LECTURE - 1 ,2

$$\begin{array}{c} \text{SECTION} & C & & \text{(Lecture-1)} \\ \text{SECTION} & C & & \\ \text{SECTION} & D & & \\ \text{SECTION} & C & & \\ \text{SECTION} & C & & \\ \text{SECTION} & D & & \\ \text{SECTION} & D & & \\ \end{array}$$

Classical Mechanics -> 14 Lecture Hours and Electrodynamics

Mechanics of Many-body Systems

→ 1-2 Lecture Hours

-> Lecture notes will be uploaded periodically on MIS

Lagrangian and Hamiltonian Equations

→ 3-6 Lecture Hours

→ Topics covered in Class are very important for exams.

Electrodynamics

-> Maxwell's Equations

- wave equation

-> Energy density, Poynting's Theorem

→ 6 - 7 Lecture Hours

- Also follow the lecture plan provided on MIS for details about Marks and weightage distribution.

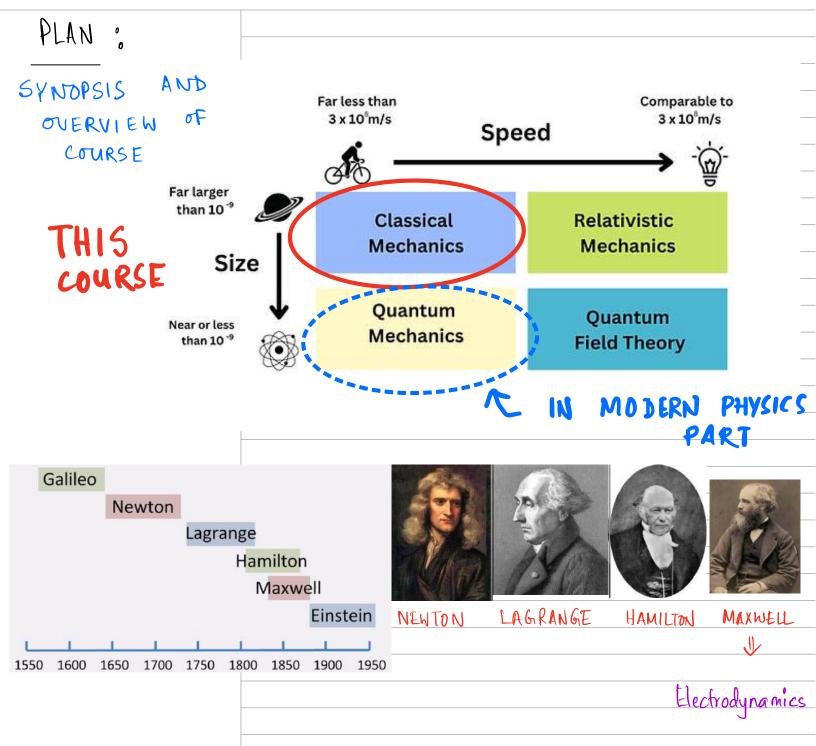
[LECTURE PLAN FOLDER]

LECTURES SHALL ALSO INCLUDE PROBLEM SOLVING -> PROBLEM SOLVING SESSIONS COULD PROVE VERY BENEFICIAL

TEXTBOOK NAMES SHALL

BE MENTIONED DURING

LECTURE:



1 Why study Classical Mechanics?

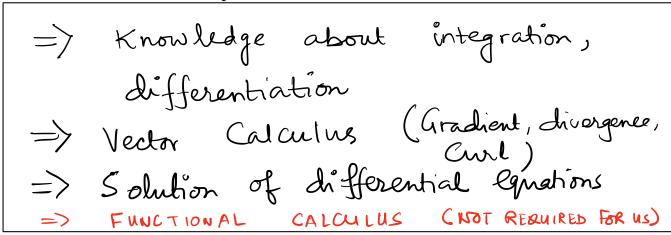
Newton's laws date back to 1687, when he published Philosophiae Naturalis Principia Mathematica, laying out his laws of motion and universal gravitation. Now, over 300 years later, we understand the world in terms of relativistic quantum field theory or even fundamental string. Classical mechanics is known to fail in at least three ways. At distances smaller than $\frac{\hbar}{mv}$, we must use quantum mechanics or quantum field theory to describe the motions and interactions of matter. Second, at velocities close to the speed of light we must make relativistic corrections, working in a unified spacetime instead of Newton's abstraction of ideal Euclidan space and universal time. Finally, the law of universal gravitation has been replaced by general relativity. Why then, study classical mechanics at all?

