

Mechanics II

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Q.1 Part(a)

The invariant space-time interval ds^2 is given by:

$$ds^2 = (dx^0)^2 - (dx^1)^2 - (dx^2)^2 - (dx^3)^2$$

For a light-like (null) trajectory, the space-time interval must satisfy $ds^2 = 0$. In the x^0 - x^1 plane, the null condition is:

$$(dx^0)^2 - (dx^1)^2 = 0 \quad \Rightarrow \quad dx^0 = \pm dx^1$$

This equation represents the path of light, where $x^0 = x^1$ or $x^0 = -x^1$, corresponding to the light cone's axes in x^0 - x^1 plane. We thus define *light cone coordinates* in the following way, to simplify the analysis of light-like trajectories.

$$x^+ = x^0 + x^1 \quad \text{and} \quad x^- = x^0 - x^1$$

In these coordinates, light moving along $x^1 = x^0$ corresponds to $x^- = 0$, and light moving along $x^1 = -x^0$ corresponds to $x^+ = 0$. We are thus aligning the coordinates axes along the light-like directions, making calculations involving light propagation easier. These coordinates are named *light cone coordinates* because they are aligned with the light cone structure of null paths.

Part(b)

The space-time invariant in terms of standard coordinates (x^0, x^1, x^2, x^3) can be written as

$$ds^2 = (dx^0)^2 - (dx^1)^2 - (dx^2)^2 - (dx^3)^2$$

Now, using the transformations for light cone coordinates:

$$x^+ = x^0 + x^1, \quad x^- = x^0 - x^1$$

Taking differential of the above Equation, we get:

$$dx^+ = dx^0 + dx^1, \quad dx^- = dx^0 - dx^1$$

Since $dx^+ dx^- = (dx^0)^2 - (dx^1)^2$, therefore the space-time invariant in terms of *light cone coordinates* (x^+, x^-, x^2, x^3) can be written as :

$$ds^2 = dx^+ dx^- - (dx^2)^2 - (dx^3)^2$$

To find the metric η in *light cone coordinates*, we can use the following identity:

$$\boxed{\eta_{\mu\nu} dx^\mu dx^\nu = ds^2 = dx^+ dx^- - (dx^2)^2 - (dx^3)^2}$$

where $dx^\mu = (dx^+, dx^-, dx^2, dx^3)$ and $\mu, \nu \in \{+, -, 2, 3\}$. On solving the above Equation we get the following equations in *light cone coordinates*:

$$(\eta_{+-} + \eta_{-+}) dx^+ dx^- = dx^+ dx^- \Rightarrow \eta_{+-} + \eta_{-+} = 1$$

$$\eta_{22} (dx^2)^2 = -(dx^2)^2 \Rightarrow \eta_{22} = -1$$

$$\eta_{33} (dx^3)^2 = -(dx^3)^2 \Rightarrow \eta_{33} = -1$$

$$\eta_{\mu\nu} = 0 \quad \forall \quad \eta_{\mu\nu} \notin \{\eta_{+-}, \eta_{-+}, \eta_{22}, \eta_{33}\}$$

Therefore the metric η in *light cone coordinates* can be written as:

$$\boxed{\eta = \begin{pmatrix} 0 & r & 0 & 0 \\ (1-r) & 0 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix}, \quad \text{where } r \in \mathbb{R}}$$

Part(c)

The special Lorentz transformation (boost along x^1) is given by:

$$x'^0 = \gamma (x^0 - \beta x^1), \quad x'^1 = \gamma (x^1 - \beta x^0), \quad x'^2 = x^2, \quad x'^3 = x^3 \quad (1)$$

where $\beta = \frac{v}{c}$ and $\gamma = \frac{1}{\sqrt{1-\beta^2}}$.

The light cone coordinates are given as:

$$x^+ = x^0 + x^1, \quad x^- = x^0 - x^1 \quad (2)$$

where the other two coordinates x^2 and x^3 remain unchanged. Now, let's compute the Lorentz transformations in terms of x^+ and x^- .

