CELESTIAL MECHANICS (Fall 2012): COMPUTER EXERCISES II - SOLUTIONS

(Heikki Salo 2.11.2012)

- 5. Calculation of R and V from orbital elements: 3D case
- 6. Calculation of orbital elements from R and V
- 7. Playing with orbits

These example routines for the solution of the exercises can be copied from

/wrk/hsalo/TM2012_idl.dir/TM2012_DEMO2.dir

Additional useful procedures:

/wrk/hsalo/TM2012_idl.dir/TM2012_DEMO1.dir -- Solutions for exercises I

/wrk/hsalo/TM2012_idl.dir/TM2012_DEMO_apu.dir -- Useful auxillary programs

5. Calculation of R and V from orbital elements elem_to_rv.pro

```
•
; elem_to_rv.pro
; solves R and V from orbital elements
; HS 6.5.1997 -> 05.11.02 / 01.11.06
****************
    pro elem_to_rv,elem,time,rad,vel,myy0=myy0,eano=eano,mano=mano
if(n_params() le 0) then begin
       print,'-----
       print,'elem_to_rv,elem,time,rad,vel,myy0=myy0'
               calculates radius- and velocity-vector at a given time
       print,'
                 from the orbital elements'
       print,'----
       print,' elem = orbital elements [a,eks,ink,ome,w,tau]'
print,' a = semimaior-aria'
                a = semimajor-axis'
eks = eccentricity (>1 -> hyperbolic)'
       print,'
                  ink = inclination'
       print,'
       print,'
                   ome = longitude of ascending node'
       print,'
                  w = argument of pericenter'
       print,'
                   tau = time of pericenter passage'
               tau = time or pericente
time = instant of time'
       print,'
                     NOTE: angles in degrees'
       print,' INPUT VIA KEYWORDS:'
print,' myy = G (m1+m2) : def=1'
print,' OUTPUT:'
       print,' rad,vel = radius and velocity vector'
       print, 'OUTPUT VIA KEYWORDS:'
       print,' eano=eano eccentric anomaly (degrees)'
print,' mano=mano mean anomaly (degrees)'
       print,' HS 6.5.1997 / 22.10.2002 / 01.11.2006 /2008'
     return
endif
      pii=atan(1.d0)*4.d0
      ra=180.d0/pii
      ar=pii/180.d0
      myy=1.d0
      if(keyword_set(myy0)) then myy=myy0
********************************
     ap=isoakseli
                       semimajor axis
     eps=eksentrisyys
                                eccentricity
     ink=inklinaatio
                                inclination
     ome=nousevan solmun pituus longitude of node w=perihelin argumentti argument of pericenter tau=periheliaika pericenter time
ap=elem(0) & eps=elem(1) & ink=elem(2)
     bp=ap*sqrt(abs(1.d0-1.d0*eps*eps))
*************************************
    Orbital vectors ii,jj,kk
     (idl-arrays start at zero: here zero-elements are dummies!)
*******************
    ii=dblarr(4) & jj=ii & kk=ii
    sino=sin(ome) & coso=cos(ome)
    sini=sin(ink) & cosi=cos(ink)
    sinw=sin(w) & cosw=cos(w)
```

```
ii(1)=cosw*coso-sinw*sino*cosi
    ii(2)=cosw*sino+sinw*coso*cosi
    ii(3)=sinw*sini
    jj(1)=-sinw*coso-cosw*sino*cosi
    jj(2)=-sinw*sino+cosw*coso*cosi
    jj(3)=cosw*sini
    kk(1)=sino*sini
    kk(2)=-coso*sini
    kk(3)=cosi
********************
;mean anomaly
*************************************
    per=2.d0*pii*sqrt(ap*ap*ap/myy)
    m=2.d0*pii/per*(time-tau)
;solve eccentric anomaly
;cheat a bit: parabolic orbit described by a highly elliptic orbit
if(eps eq 1) then begin
      eps=.999999
      a=10000.
endif
;elliptic orbit - Kepler's equation
if(eps le 1) then begin
      kepler,m,eps,e
    dedt=sqrt(myy/ap^3)/(1.d0-eps*cos(e))
    xx=ap*(cos(e)-eps)
    yy=bp*sin(e)
    vxx=-ap*sin(e)*dedt
    vyy=bp*cos(e)*dedt
    if(e lt 0.) then e=e+!pi*2.d0
;hyperbolic orbit - hyperbolic Kepler's equation
if(eps gt 1) then begin
      hkepler,m,eps,e
    dedt=sqrt(myy/ap^3)/(eps*cosh(e)-1.d0)
    xx=ap*(eps-cosh(e))
    yy=bp*sinh(e)
    vxx=-ap*sinh(e)*dedt
    vyy=bp*cosh(e)*dedt
\verb"endif"
    mano=m*ra & eano=e*ra
**************************************
      convert coordinates from orbit-coordinates
      to the reference frame
***************
    x=xx*ii(1)+yy*jj(1)
    y=xx*ii(2)+yy*jj(2)
    z=xx*ii(3)+yy*jj(3)
   vx=vxx*ii(1)+vyy*jj(1)
vy=vxx*ii(2)+vyy*jj(2)
    vz=vxx*ii(3)+vyy*jj(3)
    rad=[x,y,z]
   vel=[vx,vy,vz]
   end
```

