

# RAMAN DATA PROCESSING V6 QUICK MANUAL

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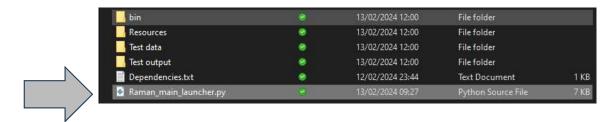


- 1. Install python >= 3.8.18
- 2. Install the following libraries:
  - Multiprocessing
  - Tendo
  - svs
  - os
  - csv
  - re
  - pillow>=10.2.0
  - numpy>=1.24.4
  - matplotlib>=3.7.4
  - numexpr>=2.8.6
  - scipy>=1.10.1
  - Imfit>=1.2.2
  - tkinter>=8.6

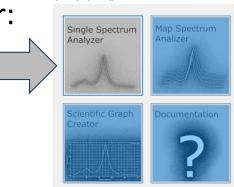


# Step 2

- Currently the software works for individual spectra taken in Horiba YobinYvon and B&WTech spectrometers.
- Data must be saved as .txt files.
- Run the main launcher:

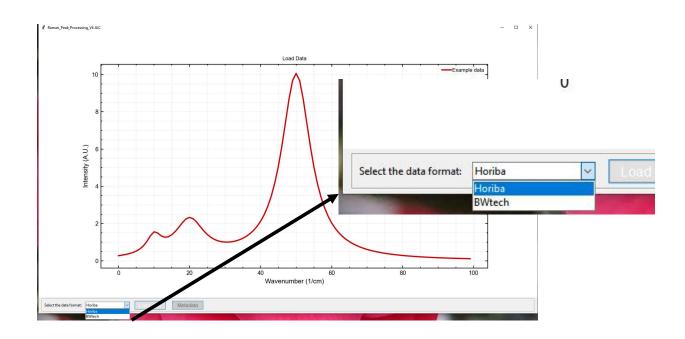


Select the Single Spectrum Analyzer:



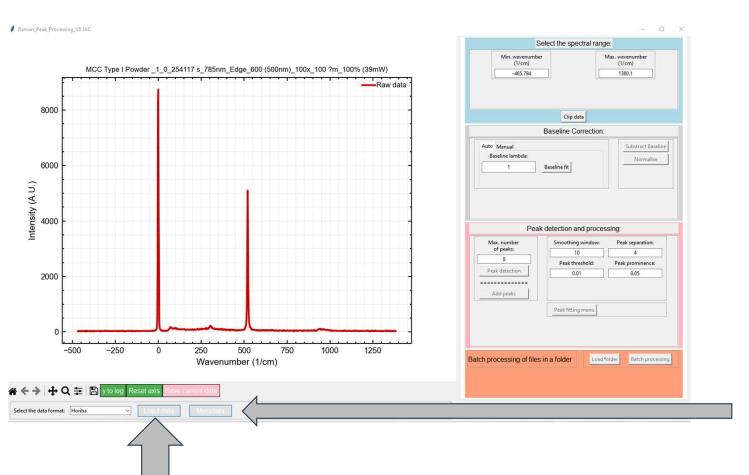


# • Select the file format:





# • Load data:





Retrieve metadata:

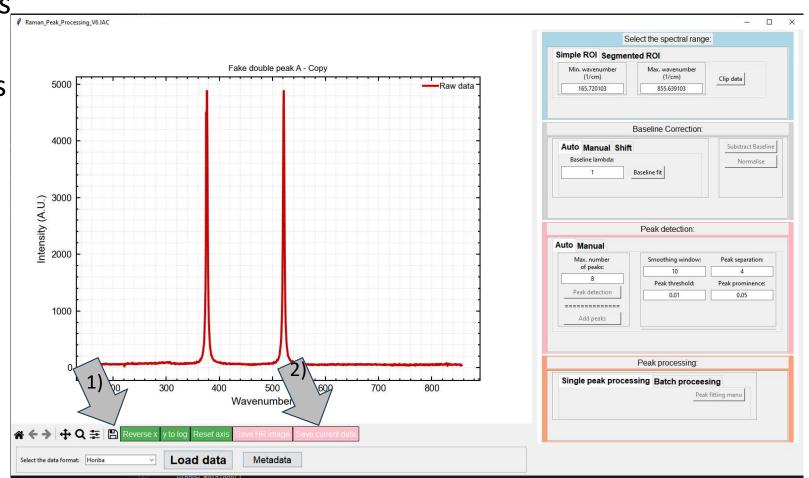


## Step 4-b

UVa

 Image can be saved as a .png clicking in 1)

 The data that appears in the image can be saved as a .csv by clicking 2)

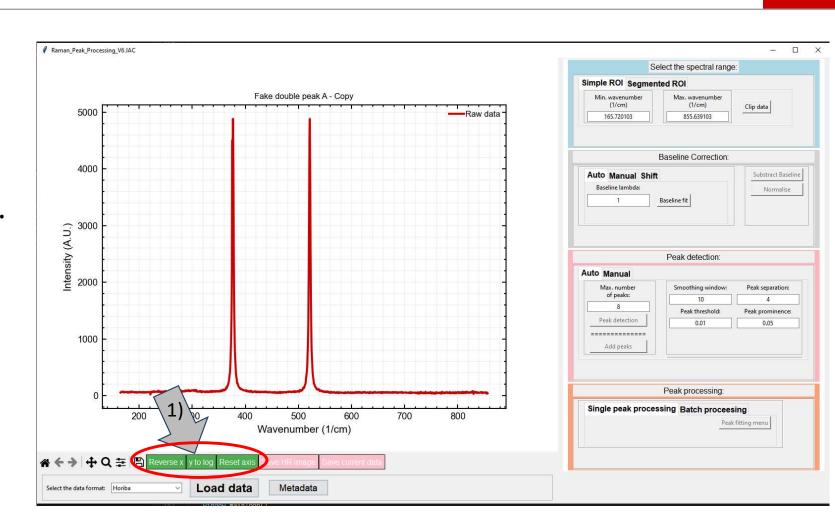




## Step 4-c

**UVa** 

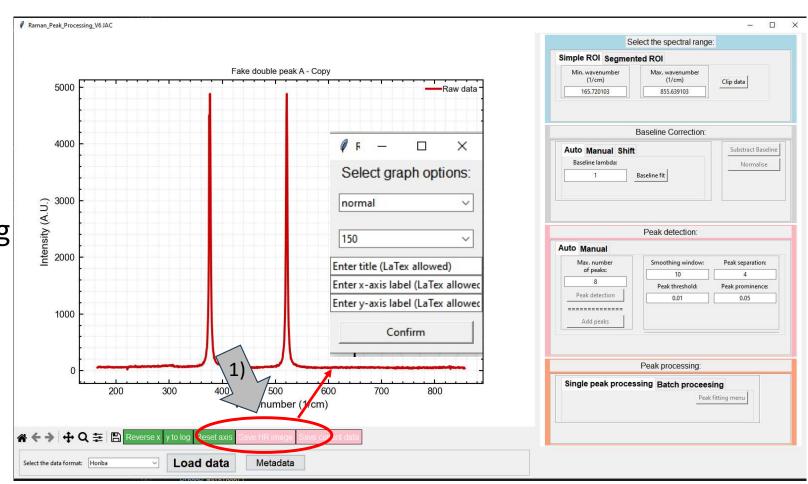
 The data can be plotted in reverse order (for IR for example). Also the intensity can be plotted in Log scale.