Note that since it is special Lorentz Transformation (boost along x^1), $x'^2 = x^2$ and $x'^3 = x^3$ in light cone coordinates. Using Eqn(1) and Eqn(2), we can write x'^+ and x'^- as:

$$x'^+ = x'^0 + x'^1 = \gamma (1 - \beta) x^0 + \gamma (1 - \beta) x^1 = \gamma (1 - \beta) (x^0 + x^1) = \gamma (1 - \beta) x^+$$

$$x'^- = x'^0 - x'^1 = \gamma (1 + \beta) x^0 - \gamma (1 + \beta) x^1 = \gamma (1 + \beta) (x^0 - x^1) = \gamma (1 + \beta) x^-$$

Thus, in matrix form, the special Lorentz transformation (boost along x^1) in light cone coordinates is given by:

$$\boxed{\begin{pmatrix} x'^+ \\ x'^- \\ x'^2 \\ x'^3 \end{pmatrix} = \begin{pmatrix} \gamma(1-\beta) & 0 & 0 & 0 \\ 0 & \gamma(1+\beta) & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} x^+ \\ x^- \\ x^2 \\ x^3 \end{pmatrix}}$$

Thus the special Lorentz Transformation matrix (boost along x^1) in light cone coordinates is given as:

$$\begin{pmatrix} \gamma(1-\beta) & 0 & 0 & 0 \\ 0 & \gamma(1+\beta) & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$$

Part(d)

In *light cone coordinates*, a general 4-vector will be of the form $A^\mu = (A^+, A^-, A^2, A^3)$. The corresponding 4-covector will be of the form $A_\mu = (A_+, A_-, A_2, A_3)$. The 4-covector is related to the 4-vector through the metric η as:

$$\begin{aligned} & \boxed{A_\mu = \eta A^\mu} \\ \Rightarrow \quad \begin{pmatrix} A_+ \\ A_- \\ A_2 \\ A_3 \end{pmatrix} &= \begin{pmatrix} 0 & r & 0 & 0 \\ (1-r) & 0 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix} \begin{pmatrix} A^+ \\ A^- \\ A^2 \\ A^3 \end{pmatrix}, \quad \text{where } r \in \mathbb{R} \end{aligned}$$

Thus in *light cone coordinates*, the 4-covector in terms of the components of 4-vector can be written as:

$$\boxed{A_\mu = (rA^-, (1-r)A^+, -A^2, -A^3), \quad \text{where } r \in \mathbb{R}}$$

Q.2 Part(a)

We begin with:

$$T = \int_{\vec{v}=\vec{0}}^{\vec{v}=\vec{u}} \frac{d}{dt} (m\gamma(v)\vec{v}) \cdot d\vec{r} = \int_{\vec{v}=\vec{0}}^{\vec{v}=\vec{u}} d(m\gamma(v)\vec{v}) \cdot \frac{d\vec{r}}{dt} = \int_{\vec{v}=\vec{0}}^{\vec{v}=\vec{u}} d(m\gamma(v)\vec{v}) \cdot \vec{v} \quad (3)$$

Note that m is the rest mass. Relativistic mass m_r is given as $m_r = m\gamma(v)$. Therefore taking differential of the term $m\gamma(v)\vec{v}$, we get:

$$d(m\gamma(v)\vec{v}) = d(m_r\vec{v}) = m_r d\vec{v} + \vec{v} dm_r \quad (4)$$

Substituting Eqn(4) in Eqn(3), we get:

$$T = \int_{\vec{v}=\vec{0}}^{\vec{v}=\vec{u}} (m_r d\vec{v} + \vec{v} dm_r) \cdot \vec{v} = \int_{\vec{v}=\vec{0}}^{\vec{v}=\vec{u}} (m_r \vec{v} \cdot d\vec{v} + v^2 dm_r) \quad (5)$$

Now we know that:

$$m_r = \frac{m}{\sqrt{1 - \frac{v^2}{c^2}}} \quad (6)$$

$$\Rightarrow m_r^2 c^2 - m_r^2 v^2 = m^2 c^2 \quad (7)$$

Taking differential of both sides in Eqn(7), we get:

$$2m_r c^2 dm_r - 2m_r^2 \vec{v} \cdot d\vec{v} - v^2 2m_r dm_r = 0 \quad (8)$$

Dividing both sides of Eqn(8) by $2m_r$, we get:

$$m_r \vec{v} \cdot d\vec{v} + v^2 dm_r = c^2 dm_r \quad (9)$$

Substituting Eqn(9) in Eqn(5), we get:

$$T = \int_{\vec{v}=\vec{0}}^{\vec{v}=\vec{u}} c^2 dm_r = c^2 \int_{m_r=m}^{m_r=m\gamma(u)} dm_r = m\gamma(u)c^2 - mc^2 = m(\gamma(u) - 1)c^2 \quad (10)$$

