

Statistical Inference for inter-arrival times of extreme events in bursty time series

Abstract

In many complex systems studied in statistical physics, inter-arrival times between events such as solar flares, trades and neuron voltages follow a heavy-tailed distribution. The set of event times is fractal-like, being dense in some time windows and empty in others, a phenomenon which has been dubbed “bursty”.

This article proposes a new model for the *inter-exceedance times* of thresholds; the threshold exceedances are modeled via the standard Peaks-Over-Threshold (POT) method. For high thresholds and infinite-mean waiting times, we show that the times between threshold crossings are Mittag-Leffler distributed, and thus form a “fractional Poisson Process” which generalizes the standard Poisson Process of threshold exceedances. We provide graphical means of estimating model parameters and assessing model fit. Along the way, we apply our inference method to a real-world bursty time series, and show how the memory of the Mittag-Leffler distribution affects the predictive distribution for the time until the next extreme event.

Keywords: heavy tails renewal process extreme value theory peaks over threshold

1. Introduction

Time series displaying temporally inhomogeneous behaviour have received strong interest in the recent statistical physics literature (Barabási 2005; Oliveira and Barabási 2005; Vasquez et al. 2006; Vazquez et al. 2007; Omi and Shinomoto 2011; Min, Goh, and Vazquez 2011; Karsai et al. 2011; Bagrow and Brockmann 2013). They have been observed in the context of earthquakes, sunspots, neuronal activity and human communication (see Karsai et al. 2012; Vajna, Tóth, and Kertész 2013; Meerschaert and Stoev 2008 for a list of references). Such time series exhibit high activity in some ‘bursty’ intervals, which alternate with other, quiet intervals. Although several mechanisms are plausible explanations for bursty behaviour (most prominently self-exciting point process by Hawkes (1971)), there seems to be one salient feature which very typically indicates the departure from temporal homogeneity: a heavy-tailed distribution of waiting times (Vasquez et al. 2006; Karsai et al. 2012; Vajna, Tóth, and Kertész 2013). As we show below in simulations, a simple renewal process with heavy-tailed waiting times can capture this type of dynamics. For many systems, the renewal property is appropriate; a simple test of the absence of correlations in a succession

of waiting times can be undertaken by randomly reshuffling the waiting times (Karsai et al. 2012).

Often a magnitude, or mark can be assigned to each event in the renewal process, such as for earthquakes, solar flares or neuron voltages. The Peaks-Over-Threshold model (POT, see e.g. Coles 2001) applies a threshold to the magnitudes, and fits a Generalized Pareto distribution to the threshold exceedances. A commonly made assumption in POT models is that times between events are either fixed or light-tailed, and this entails that the threshold crossing times form a Poisson process (Hsing, Hüsler, and Leadbetter 1988). Then as one increases the threshold and thus decreases the threshold crossing probability p , the Poisson process is rarefied, i.e. its intensity decreases *linearly* with p (see e.g. Beirlant et al. 2006).

As will be shown below, in the heavy-tailed waiting time scenario threshold crossing times form a *fractional Poisson process* (Laskin 2003; Meerschaert, Nane, and Vellaisamy 2011), which is a renewal process with Mittag-Leffler distributed waiting times. The family of Mittag-Leffler distributions nests the exponential distribution (Haubold, Mathai, and Saxena 2011), and hence the fractional Poisson process generalizes the standard Poisson process. Again as the threshold size increases and the threshold crossing probability p decreases, the fractional Poisson process is rarefied: The scale parameter of the Mittag-Leffler inter-arrival times of threshold crossing times increases, but *superlinearly*; see the Theorem below.

Maxima of events which occur according to a renewal process with heavy-tailed waiting times have been studied under the names “Continuous Time Random Maxima process” (CTRM) (Benson, Schumer, and Meerschaert 2007; Meerschaert and Stoev 2008; Hees and Scheffler 2016, 2017), “Max-Renewal process” (Silvestrov 2002; Silvestrov and Teugels 2004; Basrak and Špoljarić 2015), and “Shock process” (Esary and Marshall 1973; Shanthikumar and Sumita 1983, 1984, 1985; Anderson 1987; Gut and Hüsler 1999). The existing literature focuses on probabilistic results surrounding these models. In this work, however, we introduce a method of inference for this type of model, which is seemingly not available in the literature.

We review the marked renewal process in Section 2, and derive a scaling limit theorem for inter-exceedance times in Section 3. We give a statistical procedure to estimate model parameters via stability plots in Section 5, but to set the stage we first discuss inference for the Mittag-Leffler distribution in Section 4. In Section 6, we discuss the memory property of the Mittag-Leffler distribution, and how it affects the predictive distribution for the time until the next threshold crossing event. A simulation study of the effectiveness of our statistical procedure is given in Section 7. In Section 8 we apply our method to a real data set. Finally we close with a discussion and conclusion in Section 8. For all statistical computations we have used R (R Core Team 2018). All code and data used for the analysis in this article has been organized into an R package CTRE (<https://github.com/strakaps/CTRE>). The source code for the figures generated in this manuscript is available online at <https://github.com/strakaps/bursty-POT>.

2. Continuous Time Random Exceedances (CTRE)

As a model for extreme observations, we use a Marked Renewal Process (MRP):

Definition (MRP): Let $(W, J), (W_1, J_1), (W_2, J_2), \dots$ be i.i.d. pairs of random variables, where the $W_k > 0$ are interpreted as the *waiting times* and $J_k \in [x_L, x_R]$ as the *event magnitudes* ($x_L \in [-\infty, +\infty), x_R \in (-\infty, +\infty]$). If W and J are independent, the Marked Renewal Process is said to be *uncoupled*.

We assume that the k -th magnitude J_k occurs at time $T_k = W_1 + \dots + W_k$. Based on an MRP, we define the Continuous Time Random Exceedance model (CTRE) as follows:

Definition (CTRE): Given a threshold $\ell \in (x_L, x_R)$, consider the stopping time

$$\tau(\ell) := \min\{k : J_k > \ell\}, \quad \ell \in (x_L, x_R).$$

Define the pair of random variables $(X(\ell), T(\ell))$ via

$$X(\ell) = J_{\tau(\ell)} - \ell, \quad T(\ell) = \sum_{k=1}^{\tau(\ell)} W_k.$$

By restarting the MRP at $\tau(\ell)$, inductively define the two i.i.d. sequences $T(\ell, n)$ and $X(\ell, n)$, $n \in \mathbb{N}$, called the “inter-arrival times” and the “exceedances”, respectively. The pair sequence $(T(\ell, n), W(\ell, n))_{n \in \mathbb{N}}$ is called a Continuous Time Random Exceedance model (CTRE). If the underlying MRP is uncoupled, then the CTRE is also called uncoupled.

