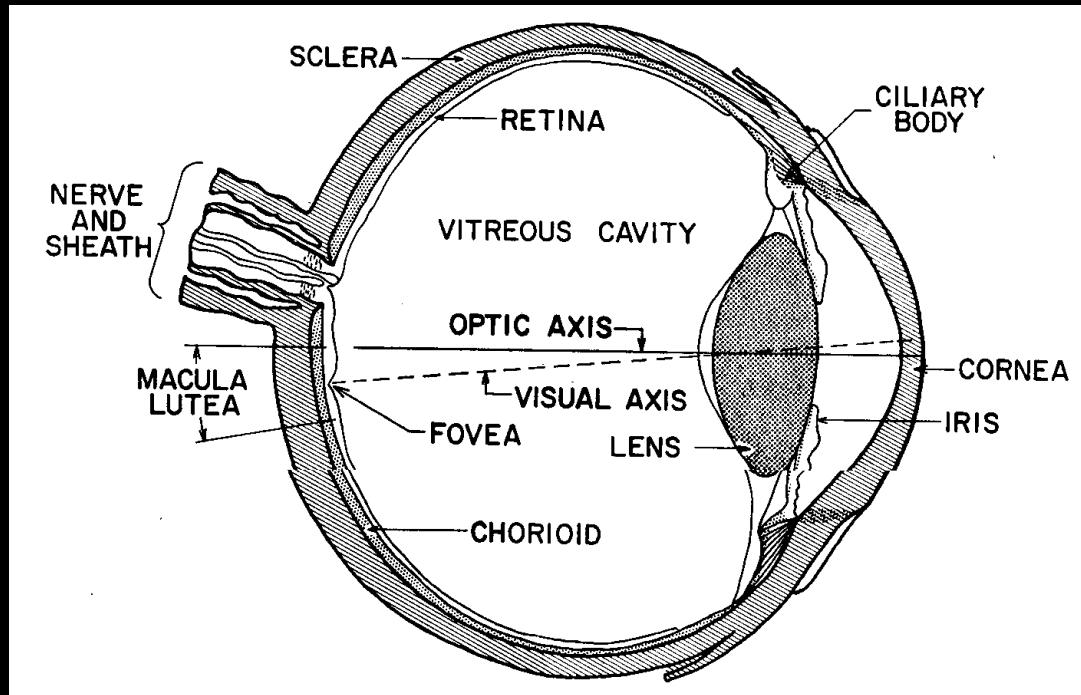
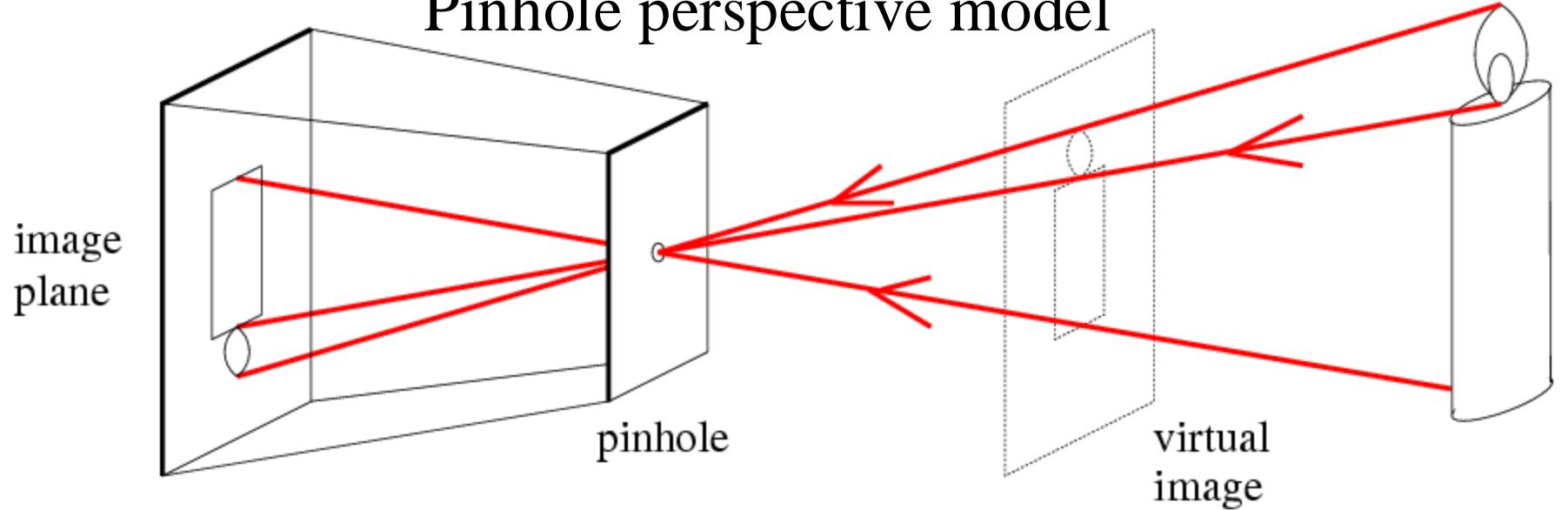
They are formed by the projection of 3D objects.



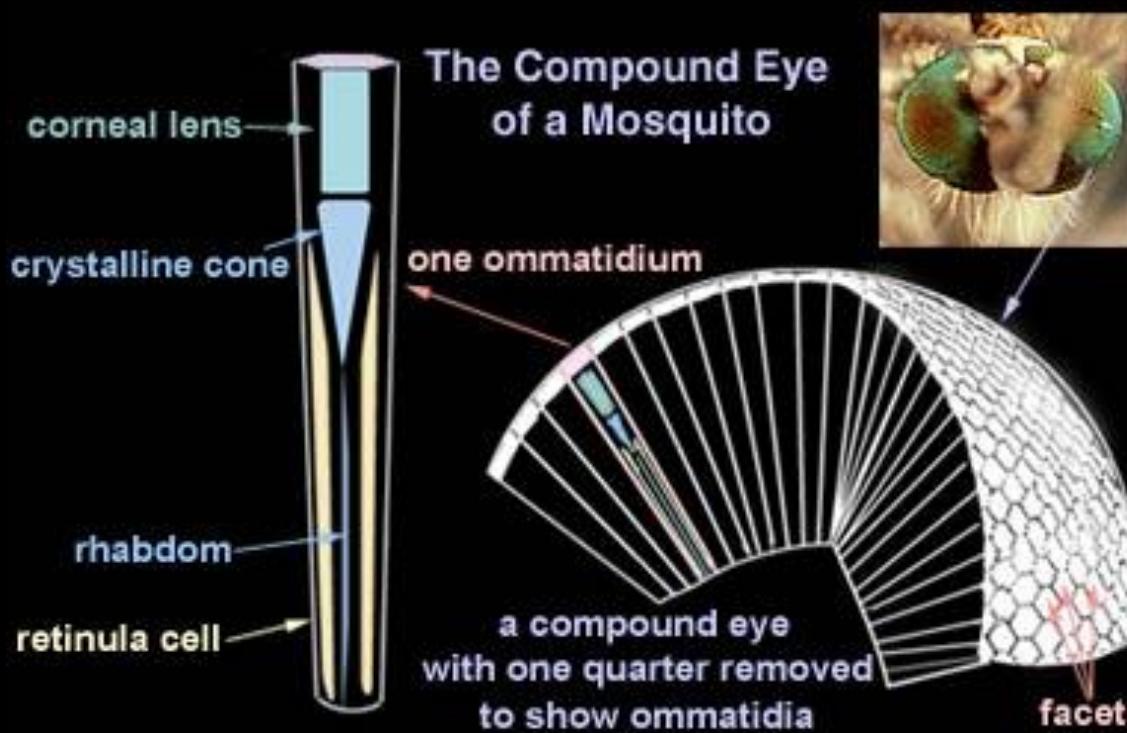
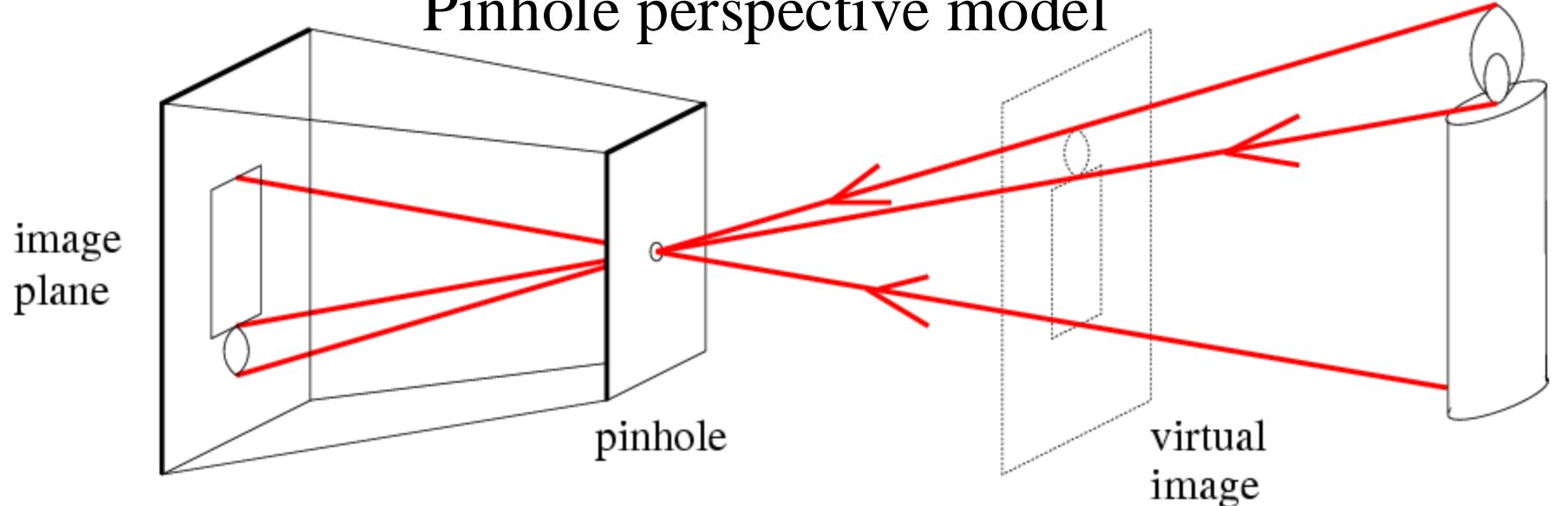
Images are two-dimensional patterns of brightness/color values



# Pinhole perspective model

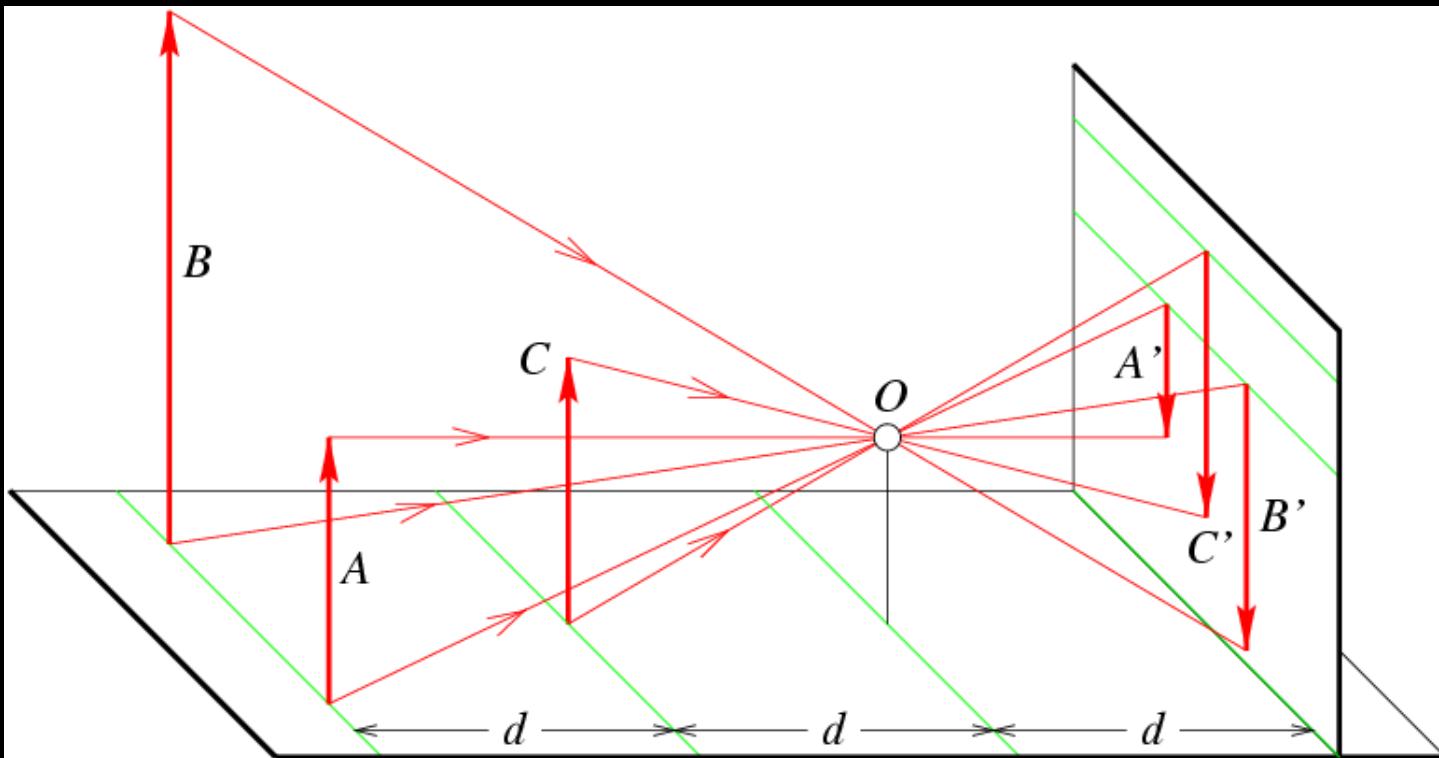
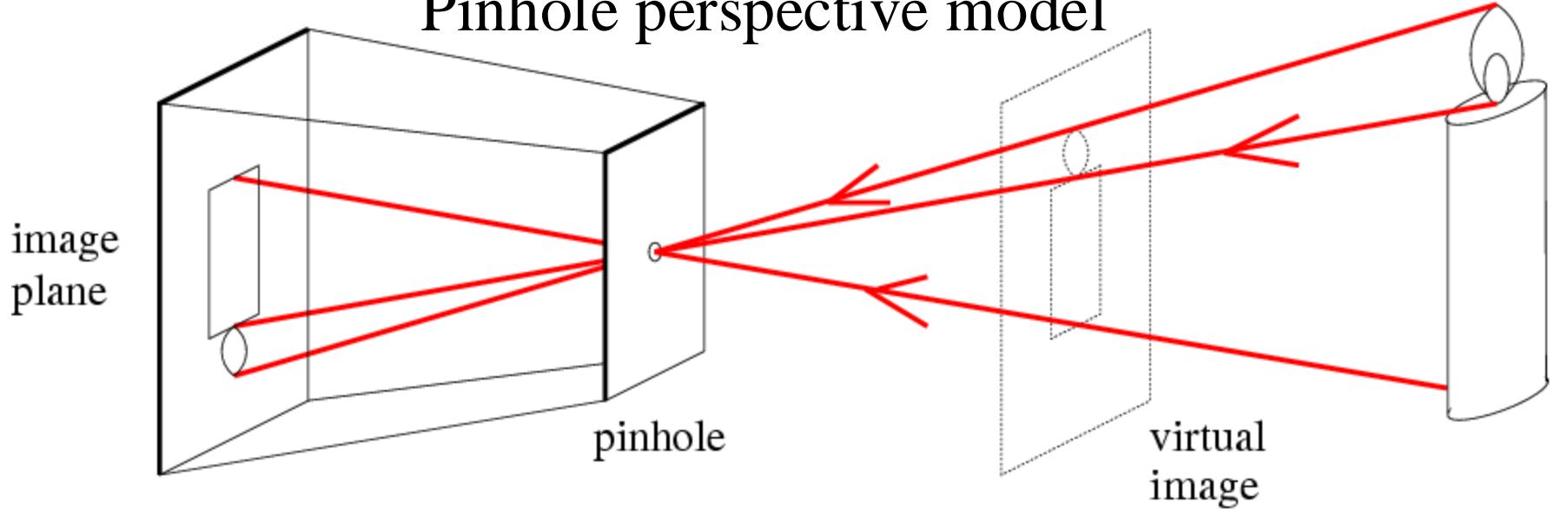


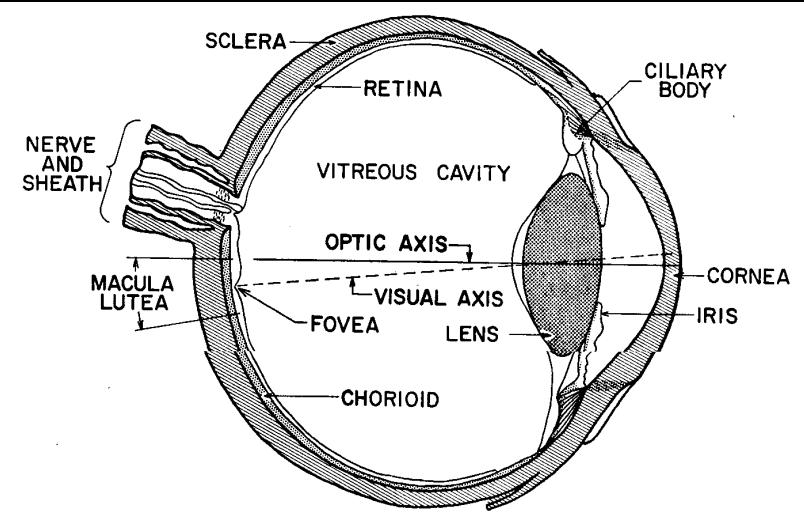
# Pinhole perspective model



Land & Nilsson  
“Animal Eyes”  
Oxford, 2012

# Pinhole perspective model

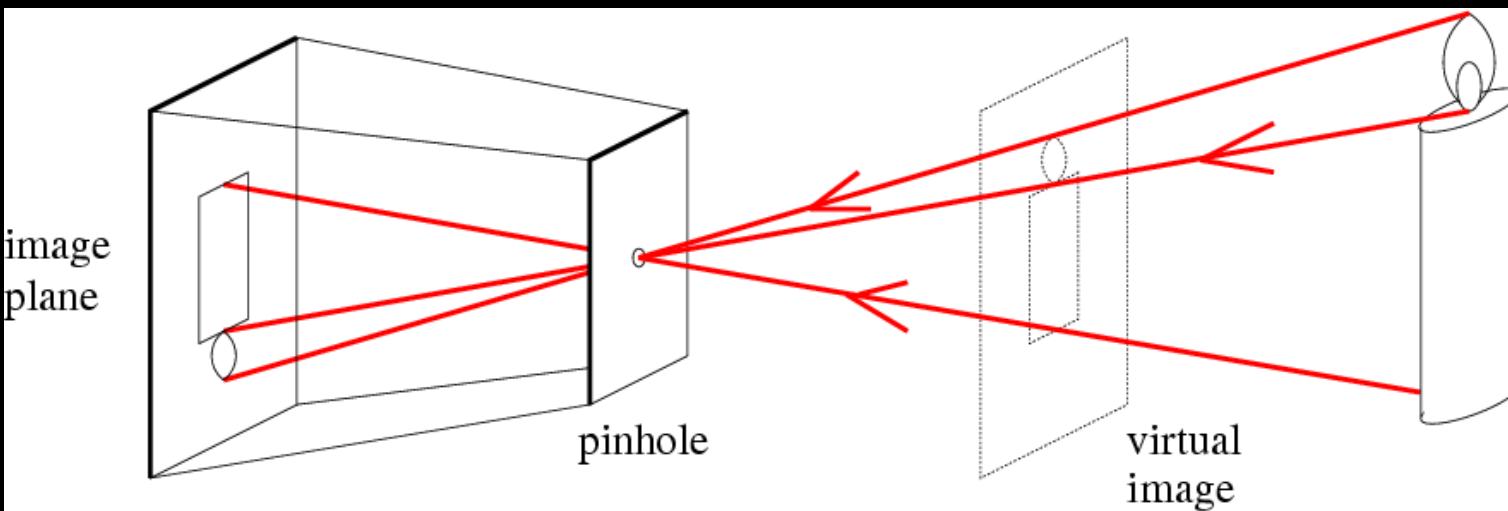




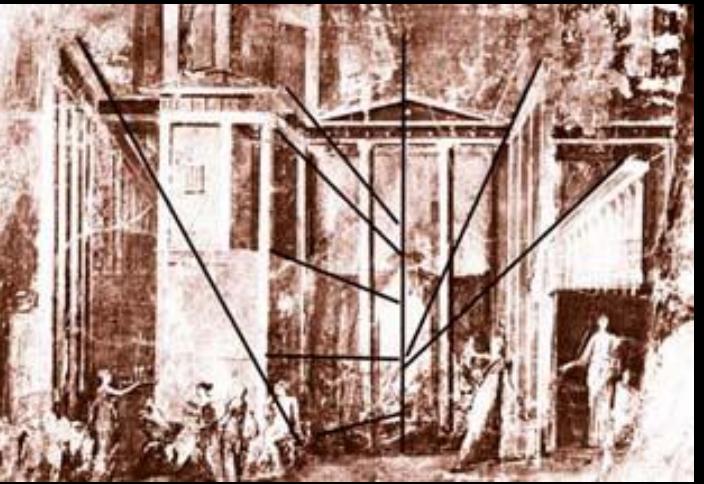
Animal eye: a looonng time ago.



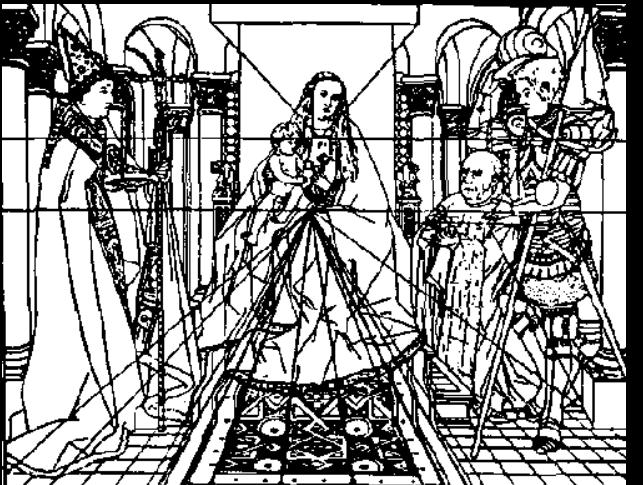
Photographic camera:  
Niepce, 1816.



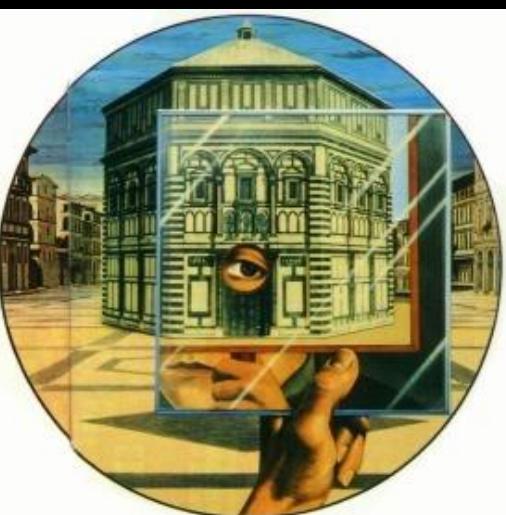
Pinhole perspective projection: Brunelleschi, XV<sup>th</sup> Century.  
Camera obscura: XVI<sup>th</sup> Century.



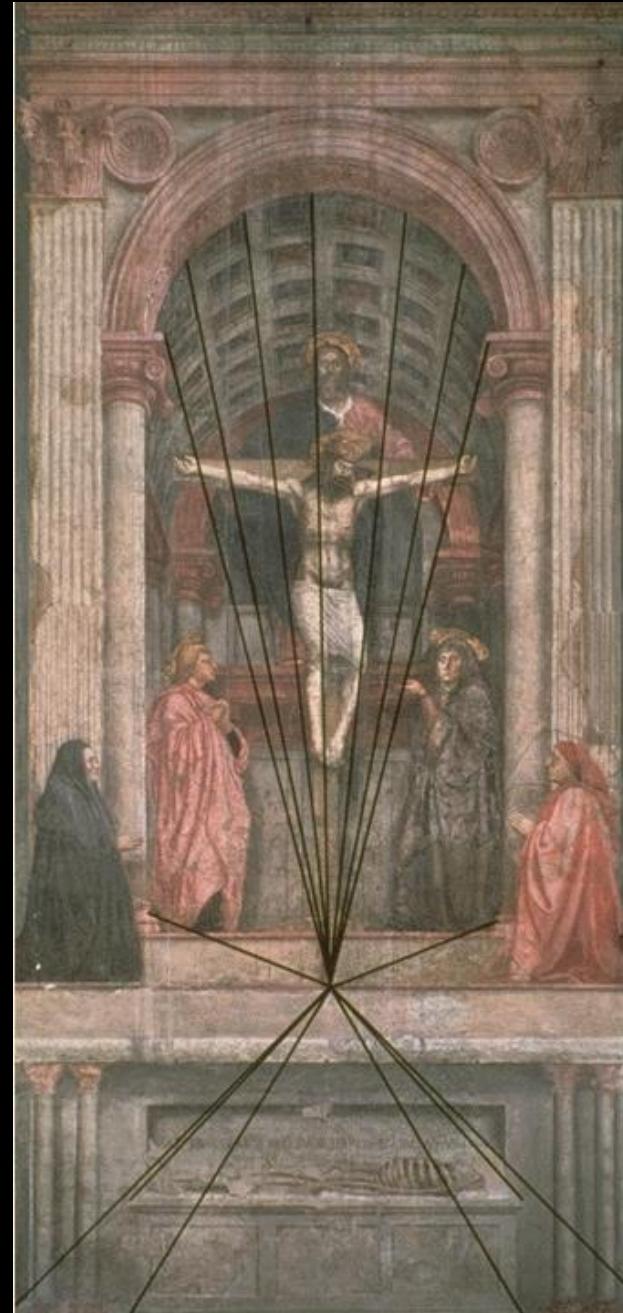
Pompeii painting, 2000 years ago



Van Eyk, XIV<sup>th</sup> Century

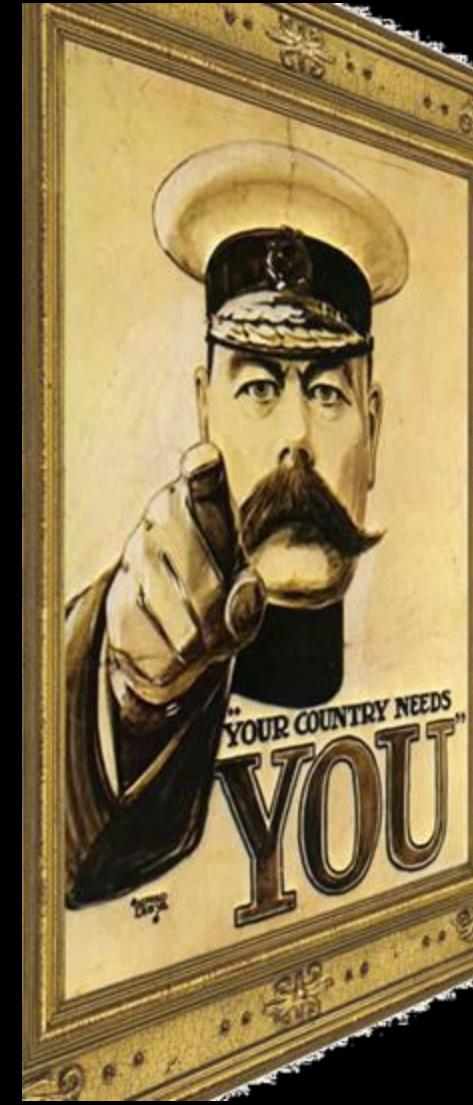


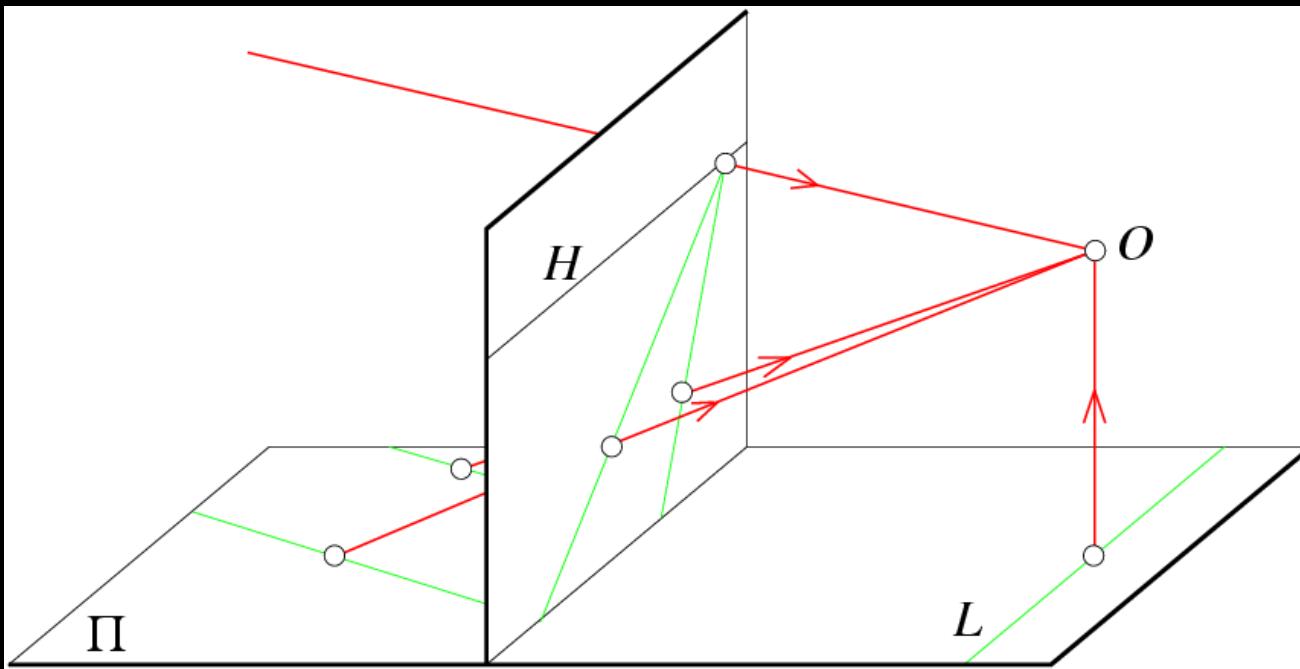
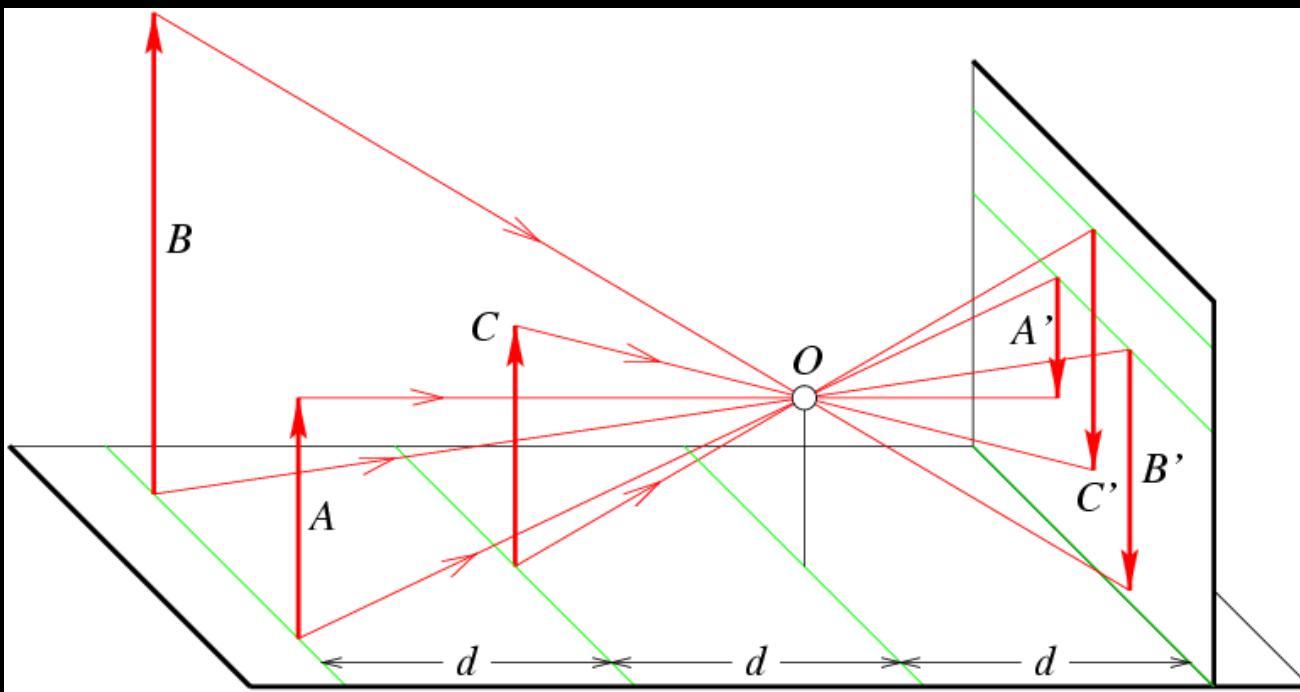
Brunelleschi, 1415



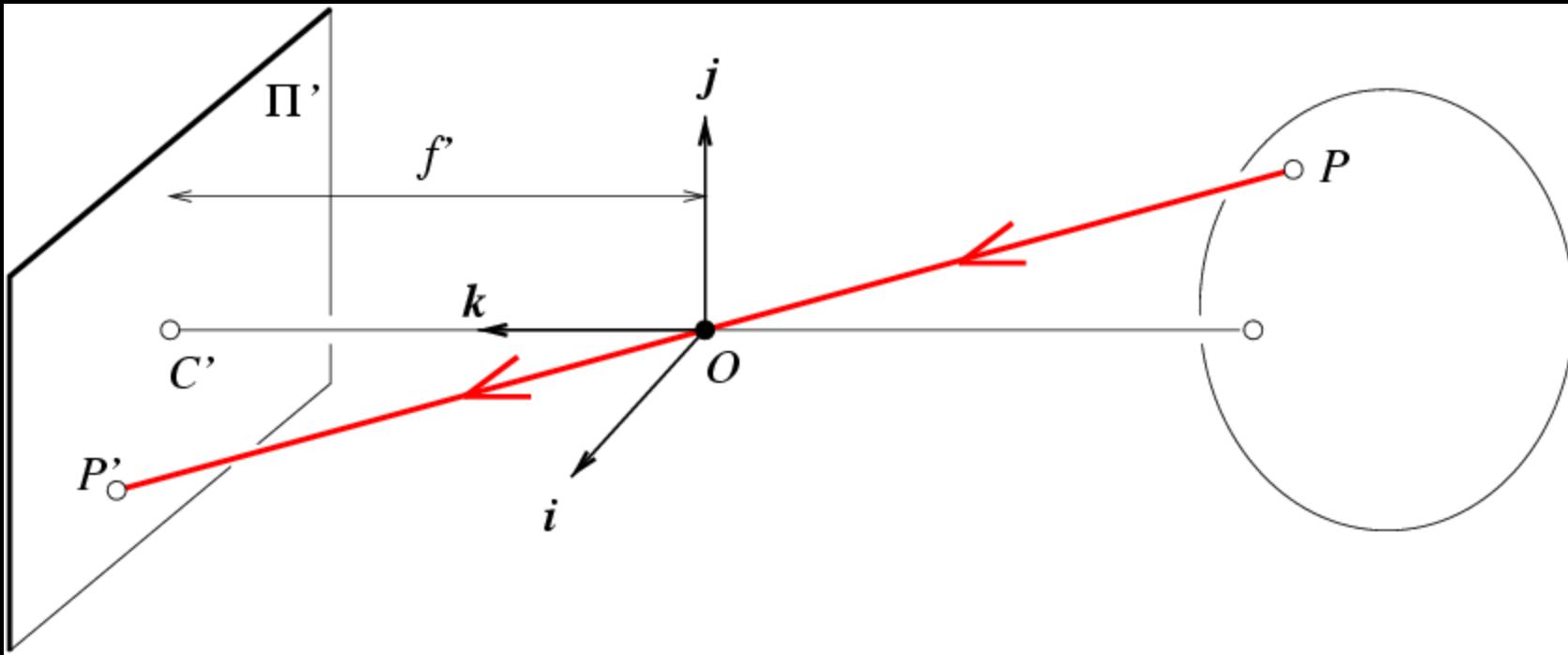
Massaccio's Trinity, 1425

# How do we see images?





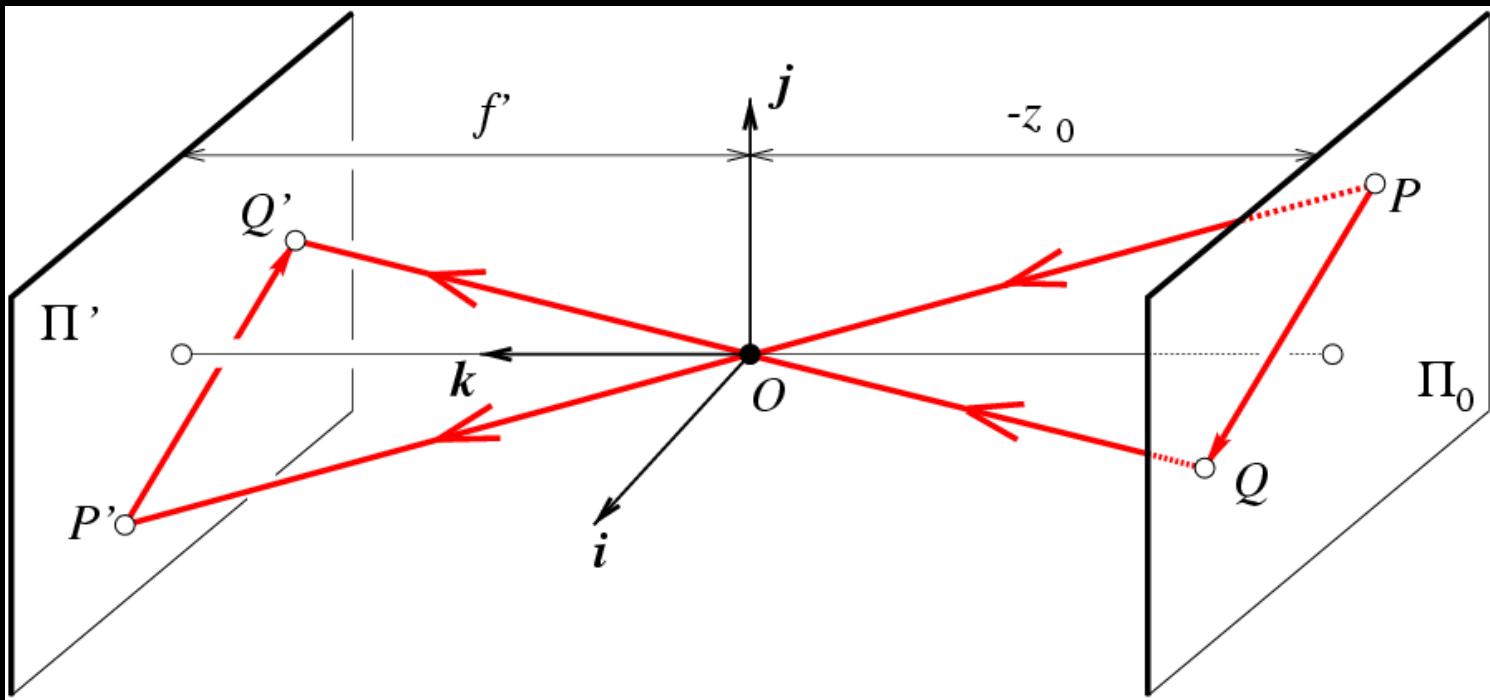
# Pinhole Perspective Equation



$$\begin{cases} x' = f' \frac{x}{z} \\ y' = f' \frac{y}{z} \end{cases}$$

NOTE:  $z$  is always negative..

## Affine projection models: Weak perspective projection

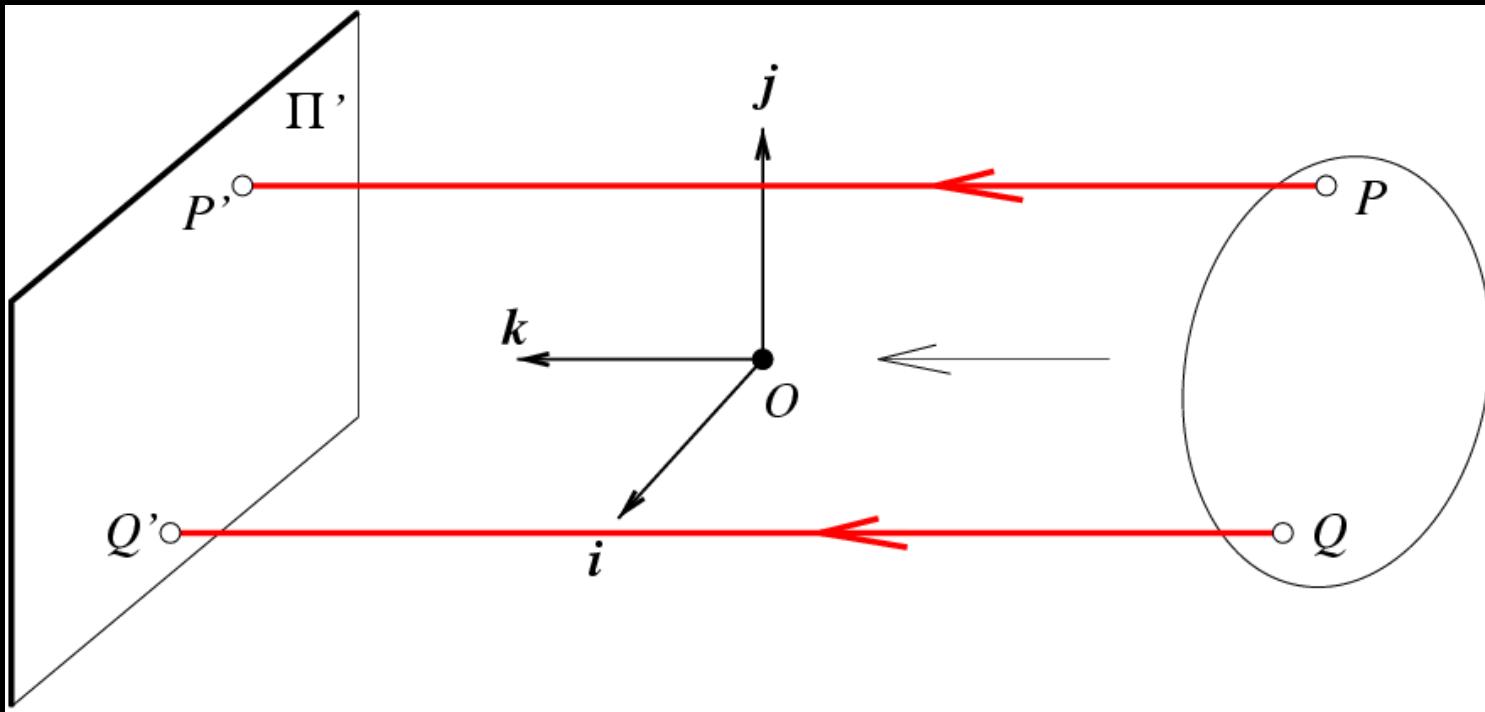


$$\begin{cases} x' = mx \\ y' = my \end{cases}$$

where  $m = \frac{f'}{-z_0}$  is the (negative) magnification.

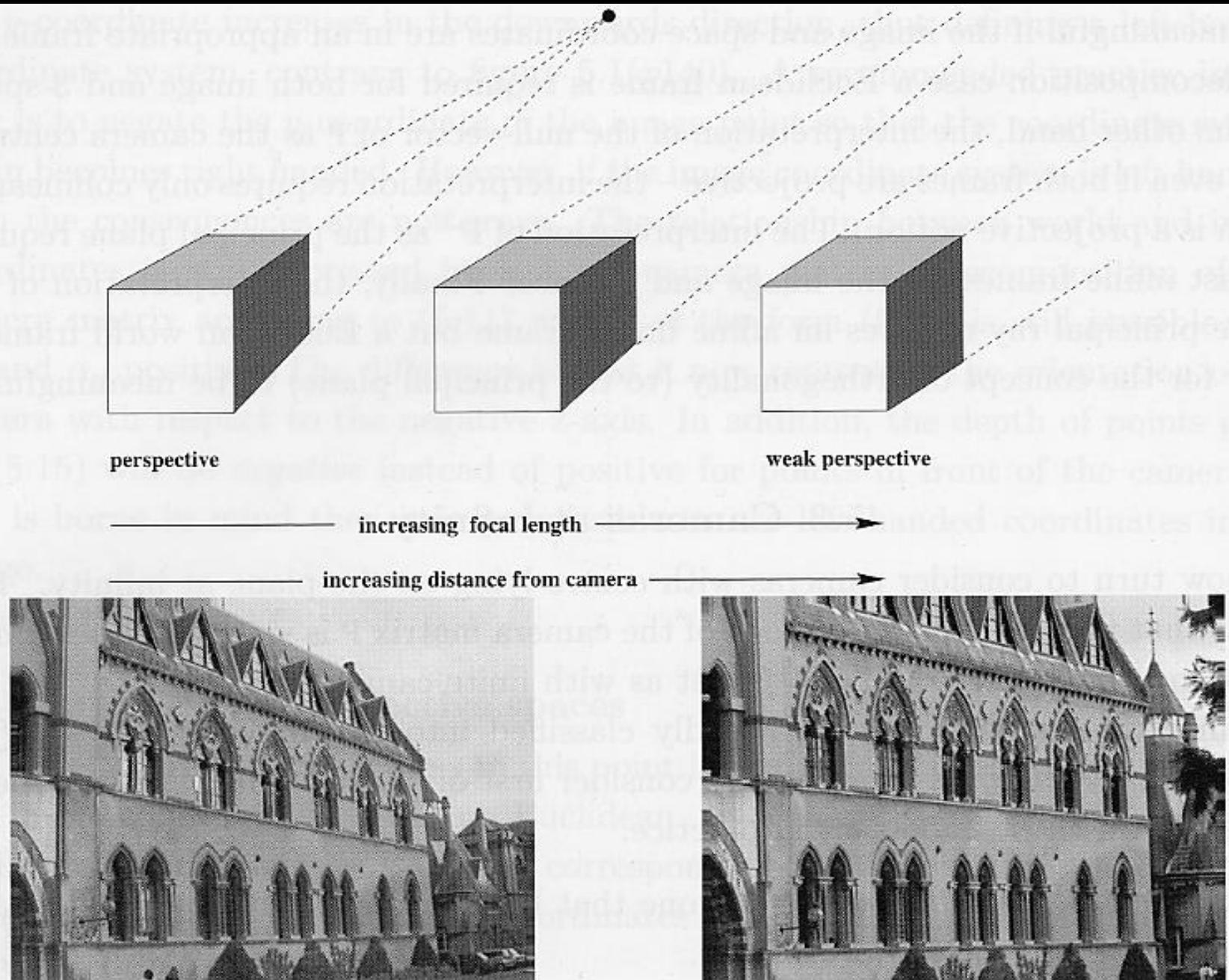
When the scene relief is small compared its distance from the Camera,  $m$  can be taken constant: weak perspective projection.

