

$$\frac{\partial \mathbf{u}_{\mathsf{C}}}{\partial t} + (\mathbf{u}_{\mathsf{C}} \cdot \nabla)\mathbf{u}_{\mathsf{C}} = \frac{q_{\mathsf{e}}}{m_{\mathsf{e}}} (\mathbf{E} + \mathbf{u}_{\mathsf{C}} \times \mathbf{B})$$

$$\frac{\partial \mathbf{B}}{\partial t} = -\nabla \times \mathbf{E}$$

$$\frac{1}{c^2} \frac{\partial \mathbf{E}}{\partial t} = \nabla \times \mathbf{B} - \mu_0 (\mathbf{j}_{\mathsf{C}} + \mathbf{j}_{\mathsf{h}})$$

$$\frac{\partial f_h}{\partial t} + \mathbf{v} \cdot \nabla f_h + \frac{q_e}{m_e} (\mathbf{E} + \mathbf{v} \times \mathbf{B}) \cdot \nabla_{\mathbf{v}} f_h = 0$$

•
$$\nabla \cdot \mathbf{E} = \frac{1}{\epsilon_0} [q_i n_i + q_e (n_c + n_h)]$$
 at $t = 0$

Linearization about an equilibrium state:

$$n_{\mathsf{C}}(\mathbf{x},t) = n_{\mathsf{C},0}(\mathbf{x}) + \tilde{n}_{\mathsf{C}}(\mathbf{x},t)$$

$$\mathbf{u}_{\mathsf{C}}(\mathbf{x},t) = \tilde{\mathbf{u}}_{\mathsf{C}}(\mathbf{x},t)$$

$$\mathbf{B}(\mathbf{x},t) = \mathbf{B}_0(\mathbf{x}) + \tilde{\mathbf{B}}(\mathbf{x},t)$$

- $\mathbf{E}(\mathbf{x},t) = \tilde{\mathbf{E}}(\mathbf{x},t)$
 - \rightarrow cold current density $\tilde{\mathbf{j}}_{C} = q_{e} (n_{c,0} + \tilde{n}_{c}) \tilde{\mathbf{u}}_{C} \approx q_{e} n_{c,0} \tilde{\mathbf{u}}_{C}$

•
$$\frac{\partial \tilde{\mathbf{j}}_{c}}{\partial t} = \epsilon_{0} \Omega_{pe}^{2} \tilde{\mathbf{E}} + \tilde{\mathbf{j}}_{c} \times \Omega_{ce}$$

• $\frac{\partial \tilde{\mathbf{B}}}{\partial t} = -\nabla \times \tilde{\mathbf{E}}$

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$$\frac{1}{c^2} \frac{\partial \tilde{\mathbf{E}}}{\partial t} = \nabla \times \tilde{\mathbf{B}} - \mu_0 (\tilde{\mathbf{j}}_{\mathsf{C}} + \mathbf{j}_{\mathsf{h}})$$

$$\frac{\partial f_h}{\partial t} + \mathbf{V} \cdot \nabla f_h + \frac{q_e}{m_e} (\mathbf{E} + \mathbf{V} \times \mathbf{B}) \cdot \nabla_{\mathbf{V}} f_h = 0$$

$$\Omega_{\mathrm{pe}}^2 = \frac{n_{\mathrm{c},0}e^2}{\epsilon_0 m_{\mathrm{e}}}$$

$$\Omega_{ce} = \frac{q \mathbf{B}_0}{m}$$

→ cold plasma model + energetic kinetic electrons

The dynamical equations conserve exactly the following energy:

$$\boldsymbol{\epsilon} := \frac{1}{2} \left[\boldsymbol{\epsilon}_0 \int \tilde{\mathbf{E}}^2 d^3 \mathbf{x} + \frac{1}{\mu_0} \int \tilde{\mathbf{B}}^2 d^3 \mathbf{x} + \frac{1}{\epsilon_0 \Omega_{pe}^2} \int \tilde{\mathbf{j}}_c^2 d^3 \mathbf{x} + m_e \int f_h v^2 d^3 \mathbf{v} d^3 \mathbf{x} \right]$$
$$= \boldsymbol{\epsilon}_{em} + \boldsymbol{\epsilon}_c + \boldsymbol{\epsilon}_h$$

Dispersion relation for ...

