

University of Oulu Astronomy and Space Physics Jürgen Schmidt Oulu, 30.09.2015 Celestial Mechanics II, Fall 2015 jurgen.a.schmidt@oulu.fi

Exercise Problems — Part 2

(Due Date: Tuesday, 06.10.15, before the lecture. You can also put the solutions in the red box outside office FY277.)

1. Atmospheric Drag

- (a) Check from a source of your choice what is the structure of the *upper* part of Earth's atmosphere. Describe briefly the different parts and altitude ranges. What densities are expected there, roughly?
- (b) Derive the equation given in the lecture

$$\frac{\mathrm{d}e}{\mathrm{d}t} = -\frac{C_{\mathrm{d}} A \rho}{m} v \left(e + \cos f\right) .$$

for the slow evolution of the eccentricity under the influence of atmospheric drag.

(c) Solve the equation for the slow evolution of the semi-major axis

$$\frac{\mathrm{d}a}{\mathrm{d}t} = -\frac{C_\mathrm{d}\,A\,\rho}{m}\frac{a^2}{\mu}\,v^3$$

for the case of constant density ρ and e = 0.

2. Advanced: Slow evolution of the longitude of the ascending node

Show that the general evolution equation for the longitude of the ascending node reads

$$\frac{\mathrm{d}\Omega}{\mathrm{d}t} = \sqrt{\frac{a(1-e^2)}{\mu}} N \frac{\sin\varphi}{1+e\cos f}.$$

Hint: Use the relations given in the lecture and express the longitude of the ascending node in terms of components of the angular momentum vector as

$$\tan\Omega = -\frac{h_x}{h_y}.$$

Also, you must express the unit vectors normal to the orbital plane and in azimuthal direction, \vec{e}_N and \vec{e}_{φ} , respectively, in terms of Ω , i, and $\varphi = \varpi + f$ as

$$ec{e}_N = \left(egin{array}{c} \sin i \, \sin \Omega \\ -\sin i \, \cos \Omega \\ \cos i \end{array}
ight), \quad and \quad ec{e}_{arphi} = \left(egin{array}{c} -\sin arphi \, \cos \Omega - \cos i \, \sin \Omega \, \cos arphi \\ -\sin arphi \, \sin \Omega + \cos i \, \cos \Omega \, \cos arphi \\ \sin i \, \cos arphi \end{array}
ight).$$

Keep in mind that the angular momentum vector points in the direction of \vec{e}_N .

1. Samophie Irag

(a) Exosphere - Foolan to 10,000 km mobiules on ballistie hajectrices, no collision, mactically

Pr10-14 kg

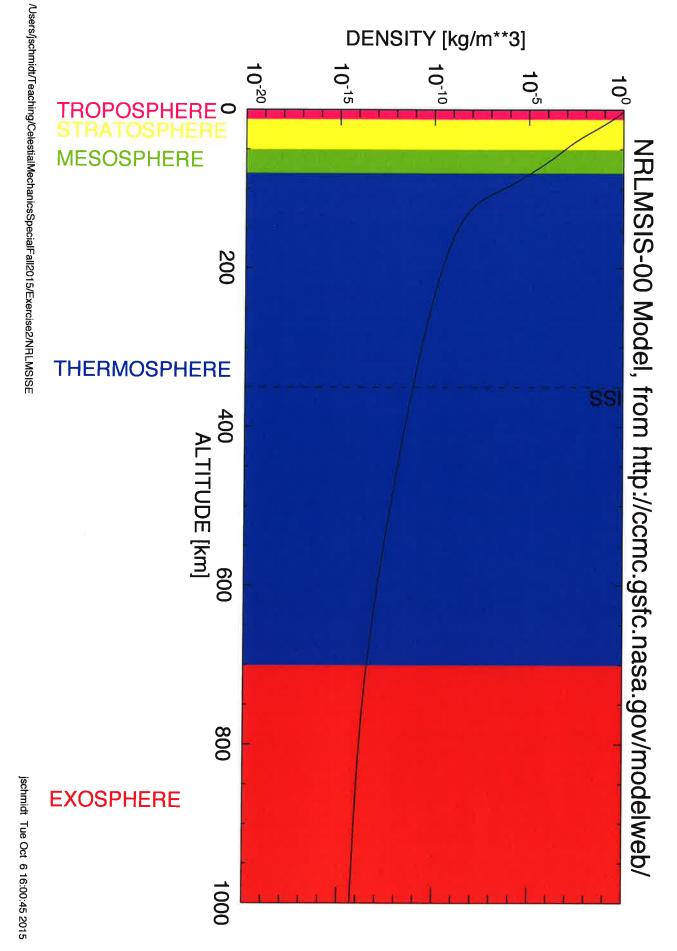
Throughphere - 80 lan to too too too too land of solar ochings