**UVa** 



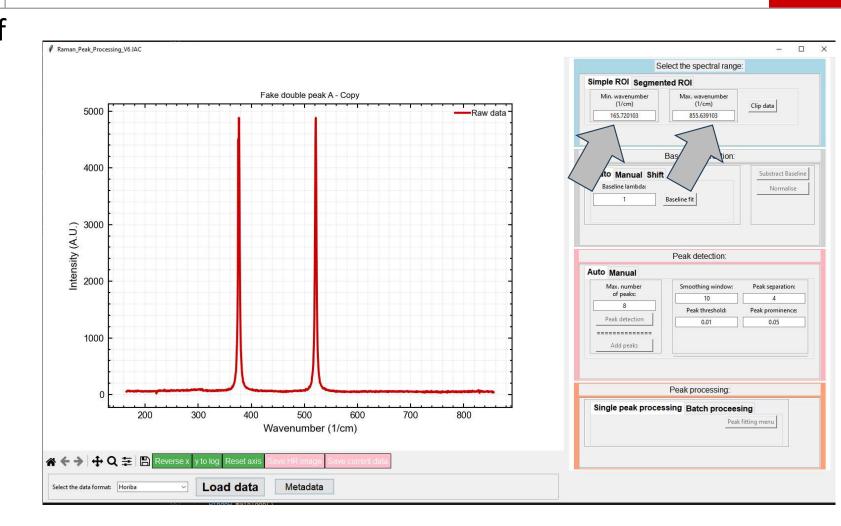
- The graphs of the window can be exported as publication ready graphs.
- Title and axis labels can be modified using LaTex syntaxis.
- The size of the image and the DPI can also be adjusted.





# Step 5

• Select the region of interest for the analysis.





## Step 6-a

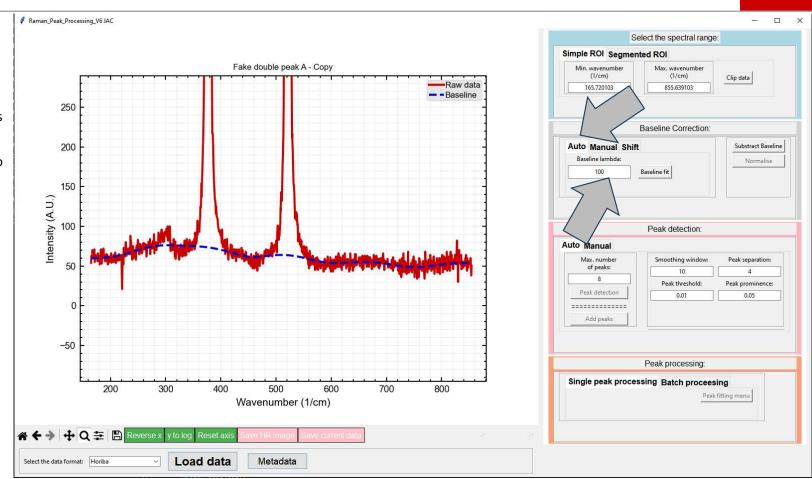
## UVa

#### • Remove baseline:

#### • Auto:

- Uses the asymmetrically reweighted penalized least squares (arPLS) baseline correction on the given spectrum.\* Adjust lambda (0-1e6), closer to 1 less sensitive to peak, higher values more sensitive to peaks
- Subtract baseline.
- Normalise the data to the highest peak if needed.

\*Sung-June Baek, Aaron Park, Young-Jin Ahna and Jaebum Choo: "Baseline correction using asymmetrically reweighted penalized least squares smoothing", Analyst, 2015,140, 250-257