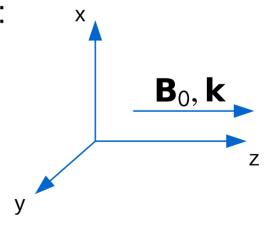
- ... a homogeneous background plasma:
- ... a uniform magnetic background field:
- ... parallel wave propagation:
- ... a uniform equilibrium hot electron distribution:

$$f_h(\mathbf{x}, \mathbf{v}, t) = f_h^0(\mathbf{v}) + \tilde{f}_h(\mathbf{x}, \mathbf{v}, t)$$

$$n_{\rm C,0} = const.$$

$$\mathbf{B}_0 = B_0 \mathbf{e}_z$$

$$\mathbf{k} \parallel \mathbf{B}_0 \parallel \mathbf{e}_z$$



Solution of the fully linearized problem by plane wave ansatz:

$$\nabla \rightarrow ik\mathbf{e}_z$$

$$\rightarrow \frac{\partial}{\partial t} \rightarrow -i\omega$$

- 3 types of solutions of which
 - > 2 correspond to electromagnetic waves with $\tilde{\mathbf{E}}, \tilde{\mathbf{B}} \perp \mathbf{k}$
 - > 1 corresponds to electrostatic waves with $\tilde{\mathbf{E}} \parallel \mathbf{k}, \quad \tilde{\mathbf{B}} = 0$

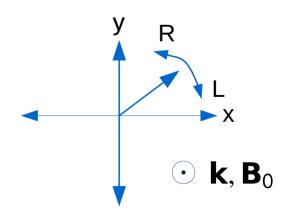


• Dispersion relation for electromagnetic waves: $\tilde{\mathbf{E}}, \tilde{\mathbf{B}} \perp \mathbf{k}$

$$D_{\text{R/L}}(k,\omega) = 1 - \frac{c^2 k^2}{\omega^2} - \frac{\Omega_{\text{pe}}^2}{\omega(\omega \pm \Omega_{\text{ce}})} + v_{\text{h}} \frac{\Omega_{\text{pe}}^2}{\omega} \int d^3 \mathbf{v} \frac{v_{\perp}}{2} \frac{\hat{G} f_{\text{h}}^0(v_{\parallel}, v_{\perp})}{\omega \pm \Omega_{\text{ce}} - k v_{\parallel}} = 0$$

$$\hat{G} = \frac{\partial}{\partial v_{\perp}} + \frac{k}{\omega} \left(v_{\perp} \frac{\partial}{\partial v_{\parallel}} - v_{\parallel} \frac{\partial}{\partial v_{\perp}} \right), \qquad v_{h} = n_{h,0}/n_{c,0}, \quad d^{3}\mathbf{v} = dv_{\parallel} dv_{\perp} v_{\perp} 2\pi$$

- Wave-particle interaction due to electrons with $kv_{\parallel} = \omega \pm \Omega_{ce}$
- Solutions correspond to
 - Right-hand circularly polarized waves (R)
 - Left-hand circularly polarized waves (L)



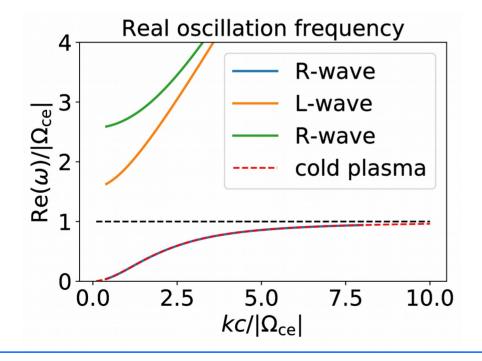


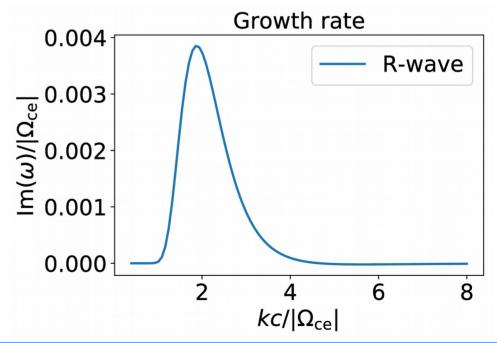
Example: anisotropic Maxwellian

$$f_{\mathsf{h}}^{0}(v_{\parallel}, v_{\perp}) = \frac{n_{\mathsf{h},0}}{(2\pi)^{3/2} w_{\parallel} w_{\perp}^{2}} \exp\left(-\frac{v_{\parallel}^{2}}{2w_{\parallel}^{2}} - \frac{v_{\perp}^{2}}{2w_{\perp}^{2}}\right)$$

