Documentation for module stacker

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1 Introduction

This document describes version 1.0 of the module stacker. The module is designed to stack interferometric data. Primarily it was designed to allow stacking in the uv domain, but supports stacking in image domain.

2 Algorithm

The algorithms used by this code are described in Lindroos et al. (2014). The image based algorithm supports either mean or median stacking, where the flux from sub images are average with either the median or a weighted mean method. It works on a pixel-by-pixel basis in a defined stamp around the stacking positions

The uv-stacking algorithm is based on aligning the phases of all visibilities to the target sources. Stacked visibilities are calculated as

$$V_{\text{stack}}(u, v, w) = V(u, v, w) \frac{\sum_{k=1}^{N} W_k \frac{1}{A_N(l_k, m_k)} e^{\frac{2\pi}{\lambda} i \left(u l_k + v m_k + w \left(\sqrt{1 - l_k^2 - m_k^2} \right) \right)}{\sum_{k=1}^{N} W_k}$$
 (1)

where (l_k, m_k) are the separation of stacking position k from the phase centre, (u, v, w) are the coordinates of the visibility in the uv plane, λ is the wavelength, $A_N(l_k, m_k)$ is the primary beam attenuation at stacking position k, and W_k is the weight of the stacking position.

3 Usage

This document provides a quick overview of the available functionalities in stacker. The module contains 3 packages: pb, image and uv. The packages image and uv provides functions to stack in the image and uv domain respectively. The package pb allows for describing the primary beam model. For a typical example of how stacker can be used see the provided example in "example/stack_testdata.py".

3.1 Coordinates

The coordinates are described by a coordList object in the module. A coordList object can be generated from a csv file with the function stacker.readCoords, i.e.,

```
coord = stacker.readCoords(<path to coordinate file>)
Syntax of coordinate file should be
x1, y1(, weight1)
x2, y2(, weight2)
etc. For more info run help(stacker.readCoords) in casapy.
    A coords object can also be built from scratch
coords = stacker.CoordList()
coords.append(stacker.Coord(x1, y1))
coords.append(stacker.Coord(x2, y2))
```

Note that coordinates here should be give in J2000 radians.

3.2 Primary beam model

The primary beam model can be defined by an instance of stacker.pb.PrimaryBeamModel or inherited classes. The c code only support stacker.pb.MSPrimaryBeamModel and stacker.pb.PrimaryBeamModel. A primary beam model can be generated from a measurement set with the function stacker.pb.guesspb.

```
pbmodel = stacker.pb.guesspb(<path to ms file>)
```

If uv stacking is run without specifying a primary beam model it will automatically attempt to use stacker.pb.guesspb.

A custom primary-beam model can be used by writing it to a casa image. The image should be centered at (0.,0.). The primary beam is loaded into casa using

```
pbmodel = stacker.pb.MSPrimaryBeamModel(<path to pb-file>)
```

3.3 uv

Submodule for stacking in the uv domain. Primarily provides two functions stacker.uv.stack and stacker.uv.noise. The first perform (stacker.uv.stack) the actual stacking, and requires an input uv-data file and a coordsList object as input.

For more info on usage run help(stacker.uv.stack) in casapy.

The second (stacker.uv.noise) calculates noise using a Monte Carlo where random positions are stacked to estimate the noise level. The function will try to recompute weights for the random positions. If you require variable weights which are not simply the primary beam or the noise in a local stamp you will have to re-implement the function.

3.4 image

Submoduls for stacking in the image domain. Provides the same functions as the submodule uv except it works fully in the image domain.

The image maps are assumed to have been primary beam corrected. The package can handle a list of images as well as an individual image. An individual image should be specified as a one element list. For more details run help(stacker.image.stack) in casapy.

The image package also provides functions to calculate local weights for positions from the stamps surrounding them or from the primary beam. For ALMA data it is most likely best to use the pb weighting mode. This mode requires a primary beam model to calculate the weights, for more details see section 3.2.

4 Stacking related techniques

This sections describes a few of techniques which are not unique to stacking, but may be useful in the context of stacking.

4.1 Removing bright sources

Bright non-target sources present in the data can significantly impact the stacking result. To reduce this effect it a model of the bright sources can be subtracted from the uv data. If the data is imaged with clean in casa the model can be subtracted using the task uvsub.

The package modsub also provides a facility to subtract a model from the uv data. As input it takes a casa component list of the model.

```
import stacker.modsub
stacker.modsub.modsub(<path to casa cl file>, <path to in ms file>, <path to out ms file>)
```

This function can also be given a custom primary-beam model defined by the pb package.

4.2 Stacking in mosaic data sets

When uv stacking in mosaics it is important that the relative weights between the different pointings are scaled correctly. This may not be the case for non contiguous mosaics. In these cases the weights can be recomputed using the casa task statwt. This calculates the weight of each visibility using the scatter over nearby visibilities.

If using image-stacking it is most likely best image each pointing separately. The function stacker.image.stack can be given a list of all pointings, and will automatically assign proper weights to each position in each pointing.

References

Lindroos L., Knudsen K. K., Vlemmings W., Conway J., Martí-Vidal I., 2014, in prep.