## Affine projection models: Orthographic projection

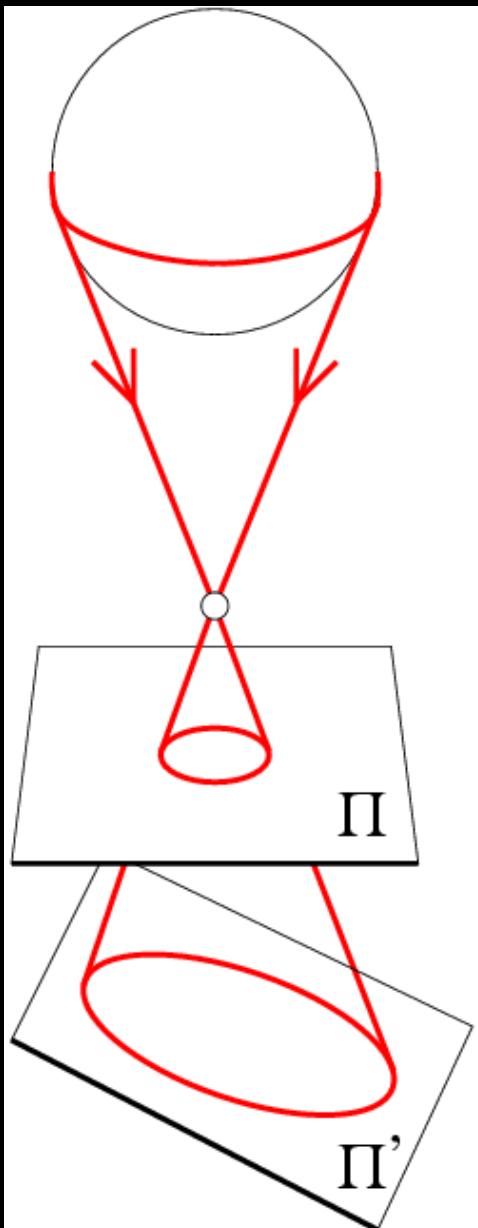


$$\begin{cases} x' = x \\ y' = y \end{cases}$$

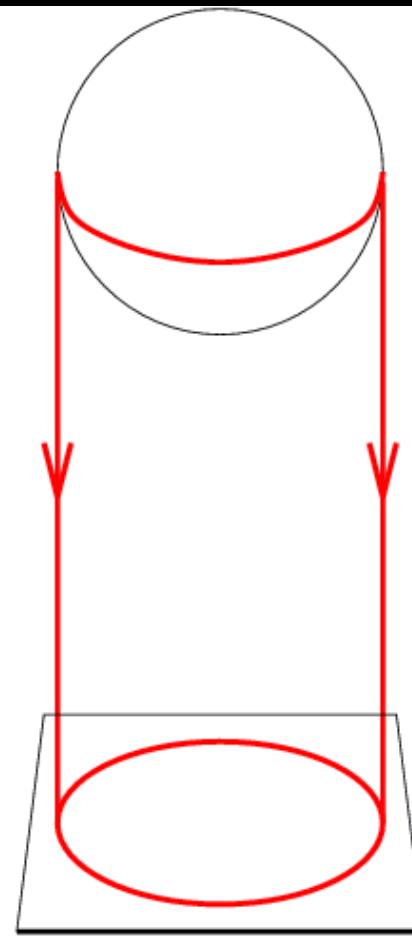
When the camera is at a  
(roughly constant) distance  
from the scene, take  $m = -1$



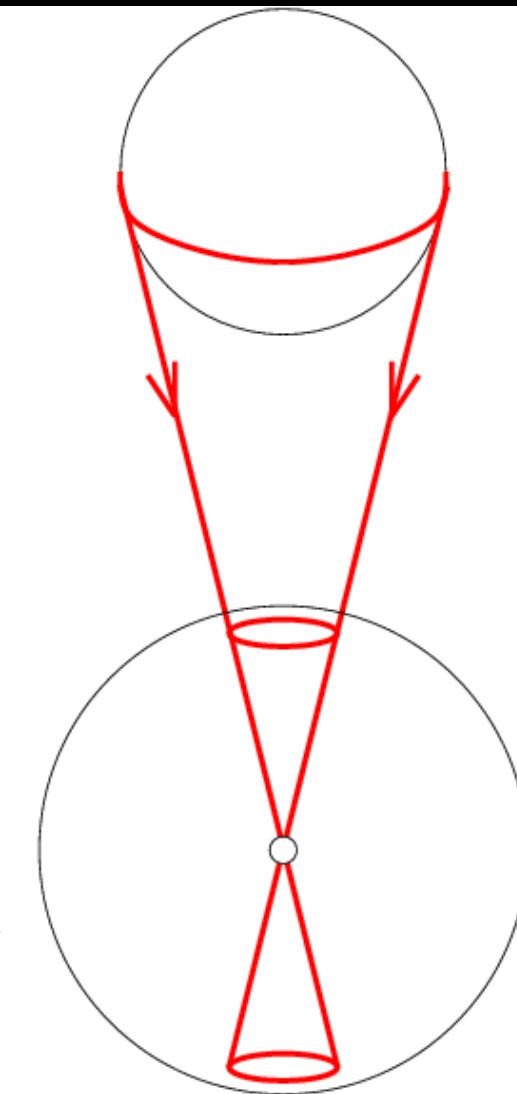
From Zisserman & Hartley



Planar pinhole  
perspective

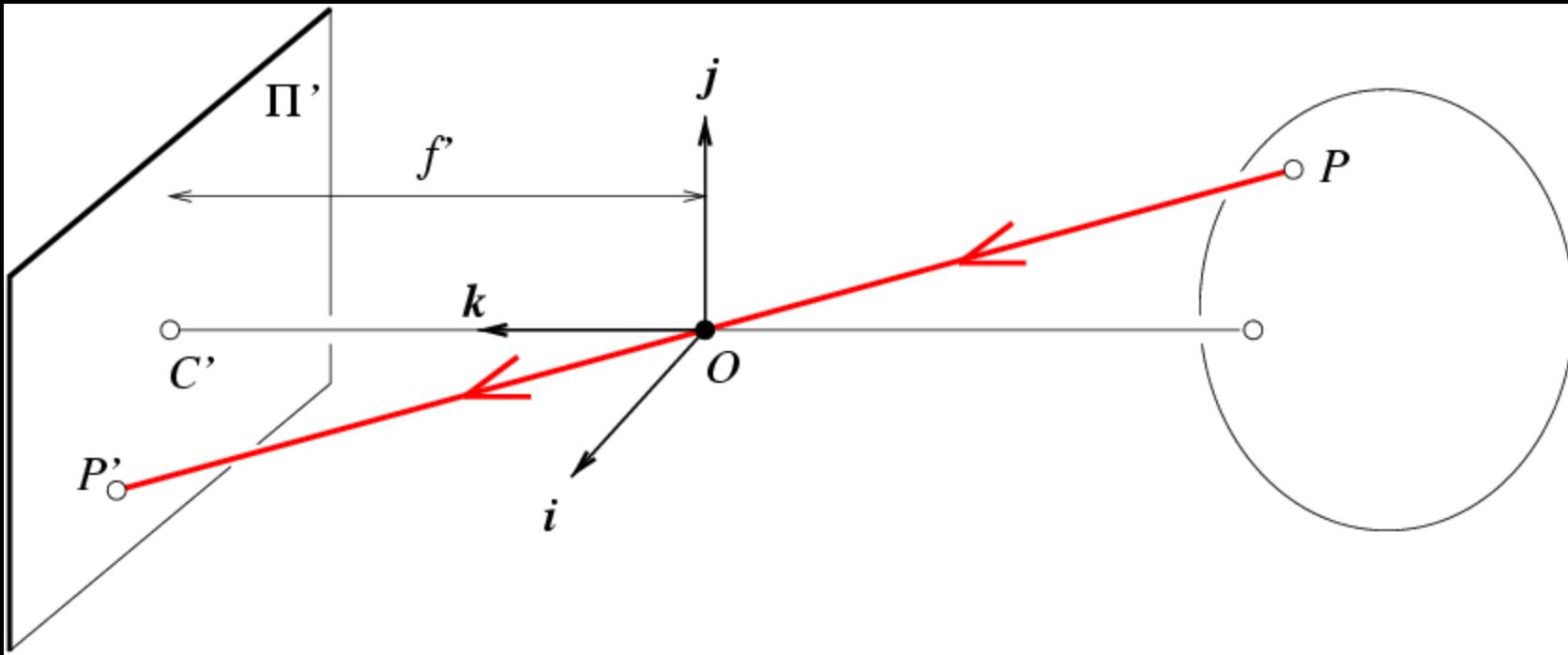


Orthographic  
projection



Spherical pinhole  
perspective

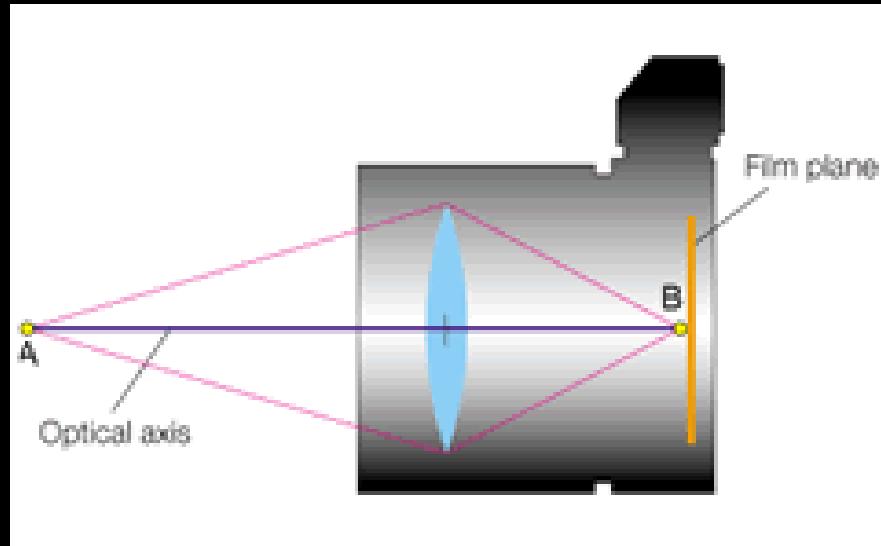
# Pinhole Perspective Equation



$$\begin{cases} x' = f' \frac{x}{z} \\ y' = f' \frac{y}{z} \end{cases}$$

NOTE:  $z$  is always negative..

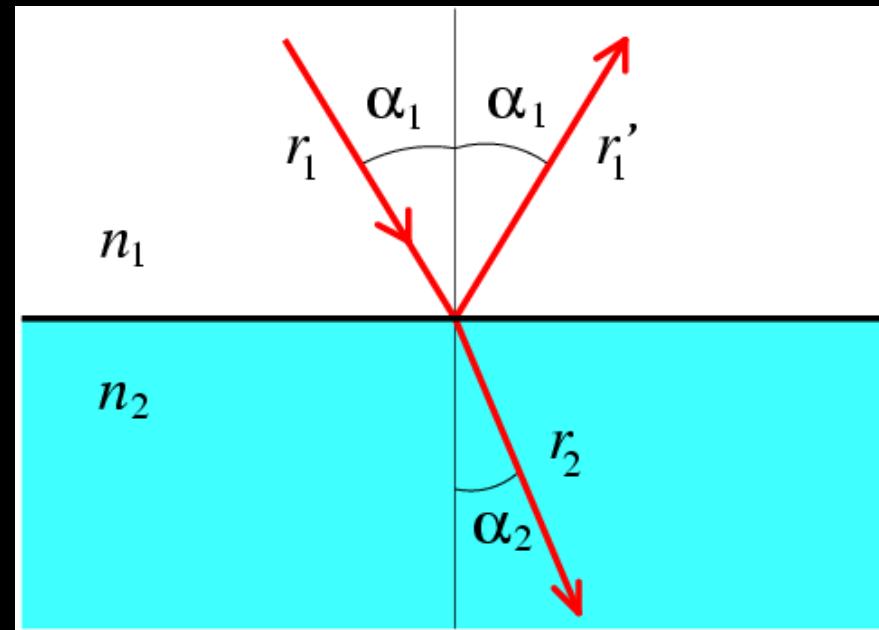
# Lenses



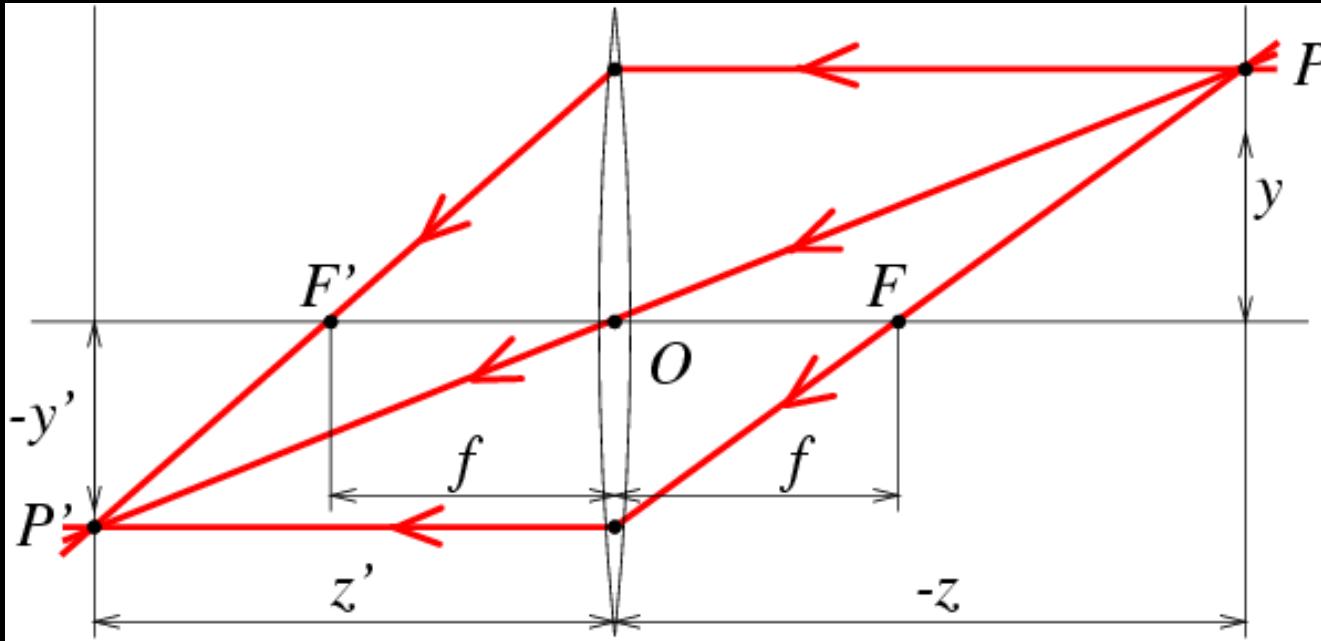
Snell's law

$$n_1 \sin \alpha_1 = n_2 \sin \alpha_2$$

(Descartes' law  
for Frenchies)



# Thin Lenses (including paraxial approximation)



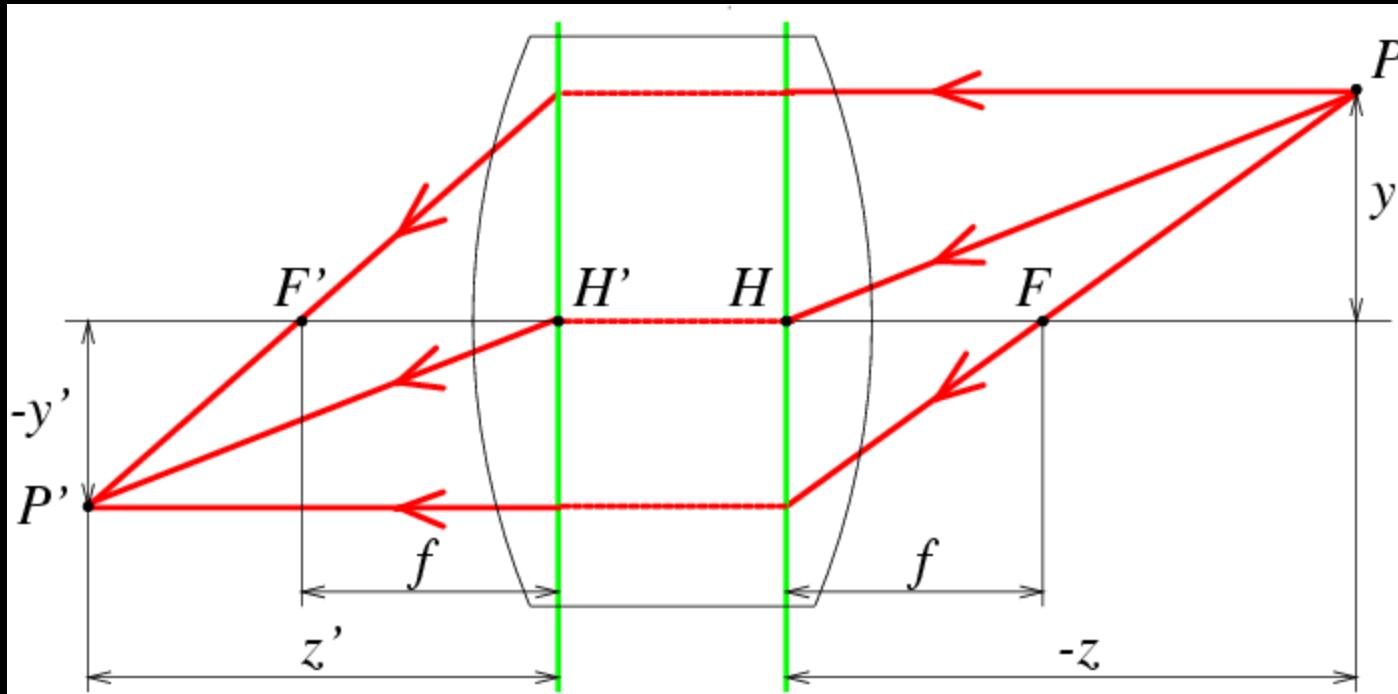
$$\begin{cases} x' = z \frac{x}{z} \\ y' = z \frac{y}{z} \end{cases}$$

where

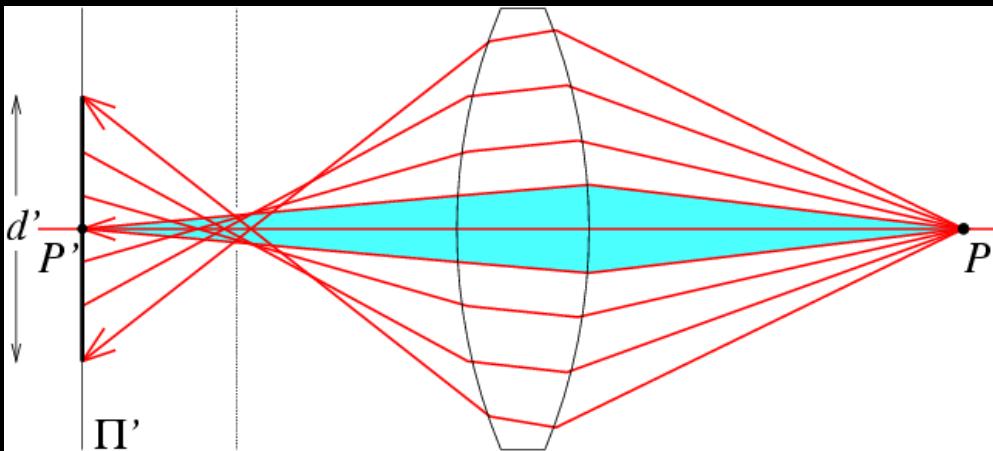
$$\frac{1}{z'} - \frac{1}{z} = \frac{1}{f}$$

$$f = \frac{R}{2(n-1)}$$

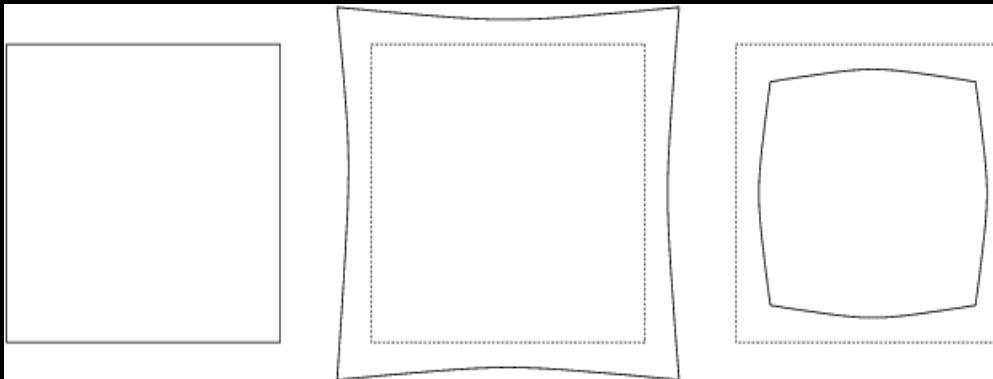
# Thick Lenses



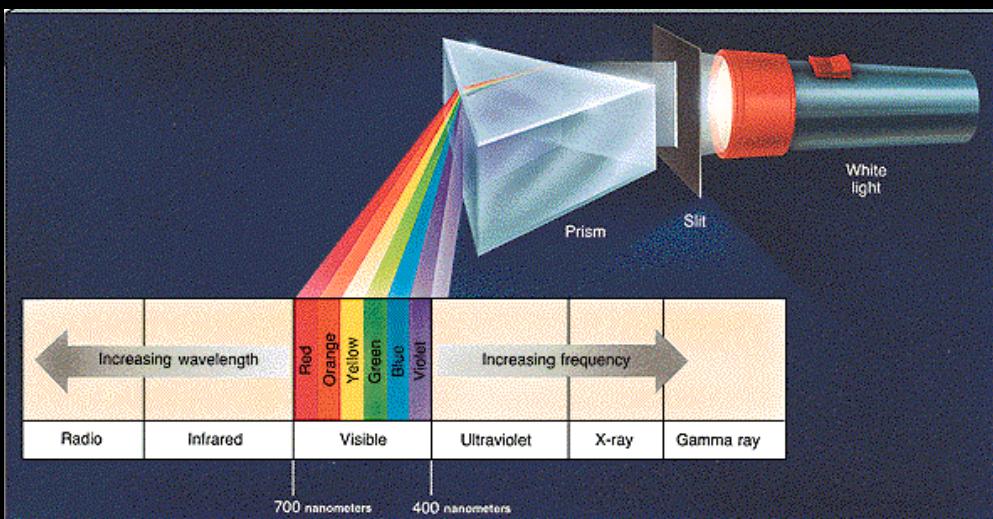
# Spherical Aberration



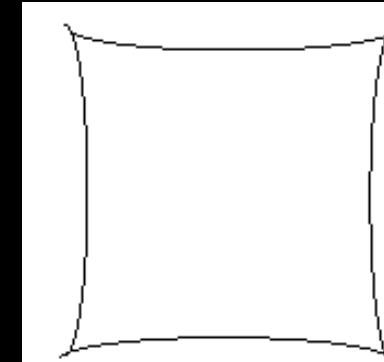
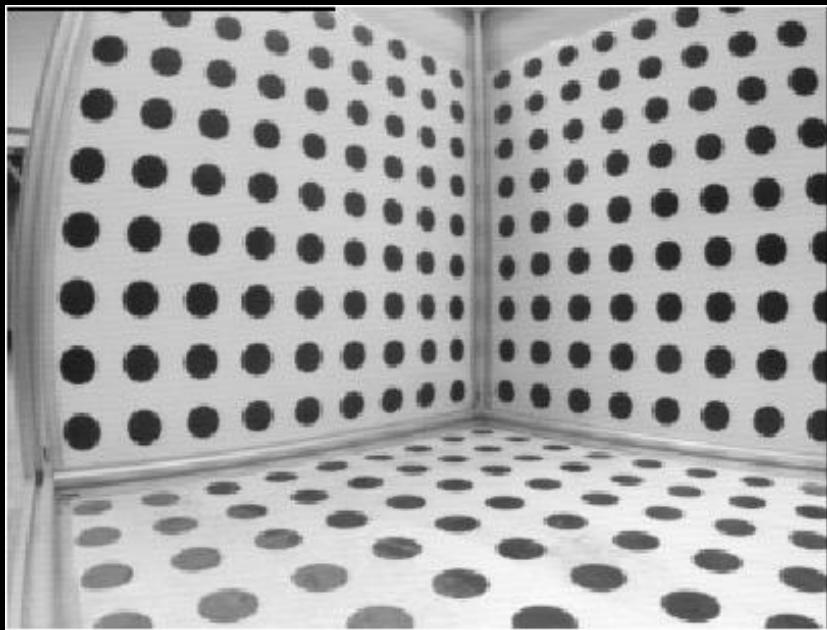
# Distortion



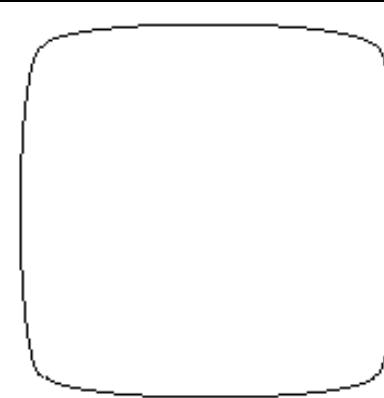
# Chromatic Aberration



# Geometric Distortion



pincushion

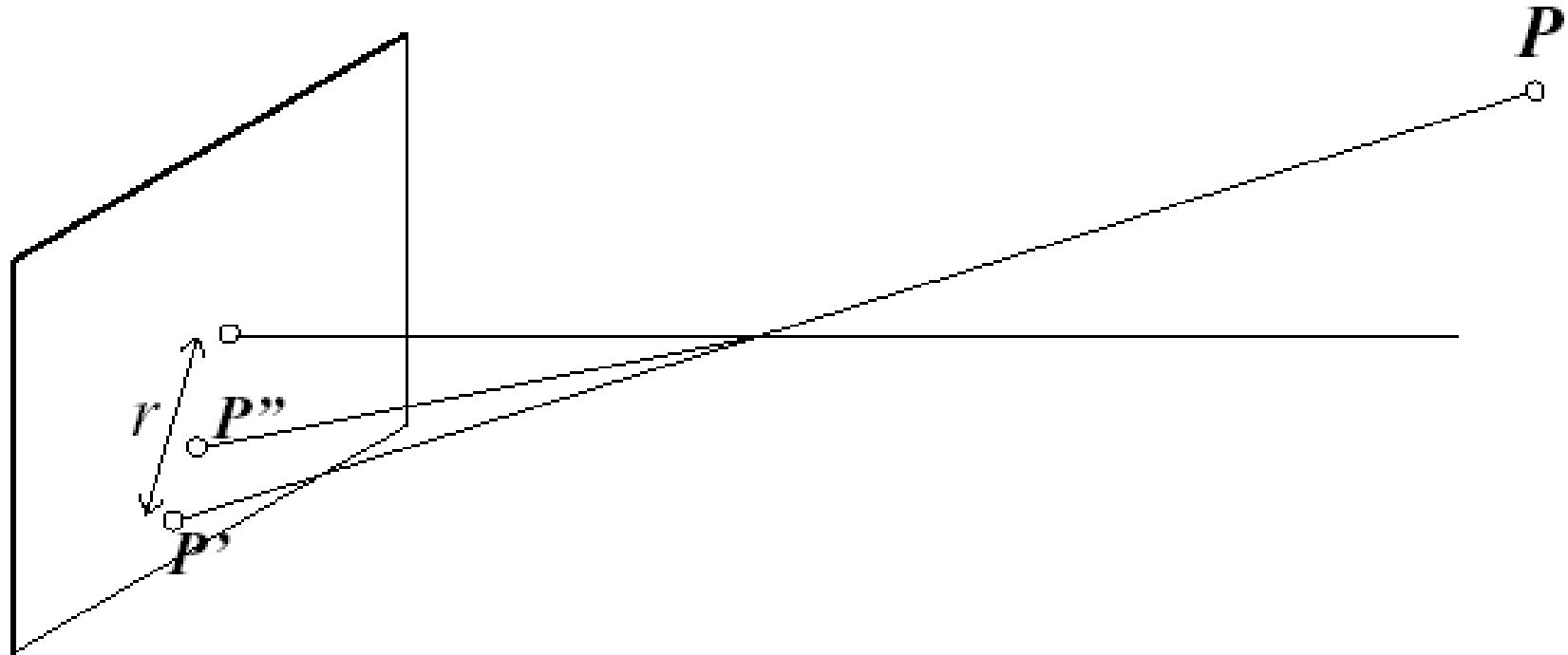


barrel



Rectification

# Radial Distortion Model



Ideal:

$$x' = f \frac{x}{z}$$

$$y' = f \frac{y}{z}$$

Distorted:

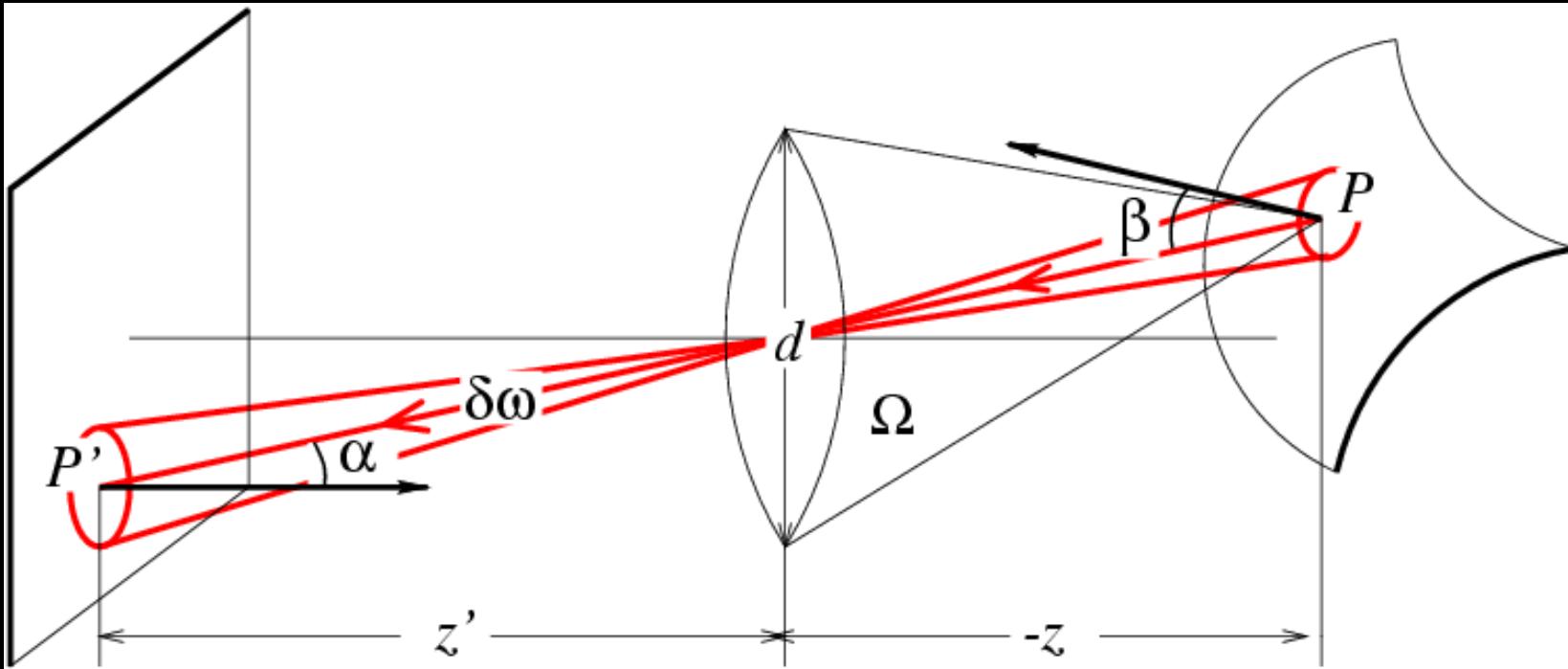
$$x'' = \frac{1}{\lambda} x'$$

$$y'' = \frac{1}{\lambda} y'$$

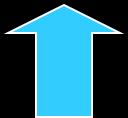
$$\lambda = 1 + k_1 r^2 + k_2 r^4 + \dots$$

A compound lens

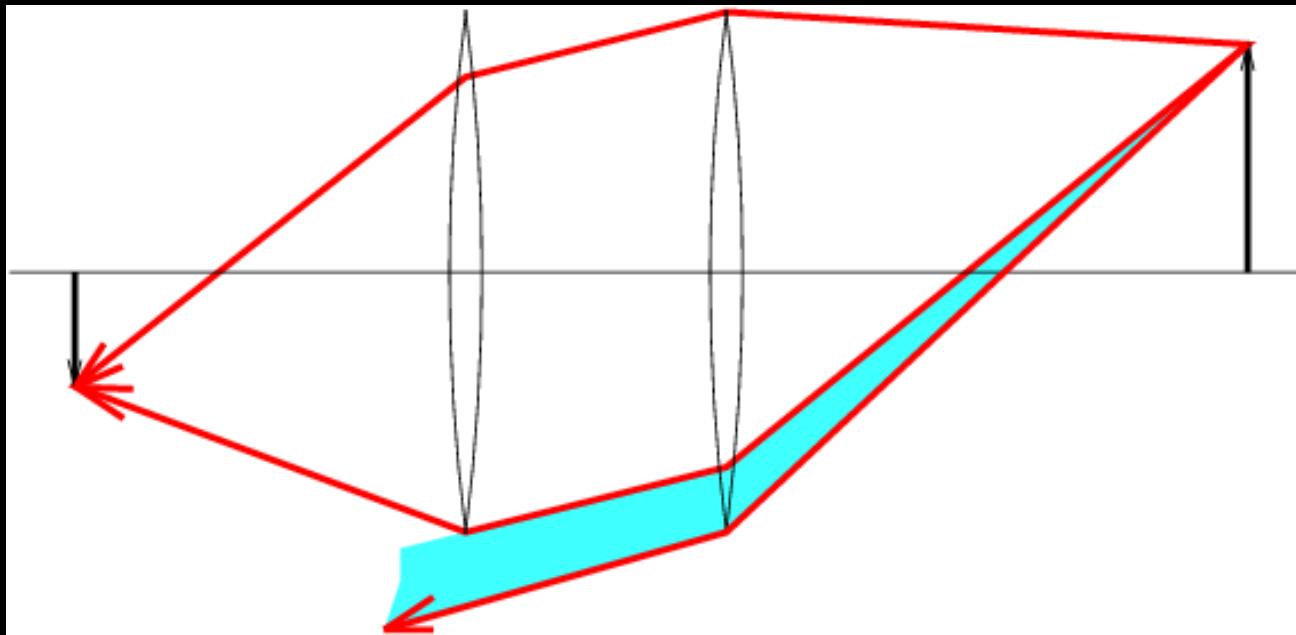




$$E = (\Pi/4) [ (d/z')^2 \cos^4 \alpha ] L$$



# Vignetting



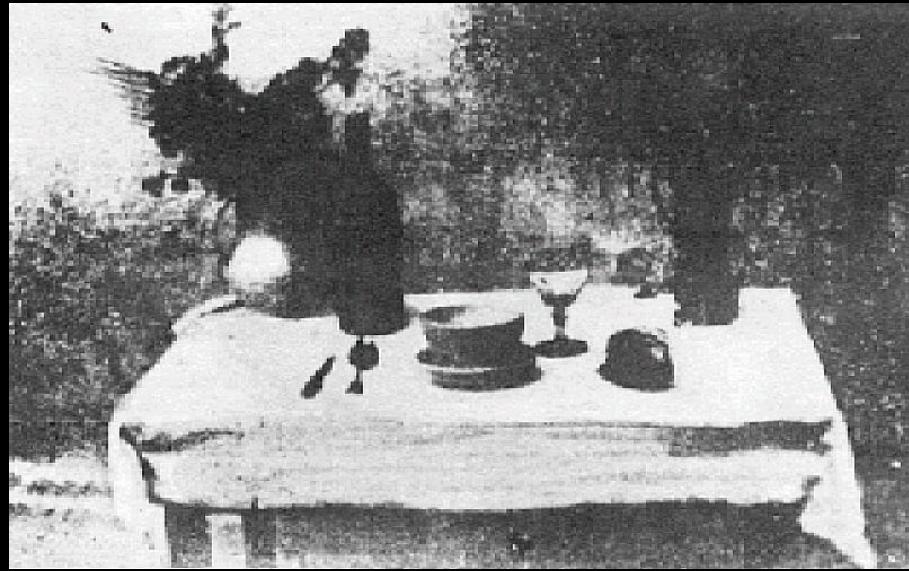


# Challenge: Illumination - What is wrong with these pictures?



## Photography

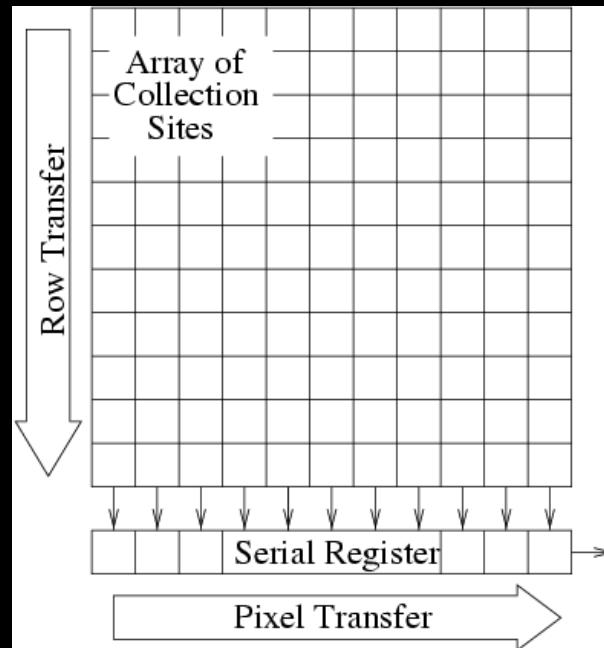
(Niepce, "La Table Servie," 1822)



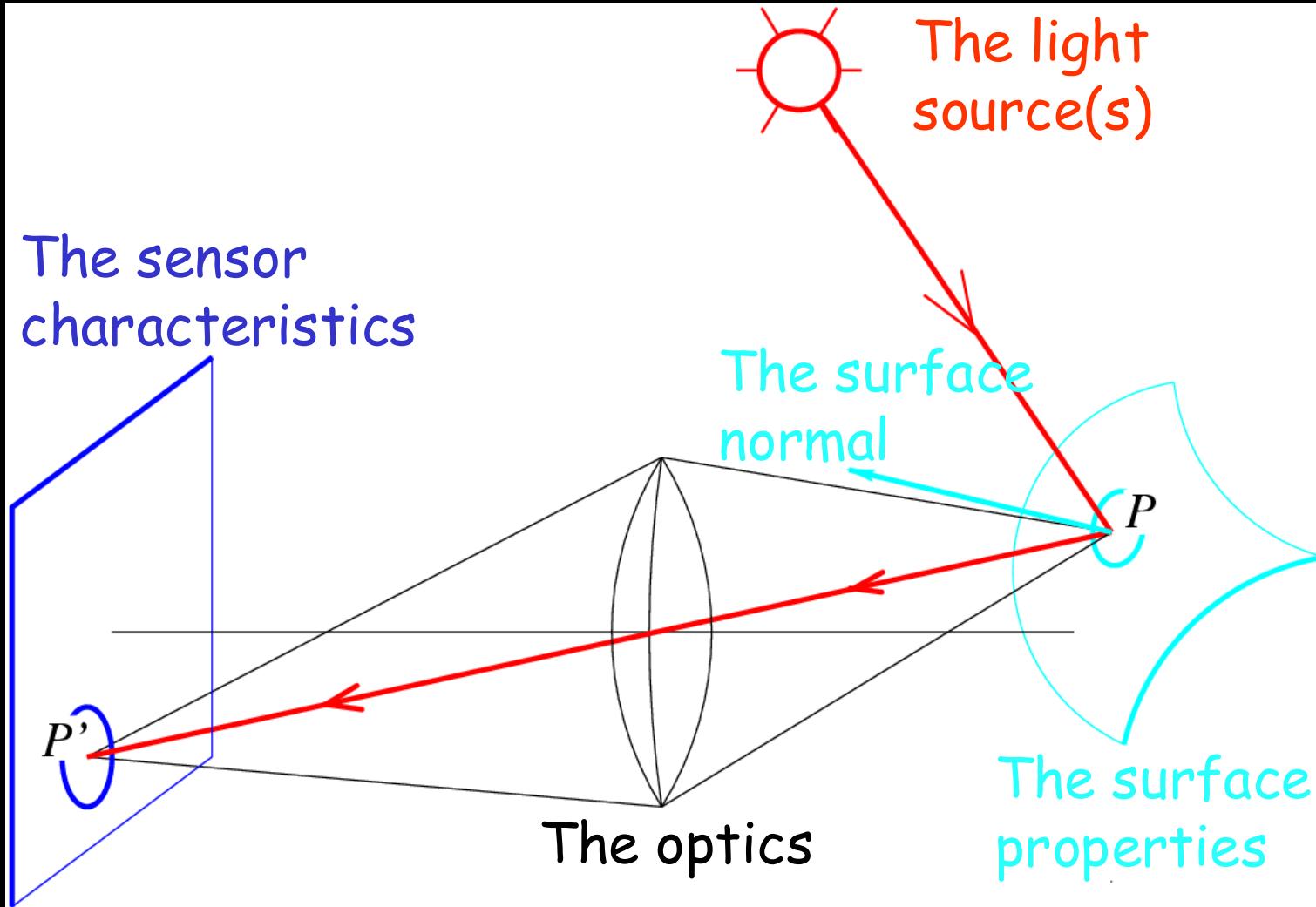
## Milestones:

- Daguerreotypes (1839)
- Photographic Film (Eastman, 1889)
- Cinema (Lumière brothers, 1895)
- Color Photography (Lumière brothers, again, 1908)
- Television (Baird, Farnsworth, Zworykin, 1920s)

CCD Devices (1970), etc.

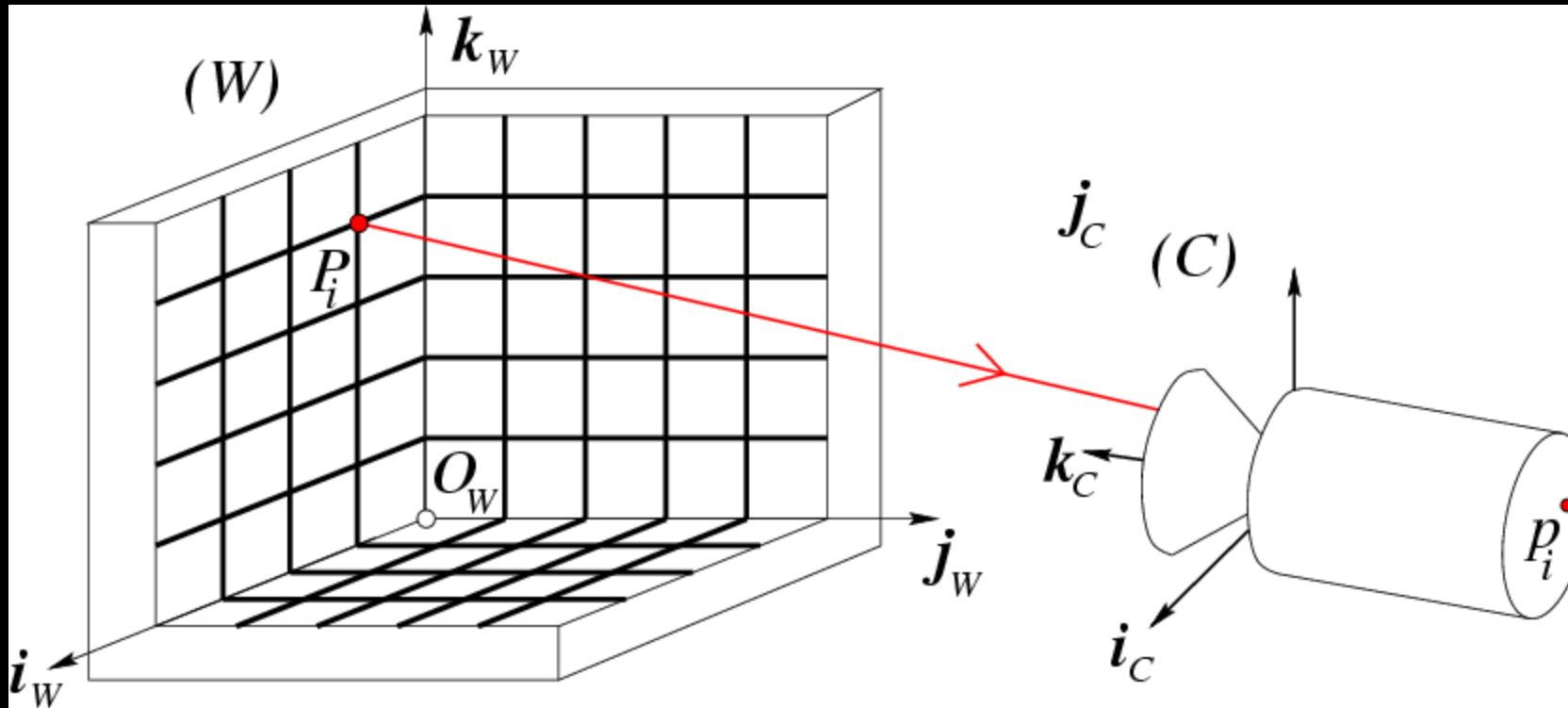


# Image Formation: Radiometry



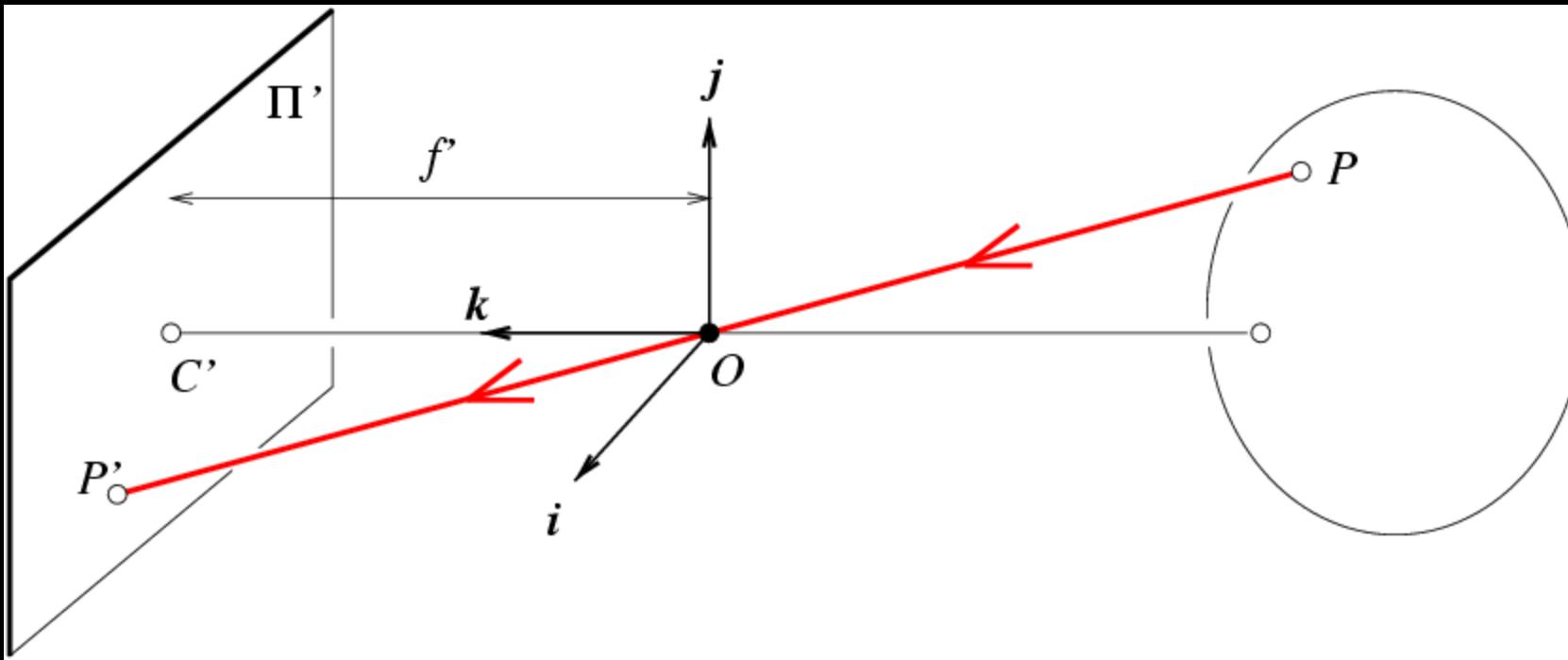
What determines the brightness of an image pixel?

