

Digital Image Processing

CS390S

Feng Jiang
2018 Spring



METROPOLITAN STATE UNIVERSITYSM
OF DENVER

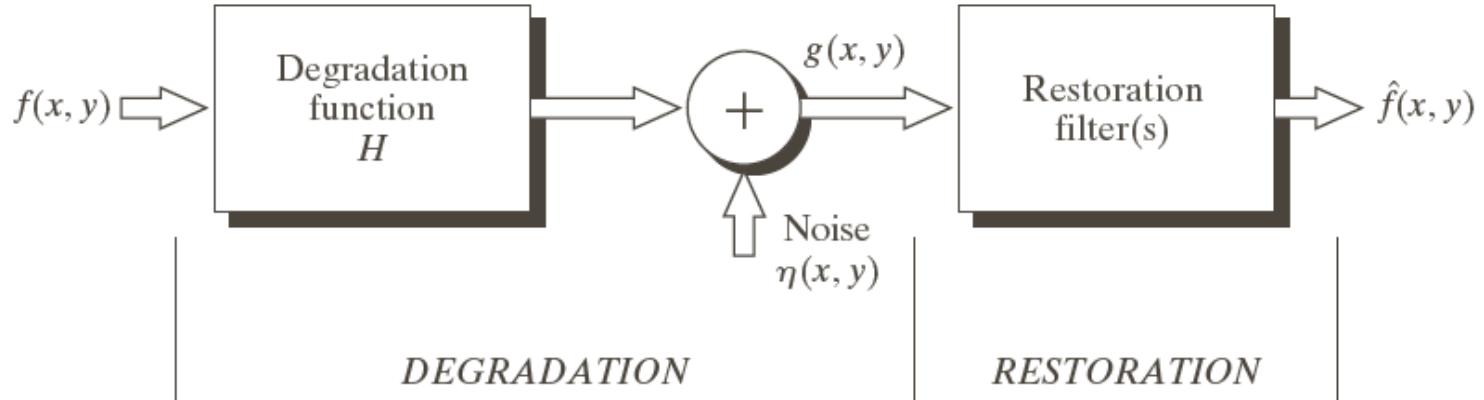
Image Restoration

- Image restoration: recover an image that has been degraded by using a prior knowledge of the degradation phenomenon.
- Model the degradation and applying the **inverse** process in order to recover the original image.

A Model of Image Degradation/Restoration Process

FIGURE 5.1

A model of the image degradation/restoration process.

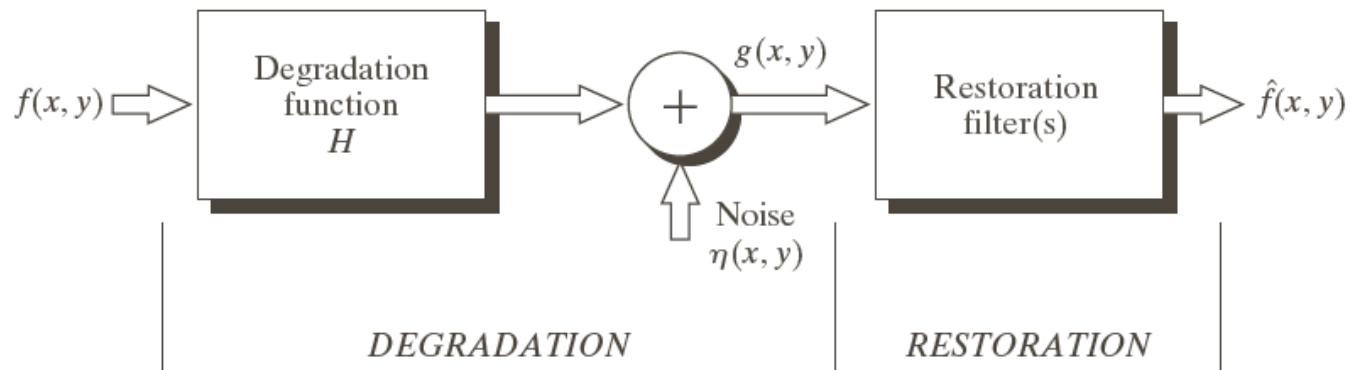


► Degradation

- Degradation function H
- Additive noise $\eta(x, y)$

A Model of Image Degradation/Restoration Process

FIGURE 5.1
A model of the image degradation/restoration process.



If H is a linear, position-invariant process, then the degraded image is given in the spatial domain by

$$g(x, y) = h(x, y) \star f(x, y) + \eta(x, y)$$

A Model of Image Degradation/Restoration Process

The model of the degraded image is given in the frequency domain by

$$G(u, v) = H(u, v)F(u, v) + N(u, v)$$

Noise Sources

- The principal sources of noise in digital images arise during **image acquisition and/or transmission**
 - ✓ Image acquisition
 - e.g., light levels, sensor temperature, etc.
 - ✓ Transmission/environment
 - e.g., lightning or other atmospheric disturbance in wireless network

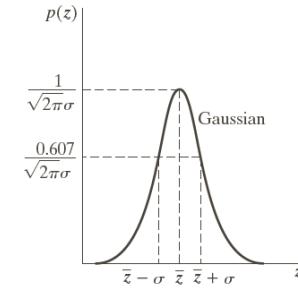
Noise Models (1)

- **Gaussian noise**

- Mean, variance; sum of the independent noises
- Probability of values follows Gaussian model
- Electronic circuit noise, sensor noise due to poor illumination and/or high temperature

- **White noise**

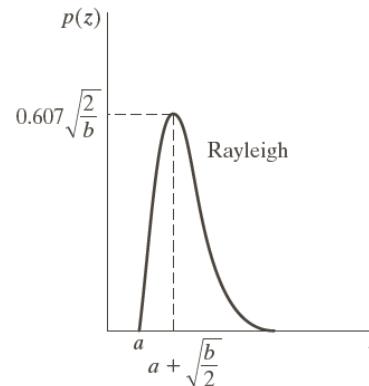
- https://en.wikipedia.org/wiki/White_noise
- In discrete time, white noise is a discrete signal whose samples are regarded as a sequence of serially uncorrelated random variables with zero mean and finite variance;
- Being uncorrelated in time does not restrict the values a signal can take. Any distribution of values is possible (although it must have zero DC component). Even a binary signal which can only take on the values 1 or -1 will be white if the sequence is statistically uncorrelated. Noise having a continuous distribution, such as a normal distribution, can of course be white.
- It is often incorrectly assumed that Gaussian noise (i.e., noise with a Gaussian amplitude distribution – see normal distribution) necessarily refers to white noise, yet neither property implies the other.
- Gaussian white noise is a good approximation of many real-world situations and generates mathematically tractable models.



Noise Models (2)

- **Rayleigh noise**

Range imaging



- The intention of HDRI is to accurately represent the wide range of intensity levels found in real scenes ranging from direct **sunlight to shadows**
- HDRI, also called HDR (High Dynamic Range) is a feature commonly found in high-end graphics and imaging software

http://en.wikipedia.org/wiki/High_dynamic_range_imaging

http://www.webopedia.com/TERM/H/High_Dynamic_Range_Imaging.html

Range Imaging: Examples (1)

http://en.wikipedia.org/wiki/High_dynamic_range_imaging



Tower Bridge in
Sacramento, CA

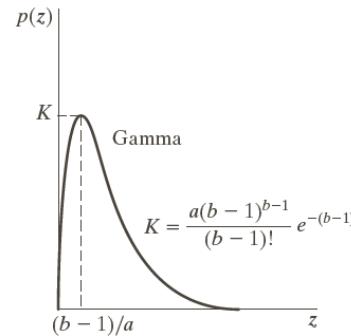
Range Imaging: Examples (2)

http://en.wikipedia.org/wiki/High_dynamic_range_imaging

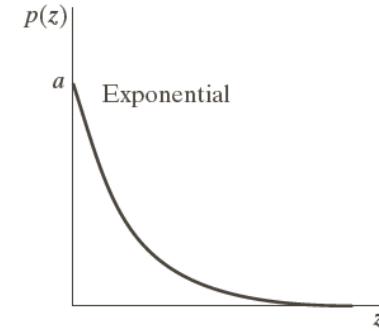
Sydney Harbour Bridge
HDRI produces greater detail and fewer shadows



Noise Models (3)

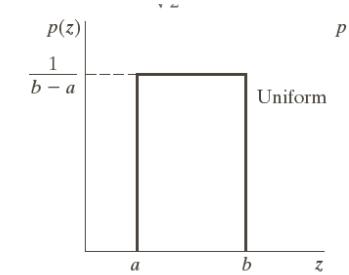


➤ **Erlang (gamma) noise:** Laser imaging

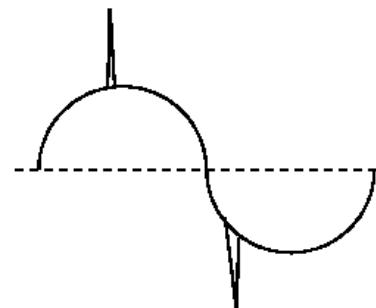


➤ **Exponential noise:** Laser imaging

➤ **Uniform noise:** Least descriptive; **quantization noise**



➤ **Impulse noise:** quick transients,
such as faulty switching



Impulse Noise

Gaussian Noise (1)

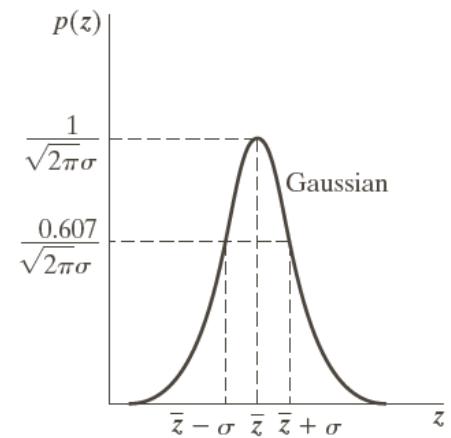
The PDF of Gaussian random variable, z , is given by

$$p(z) = \frac{1}{\sqrt{2\pi}\sigma} e^{-(z-\bar{z})^2/2\sigma^2}$$

where, z represents intensity

\bar{z} is the mean (average) value of z

σ is the standard deviation



Gaussian Noise (2)

The PDF of Gaussian random variable, z, is given by

$$p(z) = \frac{1}{\sqrt{2\pi}\sigma} e^{-(z-\bar{z})^2/2\sigma^2}$$

- 70% of its values will be in the range

$$[(\mu - \sigma), (\mu + \sigma)]$$

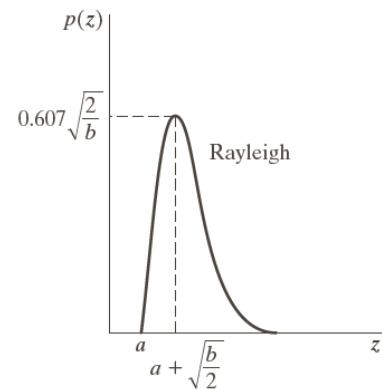
- 95% of its values will be in the range

$$[(\mu - 2\sigma), (\mu + 2\sigma)]$$

Rayleigh Noise

The PDF of Rayleigh noise is given by

$$p(z) = \begin{cases} \frac{2}{b}(z-a)e^{-(z-a)^2/b} & \text{for } z \geq a \\ 0 & \text{for } z < a \end{cases}$$



The mean and variance of this density are given by

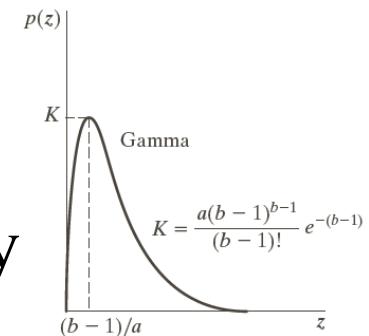
$$\bar{z} = a + \sqrt{\pi b / 4}$$

$$\sigma^2 = \frac{b(4 - \pi)}{4}$$

Erlang (Gamma) Noise

The PDF of Erlang noise is given by

$$p(z) = \begin{cases} \frac{a^b z^{b-1}}{(b-1)!} e^{-az} & \text{for } z \geq 0 \\ 0 & \text{for } z < 0 \end{cases}$$



The mean and variance of this density are given by

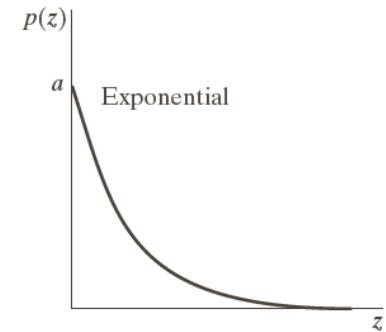
$$\bar{z} = b/a$$

$$\sigma^2 = b/a^2$$

Exponential Noise

The PDF of exponential noise is given by

$$p(z) = \begin{cases} ae^{-az} & \text{for } z \geq 0 \\ 0 & \text{for } z < a \end{cases}$$



The mean and variance of this density are given by

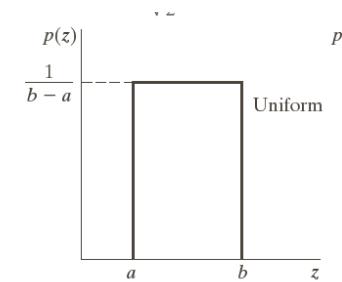
$$\bar{z} = 1/a$$

$$\sigma^2 = 1/a^2$$

Uniform Noise

The PDF of uniform noise is given by

$$p(z) = \begin{cases} \frac{1}{b-a} & \text{for } a \leq z \leq b \\ 0 & \text{otherwise} \end{cases}$$



The mean and variance of this density are given by

$$\bar{z} = (a + b) / 2$$

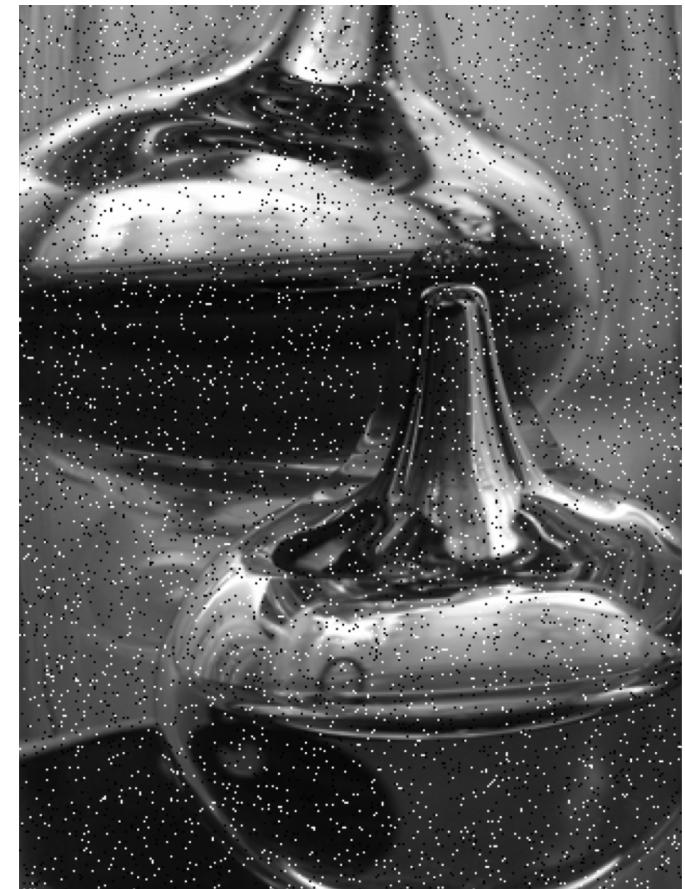
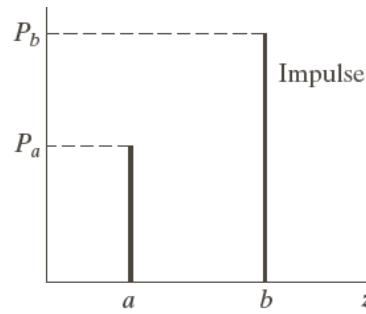
$$\sigma^2 = (b - a)^2 / 12$$

Impulse (Salt-and-Pepper) Noise

The PDF of (bipolar) impulse noise is given by

$$p(z) = \begin{cases} P_a & \text{for } z = a \\ P_b & \text{for } z = b \\ 0 & \text{otherwise} \end{cases}$$

if $b > a$, gray-level b will appear as a light dot, while level a will appear like a dark dot.



