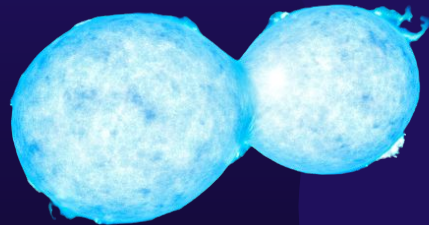


# Dynamical stability of B-type Overcontact Systems



## Internship Advisors

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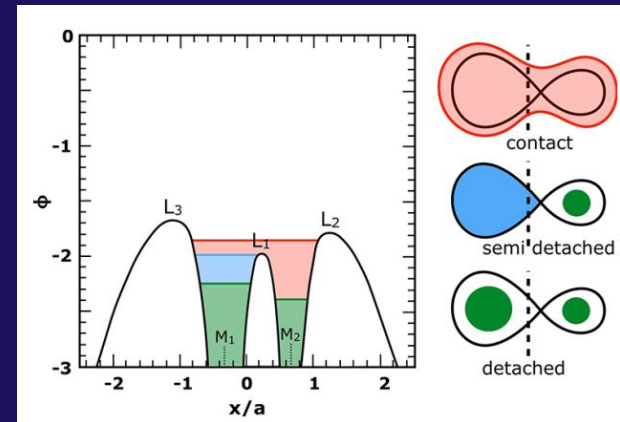
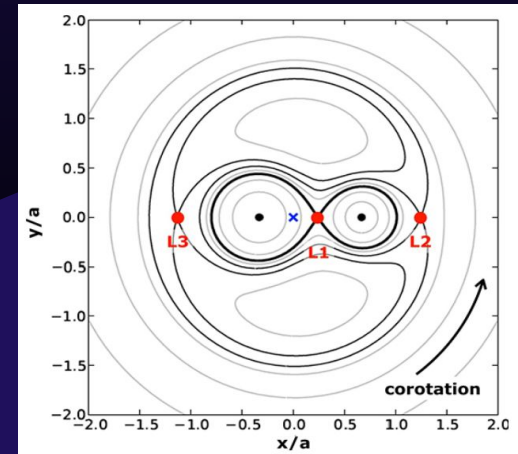
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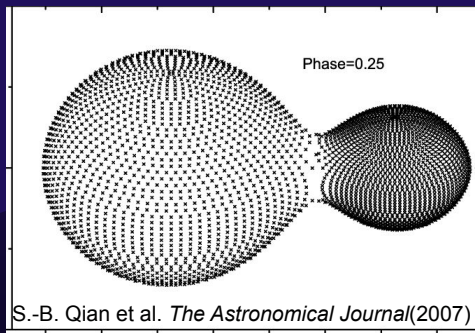
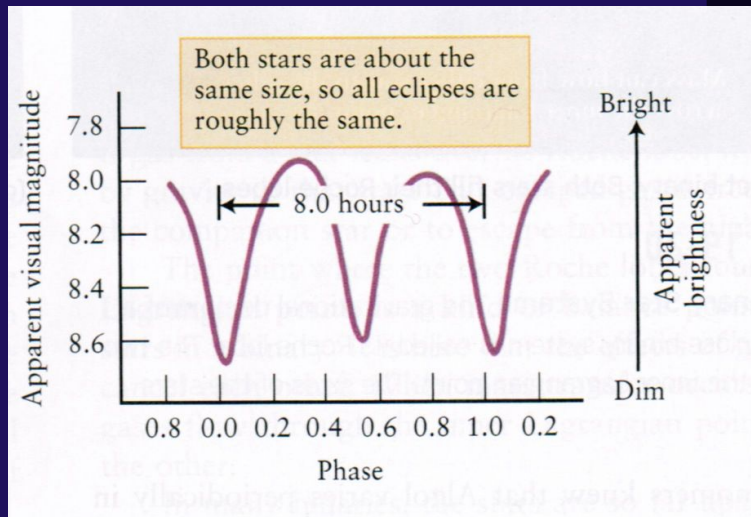
[jose.simmonds@up.ac.pa](mailto:jose.simmonds@up.ac.pa)