# Quantitative Measurements and Calibration



Euclidean Geometry

# Pinhole Perspective Equation



$$\begin{cases} x' = f' \frac{x}{z} \\ y' = f' \frac{y}{z} \end{cases}$$

$$\begin{array}{c} u = f \frac{x}{z} \\ v = f \frac{y}{z} \end{array}$$

$$p = \frac{1}{z} P$$

$$p = \begin{bmatrix} u \\ v \\ 1 \end{bmatrix}$$
$$P = \begin{bmatrix} x \\ y \\ z \end{bmatrix}$$

# Homogeneous coordinates

## Conversion

Converting to *homogeneous* coordinates

$$(x, y) \Rightarrow \begin{bmatrix} x \\ y \\ 1 \end{bmatrix}$$

homogeneous image  
coordinates

$$(x, y, z) \Rightarrow \begin{bmatrix} x \\ y \\ z \\ 1 \end{bmatrix}$$

homogeneous scene  
coordinates

Converting *from* homogeneous coordinates

$$\begin{bmatrix} x \\ y \\ w \end{bmatrix} \Rightarrow (x/w, y/w)$$

$$\begin{bmatrix} x \\ y \\ z \\ w \end{bmatrix} \Rightarrow (x/w, y/w, z/w)$$

# Homogeneous coordinates

Invariant to scaling

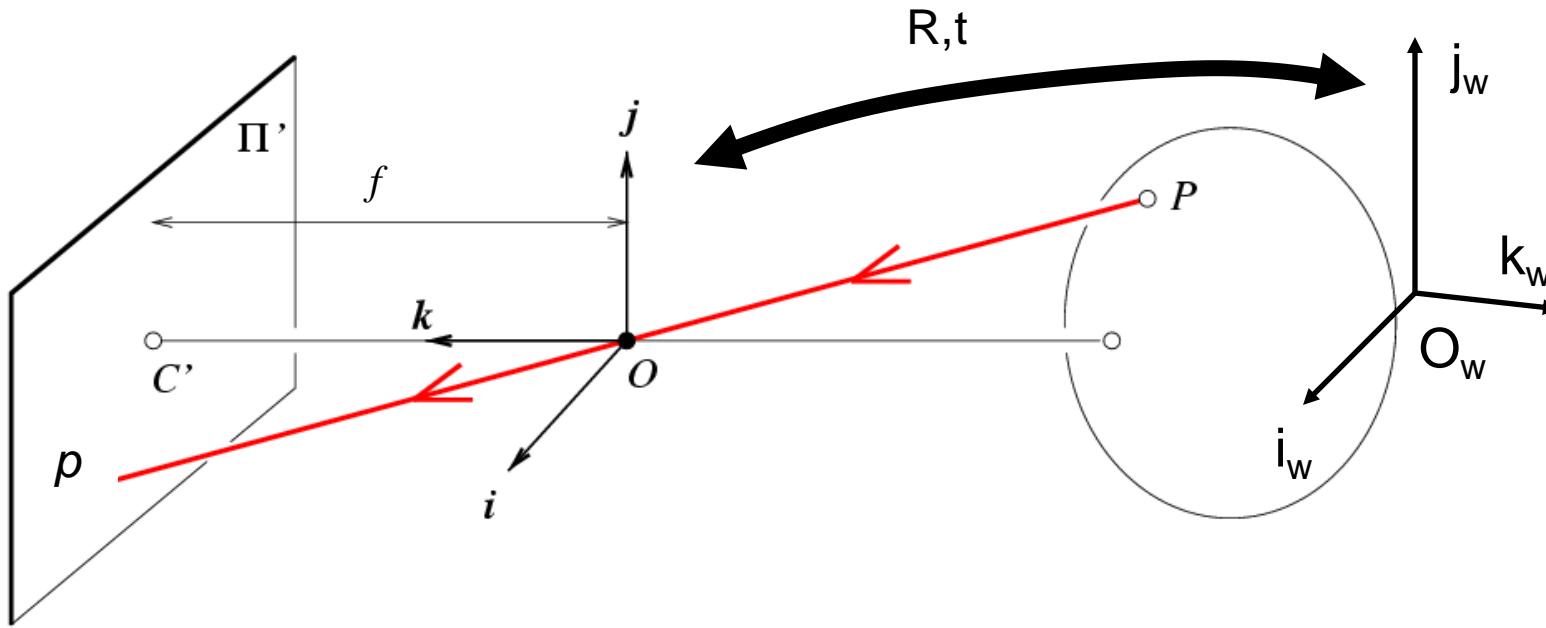
$$k \begin{bmatrix} x \\ y \\ w \end{bmatrix} = \begin{bmatrix} kx \\ ky \\ kw \end{bmatrix} \Rightarrow \begin{bmatrix} \frac{kx}{kw} \\ \frac{ky}{kw} \end{bmatrix} = \begin{bmatrix} \frac{x}{w} \\ \frac{y}{w} \end{bmatrix}$$

Homogeneous  
Coordinates

Cartesian  
Coordinates

A point in Cartesian coordinates is a ray in homogeneous ones

# Projection matrix



$$p \approx MP = K [R \ t]P$$



$$p = \lambda MP \text{ for some } \lambda \neq 0$$

$p$ : Image Coordinates:  $(u, v, 1)$

$M$ :  $3 \times 4$  projection matrix

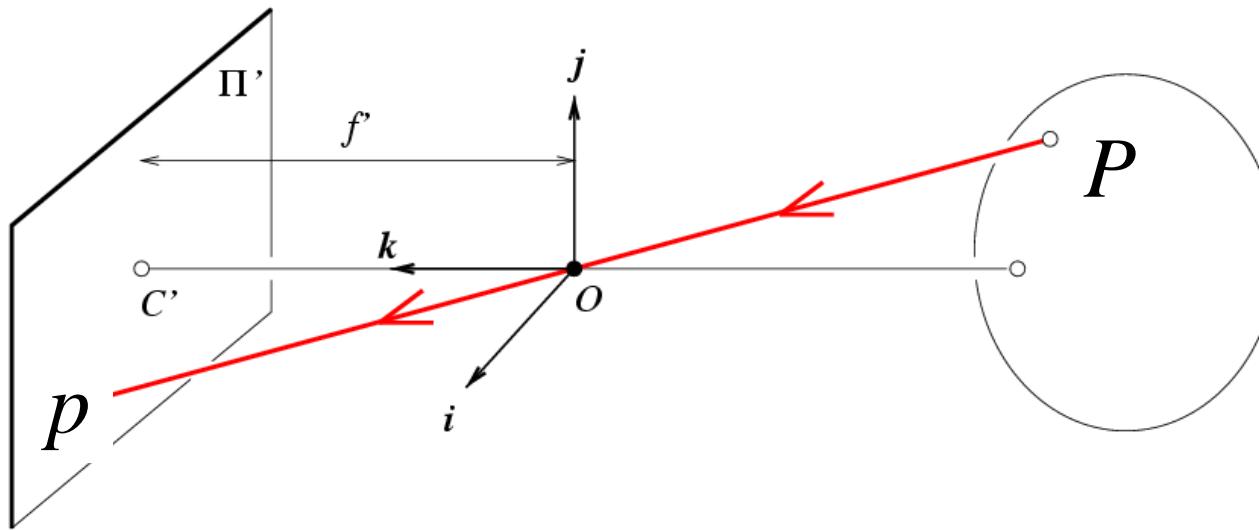
$K$ : Intrinsic Matrix ( $3 \times 3$ )

$R$ : Rotation ( $3 \times 3$ )

$t$ : Translation ( $3 \times 1$ )

$P$ : World Coordinates:  $(x, y, z, 1)$

# Projection matrix



## Intrinsic Assumptions

- Unit aspect ratio
- Image center at  $(0,0)$
- No skew

## Extrinsic Assumptions

- No rotation
- Camera at  $(0,0,0)$

$$p \approx K [I \ 0]P$$

(Note: here  $w = z$ )

$$w \begin{bmatrix} u \\ v \\ 1 \end{bmatrix} = \begin{bmatrix} f & 0 & 0 & 0 \\ 0 & f & 0 & 0 \\ 0 & 0 & 1 & 0 \end{bmatrix} \begin{bmatrix} x \\ y \\ z \\ 1 \end{bmatrix}$$

# Remove assumption: known image center

## Intrinsic Assumptions

- Unit aspect ratio
- No skew

## Extrinsic Assumptions

- No rotation
- Camera at (0,0,0)

$$p \approx K [I \ 0]P \quad \xrightarrow{\text{w}} \quad w \begin{bmatrix} u \\ v \\ 1 \end{bmatrix} = \begin{bmatrix} f & 0 & u_0 \\ 0 & f & v_0 \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} x \\ y \\ z \\ 1 \end{bmatrix}$$

# Remove assumption: square pixels

## Intrinsic Assumptions

- No skew

## Extrinsic Assumptions

- No rotation
- Camera at (0,0,0)

$$p \approx K [I \ 0]P \quad \xrightarrow{\text{blue arrow}} \quad w \begin{bmatrix} u \\ v \\ 1 \end{bmatrix} = \begin{bmatrix} \alpha & 0 & u_0 \\ 0 & \beta & v_0 \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} x \\ y \\ z \\ 1 \end{bmatrix}$$

# Remove assumption: rectangular pixels

Intrinsic Assumptions

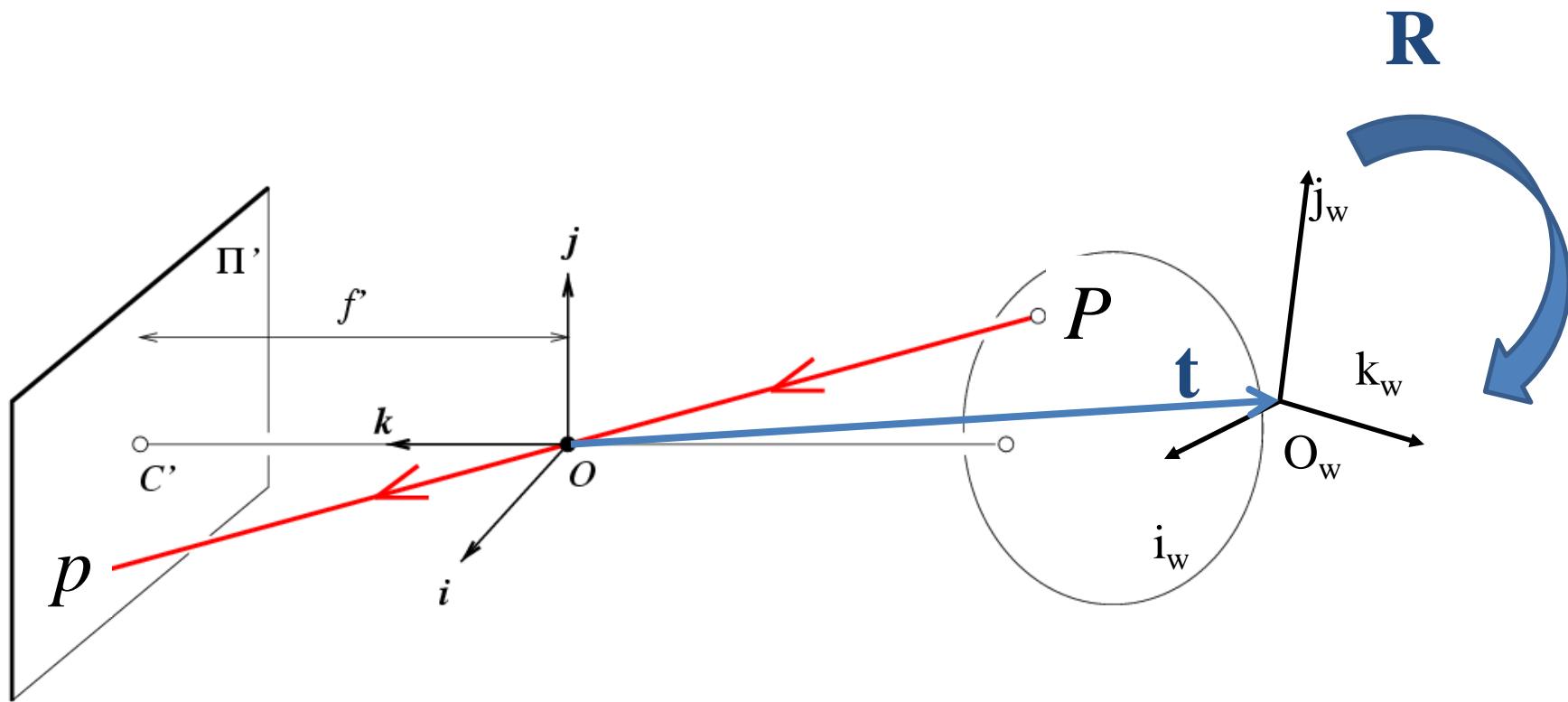
Extrinsic Assumptions

- No rotation
- Camera at (0,0,0)

$$p \approx K [I \ 0]P \rightarrow w \begin{bmatrix} u \\ v \\ 1 \end{bmatrix} = \begin{bmatrix} \alpha & s & u_0 \\ 0 & \beta & v_0 \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} x \\ y \\ z \\ 1 \end{bmatrix}$$

Note: different books use different notation for parameters

# Oriented and Translated Camera



# Allow camera translation

Intrinsic Assumptions

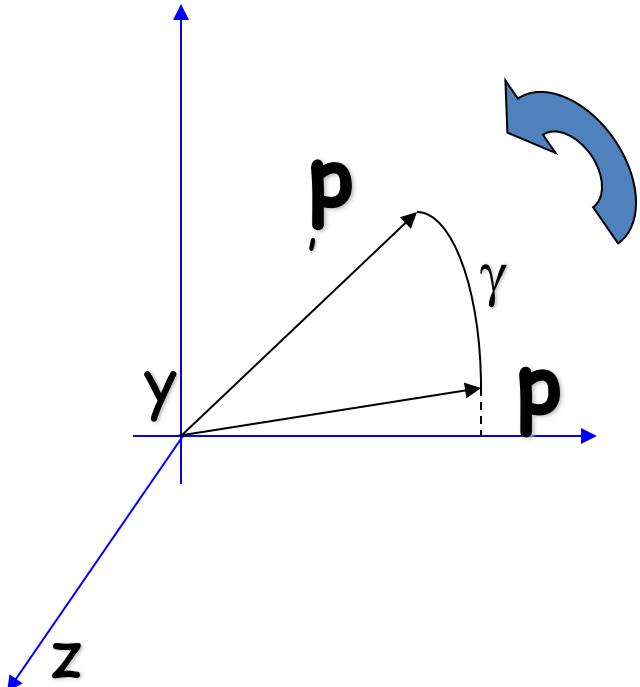
Extrinsic Assumptions

- No rotation

$$p \approx K [I \ t]P \quad \xrightarrow{\text{blue arrow}} \ w \begin{bmatrix} u \\ v \\ 1 \end{bmatrix} = \begin{bmatrix} \alpha & 0 & u_0 \\ 0 & \beta & v_0 \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} 1 & 0 & 0 & t_x \\ 0 & 1 & 0 & t_y \\ 0 & 0 & 1 & t_z \end{bmatrix} \begin{bmatrix} x \\ y \\ z \\ 1 \end{bmatrix}$$

# 3D Rotation of Points

Rotation around the coordinate axes, **counter-clockwise**:



$$R_x(\alpha) = \begin{bmatrix} 1 & 0 & 0 \\ 0 & \cos \alpha & -\sin \alpha \\ 0 & \sin \alpha & \cos \alpha \end{bmatrix}$$

$$R_y(\beta) = \begin{bmatrix} \cos \beta & 0 & \sin \beta \\ 0 & 1 & 0 \\ -\sin \beta & 0 & \cos \beta \end{bmatrix}$$

$$R_z(\gamma) = \begin{bmatrix} \cos \gamma & -\sin \gamma & 0 \\ \sin \gamma & \cos \gamma & 0 \\ 0 & 0 & 1 \end{bmatrix}$$

# Allow camera rotation

$$p \approx K [R \ t]P$$



$$w \begin{bmatrix} u \\ v \\ 1 \end{bmatrix} = \begin{bmatrix} \alpha & s & u_0 \\ 0 & \beta & v_0 \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} r_{11} & r_{12} & r_{13} & t_x \\ r_{21} & r_{22} & r_{23} & t_y \\ r_{31} & r_{32} & r_{33} & t_z \end{bmatrix} \begin{bmatrix} x \\ y \\ z \\ 1 \end{bmatrix}$$

# Degrees of freedom

$$p \approx K [R \ t]P$$



$$w \begin{bmatrix} u \\ v \\ 1 \end{bmatrix} = \begin{bmatrix} \alpha & s & u_0 \\ 0 & \beta & v_0 \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} 5 & r_{11} & r_{12} & r_{13} & t_x \\ 6 & r_{21} & r_{22} & r_{23} & t_y \\ x & r_{31} & r_{32} & r_{33} & t_z \\ y & z \\ z \\ 1 \end{bmatrix}$$

$$p \approx MP \text{ or } p = \frac{1}{z} MP$$

$$M = K [R \ t]$$

$$p = K\hat{p} \text{ and } \hat{p} = \frac{1}{z} \widehat{M}P$$

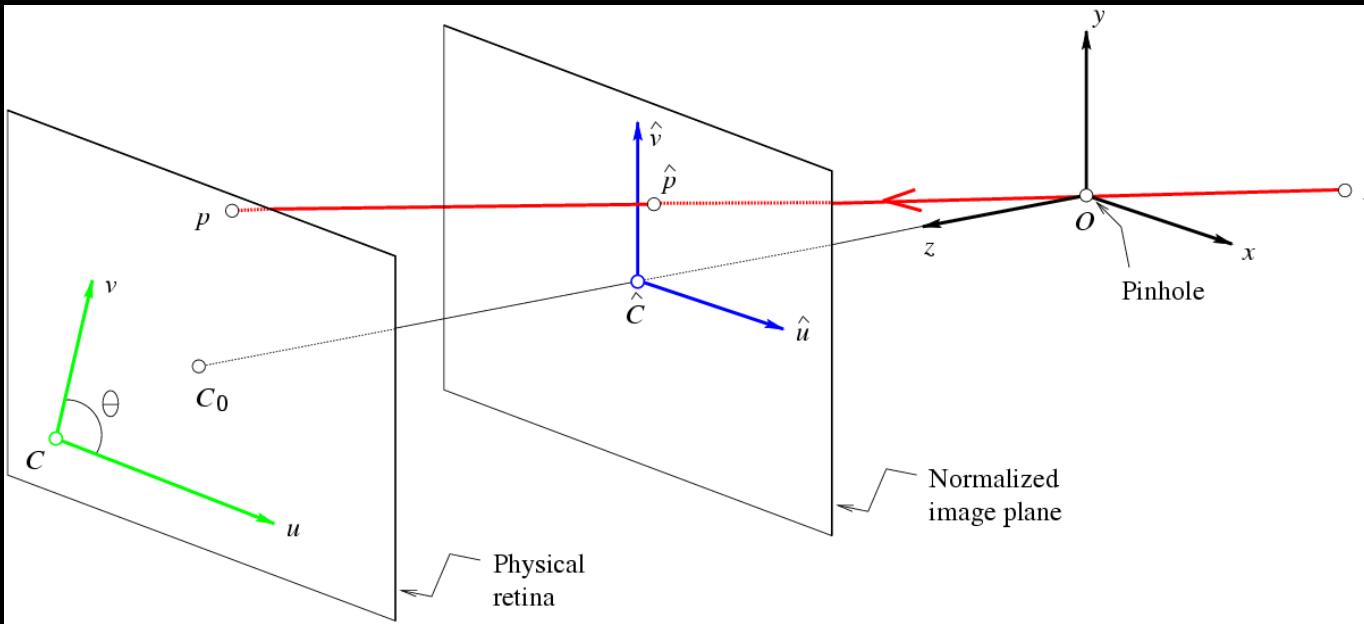
$$\widehat{M} = [R \ t]$$

normalized coordinates

Note:  $z$  is *not* independent of  $\mathcal{M}$  and  $\mathbf{P}$ :

$$\mathcal{M} = \begin{pmatrix} \mathbf{m}_1^T \\ \mathbf{m}_2^T \\ \mathbf{m}_3^T \end{pmatrix} \implies z = \mathbf{m}_3 \cdot \mathbf{P}, \quad \text{or} \quad \begin{cases} u = \frac{\mathbf{m}_1 \cdot \mathbf{P}}{\mathbf{m}_3 \cdot \mathbf{P}}, \\ v = \frac{\mathbf{m}_2 \cdot \mathbf{P}}{\mathbf{m}_3 \cdot \mathbf{P}}. \end{cases}$$

# Explicit form of the projection matrix



$$\mathcal{M} = \begin{pmatrix} \alpha \mathbf{r}_1^T - \alpha \cot \theta \mathbf{r}_2^T + u_0 \mathbf{r}_3^T & \alpha t_x - \alpha \cot \theta t_y + u_0 t_z \\ \frac{\beta}{\sin \theta} \mathbf{r}_2^T + v_0 \mathbf{r}_3^T & \frac{\beta}{\sin \theta} t_y + v_0 t_z \\ \mathbf{r}_3^T & t_z \end{pmatrix}$$

## Explicit Form of the Projection Matrix

$$\mathcal{M} = \begin{pmatrix} \alpha \mathbf{r}_1^T - \alpha \cot \theta \mathbf{r}_2^T + u_0 \mathbf{r}_3^T & \alpha t_x - \alpha \cot \theta t_y + u_0 t_z \\ \frac{\beta}{\sin \theta} \mathbf{r}_2^T + v_0 \mathbf{r}_3^T & \frac{\beta}{\sin \theta} t_y + v_0 t_z \\ \mathbf{r}_3^T & t_z \end{pmatrix}$$

Note: If  $\mathcal{M} = (\mathcal{A} \quad \mathbf{b})$  then  $|\mathbf{a}_3| = 1$ .

Replacing  $\mathcal{M}$  by  $\lambda \mathcal{M}$  in

$$\begin{cases} u = \frac{\mathbf{m}_1 \cdot \mathbf{P}}{\mathbf{m}_3 \cdot \mathbf{P}} \\ v = \frac{\mathbf{m}_2 \cdot \mathbf{P}}{\mathbf{m}_3 \cdot \mathbf{P}} \end{cases}$$

does not change  $u$  and  $v$ .

$\mathcal{M}$  is only defined up to scale in this setting!!

## Theorem (Faugeras, 1993)

Let  $\mathcal{M} = (\mathcal{A} \quad \mathbf{b})$  be a  $3 \times 4$  matrix and let  $\mathbf{a}_i^T$  ( $i = 1, 2, 3$ ) denote the rows of the matrix  $\mathcal{A}$  formed by the three leftmost columns of  $\mathcal{M}$ .

- A necessary and sufficient condition for  $\mathcal{M}$  to be a perspective projection matrix is that  $\text{Det}(\mathcal{A}) \neq 0$ .
- A necessary and sufficient condition for  $\mathcal{M}$  to be a zero-skew perspective projection matrix is that  $\text{Det}(\mathcal{A}) \neq 0$  and

$$(\mathbf{a}_1 \times \mathbf{a}_3) \cdot (\mathbf{a}_2 \times \mathbf{a}_3) = 0.$$

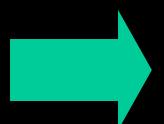
- A necessary and sufficient condition for  $\mathcal{M}$  to be a perspective projection matrix with zero skew and unit aspect-ratio is that  $\text{Det}(\mathcal{A}) \neq 0$  and

$$\begin{cases} (\mathbf{a}_1 \times \mathbf{a}_3) \cdot (\mathbf{a}_2 \times \mathbf{a}_3) = 0, \\ (\mathbf{a}_1 \times \mathbf{a}_3) \cdot (\mathbf{a}_1 \times \mathbf{a}_3) = (\mathbf{a}_2 \times \mathbf{a}_3) \cdot (\mathbf{a}_2 \times \mathbf{a}_3). \end{cases}$$

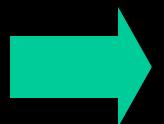
## Linear Camera Calibration

Given  $n$  points  $P_1, \dots, P_n$  with *known* positions and their images  $p_1, \dots, p_n$

Remember:  $a \cdot b = a^T b$



$$\begin{pmatrix} u_i \\ v_i \end{pmatrix} = \begin{pmatrix} \mathbf{m}_1 \cdot \mathbf{P}_i \\ \mathbf{m}_3 \cdot \mathbf{P}_i \\ \mathbf{m}_2 \cdot \mathbf{P}_i \\ \mathbf{m}_3 \cdot \mathbf{P}_i \end{pmatrix} \iff \begin{pmatrix} \mathbf{m}_1^T - u_i \mathbf{m}_3^T \\ \mathbf{m}_2^T - v_i \mathbf{m}_3^T \end{pmatrix} \mathbf{P}_i = 0$$



$$\mathcal{P}\mathbf{m} = 0 \text{ with } \mathcal{P} \stackrel{\text{def}}{=} \begin{pmatrix} \mathbf{P}_1^T & \mathbf{0}^T & -u_1 \mathbf{P}_1^T \\ \mathbf{0}^T & \mathbf{P}_1^T & -v_1 \mathbf{P}_1^T \\ \dots & \dots & \dots \\ \mathbf{P}_n^T & \mathbf{0}^T & -u_n \mathbf{P}_n^T \\ \mathbf{0}^T & \mathbf{P}_n^T & -v_n \mathbf{P}_n^T \end{pmatrix} \text{ and } \mathbf{m} \stackrel{\text{def}}{=} \begin{pmatrix} \mathbf{m}_1 \\ \mathbf{m}_2 \\ \mathbf{m}_3 \end{pmatrix} = 0$$

# Homogeneous Linear Systems

$$\begin{bmatrix} A \end{bmatrix} \begin{bmatrix} x \end{bmatrix} = \begin{bmatrix} 0 \end{bmatrix}$$

Square system:

- unique solution: 0
- unless  $\text{Det}(A)=0$

$$\begin{bmatrix} A \\ A \\ A \end{bmatrix} \begin{bmatrix} x \end{bmatrix} = \begin{bmatrix} 0 \\ 0 \\ 0 \end{bmatrix}$$

Rectangular system ??

- 0 is always a solution



Minimize  $||Ax||^2$   
under the constraint  
 $||x||^2 = 1$

How do you solve overconstrained homogeneous linear equations ?? Homogeneous linear least squares

$$E = |\mathcal{U}\mathbf{x}|^2 = \mathbf{x}^T(\mathcal{U}^T\mathcal{U})\mathbf{x}$$

- Orthonormal basis of eigenvectors:  $\mathbf{e}_1, \dots, \mathbf{e}_q$ .
- Associated eigenvalues:  $0 \leq \lambda_1 \leq \dots \leq \lambda_q$ .
- Any vector can be written as

$$\mathbf{x} = \mu_1 \mathbf{e}_1 + \dots + \mu_q \mathbf{e}_q$$

for some  $\mu_i$  ( $i = 1, \dots, q$ ) such that  $\mu_1^2 + \dots + \mu_q^2 = 1$ .

$$\begin{aligned} E(\mathbf{x}) - E(\mathbf{e}_1) &= \mathbf{x}^T(\mathcal{U}^T\mathcal{U})\mathbf{x} - \mathbf{e}_1^T(\mathcal{U}^T\mathcal{U})\mathbf{e}_1 \\ &= \lambda_1 \mu_1^2 + \dots + \lambda_q \mu_q^2 - \lambda_1 \\ &\geq \lambda_1 (\mu_1^2 + \dots + \mu_q^2 - 1) = 0 \end{aligned}$$

The solution is  $\mathbf{e}_1$ .

# Linear Camera Calibration

Given  $n$  points  $P_1, \dots, P_n$  with *known* positions and their images  $p_1, \dots, p_n$

$$\xrightarrow{\hspace{1cm}} \begin{pmatrix} u_i \\ v_i \end{pmatrix} = \begin{pmatrix} \mathbf{m}_1 \cdot \mathbf{P}_i \\ \mathbf{m}_3 \cdot \mathbf{P}_i \\ \mathbf{m}_2 \cdot \mathbf{P}_i \\ \mathbf{m}_3 \cdot \mathbf{P}_i \end{pmatrix} \iff (\mathbf{m}_1 - u_i \mathbf{m}_3) \mathbf{P}_i = 0$$

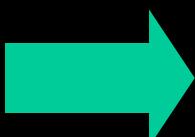
$$\xrightarrow{\hspace{1cm}} \mathcal{P}\mathbf{m} = 0 \text{ with } \mathcal{P} \stackrel{\text{def}}{=} \begin{pmatrix} \mathbf{P}_1^T & \mathbf{0}^T & -u_1 \mathbf{P}_1^T \\ \mathbf{0}^T & \mathbf{P}_1^T & -v_1 \mathbf{P}_1^T \\ \dots & \dots & \dots \\ \mathbf{P}_n^T & \mathbf{0}^T & -u_n \mathbf{P}_n^T \\ \mathbf{0}^T & \mathbf{P}_n^T & -v_n \mathbf{P}_n^T \end{pmatrix} \text{ and } \mathbf{m} \stackrel{\text{def}}{=} \begin{pmatrix} \mathbf{m}_1 \\ \mathbf{m}_2 \\ \mathbf{m}_3 \end{pmatrix} = 0$$

$\xrightarrow{\hspace{1cm}}$  Minimize  $||\mathcal{P}\mathbf{m}||^2$  under the constraint  $||\mathbf{m}||^2 = 1$

Once  $M$  is known, you still got to recover the intrinsic and extrinsic parameters !!!

This is a decomposition problem, **not** an estimation problem.

$$\rho \mathcal{M} = \begin{pmatrix} \alpha \mathbf{r}_1^T - \alpha \cot \theta \mathbf{r}_2^T + u_0 \mathbf{r}_3^T & \alpha t_x - \alpha \cot \theta t_y + u_0 t_z \\ \frac{\beta}{\sin \theta} \mathbf{r}_2^T + v_0 \mathbf{r}_3^T & \frac{\beta}{\sin \theta} t_y + v_0 t_z \\ \mathbf{r}_3^T & t_z \end{pmatrix}$$



- Intrinsic parameters
- Extrinsic parameters

# Weak-Perspective Projection Model

$$\mathbf{p} = \frac{1}{z} \mathcal{M} \mathbf{P}$$

( $\mathbf{p}$  and  $\mathbf{P}$  are in homogeneous coordinates)



$$\mathbf{p} = \mathcal{M} \mathbf{P}$$

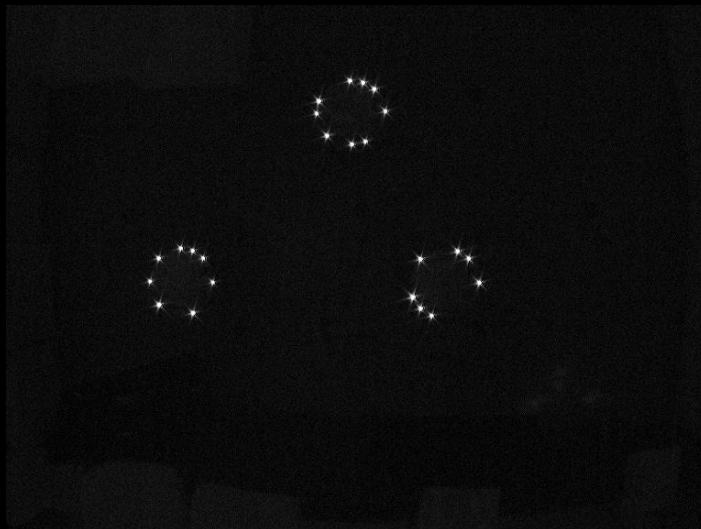
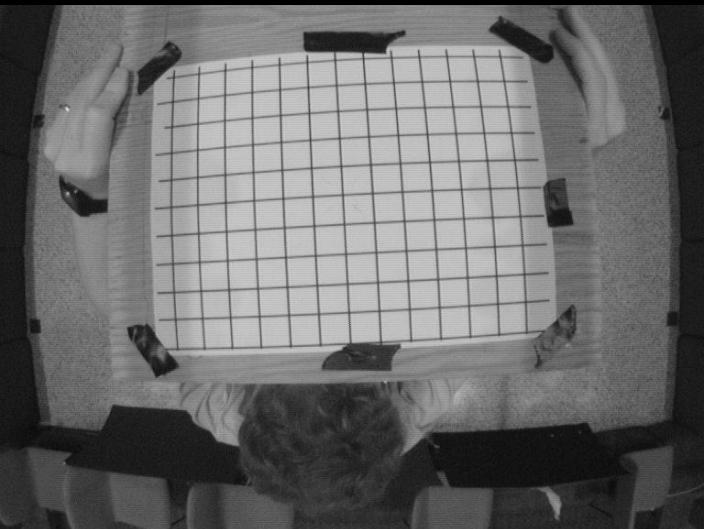
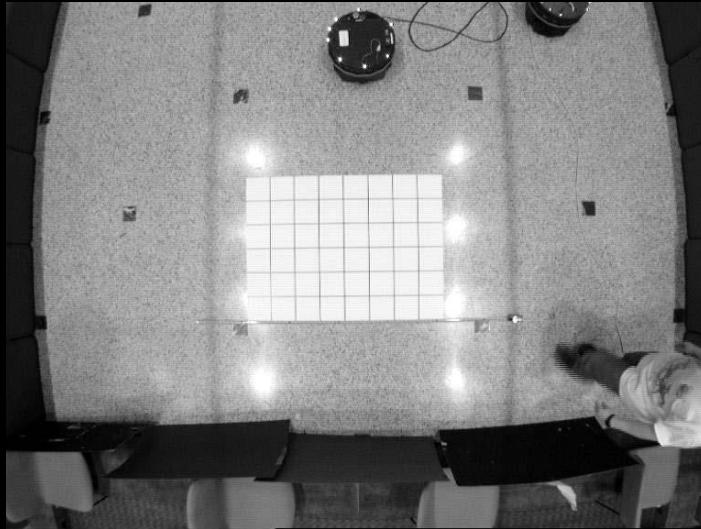
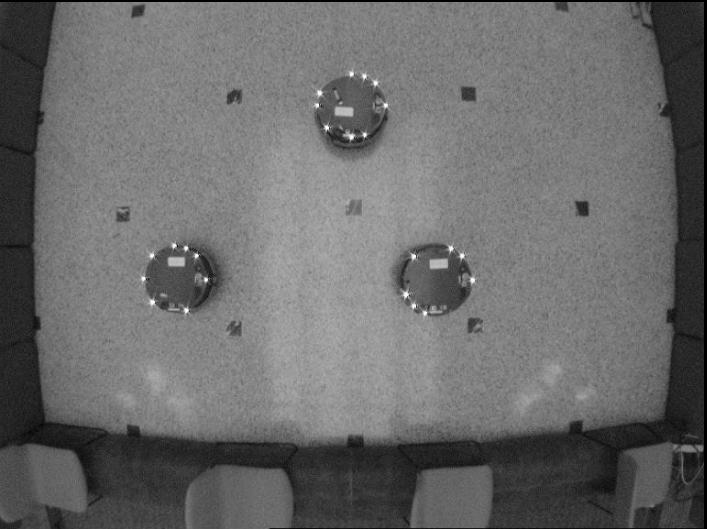
( $\mathbf{P}$  is in homogeneous coordinates)



$$\mathbf{p} = \mathbf{A} \mathbf{P} + \mathbf{b}$$

(neither  $\mathbf{p}$  nor  $\mathbf{P}$  is in hom. coordinates)

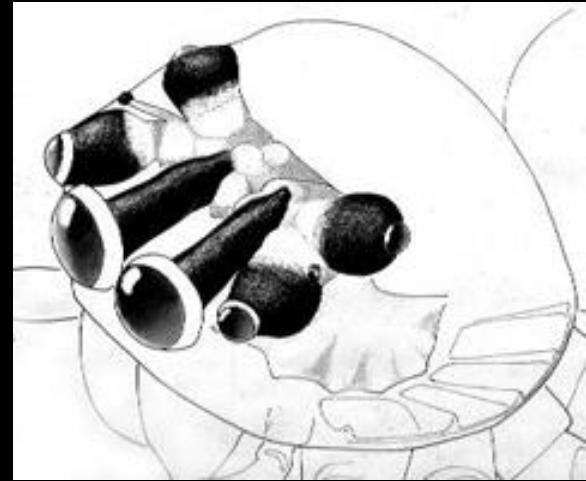
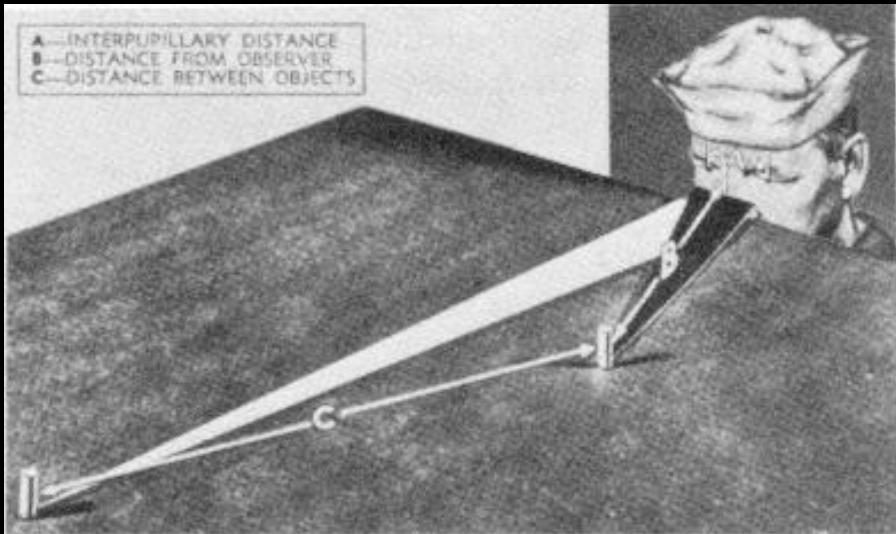
## Applications: Mobile Robot Localization (Devy et al., 1997)

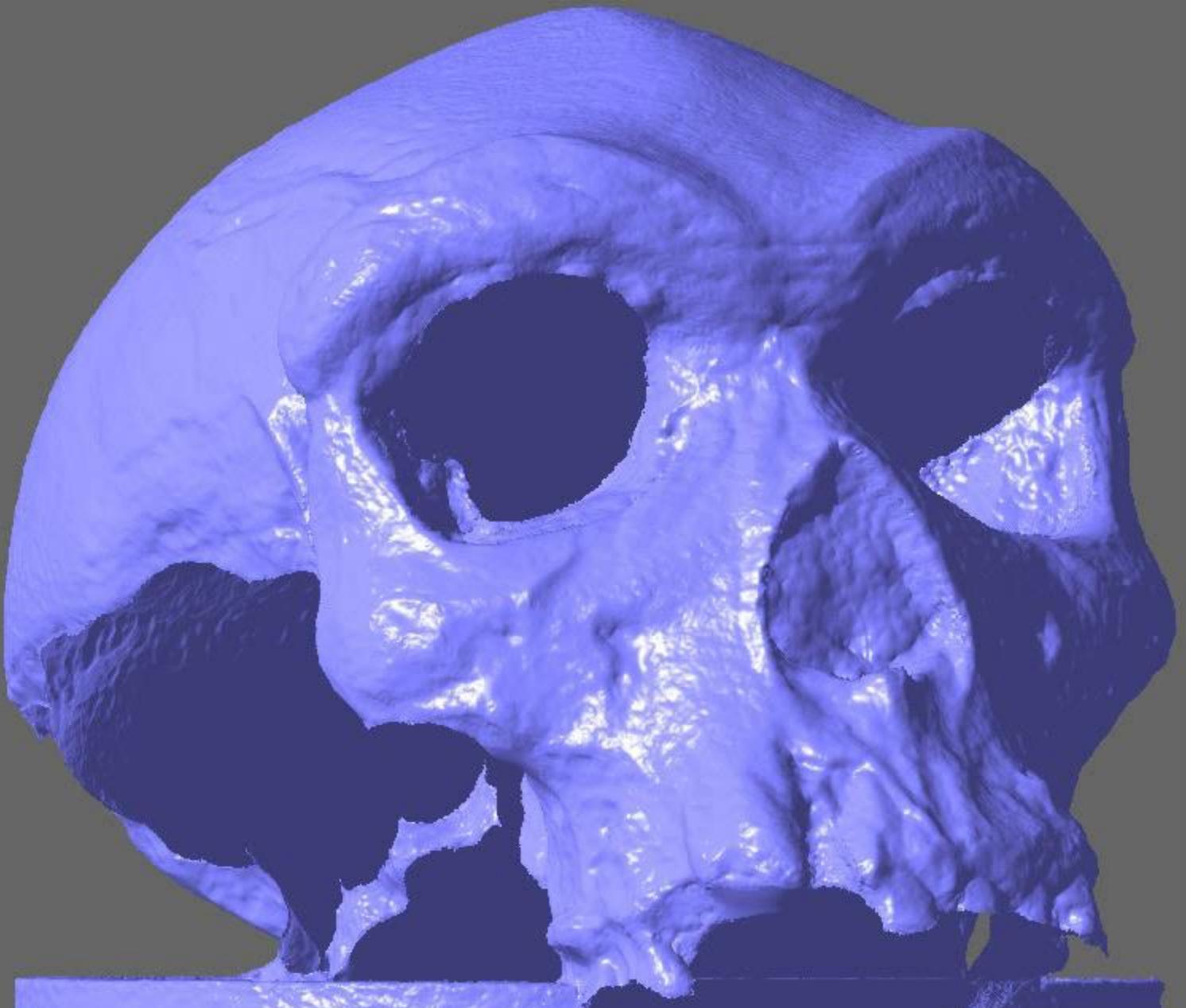




(Rothganger, Sudsang, Ponce, 2002)

# How do we perceive depth?





PMVS (Furukawa & Ponce, 2007)

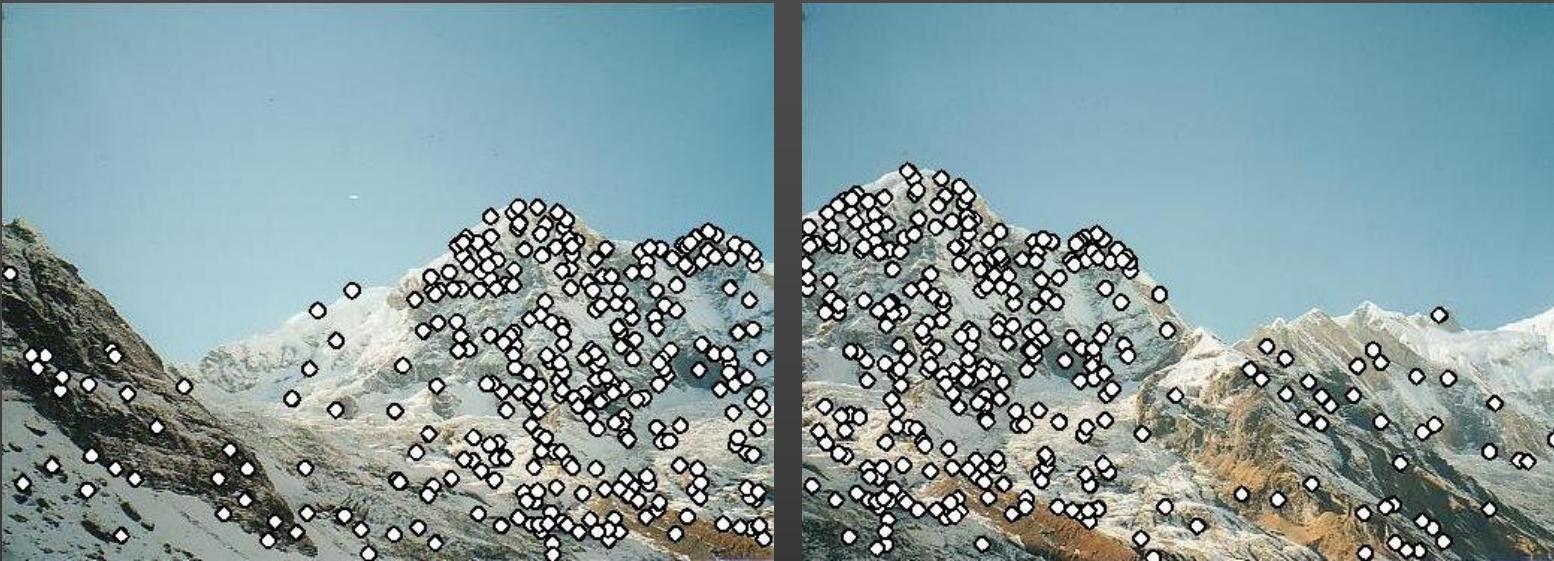
# Feature-based alignment outline

---



# Feature-based alignment outline

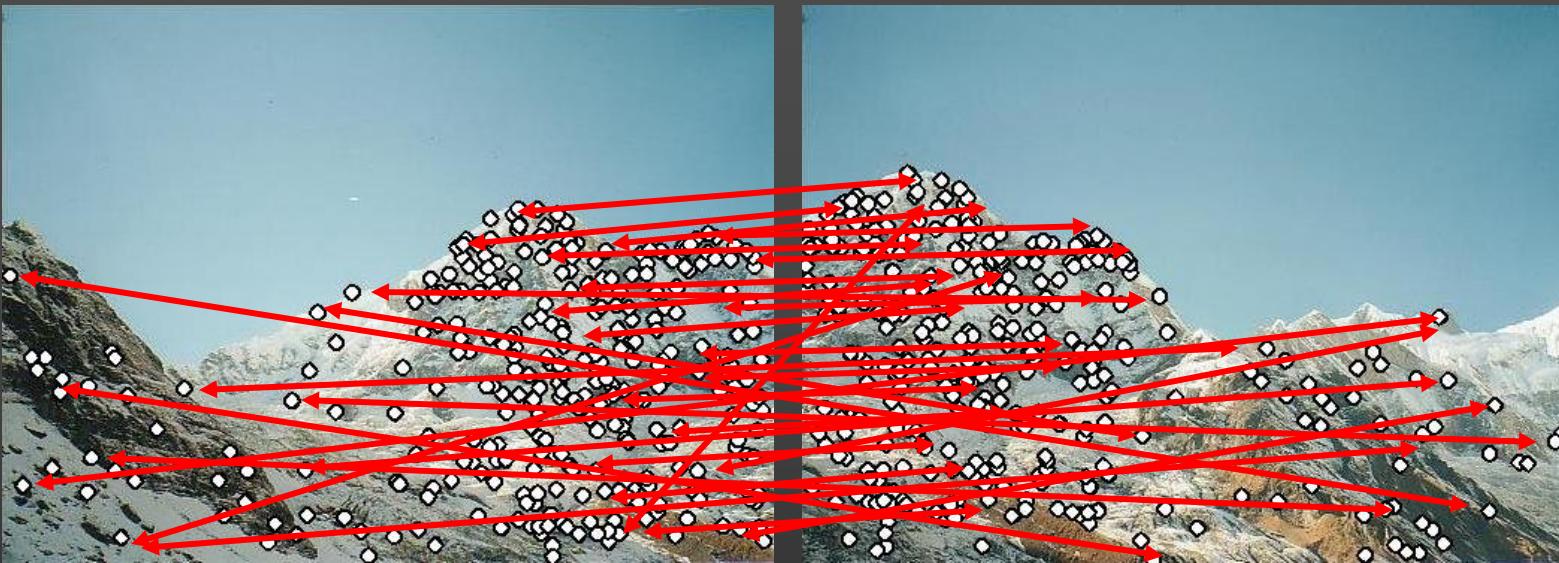
---



Extract features

# Feature-based alignment outline

---

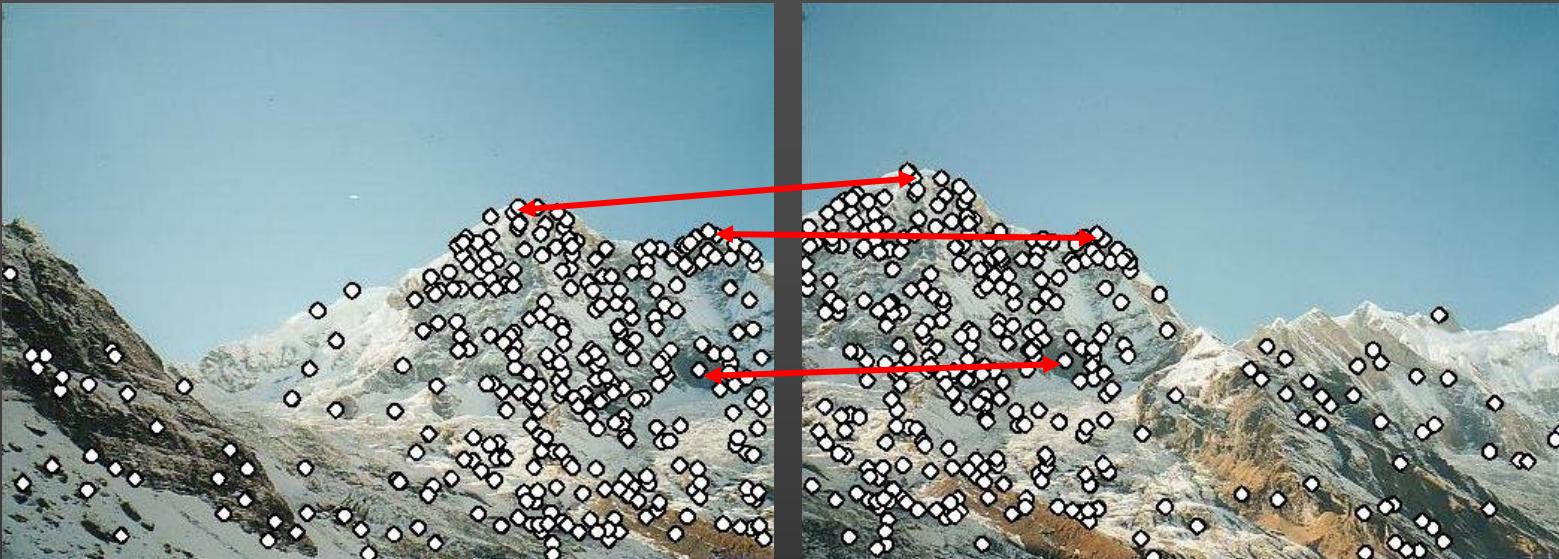


Extract features

Compute *putative matches*

# Feature-based alignment outline

---



Extract features

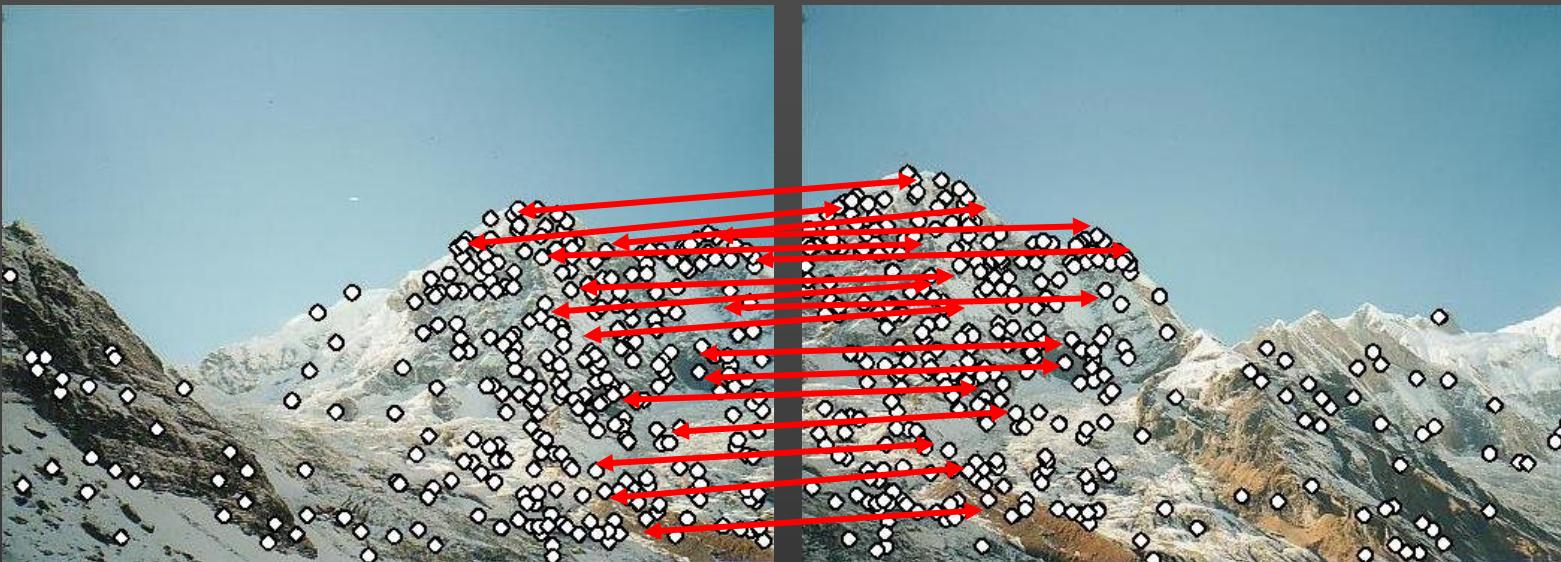
Compute *putative matches*

Loop:

