

RIEMANNIAN GEOMETRY

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0. PREFACE

0.1. **To readers.** This note is divided into several parts:

1. In the **First** part, we firstly introduce Levi-Civita connection on a (real) vector bundle E equipped with a metric, and in Riemannian geometry we mostly concern about tangent bundle. Holding a connection on E , one can construct connection on its dual bundle E^* , tensor product $E \otimes E^*$ and so on. When E is chosen to be tangent bundle, a section of tensor products is sometimes called a tensor, and tensor computation is a powerful tool of Riemannian geometry so we collect some basic properties and operations about tensor together in section 2.

However, tensor computation may be quite tough in general. To give a neat local computation for tensor, we introduce geodesic in section 3 in order to introduce normal coordinate. By the way we introduce some other properties about geodesic such as global existence of geodesic and geodesics on Lie group.

In section 4, we introduce curvature using two different views: curvature form and curvature tensor and prove Bianchi identities in these two views. We also introduce some other important curvatures such as sectional curvature, Ricci curvature and scalar curvature.

2. The **Second** part is about Bochner's technique, which is one of the most important technique in modern Riemannian geometry. Holding this technical, we can see how does bounded Ricci curvature as an obstruction to the existence of Killing fields and harmonic 1-forms. Aside these, we also introduce Hodge theory, which allows us to use harmonic 1-forms to represent elements in the first homology group, then Bochner's technique gives a kind of vanishing theorem.
3. In the **Third** part, we firstly solve the following question: "Given two points p, q , what's the length-minimizing curve connecting p, q ?". To answer this question, we need to consider the arc-length functional, and
 - (a) First variation formula implies geodesics are critical points of arc-length functional;
 - (b) Second variation formula implies if a geodesic contains no interior conjugate points, then it's locally minimum of arc-length functional.
 Along the way we develop the tools of index form and Jacobi fields, which are also quite important in the third part and fourth part.
- 4.
- 5.

0.2. Some notations and conventions.

0.2.1. Conventions.

1. We always use Einstein summation.
2. When we say M is a smooth manifold, we assume it's a real smooth manifold.

0.2.2. Notations about smooth manifolds.

1. For a smooth function $f : \mathbb{R}^n \rightarrow \mathbb{R}$, we use $\frac{\partial f}{\partial x^i}$ to denote its partial derivative with respect to x^i , where x^i are coordinates of \mathbb{R}^n .
2. For a smooth manifold M , we use TM, T^*M to denote its tangent space and cotangent space respectively, and we also use Ω_M^k to denote the bundle of k -forms, that is $\bigwedge^k T^*M$.
3. We always use X, Y, Z to denote vector fields, ω to denote 1-forms and φ, ψ to denote k -forms.
4. For a smooth map $f : N \rightarrow M$ between smooth manifolds, we use df or f_* to denote its differential.
5. Given a vector bundle $E \rightarrow M$ over a smooth manifold M , we use $C^\infty(M, E)$ to denote the set of all smooth sections of E .

0.2.3. Notations about Riemannian manifolds.

1. We use (M, g) to denote a Riemannian manifold, where M is a smooth manifold, and g is its Riemannian metric. If there is no ambiguity, we will omit g .
2. For a Riemannian metric g , we sometimes use $\langle -, - \rangle_g$ to denote it, or directly $\langle -, - \rangle$ if there is no ambiguity.

Part 1. Basic settings

1. CONNECTIONS

Connection is a very basic conception in realm of geometry of vector bundles, and there are too many definitions of it which seem to be different. This part is divided into four parts:

1. In the first section, we will introduce one approach to connection in two different ways, the first one is often used in complex geometry and the second is given by Do carmo in [Car92];
2. In the second section, we will give another characterization of connection using parallel transport, and we will see all these approaches are same in fact.
3. In the third section, we will put more restrictions on our connection, such as compatibility with metric and torsion-free;
4. In the fourth section, we will construct many new connections from a given connection, which play an important role in our later discuss.

1.1. Two different viewpoints to connection.

1.1.1. *First definition.* When I first learn Riemannian geometry or complex geometry, I'm quite confused about why we need connection, and why we define it like this? In fact, given a vector bundle $\pi : E \rightarrow M$, connections on E are arised to take “derivative” of a section $s : M \rightarrow E$ in a given direction.

It's quite natural to ask such a question, since when we learn calculus, we already know how to take derivative of a smooth function $f : M \rightarrow \mathbb{R}^m$ to obtain a 1-form, that is a section of T^*M . In another point of view, any smooth function $f : M \rightarrow \mathbb{R}^m$ can be regarded as a section of trivial vector bundle $M \times \mathbb{R}^m$, as follows

$$x \mapsto (x, f(x))$$

and we can also regard its derivative df as a section of $T^*M \otimes (M \times \mathbb{R}^m)$. So taking derivative can be seen as the following operator:

$$\nabla : C^\infty(M, M \times \mathbb{R}^m) \rightarrow C^\infty(M, T^*M \otimes (M \times \mathbb{R}^m))$$

In general, we can define a connection as follows:

Definition 1.1.1 (connection). A connection ∇ on a vector bundle E on a smooth manifold M is a linear operator

$$\nabla : C^\infty(M, E) \rightarrow C^\infty(M, T^*M \otimes E)$$

satisfying Leibniz rule $\nabla(fs) = df \otimes s + f\nabla s$, where $s \in C^\infty(M, E)$.

Remark 1.1.1 (local form). We can locally write a section s of E as $s^\alpha e_\alpha$, then Leibniz rule implies

$$\nabla(s^\alpha e_\alpha) = ds^\alpha e_\alpha + s^\alpha \nabla e_\alpha$$

If we write ∇e_α explicitly as follows

$$\nabla e_\alpha = \omega_\alpha^\beta e_\beta$$

where ω_α^β are 1-forms. So connection locally looks like $d + \omega$, where ω is a 1-form valued matrix.

Now let's see how does ω change with change of local basis. Suppose there is another local basis \tilde{e}_α , which is related by $\tilde{e}_\alpha = g_\alpha^\beta e_\beta$, then

$$\begin{aligned} \nabla \tilde{e}_\alpha &= \nabla (g_\alpha^\beta e_\beta) \\ &= g_\alpha^\beta \nabla e_\beta + dg_\alpha^\beta e_\beta \\ &= g_\alpha^\beta \omega_\beta^\gamma e_\gamma + dg_\alpha^\beta e_\beta \end{aligned}$$

So if we write in matrix notation, we have

$$\begin{aligned} \nabla \tilde{e} &= g\omega e + dg e \\ &= (g\omega g^{-1} + dg g^{-1}) \tilde{e} \end{aligned}$$

which implies $\tilde{\omega} = g\omega g^{-1} + dg g^{-1}$.

1.1.2. *Second definition.* The following is the definition given by Do carmo in [Car92].

Definition 1.1.2 (connection). A connection ∇ on a vector bundle E on a smooth manifold M is a mapping

$$\begin{aligned} \nabla : C^\infty(M, TM) \times C^\infty(M, E) &\rightarrow C^\infty(M, E) \\ (X, s) &\mapsto \nabla_X s \end{aligned}$$

satisfying the following properties:

1. $\nabla_{fX+gY}s = f\nabla_X s + g\nabla_Y s$
2. $\nabla_X(s + s') = \nabla_X s + \nabla_X s'$
3. $\nabla_X(fs) = f\nabla_X s + X(f)s$

where $X, Y \in C^\infty(M, TM)$, $f, g \in C^\infty(M)$ and $s, s' \in C^\infty(M, E)$.

Remark 1.1.2 (local form). For a given point $p \in M$ and choose a local basis $\{\frac{\partial}{\partial x^i}\}$ of TM and a local basis $\{e_\alpha\}$ of E , then we can write a vector field X and a section s of E as

$$X = X^i \frac{\partial}{\partial x^i}, \quad e = s^\alpha e_\alpha$$

Then

$$\begin{aligned} \nabla_X s &= \nabla_{X^i \frac{\partial}{\partial x^i}} s^\alpha e_\alpha \\ &= X^i \nabla_{\frac{\partial}{\partial x^i}} s^\alpha e_\alpha \\ &= X^i s^\alpha \nabla_{\frac{\partial}{\partial x^i}} e_\alpha + X^i \frac{\partial s^\alpha}{\partial x^i} e_\alpha \\ &= X^i s^\alpha \omega_{\frac{\partial}{\partial x^i}}^\beta e_\beta + X^i \frac{\partial s^\alpha}{\partial x^i} e_\alpha \end{aligned}$$

If we write $\nabla_{\frac{\partial}{\partial x^i}} s_\alpha = \Gamma_{i\alpha}^\beta e_\beta$, we can write

$$\nabla_X s = (X^i s^\alpha \Gamma_{i\alpha}^\beta + X(s^\beta)) e_\beta$$

So as we can see, $\Gamma_{i\alpha}^\beta$, which is sometimes called Christoffel symbol, completely determines our connection ∇ .

Remark 1.1.3 (The equivalence between two definitions). Locally a connection in definition 1.1.1 is a 1-form valued matrix ω , and write it as $\omega_\alpha^\beta = \Gamma_{j\alpha}^\beta dx^j$. Then

$$\begin{aligned} \nabla e_\alpha &= \omega_\alpha^\beta e_\beta \\ &= \Gamma_{i\alpha}^\beta dx^i e_\beta \end{aligned}$$

So if want to define $\nabla_{\frac{\partial}{\partial x^i}} e_\alpha$, ∇e_α need to “eat” a vector field, and luckily dx^j can eat one, so we can define it as follows

$$\begin{aligned} \nabla_{\frac{\partial}{\partial x^i}} e_\alpha &:= \Gamma_{j\alpha}^\beta dx^j \left(\frac{\partial}{\partial x^i} \right) e_\beta \\ &= \Gamma_{i\alpha}^\beta e_\beta \end{aligned}$$

From this we can see these two definitions are same.

Remark 1.1.4 (connection and covariant derivative). Some authors may also use terminology “covariant derivative”, here we make a clarify: Here we give two definitions of connection ∇ on a vector bundle E . Given a section s of E and a vector field X , we call $\nabla_X s$ the covariant derivative of s with respect to X . In fact, you can see connection and covariant derivative the same thing, just different terminology.

1.2. Parallel transport. In this section we fix a vector bundle E over M with connection ∇ , $\gamma : I \rightarrow M$ is a smooth curve. With this setting, we can define what is parallel transport along a smooth curve $\gamma(t)$.

Firstly, we can define a connection on pullback bundle γ^*E over γ as follows

$$\widehat{\nabla}_{\frac{d}{dt}} \gamma^* s := \nabla_{\gamma_* \left(\frac{d}{dt} \right)} s$$

where $s \in C^\infty(M, E)$.

Remark 1.2.1 (local form). Locally we have

$$\begin{aligned} \widehat{\nabla}_{\frac{d}{dt}} \gamma^* s &= \nabla_{\frac{d\gamma^i}{dt} \frac{\partial}{\partial x^i}} s^\alpha e_\alpha \\ &= \frac{d\gamma^i}{dt} \nabla_{\frac{\partial}{\partial x^i}} s^\alpha e_\alpha \\ &= \frac{d\gamma^i}{dt} \left(\frac{\partial s^\alpha}{\partial x^i} e_\alpha + s^\alpha \Gamma_{i\alpha}^\beta e_\beta \right) \end{aligned}$$

Definition 1.2.1 (parallel). A section s of γ^*E is called parallel along γ , if $\widehat{\nabla}_{\frac{d}{dt}} s = 0$.

From local form we can see $\widehat{\nabla}_{\frac{d}{dt}} s = 0$ is a system of ODEs locally, which can always be solved uniquely in a sufficiently short interval if we given a initial value, that's how we define parallel transport.

Definition 1.2.2 (parallel transport). For $t_0, t \in I$, parallel transport $P_{t_0, t}^\gamma$ is an isomorphism between vector spaces¹ defined by

$$\begin{aligned} P_{t_0, t}^\gamma : E_{\gamma(t_0)} &\rightarrow E_{\gamma(t)} \\ s_0 &\mapsto s(t) \end{aligned}$$

where s is the unique parallel section along γ satisfying $s(t_0) = s_0$.

Remark 1.2.2 (parallel frame). A useful tool is parallel frame: Fix a basis $\{e_\alpha\}$ of $E_{\gamma(t_0)}$, we can use parallel transport to give a family of basis $\{e_\alpha(t)\}$ of $E_{\gamma(t)}$ along γ such that $e_\alpha(0) = e_\alpha$.

