Preliminary Report

Team Member	PSU ID	Contributions	Percentages
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Part A:

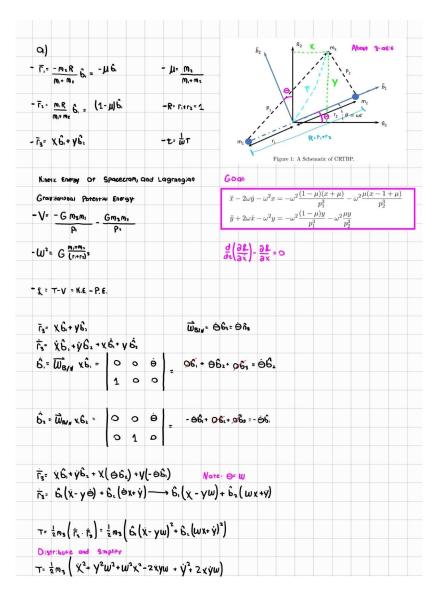


Figure 1: Finding Velocity Vector, and Kinetic Energy

$$\begin{array}{c} Gravity & QS & Q & force: (nqs) & direction) \\ \hline \overrightarrow{F_6} = \frac{-G}{G} \underbrace{m_5 m_1}_{P_1} \cdot \widehat{p}_1 - \underbrace{\frac{Gm_5 m_5}{P_2}}_{P_2} \cdot \widehat{p}_2 + \underbrace{\frac{P_1}{|\widehat{P_1}|}}_{P_1} = \frac{1}{\widehat{P_1}} \left(\left(X + \mu \right)^2 \widehat{b}_1 + V^2 \widehat{b}_2 \right) \\ \hline \overrightarrow{P_2} = \left(X - \left(1 - \mu \right) \right)^3 \widehat{b}_1 + V^2 \widehat{b}_2 \rightarrow \left(X - 1 + \mu \right)^3 \widehat{b}_1 + V^2 \widehat{b}_2 \\ \hline \overrightarrow{P_2} = \left(X - \left(1 - \mu \right) \right)^3 \widehat{b}_1 + V^2 \widehat{b}_2 \rightarrow \left(X - 1 + \mu \right)^3 \widehat{b}_1 + V^2 \widehat{b}_2 \\ \hline \overrightarrow{P_2} = \left(X - \left(1 - \mu \right) \right)^3 \widehat{b}_1 + V^2 \widehat{b}_2 \rightarrow \left(X - 1 + \mu \right)^3 \widehat{b}_1 + V^2 \widehat{b}_2 \\ \hline \overrightarrow{P_2} = \left(X - \left(1 - \mu \right) \right)^3 \widehat{b}_1 + V^2 \widehat{b}_2 \rightarrow \left(X - 1 + \mu \right)^3 \widehat{b}_1 + V^2 \widehat{b}_2 \right) \\ \hline \overrightarrow{P_6} = \frac{-Gm_5 m_1}{P_1} \cdot \frac{1}{P_1} \left(\left(X + \mu \right)^2 \widehat{b}_1 + V^2 \widehat{b}_2 \right)^{-3\lambda_1} - \frac{Gm_5 m_2}{P_2} \cdot \left(\left(X - 1 + \mu \right)^2 \widehat{b}_2 + V^2 \widehat{b}_2 \right)^{-3\lambda_2} \\ \hline \overrightarrow{P_6} = \frac{-Gm_5 m_1}{P_1} \cdot \left(X + \mu \right)^2 \widehat{b}_1 + V^2 \widehat{b}_2 \right)^{-3\lambda_2} \cdot \left(\left(X - 1 + \mu \right)^2 \widehat{b}_2 + V^2 \widehat{b}_2 \right)^{-3\lambda_2} \\ \hline \overrightarrow{B_6} = \frac{-Gm_5 m_1}{P_1} \cdot \left(X + \mu \right)^2 \widehat{b}_1 + V^2 \widehat{b}_2 \right)^{-3\lambda_1} \cdot \left(\left(X - 1 + \mu \right)^2 \widehat{b}_2 + V^2 \widehat{b}_2 \right)^{-3\lambda_2} \\ \hline \overrightarrow{B_6} = \frac{-Gm_5 m_1}{P_1} \cdot \left(X + \mu \right)^2 \widehat{b}_1 + V^2 \widehat{b}_2 \right)^{-3\lambda_1} \cdot \left(\left(X - 1 + \mu \right)^2 \widehat{b}_2 + V^2 \widehat{b}_2 \right)^{-3\lambda_2} \\ \hline \overrightarrow{B_6} = \frac{-Gm_5 m_1}{P_1} \cdot \left(X + \mu \right)^2 \widehat{b}_1 + V^2 \widehat{b}_2 \right)^{-3\lambda_1} \cdot \left(\left(X - 1 + \mu \right)^2 \widehat{b}_2 + V^2 \widehat{b}_2 \right)^{-3\lambda_2} \\ \hline \overrightarrow{B_6} = \frac{-Gm_5 m_1}{P_1} \cdot \left(X + \mu \right)^2 \widehat{b}_1 + V^2 \widehat{b}_2 \right)^{-3\lambda_1} \cdot \left(\left(X - 1 + \mu \right)^2 \widehat{b}_2 + V^2 \widehat{b}_2 \right)^{-3\lambda_2} \right)$$