- *Hypothesize* transformation  $T$  (small group of putative matches that are related by  $T$ )

# Feature-based alignment outline

---



Extract features

Compute *putative matches*

Loop:

- *Hypothesize* transformation  $T$  (small group of putative matches that are related by  $T$ )
- *Verify* transformation (search for other matches consistent with  $T$ )

# Feature-based alignment outline

---

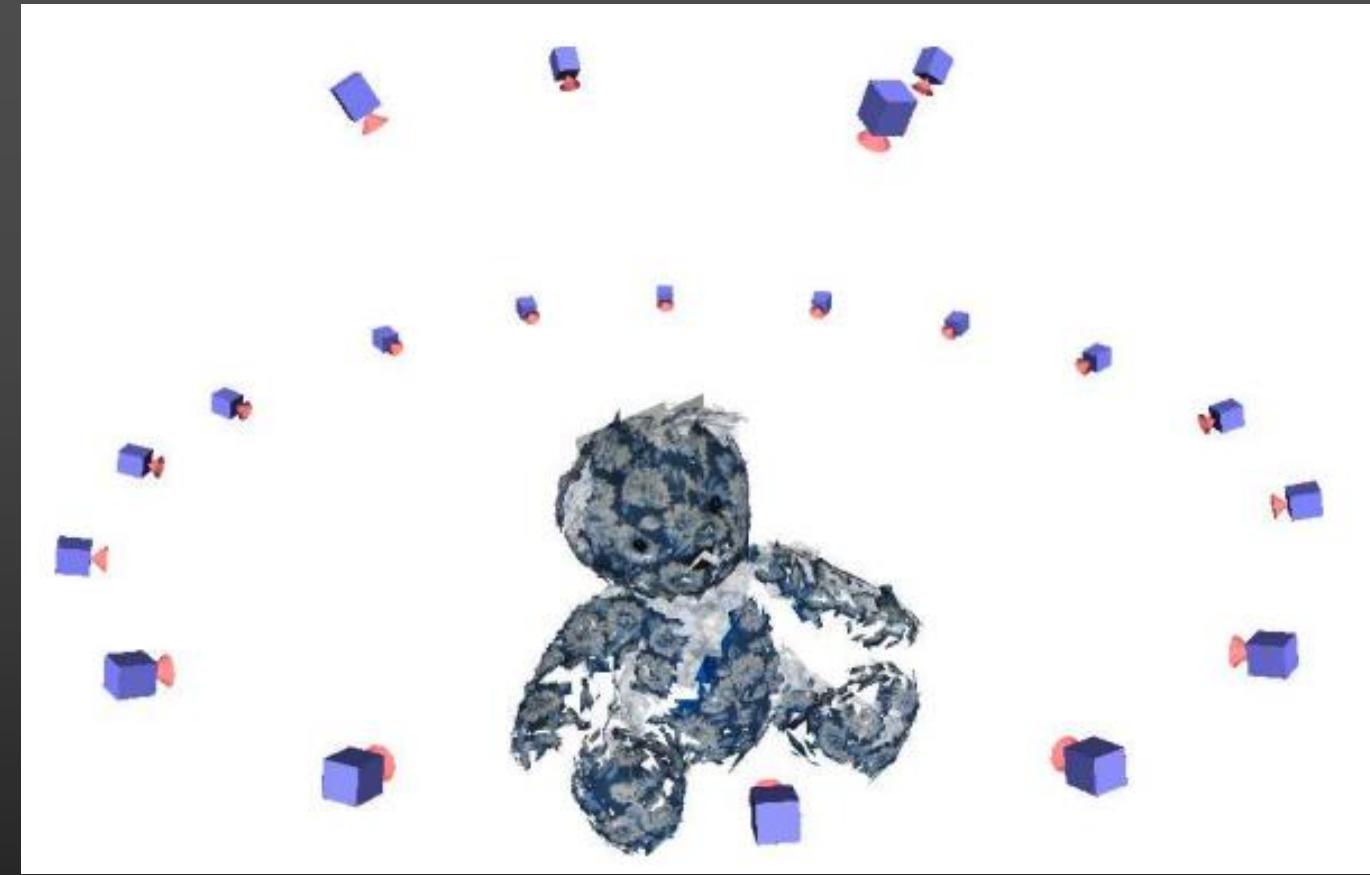


Extract features

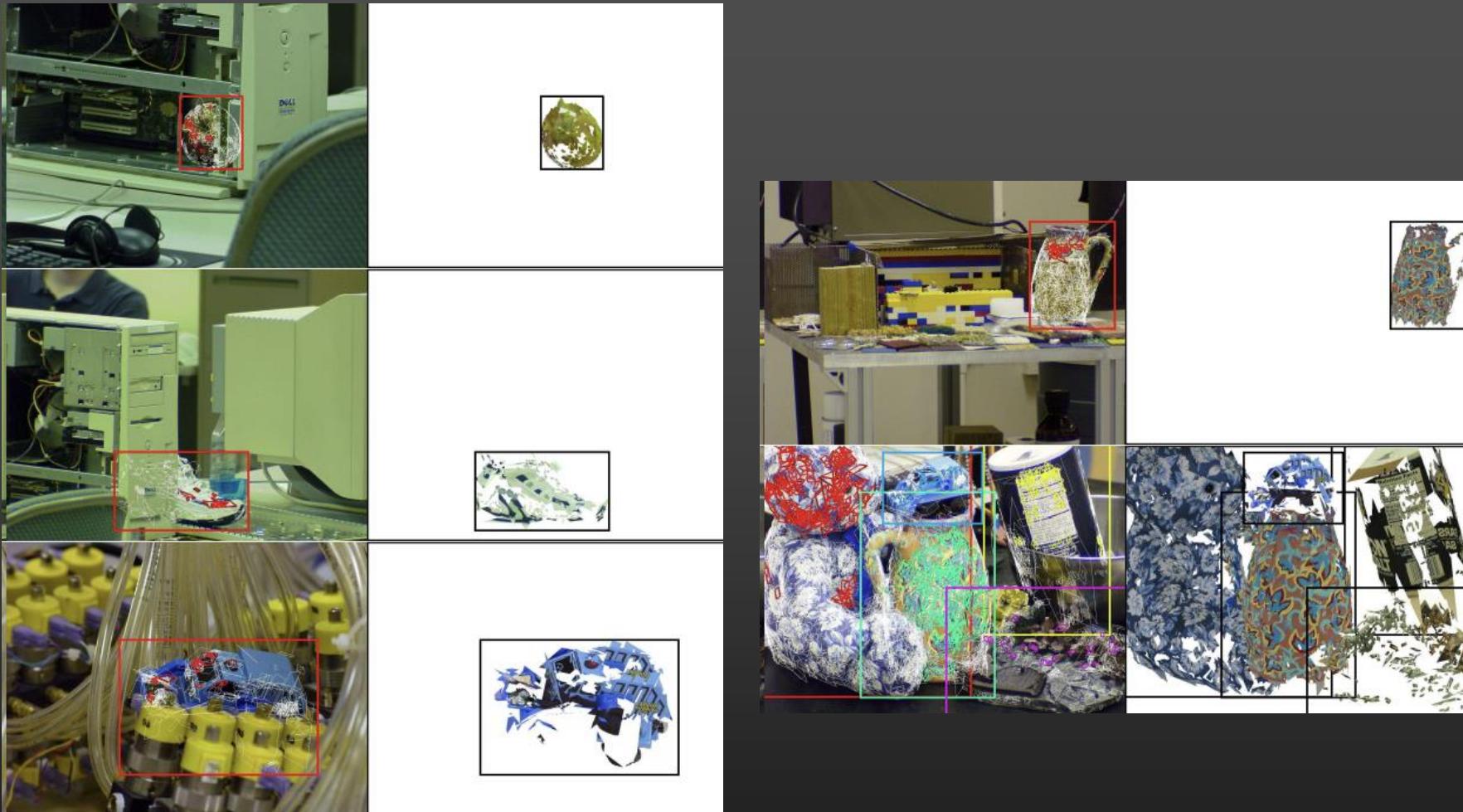
Compute *putative matches*

Loop:

- *Hypothesize* transformation  $T$  (small group of putative matches that are related by  $T$ )
- *Verify* transformation (search for other matches consistent with  $T$ )



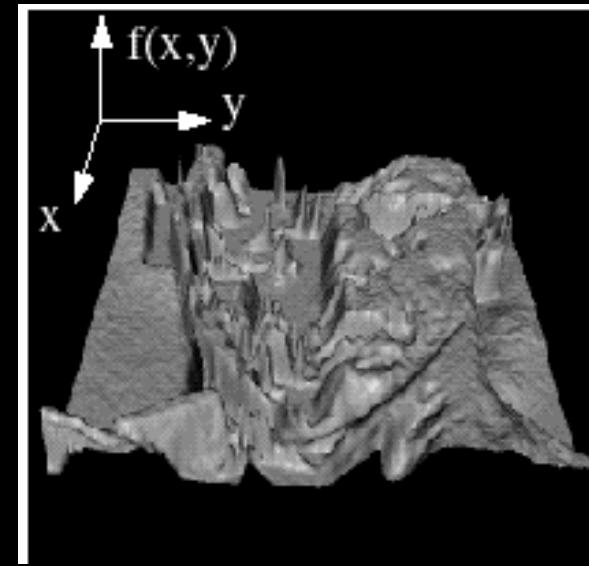
3D object modeling from multiple images  
(Rothganger et al., 2003)



Recognition examples with major clutter and partial occlusion (Rothganger et al., 2003)

# Image processing

- Filters and convolution
- Derivatives and edge detection
- The Canny edge detector
- Denoising, sparsity and dictionary learning
- Super-resolution



An image can be interpreted either as:

- a continuous function  $f(x, y)$
- a discrete array  $F_{u,v}$

# Basic Filters



# Convolution

Linear filters = Weighted averages

- Represent the weights by a rectangular array  $F$ .
- Applying the filter to an image  $G$  is equivalent to performing a **convolution**:

$$R_{ij} = (F * G)_{ij} = \sum_{u,v} F_{i-u, j-v} G_{u, v}$$

- In the continuous case:

$$(f * g)(x,y) = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} f(x-u, y-v) g(u, v) du dv$$

- Note:  $f * g = g * f$ .

Original Image

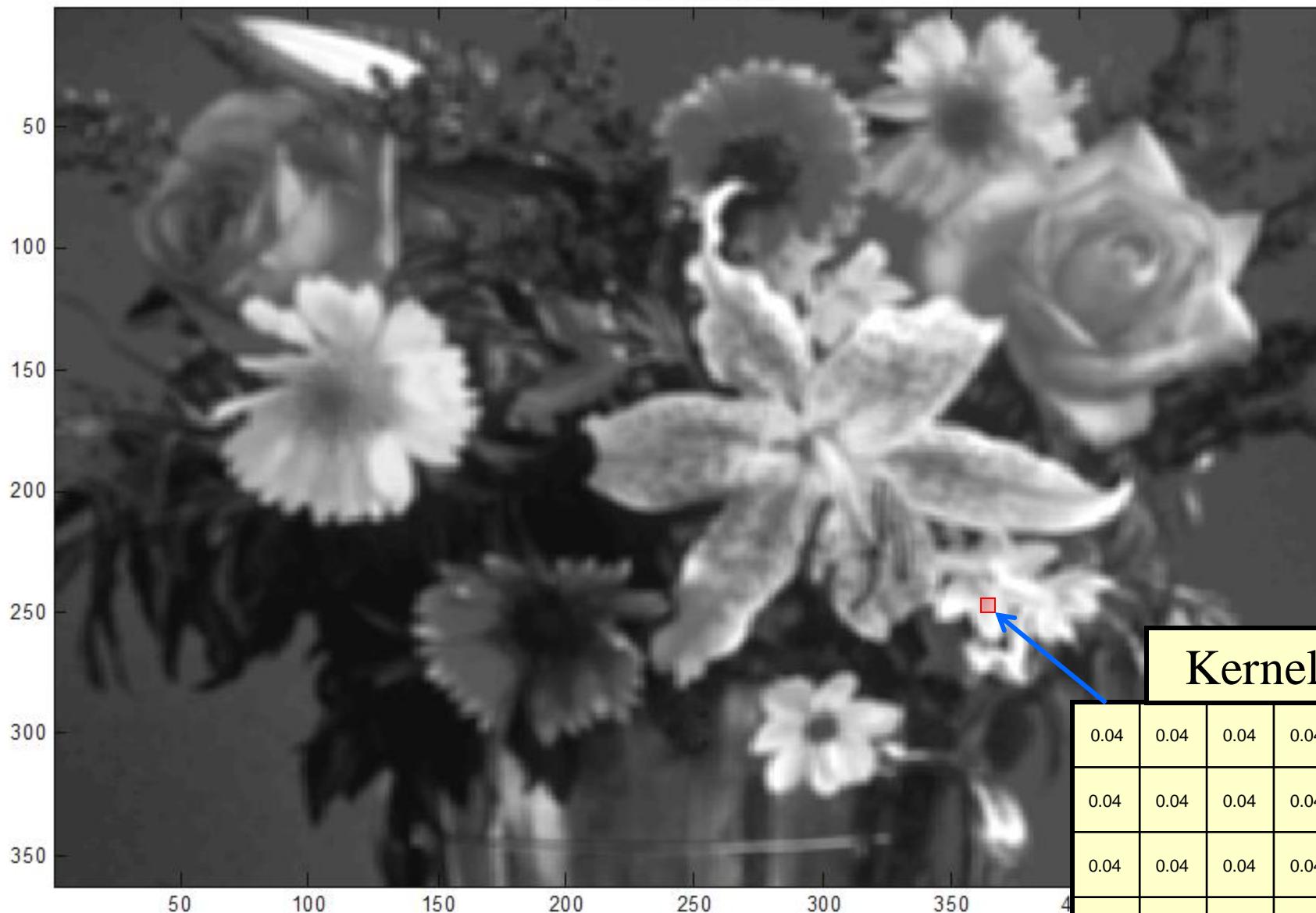


Slight Blurring



| Kernel: |     |     |
|---------|-----|-----|
| 1/9     | 1/9 | 1/9 |
| 1/9     | 1/9 | 1/9 |
| 1/9     | 1/9 | 1/9 |

More Blurring



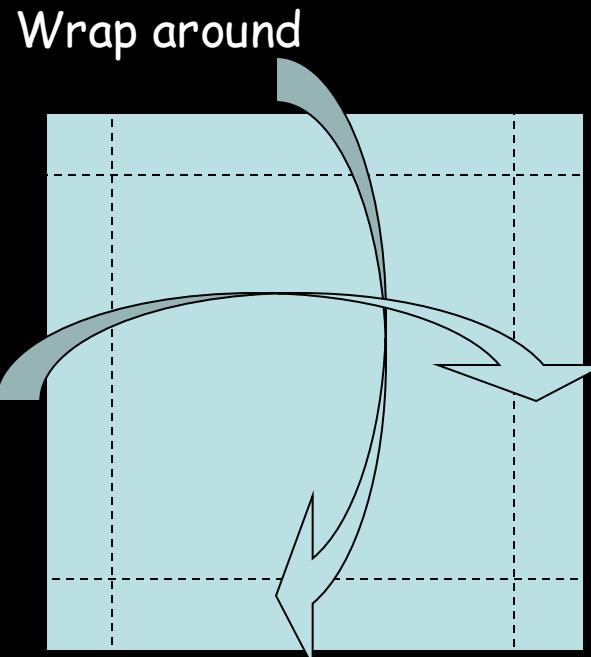
# Basic Properties

- Commutativity:  $f * g = g * f$
- Associativity:  $(f * g) * h = f * (g * h)$
- Linearity:  $(af + bg) * h = af * h + bg * h$
- Shift invariance:  $f_+ * h = (f * h)_+$
- Only operator both linear and shift invariant
- Differentiation:

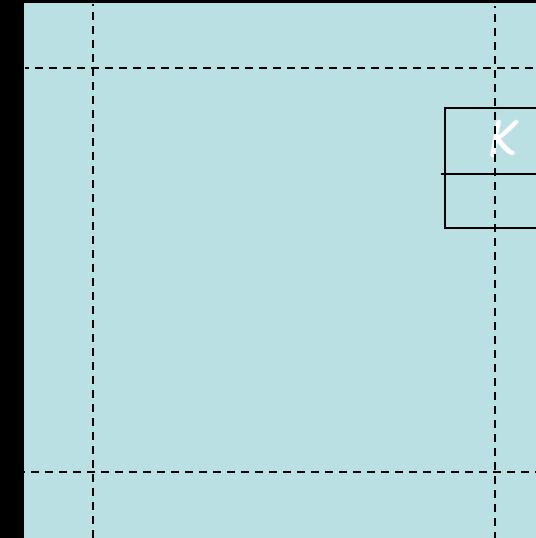
$$\frac{\partial}{\partial x} (f * g) = \frac{\partial f}{\partial x} * g$$

# Practicalities (discrete convolution)

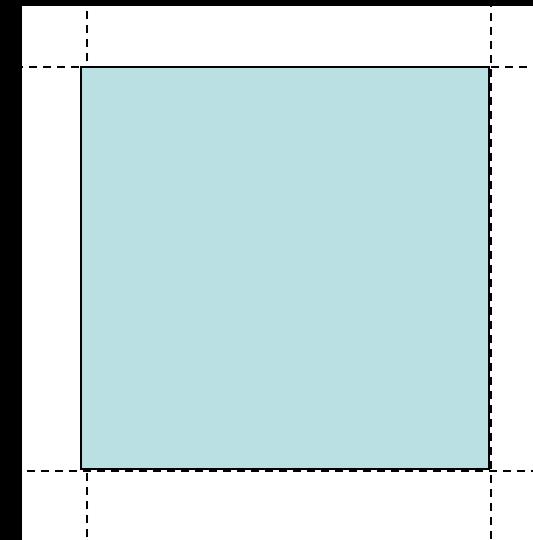
- Python: convolve function
- Border issues:
  - When applying convolution with a  $K \times K$  kernel, the result is undefined for pixels closer than  $K$  pixels from the border of the image
- Options:



Expand/Pad



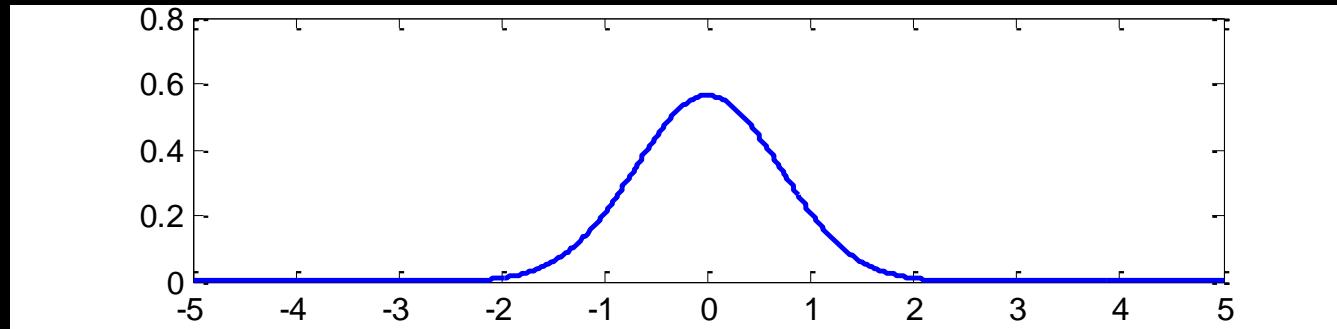
Crop



# Gaussian filters

1-D:

$$g(x) = e^{-\frac{x^2}{2\sigma^2}}$$

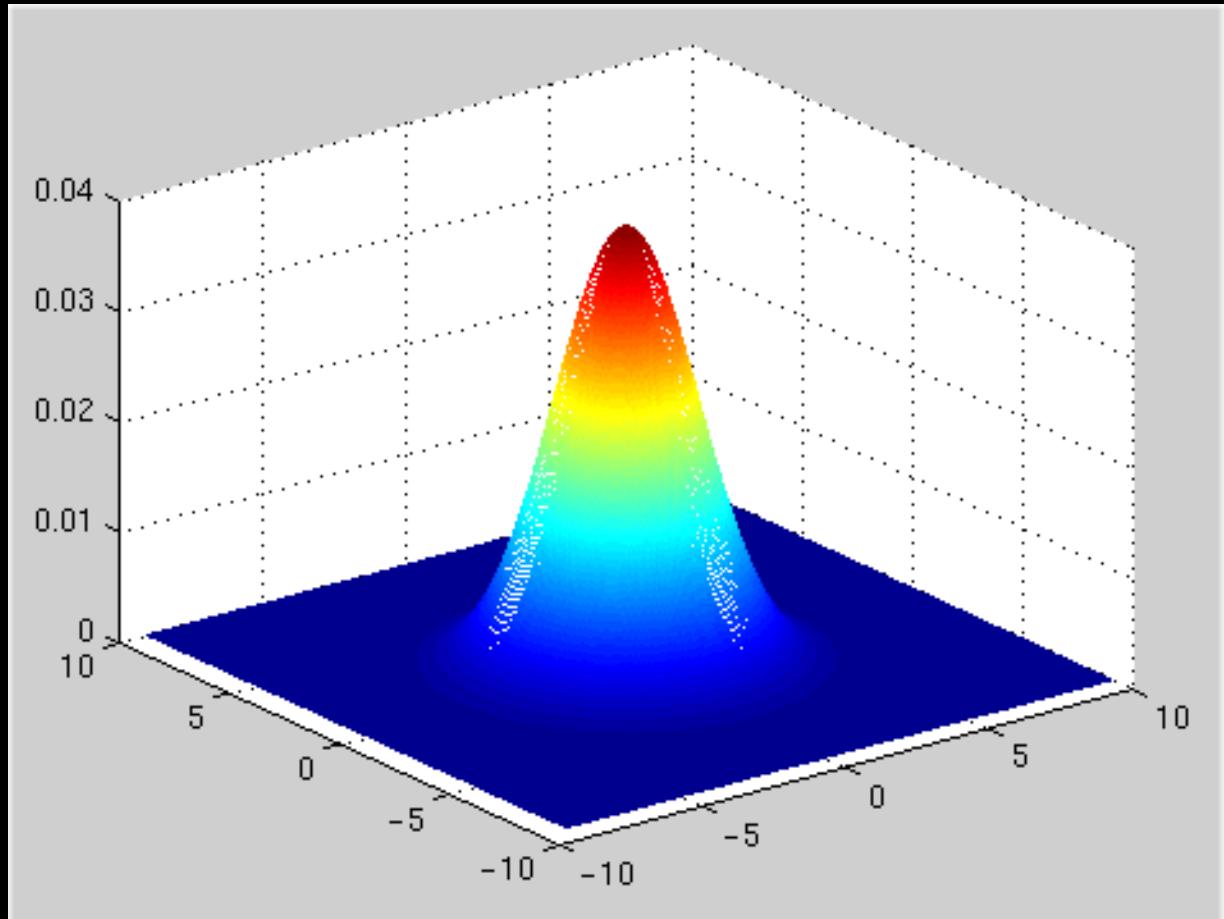


2-D:

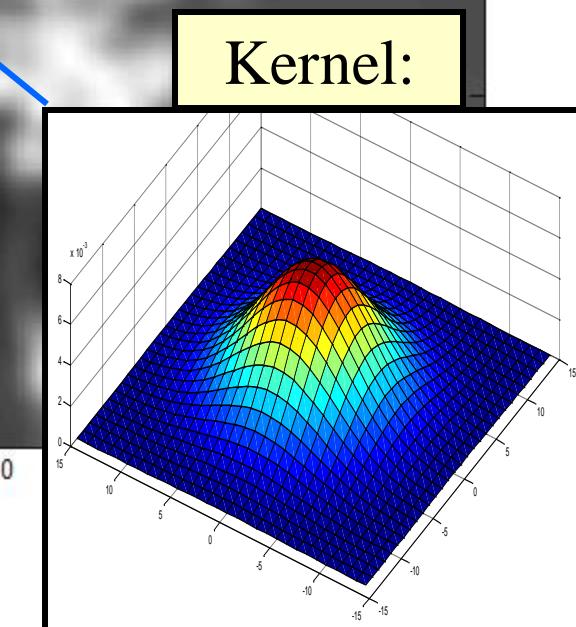
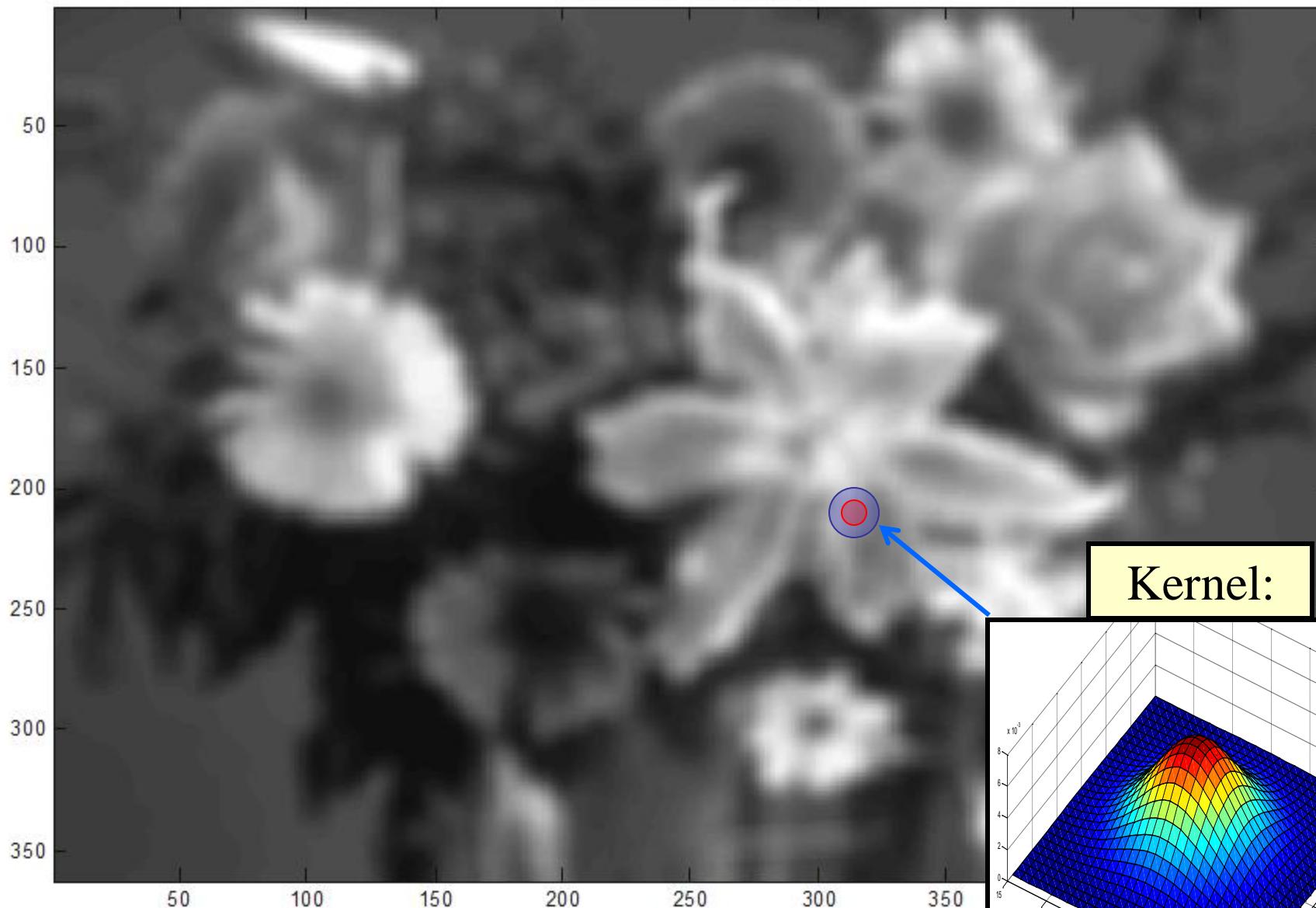
$$G(x, y) = e^{-\frac{x^2+y^2}{2\sigma^2}}$$

Slight abuse of notation:  
We ignore the normalization  
constant such that

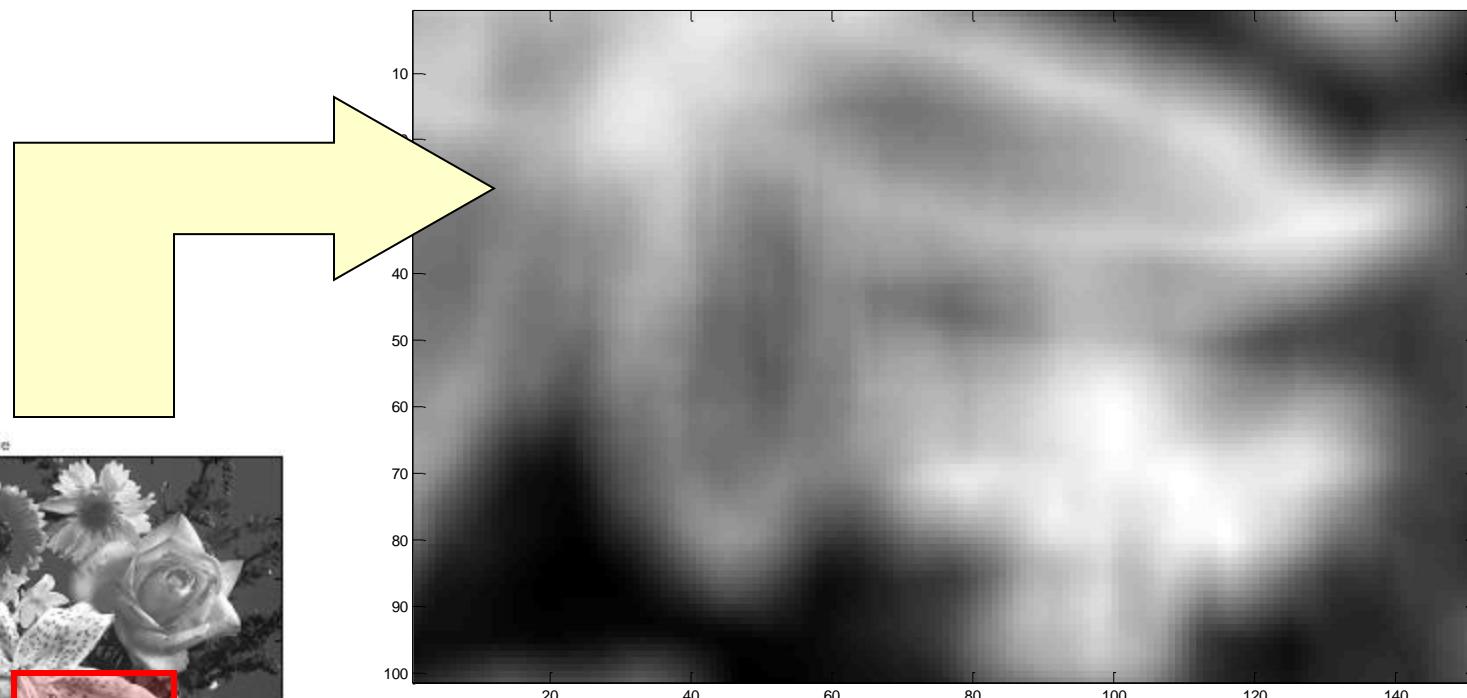
$$\int g(x)dx = 1$$



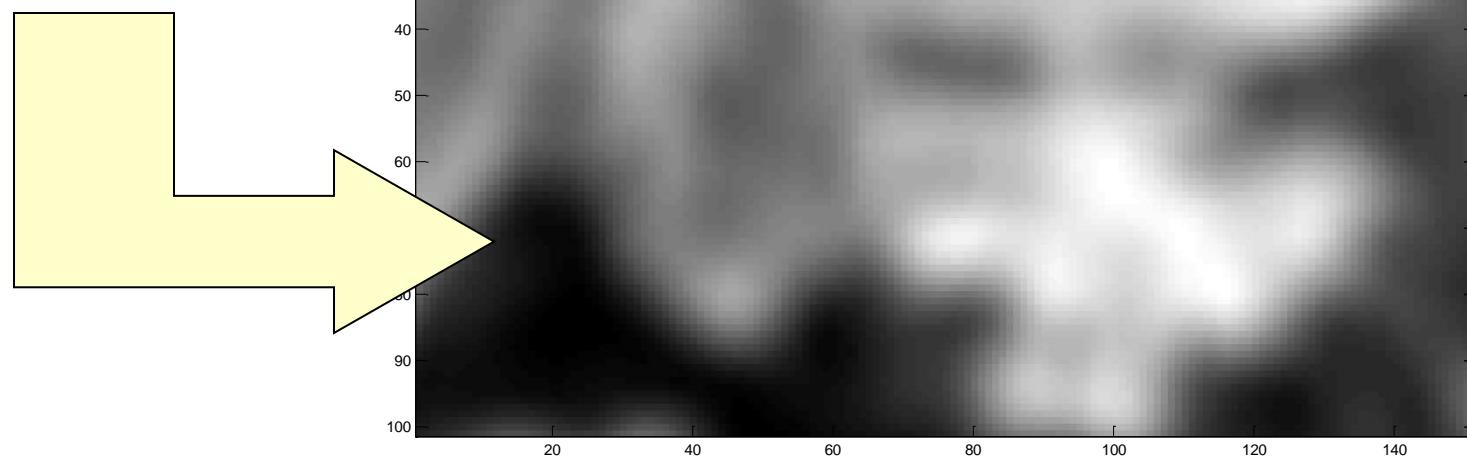
Gaussian Blurring,  $\sigma = 5$



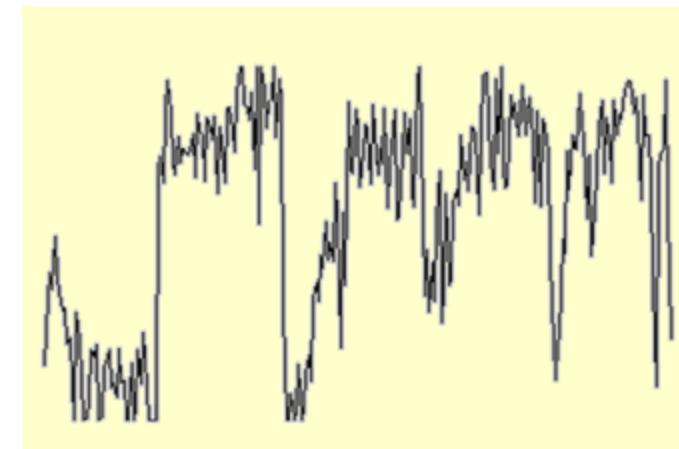
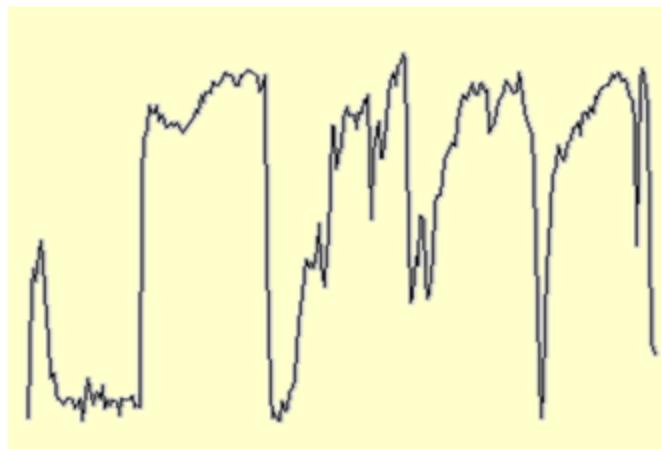
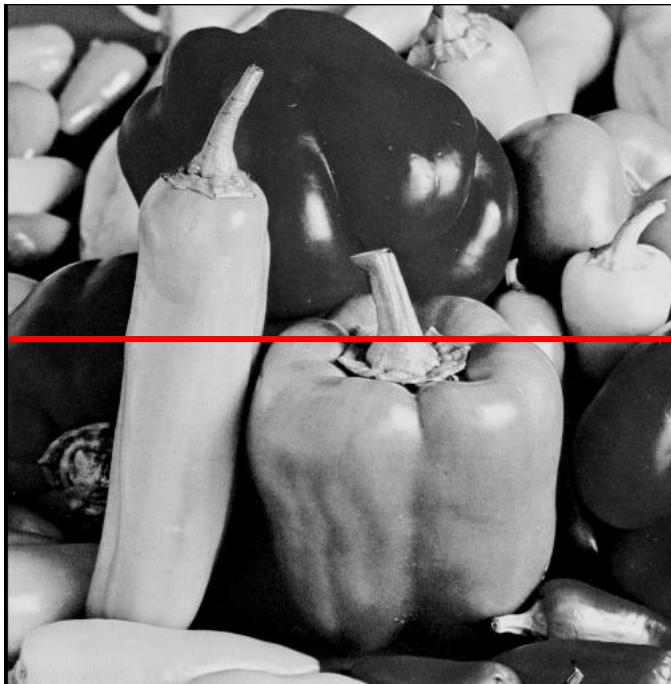
## Simple Averaging



## Gaussian Smoothing



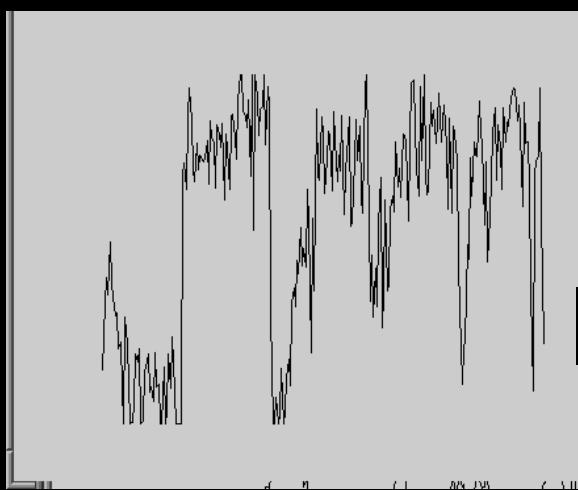
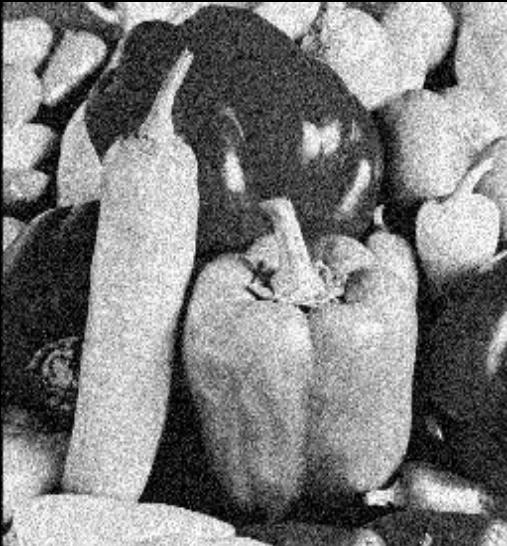
# Image Noise



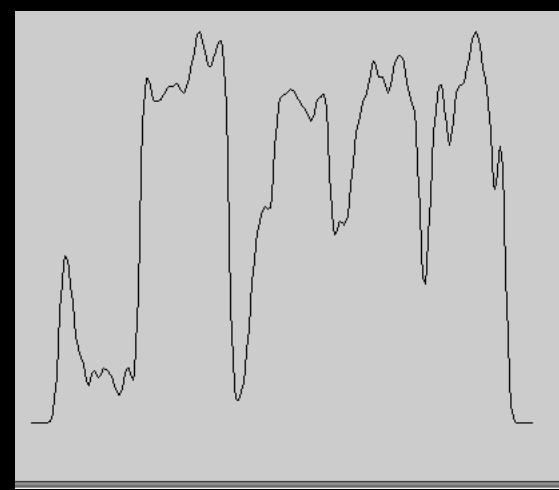
$$f(x, y) = \overbrace{\hat{f}(x, y)}^{\text{Ideal Image}} + \overbrace{\eta(x, y)}^{\text{Noise process}}$$

IID Gaussian white noise  
 $\eta(x, y) \sim N(0, \sigma)$

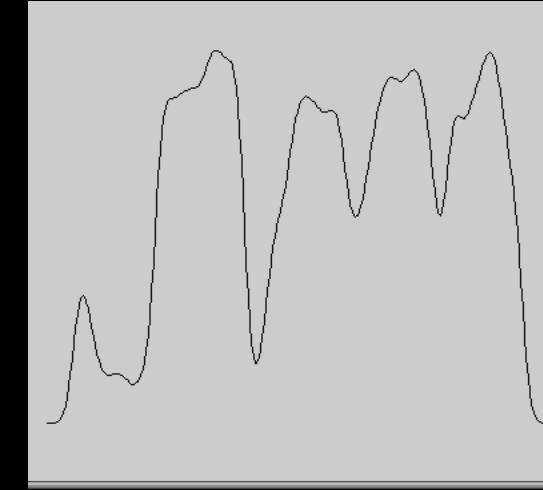
## Gaussian Smoothing to Remove Noise



No smoothing



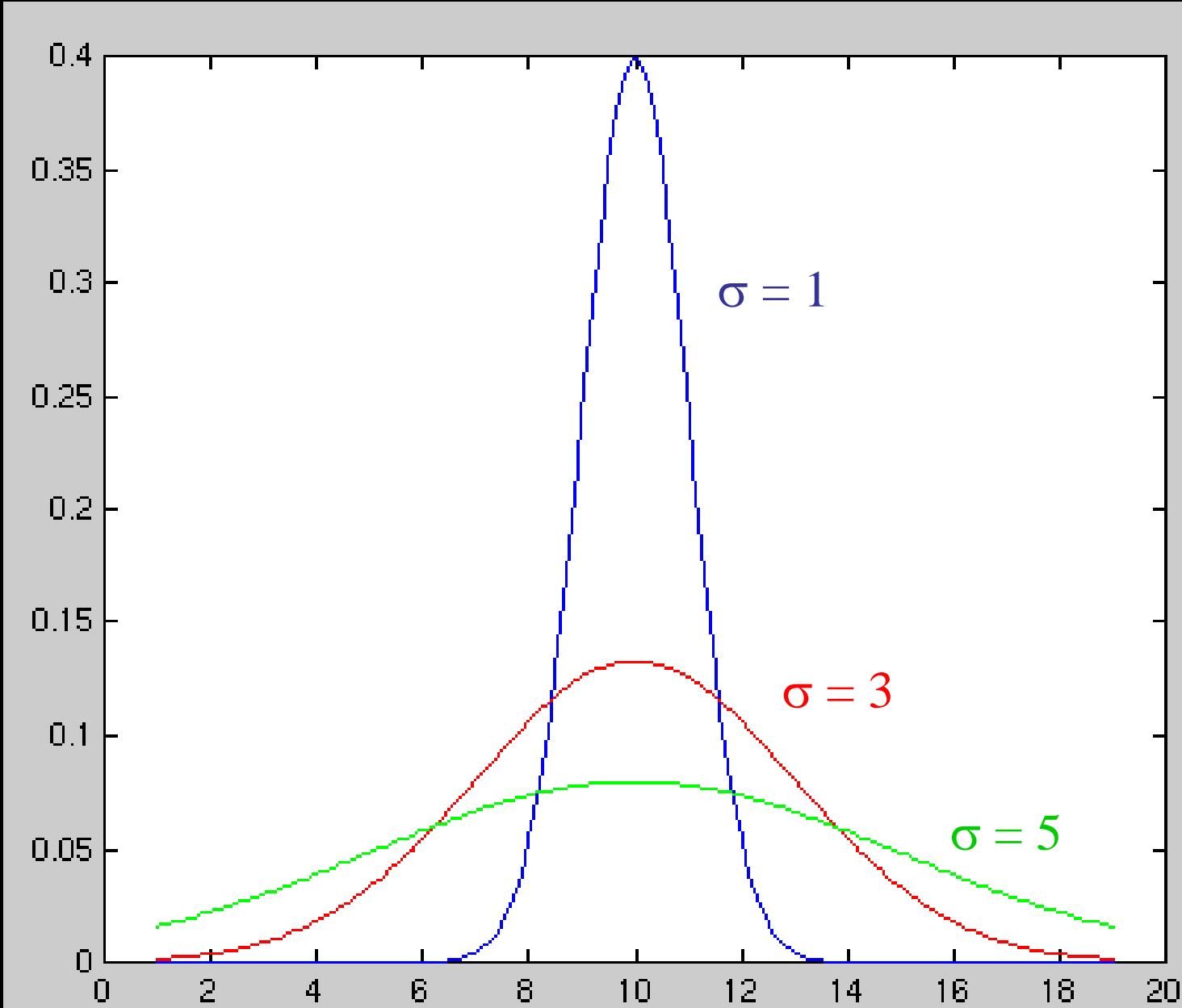
$\sigma = 2$



$\sigma = 4$

Bottom line: The standard deviation of white noise is divided by  $k^*\sigma$

# Shape of Gaussian filter as function of $\sigma$



# Basic Properties

- Gaussian removes “high-frequency” components from the image  
→ “low pass” filter
- Larger  $\sigma$  remove more details
- Combination of 2 Gaussian filters is a Gaussian filter:

$$G_{\sigma_1} * G_{\sigma_2} = G_{\sigma} \quad \sigma^2 = \sigma_1^2 + \sigma_2^2$$