Proposition 1.2.1. For any section s of E along γ and $t_0, t \in I$, we have

$$\widehat{\nabla}_{\frac{d}{dt}} P_{t, t_0}^\gamma s(t) = P_{t, t_0}^\gamma \widehat{\nabla}_{\frac{d}{dt}} s(t)$$

Proof. Assume $\{e_\alpha(t)\}$ is a parallel frame along γ . With respect to this parallel frame we can write $s(t)$ as

$$s(t) = s^\alpha(t) e_\alpha(t)$$

Thus

$$\begin{aligned} \widehat{\nabla}_{\frac{d}{dt}} P_{t, t_0}^\gamma s(t) &= \widehat{\nabla}_{\frac{d}{dt}} (s^\alpha(t) e_\alpha(t_0)) \\ &= \frac{ds^\alpha}{dt}(t) e_\alpha(t_0) \\ P_{t, t_0}^\gamma \widehat{\nabla}_{\frac{d}{dt}} s(t) &= P_{t, t_0}^\gamma \left(\frac{ds^\alpha}{dt}(t) e_\alpha(t) \right) \\ &= \frac{ds^\alpha}{dt}(t) e_\alpha(t_0) \end{aligned}$$

□

Remark 1.2.3. In fact, connection and parallel transport are the same things in different viewpoint.

1.3. Compatibility and torsion-free.

1.3.1. *Compatibility with metric.* Now consider a vector bundle E with a metric g , which can be locally written as $g_{\alpha\beta} e^\alpha \otimes e^\beta$. So if there is a connection ∇ on E , so it's natural to ask it to be compatible with our metric.

Definition 1.3.1 (compatibility). A connection ∇ on vector bundle E is compatible with metric g , if for any two section s, t of E , we have

$$dg(s, t) = g(\nabla s, t) + g(s, \nabla t)$$

¹Its inverse is P_{t, t_0}^γ .

Remark 1.3.1 (local form). Locally we can compute it as

$$\begin{aligned} dg_{\alpha\beta} &= dg(e_\alpha, e_\beta) \\ &= g(\nabla e_\alpha, e_\beta) + g(e_\alpha, \nabla e_\beta) \\ &= \omega_\alpha^\gamma g_{\gamma\beta} + g_{\alpha\gamma} \omega_\beta^\gamma \end{aligned}$$

So in matrix notation we have²

$$dg = \omega g + g \omega^t$$

In particular we have

$$\frac{\partial}{\partial x^i} g_{\alpha\beta} = \Gamma_{i\alpha}^\gamma g_{\gamma\beta} + \Gamma_{i\beta}^\gamma g_{\alpha\gamma}$$

for all i, α, β .

Proposition 1.3.1. A connection ∇ is compatible with metric if and only if for arbitrary curve $\gamma : I \rightarrow M$ and two parallel sections s_1, s_2 along γ we have $g(s_1, s_2)$ is constant.

Proof. It's clear if ∇ is compatible with metric g , then and two sections s, t are parallel along γ , we have

$$dg(s_1, s_2) = g(\nabla s_1, s_2) + g(s_1, \nabla s_2) = 0$$

which implies $g(s, t)$ is constant.

Conversely, let $\{e_\alpha(t)\}$ be a parallel orthonormal frame with respect to g along γ and write

$$s_1(t) = s_1^\alpha(t) e_\alpha, \quad s_2(t) = s_2^\alpha(t) e_\alpha(t)$$

Then we have

$$\begin{aligned} g(\nabla s_1, s_2) + g(s_1, \nabla s_2) &= \sum_\alpha \frac{ds_1^\alpha}{dt} s_2^\alpha + s_1^\alpha \frac{ds_2^\alpha}{dt} \\ &= \frac{d}{dt} \left(\sum_\alpha s_1^\alpha s_2^\alpha \right) \\ &= \frac{d}{dt} g(s_1, s_2) \end{aligned}$$

□

1.3.2. *Torsion-free.* Now let's choose our vector bundle E to be tangent bundle of a Riemannian manifold (M, g) .

Definition 1.3.2 (torsion-free). A connection ∇ of TM is torsion-free if

$$\nabla_X Y - \nabla_Y X = [X, Y]$$

where X, Y are vector fields.

²Here we need to pay more attention, although as a number $g_{\alpha\gamma} \omega_\beta^\gamma = \omega_\beta^\gamma g_{\alpha\gamma}$, we can not write this matrix notation as $dg = \omega g + \omega g^t$, since $\omega_\beta^\gamma g_{\gamma\alpha}$ is (β, α) -entry of ωg^t , but $dg_{\alpha\beta}$ and $g_{\alpha\gamma} \omega_\beta^\gamma$ are (α, β) -entries of $g \omega^t$.

Remark 1.3.2 (local form). If we choose $X = \frac{\partial}{\partial x^i}, Y = \frac{\partial}{\partial x^j}$, then we have

$$\begin{aligned}\nabla_{\frac{\partial}{\partial x^i}} \frac{\partial}{\partial x^j} - \nabla_{\frac{\partial}{\partial x^j}} \frac{\partial}{\partial x^i} &= (\Gamma_{ij}^k - \Gamma_{ji}^k) \frac{\partial}{\partial x^k} \\ &= 0\end{aligned}$$

which is equivalent to say Γ_{ij}^k is symmetric in i and j .

1.4. Levi-Civita connection. An interesting thing is that there is only one connection of TM on a Riemannian manifold (M, g) which is both compatible with Riemannian metric and torsion-free. It suffices to see such connection is completely determined, in other words, Γ_{ij}^k is completely determined by compatibility and torsion-free.

Note that compatibility implies

$$\begin{aligned}Xg(Y, Z) &= g(\nabla_X Y, Z) + g(Y, \nabla_X Z) \\ Yg(Z, X) &= g(\nabla_Y Z, X) + g(Z, \nabla_Y X) \\ Zg(X, Y) &= g(\nabla_Z X, Y) + g(X, \nabla_Z Y)\end{aligned}$$

Adding first two equations, subtract the third and use torsion-free condition, we will see

$$Xg(Y, Z) + Yg(Z, X) - Zg(X, Y) = g([X, Z], Y) + g([Y, Z], X) + g([X, Y], Z) + 2g(Z, \nabla_Y X)$$

thus

$$g(Z, \nabla_Y X) = \frac{1}{2}(Xg(Y, Z) + Yg(Z, X) - Zg(X, Y) - g([X, Z], Y) - g([Y, Z], X) - g([X, Y], Z))$$

which implies $\nabla_X Y$ is uniquely determined.

Remark 1.4.1 (local form). Firstly, compatibility implies

$$\frac{\partial g_{ij}}{\partial x^k} = \Gamma_{ki}^l g_{lj} + \Gamma_{kj}^l g_{il}$$

By permuting i, j, k we obtain the following two equations

$$\begin{aligned}\frac{\partial g_{jk}}{\partial x^i} &= \Gamma_{ij}^l g_{lk} + \Gamma_{ik}^l g_{jl} \\ \frac{\partial g_{ki}}{\partial x^j} &= \Gamma_{jk}^l g_{li} + \Gamma_{ji}^l g_{kl}\end{aligned}$$

By the symmetry of Γ_{ij}^l in i, j and symmetry of g_{ij} , we have

$$2\Gamma_{ij}^l g_{lk} = \frac{\partial g_{kj}}{\partial x^i} + \frac{\partial g_{ik}}{\partial x^j} - \frac{\partial g_{ij}}{\partial x^k}$$

If we use (g^{ij}) to denote the inverse matrix of (g_{ij}) , then we have

$$\Gamma_{ij}^l = \frac{1}{2}g^{kl}\left(\frac{\partial g_{kj}}{\partial x^i} + \frac{\partial g_{ik}}{\partial x^j} - \frac{\partial g_{ij}}{\partial x^k}\right)$$

which implies Christoffel symbol is completely determined by Riemannian metric and its partial derivatives.

1.5. Induced connection. Given a vector bundle E , you can construct many new vector bundles by algebraic method, such as considering its dual bundle E^* , tensor product $E \otimes E$ and so on. Now let's see if we already have a connection ∇ defined on E , let's construct some new connections on new vector bundles.

1.5.1. Induced connection on dual bundle. Firstly let's consider how to induce a connection on dual bundle E^* . If s is a section of E , and use s^* to denote its dual section, its natural to ask

$$d(s, s^*) = (\nabla s, s^*) + (s, \nabla s^*)$$

Here we still use ∇ to denote the induced connection on E^* . So if $\{e_\alpha\}$ is a local basis of E and $\{e^\alpha\}$ is the dual basis of E^* , then

$$\begin{aligned} 0 &= (\omega_\alpha^\gamma e_\gamma, e^\beta) + (e_\alpha, (\omega^*)^\beta_\gamma e^\gamma) \\ &= \omega_\alpha^\beta + (\omega^*)^\beta_\alpha \end{aligned}$$

which implies induced connection on E^* locally looks like $(-\omega_\alpha^\beta)$.³

Remark 1.5.1 (Another characterization for torsion-free). If we consider connection ∇ defined on TM , locally given by Christoffel symbol Γ_{ij}^k , then induced connection on T^*M locally looks like

$$\nabla dx^k = -\Gamma_{ij}^k dx^i \otimes dx^j$$

that is $\nabla dx^k \in C^\infty(M, T^*M \otimes T^*M)$.

Given a section s of T^*M , we can obtain a 2-form $ds \in C^\infty(M, \bigwedge^2 T^*M)$. Note that $\bigwedge^2 T^*M$ is just the skew-symmetrization of $T^*M \otimes T^*M$, so it's natural to require the skew-symmetrization of ∇s is ds .

If we write this down in a local basis $\{dx^i\}$ of T^*M , we have

$$\nabla dx^k = -\Gamma_{ij}^k dx^i \otimes dx^j$$

But $d^2x^k = 0$, so condition for torsion-free is equivalent to skew-symmetrization of $\nabla dx^k = 0$, that is $-\Gamma_{ij}^k dx^i \wedge dx^j = 0$, which is equivalent to say Γ_{ij}^k is symmetric in i, j .

1.5.2. Induced connections on tensor product. For any two vector bundles E, F over M , we use ∇ to denote connections on them in order to save symbols. We can define a connection ∇ on $E \otimes F$ as follows: Take s, f as sections of E and F , then

$$\nabla(s \otimes f) = \nabla s \otimes f + s \otimes \nabla f \in C^\infty(M, T^*M \otimes (E \otimes F))$$

³However, there is one thing to be care about, the upper index is row index and lower index is column index, not the same as ω_α^β . Or in other words, if a connection on E locally looks like ω , then connection induced on E^* locally looks like $-\omega^t$.

In particular, there is an induced connection ∇ on $\text{End } E$, since we have $\text{End } E \cong E \otimes E^*$. In this case, we can write it more explicitly as follows: Locally we have a basis $\{e_\alpha\}$ of E and a basis $\{e^\beta\}$ of E^* . Thus

$$\nabla(e_\alpha \otimes e^\beta) = \omega_\alpha^\gamma e_\gamma \otimes e^\beta + e_\alpha \otimes (-\omega_\gamma^\beta e^\gamma)$$

So in general a section of $E \otimes E^*$ locally takes form $s = s_\beta^\alpha e_\alpha \otimes e^\beta$, then

$$\begin{aligned} \nabla(s_\beta^\alpha e_\alpha \otimes e^\beta) &= ds_\beta^\alpha e_\alpha \otimes e^\beta + s_\beta^\alpha (\nabla e_\alpha \otimes e^\beta + e_\alpha \otimes \nabla e^\beta) \\ &= ds_\beta^\alpha e_\alpha \otimes e^\beta + s_\beta^\alpha \omega_\alpha^\gamma e_\gamma \otimes e^\beta - s_\beta^\alpha \omega_\gamma^\beta e_\alpha \otimes e^\gamma \\ &= (ds_\beta^\alpha + s_\beta^\alpha \omega_\alpha^\gamma - \omega_\gamma^\beta s_\beta^\alpha) e_\alpha \otimes e^\beta \end{aligned}$$

Thus in matrix notation we have

$$\nabla s = ds + s\omega - \omega s$$

However, there is another way to induce a connection on $E \otimes E^*$ as follows: For any section s of $E \otimes E^*$, we have a function $s(e^\alpha, e_\beta)$, so it's natural to ask

$$ds(e^\alpha, e_\beta) = \nabla s(e^\alpha, e_\beta) + s(\nabla e^\alpha, e_\beta) + s(e^\alpha, \nabla e_\beta)$$

Locally if we write $s = s_\beta^\alpha e_\alpha \otimes e^\beta$, then

$$\begin{aligned} d(s_\beta^\alpha) &= (\nabla s)_\beta^\alpha + s(-\omega_\gamma^\alpha e^\gamma, e_\beta) + s(e^\alpha, \omega_\beta^\gamma e_\gamma) \\ &= (\nabla s)_\beta^\alpha - s_\beta^\gamma \omega_\gamma^\alpha + \omega_\beta^\gamma s_\gamma^\alpha \end{aligned}$$

which implies these two ways to induce are same!

2. TENSOR

2.1. Induced connections on tensor.

Definition 2.1.1 (tensor). A section of $\bigotimes^s TM \otimes \bigotimes^r T^*M$ is called a (s, r) -tensor.

Example 2.1.1. A smooth function f is a $(0, 0)$ -tensor.

Example 2.1.2. A vector field X is a $(1, 0)$ -tensor.

Example 2.1.3. A 1-form ω is a $(0, 1)$ -tensor.