$$w_{\parallel} = 0.2 c$$

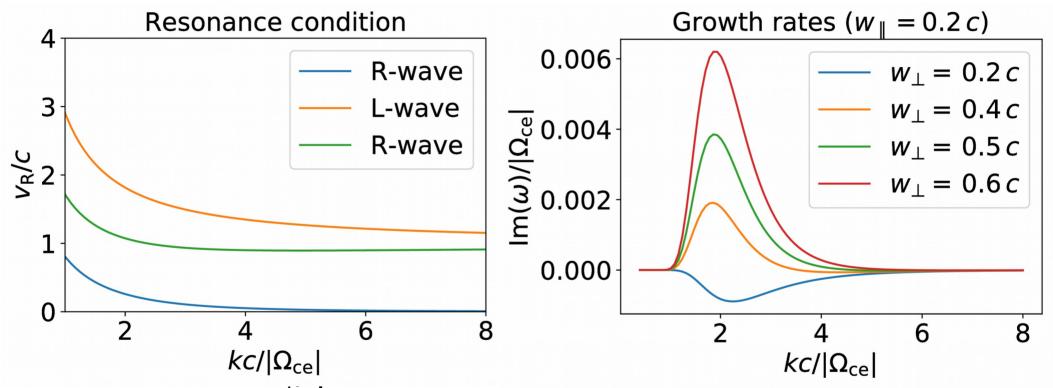
 $w_{\perp} = 0.5 c$
 $v_{h} = 5 \cdot 10^{-3}$
 $\Omega_{pe} = 2|\Omega_{ce}|$







Example: anisotropic Maxwellian



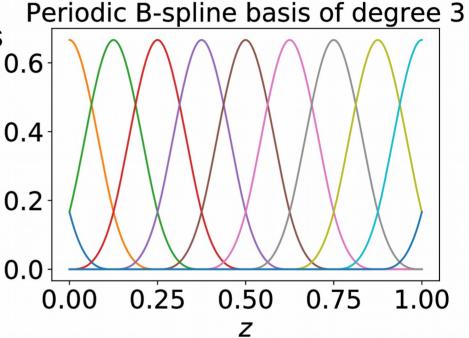
 \rightarrow no wave-particle interaction for R/L-waves with $\omega \gg |\Omega_{ce}|$



Implementation of the model in Python:

1d B-spline Finite Elements for fields

 $\mathbf{U} = (E_x, E_y, B_x, B_y, j_{c,x}, j_{c,y})$

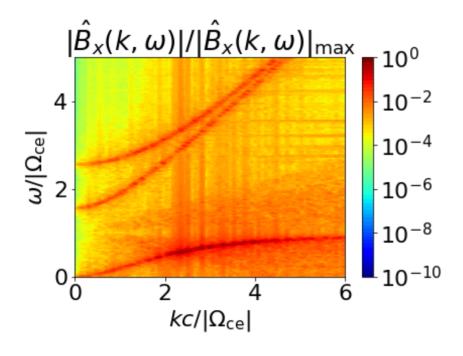


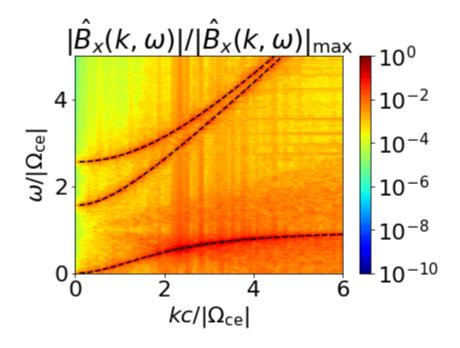
- 1d3v Particle-In-Cell with control variate and Boris particle pusher
- Implicit Crank-Nicolson time stepping scheme



Simulation Results

Test run with low density, isotropic Maxwellian to create thermal noise → no growth expected

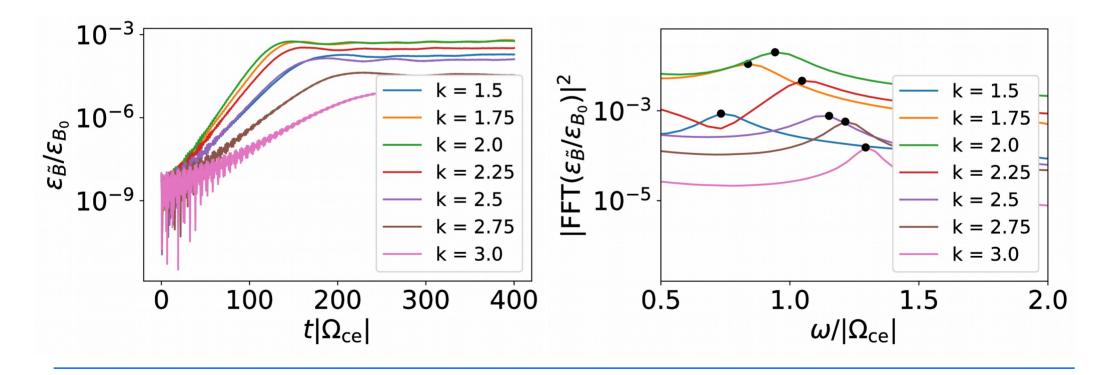






Simulation Results:

Runs with perturbed magnetic field of the form $B_x^0(z) = a \cdot \sin(kz)$ for different wavenumbers





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Runs with perturbed magnetic field of the form $B_x^0(z) = a \cdot \sin(kz)$ for different wavenumbers