Figure 2: Vectorizing Force of Gravity for Potential Energy. Defining P1 and P2 in terms of the b-frame. Solving for Lagrangian.

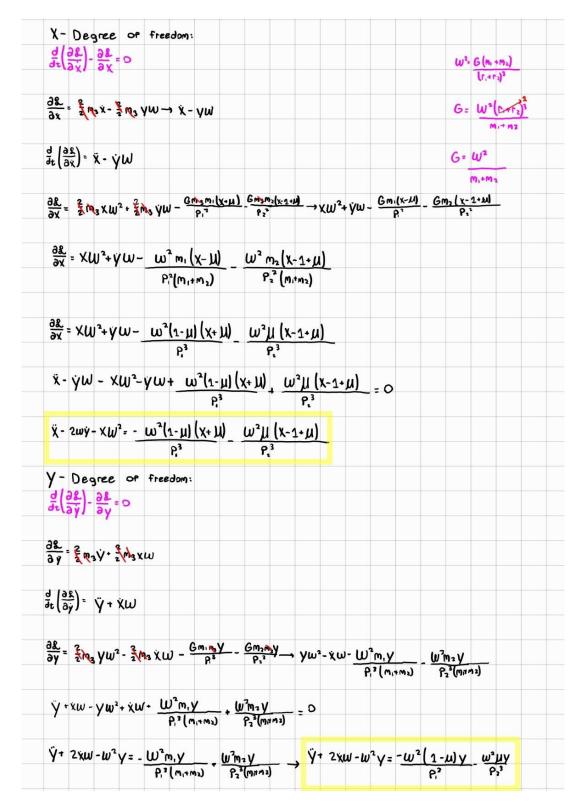


Figure 3: Using Lagrangian method to formulate two separate equations of motion around each degree of freedom.