1.1 Range of applicability

One answer lies in the wide range of applicability. Planck's constant \hbar is small, so quantum considerations usually only become important for phenomena on the scale of atoms whose size is determined by $\frac{\hbar}{mv}$ for the orbiting electrons. Similarly, special relativity gives significant corrections only when objects move at a substantial fraction of the speed of light. The fastest man-made object ever produced was not the Voyager spacecraft (35,000 mi/hr) as is often claimed, but the solar probes Helios-A and Helios-B which reached a maximum speed of 252,792 km/hr (see http://en.wikipedia.org/wiki/Helios_(spacecraft)). This is still only 0.000234c, with c the speed of light, so the relativistic corrections $\left(\sim \frac{v^2}{c^2}\right)$ are only a few parts in 10^8 . Interestingly, second place appears to be held by a nuclear powered manhole cover (nuclear testing, Pascal B, gone wrong) (see http://savvyparanoia.com/the-fastest-man-made-object-ever-a-nuclear-powered-manhole-cover-true/) which would have been traveling at about 237,500 mph. This still is only $\frac{v}{c} = 2.2 \times 10^{-4}$. These small corrections are important only for extremely fine measurements, where they are easily measured by modern atomic clocks. Finally, general relativity is important for cosmology, precise orbit predictions in planetary and satellite science, and the GPS system, but errors using Newtonian gravity are of order $\frac{GM}{Rc^2}$, where R is the distane from a mass M. This is of order 6.95×10^{-7} near the surface of Earth.

Therefore, for sizes larger than atoms and smaller than the solar system, and ordinary velocities, classical mechanics is an excellent approximation.

1.2 Mathematical techniques



1.3 First approximation (NEWTONIAN MEC)

Because the corrections to Newtonian mechanics are so small, the Newtonian solution to problems is close to the exact solution, and therefore makes a good place to start in making a perturbative approximation to the full solution. Alternatively, if we have an exact solution in general relativity or quantum mechanics, we may be able to make sense of it by comparing terms in the classical solution.

1.4 Intuition

We have a great deal of direct experience with the world, and the terms of classical mechanics line up well with this experience. We can use this familiarity to guess how a system will behave. With more precise theories, having a similar picture of what is going on becomes difficult.

2 Review of Newtonian Mechanics

Basic definitions

We define several important concepts. We picture the world as a 3-dimensional Euclidean space with points labeled by triples of numbers, often simply the Cartesian (x, y, z), but others as well. Events are parameterized by the passage of time, so a point particle is described by a curve

$$\mathbf{r}(t) = (x(t), y(t), z(t))$$
$$= (r(t), \theta(t), \varphi(t))$$

in whatever coordinates we choose, with boldface denoting a vector. Notice that the position vector $\mathbf{r}(t)$ is a *dynamical variable*, not a coordinate. As t varies, $\mathbf{r}(t)$ traces out a curve in space.