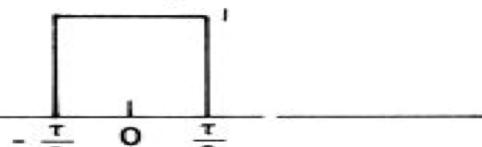
Day 12

- Review “frequency transform and filtering”
- Noise reduction / image restoration

Time Function

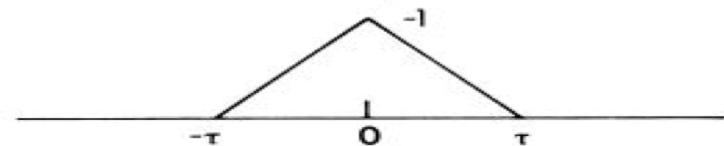
Boxcar

$$G(t) = \begin{cases} 1, & |t| < \tau/2 \\ 0, & |t| > \tau/2 \end{cases}$$



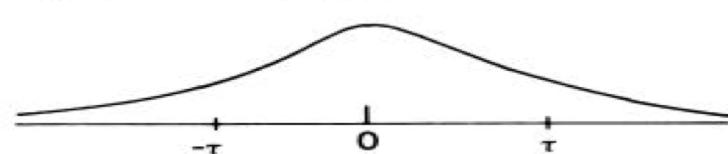
Triangle

$$G(t) = \begin{cases} 1-|t|/\tau, & |t| < \tau \\ 0, & |t| > \tau \end{cases}$$



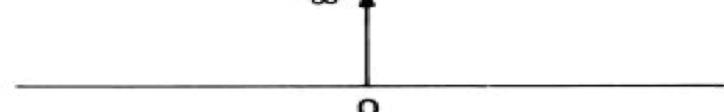
Gaussian

$$G(t) = e^{-1/2 t^2}$$



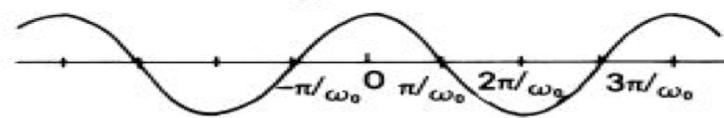
Impulse

$$G(t) = \delta(t) = 0, \quad t \neq 0$$



Sinusoid

$$G(t) = \cos \omega_0 t$$



Frequency Function

Sinc

$$S(f) = \tau \operatorname{sinc}(f\tau)$$

$$\tau = (1/\pi f) \sin(\pi f \tau)$$

$$\operatorname{sinc}^2$$

$$S(f) = \tau \operatorname{sinc}^2(f\tau)$$

$$= (1/\pi^2 f^2 \tau) \sin^2(\pi f \tau)$$

Gaussian

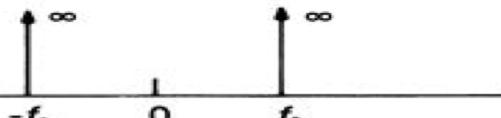
$$S(f) = \tau(2\pi)^{1/2} e^{-(\pi f \tau)^2}$$

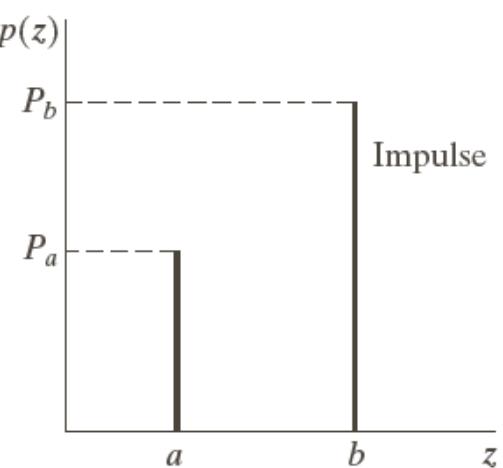
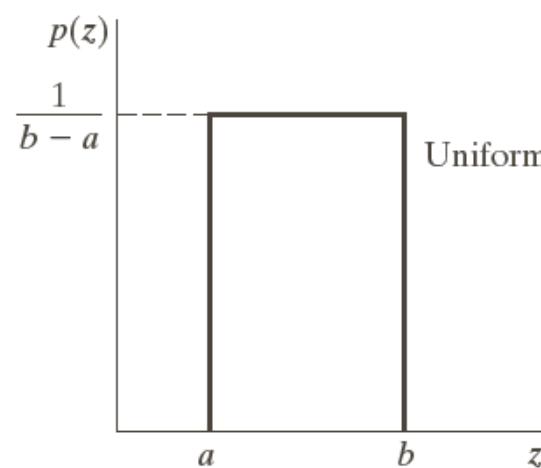
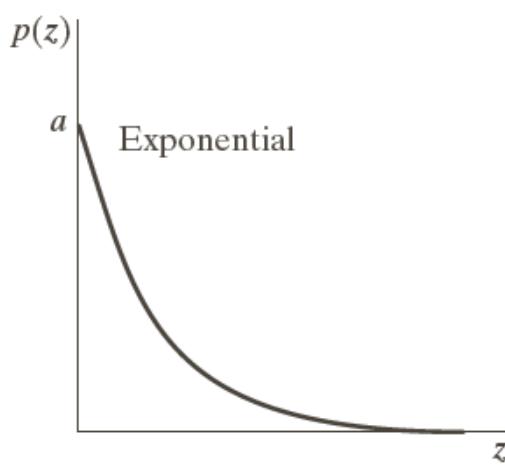
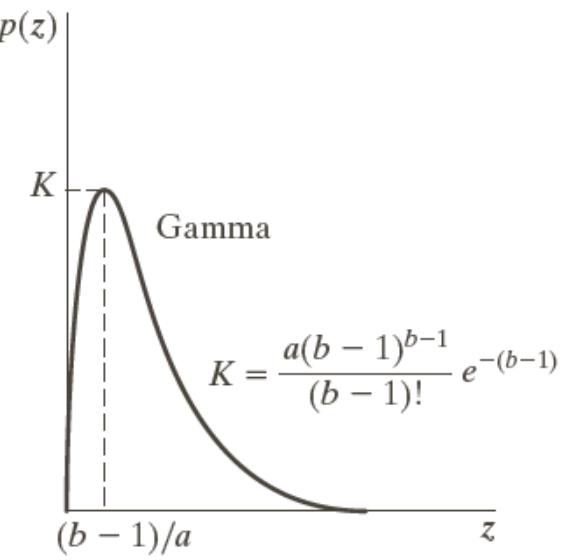
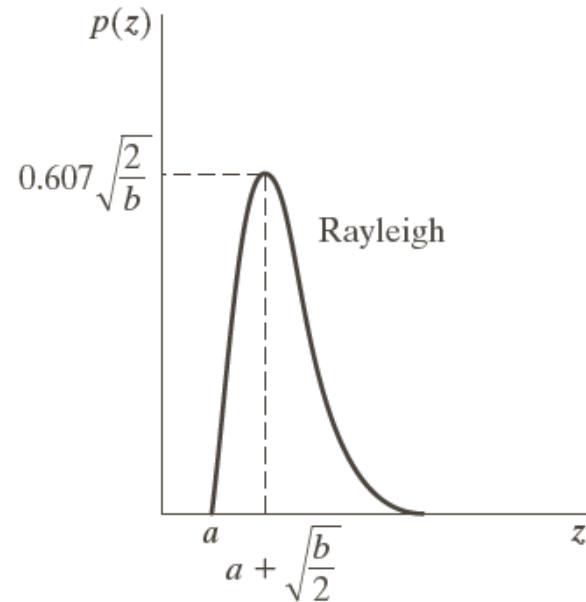
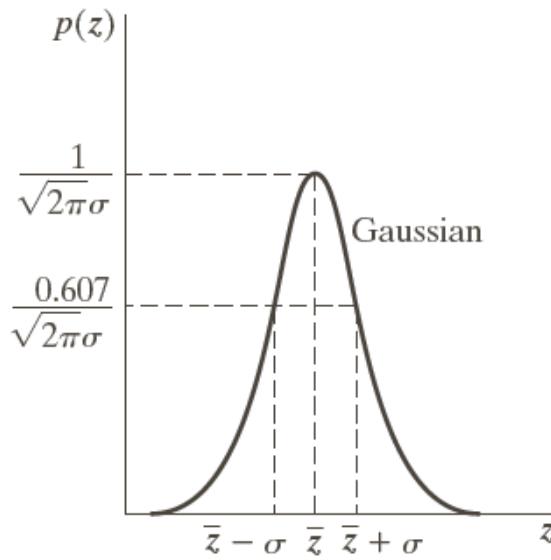
DC Shift

$$S(f) = 1$$

Single Freq.

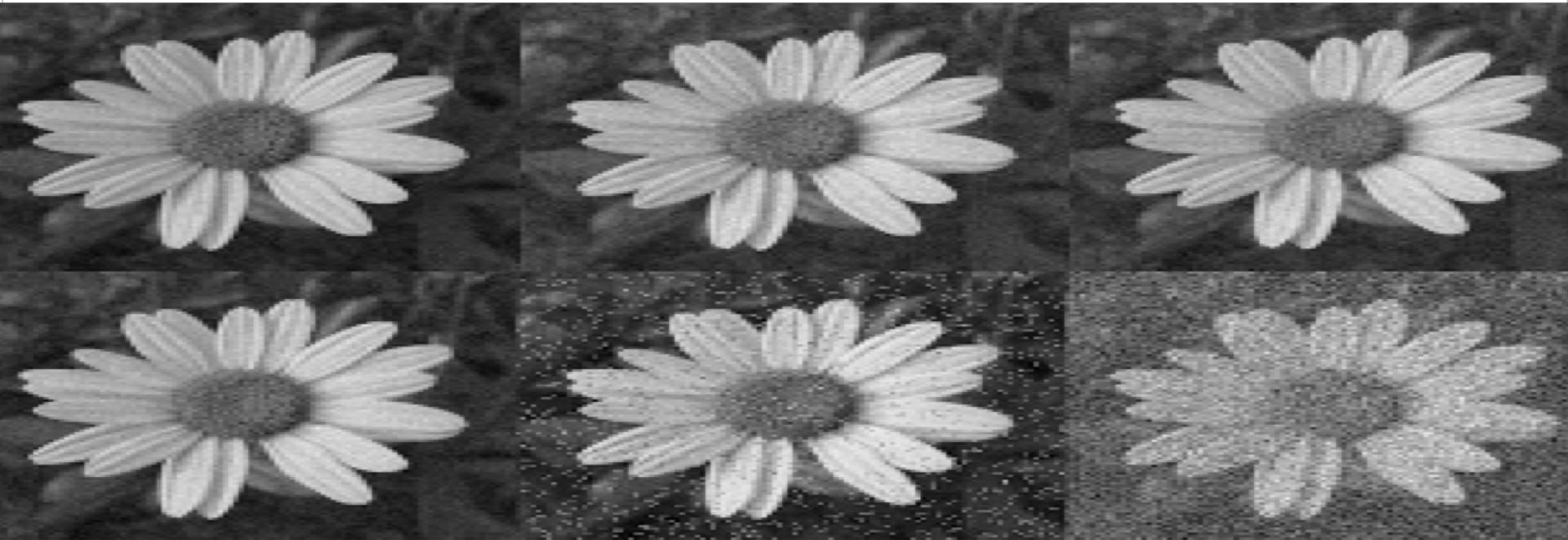
$$S(f) = 1/2 (\delta(f+f_0) + \delta(f-f_0))$$





a	b	c
d	e	f

FIGURE 5.2 Some important probability density functions.



clockwise direction: exponential noise, gamma noise, gaussian noise, rayleigh noise, salt and pepper noise, uniform noise.



clockwise direction: exponential noise, gamma noise, gaussian noise, rayleigh noise, salt and pepper noise, uniform noise.