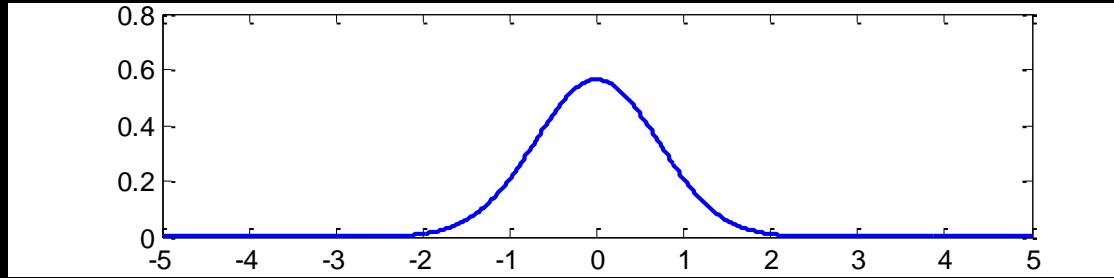
- Separable filter:

$$G_{\sigma} * f = g_{\sigma \rightarrow} * g_{\sigma \uparrow} * f$$

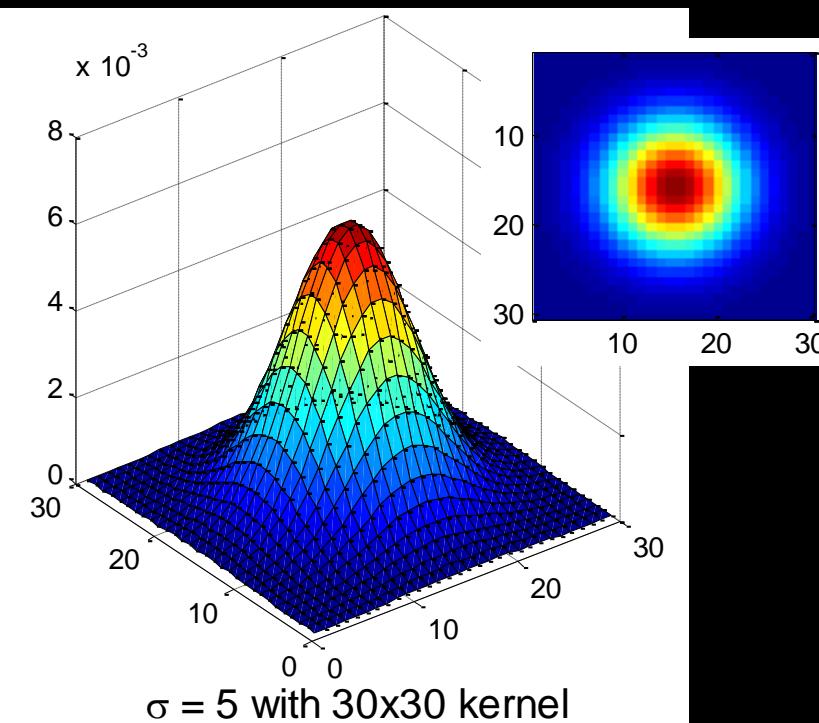
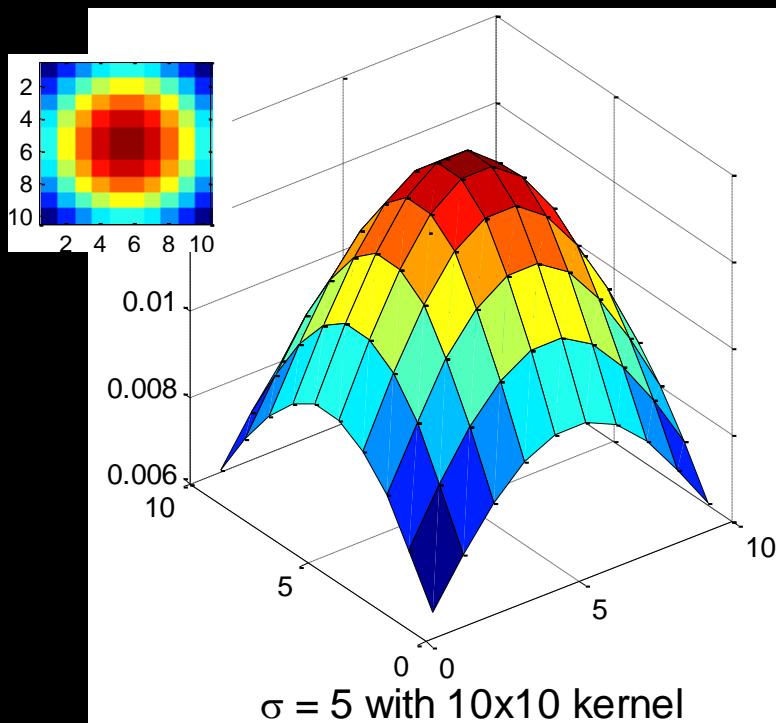
- Critical implication: Filtering with a  $N \times N$  Gaussian kernel can be implemented as two convolutions of size  $N \rightarrow$  reduction quadratic to linear → must be implemented that way

# Note about Finite Kernel Support

- Gaussian function has infinite support



- In actual filtering, we have a finite kernel size



# Image Derivatives

- We want to compute, at each pixel  $(x,y)$  the derivatives:
- In the discrete case we could take the difference between the left and right pixels:

$$\frac{\partial I}{\partial x} \approx I(i+1, j) - I(i-1, j)$$

- Convolution of the image by

$$\partial_x = \begin{bmatrix} 1 & 0 & -1 \end{bmatrix}$$

- Problem: Increases noise

$$I(i+1, j) - I(i-1, j) = \hat{I}(i+1, j) - \hat{I}(i-1, j) + n_+ + n_-$$

Difference between  
Actual image values

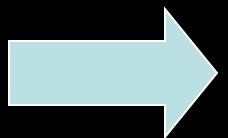
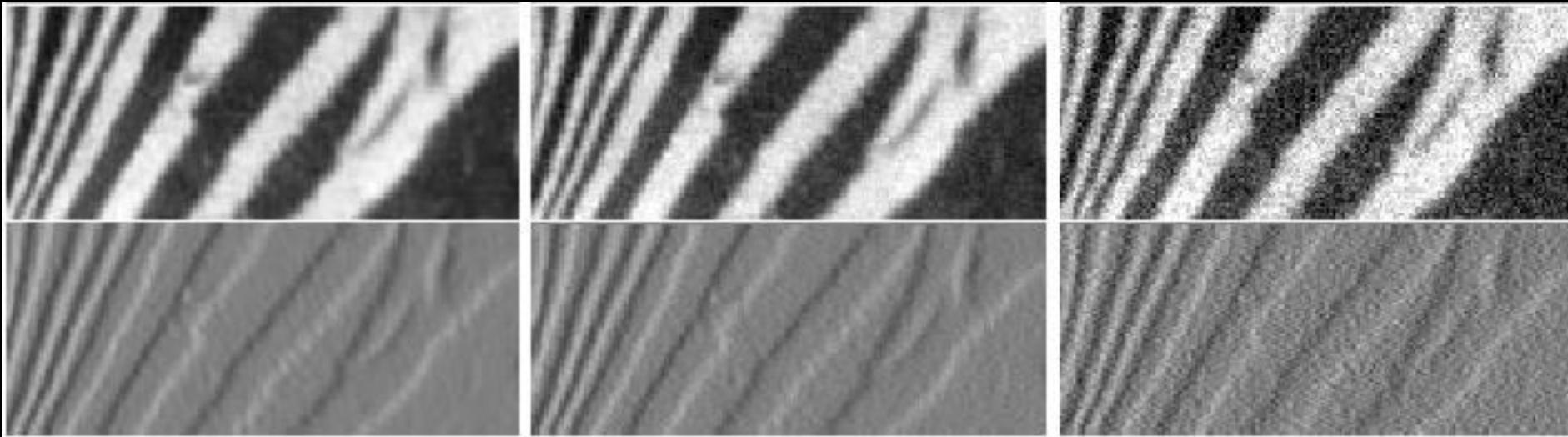
True difference  
(derivative)

Sum of the noises

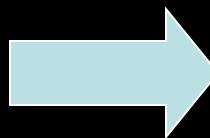
# Finite differences



# *Finite differences responding to noise*



Increasing zero-mean Gaussian noise



# Smooth Derivatives

- Solution: First smooth the image by a Gaussian  $G_\sigma$  and then take derivatives:

$$\frac{\partial f}{\partial x} \approx \frac{\partial(G_\sigma * f)}{\partial x}$$

- Applying the differentiation property of the convolution:

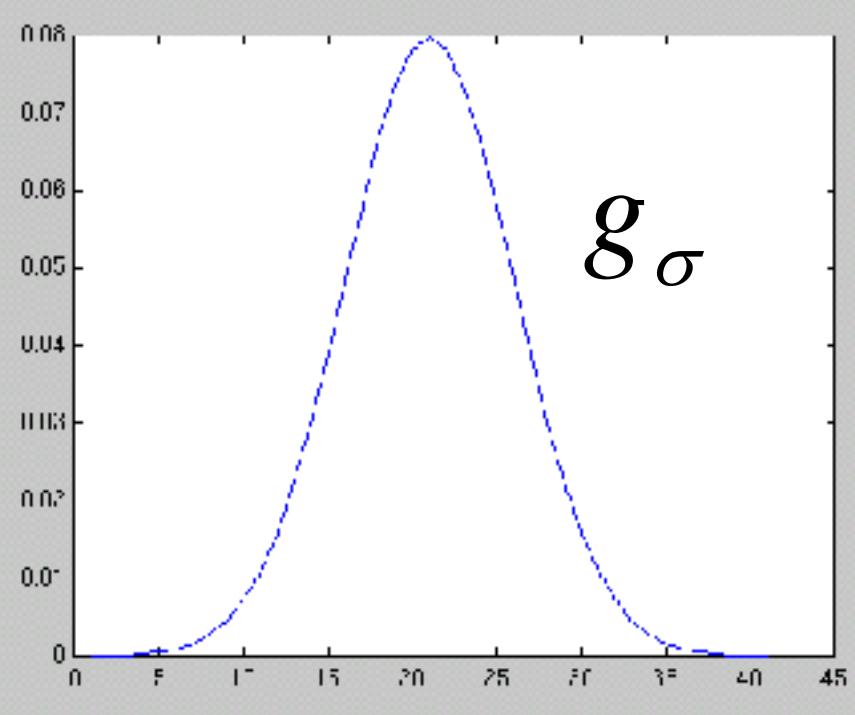
$$\frac{\partial f}{\partial x} \approx \frac{\partial G_\sigma}{\partial x} * f$$

- Therefore, taking the derivative in  $x$  of the image can be done by convolution with the derivative of a Gaussian:

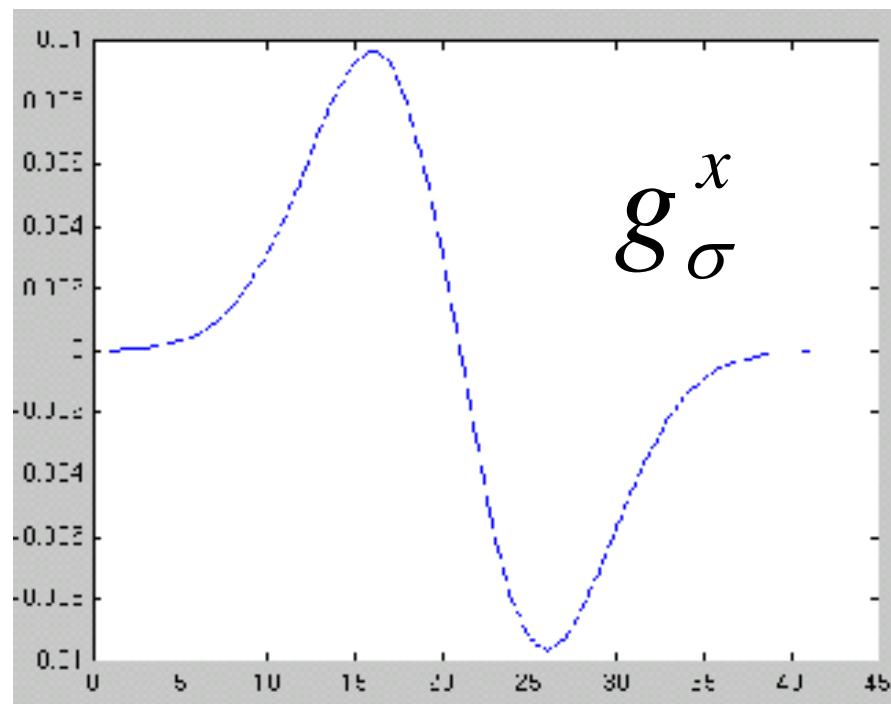
$$G_\sigma^x = \frac{\partial G_\sigma}{\partial x} = xe^{-\frac{x^2+y^2}{2\sigma^2}}$$

- Crucial property: The Gaussian derivative is also separable:

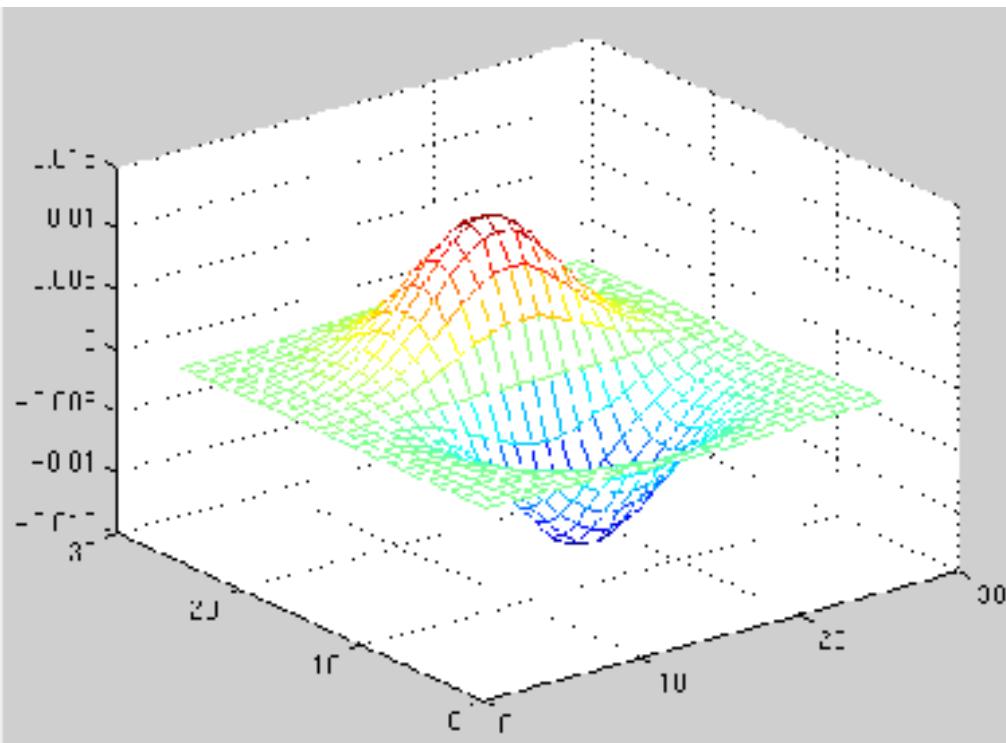
$$G_\sigma^x * f = g_\sigma^x * g_{\sigma \uparrow} * f$$



$g_\sigma$



$g_\sigma^x$

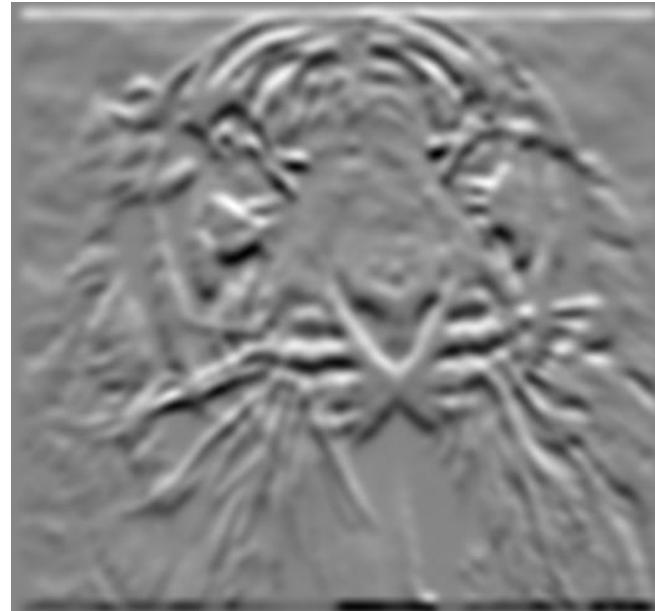


# Applying the first derivative of Gaussian

$I$



$\frac{\partial I}{\partial x}$



$$|\nabla I| = \sqrt{\frac{\partial I^2}{\partial x} + \frac{\partial I^2}{\partial y}}$$

$\frac{\partial I}{\partial y}$

There is **ALWAYS** a tradeoff between smoothing and good edge localization!

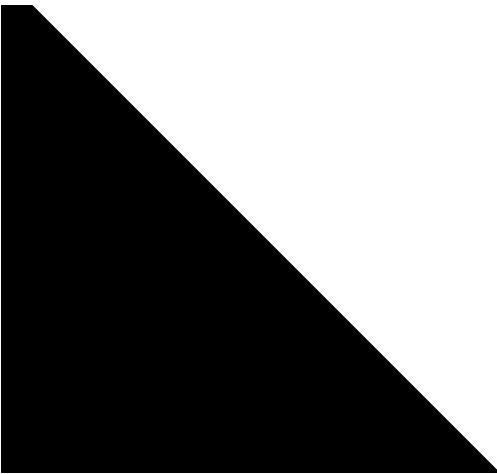
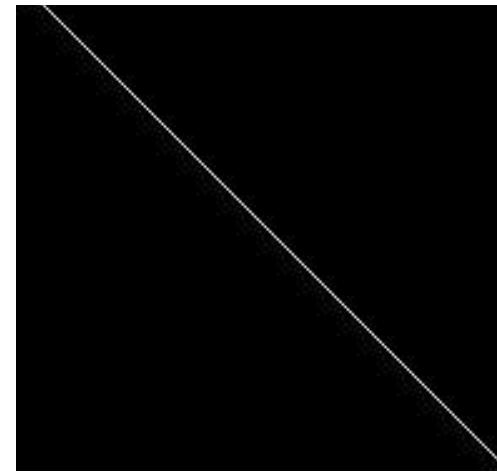


Image with Edge



Edge Location

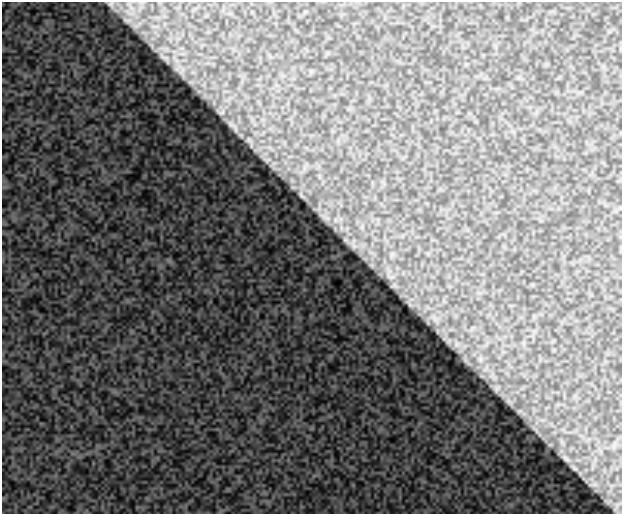
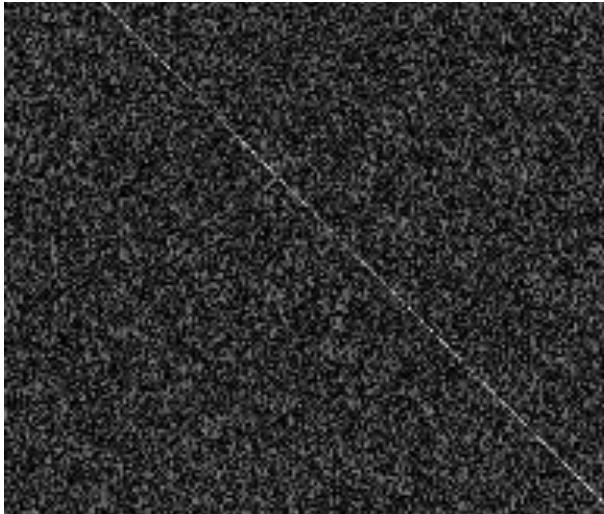
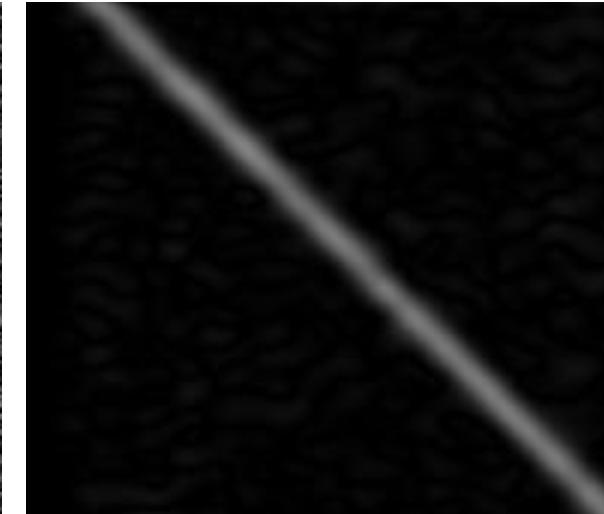


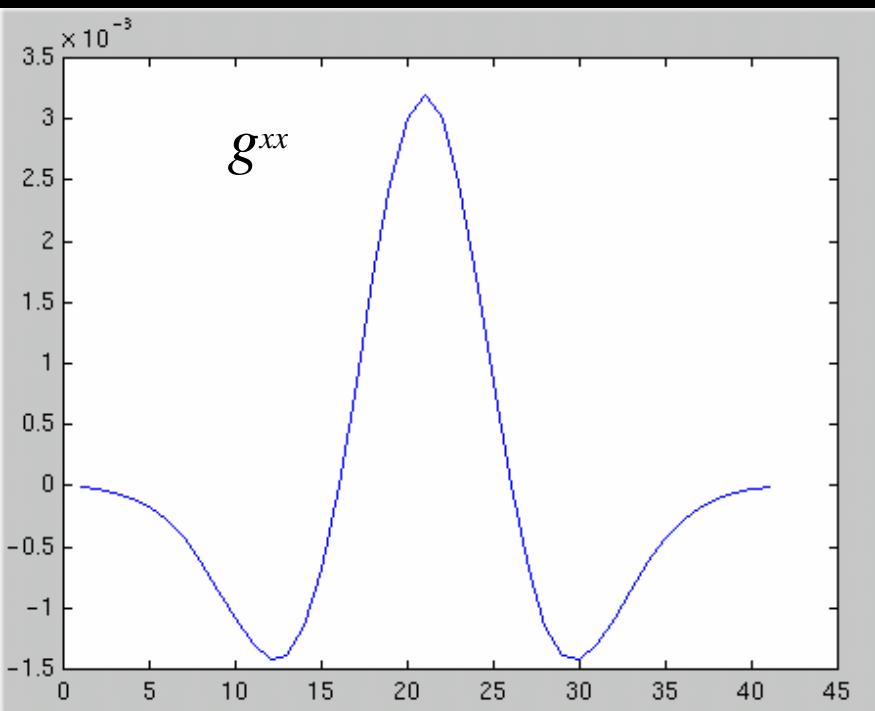
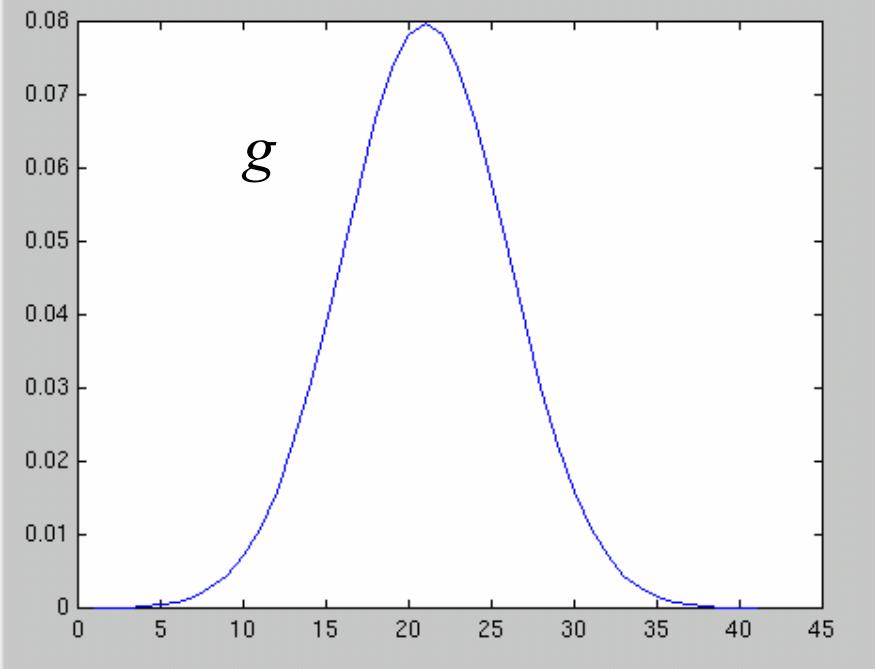
Image + Noise



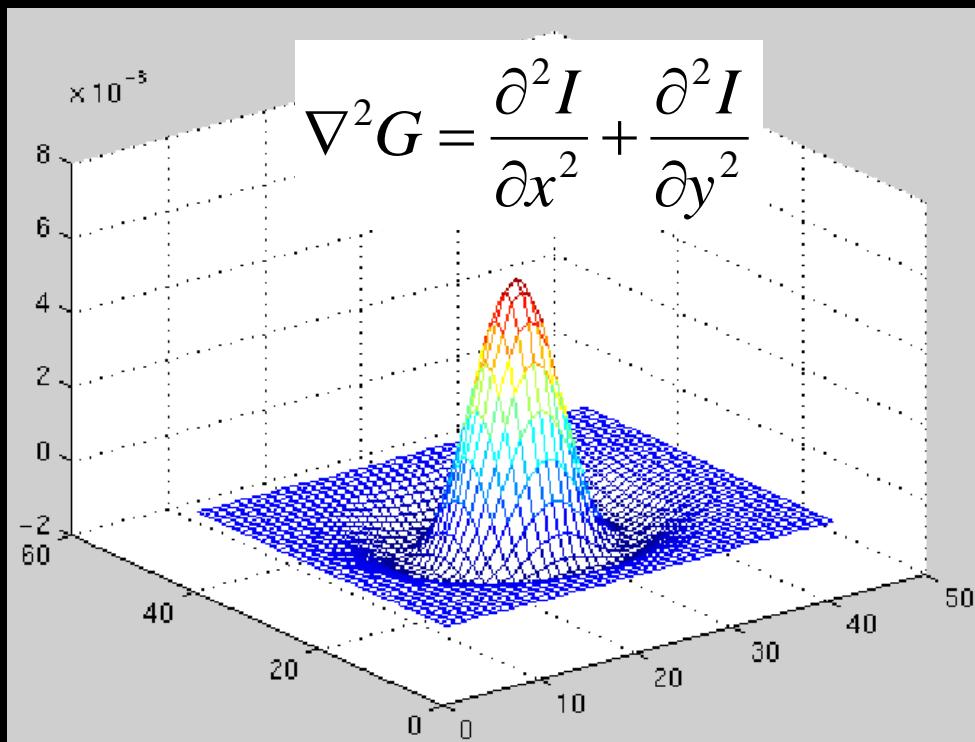
Derivatives detect  
edge *and* noise



Smoothed derivative removes  
noise, but blurs edge

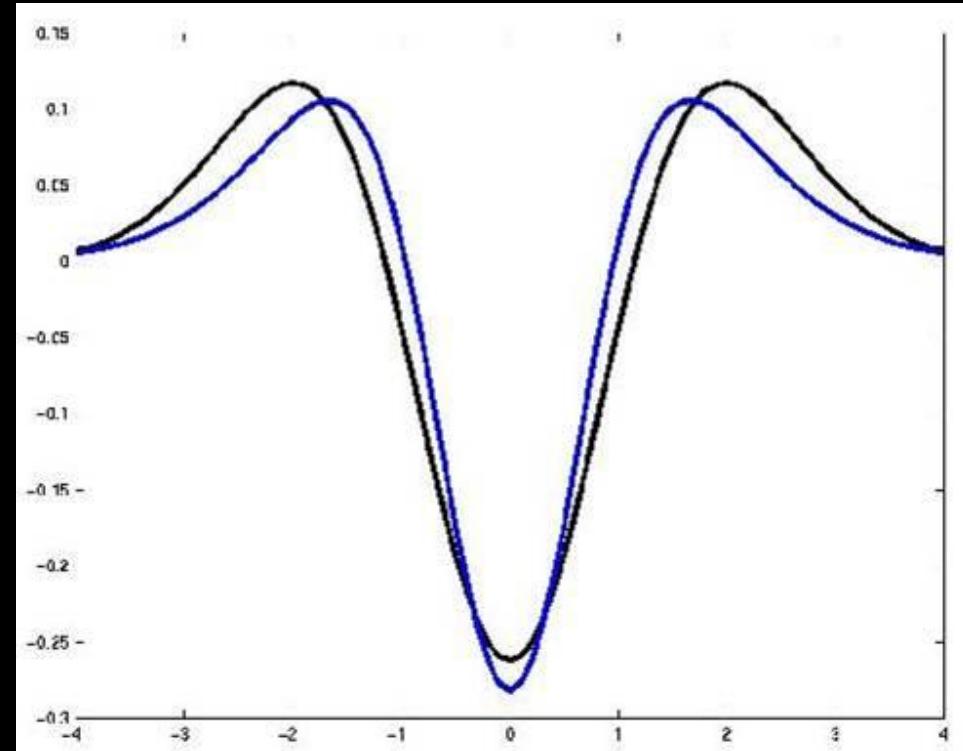
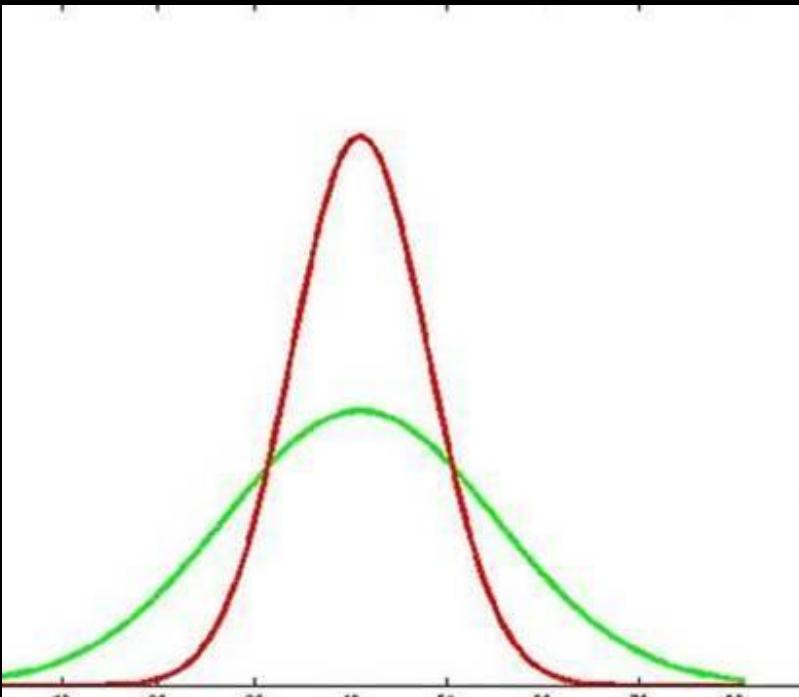


## Second derivatives: Laplacian



# DOG Approximation to LOG

$$\nabla^2 G_\sigma \approx G_{\sigma_1} - G_{\sigma_2}$$



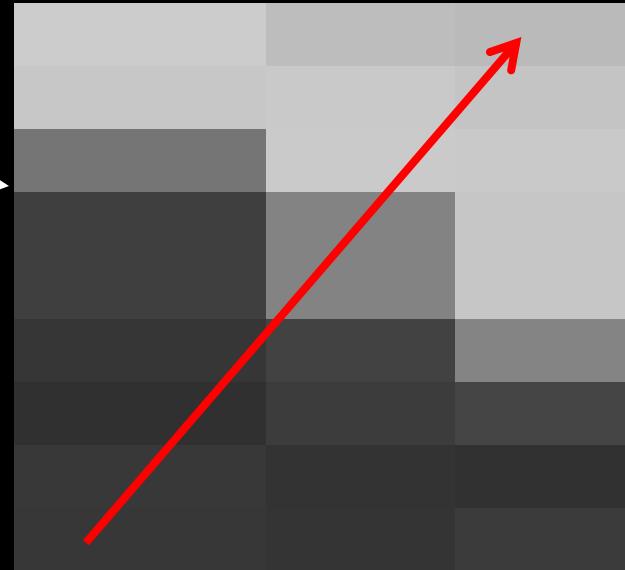
# Edge Detection

## Edge Detection

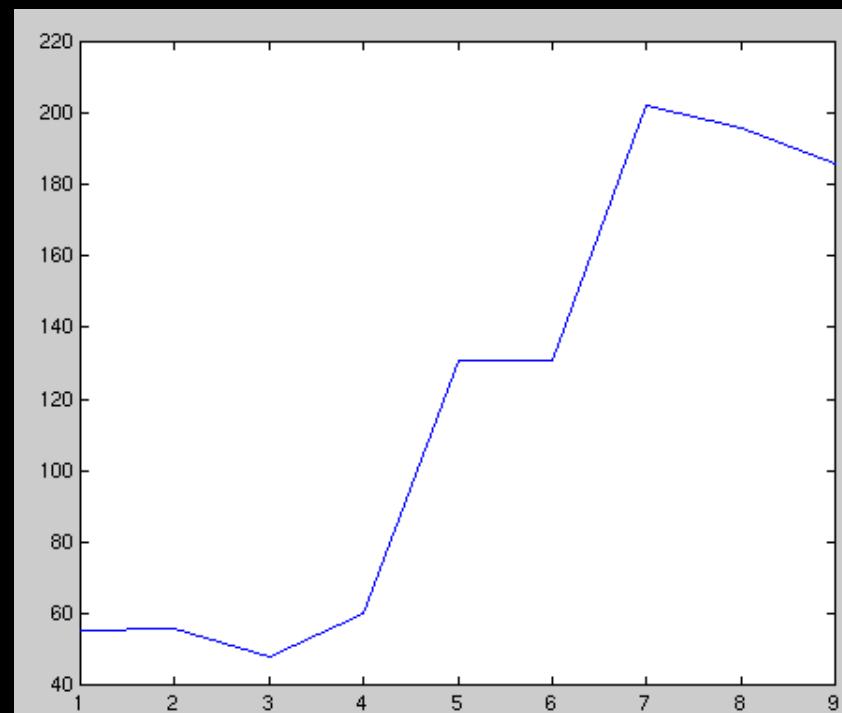
- Gradient operators
- Canny edge detectors
- Laplacian detectors



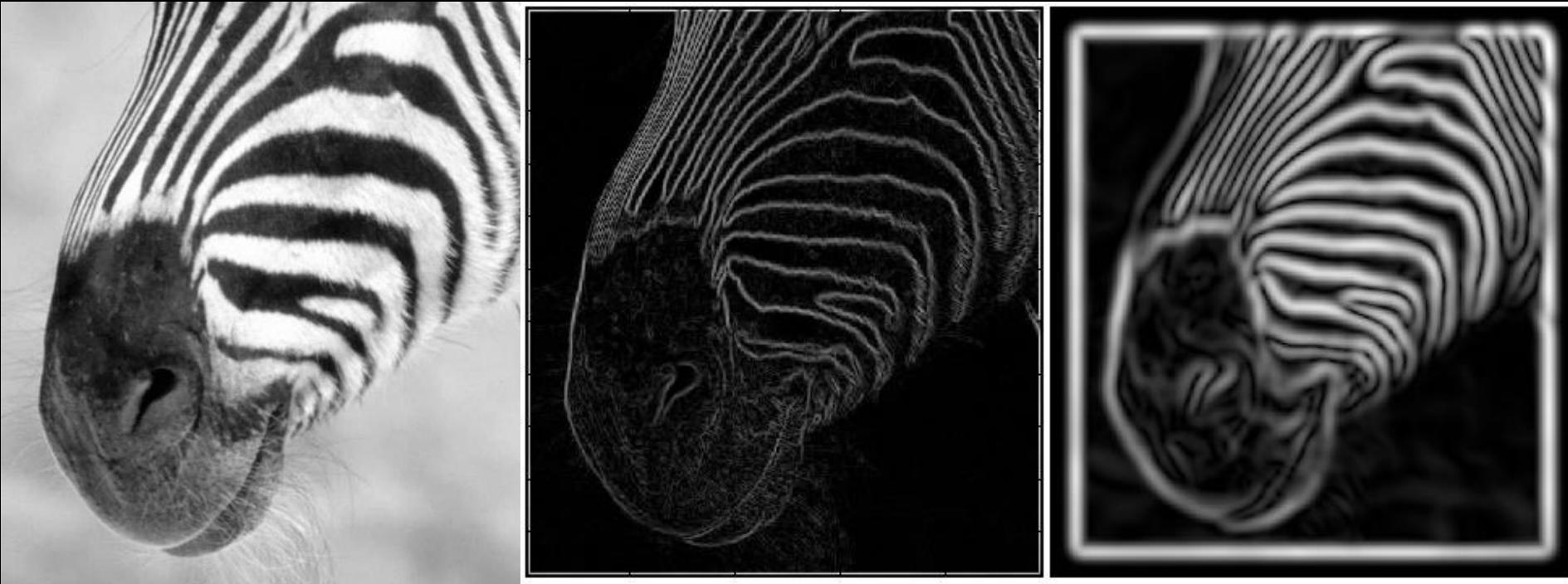
# What is an edge?



Edge = discontinuity of intensity  
in some direction.  
Could be detected by looking for  
places where the derivatives of  
the image have large values.



# Gradient-based edge detection

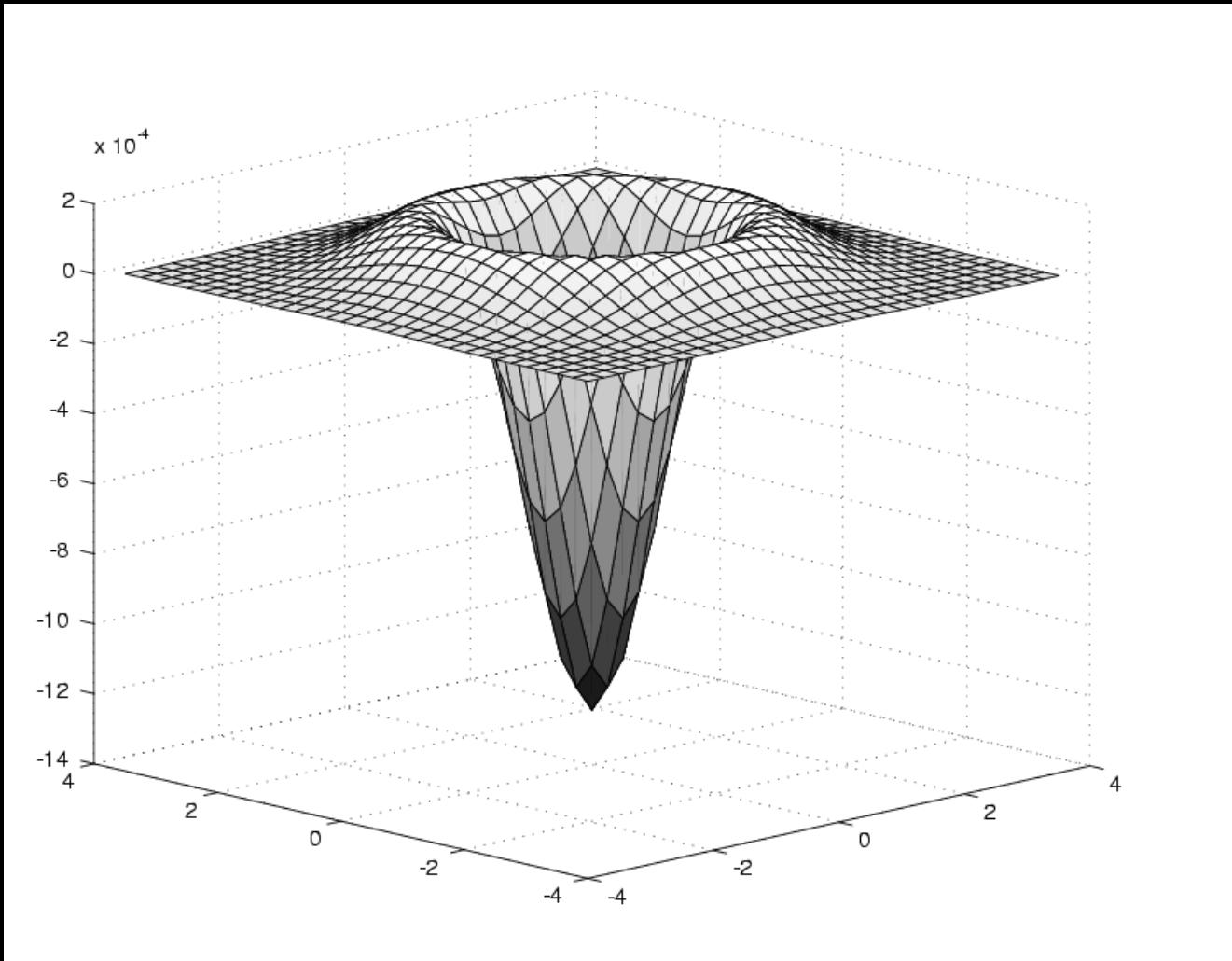


There are three major issues:

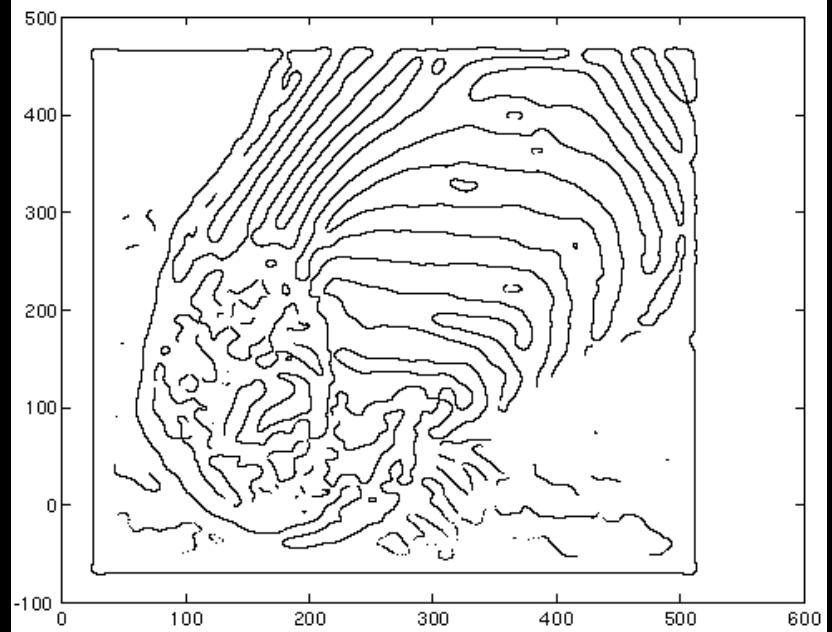
- 1) The gradient magnitudes at different scales are different;  
which one should we choose?
- 2) The gradient magnitude is large along thick trails; how  
do we identify the significant points?
- 3) How do we link the relevant points up into curves?

# The Laplacian of Gaussian (Marr-Hildreth 80)

- Another way to detect an extremal first derivative is to look for a zero second derivative.
- Appropriate 2D analogy is rotation invariant:
  - the Laplacian
$$r^2 f = \partial^2 f / \partial x^2 + \partial^2 f / \partial y^2$$
- Bad idea to apply a Laplacian without smoothing:
  - Smooth with Gaussian, apply Laplacian.
  - This is the same as filtering with a Laplacian of Gaussian filter.
- Now mark the zero points where there is a sufficiently large derivative, and enough contrast.