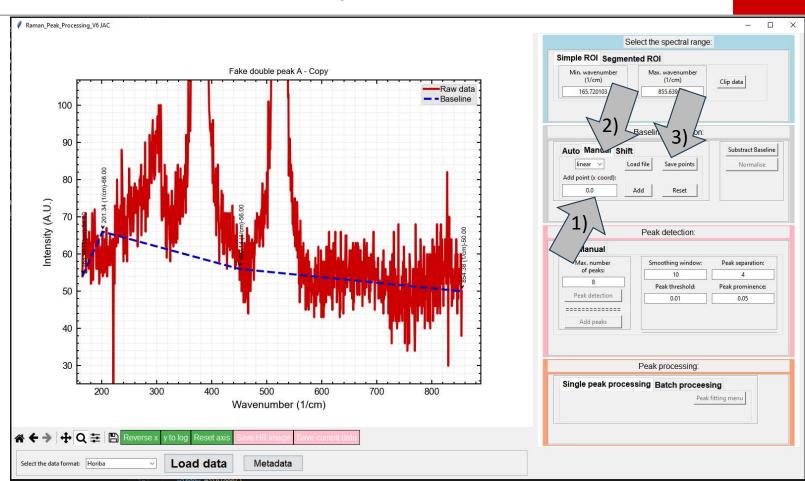
# Step 6-b

# UVa

#### • Remove baseline:

#### • Manual:

- Uses baseline defined by the user.
- Add manually as many wavenumber points as need 1).
- Choose the order of the baseline 2).
- The user can save the baseline as a .csv file and load that file for other datasets 3).
- Subtract baseline.
- Normalise the data to the highest peak if needed.

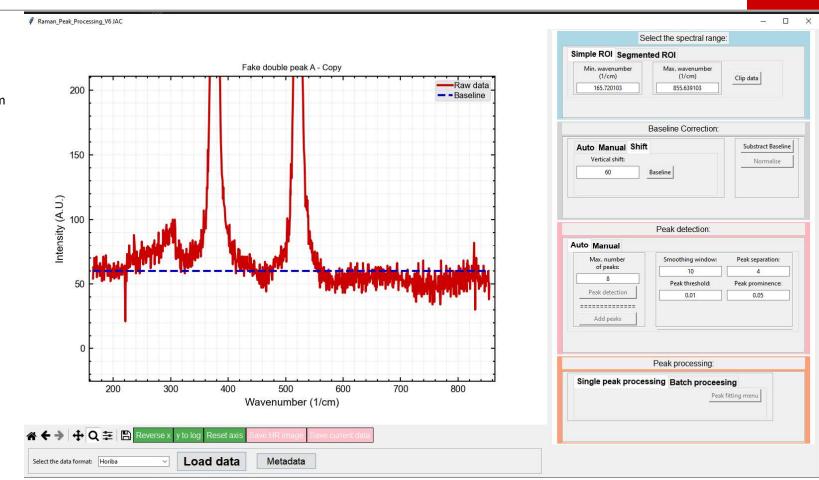


# Step 6-c



## • Remove baseline:

- Shift:
  - A constant shift is subtracted from the range of interest.





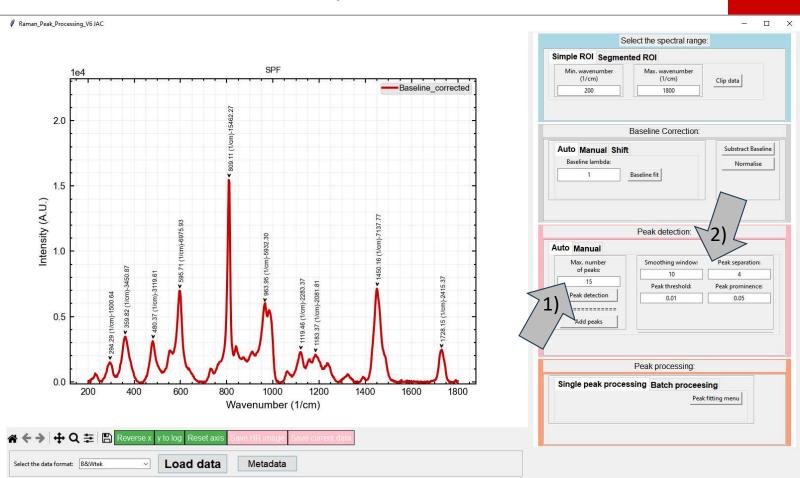
## Step 7-a



# • Peak detection:

#### • Auto:

- Select maximum number of peaks to be detected (maximum 25) 1).
- Tweak the detection parameters to find the peaks 2).
  - Smoothing window: Size of the smoothing window to produce a smooth and continuous dataset for derivative-based peak finding algorithm).
  - Peak separation: in pixels, minimum of 3pixels to detect contiguous peaks.
  - Peak Threshold: Ratio from minimum peak heigh to max peak heigh to be detected.
  - Peak prominence: Parameter to be more or less sensitive to shoulders.