Example 2.1.4. The Riemannian metric g is a $(0, 2)$ -tensor.

Definition 2.1.2 (connection on tensor). For a (s, r) -tensor T , ∇T is a $(s, r+1)$ -tensor, which is defined by

$$\begin{aligned} \nabla T(\mathrm{d}x^{j_1}, \dots, \mathrm{d}x^{j_s}, \frac{\partial}{\partial x^i}, \frac{\partial}{\partial x^{i_1}}, \dots, \frac{\partial}{\partial x^{i_r}}) &:= \frac{\partial}{\partial x^i} T(\mathrm{d}x^{j_1}, \dots, \mathrm{d}x^{j_s}, \frac{\partial}{\partial x^{i_1}}, \dots, \frac{\partial}{\partial x^{i_r}}) \\ &\quad - \sum_{l=1}^s T(\mathrm{d}x^{j_1}, \dots, \nabla_{\frac{\partial}{\partial x^i}} \mathrm{d}x^{j_l}, \dots, \mathrm{d}x^{j_s}, \frac{\partial}{\partial x^{i_1}}, \dots, \frac{\partial}{\partial x^{i_r}}) \\ &\quad - \sum_{m=1}^r T(\mathrm{d}x^{j_1}, \dots, \mathrm{d}x^{j_s}, \frac{\partial}{\partial x^{i_1}}, \dots, \nabla_{\frac{\partial}{\partial x^i}} \frac{\partial}{\partial x^{i_m}}, \dots, \frac{\partial}{\partial x^{i_r}}) \end{aligned}$$

Definition 2.1.3 (covariant derivative of tensor). For a (s, r) -tensor T , the covariant derivative of T with respect to vector field X , which is a (s, r) -tensor, is defined as

$$\nabla_X T := \nabla T(\mathrm{d}x^{j_1}, \dots, \mathrm{d}x^{j_s}, X, \frac{\partial}{\partial x^{i_1}}, \dots, \frac{\partial}{\partial x^{i_r}})$$

Remark 2.1.1 (local form). If we write a (s, r) -tensor T locally as

$$T_{i_1 \dots i_r}^{j_1 \dots j_s} \frac{\partial}{\partial x^{j_1}} \otimes \dots \otimes \frac{\partial}{\partial x^{j_s}} \otimes \mathrm{d}x^{i_1} \otimes \dots \otimes \mathrm{d}x^{i_r}$$

and $(s, r+1)$ -tensor ∇T locally as

$$\nabla_i T_{i_1 \dots i_r}^{j_1 \dots j_s} \frac{\partial}{\partial x^{j_1}} \otimes \dots \otimes \frac{\partial}{\partial x^{j_s}} \otimes \mathrm{d}x^i \otimes \mathrm{d}x^{i_1} \otimes \dots \otimes \mathrm{d}x^{i_r}$$

Then by definition we have

$$\nabla_i T_{i_1 \dots i_r}^{j_1 \dots j_s} = \frac{\partial T_{i_1 \dots i_r}^{j_1 \dots j_s}}{\partial x^i} + \sum_{l=1}^s \Gamma_{iq}^{j_l} T_{i_1 \dots i_r}^{j_1 \dots j_{l-1} q j_{l+1} \dots j_s} - \sum_{m=1}^r \Gamma_{i i_m}^q T_{i_1 \dots i_{m-1} q i_{m+1} \dots i_r}^{j_1 \dots j_s}$$

Example 2.1.5. Consider $(0, 0)$ -tensor f , that is a smooth function. Then ∇f is a $(0, 1)$ -tensor, given by

$$\nabla f = \nabla_i f \mathrm{d}x^i$$

by our definition $\nabla_i f = \frac{\partial f}{\partial x^i}$, it coincides with our usual notations.

Inductively, we can define $\nabla^2 T$ to be $\nabla(\nabla T)$, which is a $(s, r+2)$ -tensor, and locally write it as

$$\nabla^2 T = \nabla_k \nabla_i T_{i_1 \dots i_r}^{j_1 \dots j_s} \frac{\partial}{\partial x^{j_1}} \otimes \dots \otimes \frac{\partial}{\partial x^{j_s}} \otimes dx^k \otimes dx^{i_1} \otimes \dots \otimes dx^{i_r}$$

Now there is a natural question: Note that $\nabla_k \nabla_i T$ is a (s, r) -tensor, and if we regard $\nabla_i T$ as a (s, r) -tensor and take covariant derivative of it with respect to $\frac{\partial}{\partial x^k}$, does we obtain the same thing?

Example 2.1.6. For $(0, 0)$ -tensor f , by definition we have $\nabla^2 f$ is $\nabla(\nabla_i f dx^i)$, which is called the Hessian of f , denoted by $\text{Hess } f$. More explicitly

$$\begin{aligned} \text{Hess } f &= \nabla(\nabla_i f dx^i) \\ &= \frac{\partial \nabla_i f}{\partial x^k} dx^k \otimes dx^i - \nabla_i f \Gamma_{kj}^i dx^k \otimes dx^j \\ &= \left(\frac{\partial^2 f}{\partial x^k \partial x^i} - \Gamma_{ki}^j \frac{\partial f}{\partial x^j} \right) dx^k \otimes dx^i \end{aligned}$$

that is $\nabla_k \nabla_i f = \frac{\partial^2 f}{\partial x^k \partial x^i} - \Gamma_{ki}^j \frac{\partial f}{\partial x^j}$.

However, if we regard $\nabla_i f$ as a $(0, 0)$ -tensor, that is a smooth function, and take covariant derivative with respect to $\frac{\partial}{\partial x^k}$, we will obtain $\frac{\partial^2 f}{\partial x^k \partial x^i}$, and it's clear that it doesn't equal $\nabla_k \nabla_i f$, unless Christoffel symbol Γ_{ki}^j vanishes.

Proposition 2.1.1.

$$\begin{aligned} \nabla_k \nabla_i T_{i_1 \dots i_r}^{j_1 \dots j_s} &= \nabla_{\frac{\partial}{\partial x^k}} \nabla_{\frac{\partial}{\partial x^i}} T(dx^{j_1}, \dots, dx^{j_s}, \frac{\partial}{\partial x^{i_1}}, \dots, \frac{\partial}{\partial x^{i_r}}) \\ &\quad - \nabla_{\nabla_{\frac{\partial}{\partial x^k}} \frac{\partial}{\partial x^i}} T(dx^{j_1}, \dots, dx^{j_s}, \frac{\partial}{\partial x^{i_1}}, \dots, \frac{\partial}{\partial x^{i_r}}) \end{aligned}$$

Proof. By definition, we have

$$\begin{aligned}
\nabla_k \nabla_i T_{i_1 \dots i_r}^{j_1 \dots j_s} &= \nabla^2 T(dx^{j_1}, \dots, dx^{j_s}, \underbrace{\frac{\partial}{\partial x^k}, \frac{\partial}{\partial x^i}, \frac{\partial}{\partial x^{i_1}}, \dots, \frac{\partial}{\partial x^{i_r}}}_{\text{part I}}) \\
&= \nabla_{\frac{\partial}{\partial x^k}} \nabla T(dx^{j_1}, \dots, dx^{j_s}, \frac{\partial}{\partial x^i}, \frac{\partial}{\partial x^{i_1}}, \dots, \frac{\partial}{\partial x^{i_r}}) \\
&= \underbrace{\frac{\partial}{\partial x^k} \nabla T(dx^{j_1}, \dots, dx^{j_s}, \frac{\partial}{\partial x^i}, \frac{\partial}{\partial x^{i_1}}, \dots, \frac{\partial}{\partial x^{i_r}})}_{\text{part I}} \\
&\quad - \underbrace{\nabla T(dx^{j_1}, \dots, dx^{j_s}, \nabla_{\frac{\partial}{\partial x^k}} \frac{\partial}{\partial x^i}, \frac{\partial}{\partial x^{i_1}}, \dots, \frac{\partial}{\partial x^{i_r}})}_{\text{part II}} \\
&\quad - \underbrace{\sum_{l=1}^s \nabla T(dx^{j_1}, \dots, \nabla_{\frac{\partial}{\partial x^k}} dx^{j_l}, \dots, dx^{j_s}, \frac{\partial}{\partial x^i}, \frac{\partial}{\partial x^{i_1}}, \dots, \frac{\partial}{\partial x^{i_r}})}_{\text{part III}} \\
&\quad - \underbrace{\sum_{m=1}^r \nabla T(dx^{j_1}, \dots, dx^{j_s}, \frac{\partial}{\partial x^i}, \frac{\partial}{\partial x^{i_1}}, \dots, \nabla_{\frac{\partial}{\partial x^k}} \frac{\partial}{\partial x^{i_m}}, \dots, \frac{\partial}{\partial x^{i_r}})}_{\text{part IV}}
\end{aligned}$$

Note that

1. Part I+III+IV is $\nabla_{\frac{\partial}{\partial x^k}} \nabla_{\frac{\partial}{\partial x^i}} T(dx^{j_1}, \dots, dx^{j_s}, \frac{\partial}{\partial x^{i_1}}, \dots, \frac{\partial}{\partial x^{i_r}})$;
2. Part II is $-\nabla_{\nabla_{\frac{\partial}{\partial x^k}} \frac{\partial}{\partial x^i}} T(dx^{j_1}, \dots, dx^{j_s}, \frac{\partial}{\partial x^{i_1}}, \dots, \frac{\partial}{\partial x^{i_r}})$.

□

Remark 2.1.2 (Another characterization of compatibility). Note that we can regard our Riemannian metric g as a $(0, 2)$ -tensor. Recall our definition for compatibility is for any two vector fields X, Y we have

$$dg(X, Y) = g(\nabla X, Y) + g(X, \nabla Y)$$

Or more explicit for vector field Z , we have

$$Zg(X, Y) = g(\nabla_Z X, Y) + g(X, \nabla_Z Y)$$

However, by definition of ∇g we have

$$\nabla_Z g(X, Y) = Zg(X, Y) - g(\nabla_Z X, Y) - g(X, \nabla_Z Y)$$

which shows that compatibility is equivalent to $\nabla g = 0$.

2.2. Type change of tensor. In general, for a (s, r) -tensor, we can change its type into any type of $(s - k, r + k)$ for all k such that $s - k \geq 0, r + k \geq 0$, since TM is canonically isomorphic to T^*M , which is called music isomorphism.

More explicitly, for any vector field X , it gives a 1-form by

$$Y \mapsto g(X, Y)$$

where Y is a vector field. Locally we have

$$\begin{aligned} g\left(\frac{\partial}{\partial x^i}, Y\right) &= g\left(\frac{\partial}{\partial x^i}, dx^j(Y) \frac{\partial}{\partial x^j}\right) \\ &= dx^j(Y) g\left(\frac{\partial}{\partial x^i}, \frac{\partial}{\partial x^j}\right) \\ &= g_{ij} dx^j(Y) \end{aligned}$$

that is $\frac{\partial}{\partial x^i}$ can be regarded as a section $g_{ij} dx^j$ of T^*M ; Similarly we can regard dx^j as a section of $TM = T^{**}M$ by

$$\omega \mapsto g(dx^j, \omega)$$

which can be written as

$$\begin{aligned} g(dx^j, \omega) &= g(dx^j, \omega\left(\frac{\partial}{\partial x^i}\right) dx^i) \\ &= \omega\left(\frac{\partial}{\partial x^i}\right) g^{ij} \end{aligned}$$

Thus dx^j can be regarded as $g^{ij} \frac{\partial}{\partial x^i}$, a section of TM .

In a summary, we have the so-called music isomorphism locally looks like

$$\begin{aligned} \flat : TM &\rightarrow T^*M & \sharp : T^*M &\rightarrow TM \\ \frac{\partial}{\partial x^i} &\mapsto g_{ij} dx^j & dx^j &\mapsto g^{ij} \frac{\partial}{\partial x^i} \end{aligned}$$

Example 2.2.1 (dual vector field). For a smooth function f , ∇f is a $(0, 1)$ -tensor, locally written as

$$\nabla f = \frac{\partial f}{\partial x^i} dx^i$$

Then we can change its type into $(1, 0)$, that is

$$\nabla f = g^{ij} \frac{\partial f}{\partial x^i} \frac{\partial}{\partial x^j}$$

More generally, for a 1-form ω , locally looks like $\omega_i dx^i$, then we can change it into a $(1, 0)$ -tensor, called its dual vector field, and it locally looks like

$$X_\omega = g^{ij} \omega_i \frac{\partial}{\partial x^j}$$

Example 2.2.2 (Induced metric on T^*M). Recall that a Riemannian metric g is a $(0, 2)$ -tensor, locally written as

$$g = g_{ij} dx^i \otimes dx^j$$

Then we can change its type into $(2, 0)$, that is

$$g_{ij} g^{ik} g^{jl} \frac{\partial}{\partial x^k} \otimes \frac{\partial}{\partial x^l} = \delta_j^k g^{jl} \frac{\partial}{\partial x^k} \otimes \frac{\partial}{\partial x^l} = g^{kl} \frac{\partial}{\partial x^k} \otimes \frac{\partial}{\partial x^l}$$

that is a metric on T^*M .