6. Calculation of orbital elements from R and V rv.to.elem.pro

```
; rv_to_elem.pro
; solves orbital elements from radius and velocity vectors
; HS 6.5.1997 -> 05.11.02 -> 01.11.06
pro rv_to_elem,time,rad,vel,elem,$
                print=print,myy0=myy0,kvec=kvec,evec=evec,$
                ene=ene,imp=imp,eano=eano,mano=mano
if(n_params() le 0) then begin
 print,'-----'
 print,'pro rv_to_elem,time,rad,vel,elem'
          calculates orbital elements from radius and velocity vectors'
 print,'
           at a given time'
 print, '-----'
 print,' INPUT:'
print,' time = time of observation'
print,' rad = [ x, y, z] radius vector'
print,' vel = [vx,vy,vz] velocity vector'
print, ' INPUT VIA KEYWORDS:'
print,' myy = value - def: G*(m1+m2) = 1.'
print,' '
print,' OUTPUT:'
print,' elem=[a,eks,ink,ome,w,tau]'
print,' a = semimajor-axis'
           a = semimajor-axis
print,'
              eks = eccentricity (>1 -> hyperbolic)'
              ink = inclination'
ome = longitude of ascending node'
 print,'
 print,'
            w = argument of pericenter'
tau = time of pericenter passage'
 print,'
 print, OUTPUT VIA KEYWORDS:
 print,' kvec=kvector =[kx,ky,kz]'
print,' evec=evector =[ex,ey,ez]'
print,' imp=angular momentum/unit mass = |kvec|'
print,' ene=energy /unit mass'
ene=energy /unit mass'
print,' eano=eccentric anomaly (degrees)'
print,' mano=mean anomaly (degrees)'
print,'
print,' ADDITIONAL KEYWORDS:'
print,' /print -> some intermediate output'
print,' HS 6.5.1997 / 22.10.2002 /01.11.2006'
print,'
 return
endi f
;defines units
 myy=1.d0
 if(keyword_set(myy0)) then myy=myy0
 pii=atan(1.d0)*4.d0
ra=180.d0/pii
 ar=pii/180.d0
; inklinaatio ja nousevan solmun pituus
 inclination and the longitude of ascending node
************
; cross_product,A,B,C \longrightarrow C = A x B
; dot_product,A,B,c
                          --> c = A \cdot B
```

```
; here A,B,C denote 3-element vectors, c denotes a scalar
; angle=atan(y,x) returns angle for which y=sin(angle), x=cos(angle)
   cross_product,rad,vel,kvec
   dot_product,kvec,kvec,kk
   n=kvec/sqrt(kk)
   nx=n(0)
   ny=n(1)
   nz=n(2)
   ink=acos(nz)
   ome=atan(nx,-ny)
   imp=sqrt(kk)
*************************
; eksentrisyys ja perihelin argumentti: e-VEKTORI
; eccentricity and the \ensuremath{\operatorname{argument}} of pericenter
*************
   cross_product,kvec,vel,apu1
   dot_product,rad,rad,r2
   apu2=rad/sqrt(r2)
   evec=-apu1/myy-apu2
; transform evec to the coordinate system aligned % \left( 1\right) =\left( 1\right) \left( 
;with the orbital plane
   coso=cos(ome)
   sino=sin(ome)
   cosi=cos(ink)
   sini=sin(ink)
   ex=coso*evec(0)+sino*evec(1)
   ey=-sino*cosi*evec(0)+coso*cosi*evec(1)+sini*evec(2)
   eks=sqrt(ex^2+ey^2)
   w=atan(ey,ex)
;isoakseli
;semimajor axis
*************************
   dot_product,vel,vel,v2
   dist=sqrt(r2)
   h=0.5d0*v2-myy/sqrt(r2)
   ene=h
   a=abs(myy/2.d0/h)
   per=2.0d0*!dpi*sqrt(a^3/myy)
·**************
; periheliaika
; time of pericenter passage
************
;via cos(E)
;to solve the right sign:
;0 \langle E < 180 \rightarrow r \text{ is increasing} \rangle
;180 \langle E < 360 \rightarrow r is decreasing
   sign=1.
   dot_product,rad,vel,rv
   if(rv lt 0.) then sign=-1.
```

```
***********
;elliptic orbit
; just to avoid numerical problems:
if(eks le 0.) then begin
  e=0.
  tau=time
\verb"endif"
if(eks lt 1. and eks gt 0.) then begin
  cose=(1.-dist/a)/eks<1.d0>(-1.d0)
  e=acos(cose)*sign
  m=e-eks*sin(e)
  tau=(time-m*per/2./!dpi) mod per
endif
· **************
;hyperbolic orbit
;acosh-function returns the inverse of hyperbolic cosine
;(in acosh.pro)
if(eks ge 1.) then begin
  coshe=(1.d0+dist/a)/eks>1.
  e=acosh(coshe)*sign
  m=eks*sinh(e)-e
  tau=time-m*per/2./!pi
\verb"endif"
eano=e*!radeg
mano=m*!radeg
elem=[a,eks,ink*!radeg,ome*!radeg,w*!radeg,tau]
\quad \text{if(keyword\_set(print)) then begin} \\
  st=' at T='+string(time)
  print,'-----
  if(eks lt 1) then print, 'CALCULATED ELEMENTS OF ELLIPTIC ORBIT: '+st
  if(eks gt 1) then print, 'CALCULATED ELEMENTS OF HYPERBOLIC ORBIT: '+st
  print,'a = ',elem(0)
print,'eks = ',elem(1)
  print,'ink = ',elem(2)
  print, ink = ',elem(2)
print,'ome = ',elem(3)
print,'w = ',elem(4)
print,'tau = ',elem(5)
  print,''
  print,'per = ',per,m*ra
print,'q = ',elem(0)*abs(1.-elem(1))
print,''
  print,'
  print, '----'
endif
```