## Step 7-b

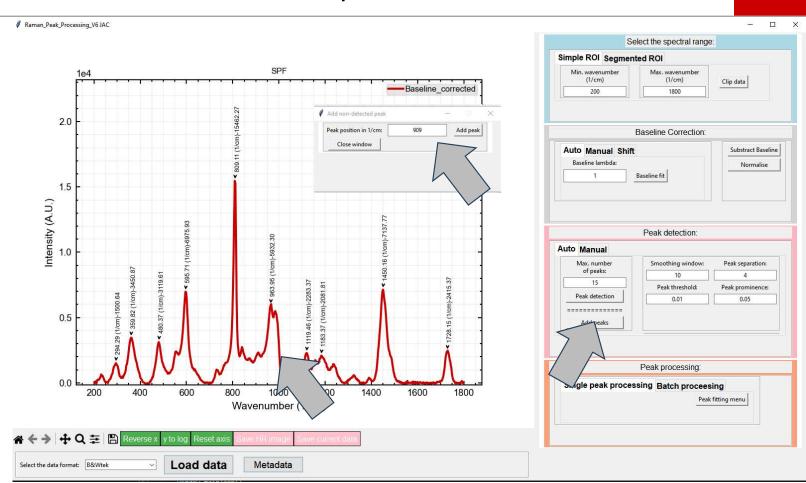


## • Peak detection:

#### • Auto:

#### • Manual:

- Automatic Peaks can be added if the automatic routine cannot find them by introducing their wavenumber.
- This is a combination of the automatic routine (7-a) and only works for the single spectrum processing. The batch processing only works with automatic detection limits or the full manual routine.





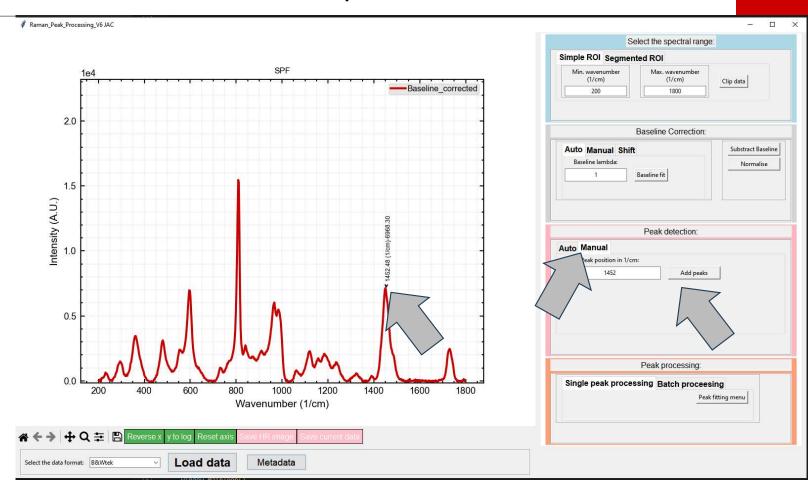
## Step 7-c

#### UVa

# • Peak detection:

#### • Manual:

- Peak positions can be specified manually.
- This positions will also be remembered for the batch processing, but if new peaks appear the will be ignored or not detected.
- Note that when swapping the tab from manual; to auto or auto to manual the previous peaks are erased.