2.3. Induced metric on tensor. If g is a Riemannian metric, then its $(2,0)$ -type is a metric on T^*M . Now we can induce a metric on $T^*M \otimes T^*M$ as follows: Take two $(0,2)$ -tensors T, S and write them locally as $T = T_{ij}dx^i \otimes dx^j, S = S_{kl}dx^k \otimes dx^l$, then

$$\begin{aligned} g(T, S) &= T_{ij}S_{kl}g(dx^i \otimes dx^j, dx^k \otimes dx^l) \\ &:= T_{ij}S_{kl}g^{ik}g^{jl} \end{aligned}$$

Remark 2.3.1. In general we also have induced metric on $\bigotimes^k T^*M$, and on Ω_M^k , which will be used later in Hodge theory.

Furthermore, we have the following compatibility:

Proposition 2.3.1. If connection ∇ on vector bundle T^*M is compatible with metric g on it, then induced connection on $T^*M \otimes T^*M$ is compatible with induced metric g on it.

Proof. It suffices to check

$$\frac{\partial}{\partial x^m}g(dx^i \otimes dx^j, dx^k \otimes dx^l) = g(\nabla_{\frac{\partial}{\partial x^m}} dx^i \otimes dx^j, dx^k \otimes dx^l) + g(dx^i \otimes dx^j, \nabla_{\frac{\partial}{\partial x^m}} dx^k \otimes dx^l)$$

By compatibility of ∇ and g , we have

$$\frac{\partial g^{ij}}{\partial x^k} = -\Gamma_{kl}^i g^{lj} - \Gamma_{kl}^j g^{il}$$

Thus direct computation shows

$$\begin{aligned} \frac{\partial}{\partial x^m}g(dx^i \otimes dx^j, dx^k \otimes dx^l) &= -(\Gamma_{mn}^i g^{nk} + \Gamma_{mn}^k g^{in})g^{jl} - g^{ik}(\Gamma_{mn}^j g^{nl} + \Gamma_{mn}^l g^{jn}) \\ g(\nabla_{\frac{\partial}{\partial x^m}} dx^i \otimes dx^j, dx^k \otimes dx^l) &= -\Gamma_{mn}^i g^{nk} g^{jl} - \Gamma_{mn}^j g^{ik} g^{nl} \\ g(dx^i \otimes dx^j, \nabla_{\frac{\partial}{\partial x^m}} dx^k \otimes dx^l) &= -\Gamma_{mn}^k g^{in} g^{jl} - \Gamma_{mn}^l g^{ik} g^{jn} \end{aligned}$$

This yields the desired result. \square

2.4. Trace of tensor. Let's see a simple example: For a $(1,1)$ -tensor T , we can define its “trace”, since there is a natural isomorphism between $TM \otimes T^*M$ and $\text{End}(TM)$, thus we can take its trace in the sense of matrix. To be explicit, if we locally write T as $T = T_j^i \frac{\partial}{\partial x^i} \otimes dx^j$, then trace of T , denoted by $\text{tr}_g T$, is defined as T_i^i .

If T is not in $(1,1)$ -type, then we change it into $(1,1)$ -type and then take trace:

1. If $T = T_{ij}dx^i \otimes dx^j$, then $T = g^{ik}T_{ij}\frac{\partial}{\partial x^k} \otimes dx^j$, thus $\text{tr}_g T = g^{ij}T_{ij}$.
2. If $T = T^{ij}\frac{\partial}{\partial x^i} \otimes \frac{\partial}{\partial x^j}$, then $T = g_{kj}T^{ij}\frac{\partial}{\partial x^i} \otimes dx^k$, thus $\text{tr}_g T = g_{ij}T^{ij}$.

In general, if a tensor of type (r, s) with $r + s = 2n$, we can change its type into (n, n) and take trace n times to obtain a number. Later we will see we obtain Ricci curvature by taking trace of curvature, and we obtain scalar curvature by taking trace of Ricci curvature.

Remark 2.4.1 (scalar Laplacian). For a smooth function $f : M \rightarrow \mathbb{R}$, $\nabla^2 f$ is a $(0, 2)$ -form, locally looks like

$$\nabla_i \nabla_j f dx^i \otimes dx^j$$

Then its trace looks like

$$\text{tr}_g \nabla^2 f = g^{ij} \nabla_i \nabla_j f$$

That's called scalar Laplacian of f , denoted by $\Delta_g f$.

Remark 2.4.2. If g is induced metric on $(0, 2)$ -tensor, then for any $(0, 2)$ -tensor T , we have

$$\begin{aligned} g(g, T) &= g(g_{ij} dx^i \otimes dx^j, T_{kl} dx^k \otimes dx^l) \\ &= g_{ij} T_{kl} g^{ik} g^{jl} \\ &= \delta_j^k g^{jl} T_{kl} \\ &= g^{kl} T_{kl} \\ &= \text{tr}_g T \end{aligned}$$

Proposition 2.4.1 (magic formula). For a $(0, 2)$ -tensor T , we have

$$X(\text{tr}_g T) = g(g, \nabla_X T)$$

Proof. From above remark we can see $\text{tr}_g T = g(g, T)$, then ∇ is compatible with metric completes the proof. \square

Remark 2.4.3 (local form). Locally we have

$$\nabla_i (g^{jk} T_{jk}) = g^{jk} (\nabla_i T_{jk})$$

that is, g^{jk} can “pass through” taking covariant derivative, which is called “magic formula”.

3. GEODESIC I: NORMAL COORDINATE

3.1. Geodesic.

Definition 3.1.1 (geodesic). A smooth curve $\gamma : (-\varepsilon, \varepsilon) \rightarrow M$ is called a geodesic, if

$$\widehat{\nabla}_{\frac{d}{dt}} \gamma_* \left(\frac{d}{dt} \right) = 0$$

Remark 3.1.1 (local form). Locally we have

$$\begin{aligned} \widehat{\nabla}_{\frac{d}{dt}} \gamma_* \left(\frac{d}{dt} \right) &:= \nabla_{\gamma_* \left(\frac{d}{dt} \right)} \gamma_* \left(\frac{d}{dt} \right) \\ &= \nabla_{\frac{d\gamma^i}{dt} \frac{\partial}{\partial x^i}} \frac{d\gamma^j}{dt} \frac{\partial}{\partial x^j} \\ &= \left(\frac{d^2 \gamma^j}{dt^2} \frac{\partial}{\partial x^j} + \frac{d\gamma^i}{dt} \frac{d\gamma^j}{dt} \Gamma_{ij}^k \frac{\partial}{\partial x^k} \right) \\ &= \left(\frac{d^2 \gamma^k}{dt^2} + \frac{d\gamma^i}{dt} \frac{d\gamma^j}{dt} \Gamma_{ij}^k \right) \frac{\partial}{\partial x^k} \end{aligned}$$

Thus condition for geodesic is a system of ODEs locally. So if we fix $\gamma(0)$, $\frac{d\gamma^k}{dt}(0)$, then the existence and uniqueness of geodesic follow from standard result of ODEs.

Theorem 3.1.1. Let (M, g) be a Riemannian manifold. For any $p \in M, v \in T_p M$, there exists $\varepsilon > 0$ and a geodesic $\gamma : (-\varepsilon, \varepsilon) \rightarrow M$ such that

$$\gamma(0) = p, \gamma'(0) = v$$

Moreover, any two such geodesics agree on their common domain.

Remark 3.1.2. Since any geodesic γ with $\gamma(0) = p, \gamma'(0) = v$ will agree on their common domain, thus by gluing these geodesics together we can obtain a unique geodesic $\gamma : I \rightarrow M$ which can not be extended to a geodesic defined on a larger interval. This unique geodesic is denoted by $\gamma_v(t)$.

Lemma 3.1.1. For each $p \in M, v \in T_p M, c, t \in \mathbb{R}$,

$$\gamma_{cv}(t) = \gamma_v(ct)$$

whenever either side is defined.

Definition 3.1.2 (exponential map). Let (M, g) be a Riemannian manifold. For any $p \in M$ we define $V_p \subseteq T_p M$ by

$$V_p := \{v \in T_p M \mid \gamma_v(1) \text{ is defined}\}$$

The exponential map at p is the map

$$\begin{aligned} \exp_p : V_p &\rightarrow M \\ v &\mapsto \gamma_v(1) \end{aligned}$$

Remark 3.1.3. Although V_p may not be the whole T_pM , it always at least contains a small neighborhood of $0 \in T_pM$ from Lemma 3.1.1. In fact, later we will see Hopf-Rinow theorem implies if M is complete as a metric space, then $V_p = T_pM$ for any $p \in M$.

Theorem 3.1.2. The exponential map \exp_p maps a neighborhood $0 \in T_pM$ diffeomorphically onto a neighborhood of $p \in M$.

Proof. Note that

$$(\mathrm{d}\exp_p)_0 : T_0(T_pM) \rightarrow T_pM$$

Since T_pM is a vector space, we can identify it with T_0T_pM . Thus $(\mathrm{d}\exp_p)_0$ then becomes a map from T_pM onto itself. To see what we need, it suffices to check $(\mathrm{d}\exp_p)_0$ is identity map. For all $v \in T_pM$,

$$\begin{aligned} (\mathrm{d}\exp_p)_0(v) &= \left. \frac{d}{dt} \right|_{t=0} \exp_p(0 + tv) \\ &= \left. \frac{d}{dt} \right|_{t=0} \gamma_{tv}(1) \\ &= \left. \frac{d}{dt} \right|_{t=0} \gamma_v(t) \\ &= \gamma'_v(0) \\ &= v \end{aligned}$$

□

Remark 3.1.4 (normal coordinate). Fix a basis $\left. \frac{\partial}{\partial x^1} \right|_p, \dots, \left. \frac{\partial}{\partial x^n} \right|_p$ of T_pM which is orthonormal with respect to Riemannian metric g , we have the following linear isomorphism

$$\begin{aligned} \Phi : T_pM &\rightarrow \mathbb{R}^n \\ v^i \left. \frac{\partial}{\partial x^i} \right|_p &\mapsto (v^1, \dots, v^n) \end{aligned}$$

Then Theorem 3.1.2 implies there exists a neighborhood U of p which is mapped by $\Phi \circ \exp_p^{-1}$ diffeomorphically onto a neighborhood of $0 \in \mathbb{R}^n$. Thus $(r := \Phi \circ \exp_p^{-1}, U)$ gives a local coordinates of M with center p , which is called normal coordinate.

Given a coordinate (ϕ, U) , we give a more explicit formula for function f on manifold, that is $f(x) := f(\phi^{-1}(x))$. In particular, if we consider normal coordinate, we have the following characterization for Riemannian metric and Christoffel symbols.

Theorem 3.1.3. In normal coordinate we have

$$\begin{aligned} g_{ij}(0) &= \delta_{ij} \\ \Gamma_{ij}^k(0) &= 0 \end{aligned}$$

Proof. Note that

$$\begin{aligned}
g_{ij}(0) &= \langle d(\exp_p \circ \Phi^{-1})_0 e_i, d(\exp_p \circ \Phi^{-1})_0 e_j \rangle_p \\
&= \langle (d \exp_p)_0 \left. \frac{\partial}{\partial x^i} \right|_p, (d \exp_p)_0 \left. \frac{\partial}{\partial x^j} \right|_p \rangle_p \\
&= \langle \left. \frac{\partial}{\partial x^i} \right|_p, \left. \frac{\partial}{\partial x^j} \right|_p \rangle_p \\
&= \delta_{ij}
\end{aligned}$$

where $e_i = (0, \dots, \underbrace{1}_{i\text{-th}}, \dots, 0) \in \mathbb{R}^n$.

For Christoffel symbol: For arbitrary $v = (v^1, \dots, v^n) \in \mathbb{R}^n$, consider geodesic $\gamma(t) = \exp_p(t\Phi^{-1}(v))$ with $\gamma(0) = p$ and $\gamma'(t) = \Phi^{-1}(v)$. In normal coordinate γ looks like $\gamma(t) = (tv^1, \dots, tv^n)$, thus geodesic equation simplifies to

$$\Gamma_{ij}^k(tv)v^i v^j = 0$$

Evaluating this expression at $t = 0$ shows $\Gamma_{ij}^k(0)v^i v^j = 0$ for arbitrary index k and every v . Now take $v = \frac{1}{2}(e_i + e_j)$ to conclude $\Gamma_{ij}^k(0) = 0$ for all i, j, k . \square

Corollary 3.1.1. In normal coordinate we have for Taylor expression of $g_{ij} : T_p M \rightarrow \mathbb{R}$ around zero as

$$g_{ij}(x) = \delta_{ij} + O(|x|^2)$$

Proof. Note that

$$\frac{\partial g_{ij}}{\partial x^k} = \Gamma_{ki}^l(0)g_{lj}(0) + \Gamma_{kj}^l(0)g_{il}(0) = 0$$

\square

3.2. Arts of computation. Tensor computation is one of the hallmarks of Riemannian geometry, but sometimes there is a way to avoid unnecessary computations. In this section we collect some useful tools which can simplify the computations.