end

Procedure for solving Kepler's equation for hyperbolic orbits heepler.pro

```
*****************
    RATKAISTAAN HYBERPOLINEN EKSENTRINEN ANOMALIA E
    ITEROIMALLA KESKIANOMALIASTA M
    RADAN EKSENTRISYYS = EKS > 1
pro hkepler,m,eks,e,itul=itul,itemax=itemax,check=check
if(n_params() le 0) then begin
     print,'-----
     print,' hkepler,M,eks,E'
     print,'-----'
     print, 'USING NEWTON-ITERATION'
     print,' input: M = mean anomaly (radians), eks=eccentricity '
print,' output: E = eccentric anomaly (radians)'
     print,' /itul -> output iteration values (E)'
print,' /check -> check accument
     print,' keywords: '
     print,' itemax = max number of iterations (def=500)'
     print,' example: eks=3, M= 45 degrees, output iteration values'
     print," hkepler,45./!radeg,3.,e,/itul & print,'E=',e*!radeg,' deg'"
     print, 'HS 04.10.02 /01.11.06'
     print,'----'
  return
endif
******************
; Newton iteration for solving f=eks*sinh(E)-E-M=0
; E_NEW = E_OLD - f(E)/(df(E)/dE)
; until abs(E_NEW-E_OLD) < 1d-10</pre>
   nstep=5000
   if(keyword_set(itemax)) then nstep=itemax
;choose the intitial guess
   if(abs(m) gt exp(1)) then e=m/abs(m)*alog(2.*abs(m)/eks)
   if(keyword_set(itul)) then print,i,e
    for i=1,nstep do begin
         fe=eks*sinh(e)-e-m
         fep=eks*cosh(e)-1.d0
         eold=e
         e=eold-fe/fep
         if(abs(e-eold) lt 1d-10*abs(e)) then goto,jump11
         if(keyword_set(itul)) then print,i,e
      endfor
      if(keyword_set(check)) then print,m-(eks*sinh(e)-e)
    return
   end
```

Auxillary routines

ACOSH acosh.pro

```
;acosh.pro
;hyperbolisen cosinin kaanteisfunktio
;y=cosh(x)=0.5*(exp(x)+exp(-x)) --> x=acosh(y)
;HS circa 1997
function acosh,y
    x=alog(y+sqrt(y^2-1))
return,x
end
```

• VECTOR PRODUCT cross_product.pro

```
*********
; cross\_product.pro
**************
pro cross_product,a,b,c
if(n_params(0) le 0) then begin
   print, '-----'
   print,' pro cross_product,A,B,C'
print,' C=A x B'
   print,'-----
   print,' case 2: A and B contain an array of 3-element vectors'
   print,' A=array(n,3) B=array(n,3) -> C=array(n,3)'
print,'-----'
   print, 'HS ca 2000'
   return
endif
;case 1
if(n_elements(a) eq 3) then begin
   c(0)=a(1)*b(2)-b(1)*a(2)
   c(1)=a(2)*b(0)-b(2)*a(0)
   c(2)=a(0)*b(1)-b(0)*a(1)
endif
;case 2
if(n_elements(a) gt 3) then begin
   n=n_elements(a)/3
   c(*,0)=a(*,1)*b(*,2)-b(*,1)*a(*,2)
   c(*,1)=a(*,2)*b(*,0)-b(*,2)*a(*,0)
   c(*,2)=a(*,0)*b(*,1)-b(*,0)*a(*,1)
return
end
```

• SCALAR PRODUCT dot_product.pro

```
;dot_product.pro
;scalar-product c = a . b
*******
pro dot_product,a,b,c
if(n_params(0) le 0) then begin
   print,'-----
print,' pro dot_product,A,B,c'
   print,' scalar product c = A . B'
print,'-----'
    print,' case 1: A and B are 3-element vectors:'
    print,' A=[ax,ay,az], B=[bx,by,bz]] -> c =ax*bx+ay*by+az*bz'
   print,' case 2: A and B contain an array of 3-element vectors' print,' A=array(n,3) B=array(n,3) -> c=array(n)' print,'-----'
    print, 'HS ca 2000'
    return
endif
;case 1
if(n_elements(a) eq 3) then begin
   c=a(0)*b(0)+a(1)*b(1)+a(2)*b(2)
if(n_elements(a) gt 3) then begin
   n=n_elements(a)/3
    c=a(*,0)
    c=a(*,0)*b(*,0)+a(*,1)*b(*,1)+a(*,2)*b(*,2)
\verb"endif"
return
end
```

Checking elem_to_rv.pro and rv_to_elem.pro elem_rv_example.pro

```
****************
;elem_rv_example.pro
***********************************
;example of using the procedures
; elem_to_rv : orbital elements -> R,V at given t
; rv_to_elem : R,V at given t -> orbital elements
;HS 22.10.2002/01.11.06
*******************************
  print, '----'
  print,'CHECK THE PROCEDURES:'
  print,' elem_to_rv : orbital elements -> R,V at given t'
print,' rv_to_elem : R,V at given t -> orbital elements '
;ELLIPTIC ORBIT
*************
  a=1.7
                         ;semimajor axis
  eks=0.1
                         ;eccentricity
  ink=10.
                         ;inclination
  ome=40.
                         ;longitude of ascending node
  w=20.
                         ;argument of pericenter
  tau=0.1
                         ;time of pericenter passage
  elem0=[a,eks,ink,ome,w,tau]
  time=0.7
  elem_to_rv,elem0,time,rad,vel
  rv_to_elem,time,rad,vel,elem1
  print,'original elements: elliptic orbit'
                                         ome
                                                            tau'
  print, elem0, f='(6f12.6)'
  print, 'new elements'
  print, elem1, f='(6f12.6)'
;HYPERBOLIC ORBIT
***************
                         ;semimajor axis
  eks=1.1
                         ;eccentricity
  ink=10.
                         ;inclination
  ome=40.
                         ;longitude of ascending node
  w=20.
                         ;argument of pericenter
  tau=0.1
                         ;time of pericenter passage
  elem0=[a,eks,ink,ome,w,tau]
  elem_to_rv,elem0,time,rad,vel
  rv_to_elem, time, rad, vel, elem1
  print,'-----'
  print,'original elements: hyperbolic orbit'
        a eks
                                                            tau'
  print,elem0,f='(6f12.6)'
  print, 'new elements'
  print,elem1,f='(6f12.6)'
  print, '-----'
```