Examples of Noise: Original Image

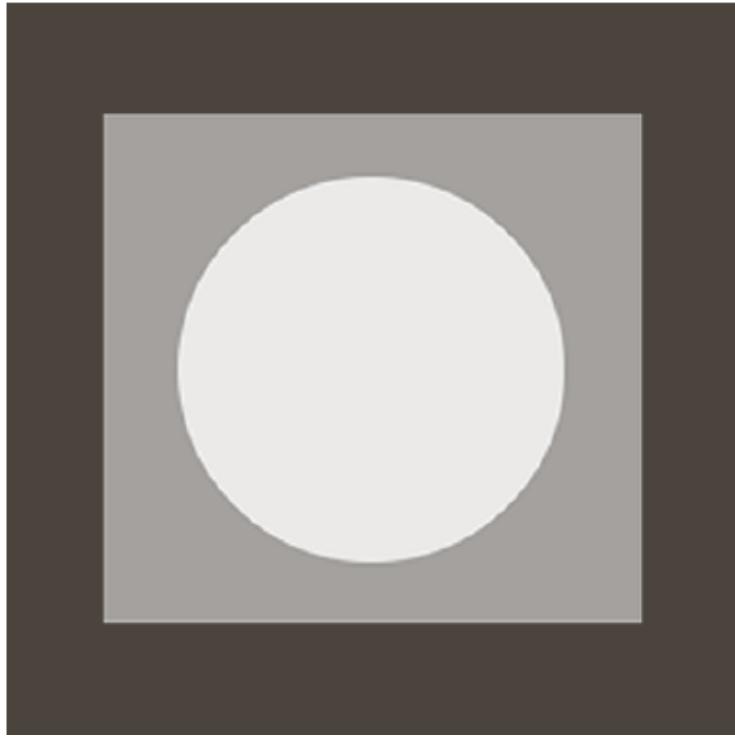
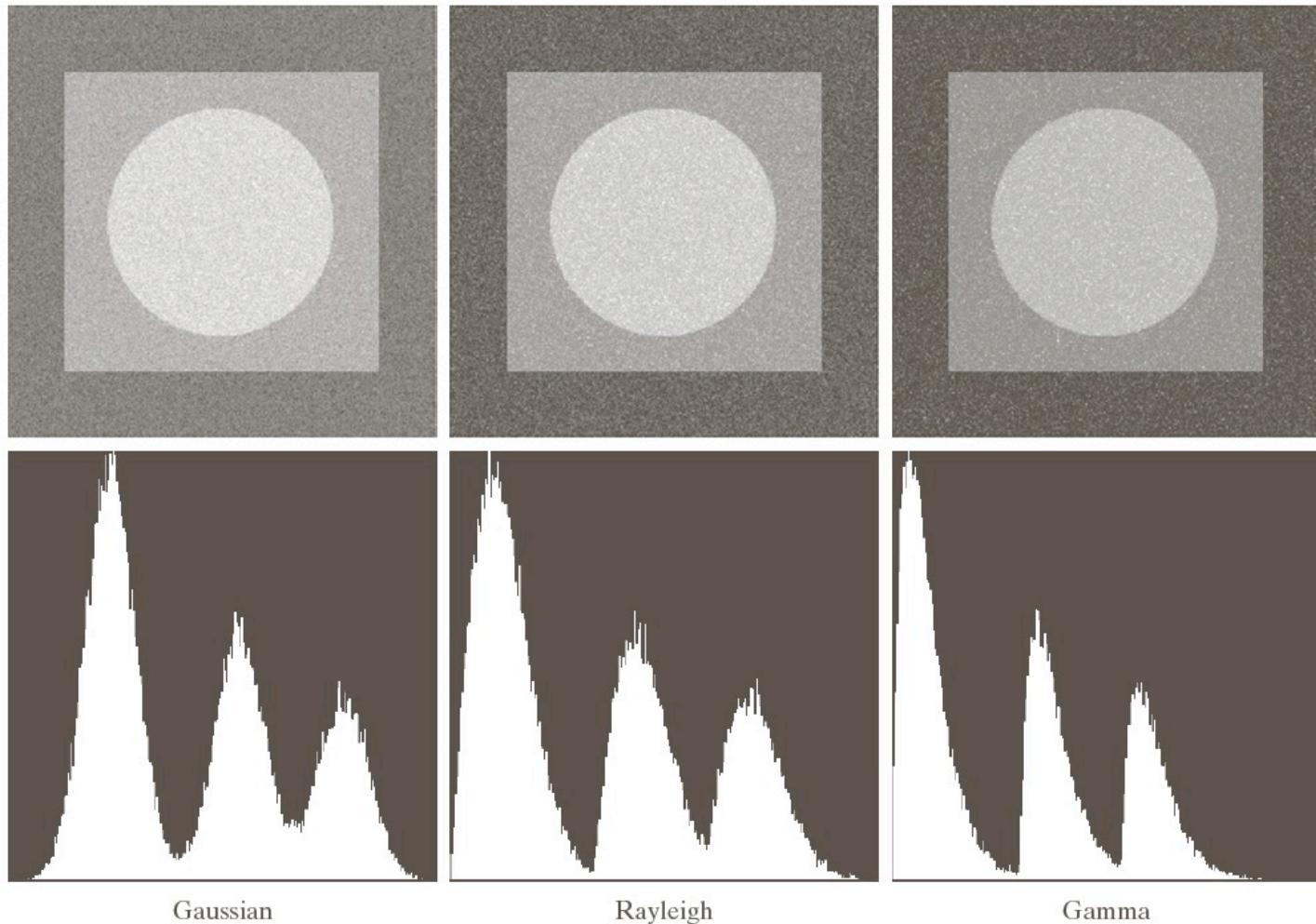


FIGURE 5.3 Test pattern used to illustrate the characteristics of the noise PDFs shown in Fig. 5.2.

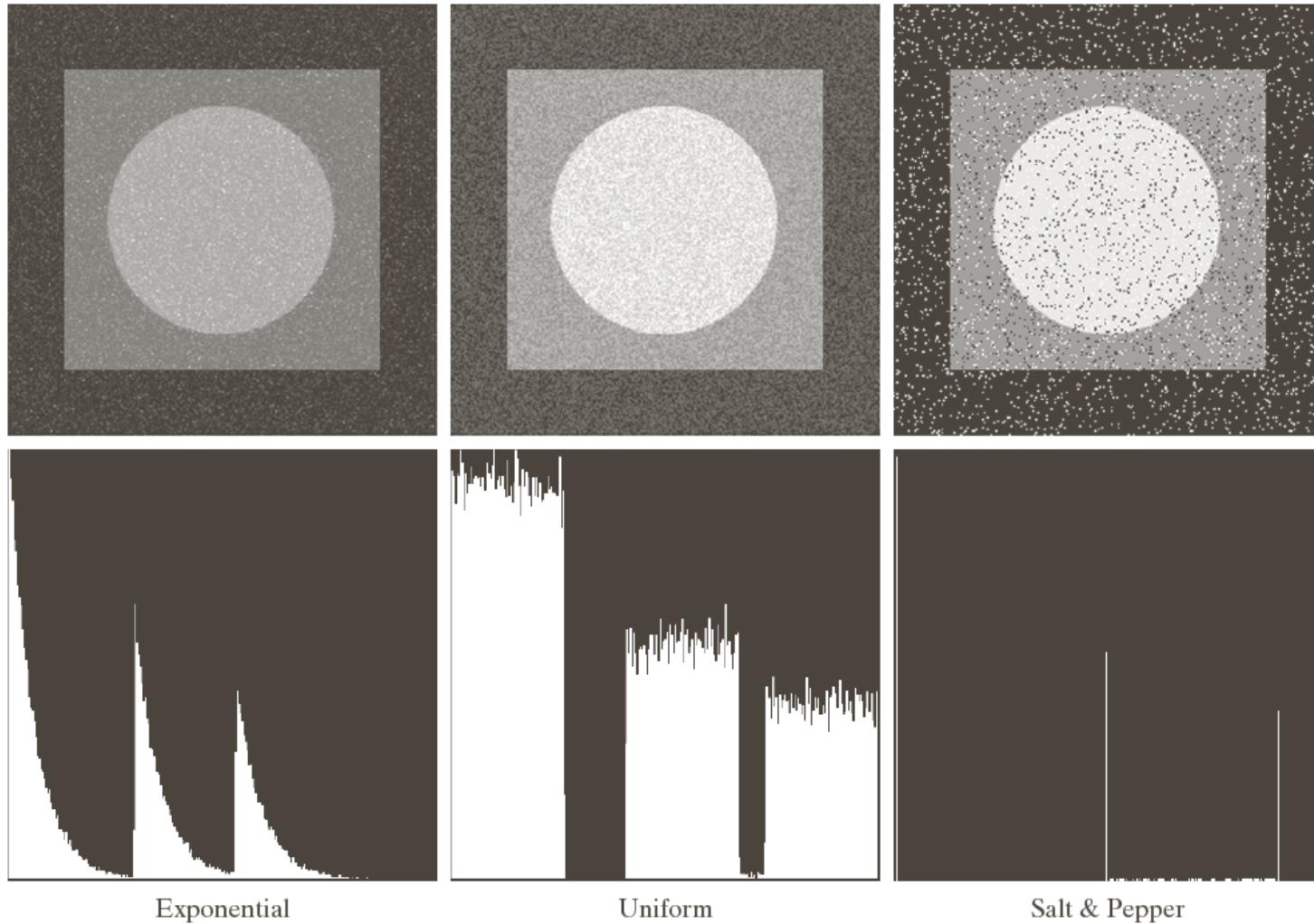
Examples of Noise: Noisy Images(1)



MI
OF I

FIGURE 5.4 Images and histograms resulting from adding Gaussian, Rayleigh, and gamma noise to the image in Fig. 5.3.

Examples of Noise: Noisy Images(2)



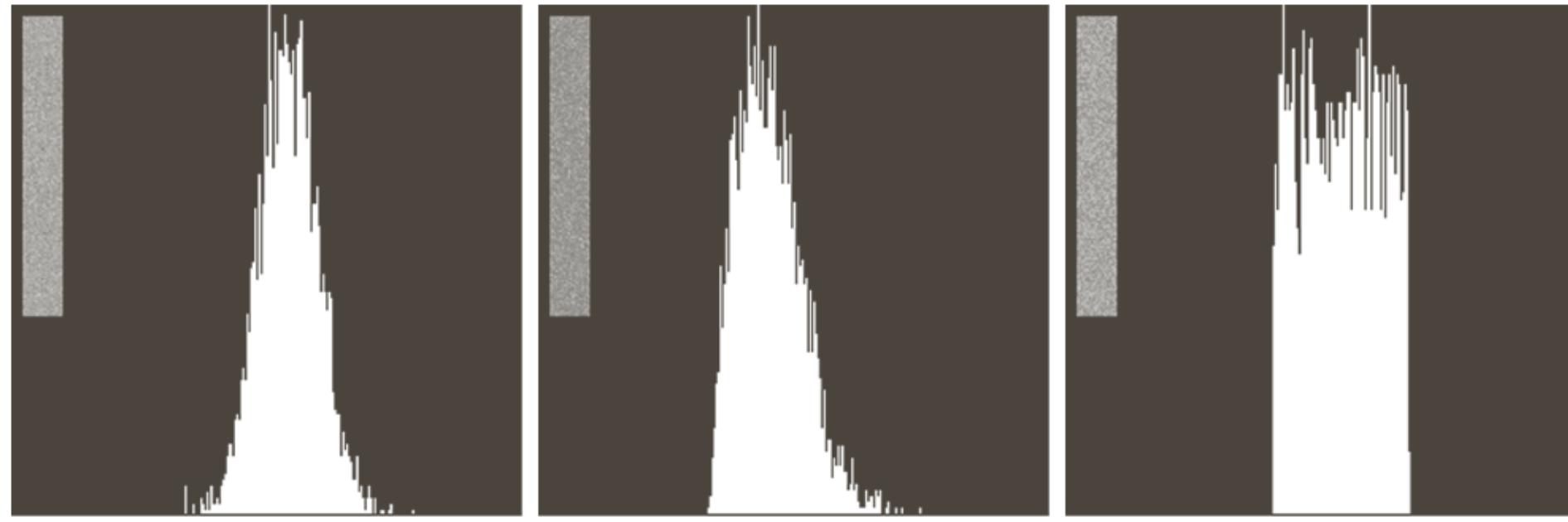
g	h	i
j	k	l



FIGURE 5.4 (Continued) Images and histograms resulting from adding exponential, uniform, and salt and pepper noise to the image in Fig. 5.3.

Estimation of Noise Parameters

The shape of the histogram identifies the closest PDF match



a b c

FIGURE 5.6 Histograms computed using small strips (shown as inserts) from (a) the Gaussian, (b) the Rayleigh, and (c) the uniform noisy images in Fig. 5.4.

Restoration in the Presence of Noise Only

– Spatial Filtering

Noise model without degradation

$$g(x, y) = f(x, y) + \eta(x, y)$$

and

$$G(u, v) = F(u, v) + N(u, v)$$

e.g. Periodic noise fixed frequency

Spatial Filtering: Mean Filters (1)

Let S_{xy} represent the set of coordinates in a rectangle subimage window of size $m \times n$, centered at (x, y) .

Arithmetical mean filter

$$\hat{f}(x, y) = \frac{1}{mn} \sum_{(s, t) \in S_{xy}} g(s, t)$$

Spatial Filtering: Mean Filters (2)

Geometric mean filter

$$\hat{f}(x, y) = \left[\prod_{(s,t) \in S_{xy}} g(s, t) \right]^{\frac{1}{mn}}$$

Generally, a geometric mean filter achieves smoothing comparable to the arithmetic mean filter, but it tends to lose less image detail in the process

Spatial Filtering: Mean Filters (2)

Geomean()

Syntax

```
m = geomean(x)  
geomean(X,dim)
```

Description

`m = geomean(x)` calculates the geometric mean of a sample. For vectors, `geomean(x)` is the geometric mean of the elements in `x`. For matrices, `geomean(X)` is a row vector containing the geometric means of each column. For N-dimensional arrays, `geomean` operates along the first nonsingleton dimension of `X`.

`geomean(X,dim)` takes the geometric mean along the dimension `dim` of `X`.