The time-rate-of-change of the position vector is called the *velocity*,

$$\mathbf{v}(t) = \frac{d\mathbf{r}(t)}{dt}$$
$$= \dot{\mathbf{r}}(t)$$

and is tangent to this curve. Notice that for simplicity we will sometimes denote the time derivative with a dot over the variable. The *acceleration* is the rate of change of velocity,

$$\mathbf{a}(t) = \frac{d\mathbf{v}(t)}{dt}$$
$$= \frac{d^{2}\mathbf{r}(t)}{dt^{2}}$$
$$= \ddot{\mathbf{r}}(t)$$

Particles are characterized by a constant called the mass, m, which reflects their resistance to change of velocity. Defining the momentum as the product

$$\mathbf{p} = m\mathbf{v}$$

we write Newton's second law as

$$\mathbf{F} = \frac{d\mathbf{p}\left(t\right)}{dt}$$

The force, \mathbf{F} , is to be taken intuitivelty and is determined by the particular problem. It is essentially that effort which produces a change of momentum. For example, if we stretch a spring it has the ability to move a mass attached to the end. Since this ability doubles if we double the stretch of the spring, we may write the force of a spring as proportional to the difference between the location of its endpoints,

$$\mathbf{F}_{spring} = -k\left(\mathbf{r}_2 - \mathbf{r}_1\right)$$

where the spring constant, k, characterizes the strength of the spring. This form of force is Hooke's Law.

There are many forces that have been identified:

$$\begin{array}{ccc} 0 & Zero\,force \\ -k\,(\mathbf{r}_2-\mathbf{r}_1) & Hooke's\,law \\ q\,(\mathbf{E}+\mathbf{v}\times\mathbf{B}) & Lorentz\,force \\ \mathbf{N} & Normal\,force \\ -\mu\mathbf{N} & Friction \\ -\frac{GMm}{r^2}\hat{\mathbf{r}} & Gravitation \\ -\frac{kQq}{Q^2}\hat{\mathbf{r}} & Coulomb's\,law \end{array}$$

Once the forces on a particle have been identified, Newton's second law becomes an ordinary, second order differential equation

$$\sum \mathbf{F} = \frac{d}{dt} \left(m \frac{d\mathbf{r}(t)}{dt} \right)$$

where the sum is over all forces on the particle. This means that two initial conditions are required to give a unique solution to a problem. If the time starts at $t = t_0$, then the initial conditions may be taken as the position and velocity at t_0 ,

$$\mathbf{r}_0 = \mathbf{r}(t_0)$$

 $\mathbf{v}_0 = \mathbf{v}(t_0)$

Conservation laws

We now develop three conservation laws.

Conservation of linear momentum

If no force acts, $\mathbf{F} = 0$ and we have

$$\frac{d\mathbf{p}\left(t\right)}{dt} = 0$$

Integrating gives

$$\mathbf{p}\left(t\right) = p_0 = m\mathbf{v}\left(t_0\right)$$

The linear momentum is therefore constant; we say that \mathbf{p} is *conserved*.

Conservation of angular momentum

We define the angular momentum of a particle at position $\mathbf{r}(t)$, relative to a fixed position \mathbf{R} , to be

$$\mathbf{L} = (\mathbf{r}(t) - \mathbf{R}) \times \mathbf{p}$$

and the torque about the same location to be

$$\mathbf{N} = (\mathbf{r}(t) - \mathbf{R}) \times \mathbf{F}$$

Then taking the cross product of the relative position, $\mathbf{r}(t) - \mathbf{R}$, with Newton's second law, we have

$$(\mathbf{r}(t) - \mathbf{R}) \times \mathbf{F} = (\mathbf{r}(t) - \mathbf{R}) \times \frac{d\mathbf{p}(t)}{dt}$$

$$\mathbf{N} = \frac{d}{dt} ((\mathbf{r}(t) - \mathbf{R}) \times \mathbf{p}(t)) - \left(\frac{d}{dt} (\mathbf{r}(t) - \mathbf{R})\right) \times \mathbf{p}(t)$$

where we use the product rule on the right. Since

$$\frac{d}{dt} (\mathbf{r}(t) - \mathbf{R}) = \frac{d\mathbf{r}(t)}{dt} - \frac{d\mathbf{R}}{dt}$$
$$= \mathbf{v}(t) - 0$$

