

Exercise Problems — Part 6, second batch

(Due Date: Tuesday, 24.11.15, *before* the lecture. For the computer exercise send me the IDL code per mail.)

2. Dynamics in the field of an axisymmetric planet — cnt'd

(a) Show that from the potential of an axisymmetric mass, including the effect of the J_2 term

$$\phi(\vec{r}) = -\frac{\mu}{r} \left[1 - \frac{1}{2} J_2 \left(\frac{R_p}{r} \right)^2 (3 \cos^2 \theta - 1) \right],$$

we get the acceleration

$$\ddot{\vec{r}} = -\vec{\nabla} \phi(\vec{r}) = -\frac{\mu}{r^2} \frac{\vec{r}}{r} + \frac{3}{2} J_2 \frac{\mu}{r^2} \left(\frac{R_p}{r} \right)^2 \left[\left(5 \frac{z^2}{r^2} - 1 \right) \frac{\vec{r}}{r} - 2 \frac{z}{r} \vec{e}_z \right].$$

Here, R_p is the (equatorial) radius of the planet and $\cos \theta = z/r$.

(d) *Advanced:* Derive from the force in part (a) the conserved energy.