The geometric mean is

$$m = \left[\prod_{i=1}^n x_i \right]^{\frac{1}{n}}$$

The arithmetic mean is greater than or equal to the geometric mean.

```
a=rand(5,1)  
a = 0.8147 0.9058 0.1270 0.9134 0.6324  
geomean(a)  
ans = 0.5581  
mean(a)  
ans = 0.6786
```

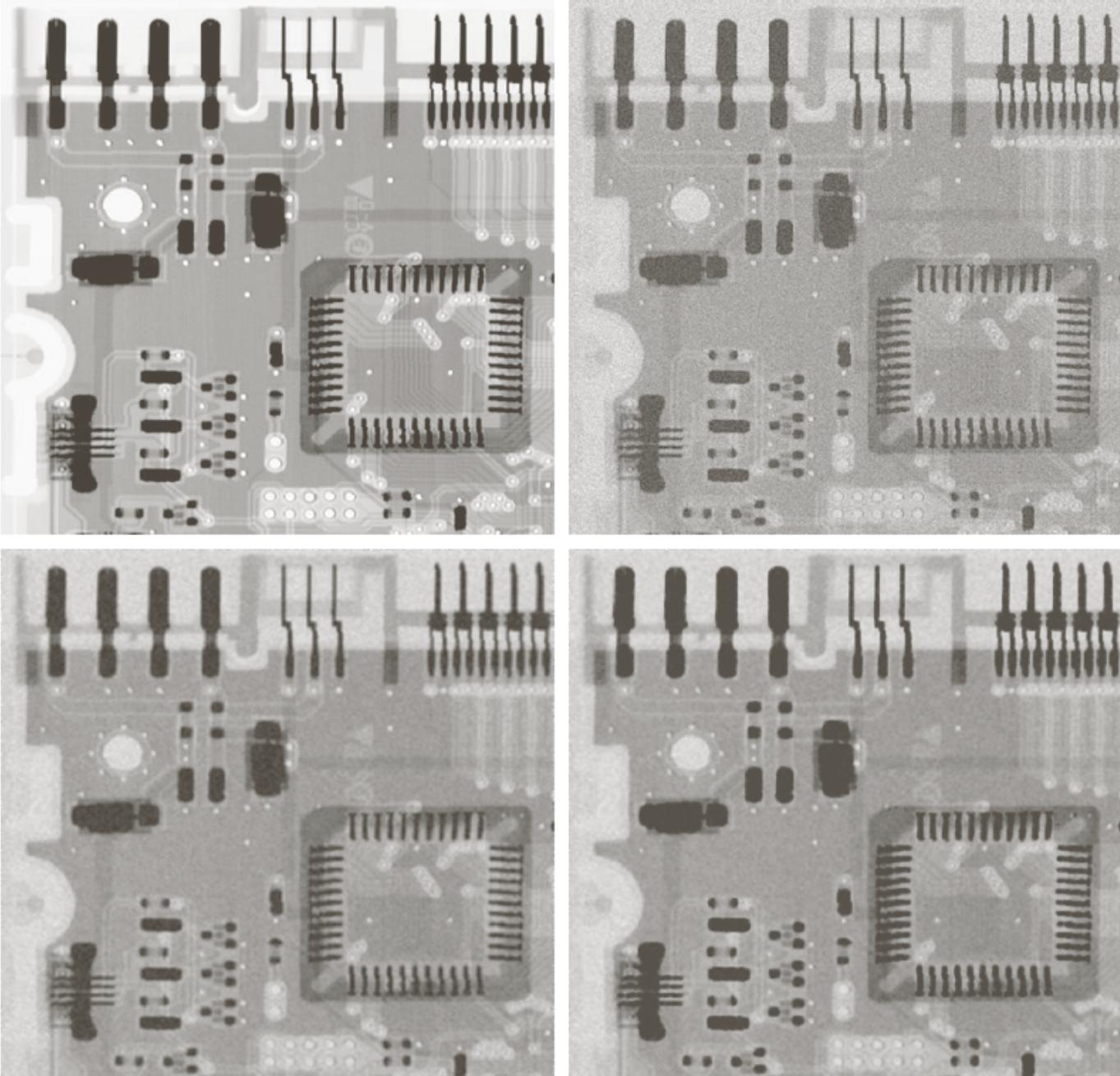
Spatial Filtering: Example (1)

a
b
c
d

FIGURE 5.7

(a) X-ray image.
(b) Image corrupted by additive Gaussian noise. (c) Result of filtering with an arithmetic mean filter of size 3×3 . (d) Result of filtering with a geometric mean filter of the same size.

(Original image courtesy of Mr. Joseph E. Pascente, Lixi, Inc.)



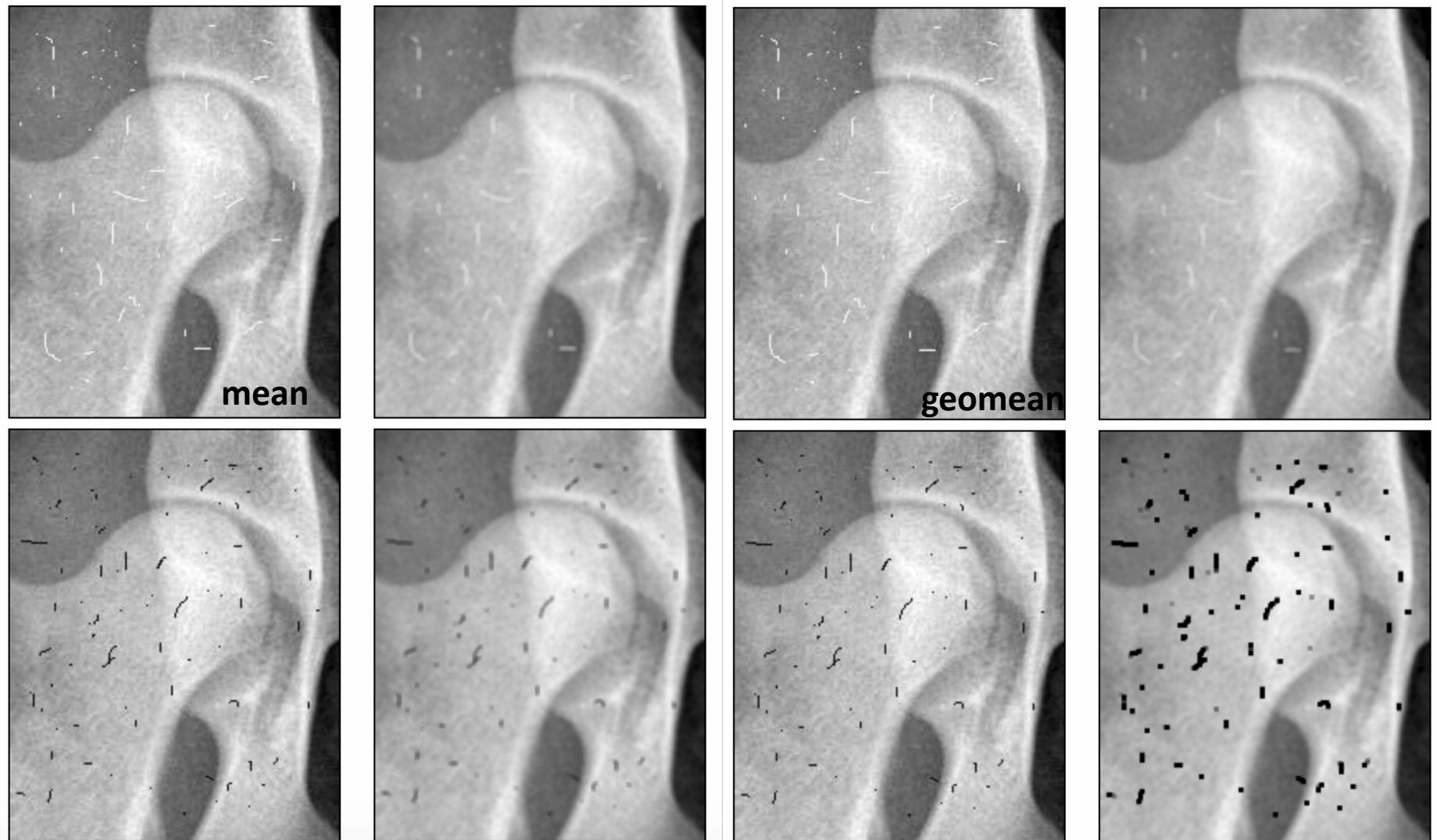
The **geometric mean filter** is better at removing Gaussian type noise and preserving edge features than the **arithmetic mean filter**.



Spatial Filtering: Example (2)

- <https://www.digimizer.com/manual/m-image-filtermean.php>
- <https://www.digimizer.com/manual/m-image-filtergeomean.php>

Spatial Filtering: Example (2)



Spatial Filtering: Mean Filters (3)

Harmonic mean filter

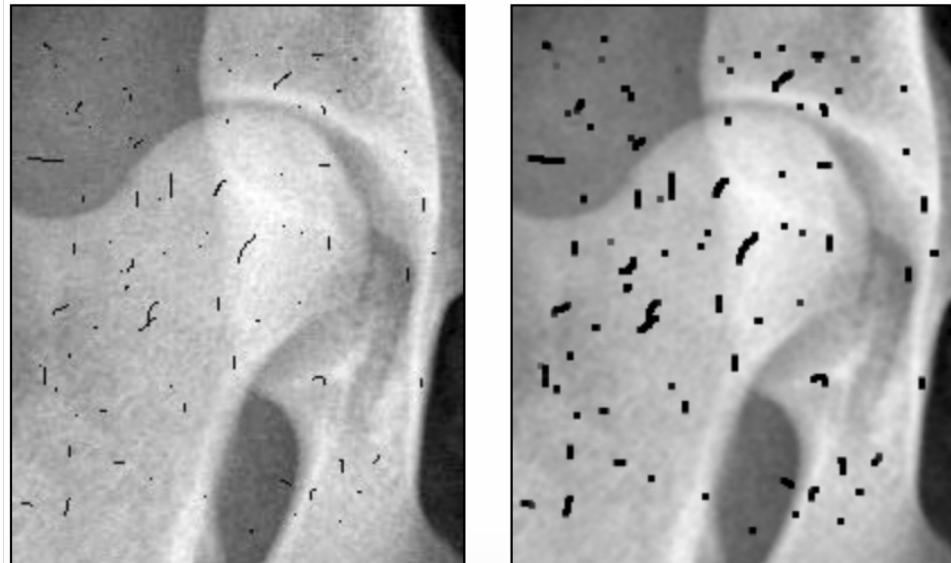
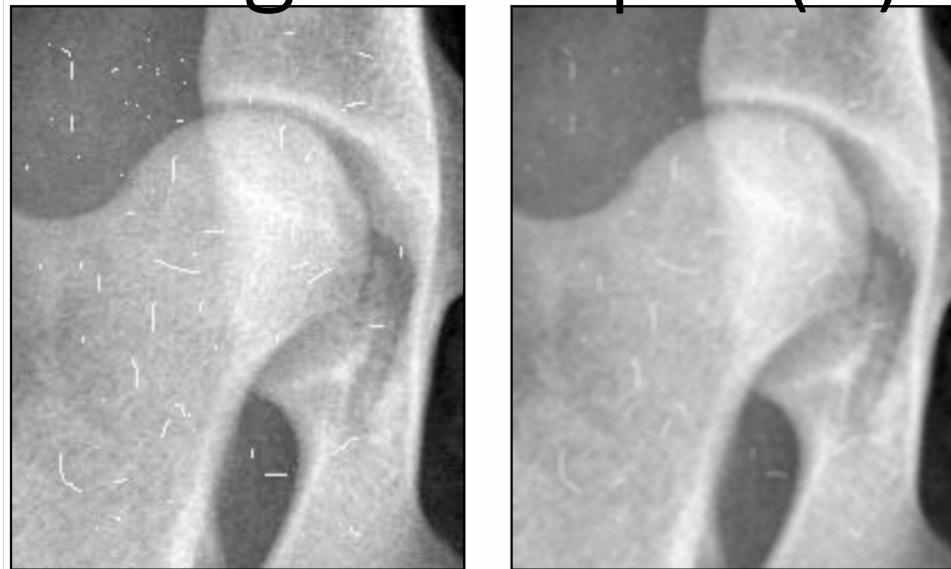
$$\hat{f}(x, y) = \frac{mn}{\sum_{(s,t) \in S_{xy}} \frac{1}{g(s, t)}}$$

$$\begin{aligned} H_n &= \frac{\sum_{i=1}^n x_i f_i}{\sum_{i=1}^n f_i} \\ &= \frac{1}{\frac{\frac{1}{x_1}m_1 + \frac{1}{x_2}m_2 + \cdots + \frac{1}{x_n}m_n}{m_1 + m_2 + \cdots + m_n}} \\ &= \frac{1}{\frac{\frac{1}{x_1} + \frac{1}{x_2} + \cdots + \frac{1}{x_n}}{n}} \\ &= \frac{n}{\sum_{i=1}^n \frac{1}{x_i}} \end{aligned}$$

“reciprocal of averaging of reciprocal”

It works **well for salt noise**, but fails for pepper noise.
It does well also with other types of noise like Gaussian noise.

Spatial Filtering: Example (2)



Spatial Filtering: Order-Statistic Filters (1)

Median filter

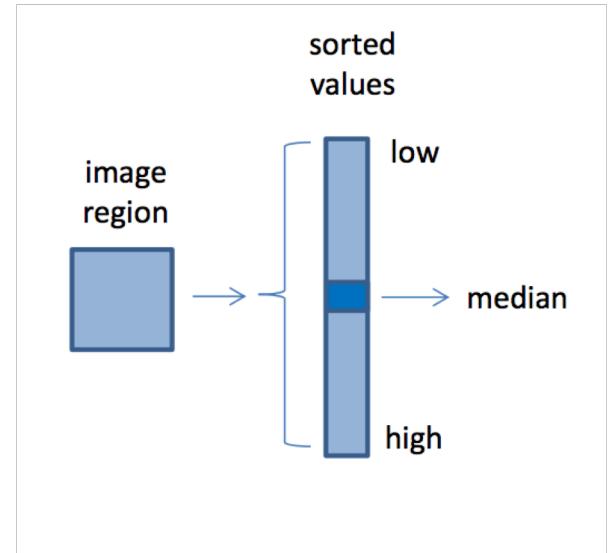
$$\hat{f}(x, y) = \underset{(s,t) \in S_{xy}}{\text{median}} \{g(s, t)\}$$

Max filter

$$\hat{f}(x, y) = \underset{(s,t) \in S_{xy}}{\max} \{g(s, t)\}$$

Min filter

$$\hat{f}(x, y) = \underset{(s,t) \in S_{xy}}{\min} \{g(s, t)\}$$



Spatial Filtering: Order-Statistic Filters (2)

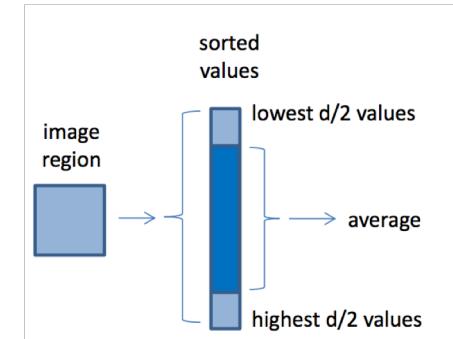
Midpoint filter

$$\hat{f}(x, y) = \frac{1}{2} \left[\max_{(s,t) \in S_{xy}} \{g(s, t)\} + \min_{(s,t) \in S_{xy}} \{g(s, t)\} \right]$$

Spatial Filtering: Order-Statistic Filters (3)