Perosphere - Solan to 80 lan, to high for space and to low population

Stratosphere 12kus to Solan;

Troposphere - Oku to 12 lan

citains about 80% of man of the



(3)

For attumphic drag (becture):

$$D = -\frac{\text{cd AP}}{2m} v \frac{h}{a(1+2)} = mif$$

$$T = -\frac{\text{cd AP}}{2m} v \frac{h}{n}$$

$$S=aaut; e=0$$

$$= > V = Ra = \sqrt{\frac{u}{q}}$$

=)
$$\frac{da}{dt} = -\frac{cdAP}{m}\frac{a^2}{a}\left(\frac{u}{a}\right)^{\frac{3}{2}}$$

(o.g., witel revineja akos)

$$\frac{dk}{dx} = -dT$$

$$\Rightarrow 2 d \left[x = -dT \right]$$

$$\Rightarrow 3 d \left[x + \frac{1}{2} \right] = 0$$

$$\Rightarrow \sqrt{x} + \frac{1}{2} = aut$$

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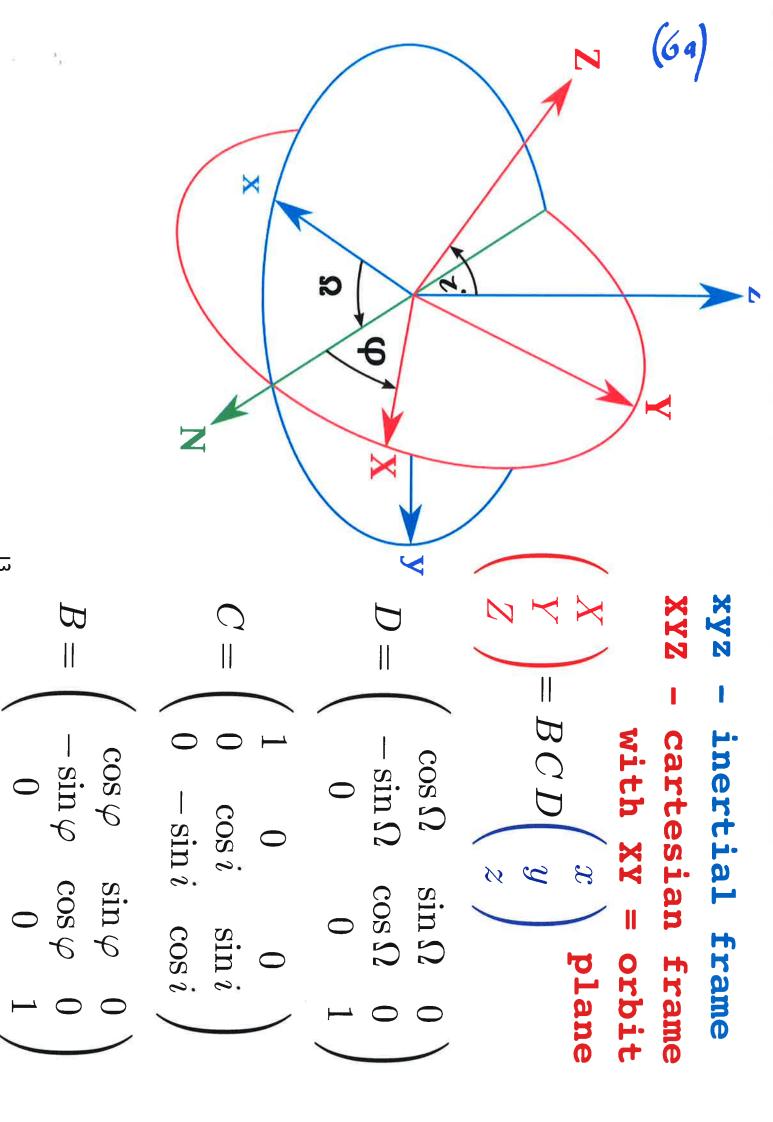
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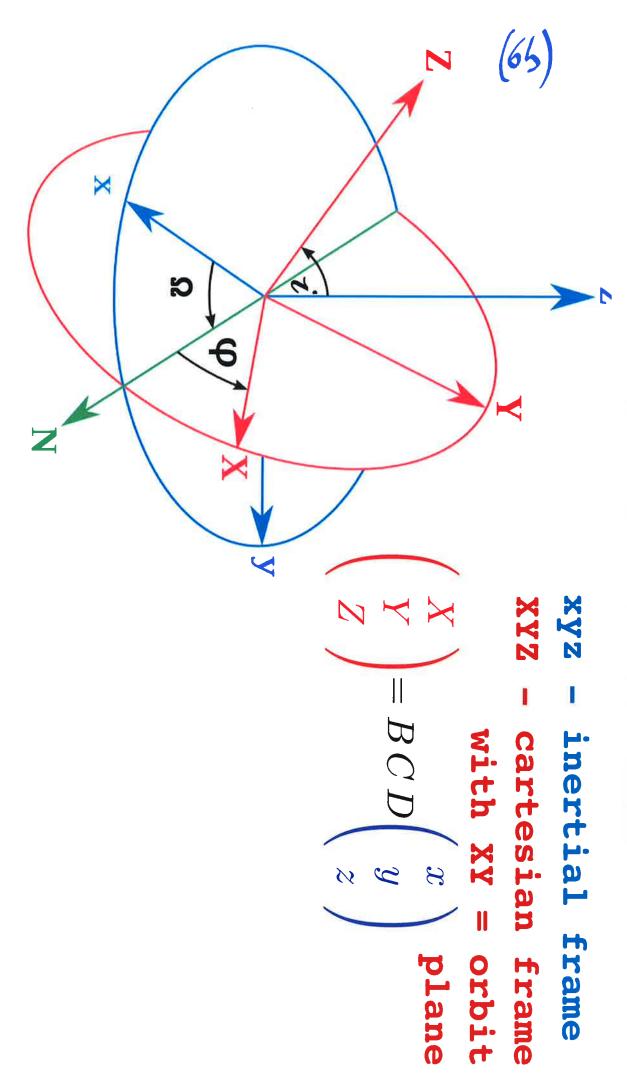
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$$\Rightarrow d \left[x + \frac{$$