and with $\mathbf{p}(t) = m\mathbf{v}(t)$, the right side becomes

$$\frac{d}{dt}\left(\left(\mathbf{r}\left(t\right)-\mathbf{R}\right)\times\mathbf{p}\left(t\right)\right)-\left(\frac{d}{dt}\left(\mathbf{r}\left(t\right)-\mathbf{R}\right)\right)\times\mathbf{p}\left(t\right) = \frac{d\mathbf{L}}{dt}-\mathbf{v}\left(t\right)\times m\mathbf{v}\left(t\right)$$

$$=\frac{d\mathbf{L}}{dt}$$

$$\Rightarrow \overrightarrow{N} = \overrightarrow{L}$$
Similarly with $NL\overrightarrow{ll}$ for an angular system

since $\mathbf{v}(t) \times \mathbf{v}(t) = 0$. We therefore have

$$\mathbf{N} = \frac{d\mathbf{L}}{dt} \qquad =) \quad \overrightarrow{\mathbf{L}} = \overrightarrow{\mathbf{O}}, \quad \text{then } \overrightarrow{\mathbf{L}} = \overrightarrow{\mathbf{O}} =) \quad \overrightarrow{\mathbf{L}} = const.$$

It follows immediately that L is conserved if the torque vanishes.

Conservation of energy

Angular momentum is conserved if no torque acts!

Suppose the force on a particle is a function of particle position, $\mathbf{F} = \mathbf{F}(\mathbf{r})$. Then we may integrate the second law by taking the dot product with the velocity:

$$\mathbf{F} = \frac{d\mathbf{p}(t)}{dt}$$

$$\mathbf{F} \cdot \mathbf{v} = \frac{d\mathbf{p}}{dt} \cdot \mathbf{v}$$

$$\mathbf{F} \cdot \frac{d\mathbf{r}}{dt} = m\frac{d\mathbf{v}}{dt} \cdot \mathbf{v}$$

$$\mathbf{F} \cdot d\mathbf{r} = m\mathbf{v} \cdot d\mathbf{v}$$

Integrating from the initial values to the values at a general time, t, we have

$$\int_{\mathbf{r}_0}^{\mathbf{r}(t)} \mathbf{F} \cdot d\mathbf{r} = m \int_{\mathbf{v}_0}^{\mathbf{v}(t)} \mathbf{v} \cdot d\mathbf{v}$$
$$= \frac{1}{2} m \mathbf{v}^2 - \frac{1}{2} m \mathbf{v}_0^2$$
$$= T - T_0$$

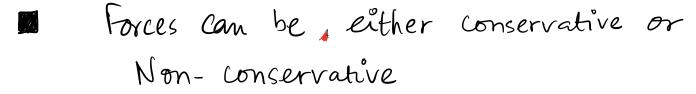
where we define the kinetic energy, $T(t) = \frac{1}{2}m\mathbf{v}^2$. In general, the integral on the left side depends on the path of integration. Such a path-dependent integral is not a function, but is called instead a functional. Along any path we define the work as

$$W_{12} = \int_{\mathbf{r}_1}^{\mathbf{r}_2} \mathbf{F} \cdot d\mathbf{r}$$

so that the work is equal to the change in kinetic energy,

$$W_{12} = T_2 - T_1$$

This is the work-energy theorem.



START OF LECTURE - 2

Conservation of Energy Work done is FORCE times DISTANCE moved in the direction of the force. [Work] _ [F] [&s] $= MLT^{-2}L$ $= ML^2T^{-2}$ = Joule Constant mass, m:

$$\overrightarrow{F} = \overrightarrow{ma}$$

$$W_{12} = \overrightarrow{m} \qquad \overrightarrow{0} \qquad \overrightarrow{a} \cdot \overrightarrow{dr}$$

$$= \overrightarrow{m} \qquad \overrightarrow{0} \qquad \overrightarrow{dr} \qquad \overrightarrow{dr}$$

$$= m \qquad \overrightarrow{0} \qquad \overrightarrow{dr} \qquad \overrightarrow{$$

Another definition:

=) Conservative: No mechanical energy
destroyed or dissipated.

Energy can be simply
written as, E=KE
of path.

e.g. gravitation, electrostatics, Hooke's law for
springs

=) Non-conservative: mechanical energy

=> Non-conservative: mechanical energy
is dissipated.