In this article, we restrict ourselves to the uncoupled case, where W and J are independent. Then the two sequences $X(\ell, n)_{n \in \mathbb{N}}$ and $T(\ell, n)_{n \in \mathbb{N}}$ are independent as well. To see why, note that $X(\ell)$ is, in distribution, simply equal to $J|J > \ell$, independent of any waiting time W_k .

Figure 1 shows a simulated dataset in the top panel, where W has a stable distribution with tail parameter $\beta = 0.8$ (and skewness 1 and location 0), and where J is from a standard Gumbel distribution. In the bottom panel, we plot a time series of solar flare intensities derived from a NASA dataset (Dennis et al. 1991) which we will later examine more closely (see Section 8). Clearly, the simulated data exhibit long intervals *without any* events, whereas in the real-world dataset events appear continuously. The threshold exceedances, however, appear to have visually similar statistical behaviour in both models. Observations below a threshold are commonly discarded in Extreme Value Theory (POT approach); likewise, the CTRE model interprets these observations as noise and discards them.

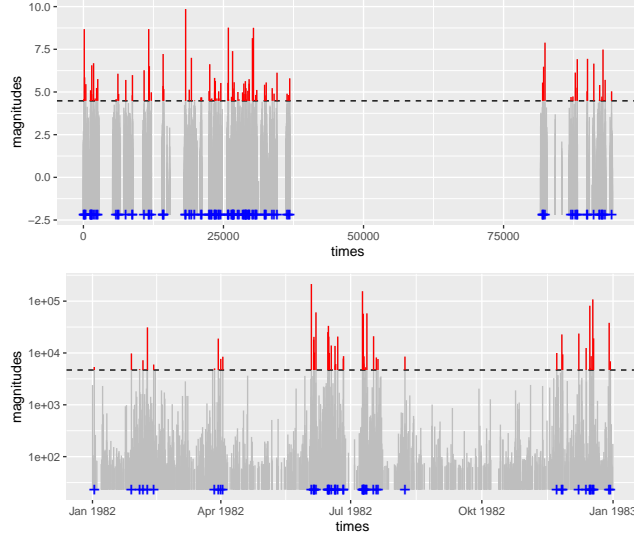


Figure 1: Exceedances (red) and times until Exceedance (durations between blue crosses) for a given threshold ℓ (dashed line).

3. Scaling limit of Exceedance Times

In this section we state and prove the key theorem, see below. For an accessible introduction to regular variation and stable limit theorems, we recommend the book by Mark M Meerschaert and Sikorskii (2012).

Theorem: Let the waiting times J_k be in the domain of attraction of a positively skewed sum-stable law with stability parameter $0 < \beta < 1$; more precisely,

$$\frac{W_1 + \dots + W_n}{b(n)} \xrightarrow{d} D, \quad n \rightarrow \infty \quad (1)$$

for a function $b(n)$ which is regularly varying at ∞ with parameter $1/\beta$, and where $\mathbf{E}[\exp(-sD)] = \exp(-s^\beta)$. Write $p := \mathbf{P}(J > \ell)$. Then the weak convergence

$$\frac{T(\ell)}{b(1/p)} \rightarrow W_\beta \quad \text{as } \ell \uparrow x_R$$

holds, where the Mittag-Leffler random variable W_β is defined on the positive real numbers via

$$\mathbf{E}[\exp(-sW_\beta)] = \frac{1}{1 + s^\beta}.$$

Proof of Theorem: We interpret the threshold crossing time $T(\ell)$ as the hitting time of the underlying CTRM (Continuous Time Random Maxima) or

“max-renewal process”, and then utilize a result by Meerschaert and Stoev (2008). The running maximum process is defined as

$$M(c) := J_1 \vee \dots \vee J_{\lfloor c \rfloor},$$

and since we assume that the J_k have a continuous distribution, there exist norming functions $a(c)$ and $d(c)$ such that

$$\mathbf{P} \left[\frac{M(c) - d(c)}{a(c)} \leq \ell^* \right] \longrightarrow F(\ell^*), \quad t \rightarrow \infty$$

where F is a generalized extreme value distribution, and ℓ^* is any value from the support of F . The CTRM process is then defined via

$$V(t) = M(N(t)), \quad t \geq 0$$

where $N(t)$ is the renewal process associated with the waiting times W_k :

$$N(t) = \max\{n : W_1 + \dots + W_n \leq t\}.$$

Now a key observation is that

$$T(\ell) = \inf\{t : V(t) > \ell\},$$

and that

$$T(\ell) > t \quad \text{if and only if} \quad V(t) \leq \ell.$$

By (Theorem 3.1, Meerschaert and Stoev 2008), we have the stochastic process convergence

$$\frac{V(ct) - d(\tilde{b}(c))}{a(\tilde{b}(c))} \xrightarrow{d} Y(t), \quad t > 0.$$

where $Y(t)$ is a time-changed (“subordinated”) extremal process, and where $\tilde{b}(c)$ is a regularly varying norming function which is *inverse* to $b(c)$, in the sense that $b(\tilde{b}(c)) \sim c \sim \tilde{b}(b(c))$.