This code should produce output:

	E PROCEDURES: _rv : orbital elements -> R,V at given t lem : R,V at given t -> orbital elements						
original elem	ents: ellipt:	ic orbit					
a	eks	ink	ome	W	tau		
1.700000	0.100000	10.000000	40.000000	20.000000	0.100000		
new elements							
1.700000	0.100000	9.999999	39.999998	19.999992	0.100000		
original elem	ents: hyperbo	olic orbit		-			
a	eks	ink	ome	W	tau		
1.700000	1.100000	10.000000	40.000000	20.000000	0.100000		
new elements							
1.700000	1.100000	9.999999	39.999998	19.999997	0.100000		

7. Playing with orbits

a) Procedure for calculating/plotting an orbit [elem_orbit.pro]

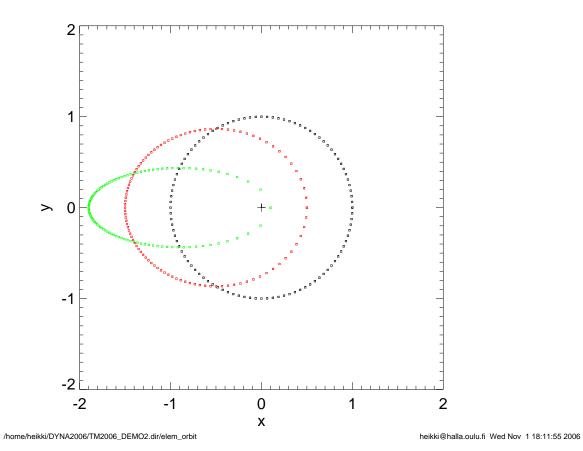
```
**************************
; elem_orbit.pro
; use elem_to_rv to calculate orbit
; HS 05.11.02 /01.11.2006
****************
pro elem_orbit,elem,timet,rr,vv,myy0=myy0,plot=plot
 if(n_params() le 0) then begin
     print,'elem_orbit,elem,timet,rr,vv,myy0=myy0'
     print,'elem = orbital elements [a,eks,ink,ome,w,tau]'
     print,'timet -> calculate RR(*,3) VV(*,3) for these times'
     print,'myy -> scaling (def=1.)'
     print,'plot=1 -> plot x,y orbit'
     print,'plot>1 -> oplot x,y orbit with col=plot'
     print,'Example:'
     print, 'timet=findgen(101)/100.*2*!pi \& !x.range=[-2,2] \& !y.range=[-2,2]' \\
     print, 'elem_orbit,[1.,.0,0.,0.,0.,0.],timet,/plot'
     print,'elem_orbit,[1.,.5,0.,0.,0.,0.],timet, plot=2'
     print,'elem_orbit,[1.,.9,0.,0.,0.,0.],timet, plot=3'
     return
 endif
;choose scaling
 if(keyword_set(myy0)) then myy=myy0
 ****************
                                .
********
;orbital position vector R and velocity vector R are calculated
;for times defined in call
;rr and vv contain position and velocity components
   ntimet=n_elements(timet)
   rr=findgen(ntimet,3)
   vv=findgen(ntimet,3)
 for i=0,ntimet-1 do begin
   time=timet(i)
   elem_to_rv,elem,time,rad,vel,myy=myy
   rr(i,0)=rad(0)
   rr(i,1)=rad(1)
   rr(i,2)=rad(2)
   vv(i,0)=vel(0)
   vv(i,1)=vel(1)
   vv(i,2)=vel(2)
 endfor
if(keyword_set(plot)) then begin
   if(plot eq 1) then begin
      nwin
       plot,rr(*,0),rr(*,1),xtitle='x',ytitle='y',/iso,psym=3
       plots, 0, 0, psym=1
   if(plot gt 1) then begin
      oplot,rr(*,0),rr(*,1),col=plot,psym=3
   endi f
endif
end
```

```
Example output form elem_orbit.pro
```

Enter in IDL workspace the following lines:

```
timet=findgen(101)/100.*2*!pi & !x.range=[-2,2] & !y.range=[-2,2]
elem_orbit,[1.,.0,0.,0.,0.,0.],timet,/plot
elem_orbit,[1.,.5,0.,0.,0.,0.],timet, plot=2
elem_orbit,[1.,.9,0.,0.,0.,0.],timet, plot=3
```

This should make a plot:



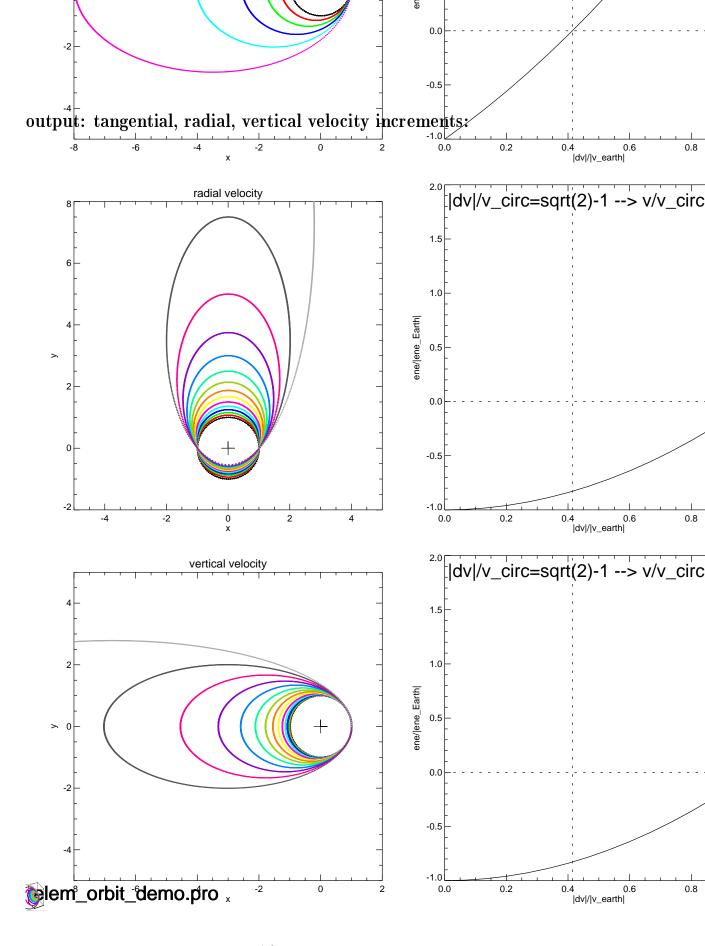
b) Satellite launched from Earth

- elem_orbit_demo_tangential.pro tangential velocity increments (max=30km/sec)
- elem_orbit_demo_radial.pro radial velocity increments
- elem_orbit_demo_vertical.pro vertical velocity increments

Just the first one is printed here, since the other two are almost similar

```
; elem_orbit_demo_tangential.pro
; use elem\_orbit to plot orbits with different elements
; elem = orbital elements [a,eks,ink,ome,w,tau]
; HS 05.11.02 /01.11.06
*****************
; Earth position (AU) and velocity (km/sec)
  assuming circular orbit with velocity 30 km/sec
     r_earth=[1.,0.,0.]
     v_earth= [0.,30.,0.]
  units: AU and year -> G(m1+m2)=4*pi^2
   orbital speed of Earth = 30 \text{ km/sec} = 2*pi \text{ AU/year}
   km/sec converted to au/year by multiplying with vscale=(2*pi)/30.
     myy=4.*!pi^2    ;yksikot AU, vuosi
vscale=(2.*!pi)/30.    ;muuntaa km/sec --> AU/vuosi
; give here different tangential velocity increments (norb different ones)
    norb=16
     dv=findgen(norb)*2
    norb=n_elements(dv)
     vx=v_earth(0)+dv*0.
     vy=v_earth(1)+dv
     vz=v_earth(2)+dv*0.
     title='tangential velocity increments, max='+string(max(dv))+'km/sec'
;1) calculate and store orbital elements
  rv_to_elem returns: elem=[a,eks,ink,ome,w,tau]
   also store the orbital energy (ene)
******************
     elem_tab=fltarr(6,norb) ; norb different orbits
     ene_tab=fltarr(norb)
     ene_earth=abs(myy/2.)
                           ; Maan energia
print,'Satellite launched from Earth orbit (v_cir=30km/sec)'
print,'-----
print,'dv_tan a eks ink ome w tau ene/|ene_earth|'
print,'--
ff='(f4.1,2x, 2f8.3, 4f8.1, 2x,f9.3)'
   for i=0,norb-1 do begin
      v=[vx(i),vy(i),vz(i)]
      r=r_earth
      v=v*vscale
;orbital elements
      rv_to_elem, 0., r, v, elem, myy=myy, ene=ene
      print,dv(i),elem,ene/ene_earth,f=ff
      elem_tab(*,i)=elem
      ene_tab(i)=ene
    endfor
```

```
***********
;2) calculate and plot orbits
**********************
  nwin
  timet=findgen(1000)/100. ; 0-9 years
  for i=0,norb-1 do begin
      elem=elem_tab(*,i)
      elem_orbit,elem,timet,rr,vv,myy=myy
      if(i eq 0) then begin
         plot,rr(*,0),rr(*,1),psym=3,xtitle='x',ytitle='y',$
             xr=[-5,5]-3, yr=[-5,5], xs=1, ys=1, /iso, title=title
         plots,0,0,psym=1
      endif
      if(i ne 0) then oplot,rr(*,0),rr(*,1),psym=3,col=i+1
   endfor
   xyouts,0.01,0.01,/normal,'elem_orbit_demo.pro',chars=.75
.*********************
;3) plot energy/|energy vs |dv|/|v_earth|
    plot,dv/30.,ene_tab/ene_earth,xtit='|dv|/|v_earth|',ytit='ene/|ene_Earth|'
;indicate the limit for hyperbolic orbit
    oplot,[0,1],[0,0],lines=1
    oplot,(sqrt(2)-1)*[1,1],[-1,2],lines=1
    xyouts,0.01,1.8,'|dv|/v_circ=sqrt(2)-1 --> v/v_circ=sqrt(2)'
    xyouts,0.01,0.01,/normal,'elem_orbit_demo.pro',chars=.75
end
```



