

13. Additional programs-utilities

In addition to the programs that were described separately, a number of short utility programs are available that assist in the processing and managing of the data.

BOUND_BSW

Description:	converts the close-coupling <i>B</i> -spline expansions to the <i>c</i> - and <i>bsw</i> -files
Input files:	bound.nnn , bound_bsw.inp
Output files:	<i>c</i> - and <i>bsw</i> -files for given states
Call:	bound_bsw < bound_bsw.inp

The results of the bound-state *B*-spline calculations are recorded in the **bound.nnn** files in the form of the corresponding close-coupling expansions (7.3). They are similar to the *l*- or *j*-files in the MCHF complex, and contain descriptions of all states for a given total term. The BOUND_BSW utility collects the information for a given bound state indicated in the file **bound_bsw.inp**, and records it as pair **name.c** and **name.bsw**. These files can be used as input target files for the BSR scattering calculations.

BOUND_TAB

Description:	produces the total list of resulting bound states
Input files:	bound.nnn
Output files:	bound.tab
Call:	bound_tab

This program sorts the energy levels in different **bound.nnn** files and prints them in atomic units, eV or cm⁻¹, relative to the lowest state. The user can define the range of partial waves and the energy range to restrict the output.

BSW_W

Description:	converts the <i>bsw</i> -file to the <i>w</i> -file
Input files:	name.bsw
Output files:	name.w
Call:	bsw_w name.bsw

CFILE

Description:	extracts the given state from <i>l</i> - or <i>j</i> -file in separate <i>c</i> -files
Input files:	name.l or name.j
Output files:	result.c
Call:	cfile name.l(j) nn 2J result.c eps_c

In the BSR complex, each target state is described by a pair **name.c** and **name.bsw**. The CI program from the MCHF complex produces the results in the *l(j)*-files as a list of states. The CFILE utility allows one to extract the expansion coefficients for a given state and create the corresponding *c*-file. The user should indicate in the command line the name of the input *l*- or *j*-file, the pointer for the given state, *nn*, the value of *2J* if it is a *j*-file, the name for the output *c*-file, and the tolerance for the expansion coefficients, *eps_c*.

PHOTO_TAB

Description:	produces the tables for channel or total cross sections and asymmetry parameters
Input files:	bsr_phot.nnn
Output files:	indicated by user
Call:	photo_tab with interactive response

This utility serves for processing the data recorded during photoionization calculations with program BSR_PHOT. The program BSR_PHOT can be run for different energy intervals and different partial waves separately. The results are accumulated in files **bsr_phot.nnn**. Then final tables for channel or total cross sections and asymmetry parameters may be generated with utility **photo_tab**. More details are given in the description of the BSR_PHOT program.

SLATER_W

Description:	transfers Slater-type orbitals to MCHF format
Input files:	name.slw
Output files:	name.w
Call:	slater_w name.slw

The Slater orbitals are supposed to be given as in the RMATRX1 code, i.e.,

$$P_{nl}(r) = \sum_{j=1}^{NCO} C(j) * r^{IRAD(j)} * e^{-ZE(j)*r}.$$

The formatted file name.slw contains the Slater-type orbital parameters in the form:

1. `Z, atom, term` - nuclear charge, and optional identification for atom and term
2. `nwf` - number of radial wave functions
3. `n, l, k` - quantum numbers for given orbital
4. `NCO`
5. `IRAD(1:NCO)`
6. `ZE(1:NCO)`
7. `C(1:NCO)`

Repeat records 3-7 `nwf` times

SUM_HH

Description:	merges the set of h.nnn files to the final H.DAT
Input files:	h.nnn
Output files:	h.dat
Call:	sum_hh klsp1 klsp2

W_BSW

Description:	converts the <i>w</i> -file to the bsw-file format
Input files:	name.w
Output files:	name.bsw
Call:	bsw_w name.w