Without loss of generality, we choose ℓ^* such that $F(\ell^*) = 1/e$, and let $\ell = a(\tilde{b}(c))\ell^* + d(\tilde{b}(c))$. We may then calculate

$$\mathbf{P} \left[\frac{T(\ell)}{b(1/p)} > t \right] = \mathbf{P}[T(\ell) > b(1/p)t] = \mathbf{P}[V(ct) \leq \ell]$$

where we have substituted $c = b(1/p)$. Moreover

$$\mathbf{P}[V(ct) \leq \ell] = \mathbf{P} \left[\frac{V(ct) - d(\tilde{b}(c))}{a(\tilde{b}(c))} \leq \frac{\ell - d(\tilde{b}(c))}{a(\tilde{b}(c))} \right] \longrightarrow \mathbf{P}[Y(t) \leq \ell^*]$$

Defining the hitting time of level ℓ^* by $Y(t)$ as $\xi_{\ell^*} := \inf\{t : Y(t) > \ell^*\}$, we then have

$$\mathbf{P}[Y(t) \leq \ell^*] = \mathbf{P}[\xi_{\ell^*} > t] = \mathbf{P}[(-\log F(\ell^*))^{-1/\beta} X^{1/\beta} D > t]$$

by (Proposition 4.2, Meerschaert and Stoev 2008), where X is an exponential random variable with mean 1. Using (Theorem 19.1, Haubold, Mathai, and Saxena 2011), we see that $X^{1/\beta}D \sim \text{ML}(\beta, 1)$, concluding the proof. \square

For a scale parameter $\sigma > 0$, we write $\text{ML}(\beta, \sigma)$ for the distribution of σW_β . The Mittag-Leffler distribution with parameter $\beta \in (0, 1]$ is a heavy-tailed positive distribution for $\beta < 1$, with infinite mean. However, as $\beta \uparrow 1$, $\text{ML}(\beta, \sigma)$ converges weakly to the exponential distribution $\text{Exp}(\sigma)$ with mean σ . This means that although its moments are all infinite, the Mittag-Leffler distribution may (if β is close to 1) be indistinguishable from the exponential distribution, for the purposes of applied statistics. For a detailed reference on the Mittag-Leffler distribution, see e.g. Haubold, Mathai, and Saxena (2011), and for algorithms, see e.g. the R package `MittagLeffler` (Gill and Straka 2017).

Remark: If $\beta = 1$, the result of the Theorem above is standard, see e.g. Equation (2.2) in Gut and Hüsler (1999). In Anderson (1987) a similar result is shown with a different choice of scaling constant.

4. Model Choice and inference for the Mittag-Leffler distribution

The classical POT approach is based on the fact that exceedances above a high threshold are asymptotically GPD distributed, hence a GPD distribution is fitted to the exceedances above a high enough threshold. Due to the Theorem in the last Section we know that for high thresholds the times between exceedances are asymptotically Mittag-Leffler distributed in case of heavy-tailed waiting times between the observations. Hence, before we talk about inference for the exceedances in the next section, we discuss inference for Mittag-Leffler distributions.

Historically, the first method proposed for the estimation of the Mittag-Leffler distribution parameters was the fractional moment estimator by Kozubowski (2001). Unlike the first moments, the fractional moments of order p for $p < \beta$ exist and are tractable. One drawback of this method is that constant priors for the tail parameter are needed for the calculation of the estimates. Cahoy, Uchaikin, and Woyczynski (2010) proposed a moment estimator of the log-transformed data, which does not require any prior. Furthermore, they performed simulation studies illustrating that the log-moment outperforms the fractional moment estimator with respect to bias and root mean squared error (RMSE).

Due to the form of the density function of a Mittag-Leffler distribution there exists no closed form for the Maximum Likelihood estimator (MLE). In the R Package `MittagLeffler` (Gill and Straka 2017), Maximum Likelihood estimation is implemented via numerical optimization. The MLE slightly outperforms the log-moment estimator regarding bias and RMSE for big enough sample sizes, but is extremely computational intensive. Figure 2 shows that both estimators perform well, even for small sample sizes.

Since the Mittag-Leffler distribution is heavy-tailed, many researchers would intuitively give the highest importance to the tail behaviour of the distribution.

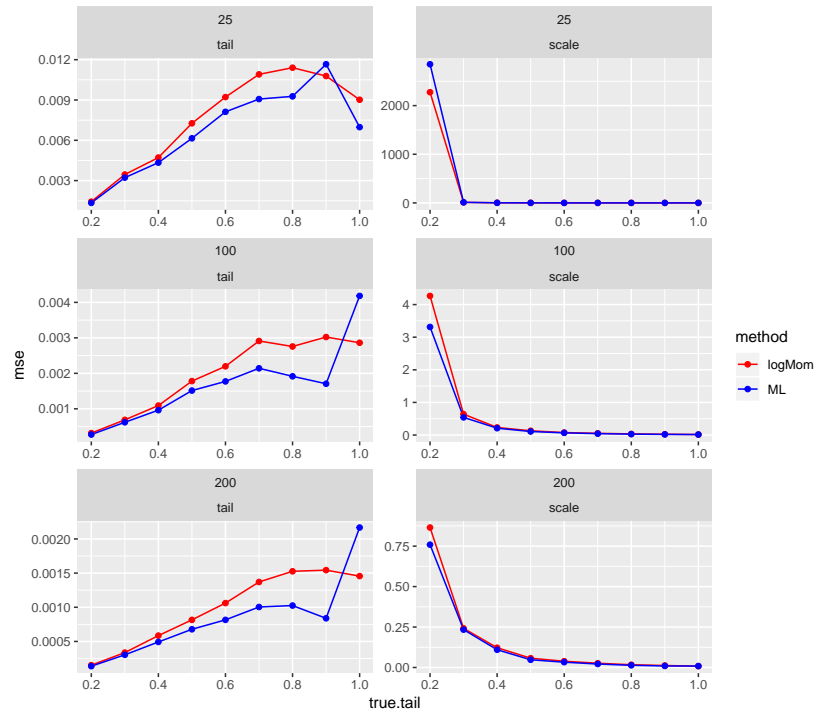


Figure 2: MSE for the estimation of tail (left column) and scale (right column) parameters via log-moment estimator and mle of a Mittag-Leffler sample with varying sample size $n=25, 100, 200$, varying tails on the x-axis and fixed scale equal to one.

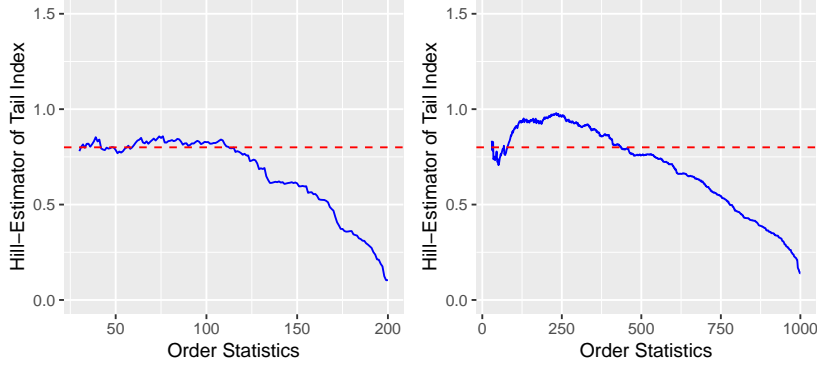


Figure 3: Hillplots for simulated Mittag-Leffler data with true tail 0.8 and sample size 200 (left) and 1000 (right), with number of upper order statistics k on which the hill estimator is based on the x-axis.