dv_tan	a	eks	ink	ome	W	tau	ene/ ene_earth
0.0	1.000	0.000	0.0	180.0	0.0	0.0	-1.000
2.0	1.160	0.138	0.0	180.0	180.0	0.0	-0.862
4.0	1.398	0.284	0.0	180.0	180.0	-0.0	-0.716
6.0	1.786	0.440	0.0	180.0	180.0	0.0	-0.560
8.0	2.528	0.604	0.0	180.0	180.0	0.0	-0.396
10.0	4.500	0.778	0.0	180.0	180.0	0.0	-0.222
12.0	25.000	0.960	0.0	180.0	180.0	0.0	-0.040
14.0	6.618	1.151	0.0	180.0	180.0	-0.0	0.151
16.0	2.848	1.351	0.0	180.0	180.0	0.0	0.351
18.0	1.786	1.560	0.0	180.0	180.0	-0.0	0.560
20.0	1.286	1.778	0.0	180.0	180.0	-0.0	0.778
22.0	0.996	2.004	0.0	180.0	180.0	-0.0	1.004
24.0	0.806	2.240	0.0	180.0	180.0	0.0	1.240
26.0	0.674	2.484	0.0	180.0	180.0	-0.0	1.484
28.0	0.575	2.738	0.0	180.0	180.0	-0.0	1.738
30.0	0.500	3.000	0.0	180.0	180.0	0.0	2.000
RADIAL	VELOCITY	INCREME	ENTS				
dv_rad	a	eks	ink	ome	W	tau	ene/ ene_earth
0.0	1.000	0.000	0.00	180.00	0.00	0.00	-1.000
2.0	1.004	0.067	0.00	180.00	90.00	-0.23	-0.996
4.0	1.018	0.133	0.00	180.00	90.00	-0.21	-0.982
6.0	1.042	0.200	0.00	180.00	90.00	-0.20	-0.960
8.0	1.077	0.267	0.00	180.00	90.00	-0.19	-0.929
10.0	1.125	0.333	0.00	180.00	90.00	-0.17	-0.889
12.0	1.190	0.400	0.00	180.00	90.00	-0.16	-0.840
14.0	1.278	0.467	0.00	180.00	90.00	-0.15	-0.782
16 0	1 200	W E33	0 00	120 00	00 00	0 1 5	_0_716

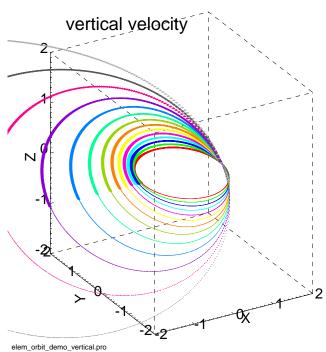
16.0	1.398	0.533	0.00	180.00	90.00	-0.15	-0.716
18.0	1.562	0.600	0.00	180.00	90.00	-0.14	-0.640
20.0	1.800	0.667	0.00	180.00	90.00	-0.13	-0.556
22.0	2.163	0.733	0.00	180.00	90.00	-0.13	-0.462
24.0	2.778	0.800	0.00	180.00	90.00	-0.12	-0.360
26.0	4.018	0.867	0.00	180.00	90.00	-0.12	-0.249
28.0	7.759	0.933	0.00	180.00	90.00	-0.11	-0.129
30.0	Infinity	1.000	0.00	180.00	90.00	-NaN	0.000

VERTICAL VELOCITY INCREMENT

dv_ve1	r a	eks	ink	ome	W	tau	ene/ ene_earth
0.0	1.000	0.000	0.00	180.00	0.00	0.00	-1.000
2.0	1.004	0.004	3.81	0.00	-0.00	-0.00	-0.996
4.0	1.018	0.018	7.59	0.00	-0.00	-0.00	-0.982
6.0	1.042	0.040	11.31	0.00	-0.00	0.00	-0.960
8.0	1.077	0.071	14.93	0.00	-0.00	0.00	-0.929
10.0	1.125	0.111	18.43	0.00	-0.00	0.00	-0.889
12.0	1.190	0.160	21.80	0.00	-0.00	0.00	-0.840
14.0	1.278	0.218	25.02	0.00	-0.00	0.00	-0.782
16.0	1.398	0.284	28.07	0.00	-0.00	-0.00	-0.716
18.0	1.562	0.360	30.96	0.00	-0.00	-0.00	-0.640
20.0	1.800	0.444	33.69	0.00	-0.00	-0.00	-0.556
22.0	2.163	0.538	36.25	0.00	-0.00	0.00	-0.462
24.0	2.778	0.640	38.66	0.00	-0.00	0.00	-0.360
26.0	4.018	0.751	40.91	0.00	-0.00	0.00	-0.249
28.0	7.759	0.871	43.03	0.00	-0.00	-0.00	-0.129
30.0	Infinity	1.000	45.00	0.00	-0.00	-NaN	0.000