# BACKGROUND

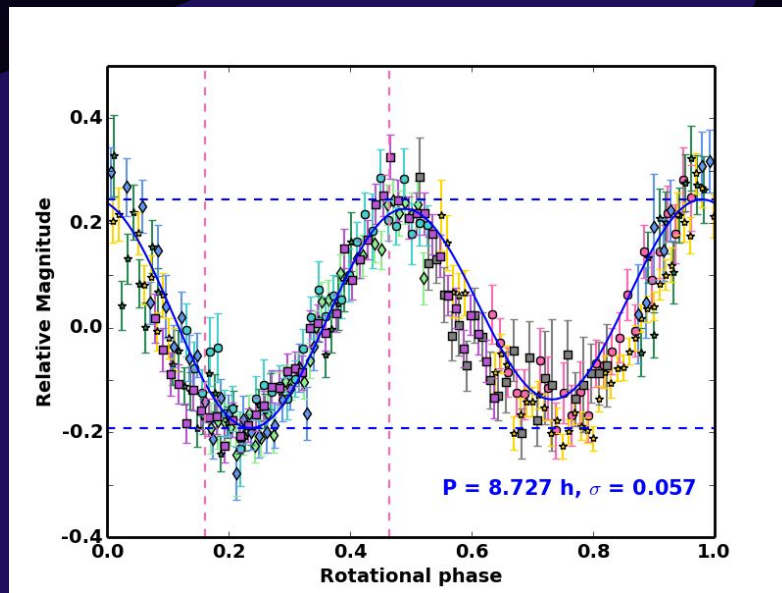
- **The Roche Model:** Region around a star in a binary system within which orbiting material is gravitationally bound to that star.
- **Tidal Effects:** Tidal forces can transfer angular momentum between the rotation of the individual stars and the orbit of the system.
- **Mass Transfer:** Mass transfer is the idea that material can move from one star to another.
- **The Overcontact Phase:** When both components of the binary systems are overfilling their Roche Lobel, are considered as an Overcontact binary.
- **Chemically Homogeneous Evolution in Massive Overcontact Binaries:** In overcontact systems they benefit from the very large angular momentum reservoir from the orbit to help maintain the high rotation rates.



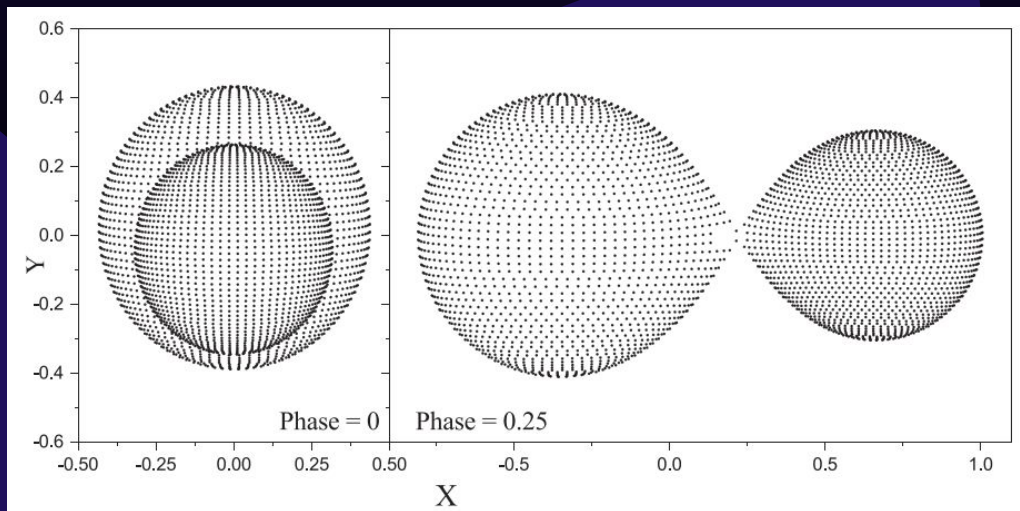
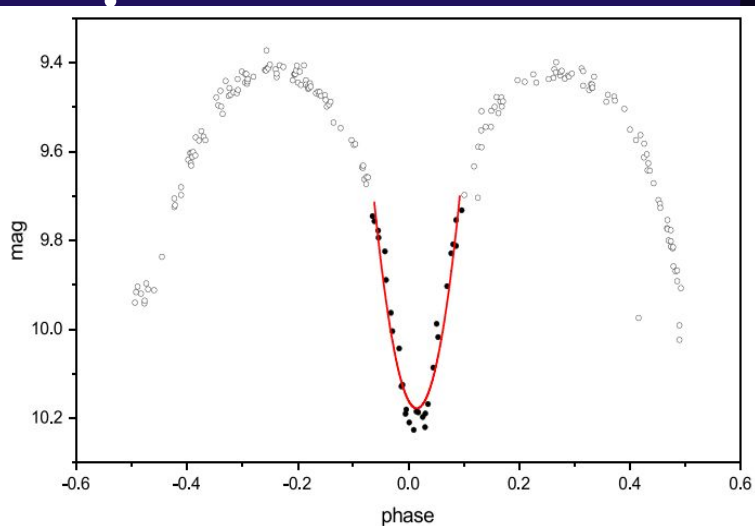
# Why this is important ?