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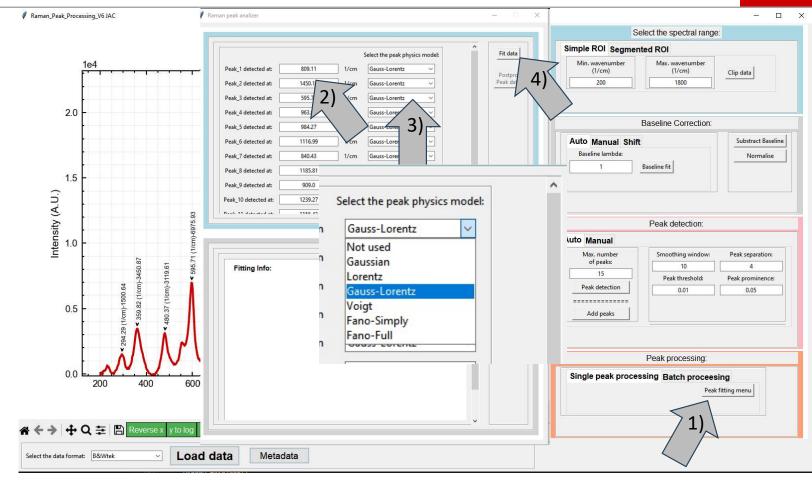


#### Step 8-a

#### UVa

# Peak fitting:

- Select peak fitting menu
  1).
- Finetune the position of detected peak if needed 2).
- Select the individual model for each peak.
- By default, it uses a Gauss-Lorentz approach (pseudo-Voigt) with a FWHM estimated following Kielkopf approach (0.02% error)\*
- Full Voigt and Voigt Fano (Fano-Full) can be used as well.

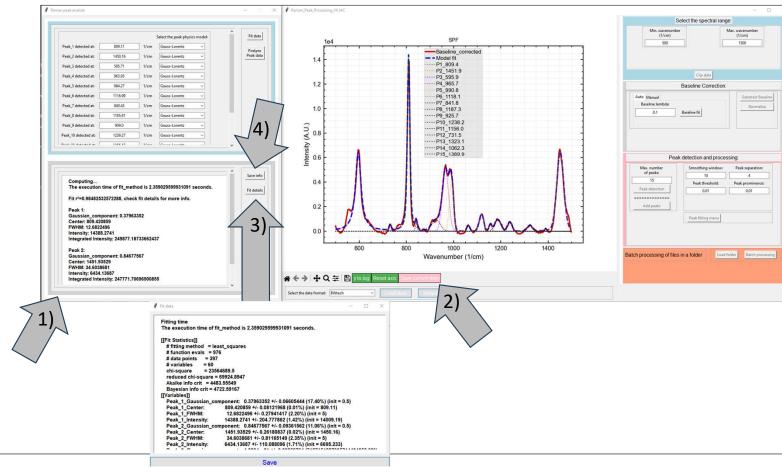


\* John F. Kielkopf (1973), "New approximation to the Voigt function with applications to spectral-line profile analysis", *Journal of the Optical Society of America*, **63** (8): 987,



# • Peak fitting:

- Review the fit data summary 1).
- The resultant fit is shown in the main window.
- Fitted peak positions are shown in the legend from higher intensity to lower intensity.
- All data displayed in the figure can be saved as a .csv file 2).
- Complete fitting details, including crosscorrelations can be retrieved and saved 3)
- Summary report can be saved as .txt 4)