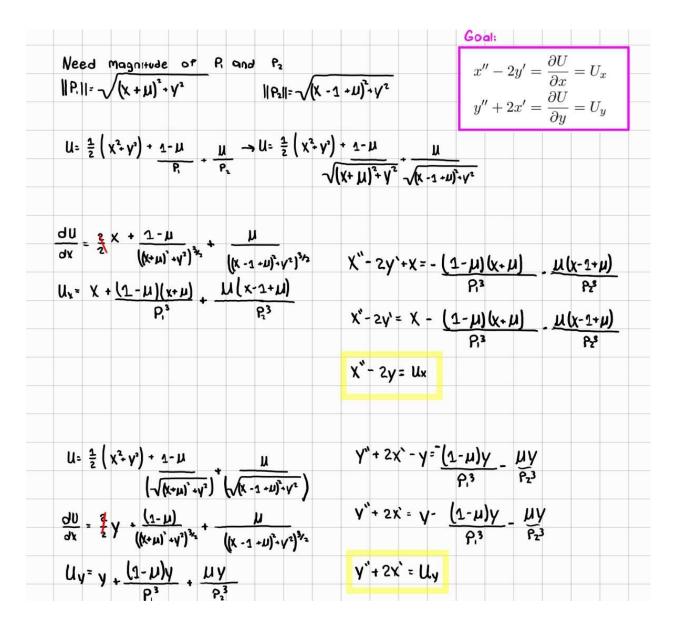


Figure 4: Solving for Ux and Uy, then rewriting Equations of Motion.

Part B:

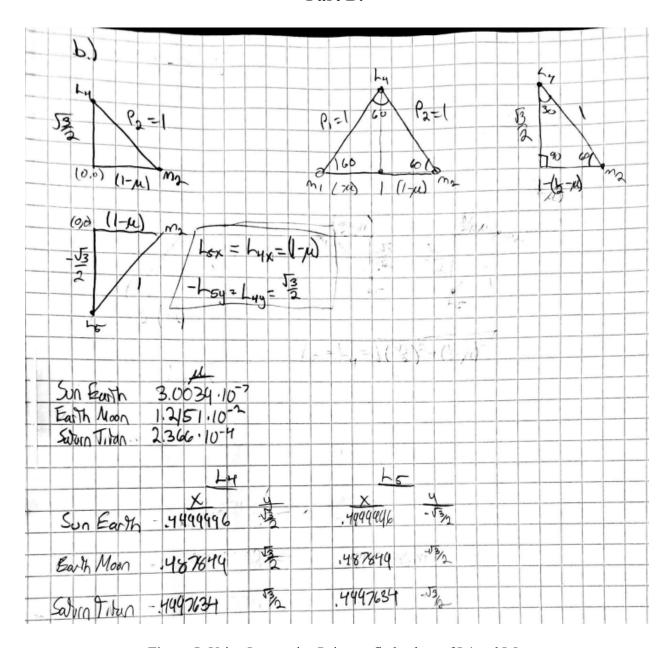


Figure 5: Using Lagrangian Points to find values of L4 and L5.

$$\lambda'' - 2y' = \frac{\partial U}{\partial x} = U_{3c} \left(y'' + 2x' = \frac{\partial U}{\partial y} = U_{y} \right)$$

$$\lim_{N \to \infty} \frac{\partial F}{\partial x} = \int_{0}^{\infty} \int_{0}^{\infty} dx + \frac{\partial F}{\partial y} = \int_{0}^{\infty} \int_{0}^{\infty} dx' + \frac{\partial F}{\partial y'} \int_{0}^{\infty} dy'' + \frac{\partial F}{\partial y''} \int_{0}^{\infty} dx'' + \frac{\partial F}{\partial y''} \int_{0}^{\infty}$$

Figure 6: Steps of linearizing equations of motion.

Note: Code Submitted in Gradescope

Table 1: Earth and Moon

L1	-2.93210331736619 + 0.000000000000000i	Unstable
	2.93210331736619 + 0.0000000000000000i	
	2.77555756156289e-17 + 2.33441574278267i	
	2.77555756156289e-17 - 2.33441574278267i	
L2	2.15875415806544 + 0.0000000000000000i	Unstable
	-2.15875415806544 + 0.000000000000000i	
	0.00000000000000 + 1.86269258455063i	
	0.0000000000000 - 1.86269258455063i	
L3	-1.43765208071578e-16 + 1.01042841824630i	Unstable
	-1.43765208071578e-16 - 1.01042841824630i	
	-0.177948963939713 + 0.000000000000000i	
	0.177948963939713 + 0.0000000000000000000000000000000000	
L4	2.07299455379228e-15 + 0.954499117832690i	Stable
	2.07299455379228e-15 - 0.954499117832690i	
	-1.63085690785270e-15 + 0.298213738879709i	
	-1.63085690785270e-15 - 0.298213738879709i	
L5	-2.07299455379228e-15 + 0.954499117832690i	Stable
	-2.07299455379228e-15 - 0.954499117832690i	
	1.63085690785270e-15 + 0.298213738879709i	
	1.63085690785270e-15 - 0.298213738879709i	