# How to determine Period ?



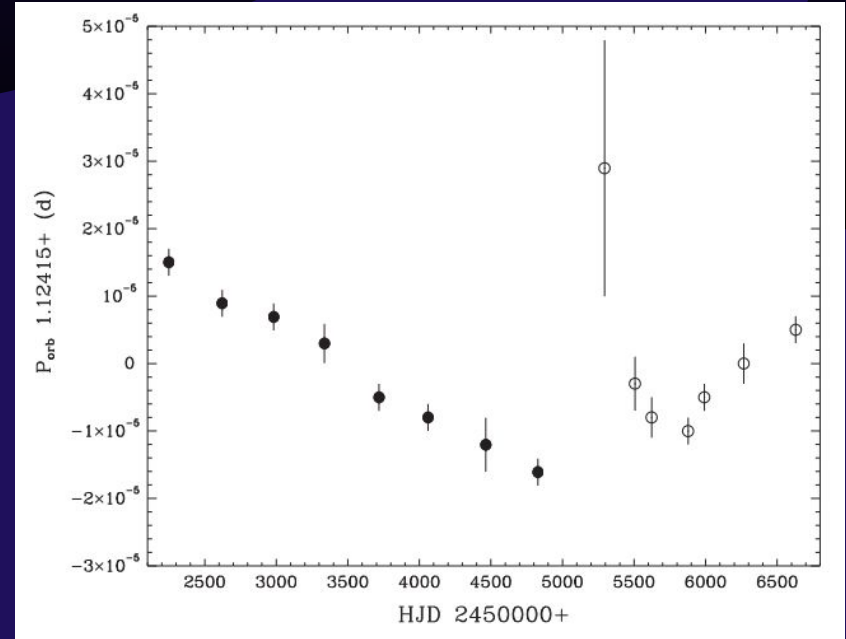
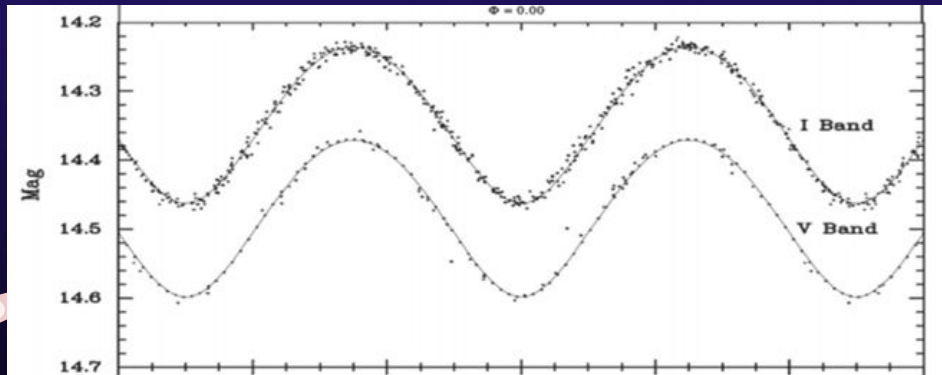
# V606 Cen



| Parameters   | Eccentric Orbit Case        | Circular Orbit Case         |
|--|-----------------------------|-----------------------------|
| Orbital Period, P (yr)                                       | 88.3(1.9)                   | 85.9(1.3)                   |
| Rate of the period change, $\dot{P}$ (day yr <sup>-1</sup> ) | -2.08(0.07)10 <sup>-7</sup> | -2.22(0.06)10 <sup>-7</sup> |
|  | Complete Data               | TESS Data                   |
| q (M2/M1)  | 0.54845(0.00074)            | 0.57434(0.0025)             |