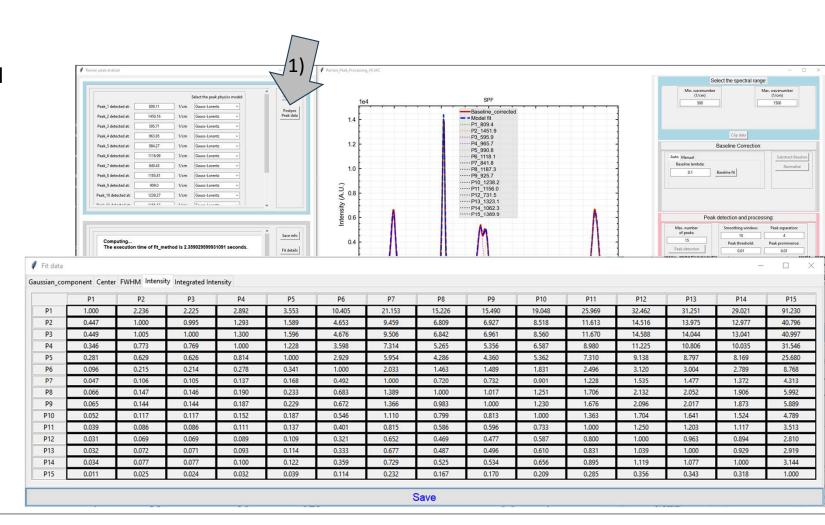
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# Step 8-c

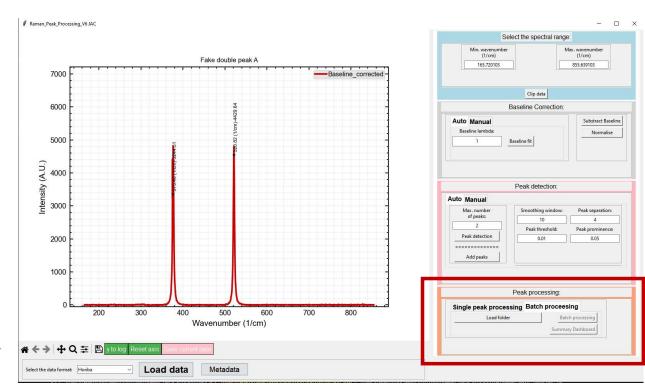
## • Peak fitting:

 Postprocessed correlations between all parameters can be calculated and saved 1)



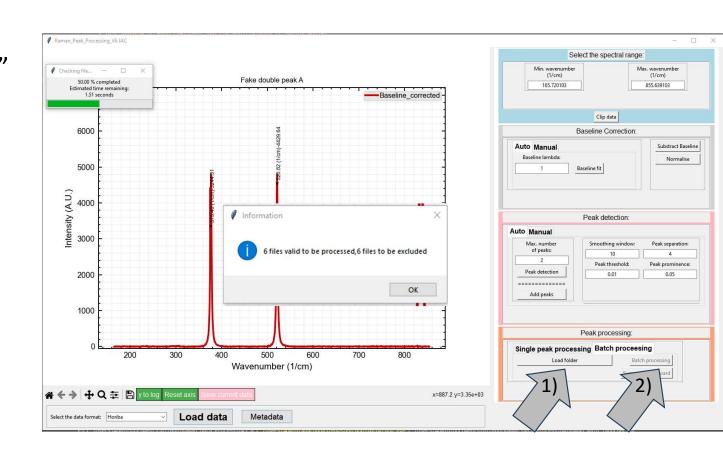


- 1. Save all the raw data files to be processed in . a common folder.
- 2. Open one data file and follow steps 3 to 7.
- 3. This fixes the following conditions to analyse all files in the folder.
  - 1. Data range
  - 2. Baseline type and properties
  - 3. Max peak detected and parameters to detect a peak.
  - 4. All peaks fitted with the default pseudo-Voigt default function.