Alpha-trimmed mean filter

$$\hat{f}(x, y) = \frac{1}{mn - d} \sum_{(s,t) \in S_{xy}} \{g_r(s, t)\}$$

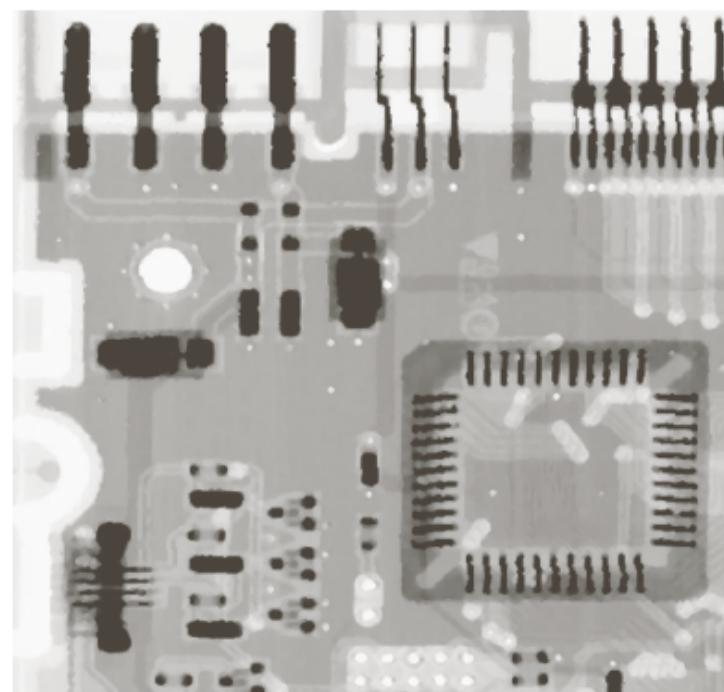
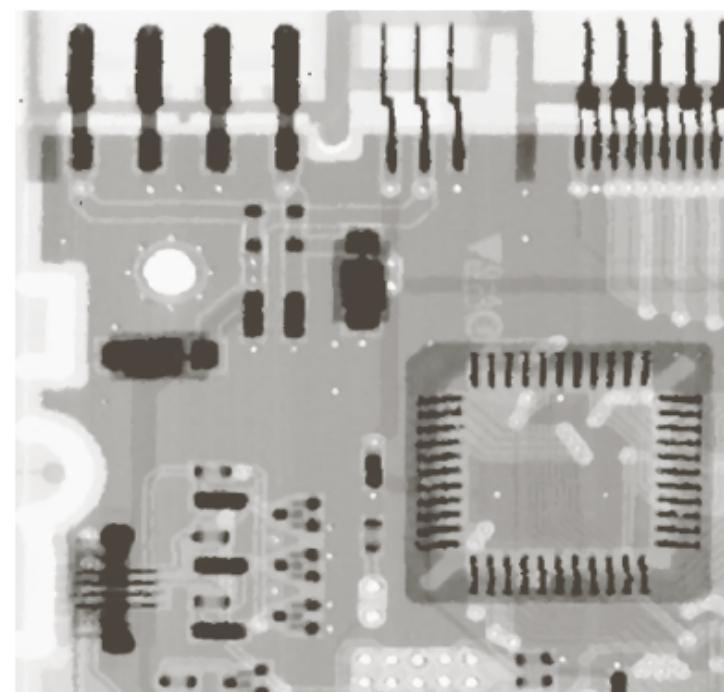
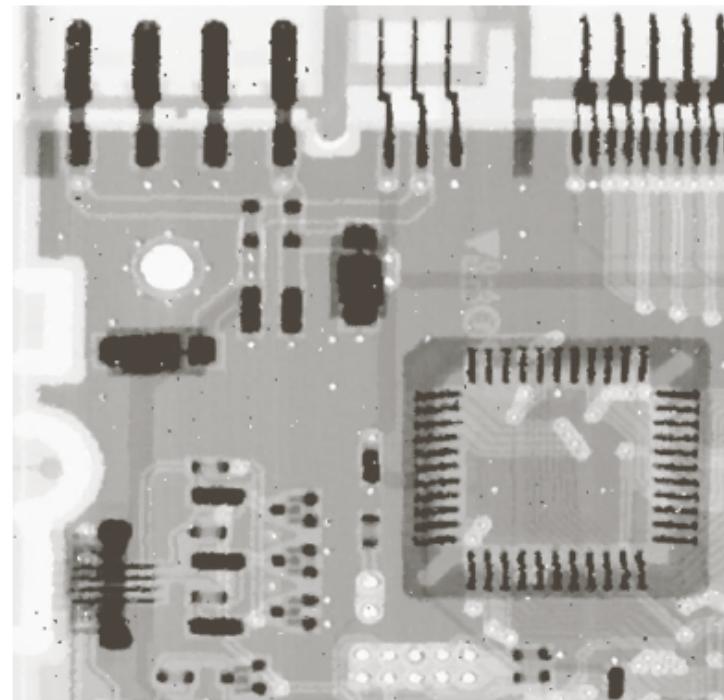
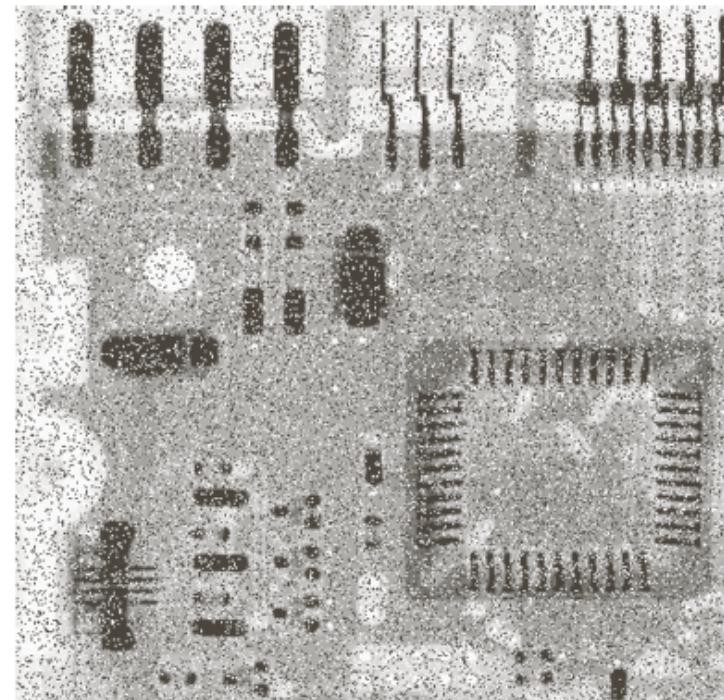


We delete the $d / 2$ lowest and the $d / 2$ highest intensity values of $g(s, t)$ in the neighborhood S_{xy} . Let $g_r(s, t)$ represent the remaining $mn - d$ pixels.

a
b
c
d

FIGURE 5.10

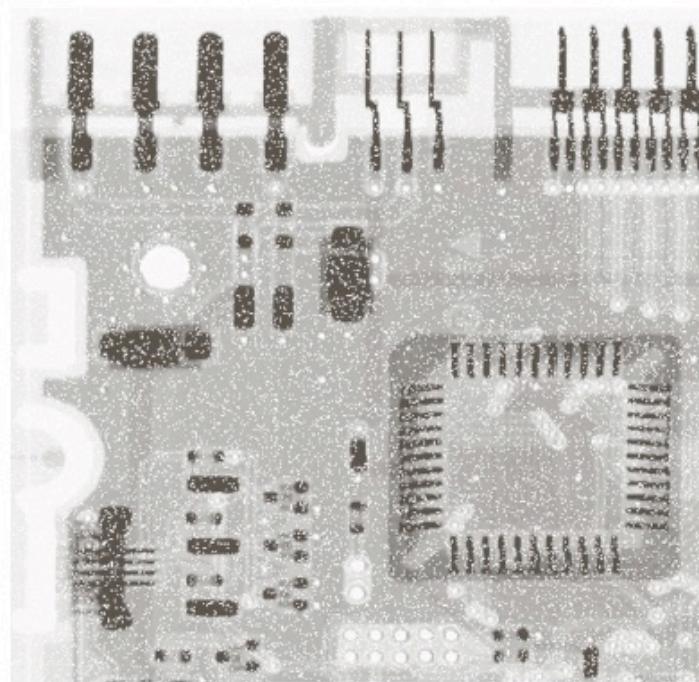
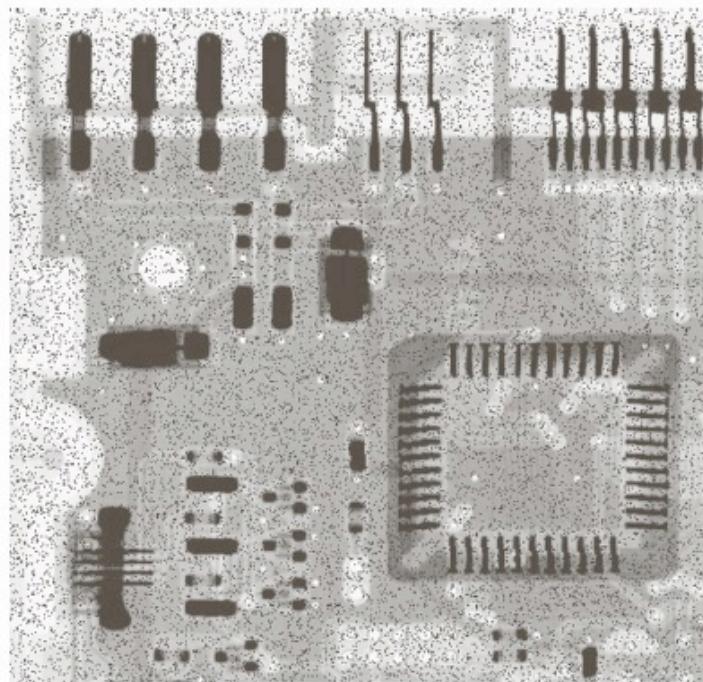
- (a) Image corrupted by salt-and-pepper noise with probabilities $P_a = P_b = 0.1$.
(b) Result of one pass with a median filter of size 3×3 .
(c) Result of processing (b) with this filter.
(d) Result of processing (c) with the same filter.
-



a	b
c	d

FIGURE 5.8

- (a) Image corrupted by pepper noise with a probability of 0.1. (b) Image corrupted by salt noise with the same probability. (c) Result of filtering (a) with a 3×3 contra-harmonic filter of order 1.5.