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B =



BCD =

 $-\sin\varphi\cos\Omega-\cos i\sin\Omega\cos\varphi$

 $-\sin\varphi\sin\Omega + \cos i\cos\Omega\cos\varphi$

 $\sin i \cos \varphi$

 $\cos i$

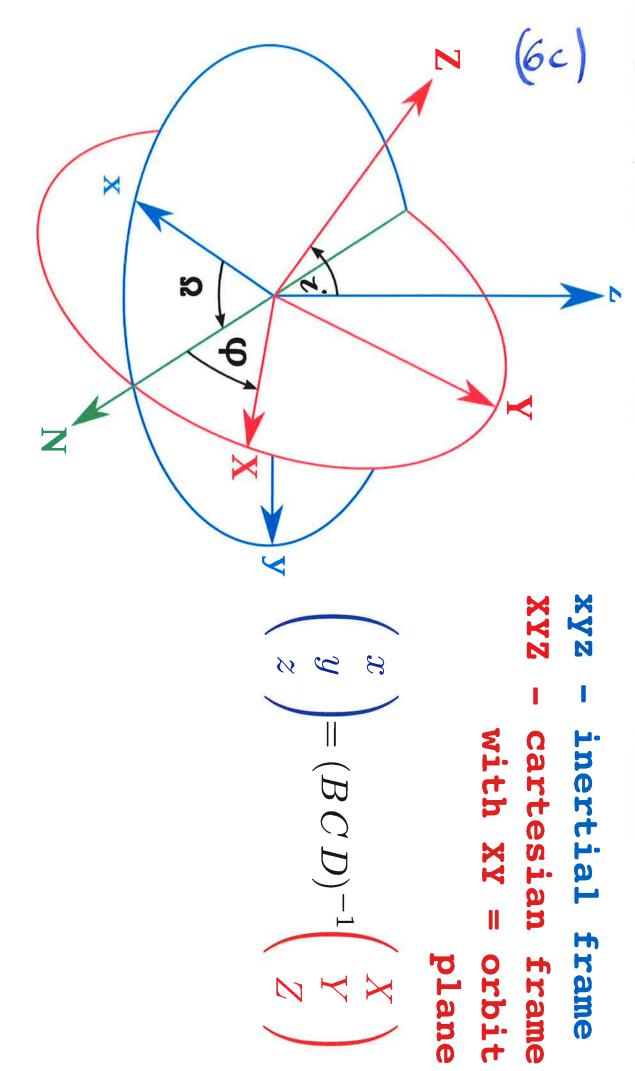
 $\sin i \sin \varphi$

 $-\sin i\cos\Omega$

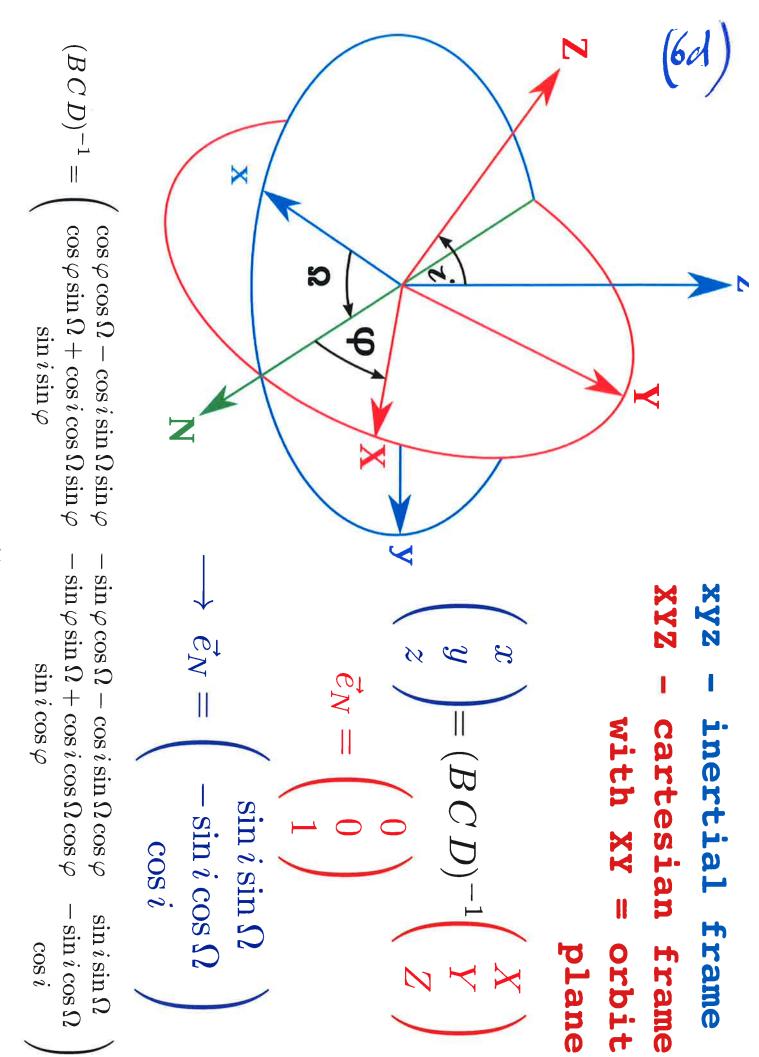
 $\cos \varphi \sin \Omega + \cos i \cos \Omega \sin \varphi$