Of course, one can also use established tail exponent estimators for the estimation of the parameter β such as the Hill estimator (Hill 1975)

$$H_{k,n} = \frac{1}{k} \sum_{i=0}^{k-1} \log \frac{X_{n-i,n}}{X_{n-k,n}}. \quad (2)$$

However, these methods are statistically less efficient since they only use a portion of the information contained in the data, as mentioned by Kozubowski (2001). Moreover, the hill estimator requires a tuning parameter, and only performs reliably well for distributions close to Pareto (see Resnick 1997). Figure 3 shows Hill plots for Mittag-Leffler simulated data, with varying sample sizes and tails. To deduce the correct tail parameter estimates from these plots is virtually impossible. Furthermore the Hill estimator of course completely fails if the inter-arrival times are exponential distributed and not heavy-tailed.

Since the exponential distribution is nested in the Mittag-Leffler family of distributions, a Likelihood-ratio Test (LRT) seems to be an appropriate way to choose between a model with exponential and Mittag-Leffler inter-exceedance times. Although the two models are nested, the asymptotic distribution is not χ_1^2 -distributed, and Wilk's Theorem does not hold: under H_0 , the parameter β of Mittag-Leffler distribution is equal to 1, and hence lies on the boundary of the parameter space $(0, 1]$. Instead, a valid approach is a bootstrapped Likelihood-ratio test (see e.g. Davison, Hinkley, and others 1997). Figure 4 displays the (simulated) power for the bootstrapped LRT for Mittag-Leffler distributions with varying tail parameters based on 1000 simulation runs. As expected, the power decreases for tail parameters close to one, since the Mittag-Leffler distribution converges as $\beta \uparrow 1$ to an exponential distribution; it becomes hard to differentiate a Mittag-Leffler distribution from an exponential.

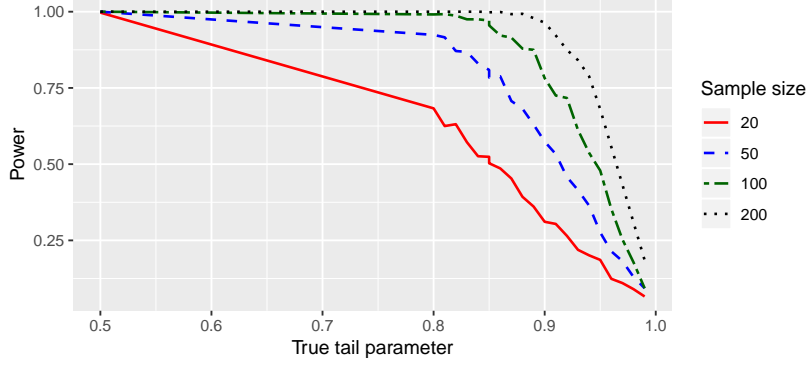


Figure 4: Power for bootstrapped LRT for different sample sizes , varying tails and scale parameter equal to 1.

5. Inference on Exceedance times

The Theorem in Section 3 implies that for a high threshold ℓ we may approximate the distribution of $T(\ell)$ with an $\text{ML}(\beta, b(1/p))$ distribution, where the function $b(c)$ varies regularly at ∞ with parameter $1/\beta$. Building on the POT (Peaks-Over-Threshold) method, we propose the following estimation procedure for the distribution of inter-exceedance time $T(\ell)$:

1. For a range of thresholds ℓ near the largest order statistics, extract datasets of exceedance times $\{T(\ell, i)\}_i$.
2. For each choice of threshold ℓ , fit a Mittag-Leffler distribution to the resulting dataset $\{T(\ell, i)\}_i$. This results in the estimates $\{\hat{\beta}(\ell)\}_\ell$ and $\{\hat{\sigma}(\ell)\}_\ell$.
3. Plot ℓ vs. $\hat{\beta}(\ell)$. As ℓ increases towards x_R , $\hat{\beta}(\ell)$ stabilizes around a constant $\hat{\beta}$. Use $\hat{\beta}$ as an estimate for the tail parameter β of the Mittag-Leffler distribution of exceedance times.
4. Approximate $p \approx |\{k : J_k > \ell\}|/n$. Recall that $b(c)$ is regularly varying with parameter $1/\beta$, and hence has the representation $b(c) = L(c)c^{1/\beta}$ for some slowly varying function $L(c)$. Assuming that the variation of $L(c)$ is negligible, we hence plot ℓ vs. $p^{1/\hat{\beta}}\hat{\sigma}(\ell)$. Again as ℓ increases towards x_R , $p^{1/\hat{\beta}}\hat{\sigma}(\ell)$ is expected to stabilize around a constant $\hat{\sigma}_0$. We then use $p^{-1/\hat{\beta}}\hat{\sigma}_0$ as an estimate of the scale parameter of the Mittag-Leffler distribution of exceedance times for the level ℓ .

The above approach benefits from the following practical adjustments (compare with Figure 10):

- We choose ℓ from the order statistics, i.e. ℓ is the k -th largest of the observations X_j , where k runs from k_{\min} , $k_{\min} + 1, \dots, k_{\max}$. The datasets are then of length $k - 1$.

- We use k rather than ℓ for the horizontal axis of our plots.
- In Step 4, rather than plotting $p^{1/\hat{\beta}}\hat{\sigma}(\ell)$ we plot $k^{1/\hat{\beta}}\hat{\sigma}(\ell)$. This changes $\hat{\sigma}_0$ by the multiplicative constant $n^{1/\hat{\beta}}$, but has the advantage that $\hat{\sigma}_0$ does not change if one pre-processes the data by removing all observations below a certain threshold.

The estimates $\hat{\beta}$ and $\hat{\sigma}_0$ give an estimate of the distribution of exceedance times, dependent on the threshold ℓ :

$$T(\ell) \sim \text{ML}(\hat{\beta}, k^{-1/\hat{\beta}}\hat{\sigma}_0),$$

where ℓ equals the k -th order statistic. For quick estimates of the Mittag-Leffler parameters we have used the method of log-transformed moments (Cahoy 2013). We have verified the validity of our estimation algorithm via simulations, see Section 6.