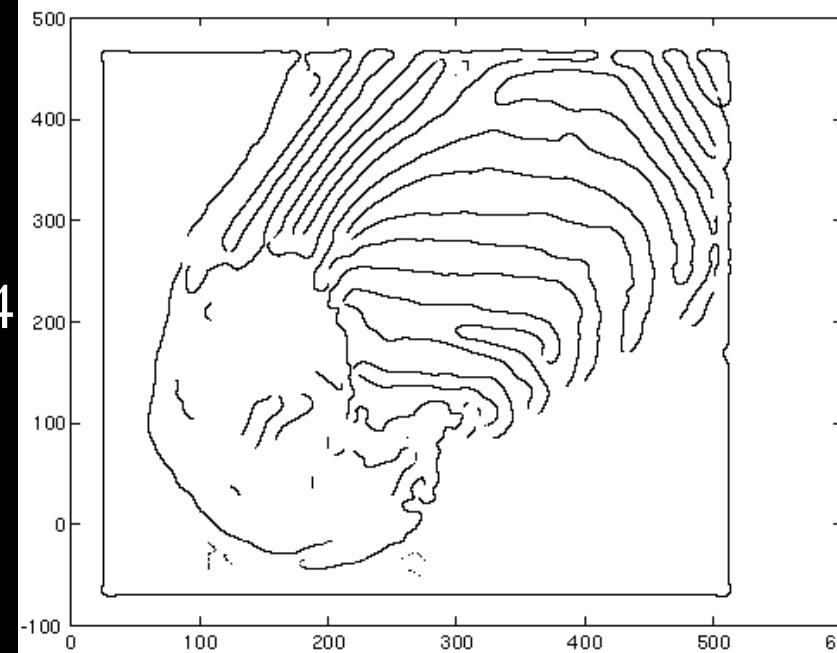


The Laplacian of a Gaussian



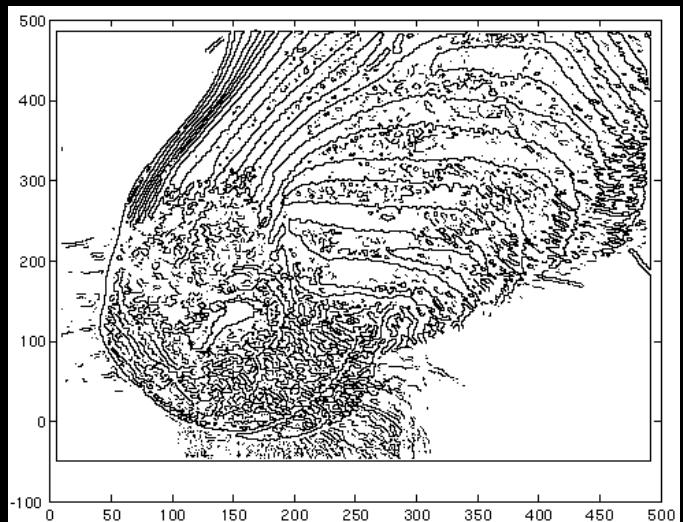
sigma=4

contrast=1

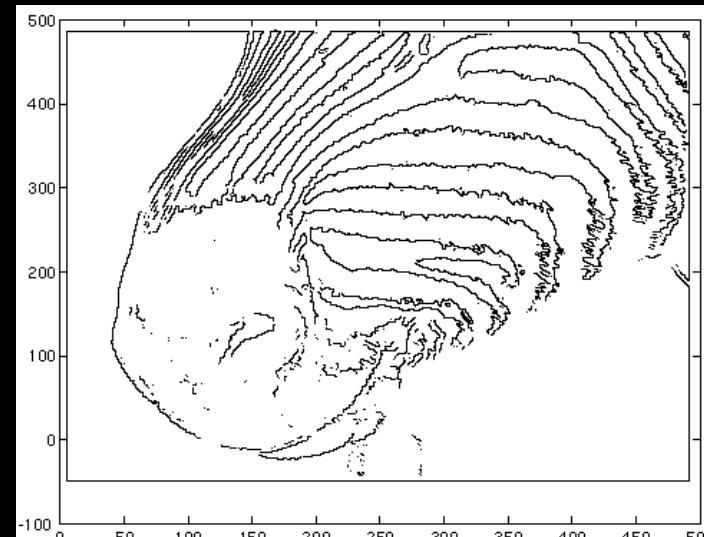


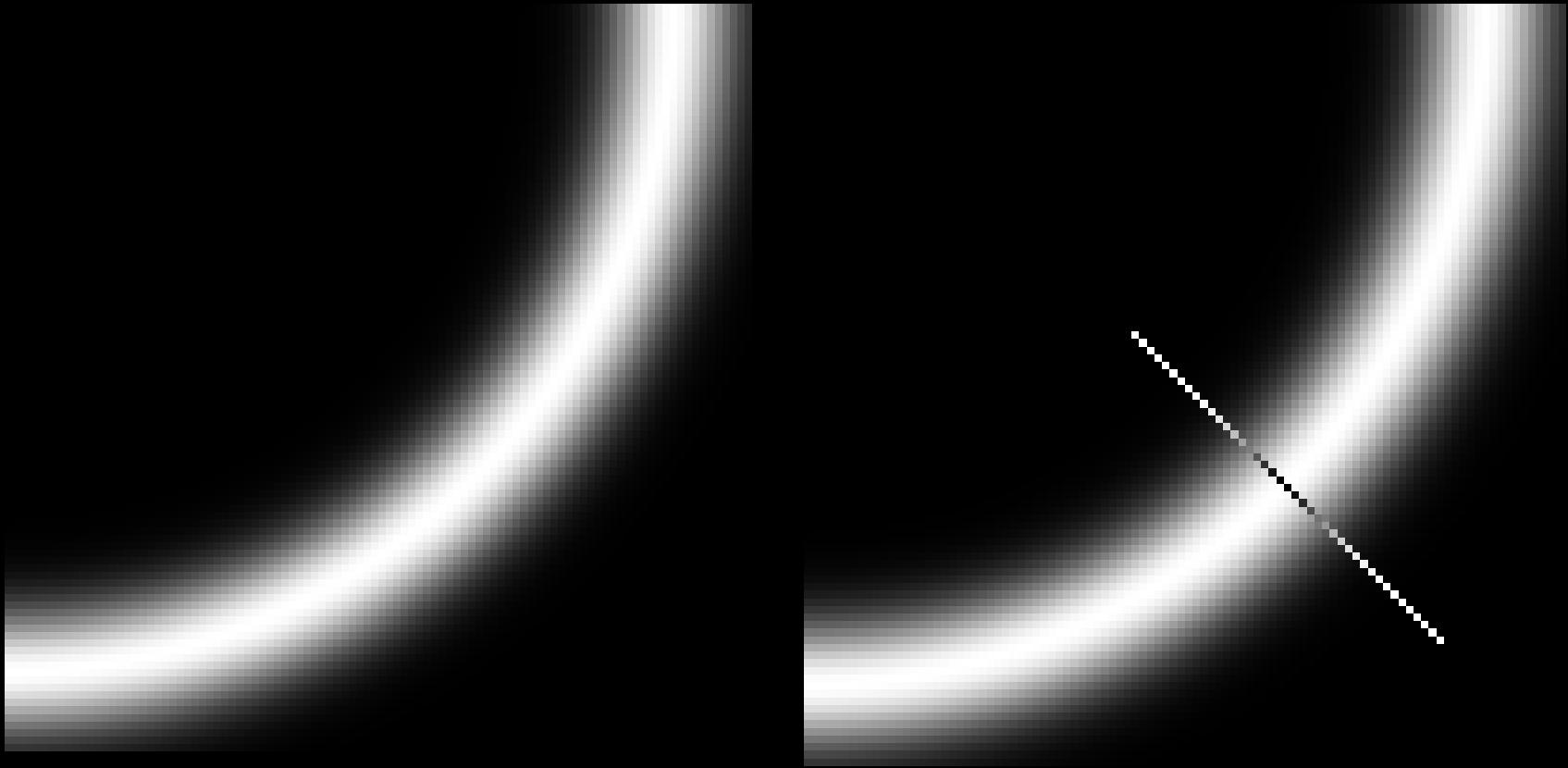
contrast=4

LOG zero crossings



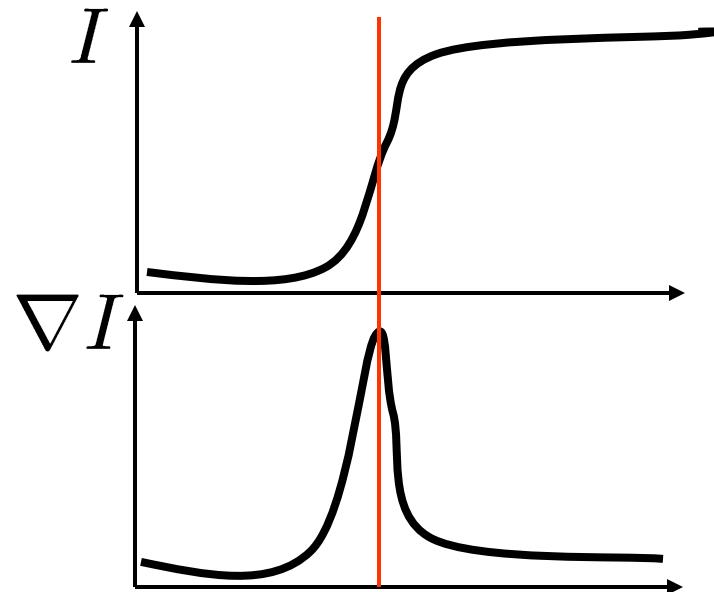
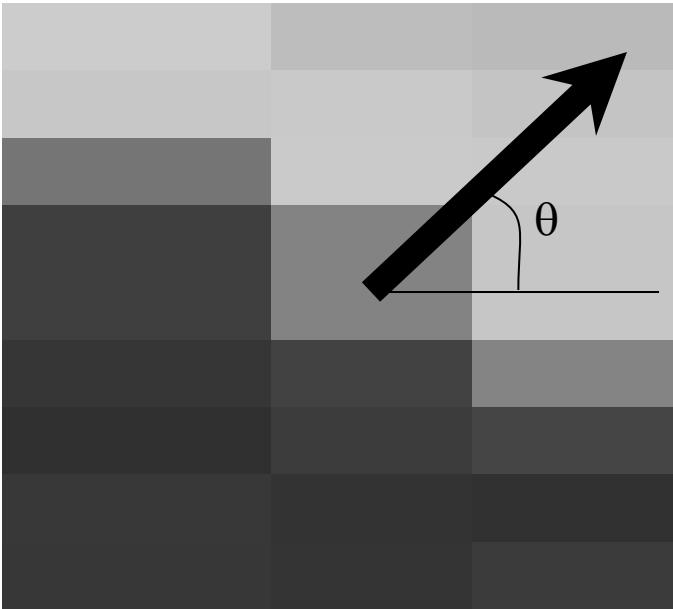
sigma=2





Gradient magnitude along an idealized curved edge.

Curved edges are locally straight: The gradient is orthogonal to the edge direction.

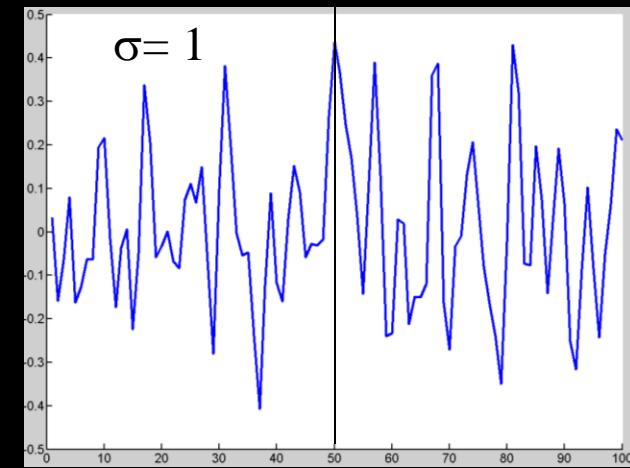
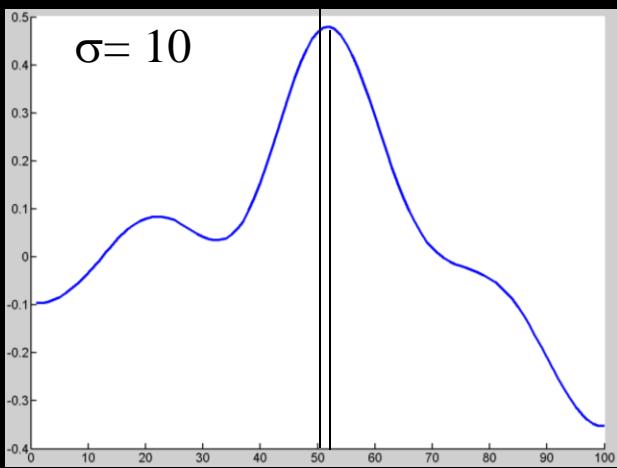
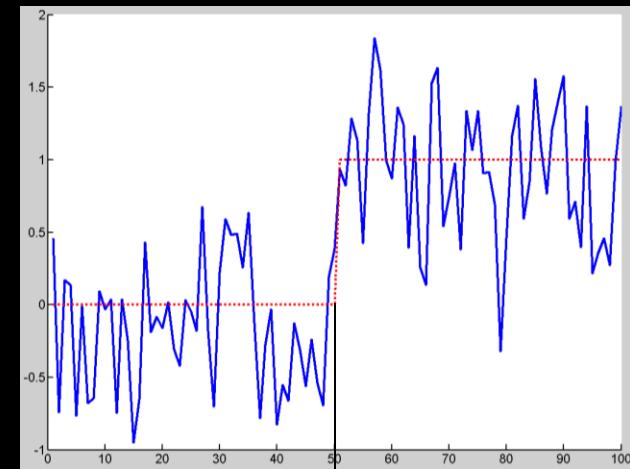
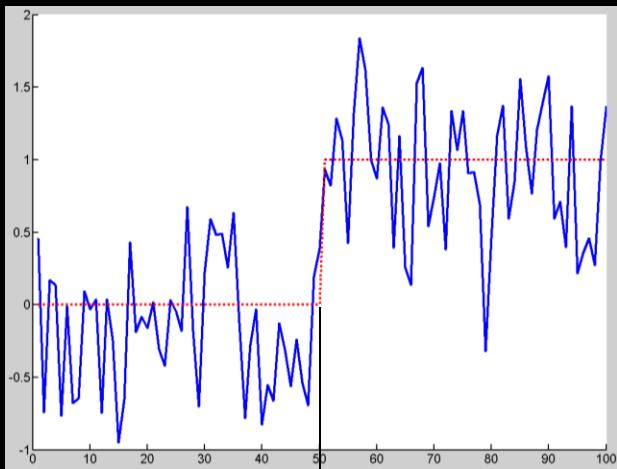


Edge pixels are at local maxima of gradient magnitude  
 Gradient computed by convolution with Gaussian derivatives  
 Gradient direction is always perpendicular to edge direction

$$\frac{\partial I}{\partial x} = G_{\sigma}^x * I$$

$$\frac{\partial I}{\partial y} = G_{\sigma}^y * I$$

$$|\nabla I| = \sqrt{\left(\frac{\partial I}{\partial x}\right)^2 + \left(\frac{\partial I}{\partial y}\right)^2} \quad \theta = \text{atan2}\left(\frac{\partial I}{\partial y}, \frac{\partial I}{\partial x}\right)$$

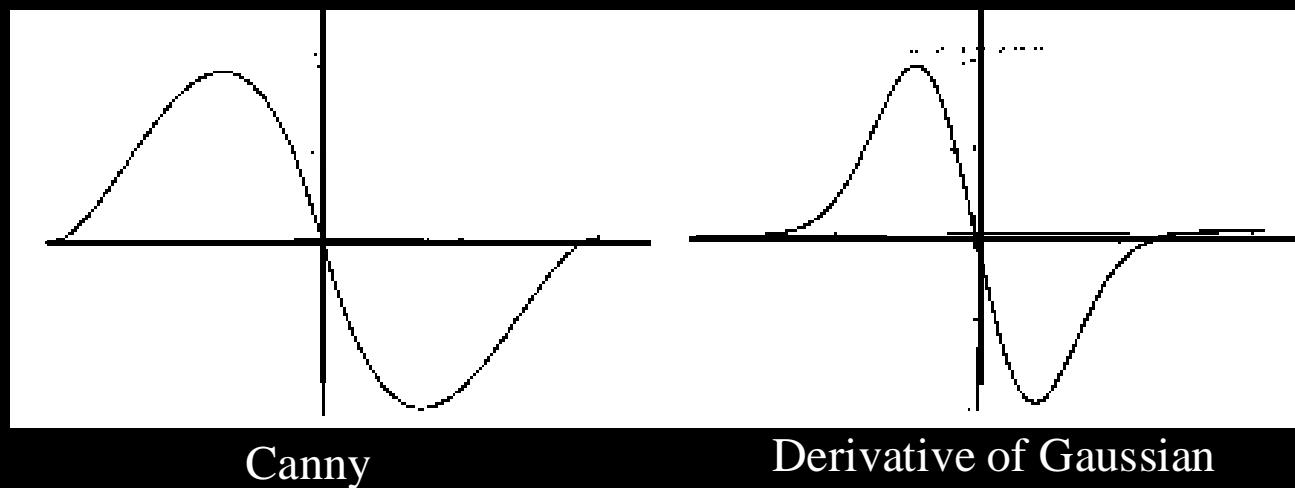


Large  $\sigma \rightarrow$  Good detection (high SNR)  
Poor localization

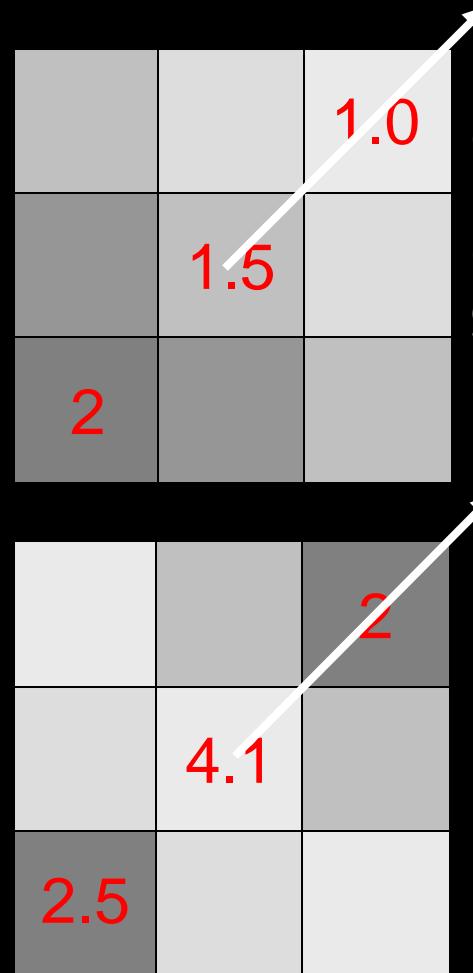
Small  $\sigma \rightarrow$  Poor detection (low SNR)  
Good localization

# Canny's Result

- Given a filter  $f$ , define the two objective functions:  
 $\Lambda(f)$  large if  $f$  produces good localization  
 $\Sigma(f)$  large if  $f$  produces good detection (high SNR)
- Problem: Find a family of filters  $f$  that maximizes the compromise criterion  
 $\Lambda(f)\Sigma(f)$   
under the constraint that a single peak is generated by a step edge
- Solution: Unique solution, a close approximation is the Gaussian derivative filter!



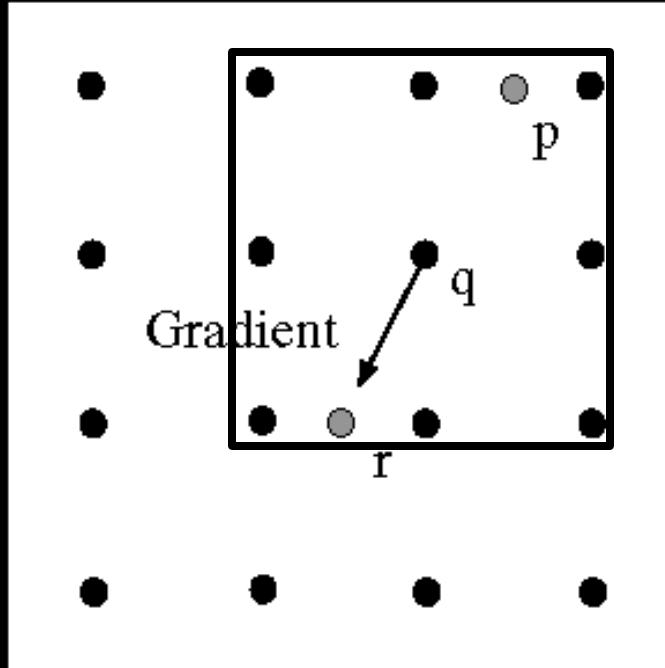
## Non-Local Maxima Suppression



$\nabla I$   
Gradient magnitude at center pixel  
is lower than the gradient magnitude  
of a neighbor in the direction of the  
gradient → Discard center pixel  
(set magnitude to 0)

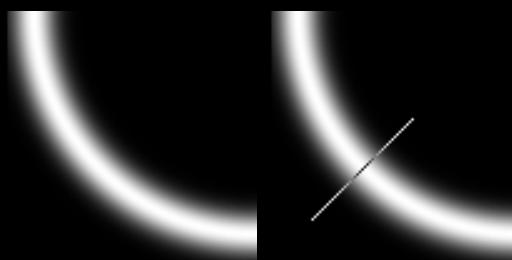


$\nabla I$   
Gradient magnitude at center pixel  
is greater than gradient magnitude  
of all the neighbors in the direction  
of the gradient  
→ Keep center pixel unchanged



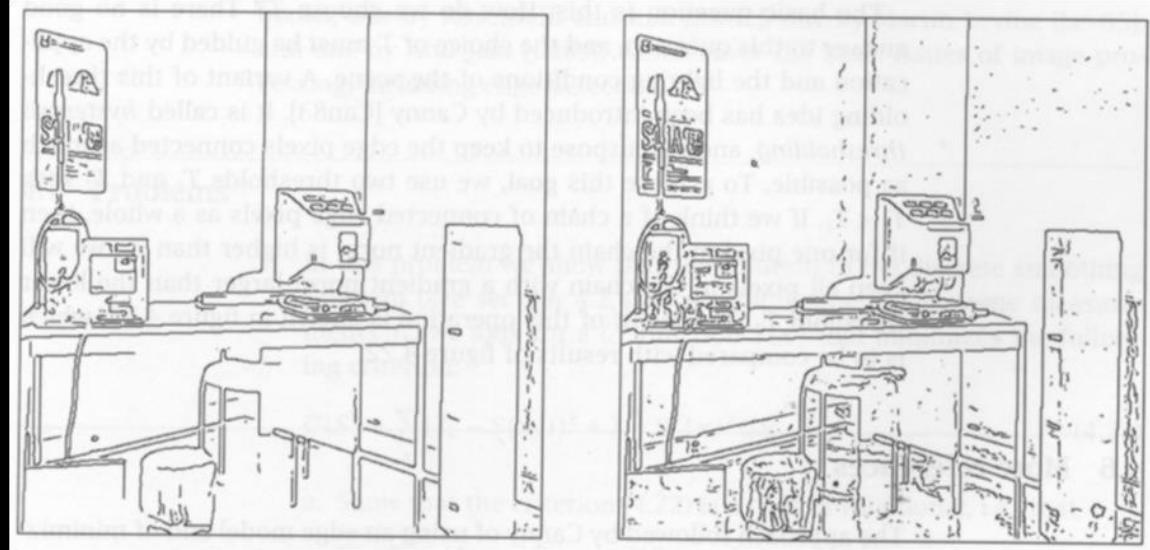
### Non-maximum suppression

At q we have a maximum if the value is larger than those at both p and at r.  
Interpolate to get these values.



Input image



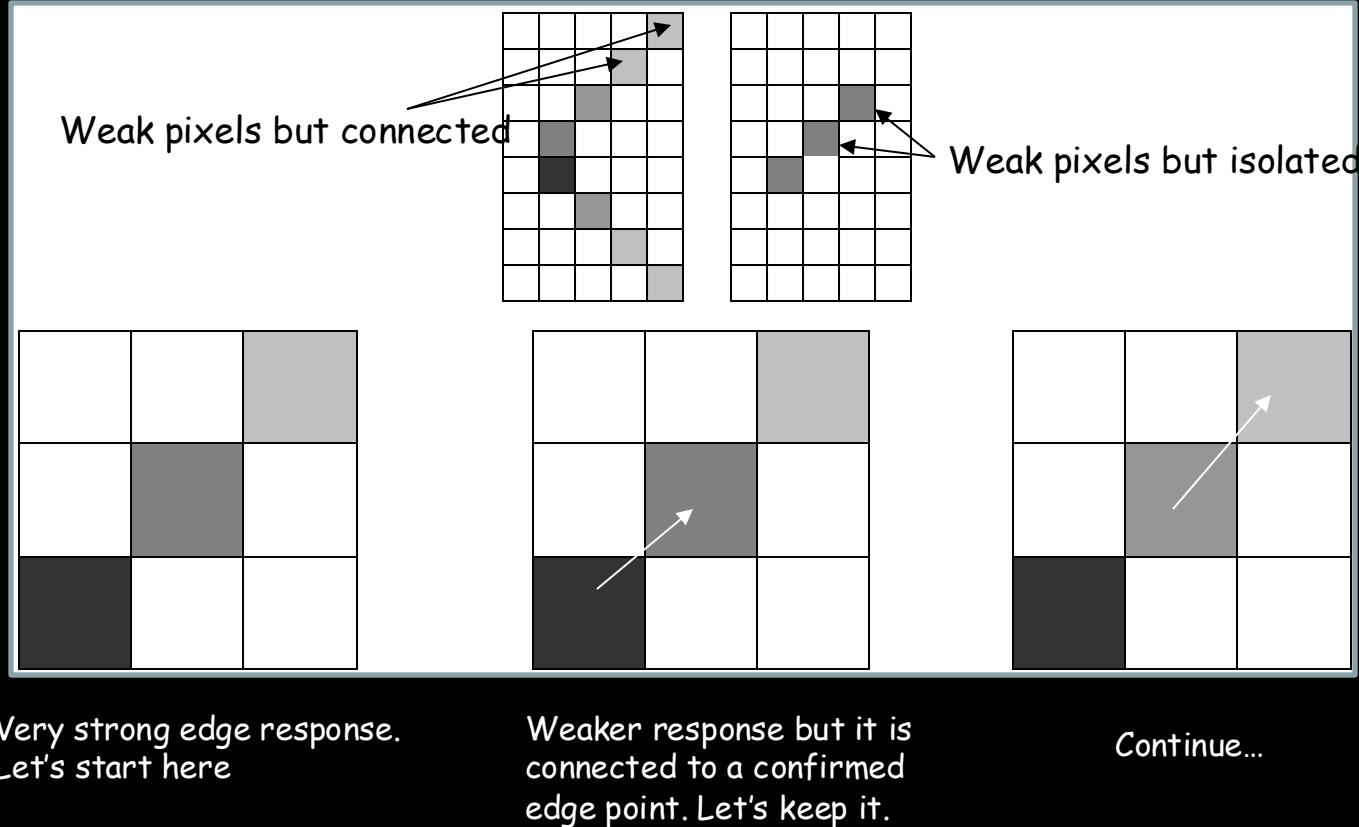


$T = 15$

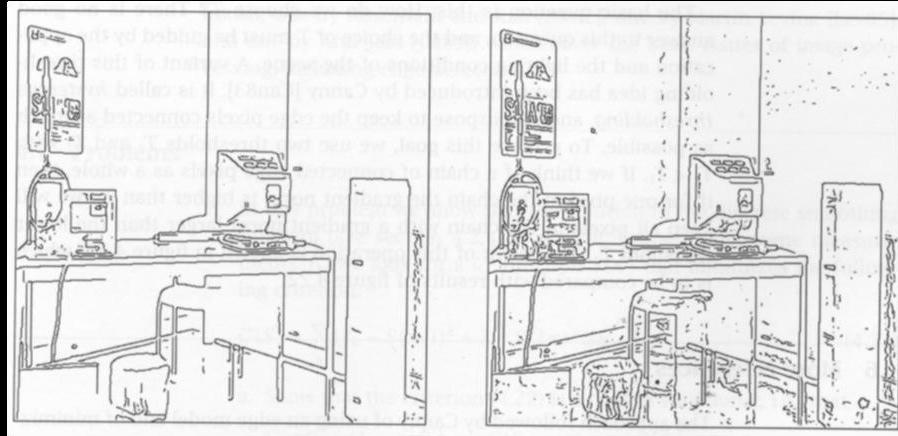
$T = 5$

Two thresholds applied to gradient magnitude

# Hysteresis Thresholding



$T=15$

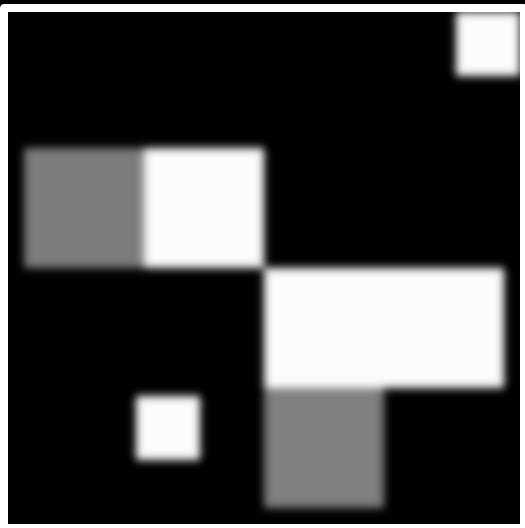


$T=5$

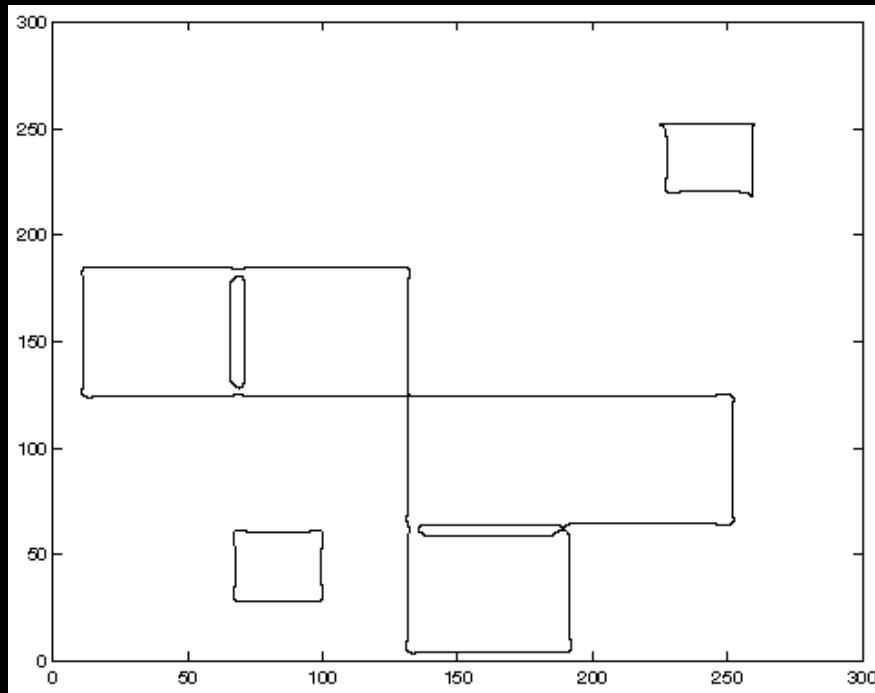
Hysteresis  
 $T_h=15 \ T_l = 5$

Hysteresis  
thresholding





We have unfortunate behaviour  
at corners



# Why machine learning for image restoration?

Reasonable physical models of image corruption

- For example:  $y = A(x) + \varepsilon$
- For example:  $A(x) = k * x$
- One can use prior knowledge
  - For example: sparsity, self similarities
- Realistic simulated training examples
- Interpretable, "functional" architectures

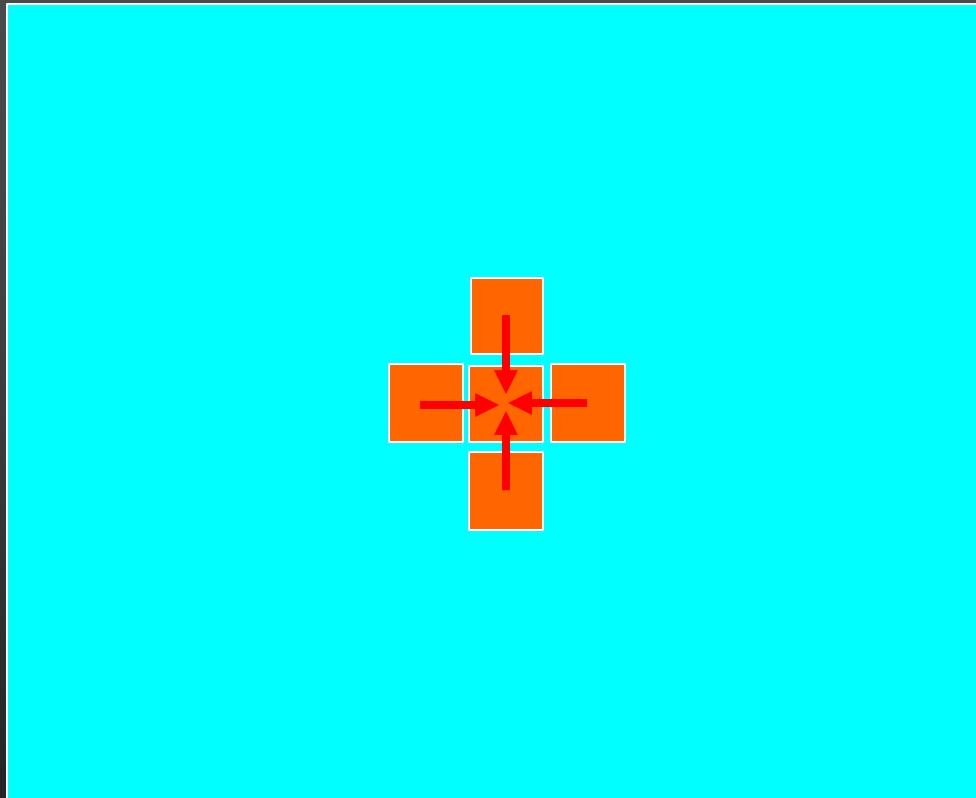
# Why machine learning for image restoration?

Reasonable physical models of image corruption

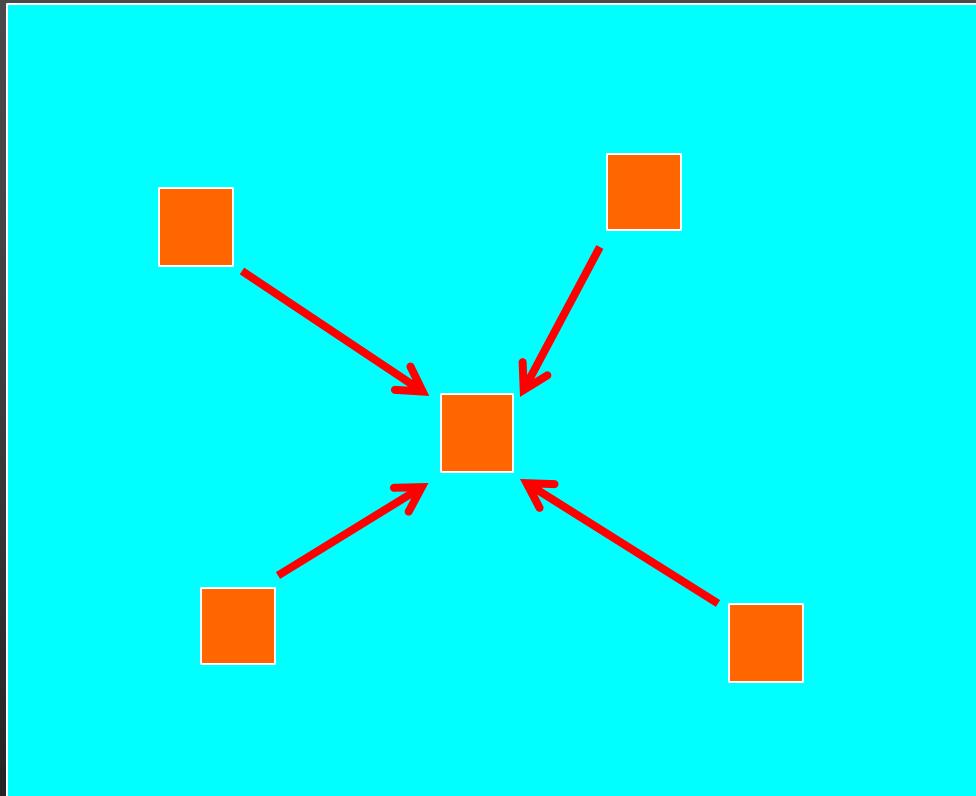
- For example:  $y = A(x) + \varepsilon$
- For example:  $A(x) = k * x$
- One can use prior knowledge
  - For example: sparsity, self similarities
- Realistic simulated training examples
- Interpretable, "functional" architectures

But where does the real ground truth come from, whether for model-based or data-driven methods?

# Let us start simple: How to denoise an image

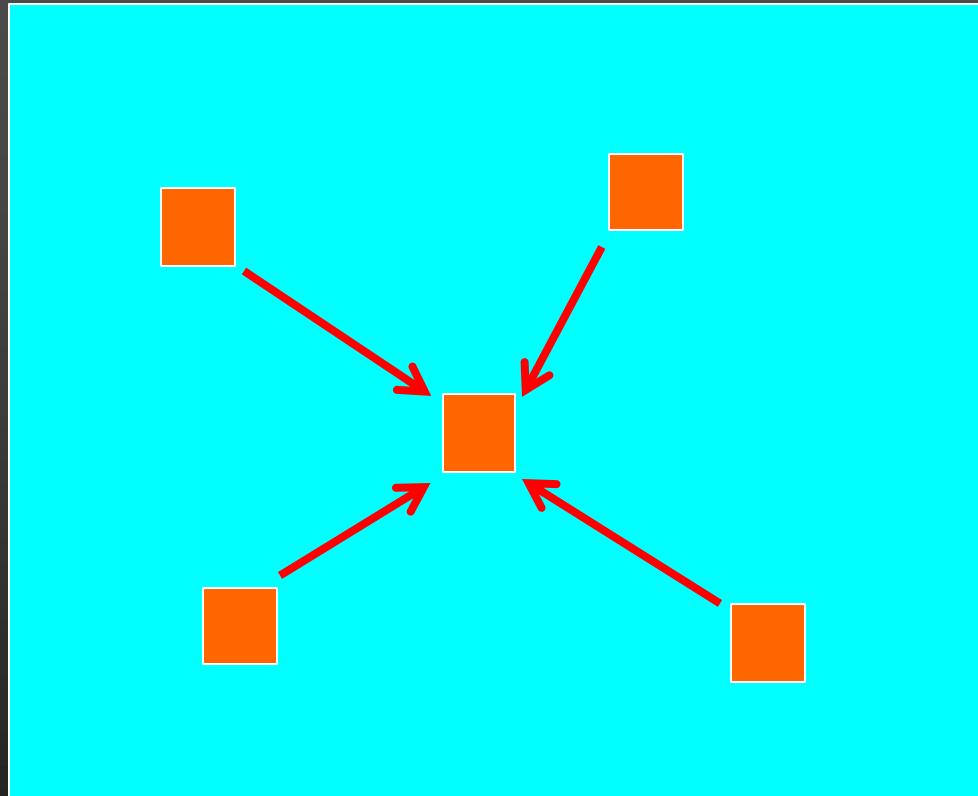


# Let us start simple: How to denoise an image

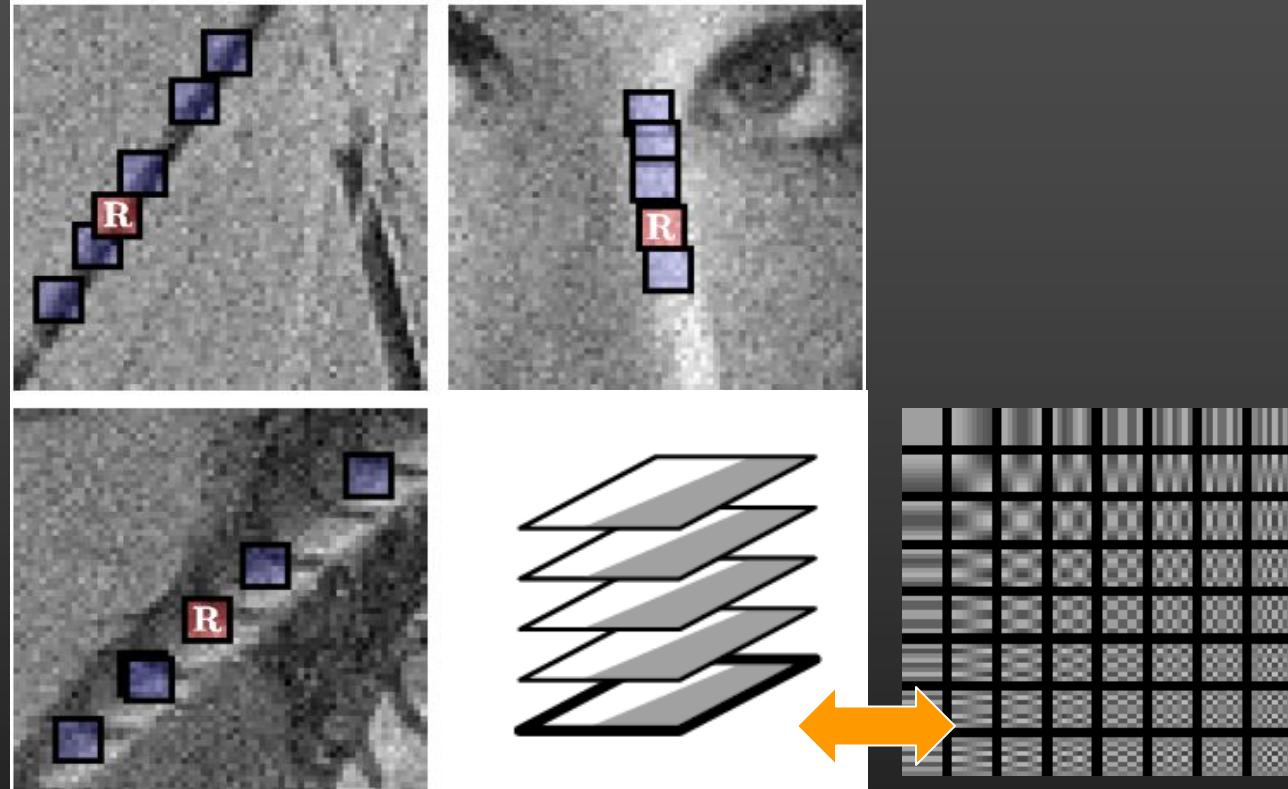


Non-local means filtering (Buades et al.'05)

# Let us start simple: How to denoise an image



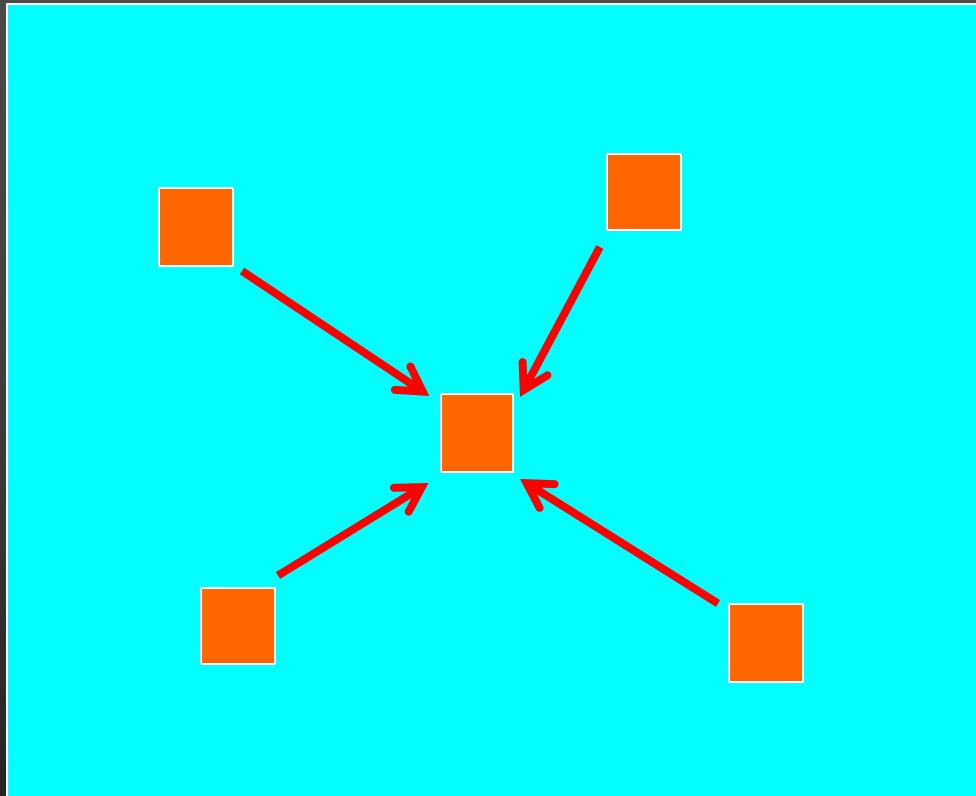
Non-local means filtering (Buades et al.'05)



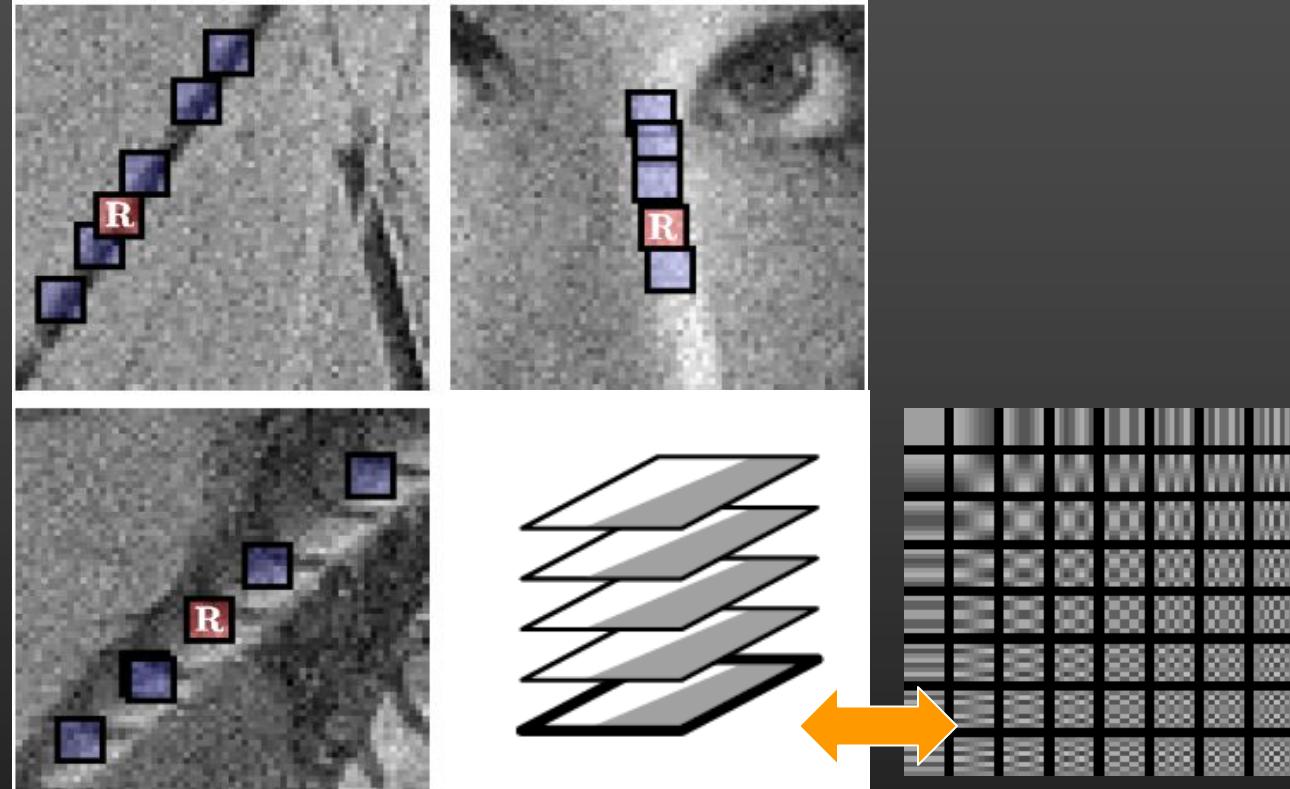
BM3D = **Sparse** representation on (3D) DCT dictionary + NLM (Dabov et al.'07)

Observation: natural image patches can be sparsely represented as linear combinations of a few elements of appropriate dictionaries, e.g., discrete cosine transform basis functions (e.g., Olshausen and Field, 1997; Chen et al., 1999; Mallat, 1999).