Note that if we want to check two tensors are same, it suffices to check pointwise. Furthermore, although for a general tensor, its value depends on the choice of coordinates, zero is independent of the choice of coordinates. So in order to check, we only need to find an appropriate coordinate.

Geodesics give us such a coordinate, that is normal coordinate, we always use (x^i, U, p) to the normal coordinate (x^i, U, p) centered at $p \in M$. According to Theorem 3.1.3, one has

$$\begin{aligned}
x^i(p) &= 0 \\
g_{ij}(p) &= \delta_{ij} \\
\Gamma_{ij}^k(p) &= 0
\end{aligned}$$

that is under normal coordinate, metric looks like standard metric in Euclidean space, which largely simplify the computations.

In fact, when we're considering a problem which is independent of the choice of coordinates, we may compute it under normal coordinate. For example

Lemma 3.2.1. Let (M, g) be a Riemannian manifold with connection ∇ . Given an arbitrary local basis $\{\frac{\partial}{\partial x^i}\}$ of TM with dual basis $\{dx^i\}$, then

$$d = dx^i \wedge \nabla \frac{\partial}{\partial x^i}$$

Proof. Firstly note that exterior derivative is independent of the choice of coordinates, and we claim so does $dx^i \wedge \nabla \frac{\partial}{\partial x^i}$. Indeed, assume $\{\frac{\partial}{\partial y^j}\}$ is another local basis with dual basis $\{dy^j\}$, and transition functions are given by

$$\begin{aligned} dy^k &= d_i^k dx^i \\ \frac{\partial}{\partial y^k} &= c_k^i \frac{\partial}{\partial x^i} \end{aligned}$$

It's clear $d_j^k c_k^i = \delta_j^i$, since $1 = dy^k(\frac{\partial}{\partial y^k}) = d_j^k dx^j(c_k^i \frac{\partial}{\partial x^i}) = d_j^k c_k^i \delta_i^j$. Thus

$$\begin{aligned} dy^k \wedge \nabla \frac{\partial}{\partial y^k} &= d_j^k dx^j \wedge \nabla c_k^i \frac{\partial}{\partial x^i} \\ &= d_j^k c_k^i dx^j \wedge \nabla \frac{\partial}{\partial x^i} \\ &= \delta_j^i dx^j \wedge \nabla \frac{\partial}{\partial x^i} \\ &= dx^i \wedge \nabla \frac{\partial}{\partial x^i} \end{aligned}$$

Now it suffices to check $d = dx^i \wedge \nabla \frac{\partial}{\partial x^i}$ in normal coordinate, that is clear, since for arbitrary k -form ω , if we write it as $f dx^1 \wedge \cdots \wedge dx^k$, then

$$\begin{aligned} dx^i \wedge \nabla \frac{\partial}{\partial x^i} \omega &= dx^i \wedge \nabla \frac{\partial}{\partial x^i} (f dx^1 \wedge \cdots \wedge dx^k) \\ &= dx^i \wedge \frac{\partial f}{\partial x^i} dx^1 \wedge \cdots \wedge dx^k \\ &= \frac{\partial f}{\partial x^i} dx^i \wedge dx^1 \wedge \cdots \wedge dx^k \\ &= d\omega \end{aligned}$$

□

Remark 3.2.1. Note that here connection ∇ is an arbitrary connection, not necessarily Levi-Civita.

3.3. Geodesics on Lie group. In this section we give a quick review of Lie groups, such as left-invariant vector fields and integral curves. Then we consider the invariant metrics on Lie groups G and we show that geodesics are exactly integral curves (or one parameter subgroup) of G , and that's why we define exponential map of Lie group by integral curves.

3.3.1. A quick review of Lie group.

Definition 3.3.1 (Lie group). A Lie group G is a smooth manifold which is also endowed with a group structure such that the multiplication map and the inverse map are smooth.

Since the multiplication map is smooth, then for any $g \in G$, there are two smooth maps L_g, R_g , defined by

$$\begin{aligned} L_g(h) &= gh \\ R_g(h) &= hg \end{aligned}$$

Furthermore, they're also diffeomorphisms with inverse $L_{g^{-1}}, R_{g^{-1}}$, since inverse maps are also smooth.

Definition 3.3.2 (invariant vector field). A vector field X on a Lie group G is called left-invariant, if

$$(L_g)_*X = X$$

for arbitrary $g \in G$.

Remark 3.3.1. It's clear there is the following isomorphism

$$\begin{aligned} \{\text{Left-invariant vector fields}\} &\rightarrow \mathfrak{g} := T_e G \\ X &\mapsto X_e \end{aligned}$$

where X_e is its value in $T_e G$. Furthermore, since Lie bracket of two left-invariant vector fields is still left-invariant, thus there is a natural Lie bracket on \mathfrak{g} .

Definition 3.3.3 (Lie algebra). The tangent space $T_e G$ of a Lie group G equipped with Lie bracket is called Lie algebra of G , denoted by \mathfrak{g} .

Definition 3.3.4 (adjoint representation). The adjoint representation is defined as follows

$$\begin{aligned} \text{Ad} : G &\rightarrow \text{GL}(\mathfrak{g}) \\ g &\mapsto (R_{g^{-1}} \circ L_g)_* \end{aligned}$$

Definition 3.3.5 (integral curve). Let X be a vector field of G and $g \in G$, then an integral curve of X through the point p is a smooth curve $\gamma : I \subseteq \mathbb{R} \rightarrow G$ such that

$$\begin{aligned} \gamma(0) &= g \\ \gamma'(t) &= X(\gamma(t)) \end{aligned}$$

Definition 3.3.6 (complete vector field). A vector field X is called complete, if its integral curve is defined for all $t \in \mathbb{R}$.

Proposition 3.3.1. Every left-invariant vector field on a Lie group G is complete.

Proof. Let X be a left-invariant vector field, γ the unique integral curve for X such that $\gamma(0) = e$, defined on $(-\varepsilon, \varepsilon)$. Then $\gamma_g := L_g\gamma$ is an integral curve for X such that $\gamma_g(0) = g$. Indeed,

$$\begin{aligned}\gamma'_g(t) &= d(L_g)_{\gamma(t)}(\gamma'(t)) \\ &= d(L_g)_{\gamma(t)}(X(\gamma(t))) \\ &= X(L_g\gamma(t)) \\ &= X(\gamma_g(t))\end{aligned}$$

In particular, for $t_0 \in (-\varepsilon, \varepsilon)$, the curve $t \mapsto \gamma(t_0)\gamma(t)$ is an integral curve for X starting at $\gamma(t_0)$. By uniqueness, this curve coincides with $\gamma(t_0 + t)$ for all $t \in (-\varepsilon, \varepsilon) \cap (-\varepsilon - t_0, \varepsilon - t_0)$. Define

$$\tilde{\gamma}(t) = \begin{cases} \gamma(t), & t \in (-\varepsilon, \varepsilon) \\ \gamma(t_0)\gamma(t), & t \in (-\varepsilon - t_0, \varepsilon - t_0) \end{cases}$$

Repeat above operations to get our desired extension. \square

Remark 3.3.2. From this proof we can see integral curve of left-invariant vector fields through identity e is just a Lie group homomorphism $\gamma : \mathbb{R} \rightarrow G$, such homomorphism is called a one parameter subgroup.

3.3.2. Riemannian geometry of Lie group.

Definition 3.3.7 (left-invariant metric). A Riemannian metric h on a Lie group G is called left-invariant if

$$L_g^*h = h$$

for arbitrary $g \in G$.

Remark 3.3.3. Similarly we can define a right invariant metric, and a Riemannian metric which is both left-invariant and right invariant is called bi-invariant metric.

Proposition 3.3.2. There is a bijective correspondence between left-invariant metrics on a Lie group G , and inner products on the Lie algebra \mathfrak{g} of G .

Proof. Given an inner product $\langle -, - \rangle_e$ on Lie algebra \mathfrak{g} , then we have an inner product on G defined as follows

$$\langle X_g, Y_g \rangle := \langle (L_{g^{-1}})_*X_g, (L_{g^{-1}})_*Y_g \rangle_e$$

where X, Y are two vector fields on G . It's left-invariant, since

$$\begin{aligned}\langle (L_h)_*X_g, (L_h)_*Y_g \rangle &= \langle (L_{(hg)^{-1}})_*(L_h)_*X_g, (L_{(hg)^{-1}})_*(L_h)_*Y_g \rangle_e \\ &= \langle (L_{g^{-1}})_*X_g, (L_{g^{-1}})_*Y_g \rangle_e\end{aligned}$$

Conversely, if we have a left-invariant inner product $\langle -, - \rangle$ on G , then it's clear we have an inner product on \mathfrak{g} , by just considering its value at identity. Furthermore, these two constructions are inverse to each other, this completes the proof. \square

Proposition 3.3.3. There is a bijective correspondence between bi-invariant metrics on a Lie group G , and Ad-invariant inner products on the Lie algebra \mathfrak{g} of G .

Proof. Given a Ad-invariant inner product $\langle -, - \rangle_e$ on the Lie algebra \mathfrak{g} , by Proposition 3.3.2, there is a left-invariant metric $\langle -, - \rangle$ on G , it suffices to check it's also right-invariant:

$$\begin{aligned} \langle (R_h)_* X_g, (R_h)_* Y_g \rangle &= \langle (L_{(hg)^{-1}})_* (R_h)_* X_g, (L_{(hg)^{-1}})_* (R_h)_* Y_g \rangle_e \\ &= \langle \text{Ad}(h^{-1})(L_{g^{-1}})_* X_g, \text{Ad}(h^{-1})(L_{g^{-1}})_* Y_g \rangle_e \\ &= \langle (L_{g^{-1}})_* X_g, (L_{g^{-1}})_* Y_g \rangle_e \\ &= \langle X_g, Y_g \rangle \end{aligned}$$

Conversely, if we start with a bi-invariant metric, then it's restriction to the Lie algebra is a Ad-invariant, since $\text{Ad}(g)$ is exactly the differential of $L_g \circ R_{g^{-1}}$. \square

Remark 3.3.4. In particular, if G is a compact connected Lie group, then it admits a bi-invariant metric, since its Killing form is a Ad-invariant inner product on \mathfrak{g} .

Lemma 3.3.1. If h is a left-invariant metric on a Lie group G , ∇ is the Levi-Civita connection, then for all left-invariant vector fields X, Y, Z , we have

$$h(X, \nabla_Y Y) = h(Y, [X, Y])$$

Proof. Recall that

$$h(X, \nabla_Y Z) = \frac{1}{2}(Yh(Z, X) + Zh(X, Y) - Xh(Y, Z) - h([Y, X], Z) - h([Z, X], Y) - h([Y, Z], X))$$

But $Yh(Z, X) = Zh(X, Y) = Xh(Y, Z) = 0$ since h is left-invariant and X, Y, Z is left-invariant, that is

$$h(X, \nabla_Y Z) = \frac{1}{2}\{h(Z, [X, Y]) + h(Y, [X, Z]) + h(X, [Z, Y])\}$$

Now take $Y = Z$ to conclude. \square

Proposition 3.3.4. If h is a bi-invariant metric on a Lie group G , then for all left-invariant vector fields X, Y, Z , we have

$$h([X, Y], Z) = h(X, [Y, Z])$$

Proof. Let y_t be the flow of Y , then

$$[X, Y] = \lim_{t \rightarrow 0} \frac{1}{t} ((y_t)_*(X) - X)$$

On the other hand, since Y is left-invariant, that is $L_g \circ y_t = y_t \circ L_g$, giving

$$y_t(g) = y_t(L_g(e)) = L_g y_t(e) = g y_t(e) = R_{y_t(e)}(g)$$

Thus $(y_t)_* = (R_{y_t(e)})_*$ and

$$[X, Y] = \lim_{t \rightarrow 0} \frac{1}{t} ((R_{y_t(e)})_*(X) - X)$$

Note that h is bi-invariant, thus

$$\begin{aligned} h(X, Z) &= h((R_{y_t(e)})_*(L_{y_t^{-1}(e)})_*X, (R_{y_t(e)})_*(L_{y_t^{-1}(e)})_*Z) \\ &= h((R_{y_t(e)})_*X, (R_{y_t(e)})_*Z) \end{aligned}$$

Differentiating the expression above with respect to t and setting $t = 0$ we conclude

$$0 = h([X, Y], Z) + h(X, [Z, Y])$$

This completes the proof. \square

Theorem 3.3.1. For every left-invariant vector field X on G , then $\nabla_X X = 0$.