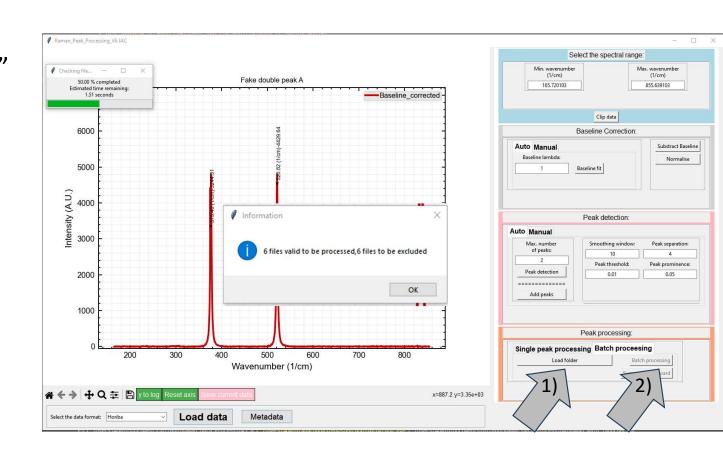


- 4. When the parameters are selected, click on "Load folder" button 1). This will scan for valid .txt data files in the folder.
- 5. Click on "Batch processing" to analyse all files in folder 2).



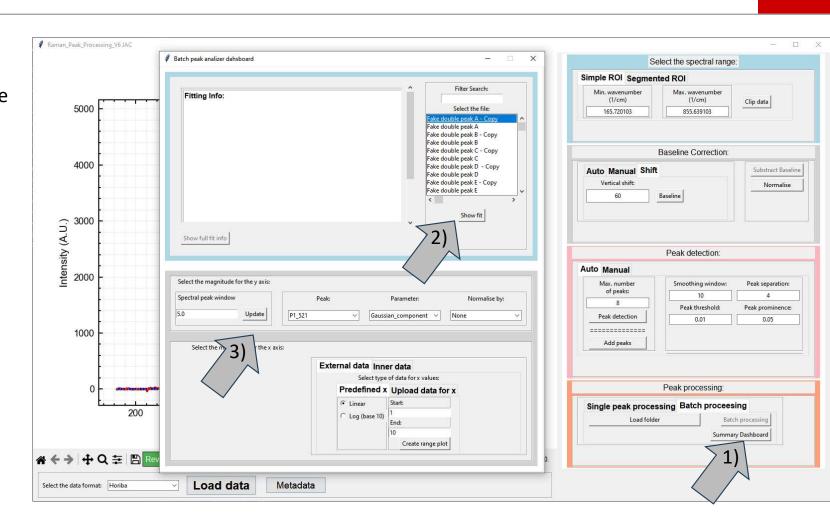


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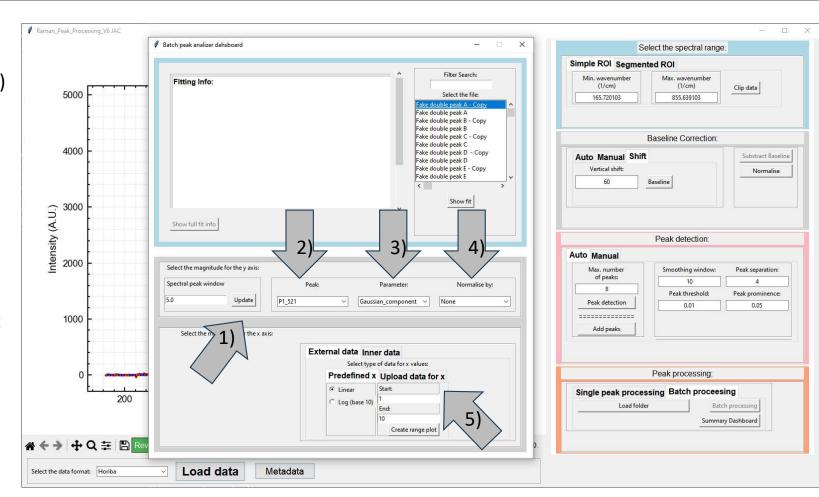


- 4. A dashboard panel will appear that allow to retrieve the fit for each file in the folder (2)
- 5. Data for each parameter in the folder can be also explored using the bottom dashboard (3).





- 4. Select the minimum separation between peaks (1)
- 5. Select the peak you want to analyse (2)
- Select the parameter you want to extract for all files in the folder (3). You can normalise that value by the same value of a different peak if needed (4)
- Select against what you want to plot the results. Can be a predefined equispaced sequence or a file containing the values for the X.



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