Documentation for module stacker

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1 Introduction

This document describes version 1.0 of the module stacker. The module is designed to stack interferometric data. Primarily it was designed to allow stacking in the uv domain, but supports stacking in image domain.

2 Algorithm

The algorithms used by this code are described in Lindroos et al. (2014). The image based algorithm supports either mean or median stacking, where the flux from subimages are average with either the median or a weighted mean method. It works on a pixel-by-pixel basis in a defined stamp around the stacking positions

The uv-stacking algorithm is based on aligning the phases of all visibilties to the target sources. Stacked visibilities are calculated as

$$V_{\text{stack}}(u, v, w) = V(u, v, w) \frac{\sum_{k=1}^{N} W_k \frac{1}{A_N(l_k, m_k)} e^{\frac{2\pi}{\lambda} i \left(u l_k + v m_k + w \left(\sqrt{1 - l_k^2 - m_k^2} \right) \right)}{\sum_{k=1}^{N} W_k}$$
 (1)

where (l_k, m_k) are the separation of stacking position k from the phase centre, (u, v, w) are the coordinates of the visibility in the uv plane, λ is the wavelength, $A_N(l_k, m_k)$ is the primary beam attenuation at stacking position k, and W_k is the weight of the stacking position.

3 Usage

This document provides a quick overview of the available functionalities in stacker. The module contains 3 packages: pb, image and uv. The packages image and uv provides functions to stack in the image and uv domain respectively. The package pb allows for describing the primary beam model. For a typical example of how stacker can be used see the provided example in "example/stack_testdata.py".

3.1 Coordinates

The coordinates are described by a coordList object in the module. A coordList object can be generated from a csv file with the function stacker.readCoords, i.e.,

```
coord = stacker.readCoords(<path to coordinate file>)
Syntax of coordinate file should be
x1, y1(, weight1)
x2, y2(, weight2)
etc. For more info run help(stacker.readCoords) in casapy.
    A coords object can also be built from scratch
coords = stacker.CoordList()
coords.append(stacker.Coord(x1, y1))
coords.append(stacker.Coord(x2, y2))
Note that coordinates here should be give in J2000 radians.
```

3.2 Primary beam model

The primary beam model can be defined by an instance of stacker.pb.PrimaryBeamModel or inherited classes. The c code only support stacker.pb.MSPrimaryBeamModel and stacker.pb.PrimaryBeamModel. A primary beam model can be generated from a measurement set with the function stacker.pb.guesspb.

```
pbmodel = stacker.pb.guesspb(<path to ms file>)
```

If uv stacking is run without specifying a primary beam model it will automatically attempt to use stacker.pb.guesspb.

3.3 uv

Submodule for stacking in the *uv* domain. Primarily provides two functions stacker.uv.stack and stacker.uv.noise. The first perform (stacker.uv.stack) the actual stacking, and requires an input *uv*-data file and a coordsList object as input.

For more info on usage run help(stacker.uv.stack) in casapy.

The second (stacker.uv.noise) calculates noise using a Monte Carlo where random positions are stacked to estimate the noise level. The function will try to recompute weights for the random positions. If you require variable weights which are not simply the primary beam or the noise in a local stamp you will have to re-implement the function.

3.4 image

Submoduls for stacking in the image domain. Provides the same functions as the submodule uv except it works fully in the image domain.

The image maps are assumed to have been primary beam corrected. The package can handle a list of images as well as an individual image. An individual image should be specified as a one element list. For more details run help(stacker.image.stack) in casapy.

The image package also provides functions to calculate local weights for positions from the stamps surrounding them.

References

Lindroos L., Knudsen K. K., Vlemmings W., Conway J., Martí-Vidal I., 2014, in prep.