Proof. From Lemma 3.3.1, we have

$$h(Y, \nabla_X X) = h(X, [Y, X])$$

where h is a bi-invariant metric. From Proposition 3.3.4, we have

$$h(X, [Y, X]) = h([Y, X], X) = -h(X, [Y, X])$$

that is $h(Y, \nabla_X X) = 0$ for arbitrary vector field Y , which implies $\nabla_X X = 0$. \square

Corollary 3.3.1. If X, Y are left-invariant vector field, then $\nabla_X Y = \frac{1}{2}[X, Y]$.

Proof. Note that

$$\begin{aligned} 0 &= \nabla_{X+Y}(X + Y) \\ &= \nabla_X Y + \nabla_Y X + \nabla_X X + \nabla_Y Y \\ &= \nabla_X Y + \nabla_Y X \\ &= 2\nabla_X Y - [X, Y] \end{aligned}$$

Division by two finally yields

$$\nabla_X Y = \frac{1}{2}[X, Y]$$

\square

Theorem 3.3.2. The geodesics on G are precisely the integral curves of left-invariant vector fields.

Proof. Let $X \in \mathfrak{g}$ be a left-invariant vector field, and $\gamma : \mathbb{R} \rightarrow G$ its integral curve. Then

$$\begin{aligned} \widehat{\nabla}_{\frac{d}{dt}} \gamma_* \left(\frac{d}{dt} \right) &= \nabla_{\gamma_* \left(\frac{d}{dt} \right)} \gamma_* \left(\frac{d}{dt} \right) \\ &= \nabla_X X \\ &= 0 \end{aligned}$$

which implies integral curves of left-invariant vector fields are geodesics.

Furthermore, since geodesics are unique, we have geodesics are precisely integral curves of left-invariant vector fields. \square

Corollary 3.3.2. The exponential map for the Lie group coincides with the exponential map of the Levi-Civita connection.

3.4. Geodesic convex neighborhood. Recall that for any $p \in M$, there exists a normal neighborhood U , that is (U, \exp_p^{-1}) is a diffeomorphism. Or in other words, any two points $p_1, p_2 \in U$, there exists a geodesic connecting these two points. However, this geodesic may not lie in U .

Definition 3.4.1 (strongly convex). A subset S of M is strongly convex if for any two points $p_1, p_2 \in \bar{S}$, there exists a geodesic connecting p_1, p_2 whose interior is contained in S .

Proposition 3.4.1 (convex neighborhood). For any $p \in M$, there exists a strongly convex normal neighborhood.

Proof. See Proposition 4.2 of page 76 in [Car92]. \square

Remark 3.4.1. Convex neighborhood is a technique tool which will be used in topology. Note that convex set is contractible and the intersection of convex sets is still convex, thus from Proposition 3.4.1, we know that there exists a open covering of M such that

1. U_α is contractible;
2. For any finite intersection $\bigcap_{i=1}^r U_{\alpha_i}$, it's still contractible.

Such covering is called a “good cover” in [RB82], and it's widely used in Mayer-Vietoris argument.

4. CURVATURE

4.1. Curvature form. Let (M, g) be a Riemannian manifold, $\nabla : C^\infty(M, E) \rightarrow C^\infty(M, T^*M \otimes E)$ a connection of vector bundles E on M . Now we're going to extend connection to something called exterior derivative defined on sections of vector bundle valued k -forms as follows

$$\begin{aligned} d^\nabla : C^\infty(M, \Omega_M^k \otimes E) &\rightarrow C^\infty(M, \Omega_M^{k+1} \otimes E) \\ \omega \otimes e &\mapsto d\omega \otimes e + (-1)^k \omega \wedge \nabla e \end{aligned}$$

Note that d^∇ on $C^\infty(M, E)$ is exactly ∇ . If we use Ω to denote $d^\nabla \circ d^\nabla$, let's see Ω locally:

$$\begin{aligned} \Omega(s^\alpha e_\alpha) &= d^\nabla(ds^\alpha e_\alpha + s^\alpha \omega_\alpha^\beta e_\beta) \\ &= -ds^\alpha \wedge \omega_\alpha^\beta e_\beta + d(s^\alpha \omega_\alpha^\beta) e_\beta - s^\alpha \omega_\alpha^\beta \wedge \omega_\beta^\gamma e_\gamma \\ &= s^\alpha (d\omega_\alpha^\beta - \omega_\alpha^\gamma \wedge \omega_\gamma^\beta) e_\beta \\ \Omega(e_\alpha) &= d^\nabla(\omega_\alpha^\beta e_\beta) \\ &= d\omega_\alpha^\beta e_\beta - \omega_\alpha^\beta \wedge \nabla e_\beta \\ &= d\omega_\alpha^\beta e_\beta - \omega_\alpha^\beta \wedge \omega_\beta^\gamma e_\gamma \\ &= (d\omega_\alpha^\beta - \omega_\alpha^\gamma \wedge \omega_\gamma^\beta) e_\beta \end{aligned}$$

This shows smooth functions commutes with Ω . This is a quite good property, from this we can conclude:

1. $\Omega(e_\alpha)$ completely determines Ω locally, thus we can say Ω locally looks like $d\omega - \omega \wedge \omega$;
2. Ω is a global section of $\Omega_M^2 \otimes \text{End } E$, that is it's compatible with change of basis. Indeed, for two local basis e, \tilde{e} such that $\tilde{e} = ge$, we will see

$$\begin{aligned} g\nabla^2 e &= \nabla^2 ge \\ &= \nabla^2 \tilde{e} \\ &= (d\tilde{\omega} - \tilde{\omega} \wedge \tilde{\omega})\tilde{e} \\ &= (d\tilde{\omega} - \tilde{\omega} \wedge \tilde{\omega})ge \end{aligned}$$

which implies

$$g^{-1}(d\tilde{\omega} - \tilde{\omega} \wedge \tilde{\omega})g = d\omega - \omega \wedge \omega$$

Definition 4.1.1 (curvature form). For a connection ∇ of a vector bundle E on M , its curvature form $\Omega \in C^\infty(M, \Omega_M^2 \otimes \text{End } E)$ is defined as above.

Remark 4.1.1 (local form). We can give a more explicit expression of Ω using Christoffel symbol: If we locally write Ω as

$$\Omega_\alpha^\beta = \Omega_{ij\alpha}^\beta dx^i \wedge dx^j$$

Then $\Omega = d\omega - \omega \wedge \omega$ can be written as

$$\begin{aligned}\Omega_{ij\alpha}^\beta dx^i \wedge dx^j &= \Omega_\alpha^\beta \\ &= d\omega_\alpha^\beta - \omega_\alpha^\gamma \wedge \omega_\gamma^\beta \\ &= d(\Gamma_{i\alpha}^\beta dx^i) - (\Gamma_{i\alpha}^\gamma dx^i) \wedge (\Gamma_{j\gamma}^\beta dx^j) \\ &= (-\partial_j \Gamma_{i\alpha}^\beta - \Gamma_{i\alpha}^\gamma \Gamma_{j\gamma}^\beta) dx^i \wedge dx^j\end{aligned}$$

4.2. Curvature tensor. In Do carmo [Car92], he defines the curvature of a connection ∇ as follows:

$$\begin{aligned}R : TM \times TM \times E &\rightarrow E \\ (X, Y, s) &\mapsto R(X, Y)s\end{aligned}$$

where $R(X, Y)s = \nabla_X \nabla_Y s - \nabla_Y \nabla_X s - \nabla_{[X, Y]}s$. It's easy to check R we defined above is a tensor, that is a section of $T^*M \otimes T^*M \otimes \text{End } E$.

Remark 4.2.1 (local form). Locally we write R as

$$R = R_{ij\alpha}^\beta dx^i \otimes dx^j \otimes e^\alpha \otimes e_\beta$$

To see $R_{ij\alpha}^\beta$, it suffices to compute

$$\begin{aligned}\nabla_{\frac{\partial}{\partial x^i}} \nabla_{\frac{\partial}{\partial x^j}} e_\alpha &= \nabla_{\frac{\partial}{\partial x^i}} (\Gamma_{j\alpha}^\beta e_\beta) \\ &= \partial_i \Gamma_{j\alpha}^\beta e_\beta + \Gamma_{j\alpha}^\gamma \Gamma_{i\gamma}^\beta e_\beta \\ &= (\partial_i \Gamma_{j\alpha}^\beta + \Gamma_{j\alpha}^\gamma \Gamma_{i\gamma}^\beta) e_\beta\end{aligned}$$

Thus

$$R_{ij\alpha}^\beta e_\beta = (\partial_i \Gamma_{j\alpha}^\beta - \partial_j \Gamma_{i\alpha}^\beta + \Gamma_{j\alpha}^\gamma \Gamma_{i\gamma}^\beta - \Gamma_{i\alpha}^\gamma \Gamma_{j\gamma}^\beta) e_\beta$$

or in other words,

$$R_\alpha^\beta = (\partial_i \Gamma_{j\alpha}^\beta - \partial_j \Gamma_{i\alpha}^\beta + \Gamma_{j\alpha}^\gamma \Gamma_{i\gamma}^\beta - \Gamma_{i\alpha}^\gamma \Gamma_{j\gamma}^\beta) dx^i \otimes dx^j$$

Recall that our curvature form Ω is a section of $\Omega_M^2 \otimes \text{End } E$, and you can regard it as a section of $T^*M \otimes T^*M \otimes \text{End } E$, that is

$$\begin{aligned}\Omega_\alpha^\beta &= (-\partial_j \Gamma_{i\alpha}^\beta - \Gamma_{i\alpha}^\gamma \Gamma_{j\gamma}^\beta) dx^i \wedge dx^j \\ &= (\partial_i \Gamma_{j\alpha}^\beta - \partial_j \Gamma_{i\alpha}^\beta + \Gamma_{j\alpha}^\gamma \Gamma_{i\gamma}^\beta - \Gamma_{i\alpha}^\gamma \Gamma_{j\gamma}^\beta) dx^i \otimes dx^j\end{aligned}$$

So if you regard curvature form as a tensor, then it's exactly curvature tensor we defined here.

If we take E to be tangent bundle, then we can regard R as a $(1, 3)$ -tensor, locally looks like

$$R_{ijk}^r dx^i \otimes dx^j \otimes dx^k \otimes \frac{\partial}{\partial x^r}$$

However, we always use its $(0, 4)$ type, that is

$$R_{ijkl} = g_{rl} R_{ijk}^r$$

Now let's give a more explicit expression about R_{ijkl} . By definition we directly have

$$\begin{aligned}
R_{ijkl} &= R\left(\frac{\partial}{\partial x^i}, \frac{\partial}{\partial x^j}, \frac{\partial}{\partial x^k}, \frac{\partial}{\partial x^l}\right) \\
&= \left\langle \nabla_{\frac{\partial}{\partial x^i}} \nabla_{\frac{\partial}{\partial x^j}} \frac{\partial}{\partial x^k} - \nabla_{\frac{\partial}{\partial x^j}} \nabla_{\frac{\partial}{\partial x^i}} \frac{\partial}{\partial x^k}, \frac{\partial}{\partial x^l} \right\rangle \\
&= \partial_i \left\langle \nabla_{\frac{\partial}{\partial x^j}} \frac{\partial}{\partial x^k}, \frac{\partial}{\partial x^l} \right\rangle - \left\langle \nabla_{\frac{\partial}{\partial x^j}} \frac{\partial}{\partial x^k}, \nabla_{\frac{\partial}{\partial x^i}} \frac{\partial}{\partial x^l} \right\rangle - (\partial_j \left\langle \nabla_{\frac{\partial}{\partial x^i}} \frac{\partial}{\partial x^k}, \frac{\partial}{\partial x^l} \right\rangle - \left\langle \nabla_{\frac{\partial}{\partial x^i}} \frac{\partial}{\partial x^k}, \nabla_{\frac{\partial}{\partial x^j}} \frac{\partial}{\partial x^l} \right\rangle) \\
&= \underbrace{\partial_i \left\langle \nabla_{\frac{\partial}{\partial x^j}} \frac{\partial}{\partial x^k}, \frac{\partial}{\partial x^l} \right\rangle - \partial_j \left\langle \nabla_{\frac{\partial}{\partial x^i}} \frac{\partial}{\partial x^k}, \frac{\partial}{\partial x^l} \right\rangle}_{\text{part I}} + \underbrace{\left\langle \nabla_{\frac{\partial}{\partial x^i}} \frac{\partial}{\partial x^k}, \nabla_{\frac{\partial}{\partial x^j}} \frac{\partial}{\partial x^l} \right\rangle - \left\langle \nabla_{\frac{\partial}{\partial x^j}} \frac{\partial}{\partial x^k}, \nabla_{\frac{\partial}{\partial x^i}} \frac{\partial}{\partial x^l} \right\rangle}_{\text{part II}}
\end{aligned}$$

For part II, we have

$$g_{rs}(\Gamma_{ik}^r \Gamma_{jl}^s - \Gamma_{jk}^r \Gamma_{il}^s)$$

For part I, note that

$$\begin{aligned}
\partial_i(\Gamma_{jk}^r g_{rl}) &= \partial_i\left(\frac{1}{2}g^{rs}(\partial_j g_{ks} + \partial_k g_{js} - \partial_s g_{jk})g_{rl}\right) \\
&= \partial_i\left(\frac{1}{2}\delta_l^s(\partial_j g_{ks} + \partial_k g_{js} - \partial_s g_{jk})\right) \\
&= \frac{1}{2}\partial_i(\partial_j g_{kl} + \partial_k g_{jl} - \partial_l g_{jk})
\end{aligned}$$

Thus we have part I is

$$\partial_i(\Gamma_{jk}^r g_{rl}) - \partial_j(\Gamma_{ik}^r g_{rl}) = \frac{1}{2}(\partial_i \partial_k g_{jl} + \partial_j \partial_l g_{ik} - \partial_i \partial_l g_{jk} - \partial_j \partial_k g_{il})$$

So we have an explicit expression for R_{ijkl}

$$R_{ijkl} = \frac{1}{2}(\partial_i \partial_k g_{jl} + \partial_j \partial_l g_{ik} - \partial_i \partial_l g_{jk} - \partial_j \partial_k g_{il}) + g_{rs}(\Gamma_{ik}^r \Gamma_{jl}^s - \Gamma_{jk}^r \Gamma_{il}^s)$$

From this expression, we can see in general curvature depends on two order partial derivatives of metric. Furthermore, there are some skew symmetries and symmetries of R_{ijkl} .