=> Work depends on path taken.

eg. Friction, magnetism (Lorentz force)

BUT: All Fundamental Forces are conservative!

Q. WHAT ARE THE CONSEQUENCES OF PATH INDEPENDENT CONSERVATIVE FORCES?

ANSWER: Scalar Potential

Path independence of work done by a conservative force implies the existence of a scalar potential.

Let's poore the following as well: Prove that if f is vindependent of the Ahro. Math. of the path joining any two points P, and P₂ in a given region, then & Food of all closed paths in the region and conversely. Closed 5 Let P, A P2 BP, be a Curre. Then, & Fodr = (F.dr) Hence, &F.dF PIA P2 BP1 = SF.dr+SF.dr as P. d. is independent of path PIAP2 P2BP1 $\int_{P_1AP_2} \overrightarrow{F} \cdot d\overrightarrow{r} - \int_{P_1BP_2} \overrightarrow{F} \cdot d\overrightarrow{r}$

Since the integral from P, to P2 through 'A' is same as the integral Inrough B'.

[Recall hypothesis that SF.d. is

midependent of path P1 joining

P1, P2. J Conversely, it (F). dF = 0, then $\Rightarrow \int \overrightarrow{F} \cdot d\overrightarrow{r} = \int \overrightarrow{F} \cdot d\overrightarrow{r} + \int \overrightarrow{F} \cdot d\overrightarrow{r} = 0$ $P_{1}AP_{2}BP_{1}$ $P_{2}AP_{2}BP_{1}$ $= \int_{A} \vec{F} \cdot d\vec{r}$ $= \int_{A} \vec{F} \cdot d\vec{r}$ Implies :

W12 = \int \vec{F} \cdot \vec{dr} = \int \vec{F} \cdot \vec{dr} = 0

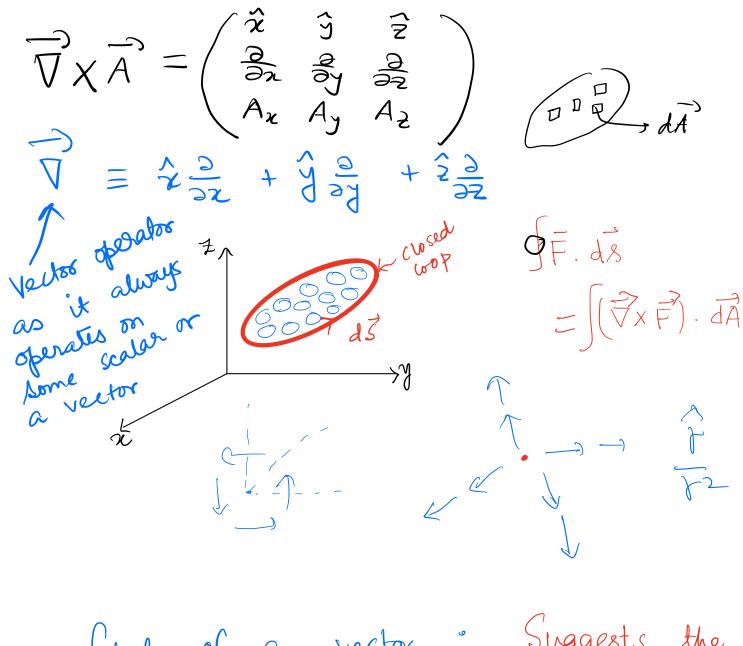
Closed

Closed

Coop Work done is

Zero for a => NECESSARY conservative conservative force along a closed path. -> Path independence implies the result depends on the START and END point only. Work done also depends on the difference blu the Kinetic energies (KE), and, if there is no dissipation while a particle traverses from point (1) to point (2), then the KE's do not Change. =) \$\vec{F}, d\vec{F} =0 W12 = \(\bar{F}' \cdot d\frac{7}{8} STOKES' THEOREM: Path Stokes' Theorem $\int \left(\overrightarrow{\nabla} \times \overrightarrow{F} \right) \cdot d\overrightarrow{A} = 0$ Area bounded by original loop. Conservative forces => $\overrightarrow{\nabla} \times \overrightarrow{F} = \overrightarrow{O}$ have no curl!