An attempt to visualize 3D-orbits: (extract from elem_orbit_demo_vertical.pro)

```
;4) calculate and plot orbits, using 3d-plot
;use shade-surface to make the viewing transformation
;and draw a box with the size lim
 lim=2
 x=findgen(6)/5.*20
 surface,x#x,x,x,/save,xr=[-1,1]*lim,yr=[-1,1]*lim,zr=[-1,1]*lim,/nod,$
 xs=1,ys=1,zs=1,pos=[.15,.1,.85,.9],xtit='X',ytit='Y',ztit='Z',chars=1.5,$
 title=title
 x = lim^*[-1,1,1,-1,-1,1,1,-1] \ \& \ y = lim^*[-1,-1,1,1,-1,-1,1,1] \ \& \ z = lim^*[-1,-1,-1,-1,1,1,1]
 front=[0,1,2,3,0] & back=[4,5,6,7,4]
 1s=2
 plots,x(front),y(front),z(front),/t3d,/data,linestyle=ls
 plots,x(back),y(back),z(back),/t3d,/data,linestyle=2
 for i=0,3 do begin
   side=[i,i+4]
   plots,x(side),y(side),z(side),/t3d,/data,linestyle=2
 endfor
;calculate and plot orbits
 timet=findgen(1000)/100.
                           ; 0-9 years
for i=0,npart-1 do begin
  elem=elem_tab(*,i)
  elem_orbit,elem,timet,rr,vv,myy=myy
  plots,rr(*,0),rr(*,1),rr(*,2),psym=3,/t3d,col=i+1
  index=where(rr(*,2) gt 0, count)
  if(count gt 0) then plots,rr(index,0),rr(index,1),rr(index,2),psym=6,$
                /t3d,col=i+1,syms=.5
endfor
```



• Launch satellite at different angles (HARJ II.1) elem_rv_demo1.pro

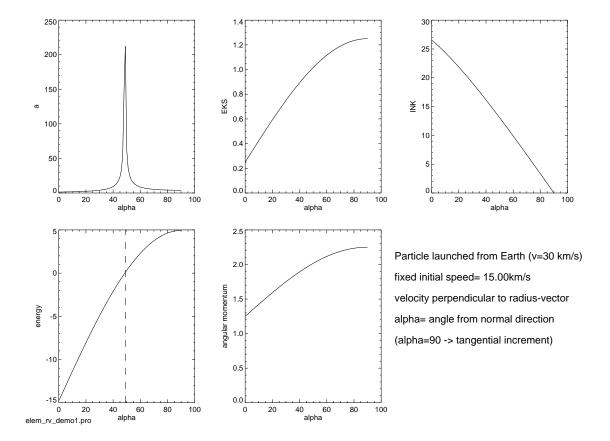
```
*********
; elem_rv_demo1.pro = HARJ II.1
; HS 05.11.02/01.11.06
**********
;Earth position (AU) and velocity (km/sec)
;assuming circular orbit with velocity 30 km/sec, f=45 degrees
  sinf=sin(45/!radeg)
  cosf=sin(45/!radeg)
  r_earth=[cosf,sinf,0.]
  v_{earth=[-sinf,cosf,0.]*30.}
*******************
;units: AU and year -> G(m1+m2)=4*pi^2
;orbital speed of Earth = 30 km/sec = 2*pi AU/year
; km/sec converted to au/year by multiplying with vscale=(2*pi)/30.
myy=4.*!pi^2
  vscale=(2.*!pi)/30.
;a) satellite launched from Earth, with velocity dv (km/sec),
; in perpendicular direction N
  dv=[0.,0.,15.]
  r=r_earth
  v=v_earth+dv
  v=v*vscale
  print,'PERPENDICULAR INITIAL VELOCITY'
  rv_to_elem, 0., r, v, elem, myy=myy,/print
;b) satellite launched from Earth, with velocity dv (km/sec),
  dv perpendicular to radius vector,
  with angle alpha to N and 90-alpha to V
  alpha=45.
  sina=sin(45/!radeg)
  cosa=sin(45/!radeg)
  vtan=30.+sina*15.
  vnor=cosa*15.
  v=[-sinf*vtan,cosf*vtan,vnor]
  r=r_earth
  v=v*vscale
  print,'-----'
  print,'PERPENDICULAR INITIAL VELOCITY, ALPHA=45'
  rv_to_elem,0.,r,v,elem,myy=myy,/print
;c) satellite launched from Earth, with velocity dv (km/sec),
  dv perpendicular to radius vector,
  with angle alpha to N
```

```
alpha=89.99
                              ; in the direction of velocity vector
  sina=sin(alpha/!radeg)
  cosa=cos(alpha/!radeg)
  vtan=30.+sina*15.
  vnor=cosa*15.
  v=[-sinf*vtan,cosf*vtan,vnor]
  r=r earth
  v=v*vscale
  print,'-----
  print, 'INITIAL VELOCITY in the direction of Earth velocity'
  print,'R [AU] =',r
print,'V [AU/year]=',v
                  =',r
  rv_to_elem, 0., r, v, elem, myy=myy, /print
·
;d) Calculate the same with different directions of the
;initial velocity alpha=0,1,...,90 degrees
;dv is the absolute value of velocity increment
dv=15.
;plot a,eks,ink, ene, amom as a function of alpha
;make arrays for alpha, a,eks,ink
;rv_to_elem returns: elem=[a,eks,ink,ome,w,tau]
  alpha_tab=findgen(91)*1.
  a_tab=alpha_tab*0.
  e_tab=alpha_tab*0.
  i_tab=alpha_tab*0.
  for i=0,n_elements(alpha_tab)-1 do begin
      alpha=alpha_tab(i)
      sina=sin(alpha/!radeg)
      cosa=cos(alpha/!radeg)
      vtan=30.+sina*dv
      vnor=cosa*dv
      v=[-sinf*vtan,cosf*vtan,vnor]
      r=r_earth
      v=v*vscale
;orbital elements
      rv_to_elem,0.,r,v,elem,myy=myy
      a_tab(i)=elem(0)
      e_tab(i)=elem(1)
      i_tab(i)=elem(2)
  endfor
; calculate also angular momentum and energy,
;take into account the sign for elliptic and hyperbolic
  k_tab=a_tab*abs(1.-e_tab^2)
  h_tab=-myy/(2*a_tab)
  index=where(e_tab gt 1,count)
  if(count ge 1) then h_tab(index)=-h_tab(index)
  !p.multi=[0,3,2]
  nwin
  plot,alpha_tab,a_tab,xtitle='alpha',ytitle='a'
  plot,alpha_tab,e_tab,xtitle='alpha',ytitle='EKS'
  plot,alpha_tab,i_tab,xtitle='alpha',ytitle='INK'
plot,alpha_tab,h_tab,xtitle='alpha',ytitle='energy'
  oplot,alpha_tab(index(0))*[1,1],[-30,30],lines=2
```