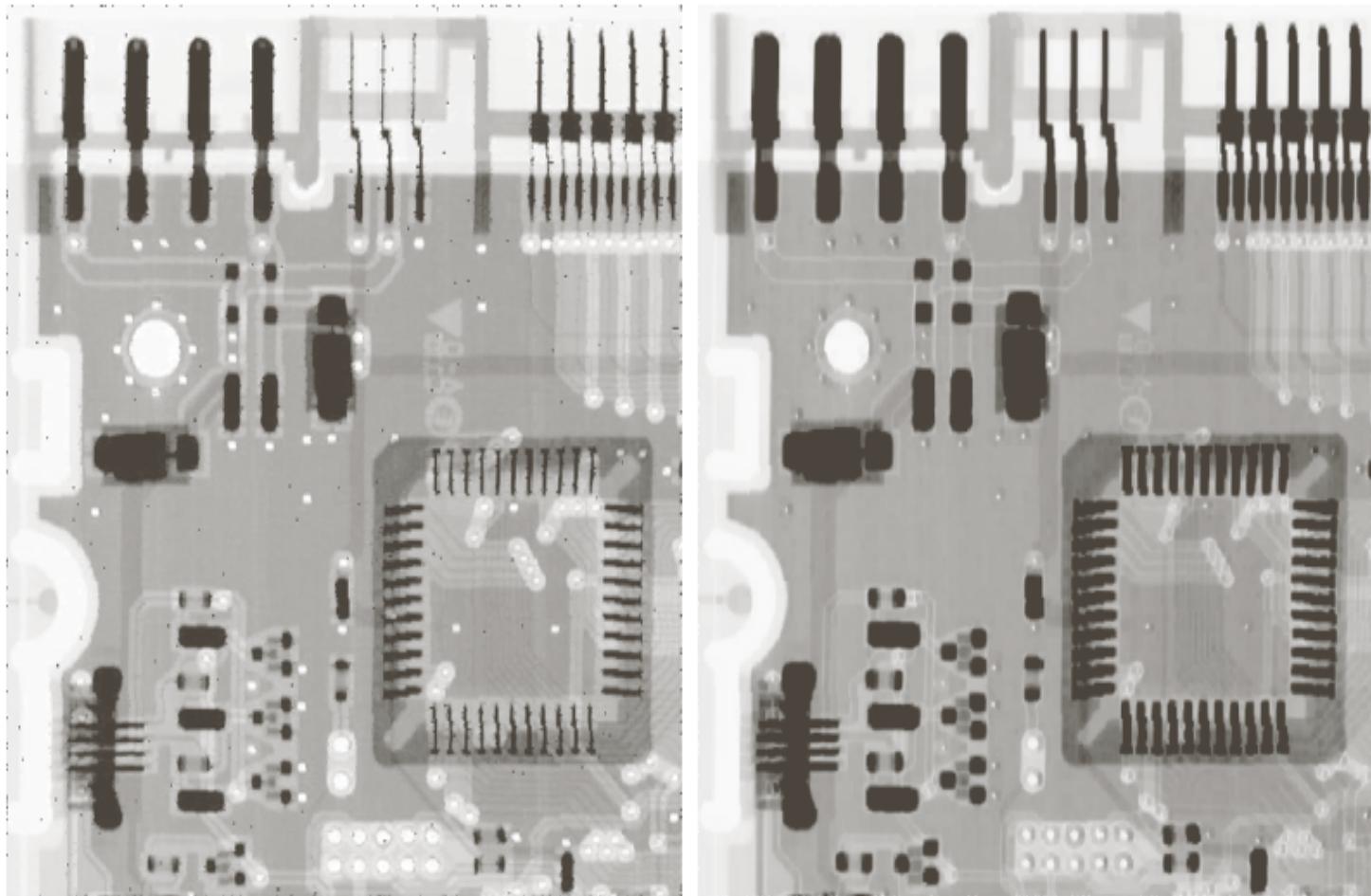


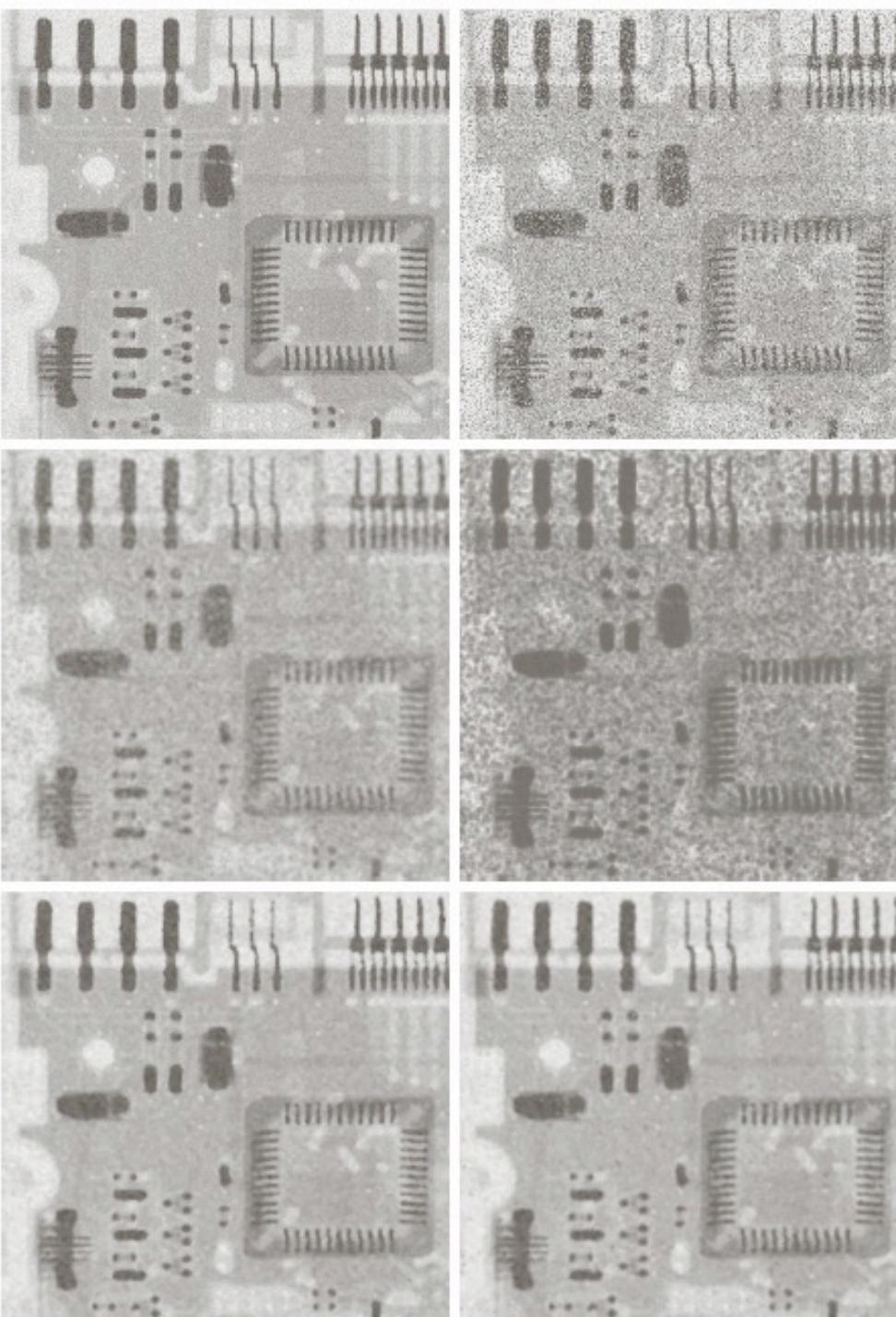
a b

FIGURE 5.11

(a) Result of filtering
Fig. 5.8(a) with a
max filter of size
 3×3 .

(b) Result of filtering Fig. 5.8(b)
with a min filter of the same size.





a
b
c
d
e
f

FIGURE 5.12

- (a) Image corrupted by additive uniform noise.
(b) Image additionally corrupted by additive salt-and-pepper noise.
Image (b) filtered with a 5×5 ;
(c) arithmetic mean filter;
(d) geometric mean filter;
(e) median filter;
and (f) alpha-trimmed mean filter with $d = 5$.

Spatial Filtering: Adaptive Filters (1)

Adaptive filters

The behavior **changes based on statistical characteristics of the image** inside the filter region defined by the $m \times n$ rectangular window.

The performance is superior to that of the filters discussed

Adaptive Filters:

Adaptive, Local Noise Reduction Filters (1)

S_{xy} : local region

The response of the filter at the center point (x,y) of S_{xy} is based on four quantities:

- (a) $g(x, y)$, the value of the noisy image at (x, y) ;
- (b) σ_η^2 , the variance of the noise corrupting $f(x, y)$ to form $g(x, y)$;
- (c) m_L , the local mean of the pixels in S_{xy} ;
- (d) σ_L^2 , the local variance of the pixels in S_{xy} .

Adaptive Filters:

Adaptive, Local Noise Reduction Filters (2)

$$\hat{f}(x, y) = g(x, y) - \frac{\sigma_{\eta}^2}{\sigma_L^2} [g(x, y) - m_L]$$

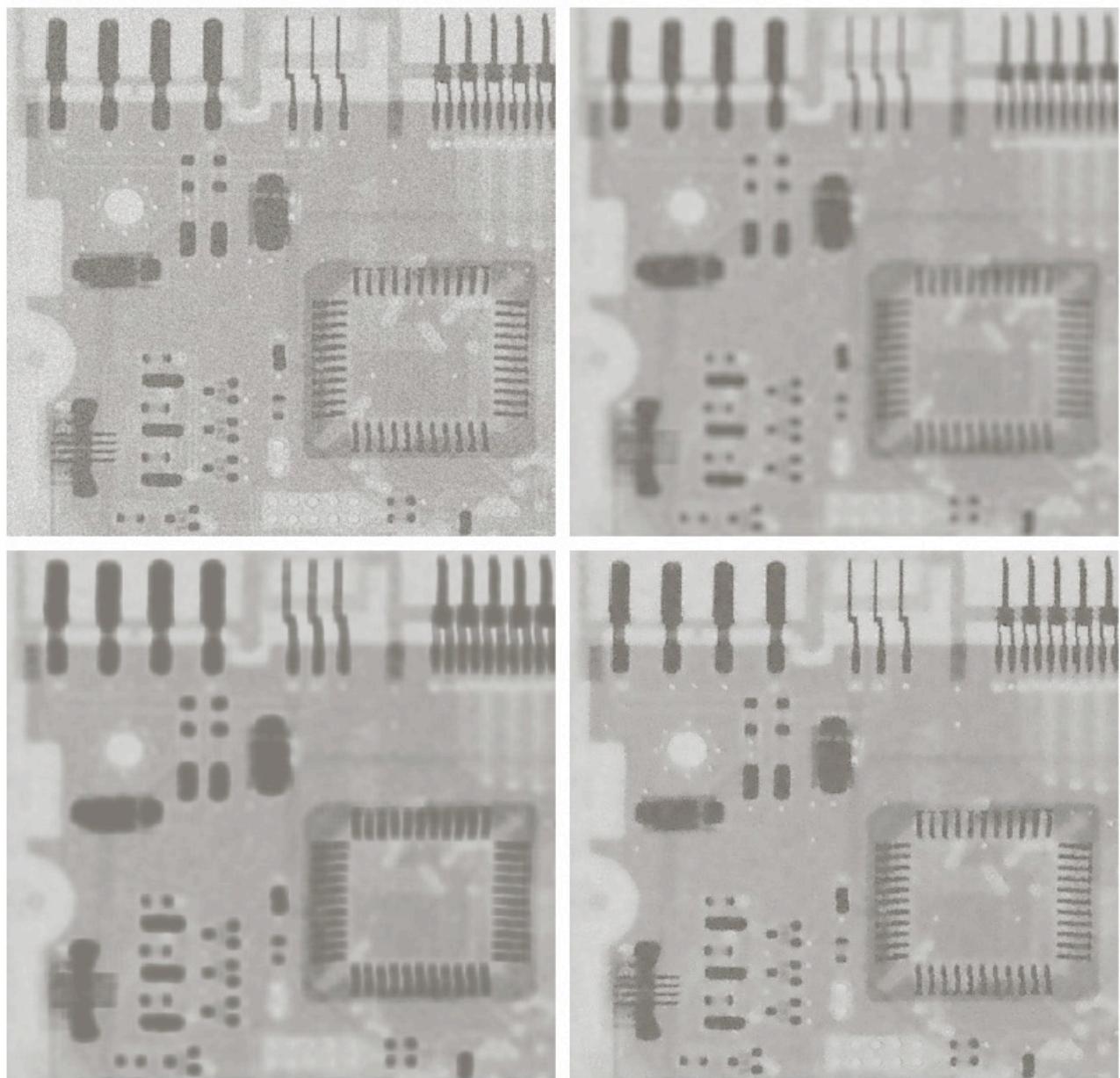
The behavior of the filter.

- (a) if σ_{η}^2 is zero, the filter should return simply the value of $g(x, y)$.
- (b) if the local variance is high relative to σ_{η}^2 , the filter should return a value close to $g(x, y)$;
- (c) if the two variances are equal, the filter returns the arithmetic mean value of the pixels in S_{xy} .

a
b
c
d

FIGURE 5.13

- (a) Image corrupted by additive Gaussian noise of zero mean and variance 1000.
(b) Result of arithmetic mean filtering.
(c) Result of geometric mean filtering.
(d) Result of adaptive noise reduction filtering. All filters were of size 7×7 .



Adaptive Filters:

Adaptive Median Filters (1)

The notation:

z_{\min} = minimum intensity value in S_{xy}

z_{\max} = maximum intensity value in S_{xy}

z_{med} = median intensity value in S_{xy}

z_{xy} = intensity value at coordinates (x, y)

S_{\max} = maximum allowed size of S_{xy}

Adaptive Filters:

Adaptive Median Filters (1)

The adaptive median-filtering works in two stages:

Stage A:

$$A1 = z_{\text{med}} - z_{\min}; \quad A2 = z_{\text{med}} - z_{\max}$$

if $A1 > 0$ and $A2 < 0$, go to stage B

Else increase the window size

if window size $\leq S_{\max}$, repeat stage A; Else output z_{med}

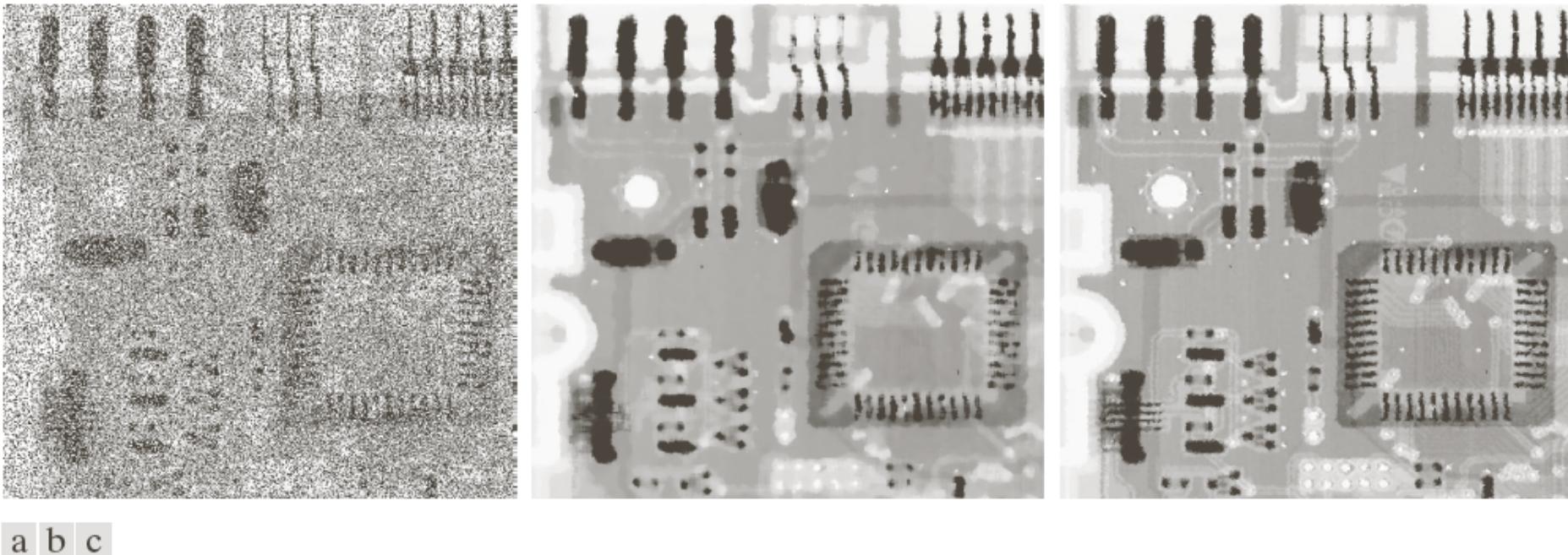
Stage B:

$$B1 = z_{xy} - z_{\min}; \quad B2 = z_{xy} - z_{\max}$$

if $B1 > 0$ and $B2 < 0$, output z_{xy} ; Else output z_{med}

Adaptive Filters:

Adaptive Median Filters (1)



a b c

FIGURE 5.14 (a) Image corrupted by salt-and-pepper noise with probabilities $P_a = P_b = 0.25$. (b) Result of filtering with a 7×7 median filter. (c) Result of adaptive median filtering with $S_{\max} = 7$.