Table 2: Sun and Earth

L1	-2.51927176915253 + 0.00000000000000i	Unstable
	2.51927176915253 + 0.000000000000000i	
	-1.11022302462516e-16 + 2.07828726022754i	
	-1.11022302462516e-16 - 2.07828726022754i	
L2	-2.50766802961189 + 0.00000000000000i	Unstable
	2.50766802961189 + 0.000000000000000i	
	1.66533453693773e-16 + 2.07121742796525i	
	1.66533453693773e-16 - 2.07121742796525i	
L3	-1.51788304147971e-18 + 0.999970636745927i	Stable
	-1.51788304147971e-18 - 0.999970636745927i	
	3.68628738645072e-18 + 0.00938545387897627i	
	3.68628738645072e-18 - 0.00938545387897627i	
L4	9.43689570931383e-16 + 0.999998986181485i	Stable
	9.43689570931383e-16 - 0.999998986181485i'	
	-8.55153397784201e-16 + 0.00142395084250021i	
	-8.55153397784201e-16 - 0.00142395084250021i	
L5	-9.43689570931383e-16 + 0.999998986181485i	Stable
	-9.43689570931383e-16 - 0.999998986181485i	
	8.55153397784201e-16 + 0.00142395084250021i	
	8.55153397784201e-16 - 0.00142395084250021i	

Table 3: Saturn and Titan

L1	2.61513098346219 + 0.0000000000000000i	Unstable
	-2.61513098346219 + 0.00000000000000i	
	0.00000000000000 + 2.13698484537958i	
	0.0000000000000 - 2.13698484537958i	
L2	-2.47314812523973 + 0.00000000000000i	Unstable
	2.47314812523973 + 0.000000000000000i	
	-4.99600361081320e-16 + 2.05023249232707i	
	-4.99600361081320e-16 - 2.05023249232707i	
L3	2.08166817117217e-17 + 1.00020269334695i	Unstable
	2.08166817117217e-17 - 1.00020269334695i	
	-0.0246621944870533 + 0.000000000000000i	
	0.0246621944870533 + 0.000000000000000i	
L4	1.38777878078145e-16 + 0.999200065215411i	Stable
	1.38777878078145e-16 - 0.999200065215411i	
	-3.98921177937727e-16 + 0.0399903697596923i	
	-3.98921177937727e-16 - 0.0399903697596923i	
L5	-1.38777878078145e-16 + 0.999200065215411i	Stable
	-1.38777878078145e-16 - 0.999200065215411i	
	3.98921177937727e-16 + 0.0399903697596923i	
	3.98921177937727e-16 - 0.0399903697596923i	

Part D and E:

Future Approach:

- Uxx = Uxx(mu,x,y): used to calculate the second derivative U versus x. The function takes x and y as input, where x is the x-position of the Lagrange point and y is the y-position of the Lagrange point.
- Uyy = Uyy(mu,x,y): Used to calculate the second derivative of U with respect to y.
- Uxy = Uxy(mu,x,y): calculates the first derivative of U with respect to x, then for y.

The function ev takes x and y as input, where x is the x location of the Lagrange point and y is the y location of the Lagrange point and mu for different planets.

•Matrix A uses the equations found above and then the eig function produces the eigenvalues.

Procedure:

- Matlab function ode45 will be used to integrate.
- The x and y positions obtained from the integration will then be plotted.
- These x and y positions will then be converted from the body frame to the inertial frame using a loop.
- These new points will be then plotted using matlab functions.
- Simulating orbits in Earth and Moon system by adding the initial conditions found in files 'EML2304P1.mat'.
- Ode45 was reused, but with a different initial condition for Earth's sun system.