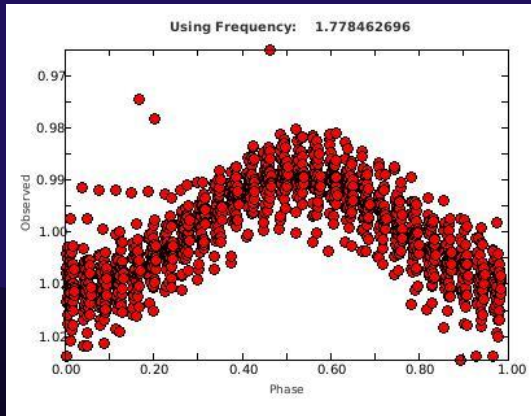
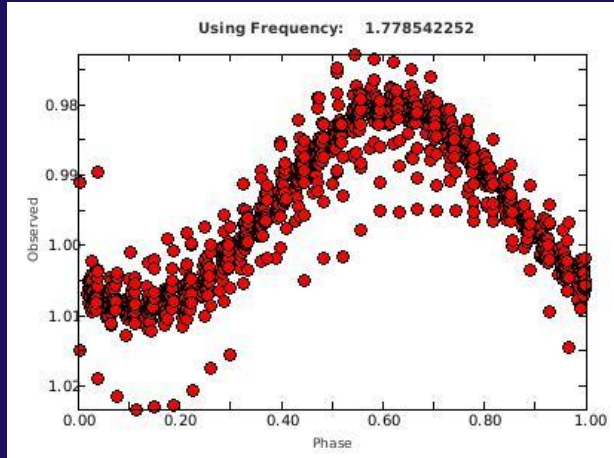
# VFTS352

| Parameters   |           |
|--|-----------|
| Orbital Period, P (yr)                                   | 1.124152  |
| Rate of the period change, $\dot{P}(\text{day yr}^{-1})$ | -0.050.10 |
| q (M2/M1)  | 0.98      |
| M2   | 25.1      |
| M1   | 25.6      |
| Period (days)  | 1.120     |



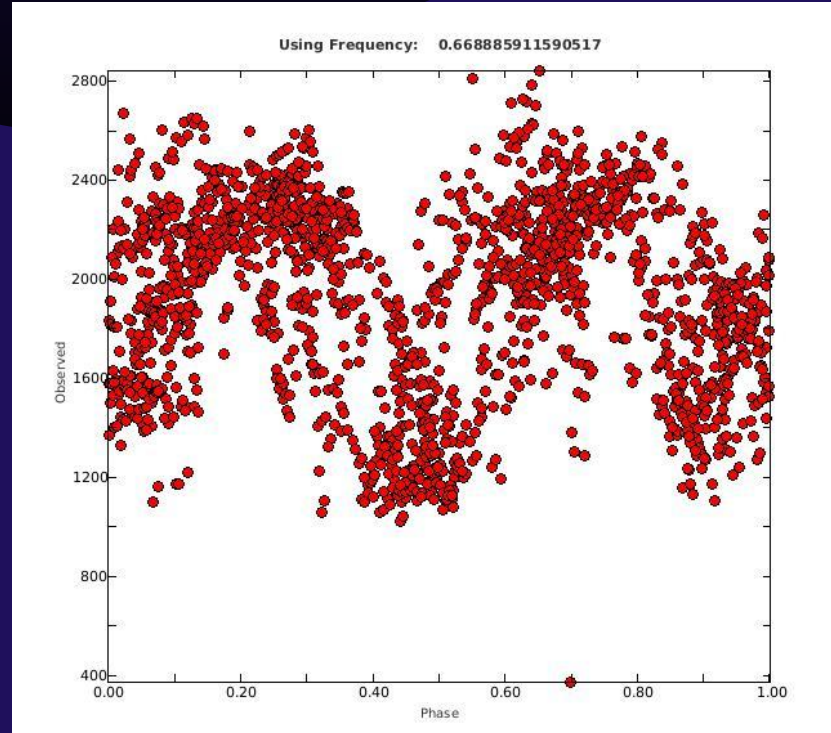
Almeida, L. A., Sana, H., Mink, S. E. D. 2015, The Astrophysical Journal, 812 (The Astrophysical Journal), 102

Using Period04 for plot the  
Light Curves of VFTS352



## Results

Using Period04 for plot the  
Light Curves of V606 Cen



Orbital period of 1.4950935 days. When do  
 $1/1.49 = 0.6711$ , approximately light curves are obtained

# Conclusions

- Several systems have period changes suggesting that they are evolving towards a lower mass ratio.
- That being said, these findings may suggest that the observed overcontact binaries are originating from systems with longer initial periods
- We have performed a period stability study of known O+O type overcontact systems. Using archival photometric data and the software package PERIOD04, we calculated the periods of the systems over a time span of tens of years. For each system in our sample, we determined the rate at which the period is changing via a linear regression through the period measurements of each data subset.



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