6. Simulation study

To test our inference method via stability plots, we have simulated independent waiting time and magnitude pairs (W_k, J_k) for waiting times that follow

- (i) a stable distribution,
- (ii) a Pareto distribution and
- (iii) an Exponential distribution.

In order to have exact analytical values available for β and σ_0 , a distribution for W_k needs to be chosen for which $b(n)$ from (1) is known. For (i) we choose $W_k \stackrel{d}{=} D$, where D is as in (1), then due to the stability property we have the *equality* of distribution $W_1 + \dots + W_n \stackrel{d}{=} b(n)D$, for $b(n) = n^{1/\beta}$. Using the parametrisation of Samorodnitsky and Taqqu (1994), a few lines of calculation (see e.g. the vignette on parametrisation in Gill and Straka 2017) show that D must have the stable distribution $S_\beta(\cos(\pi\beta/2)^{1/\beta}, +1, 0)$, which is implemented in the R package `stabledist` by Wuertz, Maechler, and members. (2016). By the Theorem, the distribution of $T(\ell)$ is approximately

$$\text{ML}(\beta, p^{-1/\beta}) = \text{ML}(\beta, k^{-1/\beta}n^{1/\beta}),$$

which means $\sigma_0 = n^{1/\beta}$. In the Pareto example we choose $P(W > t) = Ct^{-\beta}$ with $C = (1/\Gamma(1-\beta))^{1/\beta}$. We have chosen $\beta = 0.8$ in the stable as well as in the Pareto case. As third example we choose Exponential distributed waiting times with a rate of 1. Figure 5 shows the stability plots for the estimation of the tail parameter. Therefore, $\hat{\beta}(k)$ vs. k is plotted for each of the $m = 100$ simulation runs (grey thin lines). In the left column one can see the estimates based on the log-moment estimator, the stability plots in the middle are based on the maximum-likelihood estimator and on the right hand on the hill estimator. Recall that k is the index of the order statistics of J_k at which the threshold

ℓ is placed. Note, that the hill estimator is here based on the waiting times, not on the inter-exceedance times, hence the estimation is based on much more data. If one would base the hill-estimation only on the inter-exceedance times, the sample size would be too small and it would be impossible to deduce an estimate for the tail parameter from the plots (compare Figure 3). The k on the x -axis is in the plots for the hill estimator the k which defines on which orders stats the hill estimator is based on (compare (2)). The bigger sample size results in the middle in a lower bias. But, one can also see, that there is much more variance between the stability plots for the different simulation runs in case of the hill estimator compared to the other two estimators. The black line is the mean value of the point estimators for each k . However, if one wants to get a single estimate via such a stability plot, it would be much more difficult to deduce it from a hill plot compared to the other two estimators. In case that one has only thresholded data and hence a much smaller sample size, the hill estimator gets useless (see Section 4). Another advantage of the other estimators compared to the hill estimator is, that in the case one has exponential waiting times, the log-moment as well as the mle can detect this by getting an estimate round about 1. The hill estimator is just a tail parameter estimator for Pareto like tails and hence of course fails in case of Exponential inter-exceedance times. Figure 6 shows the stability plots for the different estimators for the scale parameter. Here of course only for the log-moment (on the right hand) and the mle (on the left hand), since the hill estimator is just a tail parameter estimate. Here again both estimators show good performance. The stability plots in the mle case are a little bit closer to the true value, therefore, the mle is much more computational intensive.

7. Data example

We now want to apply the proposed method to a real data example, the solar flare data which was already mentioned in Section 1 and can be seen in Figure 1. The data were extracted from the “complete Hard X Ray Burst Spectrometer event list”, a comprehensive reference for all measurements of the Hard X Ray Burst Spectrometer on NASA’s Solar Maximum Mission from the time of launch on Feb 14, 1980 to the end of the mission in Dec 1989. 12,776 events were detected, with the “vast majority being solar flares”. The list includes the start time, peak time, duration, and peak rate of each event. We have used “start time” as the variable for event times, and “peak rate” as the variable for event magnitudes.

Before we apply the POT approach described in Section 5 to the solar flare data, we first have to check if all model assumptions are full-filled. The CTRE model is based on three main assumptions, which are repeated below. For each assumption, we suggest one means of checking if it holds:

i.i.d.: After removing the “noise observations” below the smallest threshold ℓ_0 , the pair sequence $(T(\ell_0, i), X(\ell_0, i))$ is i.i.d. An indication if this is true