=> Existence of 3 constraint equations.



Curl of a vector: Suggests the rotation of a vector field at a point in Space.

MORE ON THIS PART BEFORE
THE START OF ELECTRICITY AND
MAGNETISM.

Constraint Equations:

$$\begin{pmatrix}
\frac{\partial}{\partial x} \\
\frac{\partial}{\partial y}
\end{pmatrix}$$

$$X$$

$$\begin{cases}
F_{x} \\
F_{y}
\end{cases}$$

$$\Rightarrow \begin{cases}
\frac{\partial}{\partial x} & \frac{\partial}{\partial y} & \frac{\partial}{\partial z} \\
F_{x} & F_{y}
\end{cases}$$

$$\begin{cases}
F_{x} \\
F_{z}
\end{cases}$$

$$\frac{\partial F_{y}}{\partial z} - \frac{\partial F_{z}}{\partial y} = 0 \quad j \quad \frac{\partial F_{z}}{\partial z} - \frac{\partial F_{z}}{\partial z} = 0 \quad j$$

$$\frac{\partial F_{x}}{\partial y} - \frac{\partial F_{y}}{\partial z} = 0$$

Encapsulate the above mathematics into one function => Scalar potential!

Define a potential V(r) such that,

We need!
$$\overrightarrow{\nabla} \times \overrightarrow{F} = \overrightarrow{O}$$

(Use the above)

$$= 7 \quad \overrightarrow{\nabla} \times \overrightarrow{\nabla} V = \overrightarrow{O}$$

Mathematical Identity.

Examples:

$$F_{G_1} = G_1 \quad \frac{m_1 m_2}{r^2}$$

$$V_G \propto \frac{1}{V}$$

$$F_{el} = \frac{9_1 a_2}{4\pi\epsilon_0 r^2}$$

Non-conservative forces:

$$F_{\text{mag}} = q \left(\overrightarrow{0} \times \overrightarrow{B} \right)$$

(Kinetic) Frisons = - k 70

-. No sealar potential energies for these forces.

depends on the relative velocity b/w two surfaces in contact with each other.

Potential Energy

Scalar potential $V(\vec{r})$ only defined up to a constant.

Vet's define,
$$\nabla (\vec{r}) = V(\vec{r}) + V_0$$

Then, $\vec{F} = -\nabla (\nabla (\vec{r}) - V_0)$
 $= -\nabla \nabla (\vec{r})$
 $= -\nabla \nabla (\vec{r})$

Same force (and honce, same physics)

widependent of added constant Vo.

** FOR INTERESTED STUDENTS

This is like "GAUGE SYMMETRY".

(Unknown

we will encounter this in upto a gauge

** ELECTRODYNAMICS part field)

Absolute SCALAR POTENTIAL is ambiguous.

LET'S SEE HOW THIS WORKS IN

PRACTICE

$$W_{12} = \int \vec{F} \cdot d\vec{r} = -\int \vec{\nabla} V \cdot d\vec{r} = V_1 - V_2$$

$$d\vec{r} \cdot dx \hat{i} + dy \hat{j} + dz \hat{k} ; \quad \frac{\partial V}{\partial x} \hat{i} + \frac{\partial V}{\partial y} \hat{j} + \frac{\partial V}{\partial z} \hat{k}$$
Check this out in the I dimensional version:
$$-\int_{x_1}^{x_2} dx V(x) dx = -\left(V(x_2) - V(x_1)\right)$$

 $W_{12} = T_2 - T_1 = V_1 - V_2 \Leftarrow Conservative$ force

$$\longrightarrow V_1 + T_1 = V_2 + T_2$$

Total energy at the Start of journey at pt. (1) is equal to the total energy at the end of journey at pt. (2).

=> CONSERVATION OF ENERGY