1. $R_{ijkl} = -R_{jikl}$;
2. $R_{ijkl} = -R_{ijlk}$;
3. $R_{ijkl} = R_{klij}$.

Proposition 4.2.1. In Riemannian normal coordinate we have

$$g_{ij} = \delta_{ij} - \frac{1}{3}R_{iklj}(0)x^k x^l + O(|x|^3)$$

Proof. Recall we already have

$$\frac{\partial g_{ij}}{\partial x^k} = \Gamma_{ki}^m g_{mj} + \Gamma_{kj}^m g_{mi}$$

We differential the equation with respect to x^l , evaluate at 0 and use the fact that Christoffel symbol vanishes to have

$$\frac{\partial^2 g_{ij}}{\partial x^l \partial x^k} = \frac{\partial \Gamma_{ki}^m}{\partial x^l}(0) g_{mj}(0) + \frac{\partial \Gamma_{kj}^m}{\partial x^l}(0) g_{mi}(0)$$

Here we claim

$$\frac{\partial \Gamma_{ij}^k}{\partial x^l}(0) + \frac{\partial \Gamma_{li}^k}{\partial x^j}(0) + \frac{\partial \Gamma_{jl}^k}{\partial x^i}(0) = 0$$

Indeed, in normal coordinate we have

$$0 = \Gamma_{ij}^k(tx) x^i x^j$$

Then differential this with respect to t and evaluate at $t = 0$ we have

$$0 = \frac{\partial \Gamma_{ij}^k}{\partial x^l}(0) x^i x^j x^l$$

which implies

$$\sum_{\sigma \in S_3} \frac{\partial \Gamma_{\sigma(i)\sigma(j)}^k}{\partial x^{\sigma(l)}}(0) = 0$$

Then use symmetry of Christoffel symbol in term i, j to conclude.

From $R_{ijk}^l(0) = \frac{\partial \Gamma_{jk}^l}{\partial x^i} - \frac{\partial \Gamma_{ik}^l}{\partial x^j}$ we have

$$\begin{aligned} R_{ijkl}(0) &= \left(\frac{\partial \Gamma_{jk}^m}{\partial x^i}(0) - \frac{\partial \Gamma_{ik}^m}{\partial x^j}(0) \right) g_{ml}(0) \\ &= - \left(\frac{\partial \Gamma_{ij}^m}{\partial x^k}(0) + \frac{\partial \Gamma_{ki}^m}{\partial x^j}(0) + \frac{\partial \Gamma_{ik}^m}{\partial x^j}(0) \right) g_{ml}(0) \\ &= - \left(\frac{\partial \Gamma_{ij}^m}{\partial x^k}(0) + 2 \frac{\partial \Gamma_{ki}^m}{\partial x^j}(0) \right) g_{ml}(0) \end{aligned}$$

Thus we have

$$\begin{aligned} 2R_{ikjl}(0) x^k x^l &= - (R_{iklj}(0) + R_{jlki}(0)) x^k x^l \\ &= \left(\frac{\partial \Gamma_{ik}^m}{\partial x^l}(0) + 2 \frac{\partial \Gamma_{il}^m}{\partial x^k}(0) \right) g_{mj}(0) x^k x^l \\ &\quad + \left(\frac{\partial \Gamma_{jl}^m}{\partial x^k}(0) + 2 \frac{\partial \Gamma_{jk}^m}{\partial x^l}(0) \right) g_{mi}(0) x^k x^l \\ &= 3 \frac{\partial g_{ij}}{\partial x^k \partial x^l}(0) x^k x^l \end{aligned}$$

Thus we get for the second term in the Taylor expansion

$$\begin{aligned} \frac{1}{2} \frac{\partial^2 g_{ij}}{\partial x^k \partial x^l}(0) x^k x^l &= \frac{1}{3} R_{ikjl}(0) x^k x^l \\ &= -\frac{1}{3} R_{iklj}(0) x^k x^l \end{aligned}$$

that is

$$g_{ij} = \delta_{ij} - \frac{1}{3} R_{iklj}(0) x^k x^l + O(|x|^3)$$

□

Corollary 4.2.1. In Riemannian normal coordinate we have

1. $g^{ij} = \delta_{ij} + \frac{1}{3}R_{iklj}(0)x^k x^l + O(|x|^3)$
2. $\det(g_{ij}) = 1 - \frac{1}{3}R_{kl}x^k x^l + O(|x|^3)$
3. $\sqrt{\det(g_{ij})} = 1 - \frac{1}{6}R_{kl}x^k x^l + O(|x|^3)$

Proof. For (1). Note that g^{ij} gives a Riemannian metric on T^*M , and Levi-Civita connection ∇ on T^*M with respect to g^{ij} is exactly the induced connection from the one on TM . Note that

$$\nabla dx^k = -\Gamma_{ij}^k dx^i \otimes dx^j$$

where Γ_{ij}^k is the Christoffel symbol for Levi-Civita connection on TM , we have curvature form in this case differs a sign since

$$R_{ijk}^l(0) = \frac{\partial \Gamma_{jk}^l}{\partial x^i} - \frac{\partial \Gamma_{ik}^l}{\partial x^j}$$

Thus all computations are same as proof above, but result differs a sign in curvature.

For (2). By Jacobi's formula, we have

$$\frac{\partial \det(g_{ij})}{\partial x^k} = \det(g_{ij}) g^{ij} \frac{\partial g_{ij}}{\partial x^k}$$

Thus $\frac{\partial \det(g_{ij})}{\partial x^k}(0) = 0$, since first-order partial derivatives of g_{ij} vanishes. Furthermore, since first-order partial derivatives of g^{ij} also vanishes, we have

$$\begin{aligned} \frac{1}{2} \frac{\partial^2 \det(g_{ij})}{\partial x^l \partial x^k} &= \det(g_{ij}) g^{ij} \frac{1}{2} \frac{\partial^2 g_{ij}}{\partial x^l \partial x^k} \\ &= \det(g_{ij}) g^{ij} \left(-\frac{1}{3} R_{iklj} x^k x^l \right) \\ &= -\frac{1}{3} \det(g_{ij}) R_{kl} x^k x^l \end{aligned}$$

which implies

$$\det(g_{ij}) = 1 - \frac{1}{3} R_{kl} x^k x^l + O(|x|^3)$$

For (3). It follows from (2) directly. □

4.3. Bianchi identities. There are two famous Bianchi identities in Riemannian geometry, in Do carmo [Car92] they are stated as follows

1. First Bianchi identity: $R(X, Y, Z, W) + R(Y, Z, X, W) + R(Z, X, Y, W) = 0$;
2. Second Bianchi identity: $\nabla_X R(Y, Z, W, R) + \nabla_Y R(Z, X, W, R) + \nabla_Z R(X, Y, W, R) = 0$.

4.3.1. *First Bianchi.* Locally we have first Bianchi identity as

$$R_{ijkl} + R_{jkil} + R_{kijl} = 0$$

In order to compute we use (1, 3) type as follows

$$R_{ijk}^r + R_{jki}^r + R_{kij}^r = 0$$

since we have

$$R_{ijk}^r = \underbrace{\partial_i \Gamma_{jk}^r - \partial_j \Gamma_{ik}^r}_{\text{part I}} + \underbrace{\Gamma_{jk}^s \Gamma_{is}^r - \Gamma_{ik}^s \Gamma_{js}^r}_{\text{part II}}$$

1. For the first part, if we permuting i, j, k , we have

$$\partial_i \Gamma_{jk}^r - \partial_j \Gamma_{ik}^r + \partial_j \Gamma_{ki}^r - \partial_k \Gamma_{ji}^r + \partial_k \Gamma_{ij}^r - \partial_i \Gamma_{kj}^r = 0$$

since $\Gamma_{ij}^r = \Gamma_{ji}^r$ by torsion-free.

2. For the second part, if we permuting i, j, k , we have

$$\Gamma_{jk}^s \Gamma_{is}^r - \Gamma_{ik}^s \Gamma_{js}^r + \Gamma_{ki}^s \Gamma_{js}^r - \Gamma_{ji}^s \Gamma_{ks}^r + \Gamma_{ij}^s \Gamma_{ks}^r - \Gamma_{kj}^s \Gamma_{is}^r = 0$$

by the same reason.

Thus we obtain first Bianchi identity, which is just a consequence of torsion-free.

Remark 4.3.1. If we consider connection on arbitrary vector bundle E , there is no first Bianchi identity, since e_α is just a section of E , not a section of TM , so $R(e_\alpha, \cdot)$ or $R(\cdot, e_\alpha)$ is nonsense.

4.3.2. *Second Bianchi.* In fact, we can write second Bianchi identity for arbitrary vector bundle E as follows

$$\nabla_X R(Y, Z, s, t) + \nabla_Y R(Z, X, s, t) + \nabla_Z R(X, Y, s, t) = 0$$

where $s, t \in C^\infty(M, E)$, $X, Y, Z \in C^\infty(M, TM)$. It's clear that it's equivalent to

$$\nabla_i R_{jk\alpha\beta} + \nabla_j R_{ki\alpha\beta} + \nabla_k R_{ij\alpha\beta} = 0$$

To prove it, here we choose normal coordinate, that is $\nabla_{\frac{\partial}{\partial x^i}} \frac{\partial}{\partial x^j} = 0$. Then

$$\nabla_{\frac{\partial}{\partial x^i}} g(\nabla_{\frac{\partial}{\partial x^j}} \nabla_{\frac{\partial}{\partial x^k}} \frac{\partial}{\partial x^l} - \nabla_{\frac{\partial}{\partial x^k}} \nabla_{\frac{\partial}{\partial x^j}} \frac{\partial}{\partial x^l}, \frac{\partial}{\partial x^m}) = g(\nabla_{\frac{\partial}{\partial x^i}} \nabla_{\frac{\partial}{\partial x^j}} \nabla_{\frac{\partial}{\partial x^k}} \frac{\partial}{\partial x^l} - \nabla_{\frac{\partial}{\partial x^i}} \nabla_{\frac{\partial}{\partial x^k}} \nabla_{\frac{\partial}{\partial x^j}} \frac{\partial}{\partial x^l}, \frac{\partial}{\partial x^m})$$

By permuting i, j, k we have

$$\begin{aligned} & \nabla_{\frac{\partial}{\partial x^i}} \nabla_{\frac{\partial}{\partial x^j}} \nabla_{\frac{\partial}{\partial x^k}} \frac{\partial}{\partial x^l} - \nabla_{\frac{\partial}{\partial x^i}} \nabla_{\frac{\partial}{\partial x^k}} \nabla_{\frac{\partial}{\partial x^j}} \frac{\partial}{\partial x^l} \\ & + \nabla_{\frac{\partial}{\partial x^j}} \nabla_{\frac{\partial}{\partial x^k}} \nabla_{\frac{\partial}{\partial x^i}} \frac{\partial}{\partial x^l} - \nabla_{\frac{\partial}{\partial x^j}} \nabla_{\frac{\partial}{\partial x^i}} \nabla_{\frac{\partial}{\partial x^k}} \frac{\partial}{\partial x^l} \\ & + \nabla_{\frac{\partial}{\partial x^k}} \nabla_{\frac{\partial}{\partial x^i}} \nabla_{\frac{\partial}{\partial x^j}} \frac{\partial}{\partial x^l} - \nabla_{\frac{\partial}{\partial x^k}} \nabla_{\frac{\partial}{\partial x^j}} \nabla_{\frac{\partial}{\partial x^i}} \frac{\partial}{\partial x^l} \\ & = R(\frac{\partial}{\partial x^i}, \frac{\partial}{\partial x^j}) \nabla_{\frac{\partial}{\partial x^k}} \frac{\partial}{\partial x^l} + R(\frac{\partial}{\partial x^j}, \frac{\partial}{\partial x^k}) \nabla_{\frac{\partial}{\partial x^i}} \frac{\partial}{\partial x^l} + R(\frac{\partial}{\partial x^k}, \frac{\partial}{\partial x^i}) \nabla_{\frac{\partial}{\partial x^j}} \frac{\partial}{\partial x^l} \\ & = 0 \end{aligned}$$

This completes the computation of second Bianchi identity.