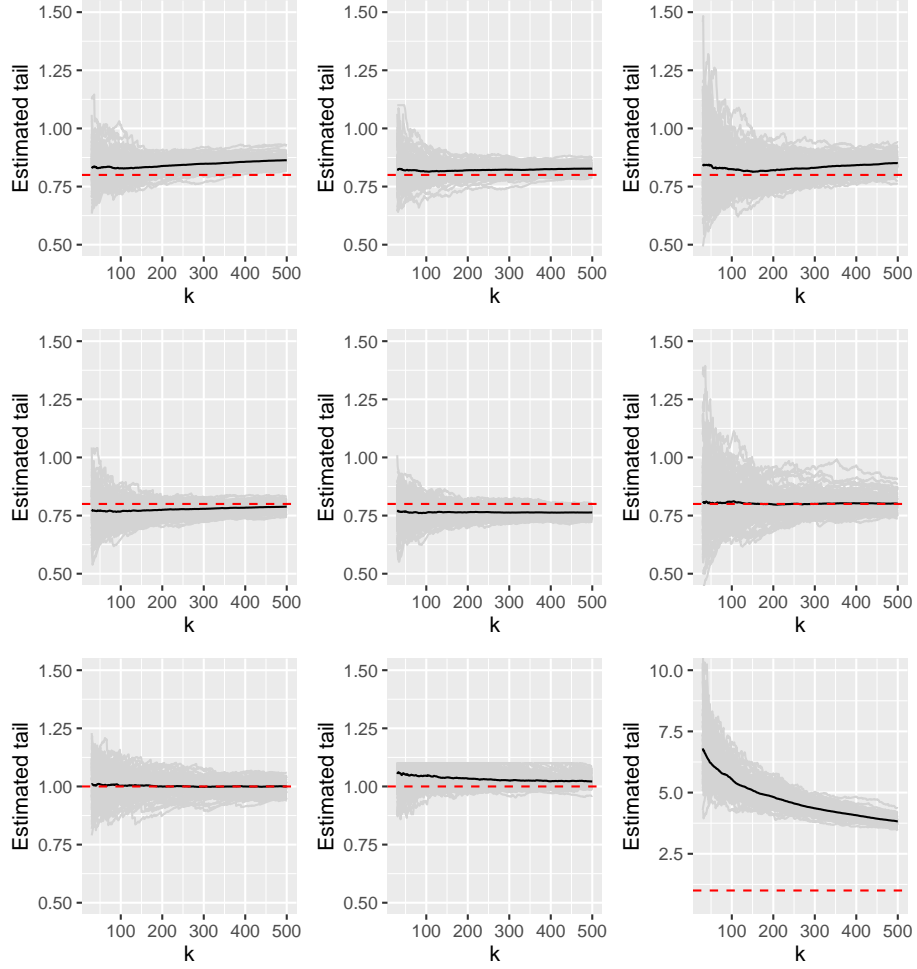


Figure 5: Stability Plots for $m=100$ simulation runs for stable distributed waiting times with a tail parameter of 0.8 (top row), Pareto distributed waiting times with tail parameter 0.8 (middle row) and Exponential distributed waiting times (lower row). Left column: log-moment estimator, middle column: MLE, right column: Hill estimator. The grey thin lines are the stability plots for the different simulation runs and the dark lines are their means. The red dotted line shows the true tail parameter.

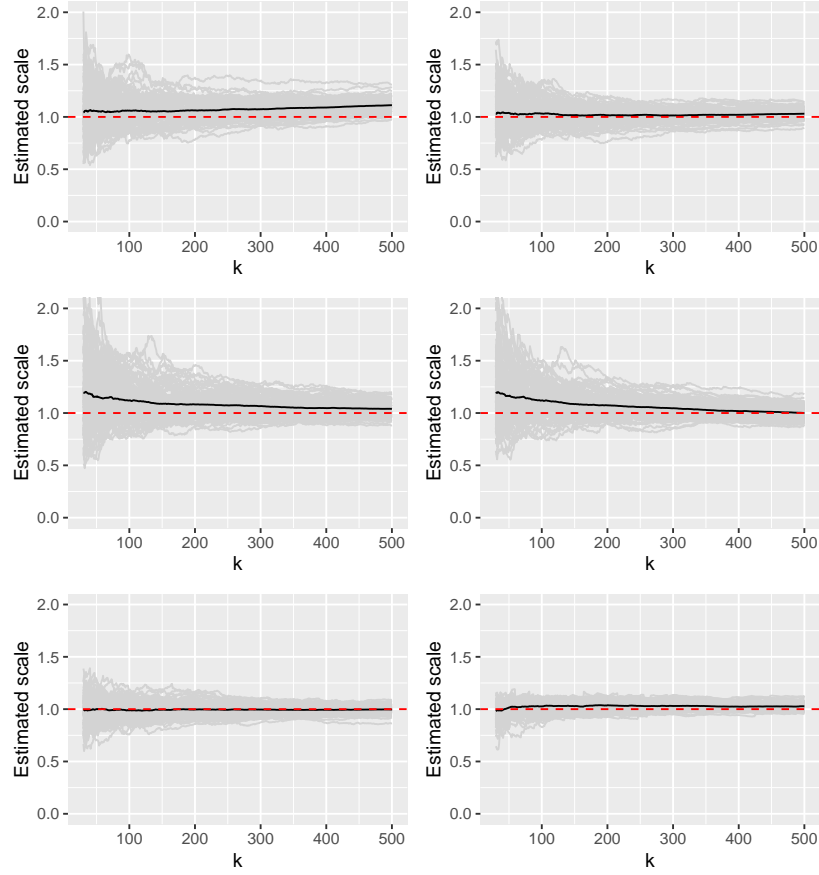


Figure 6: Stability Plots for $m=100$ simulation runs for stable distributed waiting times with a tail parameter of 0.8 (top row), Pareto distributed waiting times with tail parameter 0.8 (middle row) and Exponential distributed waiting times (lower row). Left column: log-moment estimator, right column: MLE. The grey thin lines are the stability plots for the different simulation runs and the dark lines are their means. The red dotted line shows the true tail parameter.

is given by an auto-correlation plot for the logarithms (to ensure finite moments) of the two time series.

Uncoupled: Each $T(\ell, i)$ is independent of each $X(\ell, i)$. We propose an empirical copula plot to check for any dependence.

ML(β, σ) distribution of $T(\ell, i)$: Apply a cutoff at the lowest threshold ℓ_0 , extract the threshold crossing times, and create a QQ Plot for the Mittag-Leffler distribution. Use a log-moment estimate of the tail parameter for the theoretical / population quantiles of the plot.

Figures 7, 8 and 9 show the diagnostic plots for a minimum threshold chosen at the 200th order statistic. There is some residual autocorrelation for the sequence of threshold exceedance times that is not accounted for by the CTRE model.

Figure 10 shows the stability plots for the solar flare data, on the left for the tail parameter and on the right for the scale parameter. Dotted lines show 95% confidence intervals, which are derived from the asymptotic normality of the log-moments estimators (Cahoy 2013) and the δ -method [Gill and Straka (2017)], dashed lines show the actual values of β resp. σ_0 . The stability plot for the tail stabilizes nicely round about 0.85 (dotted line), while it is a little bit harder to deduce an estimate for the scale parameter. However, the stability plot for the scale parameter seems to stabilize round about 3×10^7 .

The fit with a Mittag-Leffler distribution ($\beta = 0.85$) is good (see Figure 9), though there are signs that the power-law tail tapers off for very large inter-threshold crossing times. There is no apparent dependence between threshold exceedance times and event magnitudes seen in the copula plot (see Figure 8). We also conduct a bootstrapped LRT for the null hypothesis of exponential distributed inter-arrival times and received a p -value of $p < 0.01$.