# Let us start simple: How to denoise an image



Non-local means filtering (Buades et al.'05)

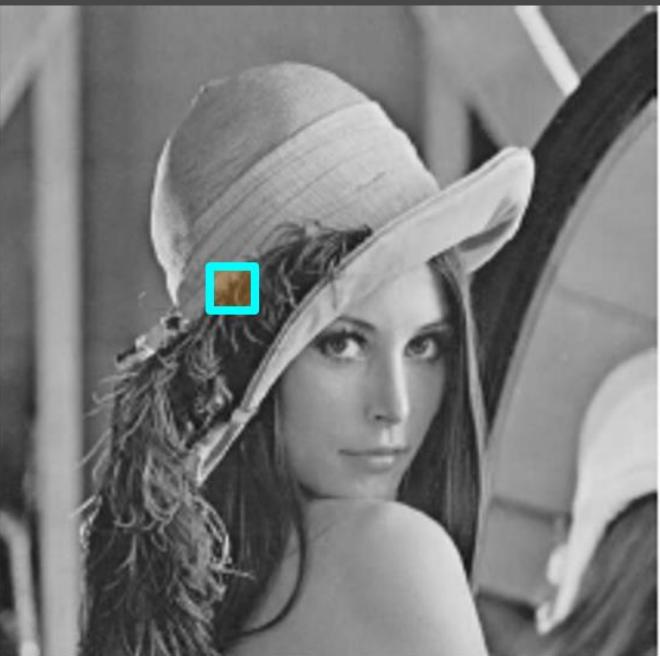


BM3D = **Sparse** representation on (3D) DCT dictionary + NLM (Dabov et al.'07)

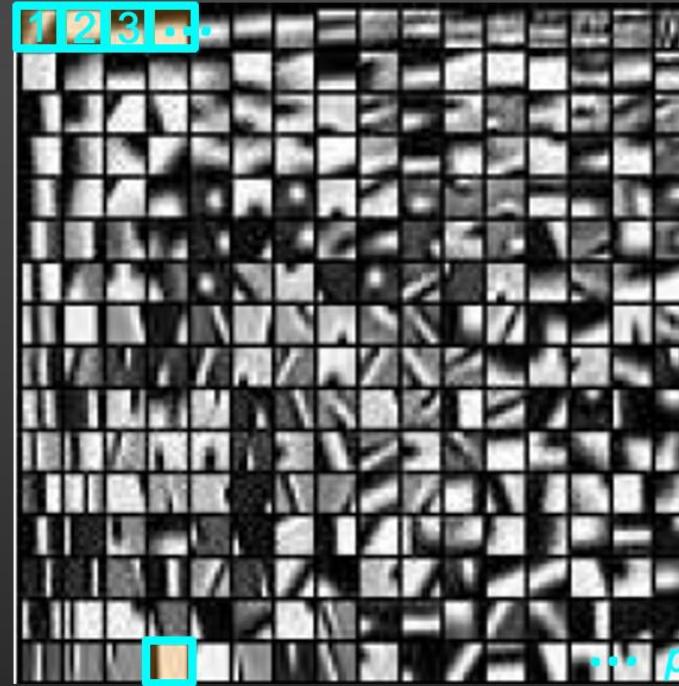
Take  $x \approx \sum_j \alpha^j d^j = D\alpha$  but limit the number of nonzero coefficients  $\|\alpha\|_0 \leq k$

# Linear signal models

Signal:  $x \in \mathbb{R}^m$



Dictionary:  
 $D = [d_1, \dots, d_p] \in \mathbb{R}^{m \times p}$

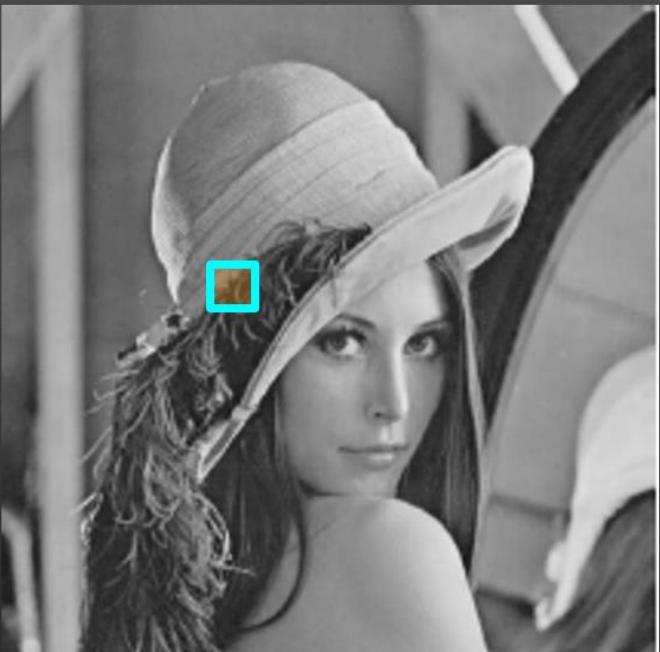


$$x \approx \alpha_1 d_1 + \alpha_2 d_2 + \dots + \alpha_p d_p = D\alpha, \text{ with } \alpha \in \mathbb{R}^p$$

- Note:
- > In general  $p \geq m$ . Here  $p=256, m=100$ .
  - > The dictionary has no spatial structure

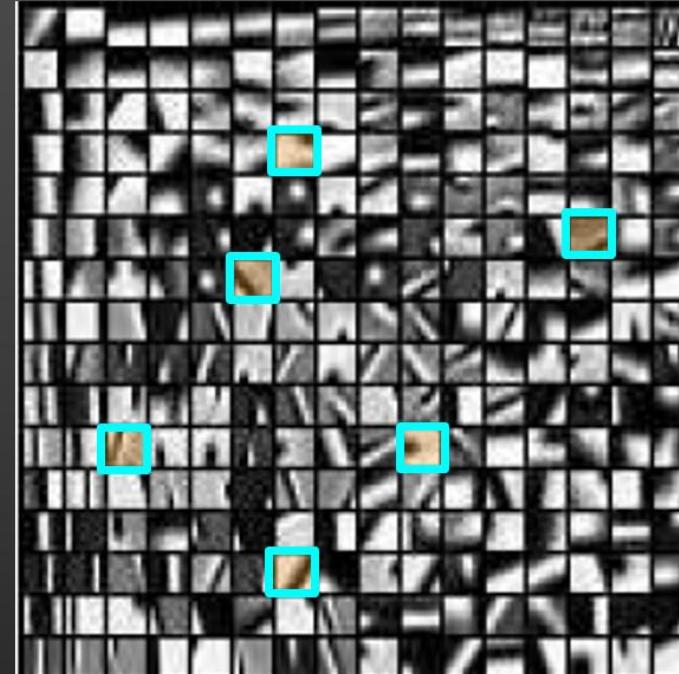
# Sparse linear models

Signal:  $x \in \mathbb{R}^m$



Dictionary:

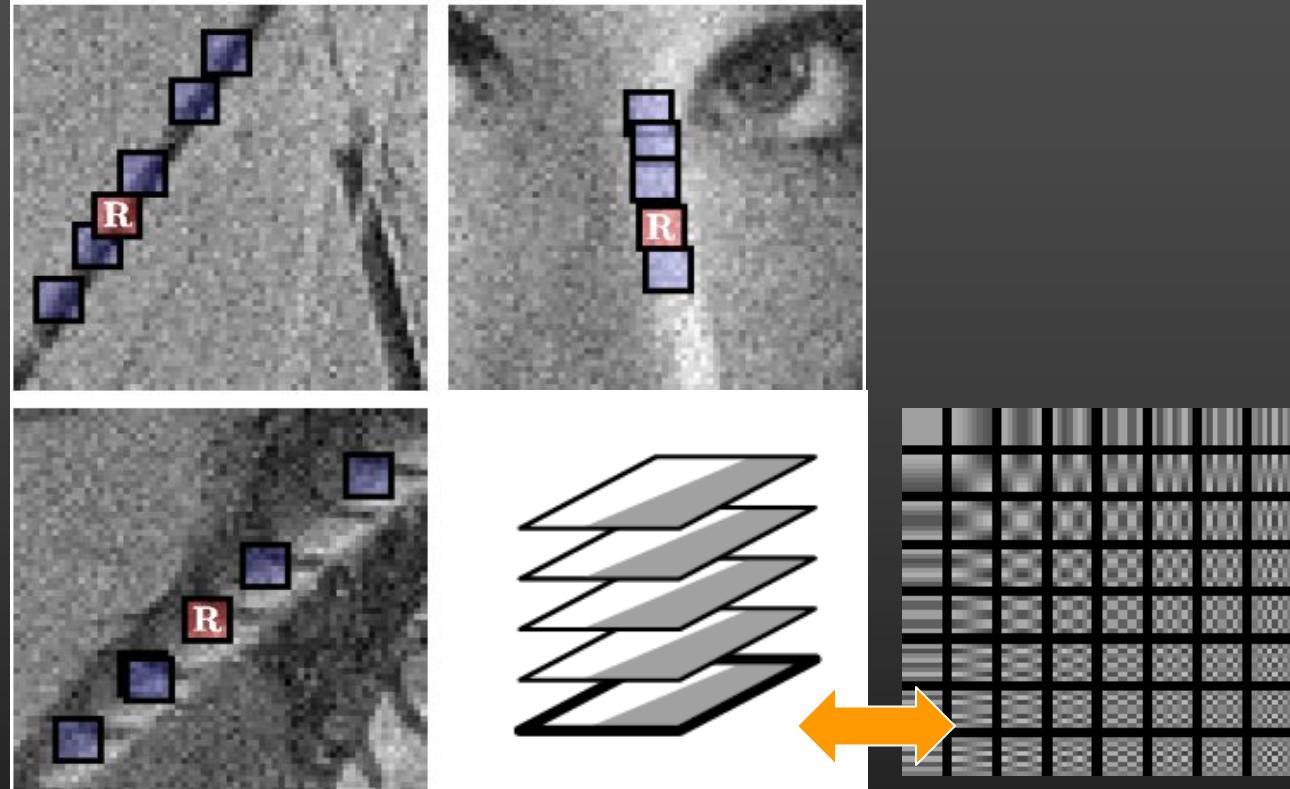
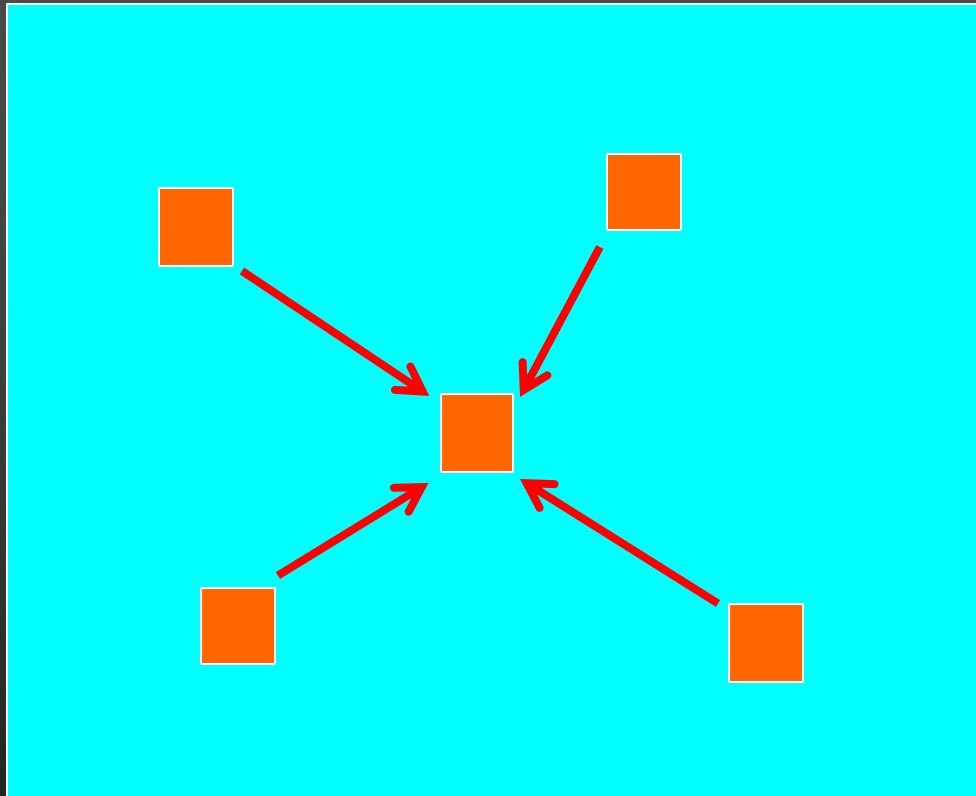
$D = [d_1, \dots, d_p] \in \mathbb{R}^{m \times p}$



$$x \approx \alpha_1 d_1 + \alpha_2 d_2 + \dots + \alpha_p d_p = D\alpha, \text{ with } |\alpha|_0 \ll p$$

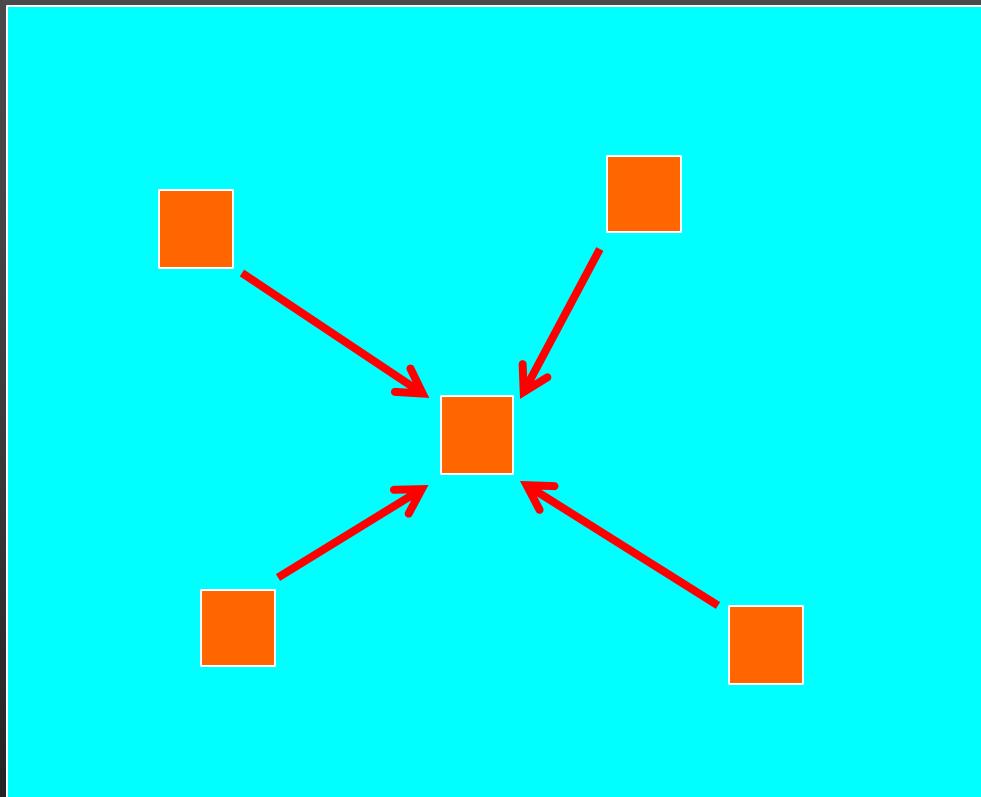
(Olshausen and Field, 1997; Chen et al., 1999; Mallat, 1999; Elad and Aharon, 2006)  
(Kavukcuoglu et al., 2009; Wright et al., 2009; Yang et al., 09; Boureau et al., 2010)

# Let us start simple: How to denoise an image



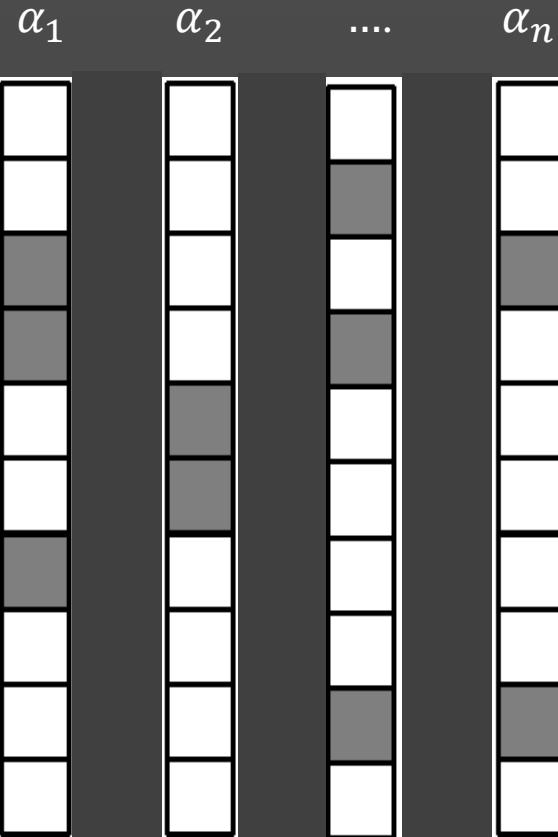
Take  $x \approx \sum_j \alpha^j d^j = D\alpha$  but limit the number of nonzero coefficients  $\|\alpha\|_0 \leq k$

# Let us start simple: How to denoise an image



Non-local means filtering (Buades et al.'05)

$$\min_{D, \alpha_1, \dots, \alpha_n} \sum \left\| x_i - D\alpha_i \right\|_F^2 + \lambda \left\| \alpha_i \right\|_1$$



LSC: **Dictionary learning** with sparsity  
(Elad & Aharon'06; Mairal et al.'08)

# Sparse coding and dictionary learning: A hierarchy of optimization problems

$$\min_{\alpha} \frac{1}{2} \|x - D\alpha\|_2^2$$

Least squares

$$\min_{\alpha} \frac{1}{2} \|x - D\alpha\|_2^2 + \lambda \|\alpha\|_0$$

Sparse coding

$$\min_{\alpha} \frac{1}{2} \|x - D\alpha\|_2^2 + \lambda \psi(\alpha)$$

Dictionary learning

$$\min_{D \in C, \alpha_1, \dots, \alpha_n} \sum_{1 \leq i \leq n} [ \frac{1}{2} \|x_i - D\alpha_i\|_2^2 + \lambda \psi(\alpha_i) ]$$

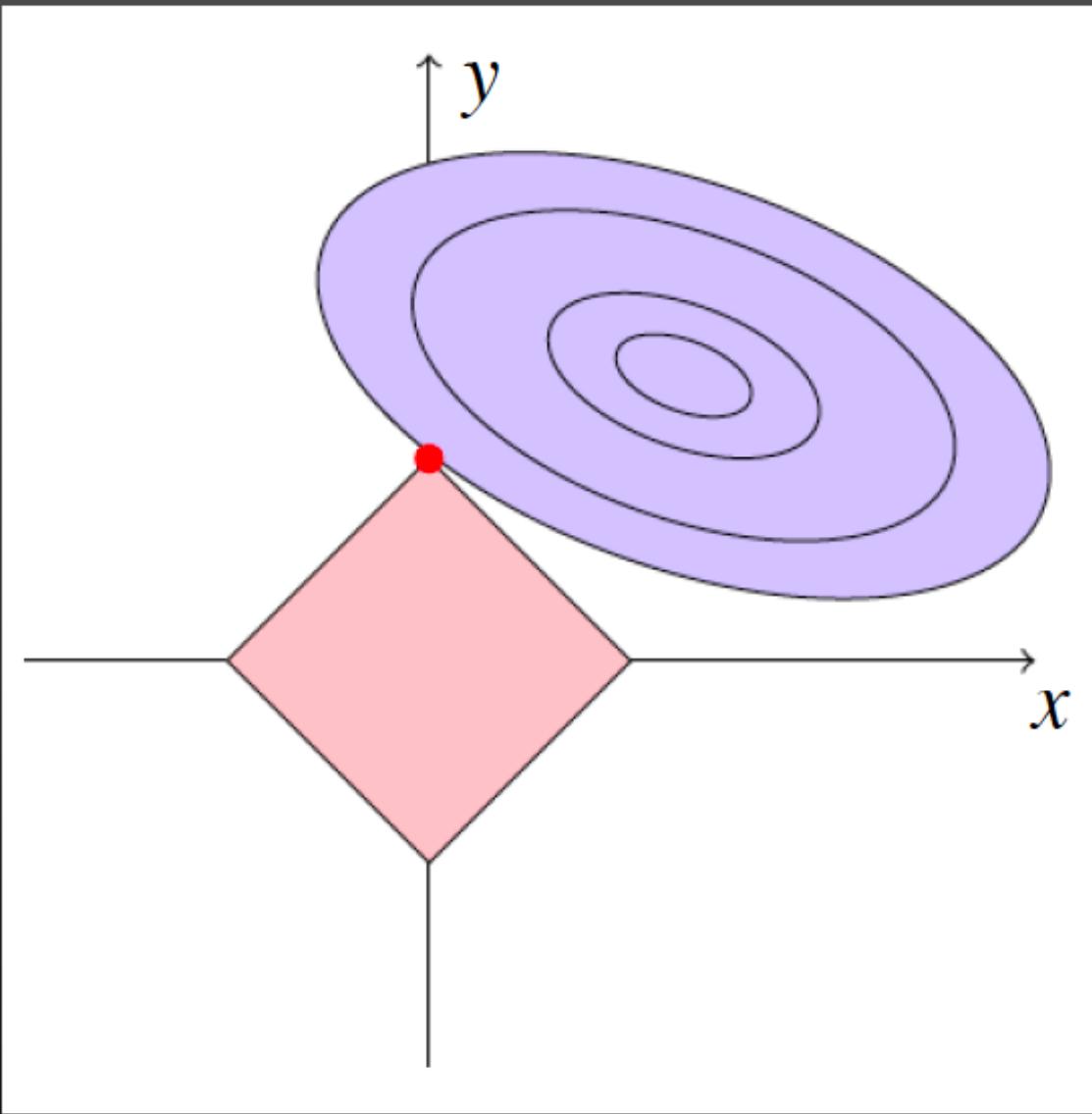
Learning for a task

$$\min_{D \in C, \alpha_1, \dots, \alpha_n} \sum_{1 \leq i \leq n} [ f(x_i, D, \alpha_i) + \lambda \psi(\alpha_i) ]$$

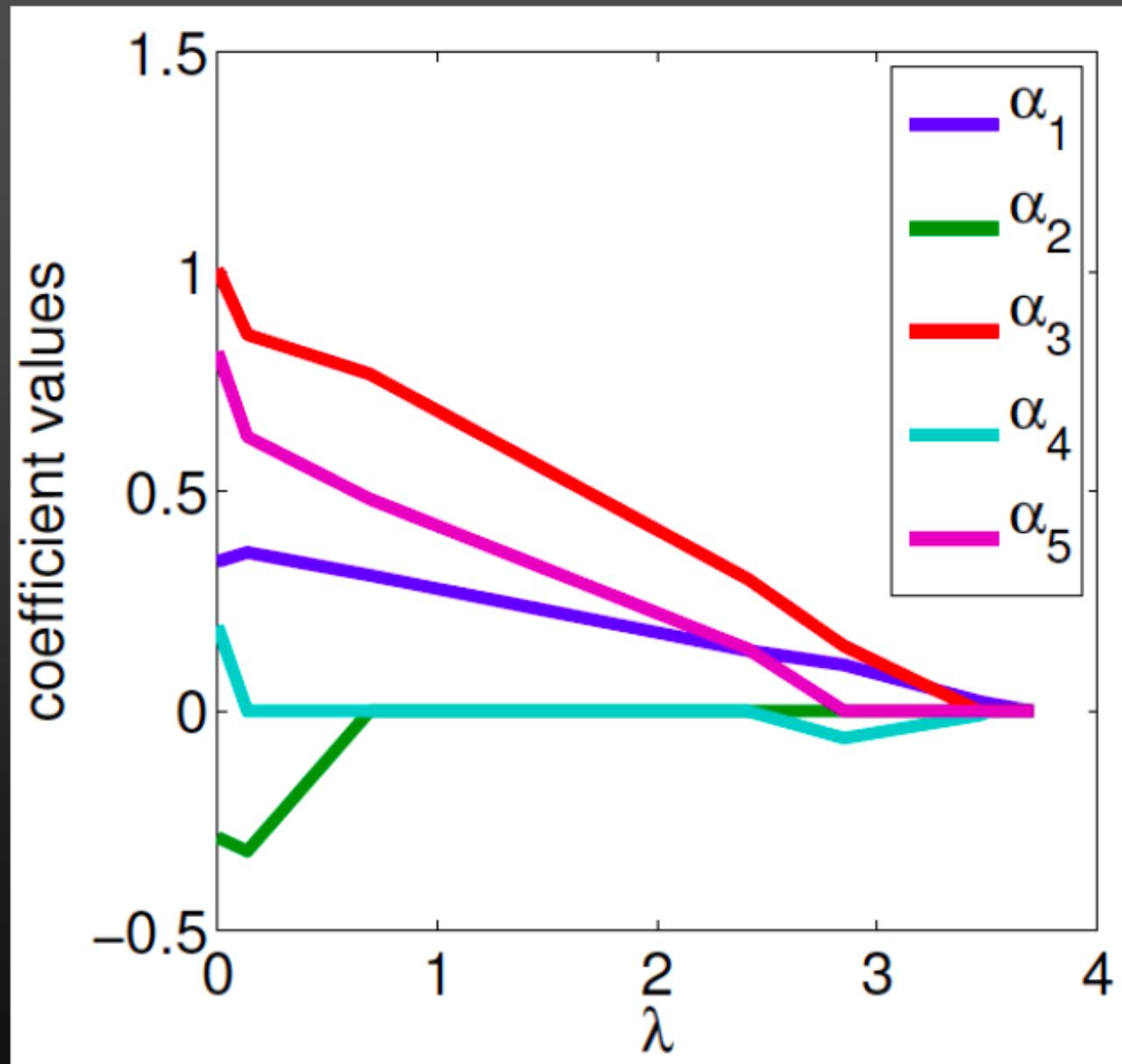
Learning structures

$$\min_{D \in C, \alpha_1, \dots, \alpha_n} \sum_{1 \leq i \leq n} [ f(x_i, D, \alpha_i) + \lambda \sum_{1 \leq k \leq q} \psi(d_k) ]$$

# The $\ell_1$ norm and sparsity



# LARS (Efron et al., 2004)



# Dictionary learning

- Given some loss function, e.g.,

$$L(x, D) = 1/2 \|x - D\alpha\|_2^2 + \lambda \|\alpha\|_1$$

- One usually minimizes, given some data  $x_i$ ,  $i = 1, \dots, n$ , the empirical risk:

$$\min_D f_n(D) = \sum_{1 \leq i \leq n} L(x_i, D)$$

- But, one would really like to minimize the expected one, that is:

$$\min_D f(D) = \mathbb{E}_x [L(x, D)]$$

# Online sparse matrix factorization

(Mairal, Bach, Ponce, Sapiro, ICML'09, JMLR'10)

Problem:

$$\min_D f(D) = E_x [L(x, D)]$$

$$L(x, D) = 1/2 \|x - D\alpha\|_2^2 + \lambda \|\alpha\|_1$$

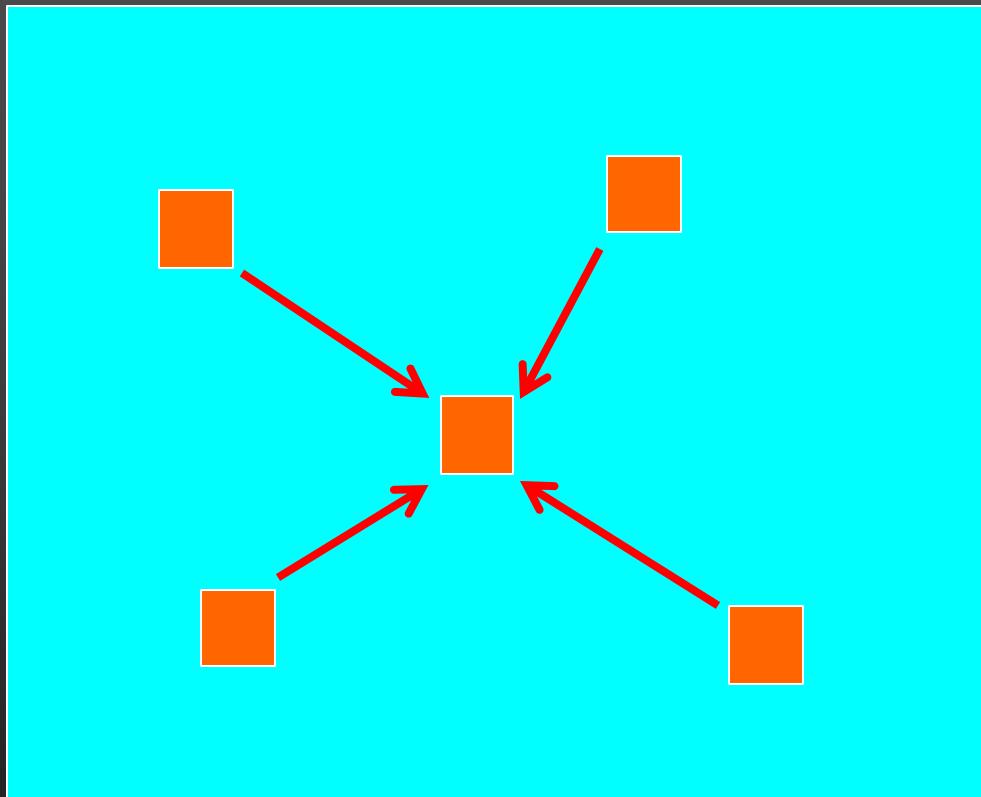
Algorithm:

Iteratively draw one random training sample  $x_t$  and minimize the quadratic surrogate function:

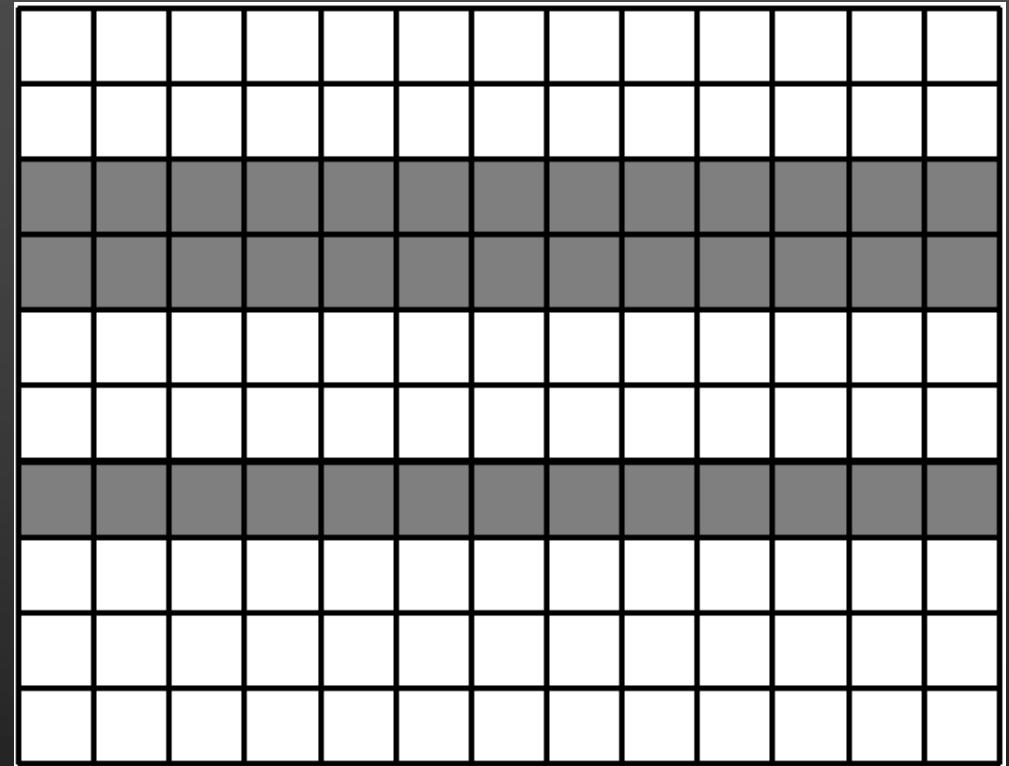
$$g_t(D) = 1/t \sum_{1 \leq i \leq t} [1/2 \|x_i - D\alpha_i\|_2^2 + \lambda \|\alpha_i\|_1]$$

(Lars/Lasso for sparse coding, block-coordinate descent with warm restarts for dictionary updates, mini-batch extensions, etc.)

# Let us start simple: How to denoise an image



Non-local means filtering (Buades et al.'05)



LSSC: Dictionary learning with **structured sparsity** (Mairal et al.'09)

$$\min_{D,A} \sum_i \|X_i - DA_i\|_F^2 + \lambda \|A_i\|_{1,2} \text{ where } \|A\|_{1,2} = \sum_r \|\alpha^r\|_2$$



LSSC



# Real noise is complicated

- Noise = shot noise (physics) plus read noise (electronics)
- Random variable following a Gaussian distribution with zero mean and signal-dependent standard deviation function (Foi et al., 2008)

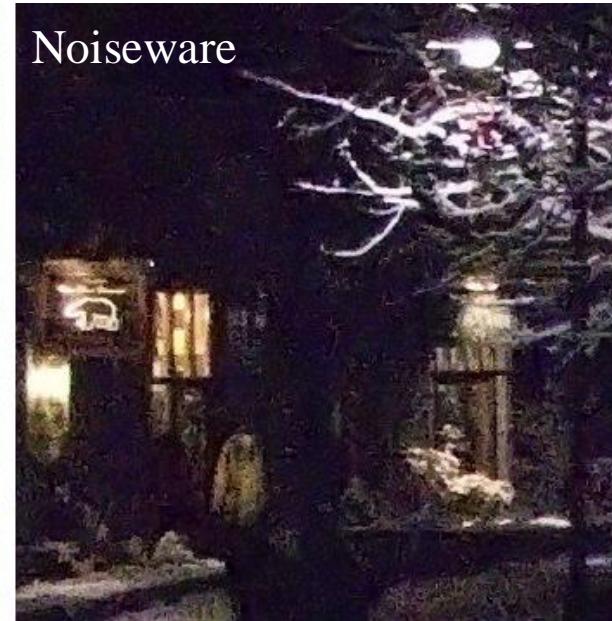
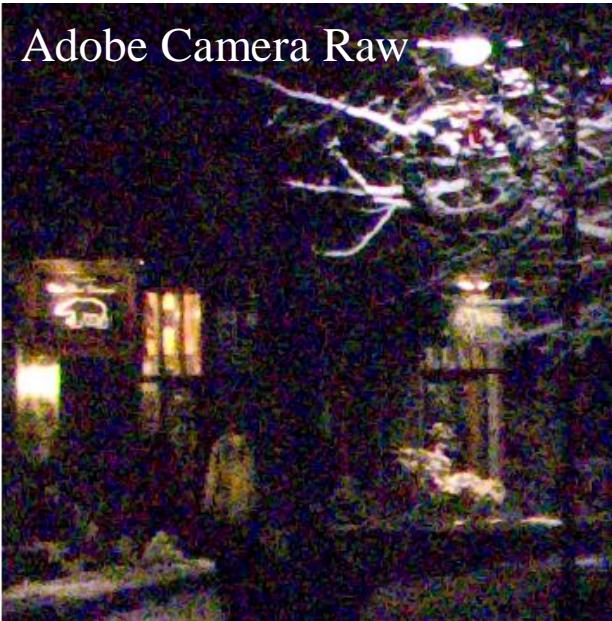
$$s(u) = \sqrt{\alpha y(u) + \beta}$$

whose parameters  $\alpha$  and  $\beta$  can be determined for a given camera

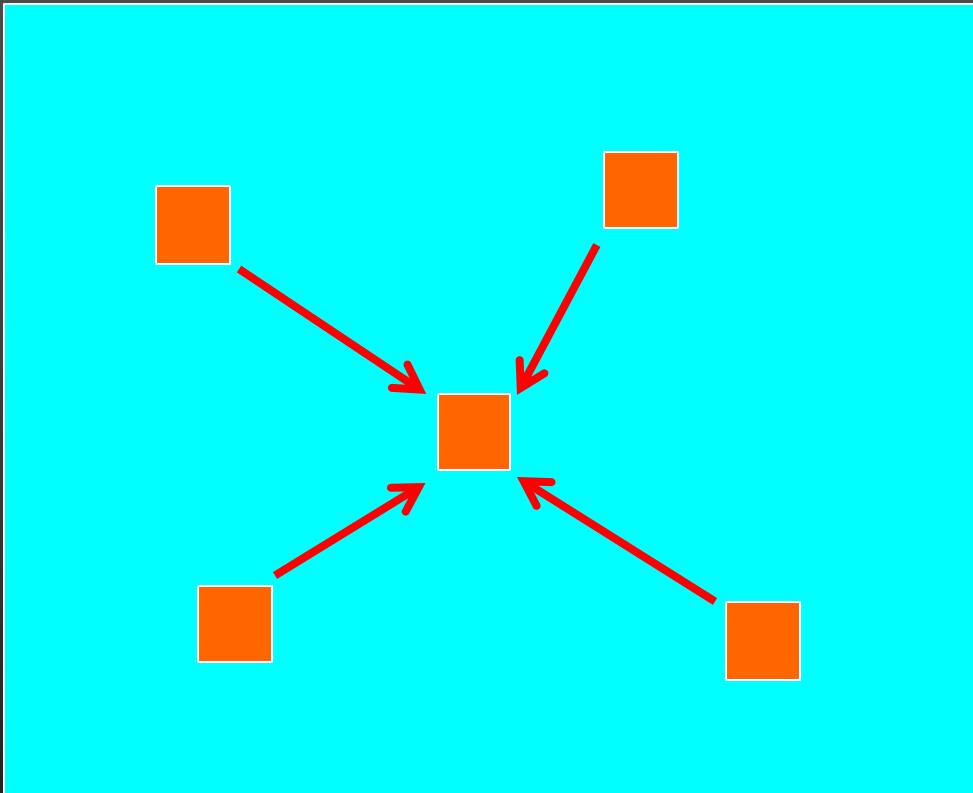
- This is only true for raw images. More on that later

A. Foi, M. Trimeche, V. Katkovnik, K. Egiazarian, "Practical Poissonian-Gaussian noise modeling and fitting for single-image raw-data", IEEE TIP 17(10):1737-1754 (2008).

# Real noise (Canon Powershot G9, 1600 ISO)



# Let us start simple: How to denoise an image



Non-local means filtering (Buades et al.'05)

**Self-attention** (Vaswani et al., 2017)

$$X_i = S_i V_i, \text{ where } S_i = \text{softmax}\left(\frac{1}{\tau} K_i Q_i^T\right)$$

where

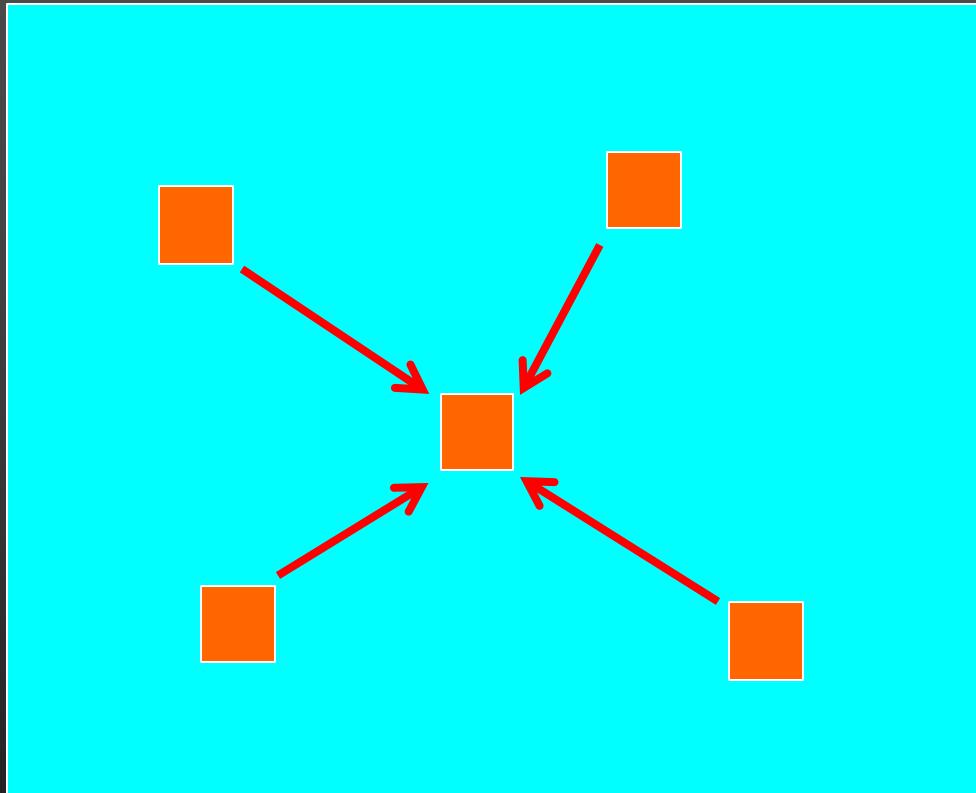
$$K_i = X_{i-1} A_i, \quad Q_i = X_{i-1} B_i, \quad \text{and} \quad V_i = X_{i-1} C_i$$



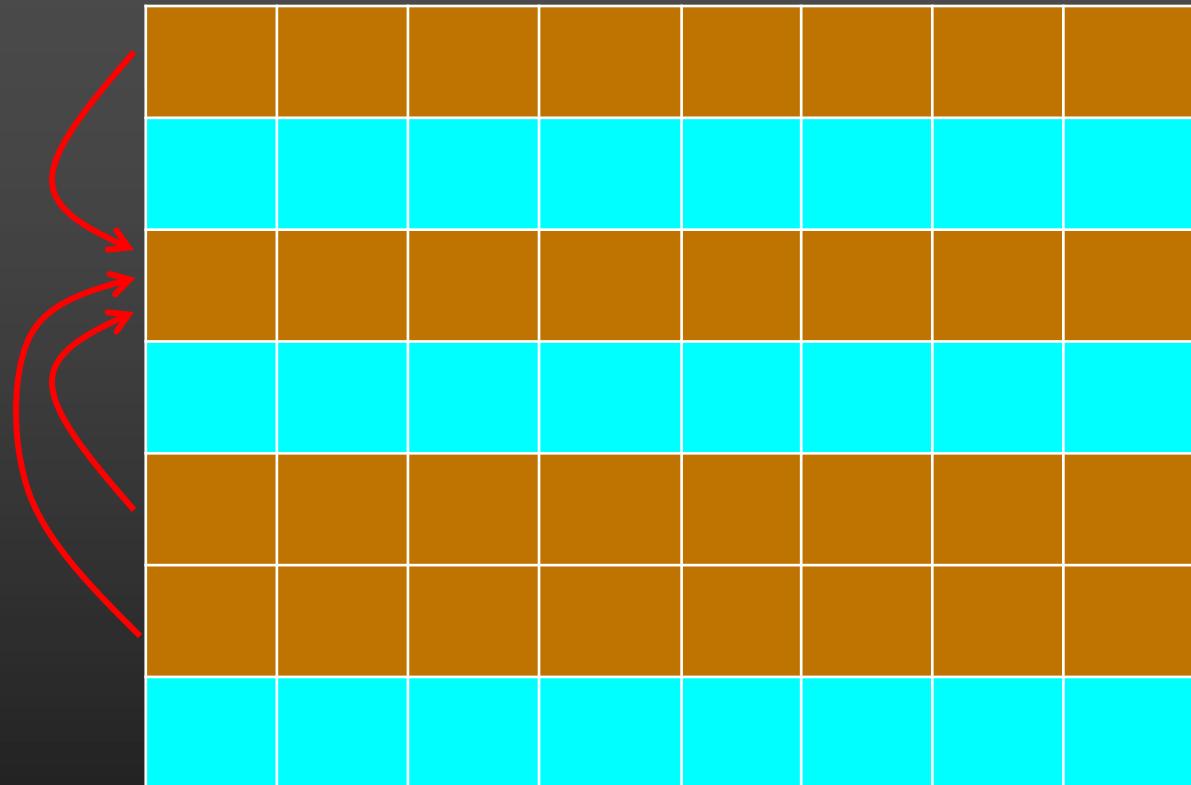
$$\begin{aligned} X' &= X_k = S_k X_{k-1} C_k = S_k (S_{k-1} X_{k-2} C_{k-1}) C_k \\ &= (S_k \dots S_1) X (C_1 \dots C_k) = T_k X D_k, \end{aligned}$$

Note:  $T_k$  is a stochastic matrix, thus the rows of  $T_k X$  are barycentric combinations of all the rows of  $X$ , weighted in a complex way by their affinities  $X_{i-1} A_i B_i^T X_{i-1}^T$ . (See also Andrej Karpathy's talk.)

# Let us start simple: How to denoise an image



Non-local means filtering (Buades et al.'05)



Note:  $T_k$  is a stochastic matrix, thus the rows of  $T_k X$  are barycentric combinations of all the rows of  $X$ , weighted in a complex way by their affinities  $X_{i-1} A_i B_i^T X_{i-1}^T$ . (See also Andrej Karpathy's talk.)

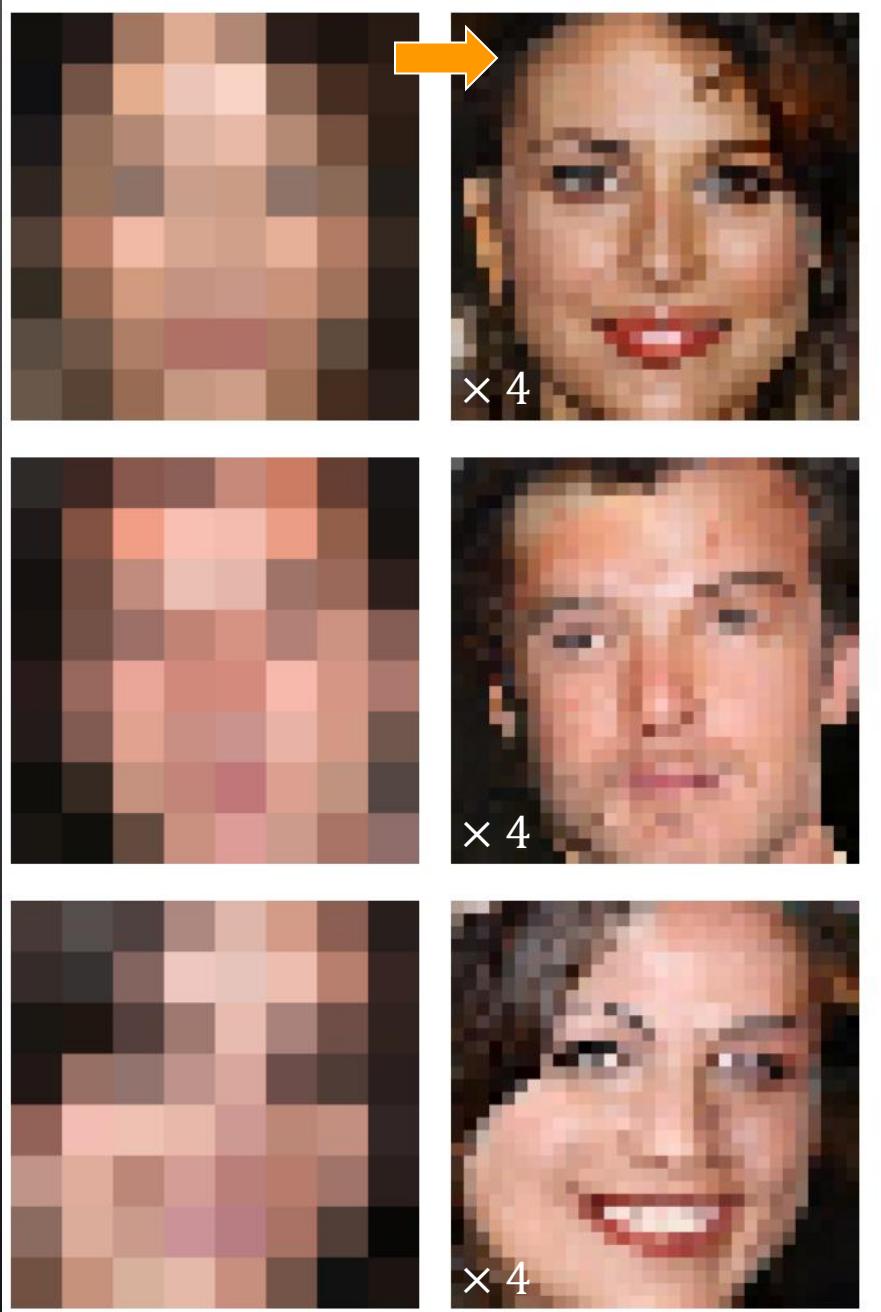


Image interpolation  
aka  
Depixellation  
aka  
Example-based super-resolution  
aka  
Single-image super-resolution

(Dahl et al., 2017)

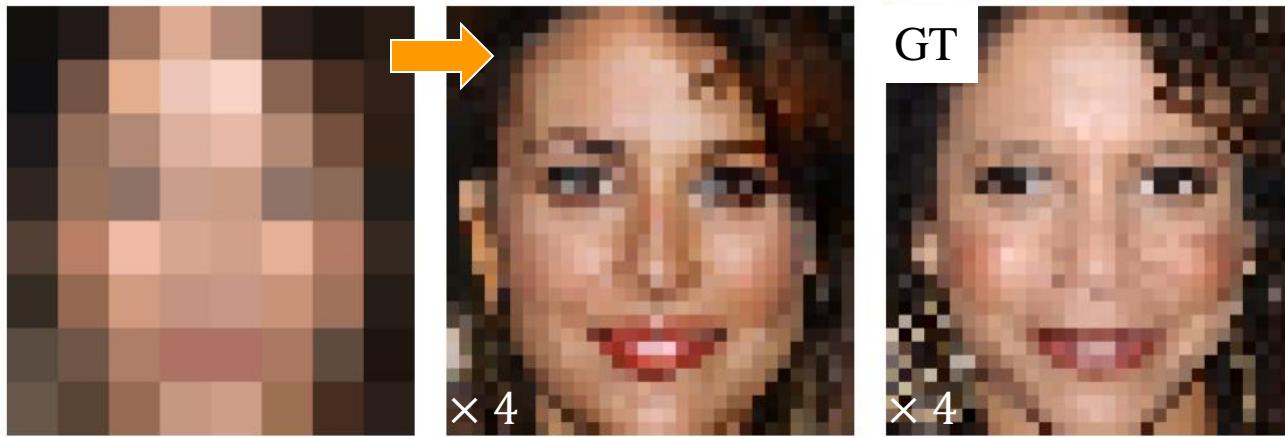
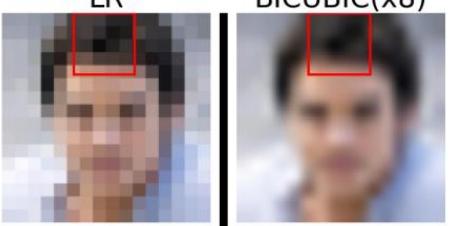


Image interpolation  
aka  
Depixellisation  
aka  
Example-based super-resolution  
aka  
Single-image super-resolution  
(Dahl et al., 2017)

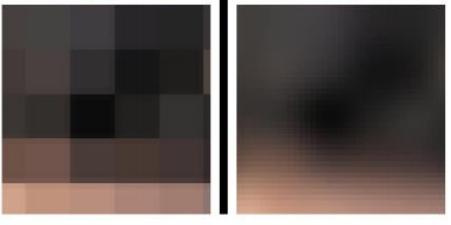
Problem: Not enough information  
in a single image: details must be  
hallucinated

(FSRNET Chen et al., 2017)  
(FSRGAN Zhu et al., 2020)  
(PULSE Menon et al., 2020)

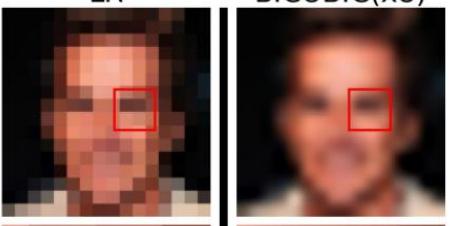
LR



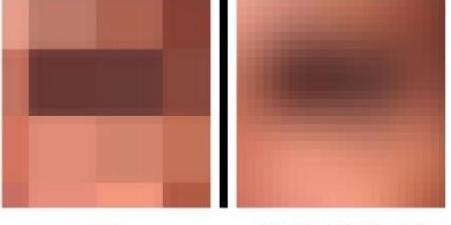
BICUBIC(x8)



LR



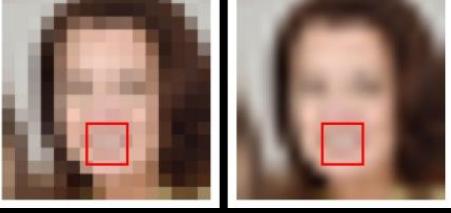
BICUBIC(x8)