plot,alpha_tab,k_tab,xtitle='alpha',ytitle='angular momentum'

```
xyouts,0.67,0.40,/normal, 'Particle launched from Earth (v=30 km/s)',chars=0.7
xyouts,0.67,0.35,/normal,'fixed initial speed='+string(dv,'(f6.2)')+'km/s',chars=0.7
xyouts,0.67,0.30,/normal, 'velocity perpendicular to radius-vector',chars=0.7
xyouts,0.67,0.25,/normal, 'alpha= angle from normal direction',chars=0.7
xyouts,0.67,0.20,/normal, '(alpha=90 -> tangential increment)',chars=0.7

xyouts,0.01,0.01,'elem_rv_demo1.pro',/normal,chars=.5
!p.charsize=1
charsize,1
!p.multi=0
end
```



```
PERPENDICULAR INITIAL VELOCITY
CALCULATED ELEMENTS OF ELLIPTIC ORBIT: at T= 0.00000
      1.3333330
eks =
        0.25000000
        26.565052
45.000000
0.0000000
ink =
ome =
w =
tau = -0.00018333057
per =
          1.5396001
                      0.042867628
         0.99999977
q =
PERPENDICULAR INITIAL VELOCITY, ALPHA=45
CALCULATED ELEMENTS OF ELLIPTIC ORBIT: at T= 0.00000
           23.313564
eks =
          0.95710689
      14.638788
45.000000
0.0000000
ink =
ome =
w =
tau = -0.00067585304
                      0.0021614330
          112.56749
per =
q =
        0.99999131
INITIAL VELOCITY in the direction of Earth velocity
R [AU] = 0.707107 0.707107
              -6.66432
                          6.66432 0.000548141
V [AU/year]=
CALCULATED ELEMENTS OF HYPERBOLIC ORBIT: at T= 0.00000
a = 4.0000027
         1.2499998
eks =
ink =
           0.0000000
          45.000000
ome =
w = 0.0000000
tau = -9.3926329e-05
                      0.0042266808
per =
          8.0000078
          0.99999972
q =
```

• Evolution of a swarm of satellites | elem_rv_demo2.pro

```
•
;elem_rv_demo2.pro
*******
;plot the evolution of a swarm of particles launched from Earth
......
:Earth position (AU) and velocity (km/sec)
; assuming circular orbit with velocity 30 km/sec, f=0
************************************
 r_earth=[1.,0.,0.]
 v_earth= [0.,30.,0.]
;initial launch speed dv, in different directions
;in the plane of Earth's orbit, angle is the direction with
;respect to x-axis
;try different values!
 dv=5
 pscale=3
; dv=1.
;pscale=2.
; choose the number of particles,
;corresponding launching angles
 npart=3600*.5
 angle_tab=findgen(npart)*360./npart
**********************************
;units: AU and year -> G(m1+m2)=4*pi^2
;orbital speed of Earth = 30 km/sec = 2*pi AU/year
;km/sec converted to au/year by multiplying with vscale=(2*pi)/30.
***********************************
 myy=4.*!pi^2
 vscale=(2.*!pi)/30.
;1) calculate and store orbital elements
  rv_to_elem returns: elem=[a,eks,ink,ome,w,tau]
elem_tab=fltarr(6,npart) ; npart different particles
 nwin
 for i=0,n_elements(angle_tab)-1 do begin
    angle=angle_tab(i)
    sina=sin(angle/!radeg)
    cosa=cos(angle/!radeg)
    v=v_earth+[cosa,sina,0]*dv; rad,tan
    r=r_earth
    v=v*vscale
;orbital elements
    rv_to_elem, 0., r, v, elem, myy=myy
    elem_tab(*,i)=elem
 endfor
· **************
;2) start calculating positions
; plot for different times
*************
;store positions to arrays
 xpos=angle_tab*0.
 ypos=angle_tab*0.
;plotting times
```

```
ntimes=1000
  times=findgen(ntimes)*.05
; ntimes=9
; times=findgen(ntimes)*.25+.25
  for j=0,ntimes-1 do begin
      time=times(j)
;earth position (circular orbit)
      x_earth=cos(2.*!pi*time)
      y_earth=sin(2.*!pi*time)
      for i=0,n_elements(angle_tab)-1 do begin
          elem=elem_tab(*,i)
          elem_to_rv,elem,time,rad,vel,myy=myy
          xpos(i)=rad(0)
          ypos(i)=rad(1)
      endfor
      \verb|plot,xpos,ypos,psym=3,xr=[-1,1]*|pscale,yr=[-1,1]*|pscale,xs=1,ys=1,/iso,\$|
        title='dv='+string(dv,'(f6.2)')+'km/sec T/PER ='+string(time,'(f8.3)')
      plots,x_earth,y_earth,psym=1,symsize=1
      plots,0,0,psym=2,symsize=1
;if you want to stop between plots - uncomment the next lines
; answ=''
; read, answ
 \quad \hbox{endfor} \quad
                                 ;end loop over times
end
```