8. Predicting the time of the next threshold crossing

According to Figure 10, for a threshold ℓ at the k -th order statistic, the fitted threshold exceedance time distribution is

$$T(\ell) \sim \text{ML}(\beta, k^{-1/\beta} \sigma_0),$$

where $\beta = 0.85$ and $\sigma_0 = 3.0 \times 10^7 \text{sec}$. Unlike the exponential distribution, the Mittag-Leffler distribution is not memoryless, and the probability density of the time t until the next threshold crossing will depend on the time t_0 elapsed since the last threshold crossing. This density equals

$$p(t|\beta, \sigma_0, \ell, t_0) = \frac{f(t + t_0|\beta, k^{-1/\beta} \sigma_0)}{\mathbf{P}[T_\ell > t_0]}$$

where $f(\cdot|\beta, k^{-1/\beta} \sigma_0)$ is the probability density of $\text{ML}(\beta, k^{-1/\beta} \sigma_0)$. The more time has passed without a threshold crossing, the more the probability distribution shifts towards larger values for the next crossing (see Figure 11, left panel).

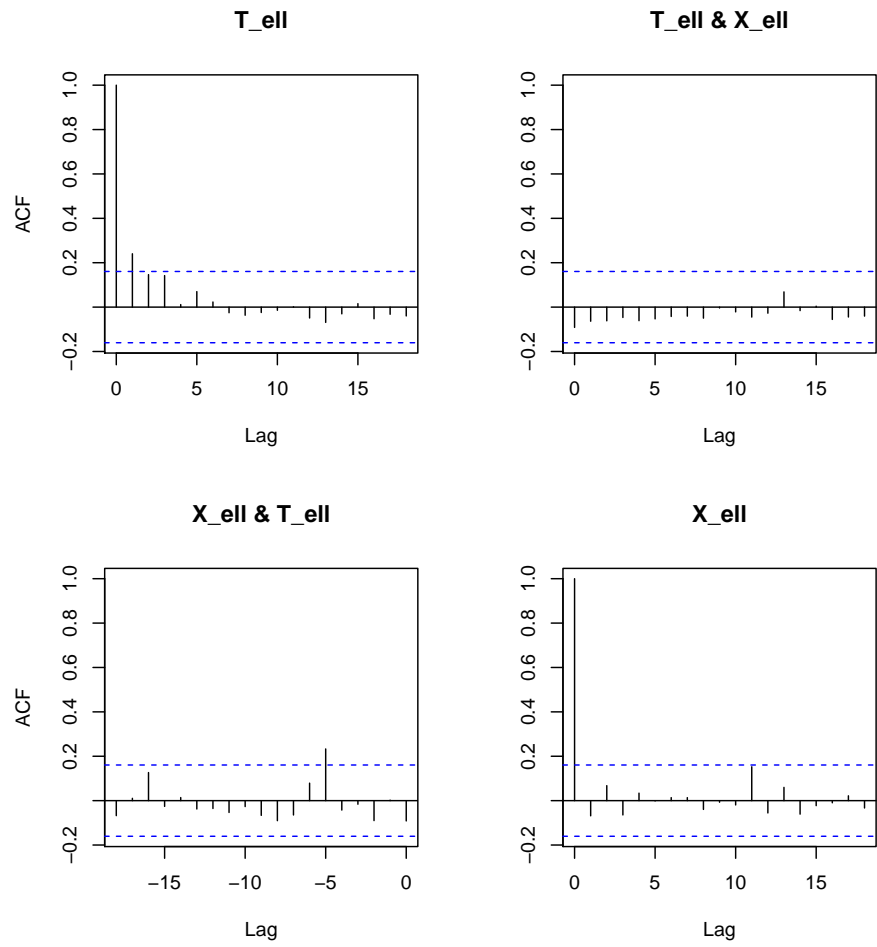


Figure 7: Diagnostic plots for the solar flare data: auto-correlation function.

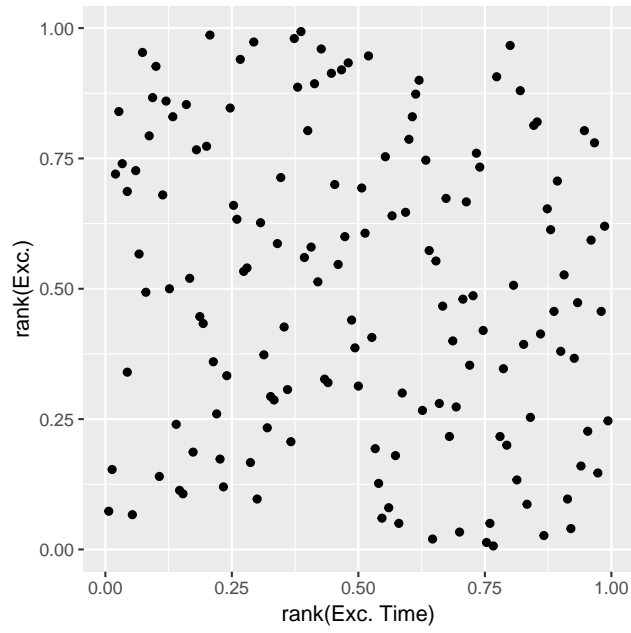


Figure 8: Diagnostic plots for the solar flare data: empirical copula and QQ Plot.

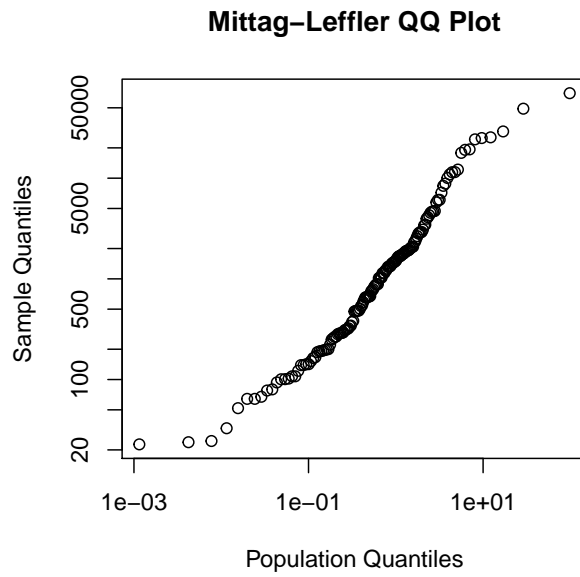


Figure 9: Diagnostic plots for the solar flare data: empirical copula and QQ Plot.