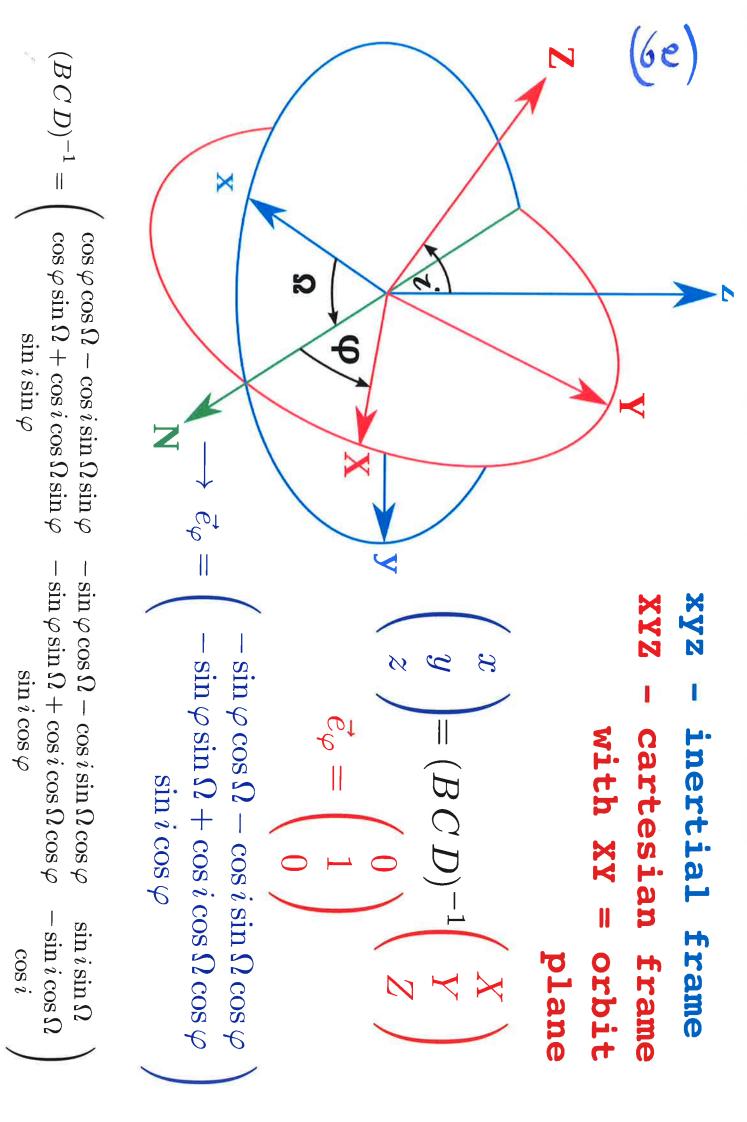
 $\sin i \sin \Omega$

 $\cos \varphi \cos \Omega - \cos i \sin \Omega \sin \varphi$



 $(BCD)^{-1} =$





hx = hen; hy = hen; from h = hen

$$e^{x} = -\sin i \cos \Omega + e^{x} - \sin i \sin \Omega$$

$$e^{x} = -\sin i \cos \Omega - \sin i \Omega \text{ and}$$

$$e^{y} = -\sin i \sin \Omega + \sin \Omega \cos \Omega \text{ conf}$$

$$e^{y} = -\sin i \sin \Omega + \sin \Omega \cos \Omega \cos \Omega$$

=
$$\frac{dR}{dt} = \frac{co^{2} \Omega}{h \sin^{2}(co\Omega)} + \frac{1}{\sin^{2}(co\Omega)} + \frac{1}{\cos^{2}(co\Omega)} + \frac{1}{\cos^$$

$$= \int d\Omega = \int \frac{\alpha(1-e^{-1})}{\alpha} \sqrt{\frac{ni}{1+ecor}}$$