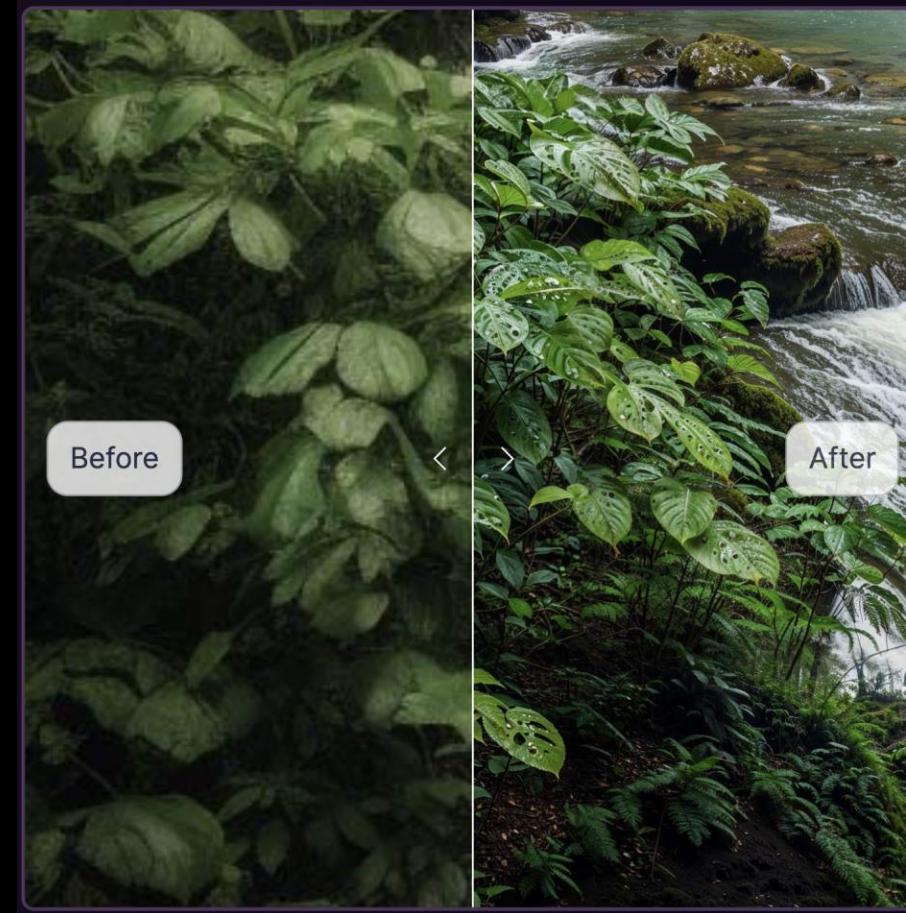
LR



BICUBIC(x8)

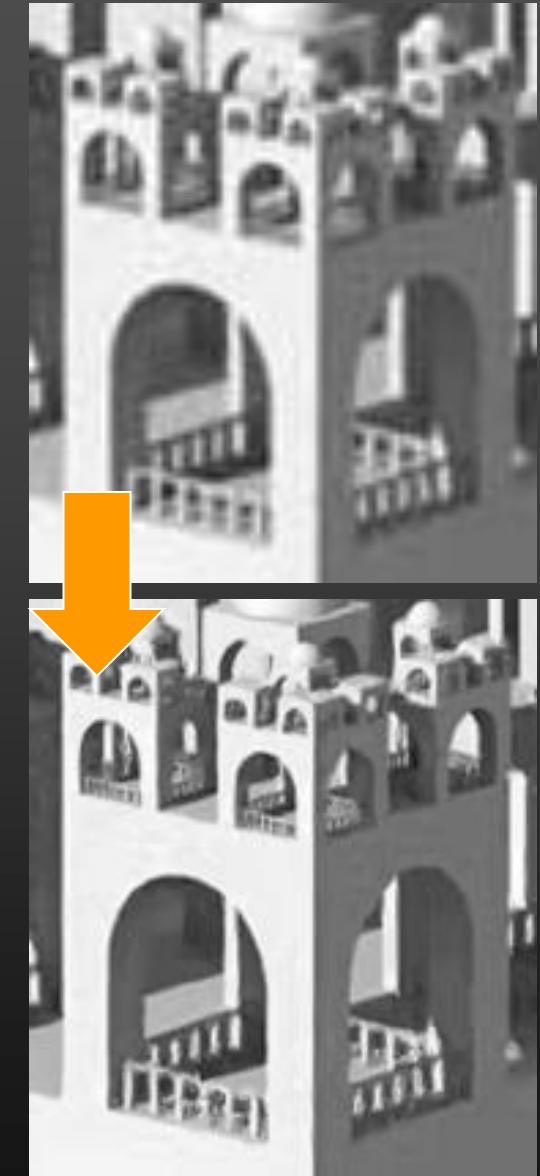
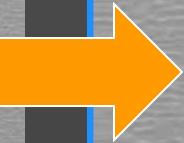


# Generative vision

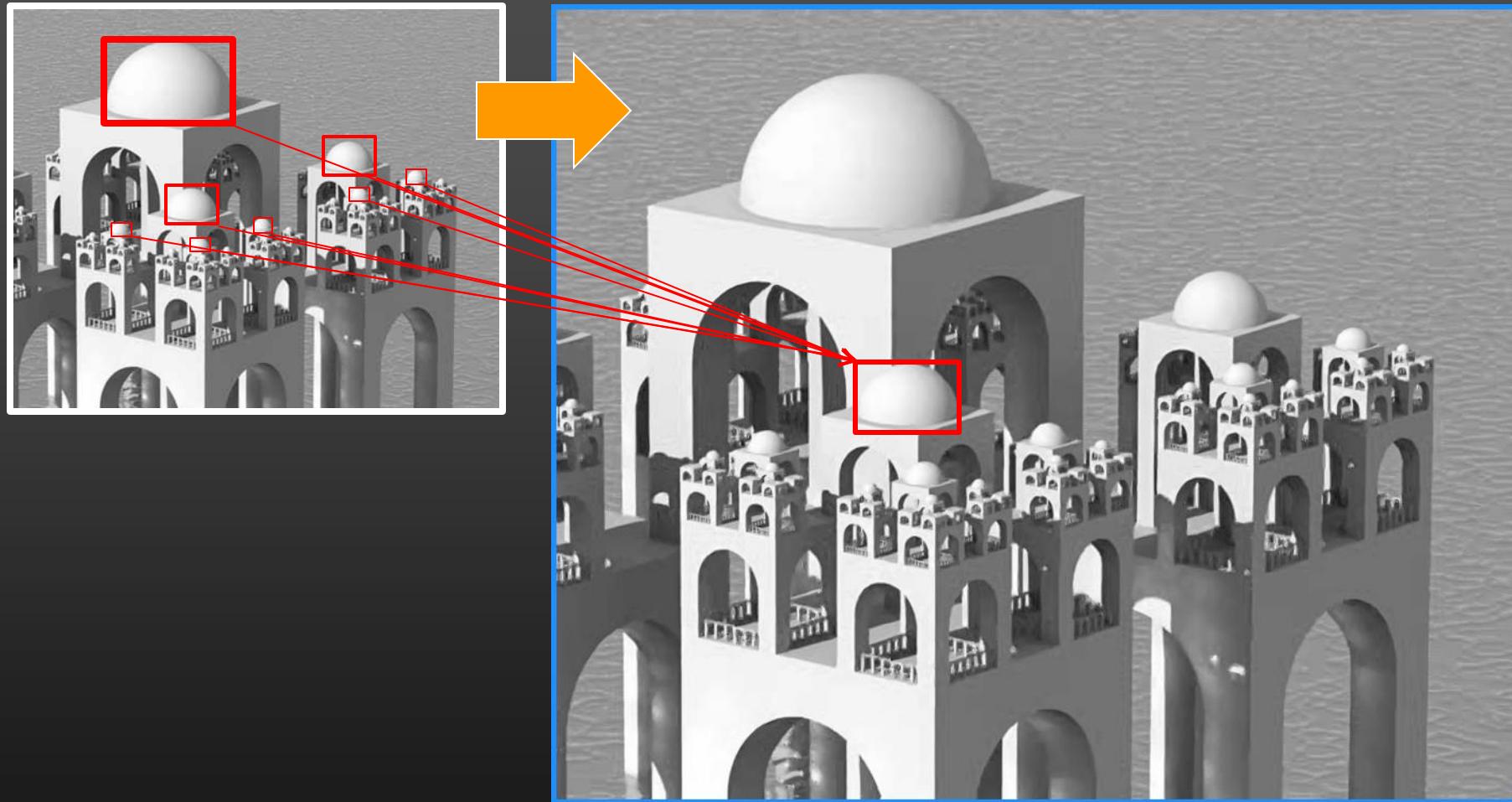


<https://magnific.ai/>

# Super-Resolution from a Single Image (Glasner, Bagon, Irani, ICCV'09)



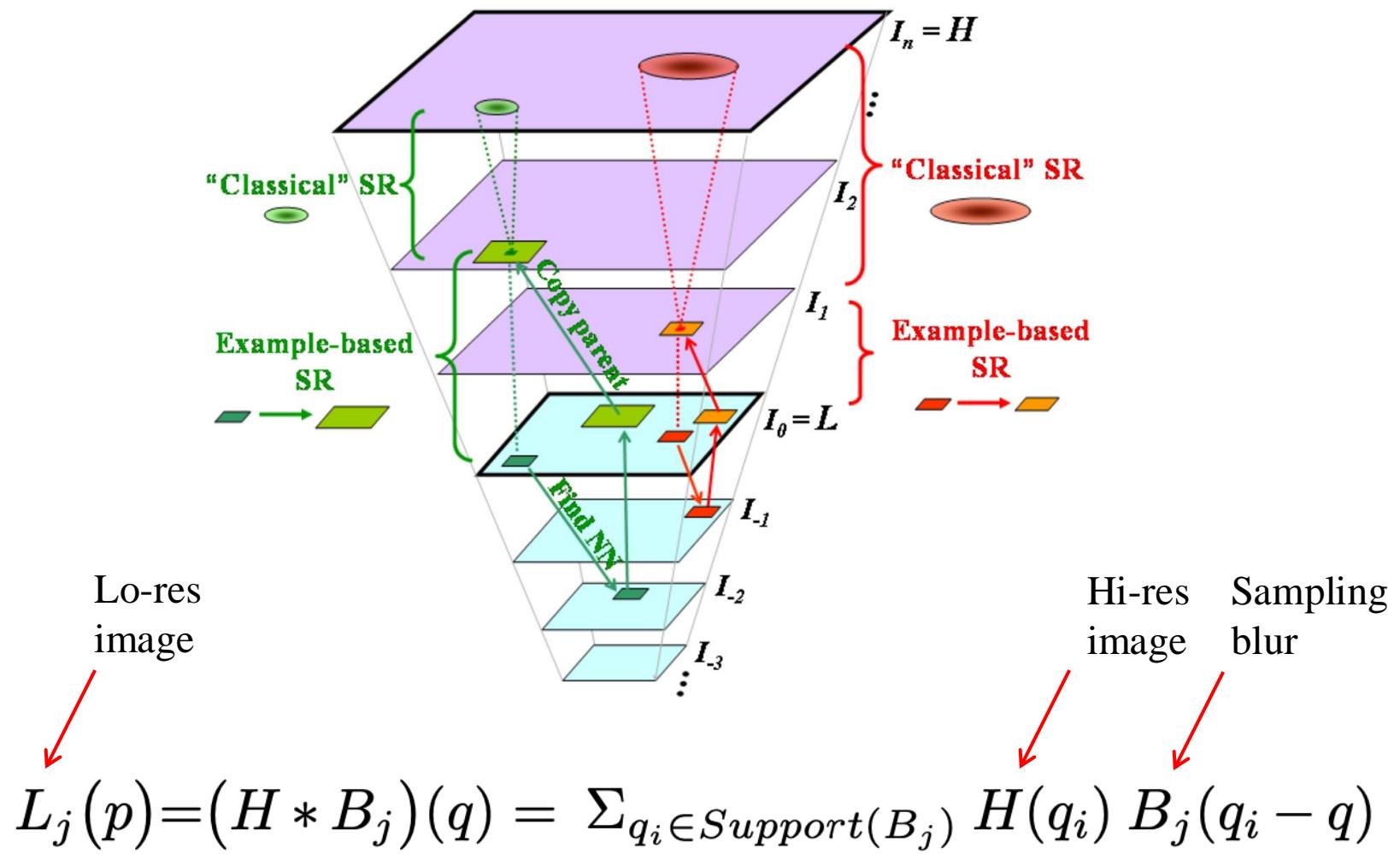
# Super-Resolution from a Single Image (Glasner, Bagon, Irani, ICCV'09)



Key idea: exploit internal self-similarities

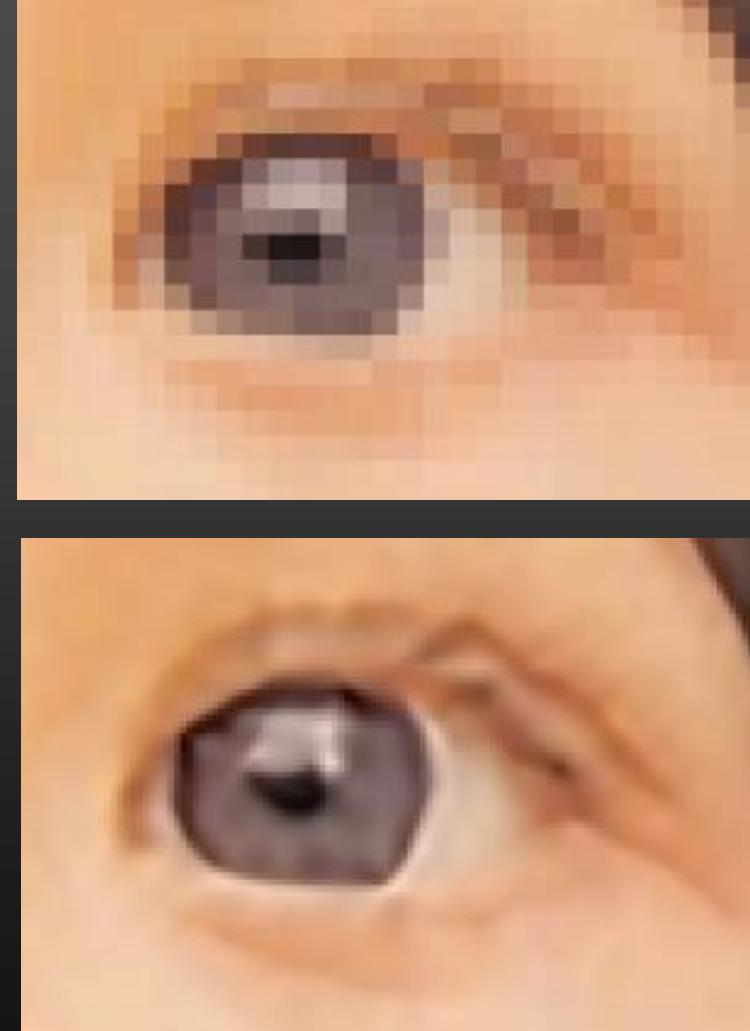
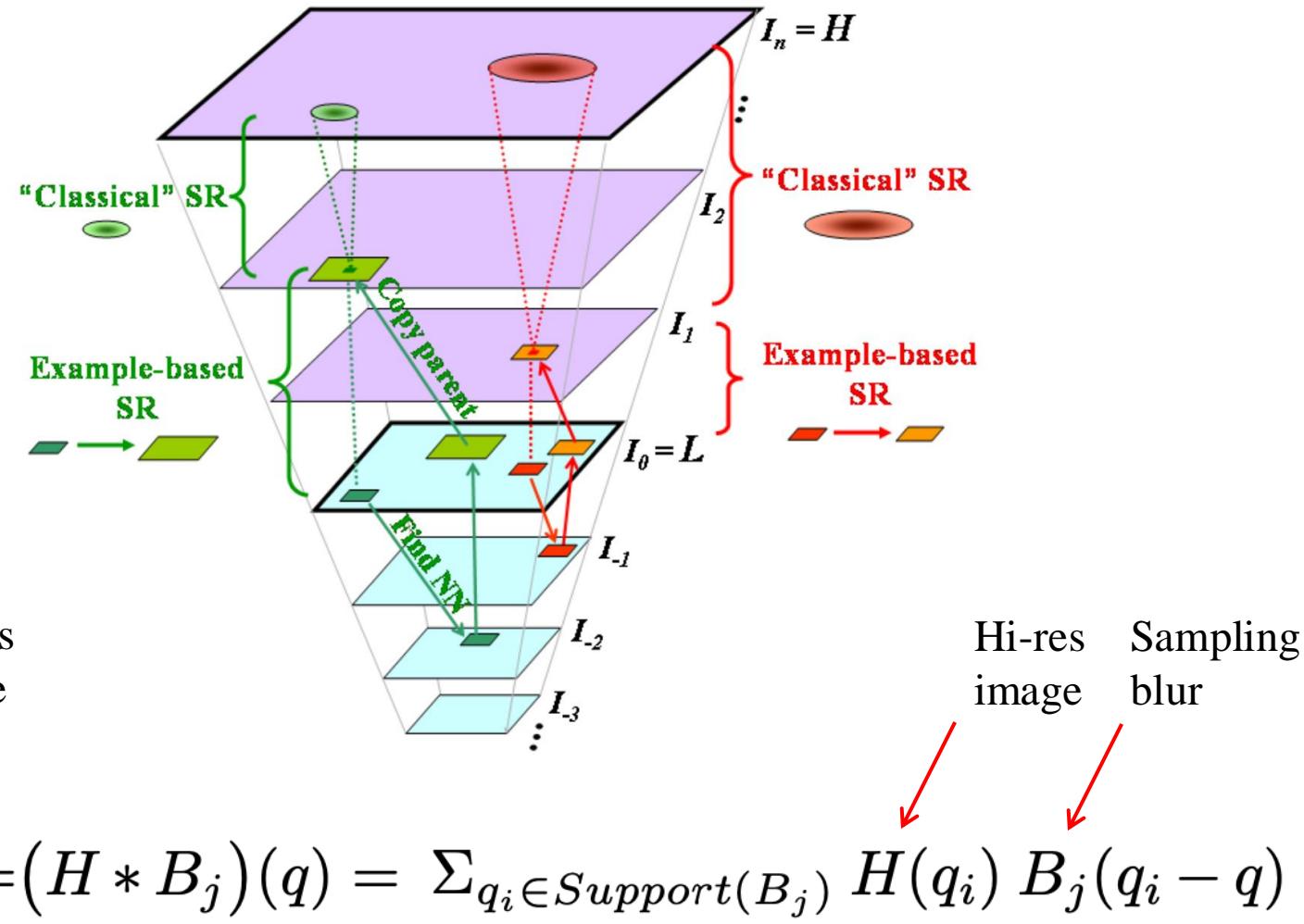


# Super-Resolution from a Single Image (Glasner, Bagon, Irani, ICCV'09)



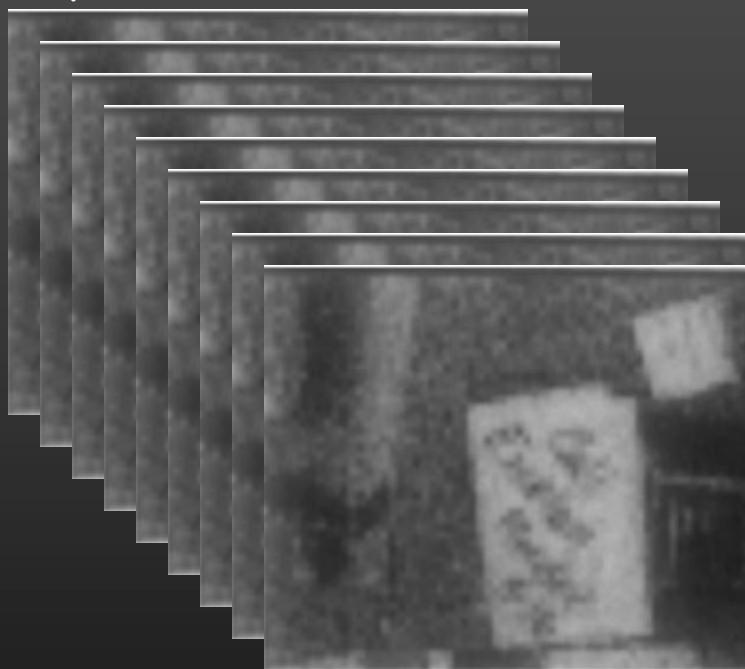
# Super-Resolution from a Single Image

(Glasner, Bagon, Irani, ICCV'09)



True (= multi-frame) super-resolution: no need to hallucinate details

Input burst



Initial estimate

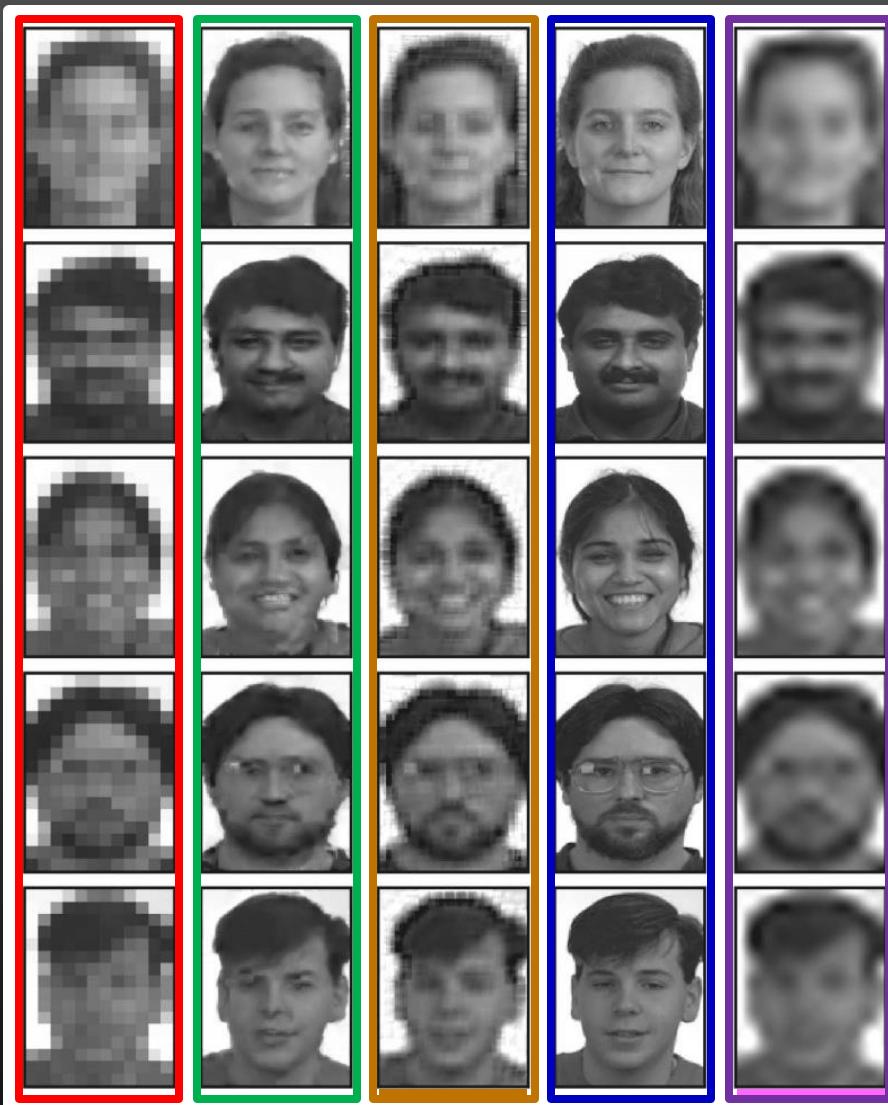


Refined estimate



(Irani & Peleg, 1991; see also Tsai & Huang, 1984)

# Super-resolution with “hallucination/recogstruction”



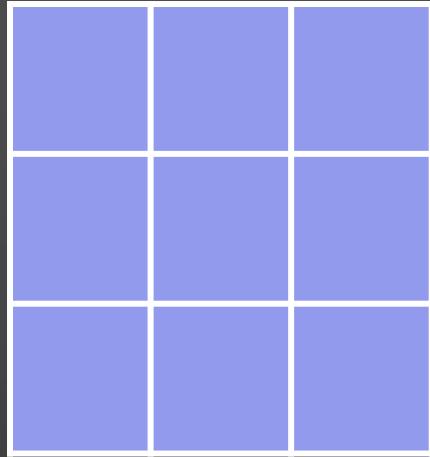
- LR input image (1 of 4)
- Recogstruction
- Ground-truth HR image
- (Hardie et al., 1997)
- Bicubic interpolation

× 4, alignment  
known exactly

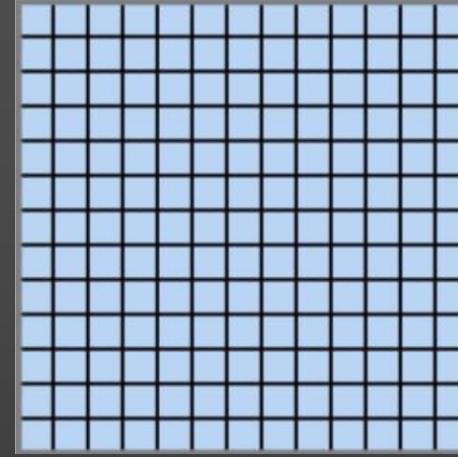
Key idea: learn a prior on the spatial distribution of the image gradient for frontal images of faces

(Baker and Kanade, 2002)

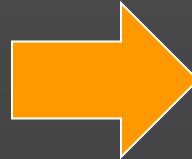
1 LR RGB image



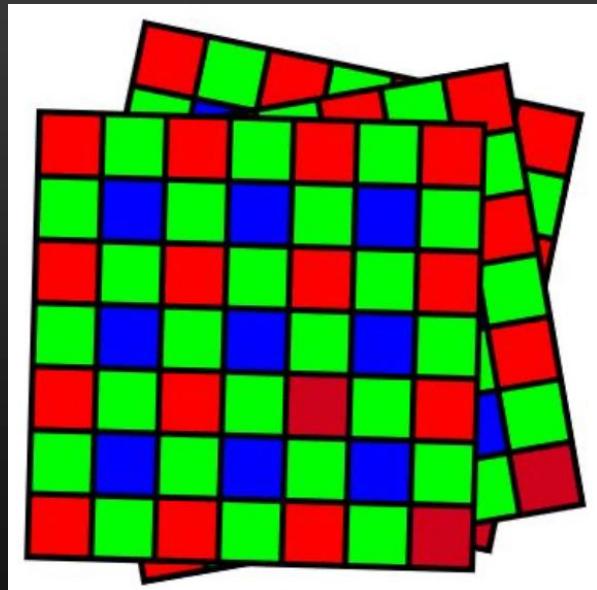
1 HR RGB image



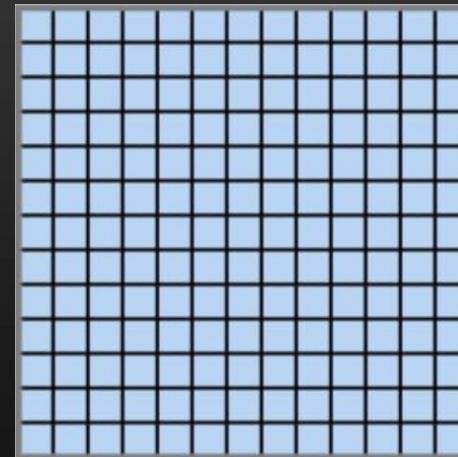
Single-image  
interpolation



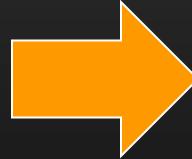
20 LR raw images = burst



1 HR RGB image

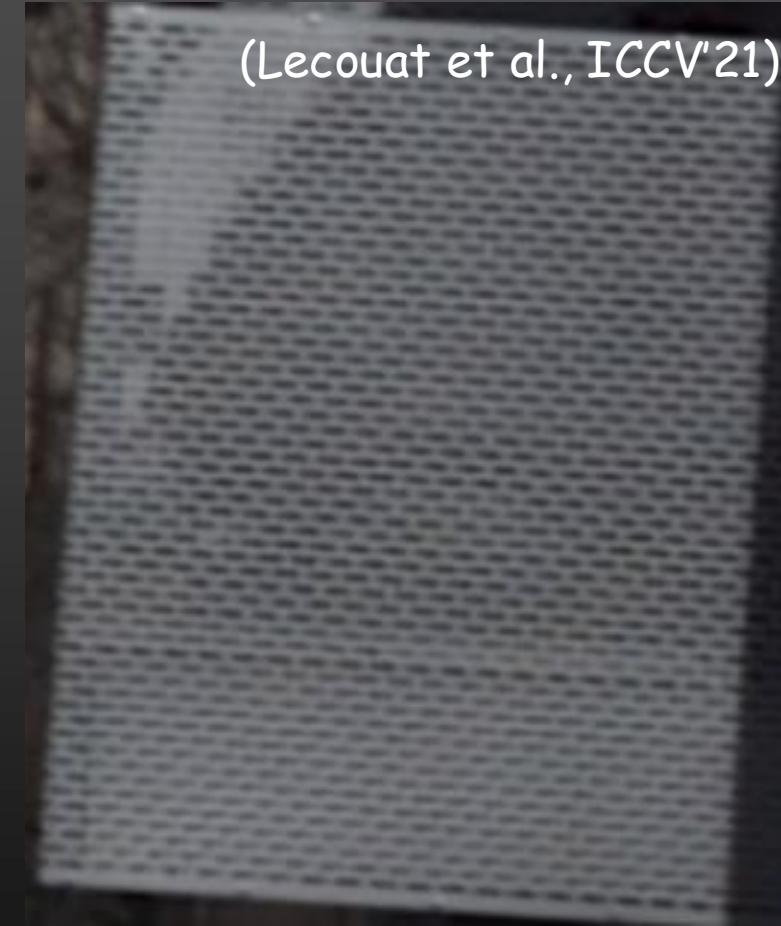
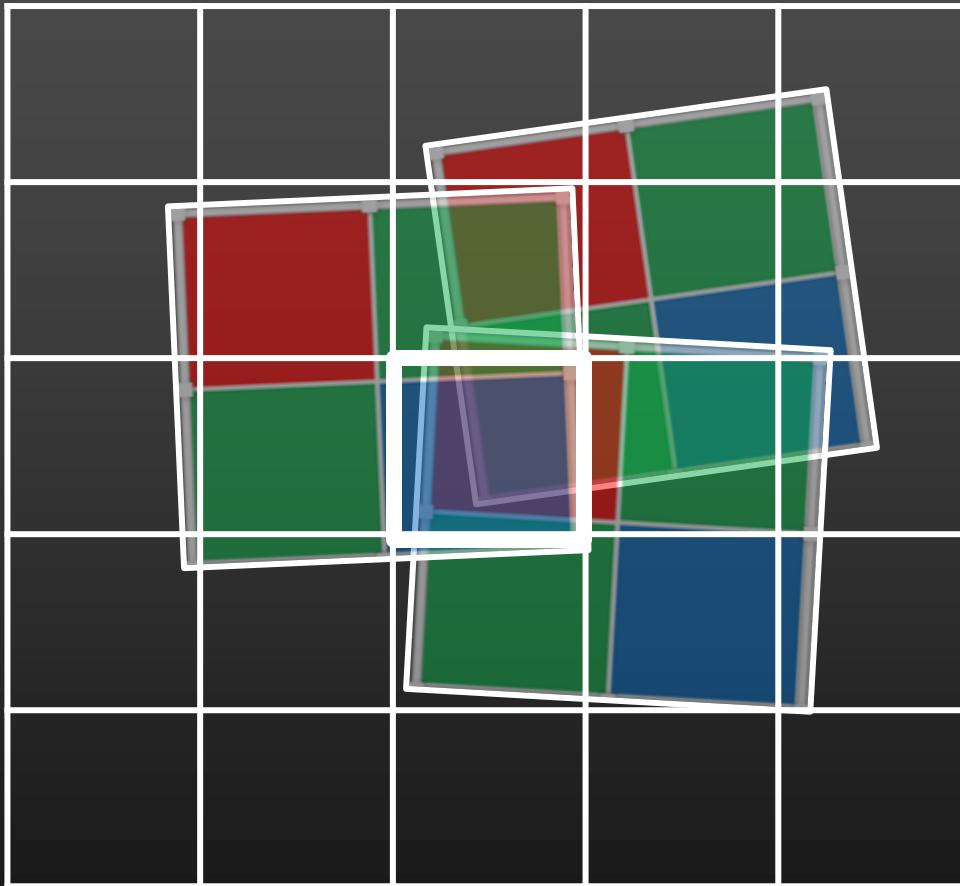


Super-resolution



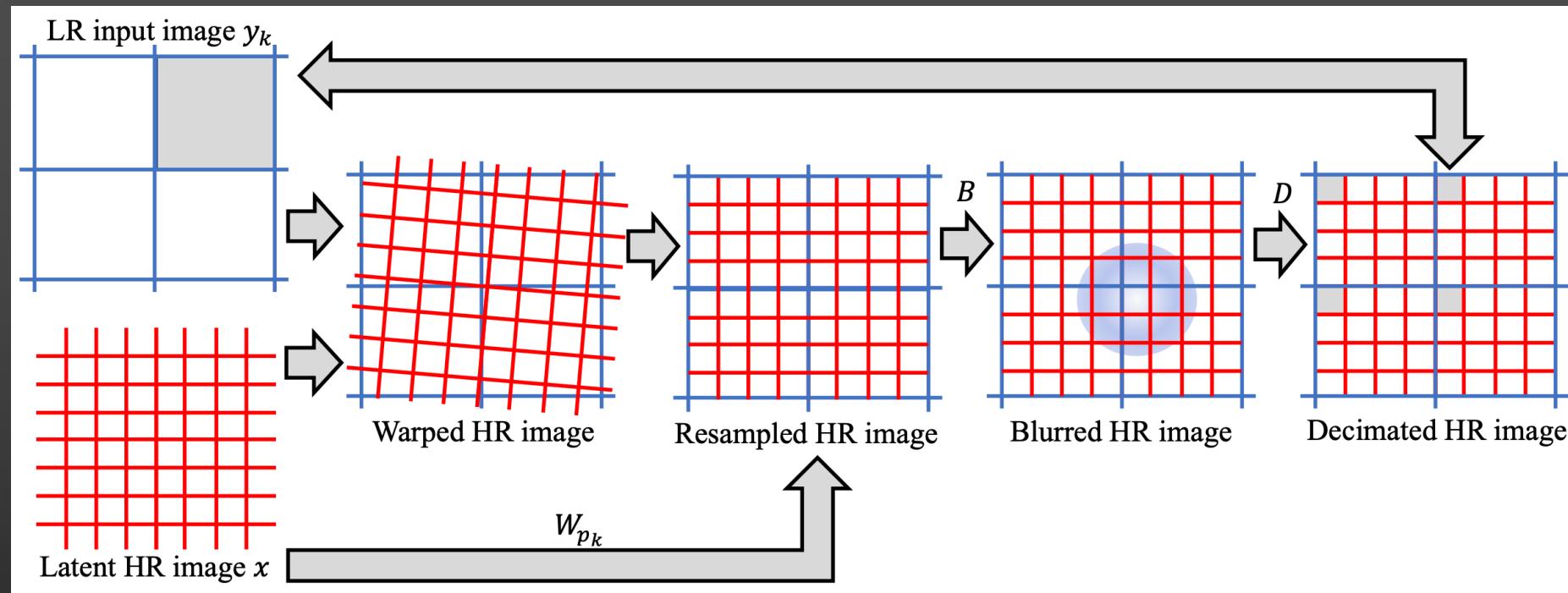
# Handheld Multi-Frame Super-Resolution

(Wronski, Garcia-Dorado, Ernst, Kelly, Kainin, Liang, Levoy, Milanfar, SIGGRAPH'19)



Key idea: exploit natural hand tremor and avoid single-image demosaicing altogether

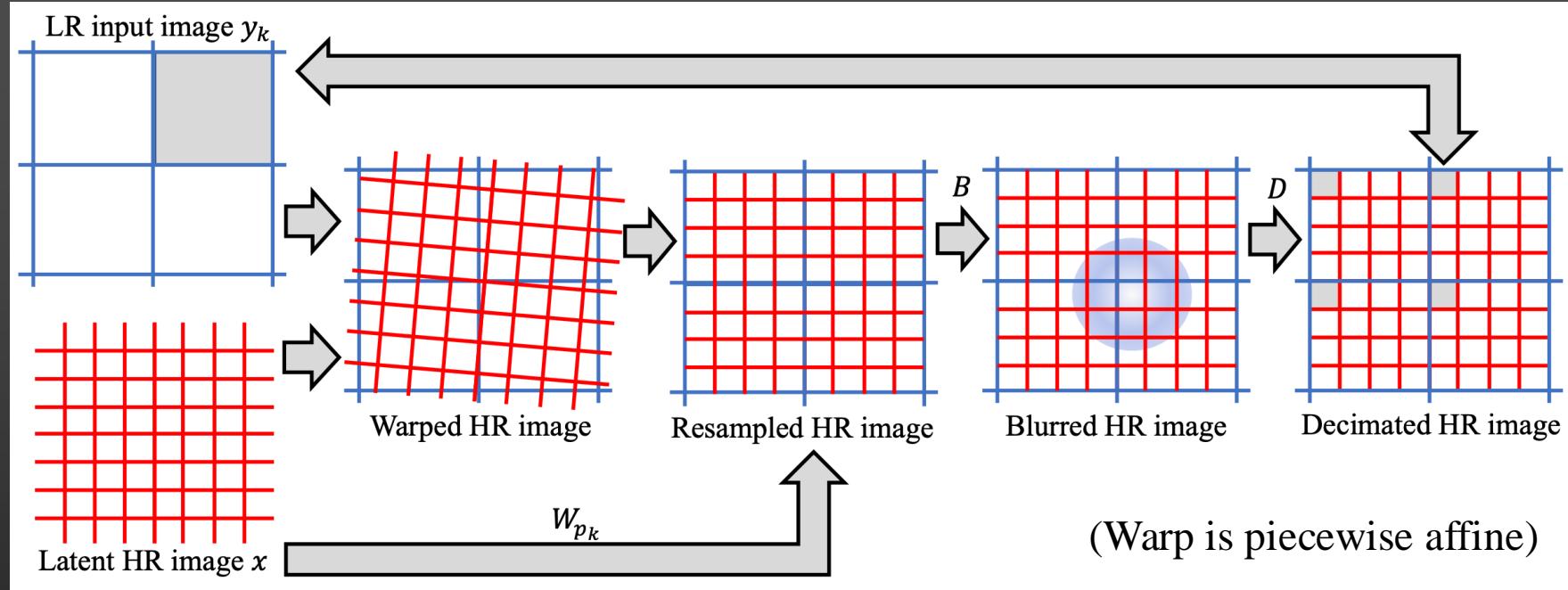
# Super-resolution as an inverse problem



- Forward model:  $y_k = U_{p_k} x + \varepsilon_k$  for  $k = 1, \dots, K$  with  $U_{p_k} = DBW_{p_k}$

- Solve  $\min_{x,p} \frac{1}{2} \|y - U_p x\|^2 + \lambda \varphi(x)$  where  $y = \begin{bmatrix} y_1 \\ \vdots \\ y_K \end{bmatrix}$  and  $U_p = \begin{bmatrix} U_{p_1} \\ \vdots \\ U_{p_K} \end{bmatrix}$

# Lucas-Kanade Reloaded: End-to-End Super-Resolution from Raw Image Bursts (Lecouat, Ponce, Mairal, ICCV'21)



- $y_k = U_{p_k} x + \varepsilon_k$  for  $k = 1, \dots, K$  with  $U_{p_k} = DBW_{p_k}$
- Define  $x_\theta(y) = \operatorname{argmin}_{x,p} \frac{1}{2} \|y - U_p x\|^2 + \lambda \varphi_\theta(x)$
- Minimize wrt  $\theta$  the objective  $\frac{1}{n} \sum_{1 \leq i \leq n} \|x_i - x_\theta(y_i)\|_1$

Note:

- Almost impossible to get real training data
- “Semi-synthetic” training data constructed using “ISP inversion” (Brooks et al., 2019) with a realistic noise model

# Optimization: unrolled iterative algorithm

$$\min_{\mathbf{x}, \mathbf{p}} \frac{1}{2} \|\mathbf{y} - U_{\mathbf{p}} \mathbf{x}\|^2 + \lambda \phi_{\theta}(\mathbf{x})$$



$$\min_{\mathbf{x}, \mathbf{p}, \mathbf{z}} E_{\mu}(\mathbf{x}, \mathbf{z}, \mathbf{p}) = \frac{1}{2} \|\mathbf{y} - U_{\mathbf{p}} \mathbf{z}\|^2 + \frac{\mu}{2} \|\mathbf{z} - \mathbf{x}\|^2 + \lambda \phi_{\theta}(\mathbf{x})$$

Quadratic penalty (aka HQS) method  
(three iterations)

➤  $\mathbf{z}^t \leftarrow \mathbf{z}^{t-1} - \eta_t [U_{\mathbf{p}^{t-1}}^\top (U_{\mathbf{p}^{t-1}} \mathbf{z}^{t-1} - \mathbf{y}) + \mu (\mathbf{z}^{t-1} - \mathbf{x}^{t-1})]$

One step of gradient descent (or a few)

➤  $\min_{\mathbf{p}_k} \frac{1}{2} \|\mathbf{y}_k - DBW_{\mathbf{p}_k} \mathbf{z}^t\|^2$

Gauss-Newton (aka Lucas-Kanade)

➤  $\mathbf{x}^t \leftarrow \arg \min_{\mathbf{x}} \frac{\mu_{t-1}}{2} \|\mathbf{z}^t - \mathbf{x}\|^2 + \lambda \phi_{\theta}(\mathbf{x})$

Proximal update

➤ Increment  $\mu$

# Optimization: unrolled iterative algorithm

$$\min_{\mathbf{x}, \mathbf{p}} \frac{1}{2} \|\mathbf{y} - U_{\mathbf{p}} \mathbf{x}\|^2 + \lambda \phi_{\theta}(\mathbf{x})$$



$$\min_{\mathbf{x}, \mathbf{p}, \mathbf{z}} E_{\mu}(\mathbf{x}, \mathbf{z}, \mathbf{p}) = \frac{1}{2} \|\mathbf{y} - U_{\mathbf{p}} \mathbf{z}\|^2 + \frac{\mu}{2} \|\mathbf{z} - \mathbf{x}\|^2 + \lambda \phi_{\theta}(\mathbf{x})$$

Quadratic penalty (aka HQS) method  
(three iterations)

➤  $\mathbf{z}^t \leftarrow \mathbf{z}^{t-1} - \eta_t [U_{\mathbf{p}^{t-1}}^\top (U_{\mathbf{p}^{t-1}} \mathbf{z}^{t-1} - \mathbf{y}) + \mu (\mathbf{z}^{t-1} - \mathbf{x}^{t-1})]$

One step of gradient descent (or a few)

➤  $\mathbf{p}_k^t \leftarrow \mathbf{p}_k^{t-1} - (\mathbf{J}_k^{t\top} \mathbf{J}_k^t)^{-1} \mathbf{J}_k^{t\top} \mathbf{r}_k^t$  (3 times)

Gauss-Newton (aka Lucas-Kanade)

➤  $\mathbf{x}^t \leftarrow \arg \min_{\mathbf{x}} \frac{\mu_{t-1}}{2} \|\mathbf{z}^t - \mathbf{x}\|^2 + \lambda \phi_{\theta}(\mathbf{x})$

Proximal update

➤ Increment  $\mu$

# Optimization: unrolled iterative algorithm

$$\min_{\mathbf{x}, \mathbf{p}} \frac{1}{2} \|\mathbf{y} - U_{\mathbf{p}} \mathbf{x}\|^2 + \lambda \phi_{\theta}(\mathbf{x})$$



$$\min_{\mathbf{x}, \mathbf{p}, \mathbf{z}} E_{\mu}(\mathbf{x}, \mathbf{z}, \mathbf{p}) = \frac{1}{2} \|\mathbf{y} - U_{\mathbf{p}} \mathbf{z}\|^2 + \frac{\mu}{2} \|\mathbf{z} - \mathbf{x}\|^2 + \lambda \phi_{\theta}(\mathbf{x})$$

Quadratic penalty (aka HQS) method  
(three iterations)

➤  $\mathbf{z}^t \leftarrow \mathbf{z}^{t-1} - \eta_t [U_{\mathbf{p}^{t-1}}^\top (U_{\mathbf{p}^{t-1}} \mathbf{z}^{t-1} - \mathbf{y}) + \mu (\mathbf{z}^{t-1} - \mathbf{x}^{t-1})]$

One step of gradient descent (or a few)

➤  $\mathbf{p}_k^t \leftarrow \mathbf{p}_k^{t-1} - (\mathbf{J}_k^{t\top} \mathbf{J}_k^t)^{-1} \mathbf{J}_k^{t\top} \mathbf{r}_k^t$  (3 times)

Gauss-Newton (aka Lucas-Kanade)

➤  $\mathbf{x}^t \leftarrow f_{\theta}(\mathbf{z}_t)$

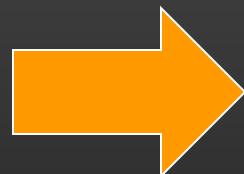
Plug-and-play approach  
(small residual U-net)

➤ Increment  $\mu$

# Example



Raw image burst (Lumix GX9)



High-quality picture



(Small crop of) Burst of raw pictures



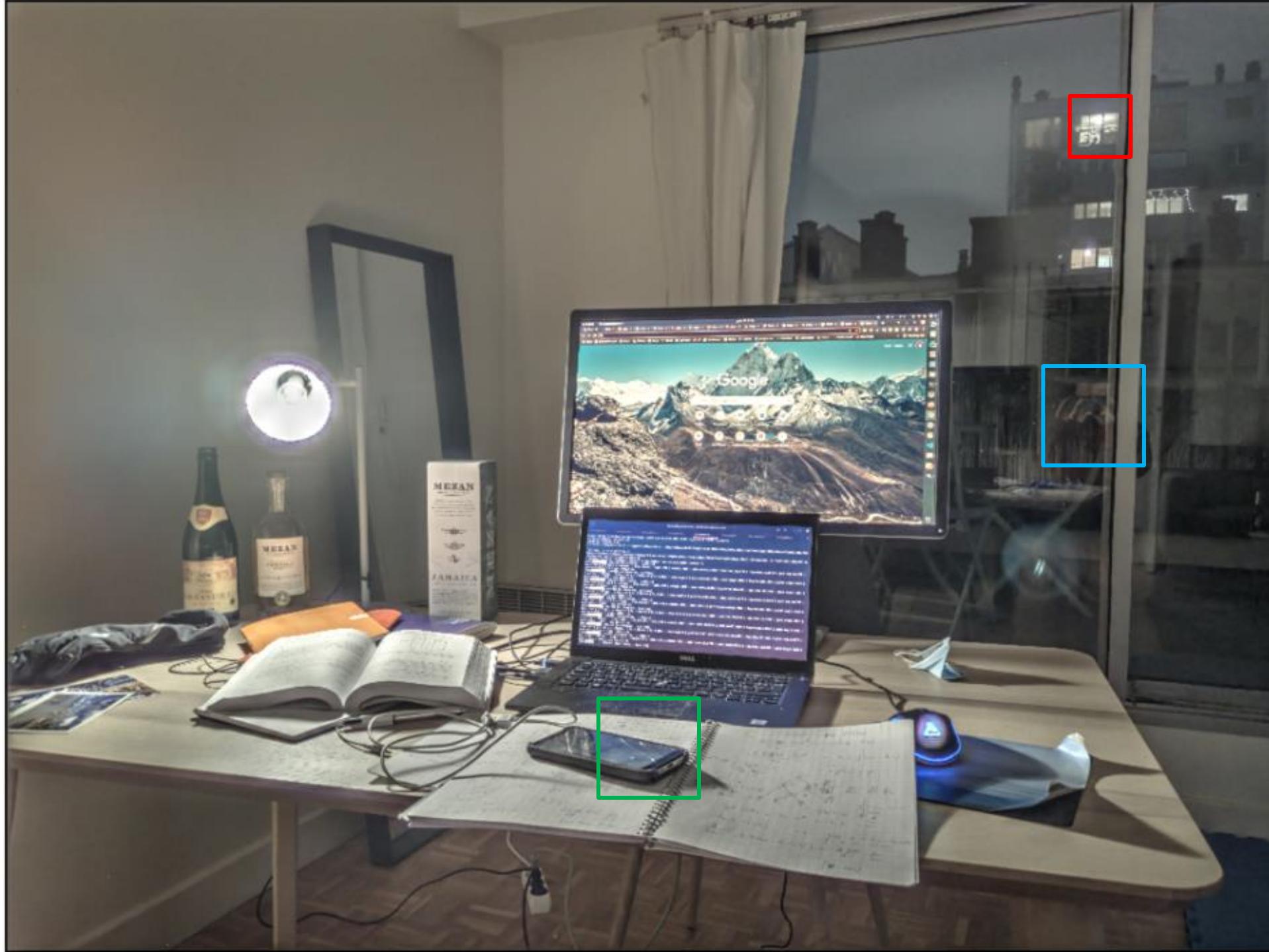
(Lecouat et al., ICCV'21)











Application: Thermal imaging - denoising + x4 super-resolution, 20 frames  
80x62 waveshare IR camera, less than 150Euro

