

Physics-Informed Deep-Learning Emulators for Type Ia Supernovae

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Background

Type Ia supernovae, the thermonuclear explosions of Carbon-Oxygen white dwarfs, **remain poorly understood** despite decades of investigation. A variety of sub-classes of Type Ia supernovae exist and the connection between these populations is an active area of research.

What is the cause of the observational differences between Type Ia sub-classes?

Do different Type Ia sub-classes share a common progenitor scenario?

Massive computationally intensive hydrodynamic simulations have been developed to model an ever-growing number of theoretical explosion scenarios with increasing computational fidelity. However, **connecting individual supernovae observations to theoretical models** through physical simulations **remains computationally intractable**.

Solution

Radiative Transfer Emulators
+
Bayesian Inference

Probabilistically Inferring the Abundance Distributions of Type Ia Supernovae Ejecta from Spectral Observations

