falPot. GalaxyPotential generating code

This documentation is just here as a basic guide to explain how to use falPot. See Section 2.3 of Dehnen & Binney (1998) for a fuller explanation of how it works.

fal Pot generates the potential due to a sum of axisymmetric disks and spheroids. It's basic functions are probably best understood by looking at the executable provided – testfal Pot.

N.B. In all cases here, the length unit is kpc, the time unit is Myr, and the mass unit is M_{\odot} . For reference, $G = 4.49866 \times 10^{-12} \mathrm{kpc^3 Myr^{-2} M_{\odot}^{-1}}$. A velocity of 1 kpc Myr⁻¹ = 977.77km s⁻¹ (See Units.h for further).

1 Input

The file that is given as the input file for the potential must be of the following form:

```
# No. Disks (0 \le \text{No.disks} \le 3)
                                          # Param. for each of the three disks.
8.90e7
         1.8
                0.04
3.50e7
         1.2
                -0.15
                        0
                                         # N.B essential to give 5 param.
1.30e7
                                         \# No. lines of param. = No. disks
         1.1
                0.25
                                          # No. Spheroids (0 \le \text{No.sph.} \le 2)
^{2}
4.0e7
         0.8
                -0.5
                              10
2.0e7
                                         # essential to give 6 parameters
                -1.5
```

1.1 Disk

The disk parameters, as given in that file are (in order) Σ_0 , R_d , z_d , R_0 , ϵ .

 Σ_0 is the disk's central surface density (in $M_{\odot} kpc^{-2}$ and the absence of a cutoff), R_d is the disk scale radius, z_d its scale height (though note there is a difference between negative and positive values), and R_0 is an inner cutoff radius (all in kpc). The term ϵ can provide a small perturbation to the surface density. In numerical terms (cylindrical polars), these give a surface density

$$\Sigma(R) = \Sigma_0 \exp\left(-\frac{R_0}{R} - \frac{R}{R_d} + \epsilon \cos\left(\frac{R}{R_d}\right)\right),\tag{1}$$

i.e. a standard exponential disk with optional hole in the middle and/or ϵ term modulation.

The vertical structure of the density is either an exponential in |z| (if $z_d > 0$), or isothermal (if $z_d < 0$), i.e.

$$\rho_d(R, z) = \begin{cases}
\frac{\Sigma(R)}{2z_d} \exp\left(\frac{-|z|}{z_d}\right) & \text{for } z_d > 0 \\
\frac{\Sigma(R)}{4(-z_d)} \operatorname{sech}^2\left(\frac{z}{2z_d}\right) & \text{for } z_d < 0.
\end{cases}$$
(2)

N.B. There is a factor of two in the denominator of the sech² profile. That is so that they tend to the same thing as $z \to \infty$.

1.2 **Spheroids**

The spheroid parameters are (in order) ρ_0 , q, γ , β , r_0 , r_{cut} . ρ_0 is a scale density (in $M_{\odot} \text{kpc}^{-3}$; q is the axis ratio (q < 1 is oblate, q > 1 is prolate); γ is the inner density slope, β the outer density slope; r_0 is a scale radius and r_{cut} is a cutoff radius (both in kpc). This corresponds to a density

$$\rho_s = \frac{\rho_0}{(r'/r_0)^{\gamma} (1 + r'/r_0)^{\beta - \gamma}} \exp\left[-\left(r'/r_{cut}\right)^2\right],\tag{3}$$

where, in cylindrical coordinates

$$r' = \sqrt{R^2 + (z/q)^2} \tag{4}$$

References

Dehnen W., Binney J., 1998, MNRAS, 294, 429