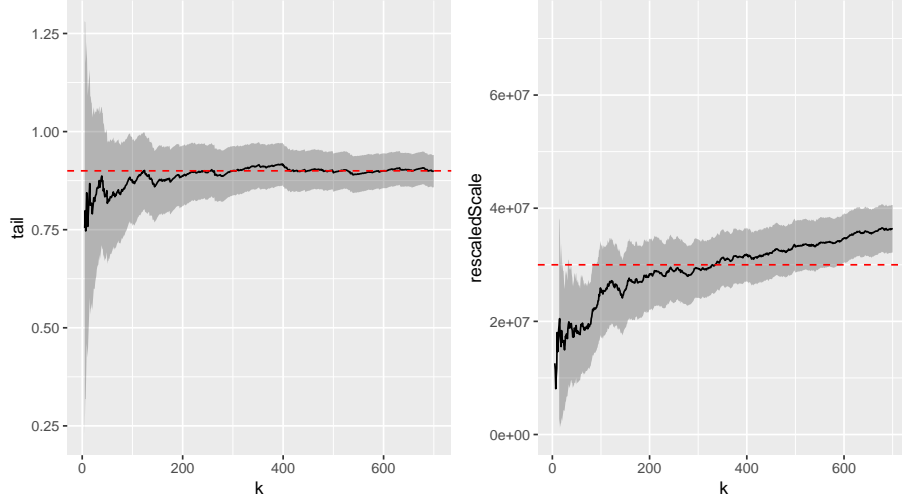


Figure 10: Stability plots for the tail and scale parameter of the Mittag-Leffler distribution of the Solar Flare dataset. Dotted horizontal lines are at $\beta = 0.85$ and $\sigma_0 = 3 \times 10^7$ seconds ≈ 0.95 years.

The hazard rate

$$h(t) = \frac{f(t|\beta, k^{-1/\beta}\sigma_0)}{\int_t^\infty f(\tau|\beta, k^{-1/\beta}\sigma_0) d\tau}$$

represents the risk of a threshold crossing per unit time, and is a decreasing function for the Mittag-Leffler distribution.

The closer β is to 1, the more the hazard rate mimics that of an exponential distribution (a constant function, see Figure 11, right panel).

9. Discussion & Conclusion

We have extended the POT (Peaks-over-Threshold) model, a mainstay of extreme value theory, to “bursty” time series, which have been studied intensively in statistical physics. Burstiness is characterized by power-law waiting times between events, and we have shown that the Mittag-Leffler distribution arises naturally as a scaling limit for the inter-exceedance times of high thresholds. Moreover, we have derived the following non-linear scaling behaviour: $\sigma \sim p^{-1/\beta}$, where σ is the scale parameter of the distribution of threshold exceedance times, p is the fraction of magnitudes above the threshold, and β the exponent of the power law. This “anomalous” scaling behaviour in the bursty setting entails two phenomena:

- i) a heavy tail of the inter-arrival time distribution of threshold crossings (long rests), and
- ii) a high propensity for more threshold crossing events immediately after each threshold crossing event (bursts).

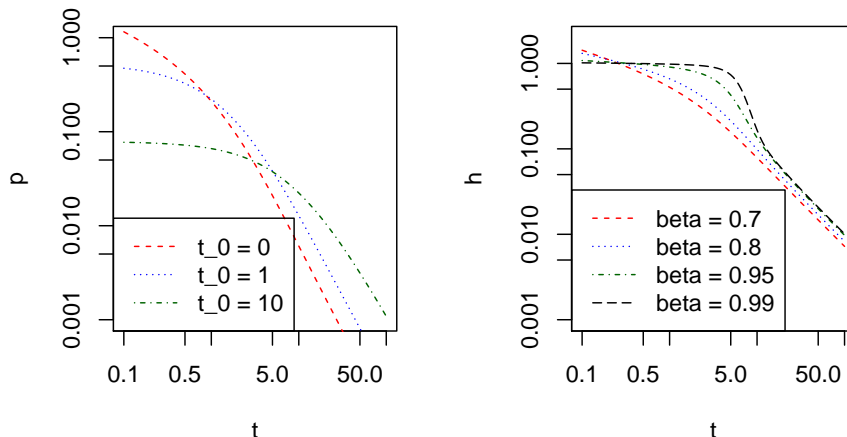


Figure 11: Left: Conditional distribution of time until next threshold crossing, depending on elapsed time t_0 since last crossing ($\beta = 0.8$, $\sigma_0 = 1$). Right: Hazard rate depending on tail parameter β .

The Mittag-Leffler distribution captures both phenomena, due to its heavy tail as well as its stretched exponential (peaked) asymptotics for small times. It generalizes the exponential distribution, and in the solar flare data example, this generalization is warranted, because the likelihood-ratio test is strongly significant.

When we introduced the CTRE model, we assumed that all events are i.i.d. This assumption is likely sufficient but not necessary for our limit theorem to hold. Moreover, any data below a (minimum) threshold ℓ_0 is discarded for CTREs, and hence need not satisfy the i.i.d. assumption. For the purposes of statistical inference, we merely require that the inter-threshold-crossing times are i.i.d.

The bursty CTRE approach to model “non-Poissonian” threshold crossing times should be contrasted with the (now standard) approach of clusters of extremes, see e.g. Ferro and Segers (2003). In this approach, i.i.d. event sequences of magnitudes are generalized to stationary sequences of event magnitudes (subject to a mixing condition). The two approaches are fundamentally different: A clustering model assumes that each event belongs to one particular (latent) group of events. For bursts, however, the aim is to identify an underlying scale-free pattern in the event dynamics, which is often characteristic of complex systems. It is an interesting open problem to develop quality criteria, based e.g. on measures of surprise (Lee, Fan, and Sisson 2015), which guide an applied statistician in the choice between a clustering and a CTRE approach for a particular problem. Moreover, we believe it may be possible to unify the two

approaches by considering CTREs based on MRPs with a *stationary*, rather than i.i.d., sequence of magnitudes.

Finally, a purely scale-free pattern for event times may be too rigid as an assumption for some bursty time series, because often the heavy-tailed character of the inter-arrival time distribution does not hold at all time scales; rather, it applies at short and intermediate time scales, and is truncated (or tempered, reverting to an exponential distribution) at very long time scales (see e.g. Mark M. Meerschaert, Roy, and Shao 2012; and Aban, Meerschaert, and Panorska 2006). In such situations, a “tempered” Mittag-Leffler distribution may provide a more realistic fit, which we aim to introduce in follow-up work.

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