Periodic Noise

- ▶ Periodic noise in an image arises typically from **electrical** or **electromechanical** interference during image acquisition.
- ▶ It is a type of spatially dependent noise
- ▶ Periodic noise can be reduced significantly via frequency domain filtering

Periodic Noise Reduction by Frequency Domain Filtering

The basic idea

Periodic noise appears as **concentrated bursts of energy** in the Fourier transform, at locations corresponding to the frequencies of the periodic interference

Approach

A **selective filter** is used to isolate the noise

Perspective Plots of Bandreject Filters



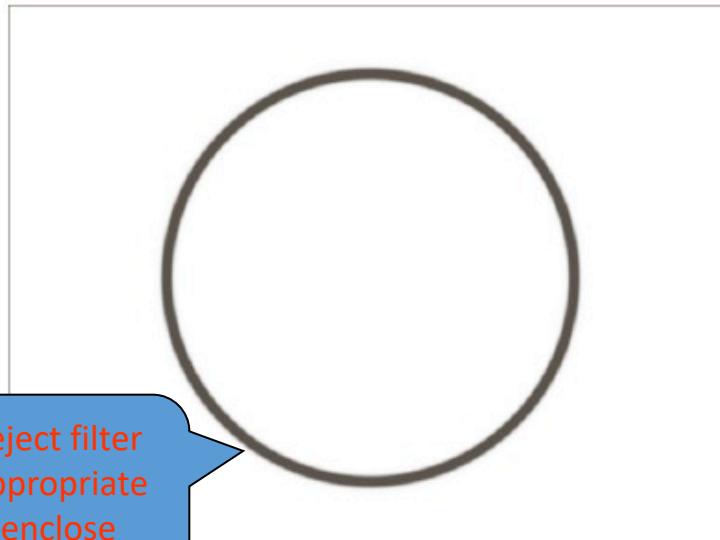
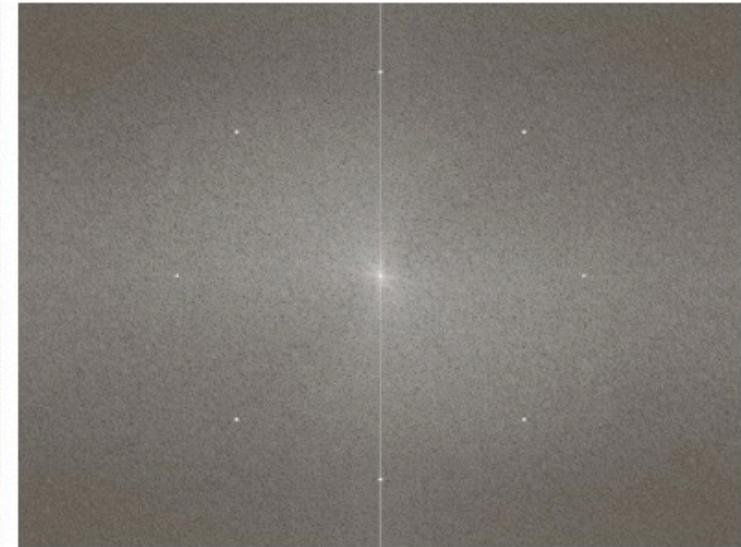
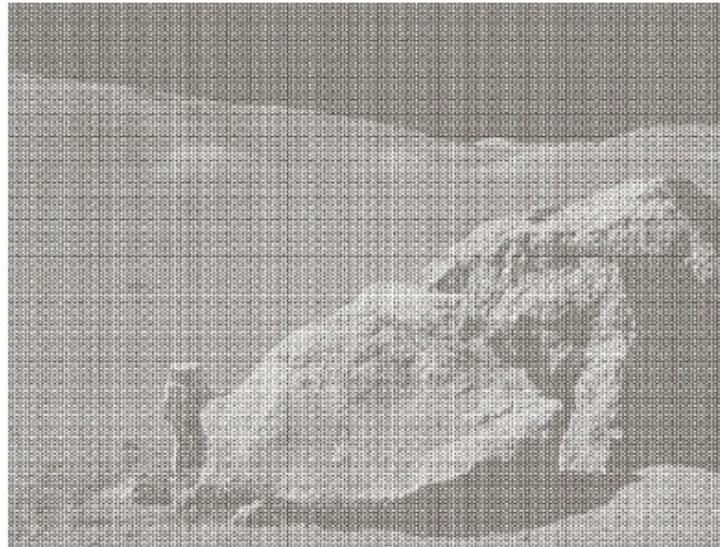
a b c

FIGURE 5.15 From left to right, perspective plots of ideal, Butterworth (of order 1), and Gaussian bandreject filters.

a
b
c
d

FIGURE 5.16

- (a) Image corrupted by sinusoidal noise.
(b) Spectrum of (a).
(c) Butterworth bandreject filter (white represents 1). (d) Result of filtering.
(Original image courtesy of NASA.)



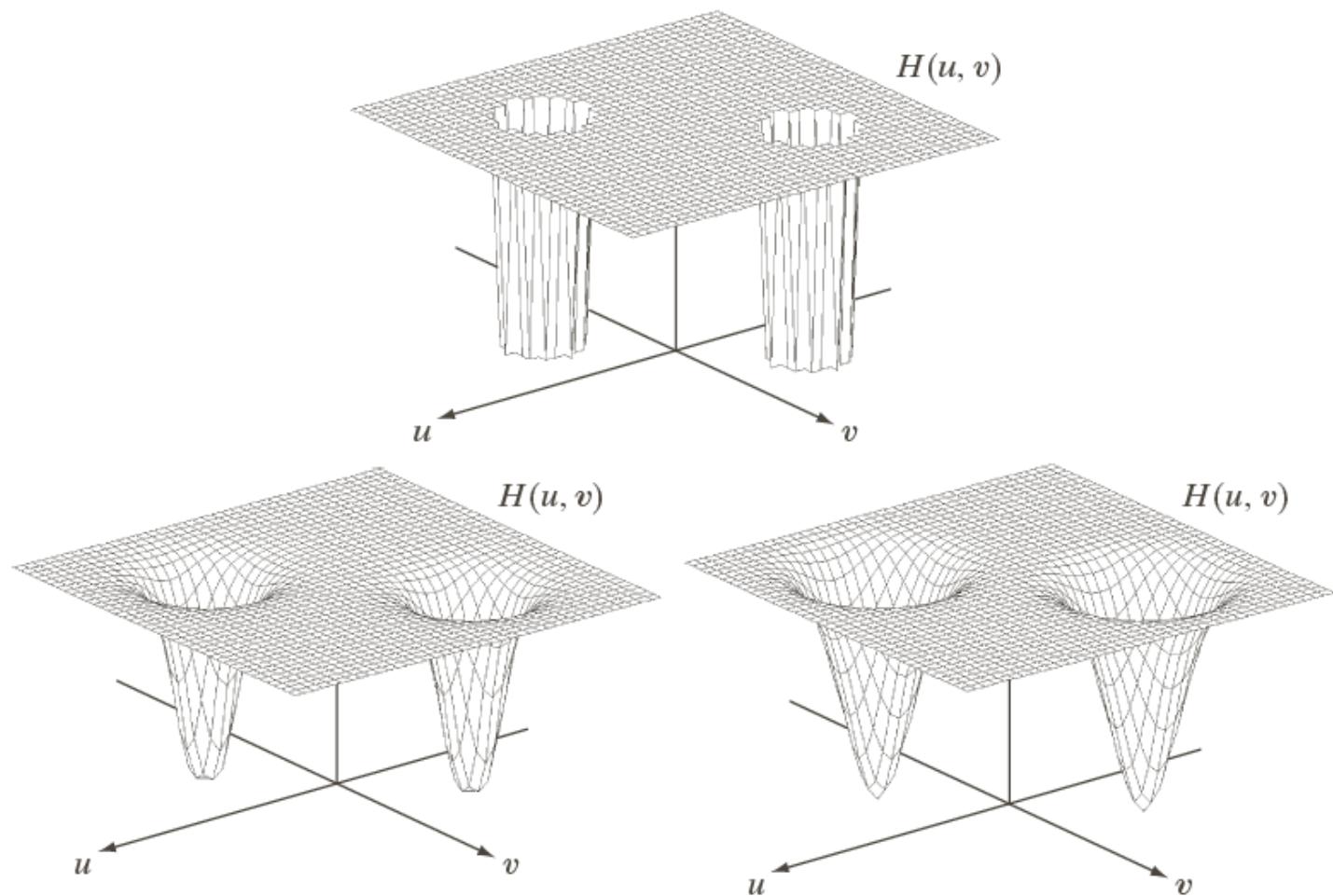
A Butterworth bandreject filter of order 4, with the appropriate radius and width to enclose completely the noise impulses

Perspective Plots of Notch Filters

a
b | c

FIGURE 5.18

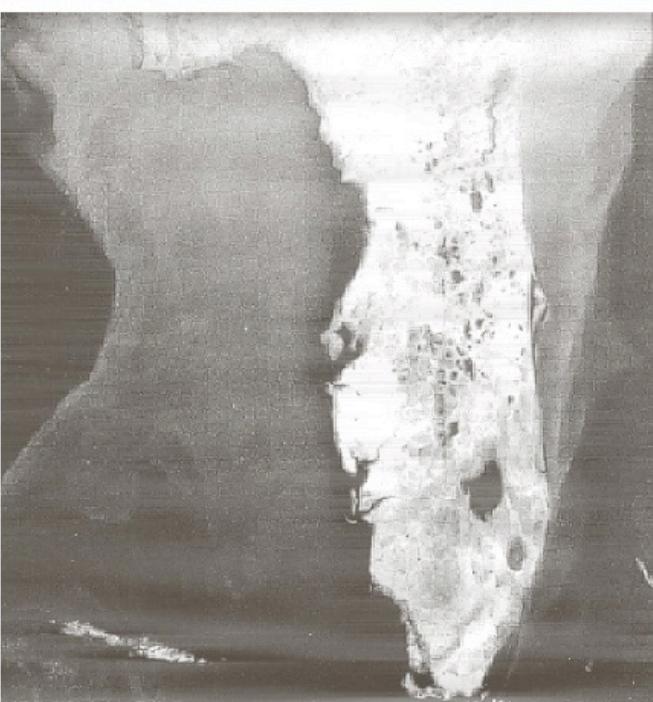
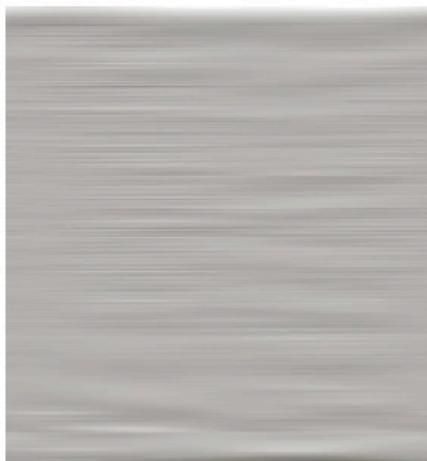
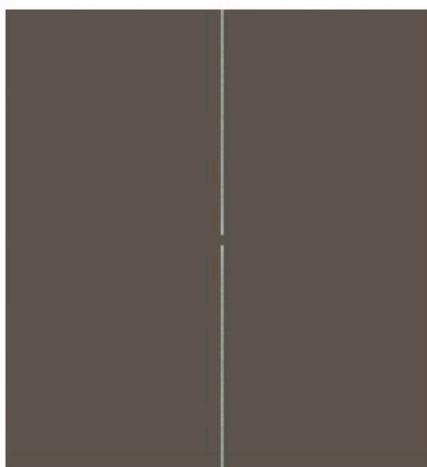
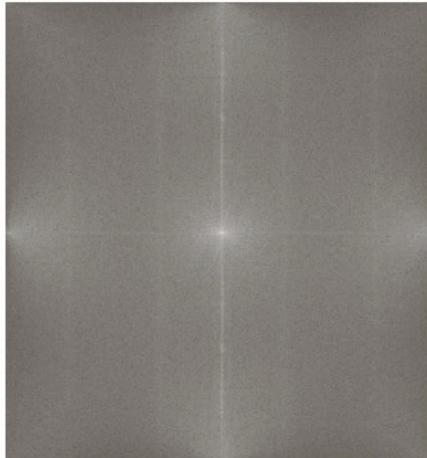
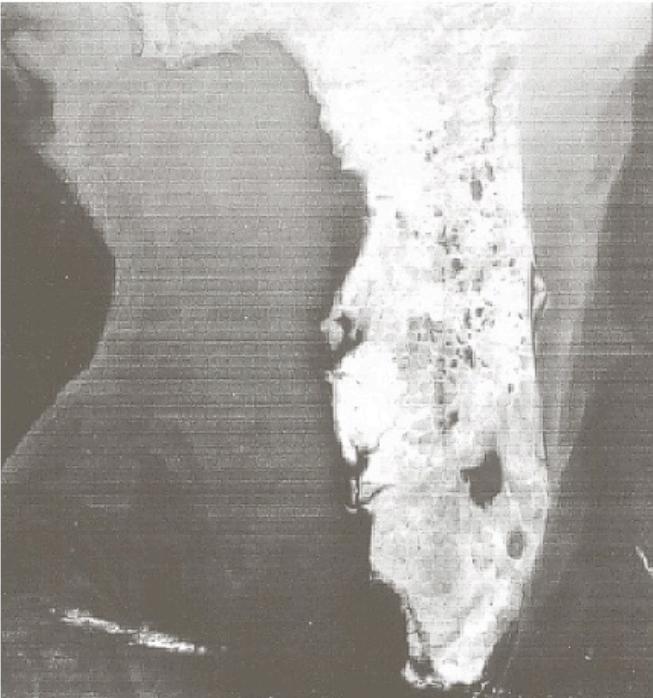
Perspective plots of (a) ideal, (b) Butterworth (of order 2), and (c) Gaussian notch (reject) filters.



a
b
c
d

FIGURE 5.19

- (a) Satellite image of Florida and the Gulf of Mexico showing horizontal scan lines.
(b) Spectrum. (c) Notch pass filter superimposed on (b). (d) Spatial noise pattern. (e) Result of notch reject filtering.
(Original image courtesy of NOAA.)



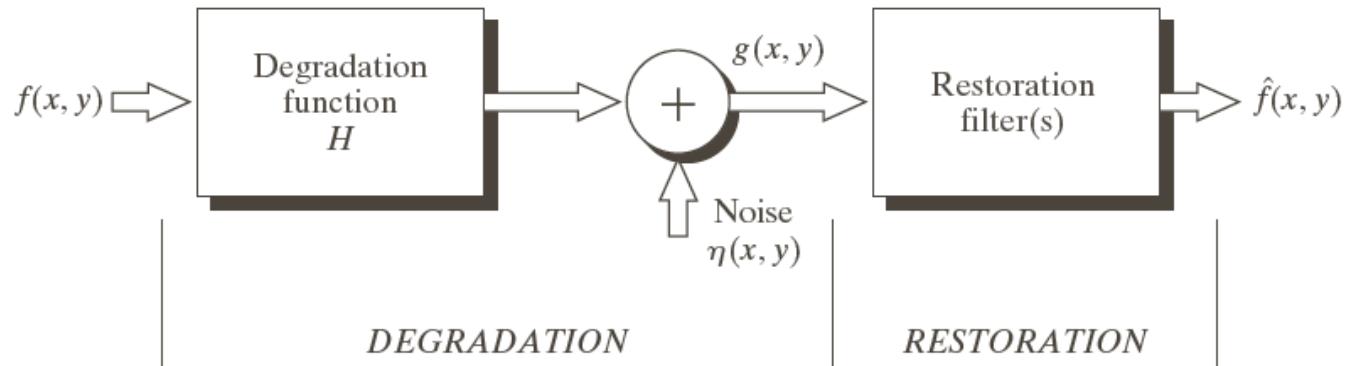
Exercise

- Select one of your own image with lots of edges
- Plot the frequency transform output
- Perform notch filtering in frequency and perform inverse transform
- Analyze your result
- You could define any information as “noise”
- Use `fft2_show.m` as an example code

Linear, Position-Invariant Degradations

FIGURE 5.1

A model of the image degradation/restoration process.



$$g(x, y) = H[f(x, y)] + \eta(x, y)$$

Linear, Position-Invariant Degradations

H is linear

$$H[af_1(x, y) + bf_2(x, y)] = aH[f_1(x, y)] + bH[f_2(x, y)]$$

f_1 and f_2 are any two input images.

An operator having the input-output relationship

$g(x, y) = H[f(x, y)]$ is said to be position invariant

if

$$H[f(x - \alpha, y - \beta)] = g(x - \alpha, y - \beta)$$

for any $f(x, y)$ and any α and β .