Hence the relativistic expression for kinetic energy for a point particle is given as:

$$\boxed{T = m(\gamma(u) - 1)c^2} \quad (11)$$

Part(b)

We know that:

$$\gamma(u) = \frac{1}{\sqrt{1 - \frac{u^2}{c^2}}} = \left(1 - \frac{u^2}{c^2}\right)^{-\frac{1}{2}} \quad (12)$$

When $u \ll c$, $\left|\frac{u^2}{c^2}\right| < 1$ and thus we can use binomial expansion and write Eqn(12) as:

$$\gamma(u) = 1 + \frac{u^2}{2c^2} + \frac{3u^4}{8c^4} + \dots \quad (13)$$

Substituting Eqn(13) in Eqn(11), we get:

$$T = m \left(1 + \frac{u^2}{2c^2} + \frac{3u^4}{8c^4} + \dots - 1\right) c^2 = m \left(\frac{u^2}{2c^2} + \frac{3u^4}{8c^4} + \dots\right) c^2 \quad (14)$$

When $u \ll c$, we just need to consider the first term in Eqn(14) as other higher order terms would be negligibly small. Hence Eqn(14) reduces to:

$$\boxed{T = \frac{1}{2}mu^2} \quad (15)$$

Hence the relativistic kinetic energy expression for a point particle reduces to the familiar Newtonian formula when $u \ll c$.

The next order correction to the classical formula is given by $\frac{3mu^4}{8c^2}$ which is the second term in Eqn(14) for the relativistic kinetic energy expression.

Let T_N denote the kinetic energy measured using Newtonian formula and let T denote the kinetic energy measured using relativistic formula. Then percentage error δ incurred if we were to use Newtonian formula rather than relativistic formula is given as:

$$\delta = \frac{|T_N - T|}{T} \times 100\% = \frac{\left|\frac{1}{2}mu^2 - m(\gamma(u) - 1)c^2\right|}{m(\gamma(u) - 1)c^2} \times 100\% = \frac{\left|\frac{1}{2}u^2 - (\gamma(u) - 1)c^2\right|}{(\gamma(u) - 1)c^2} \times 100\%$$

Case(i): $u = 0.01c$

$$\delta = \frac{\left| \frac{1}{2}(0.01c)^2 - \left(\frac{1}{\sqrt{1-(0.01)^2}} - 1 \right) c^2 \right|}{\left(\frac{1}{\sqrt{1-(0.01)^2}} - 1 \right) c^2} \times 100\% = \frac{\left| \frac{1}{2}(0.01)^2 - \left(\frac{1}{\sqrt{1-(0.01)^2}} - 1 \right) \right|}{\left(\frac{1}{\sqrt{1-(0.01)^2}} - 1 \right)} \times 100\% \approx 0.0075\%$$

Case(ii): $u = 0.1c$

$$\delta = \frac{\left| \frac{1}{2}(0.1c)^2 - \left(\frac{1}{\sqrt{1-(0.1)^2}} - 1 \right) c^2 \right|}{\left(\frac{1}{\sqrt{1-(0.1)^2}} - 1 \right) c^2} \times 100\% = \frac{\left| \frac{1}{2}(0.1)^2 - \left(\frac{1}{\sqrt{1-(0.1)^2}} - 1 \right) \right|}{\left(\frac{1}{\sqrt{1-(0.1)^2}} - 1 \right)} \times 100\% \approx 0.75\%$$

Case(iii): $u = 0.5c$

$$\delta = \frac{\left| \frac{1}{2}(0.5c)^2 - \left(\frac{1}{\sqrt{1-(0.5)^2}} - 1 \right) c^2 \right|}{\left(\frac{1}{\sqrt{1-(0.5)^2}} - 1 \right) c^2} \times 100\% = \frac{\left| \frac{1}{2}(0.5)^2 - \left(\frac{1}{\sqrt{1-(0.5)^2}} - 1 \right) \right|}{\left(\frac{1}{\sqrt{1-(0.5)^2}} - 1 \right)} \times 100\% \approx 19.20\%$$