Linear, Position-Invariant Degradations

In the presence of additive noise,
if H is a linear operator and position invariant,

$$\begin{aligned} g(x, y) &= \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} f(\alpha, \beta) h(x - \alpha, y - \beta) d\alpha d\beta + \eta(x, y) \\ &= h(x, y) \star f(x, y) + \eta(x, y) \end{aligned}$$

$$G(u, v) = H(u, v)F(u, v) + N(u, v)$$

Estimating the Degradation Function

- ▶ Three principal ways to estimate the degradation function

1. Observation

2. Experimentation

3. Mathematical Modeling

Mathematical Modeling (1)

- ▶ Environmental conditions cause degradation

A model about atmospheric turbulence

$$H(u, v) = e^{-k(u^2 + v^2)^{5/6}}$$

k : a constant that depends on
the nature of the turbulence

a
b
c
d

FIGURE 5.25

Illustration of the atmospheric turbulence model.

(a) Negligible turbulence.

(b) Severe turbulence,
 $k = 0.0025$.

(c) Mild turbulence,
 $k = 0.001$.

(d) Low turbulence,
 $k = 0.00025$.

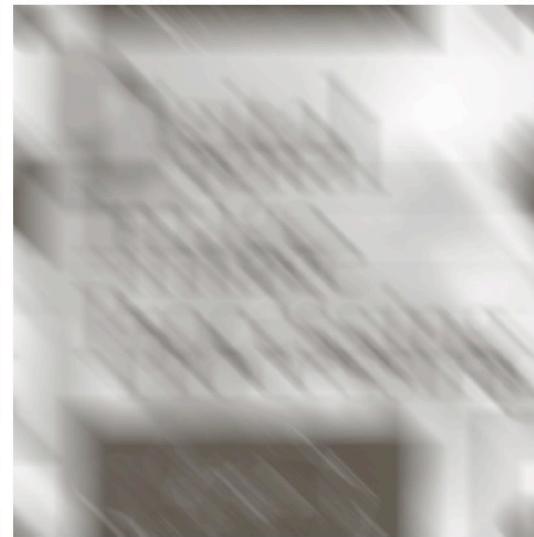
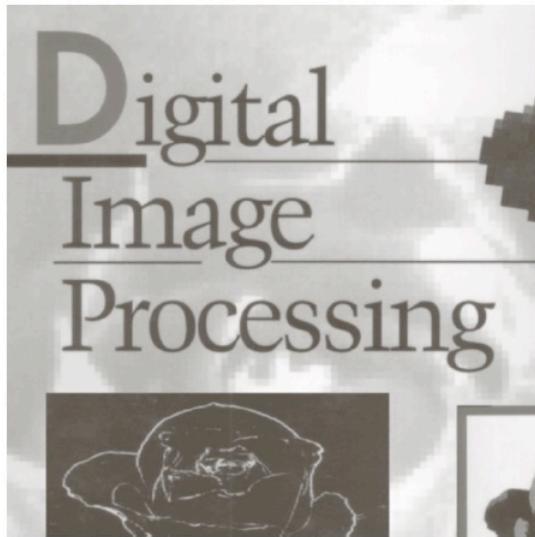
(Original image courtesy of NASA.)



Mathematical Modeling (2)

- ▶ Derive a mathematical model from basic principles

E.g., An image blurred by uniform linear motion between the image and the sensor during image acquisition



Mathematical Modeling (3)

Suppose that an image $f(x, y)$ undergoes planar motion, $x_0(t)$ and $y_0(t)$ are the time-varying components of motion in the x - and y -directions, respectively.

The optical imaging process is perfect. T is the duration of the exposure. The blurred image $g(x, y)$

$$g(x, y) = \int_0^T f[x - x_0(t), y - y_0(t)] dt$$

Mathematical Modeling (4)

$$g(x, y) = \int_0^T f[x - x_0(t), y - y_0(t)] dt$$

$$G(u, v) = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} g(x, y) e^{-j2\pi(ux+vy)} dx dy$$

$$= \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \left[\int_0^T f[x - x_0(t), y - y_0(t)] dt \right] e^{-j2\pi(ux+vy)} dx dy$$

$$= \int_0^T \left[\int_{-\infty}^{\infty} \int_{-\infty}^{\infty} f[x - x_0(t), y - y_0(t)] e^{-j2\pi(ux+vy)} dx dy \right] dt$$

$$= \int_0^T F(u, v) e^{-j2\pi[ux_0(t)+vy_0(t)]} dt$$

$$= F(u, v) \int_0^T e^{-j2\pi[ux_0(t)+vy_0(t)]} dt$$

Mathematical Modeling (4)

$$H(u, v) = \int_0^T e^{-j2\pi[ux_0(t)+vy_0(t)]} dt$$

Suppose that the image undergoes uniform linear motion in the x -direction only, at a rate given by $x_0(t) = at / T$.

$$\begin{aligned} H(u, v) &= \int_0^T e^{-j2\pi ux_0(t)} dt \\ &= \int_0^T e^{-j2\pi uat/T} dt \\ &= \frac{T}{\pi ua} \sin(\pi ua) e^{-j\pi ua} \end{aligned}$$

Mathematical Modeling (5)

Suppose that the image undergoes uniform linear motion in the x -direction and y -direction, at a rate given by

$$x_0(t) = at / T \text{ and } y_0(t) = bt / T$$

$$\begin{aligned} H(u, v) &= \int_0^T e^{-j2\pi[u x_0(t) + v y_0(t)]} dt \\ &= \int_0^T e^{-j2\pi[ua + vb]t/T} dt \end{aligned}$$

Inverse Filtering

An estimate of the transform of the original image

$$\hat{f}(x, y) = \frac{G(u, v)}{H(u, v)}$$

$$\begin{aligned}\hat{f}(x, y) &= \frac{F(u, v)H(u, v) + N(u, v)}{H(u, v)} \\ &= F(u, v) + \frac{N(u, v)}{H(u, v)}\end{aligned}$$

Inverse Filtering

$$\hat{f}(x, y) = F(u, v) + \frac{N(u, v)}{H(u, v)}$$

1. We can't exactly recover the undegraded image because $N(u, v)$ is not known.

Minimum Mean Square Error (Wiener) Filtering

- **N. Wiener (1942)**

- **Objective**

Find an estimate of the uncorrupted image such that the mean square error between them is minimized

$$e^2 = E \left\{ (f - \hat{f})^2 \right\}$$

Minimum Mean Square Error (Wiener) Filtering

The minimum of the error function is given in the frequency domain by the expression

$$\begin{aligned}\hat{f}(x, y) &= \left[\frac{H^*(u, v) S_f(u, v)}{S_f(u, v) |H(u, v)|^2 + S_\eta(u, v)} \right] G(u, v) \\ &= \left[\frac{H^*(u, v)}{|H(u, v)|^2 + S_\eta(u, v) / S_f(u, v)} \right] G(u, v) \\ &= \left[\frac{1}{H(u, v)} \frac{|H(u, v)|^2}{|H(u, v)|^2 + S_\eta(u, v) / S_f(u, v)} \right] G(u, v)\end{aligned}$$

Minimum Mean Square Error (Wiener) Filtering

$$\hat{f}(x, y) = \left[\frac{1}{H(u, v)} \frac{|H(u, v)|^2}{|H(u, v)|^2 + S_\eta(u, v) / S_f(u, v)} \right] G(u, v)$$

$H(u, v)$: degradation function

$H^*(u, v)$: complex conjugate of $H(u, v)$

$$|H(u, v)|^2 = H^*(u, v)H(u, v)$$

$S_\eta(u, v) = |N(u, v)|^2$ = power spectrum of the noise

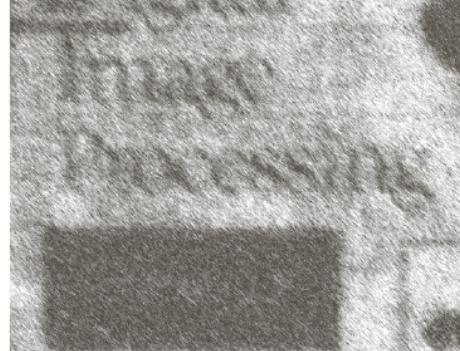
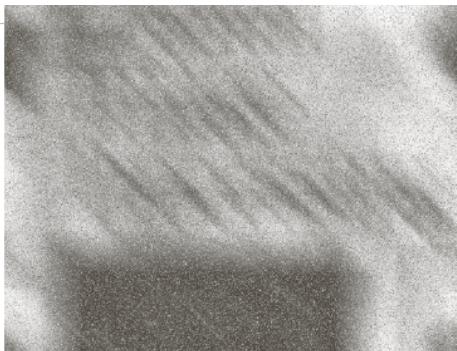
$S_f(u, v) = |F(u, v)|^2$ = power spectrum of the undegraded image

Minimum Mean Square Error (Wiener) Filtering

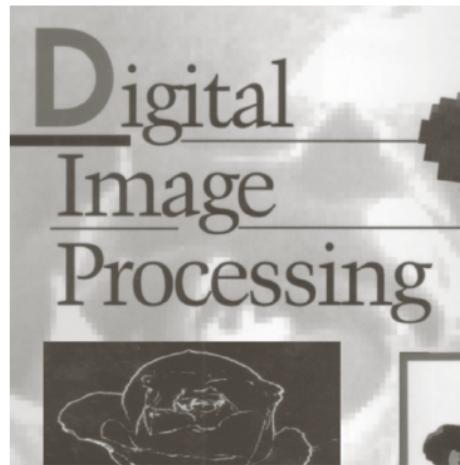
$$\hat{f}(x, y) = \left[\frac{1}{H(u, v)} \frac{|H(u, v)|^2}{|H(u, v)|^2 + K} \right] G(u, v)$$

K is a specified constant. Generally, the value of K is chosen interactively to yield the best visual results.

Left:
degraded
image



Right: Wiener
filtering



Geometric Mean Filter

$$\hat{f}(x, y) = \left[\frac{H^*(u, v)}{|H(u, v)|^2} \right]^\alpha \left[\frac{|H(u, v)|^2}{|H(u, v)|^2 + \beta [S_\eta(u, v) / S_f(u, v)]} \right]^{1-\alpha} G(u, v)$$

$\alpha = 1$: inverse filter

$\alpha = 0$: parametric Wiener filter

$\alpha = 1/2$: geometric mean filter

Exercise

- Gaussian noise
- Uniform noise
- Salt & Pepper
- Averaging filter, LPF
- Averaging filter, LPF
- Median filter

Exercise

- Gaussian noise
 - Uniform noise
 - Rayleigh noise
 - Exponential noise
 - Salt & Pepper
 - Randn()
 - Rand()
 - $a + \sqrt{-b \cdot \ln(1-U(0,1))}$
 - $-(1/a) \cdot \ln(1-U(0,1))$
 - ?
-
- “Log” inn matlab is Natural logarithm

Syntax

```
J = imnoise(I,type)
J = imnoise(I,type,parameters)
J = imnoise(I,'gaussian',M,V)
J = imnoise(I,'localvar',V)
J = imnoise(I,'localvar',image_intensity,var)
J = imnoise(I,'poisson')
J = imnoise(I,'salt & pepper',d)
J = imnoise(I,'speckle',v)
gpuarrayJ = imnoise(gpuarrayI, __)
```

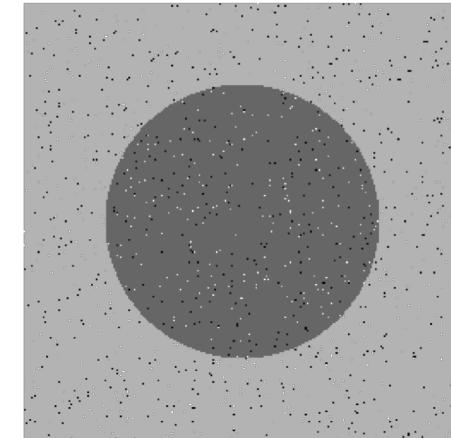
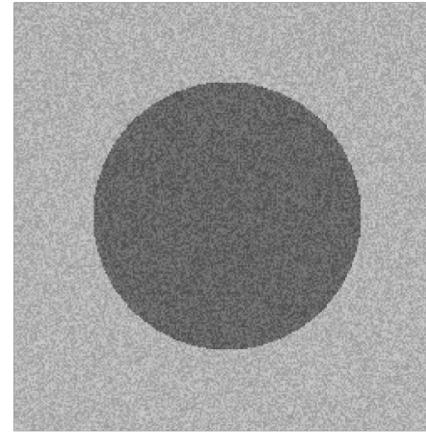
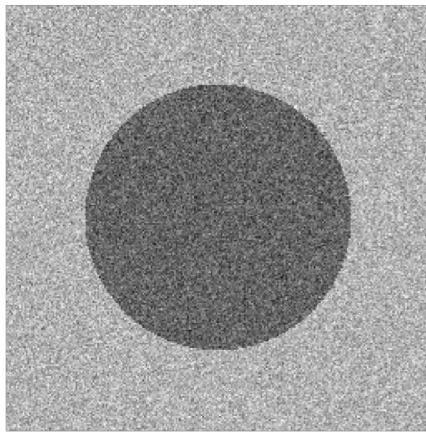
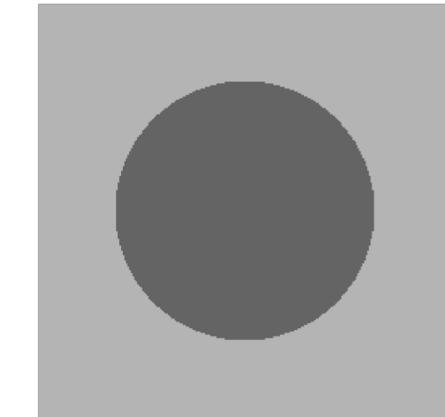
Description

`J = imnoise(I,type)` adds noise of a given type to the intensity image `I`. `type` is a string that specifies any of the following types of noise. Note that certain types of noise support additional parameters. See the related syntax.

Value	Description
'gaussian'	Gaussian white noise with constant mean and variance
'localvar'	Zero-mean Gaussian white noise with an intensity-dependent variance
'poisson'	Poisson noise
'salt & pepper'	On and off pixels
'speckle'	Multiplicative noise

Exercise

- 256 by 256 double type [0,1]
 - $\sqrt{(i-128)^2 + (j-128)^2} < 80$ Center part set 0.4
 - Other part set 0.7
-
- 1, Gaussian noise (0, 0.01)
 - 2, uniform noise [-0.05 0.05]
 - 3, Salt and Pepper Ps=Pp=0.02
-
- Plot histograms
 - Apply averaging and median filter



Exercise