From another approach, recall that our curvature form Ω is a section of $\Omega_M^2 \otimes \text{End } E$, which can be written as $\Omega_\beta^\alpha e_\alpha \otimes e^\beta$ locally. Then we have $\nabla\Omega$ can be written as

$$\nabla\Omega = d\Omega + \Omega \wedge \omega - \omega \wedge \Omega$$

However, $\nabla\Omega = 0$, since

$$\begin{aligned} \nabla\Omega &= d\Omega + \Omega \wedge \omega - \omega \wedge \Omega \\ &= d(d\omega - \omega \wedge \omega) + (d\omega - \omega \wedge \omega) \wedge \omega - \omega \wedge (d\omega - \omega \wedge \omega) \\ &= d^2\omega - d\omega \wedge \omega + \omega \wedge d\omega + d\omega \wedge \omega - \omega \wedge \omega \wedge \omega - \omega \wedge d\omega + \omega \wedge \omega \wedge \omega \\ &= 0 \end{aligned}$$

If we back to local form, we have

$$d\Omega_\alpha^\beta + \Omega_\alpha^\gamma \wedge \omega_\gamma^\beta - \omega_\alpha^\gamma \wedge \Omega_\gamma^\beta = 0$$

More explicitly, if we write $\Omega_\alpha^\beta = \Omega_{ij\alpha}^\beta dx^i \wedge dx^j$, we obtain

$$(\partial_k \Omega_{ij\alpha}^\beta + \Omega_{ij\alpha}^\gamma \Gamma_{k\gamma}^\beta - \Gamma_{k\alpha}^\gamma \Omega_{ij\gamma}^\beta) dx^k \wedge dx^i \wedge dx^j = 0$$

In other words

$$\begin{aligned} &\partial_k \Omega_{ij\alpha}^\beta + \Omega_{ij\alpha}^\gamma \Gamma_{k\gamma}^\beta - \Gamma_{k\alpha}^\gamma \Omega_{ij\gamma}^\beta \\ &+ \partial_i \Omega_{jk\alpha}^\beta + \Omega_{jk\alpha}^\gamma \Gamma_{i\gamma}^\beta - \Gamma_{i\alpha}^\gamma \Omega_{jk\gamma}^\beta \\ &+ \partial_j \Omega_{ki\alpha}^\beta + \Omega_{ki\alpha}^\gamma \Gamma_{j\gamma}^\beta - \Gamma_{j\alpha}^\gamma \Omega_{ki\gamma}^\beta = 0 \end{aligned}$$

Note that $2\Omega_{ij\alpha}^\beta = R_{ij\alpha}^\beta$, and

$$\nabla_k R_{ij\alpha}^\beta = \partial_k R_{ij\alpha}^\beta + \Gamma_{k\gamma}^\beta R_{ij\alpha}^\gamma - \Gamma_{k\alpha}^\gamma R_{ij\gamma}^\beta$$

So $\nabla\Omega = 0$ locally looks like

$$\nabla_k R_{ij\alpha}^\beta + \nabla_i R_{jk\alpha}^\beta + \nabla_j R_{ki\alpha}^\beta = 0$$

This shows two Bianchi identities are same.

4.4. Other curvatures.

4.4.1. Sectional curvature. Closely related to curvature is sectional curvature that we're going to define, which is used to characterize a two dimensional subspace of tangent space.

Fix $p \in M$ and let x, y are two linearly independent tangent vectors in $T_p M$, then sectional curvature for these two vectors are defined as

$$K_p(x, y) = \frac{R(x, y, y, x)}{g(x, x)g(y, y) - g(x, y)^2}$$

In order to show it's a invariant defined for a two dimensional subspace, we need to check if $\text{span}_{\mathbb{R}}\{x, y\} = \text{span}_{\mathbb{R}}\{z, w\}$, then

$$K_p(x, y) = K_p(z, w)$$

Indeed, if we write

$$\begin{cases} z = ax + by \\ w = cx + dy \end{cases}$$

Then by symmetry and skew symmetry properties of R we have

$$\begin{aligned} R(z, w, w, z) &= R(ax + by, cx + dy, cx + dy, ax + by) \\ &= R(ax, dy, dy, ax) + R(ax, dy, cx, by) + R(by, cx, dy, ax) + R(by, cx, cx, by) \\ &= a^2 d^2 R(x, y, y, x) - abcd R(x, y, y, x) - abcd R(x, y, y, x) + b^2 c^2 R(x, y, y, x) \\ &= (ad - bc)^2 R(x, y, y, x) \end{aligned}$$

And by the same computations we have

$$g(z, z)g(w, w) - g(z, w)^2 = (ad - bc)^2 \{g(x, x)g(y, y) - g(x, y)^2\}$$

Thus

$$K_p(x, y) = K_p(z, w)$$

So the following definition is well-defined:

Definition 4.4.1 (sectional curvature). The sectional curvature $K_p(\sigma)$ for two dimensional subspace $\sigma \subseteq T_p M$ is defined as

$$K_p(\sigma) := K_p(x, y)$$

where $\{x, y\}$ is a basis of σ .

Definition 4.4.2 (isotropic). A Riemannian manifold (M, g) is called isotropic, if for each point $p \in M$, the sectional curvature $K_p(\sigma)$ is independent of σ .

Definition 4.4.3 (constant sectional curvature). A Riemannian manifold (M, g) has constant sectional curvature, if $K_p(\sigma)$ is constant for arbitrary $\sigma \subset T_p M, p \in M$.

Remark 4.4.1. By definition, we can see if a Riemannian manifold has constant sectional curvature, then it must be isotropic; Conversely, we will see if the dimension of a Riemannian manifold ≥ 3 , then isotropic is equivalent to constant sectional curvature.

Lemma 4.4.1.

$$\begin{aligned} -6R(X, Y, Z, W) &= \frac{\partial^2}{\partial s \partial t} \Big|_{s=t=0} \{R(X + sZ, Y + tW, Y + tW, X + sZ) \\ &\quad - R(X + sW, Y + tZ, Y + tZ, X + sW)\} \end{aligned}$$

where X, Y, Z, W are vector fields.

Proof. It suffices to compute coefficients of st of $R(X + sZ, Y + tW, Y + tW, X + sZ)$ and exchange Z with W to obtain coefficients of st of $R(X + sW, Y + tZ, Y + tZ, X + sW)$.

It's easy to see coefficients of st of $R(X + sZ, Y + tW, Y + tW, X + sZ)$ is

$$R(Z, W, Y, X) + R(Z, Y, W, X) + R(X, W, Y, Z) + R(X, Y, W, Z)$$

So coefficients of st of $R(X + sZ, Y + tW, Y + tW, X + sZ)$ is

$$R(W, Z, Y, X) + R(W, Y, Z, X) + R(X, Z, Y, W) + R(X, Y, Z, W)$$

Thus the right hand of our desired identity is

$$-4R(X, Y, Z, W) - (R(Y, Z, W, X) + R(W, Y, Z, X)) - (R(W, X, Y, Z) + R(W, Y, Z, X))$$

By first Bianchi identity we have

$$\begin{aligned} R(Y, Z, W, X) + R(W, Y, Z, X) &= R(Y, Z, W, X) + R(Z, X, W, Y) \\ &= R(X, Y, Z, W) \end{aligned}$$

$$\begin{aligned} R(W, X, Y, Z) + R(W, Y, Z, X) &= R(Y, Z, W, X) + R(Z, X, W, Y) \\ &= R(X, Y, Z, W) \end{aligned}$$

This completes the proof. \square

Notation 4.4.1. For convenience, we use $R_0(X, Y, Z, W)$ to denote

$$R_0(X, Y, Z, W) = g(X, W)g(Y, Z) - g(X, Z)g(Y, W)$$

where X, Y, Z, W are vector fields. Then we can write sectional curvature as

$$K_p(\sigma) = \frac{R(x, y, y, x)}{R_0(x, y, y, x)}$$

where $\sigma \subset T_p M$ is spanned by x, y .

Proposition 4.4.1. A Riemannian manifold has constant sectional curvature K_p at point $p \in M$ if and only if $R = K_p R_0$, where K_p is a constant (may depend on p), R is curvature tensor.

Proof. If $R = K_p R_0$, then for a arbitrary x, y , we have

$$K_p(x, y) = \frac{R(x, y, y, x)}{R_0(x, y, y, x)} = K_p$$

Conversely, if $K(\sigma)$ is constant at point $p \in M$, that is for arbitrary x, y we have

$$\frac{R(x, y, y, x)}{R_0(x, y, y, x)} = K_p$$

If we denote

$$\begin{aligned} F(s, t) &= R(x + sz, y + tw, y + tw, x + sz) - R(x + sw, y + tz, y + tz, x + sw) \\ F_0(s, t) &= R_0(x + sz, y + tw, y + tw, x + sz) - R_0(x + sw, y + tz, y + tz, x + sw) \end{aligned}$$

we still have $F(s, t) = K_p F_0(s, t)$. By Lemma 4.4.1, we have

$$R(x, y, z, w) = -\frac{1}{6} \frac{\partial^2}{\partial s \partial t} \Big|_{s=t=0} F(s, t)$$

and it's easy to see

$$R_0(x, y, z, w) = -\frac{1}{6} \frac{\partial^2}{\partial s \partial t} \Big|_{s=t=0} F_0(s, t)$$

This completes the proof. \square

Corollary 4.4.1. A Riemannian manifold is isotropic if and only if $R = KR_0$, where K is a smooth function.

Corollary 4.4.2. A Riemannian manifold has constant sectional curvature K if and only if $R = KR_0$, where K is a constant.

Remark 4.4.2. An important corollary is that curvature tensor of Riemannian manifold with constant sectional curvature K is quite simple, since

$$R_{ijkl} = K(g_{il}g_{jk} - g_{ik}g_{jl})$$

that is, curvature is completely determined by zero order partial derivatives of metric, not two order in general.

Remark 4.4.3. Suppose the dimension of Riemannian manifold (M, g) is 2, and $\{e_1, e_2\}$ is a basis of $T_p M$. Then

$$K_p = K_p(e_1, e_2) = \frac{R(e_1, e_2, e_2, e_1)}{|e_1|^2|e_2|^2 - |g(e_1, e_2)|^2}$$

is exactly Gauss curvature we learnt in theory of surface.

4.4.2. *Ricci curvature and scalar curvature.*

Definition 4.4.4 (Ricci curvature). For a Riemannian manifold (M, g) , the Ricci curvature is defined to be

$$\text{Ric}(X, Y) := \text{tr}_g(Z \mapsto R(Z, X)Y)$$

where X, Y are vector fields.

Remark 4.4.4 (local form). The trace of above endomorphism is exactly R_{ijk}^i , and it can be written as

$$g^{il}R_{ijkl}$$

In other words, Ricci curvature tensor is the contracted tensor of curvature with respect to the first and fourth index.

Definition 4.4.5 (Ricci curvature in one direction). For a point $p \in M$, and $x \in T_p M$, Ricci curvature in the direction x is defined as

$$\text{Ric}_p(x) := \text{Ric}(x, x)$$

Remark 4.4.5. For $x \in T_p M$, we can write it as $x = x^i e_i$, where $\{e_1, \dots, e_n\}$ is a basis of $T_p M$, then

$$\text{Ric}_p(x) = R_{jk} x^j x^k$$

Definition 4.4.6 (scalar curvature). For a Riemannian manifold (M, g) , the scalar curvature S at $p \in M$ is defined as

$$S(p) := \sum_{i=1}^n \text{Ric}_p(e_i)$$

where $\{e_1, \dots, e_n\}$ is an orthonormal basis of $T_p M$.

Remark 4.4.6 (local form). Locally we have

$$S = g^{jk} R_{jk}$$

Proposition 4.4.2 (contracted Bianchi identity).

$$g^{jk} \nabla_k R_{ij} = \frac{1}{2} \nabla_i S$$

where R_{ij} is Ricci curvature and S is scalar curvature.

Proof. Direct computation shows

$$\begin{aligned} g^{jk} \nabla_k R_{ij} &= g^{jk} \nabla_k g^{pq} R_{pijq} \\ &= g^{jk} g^{pq} \nabla_k R_{pijq} \\ &= g^{jk} g^{pq} (-\nabla_p R_{ikjq} - \nabla_i R_{kpjq}) \\ &= -g^{pq} \nabla_p R_{iq} + \nabla_i S \\ &= -g^{jk} \nabla_k R_{ij} + \nabla_i S \end{aligned}$$

This completes the proof. \square

Proposition 4.1. The scalar curvature S at $p \in M$ is given by

$$S(p) = \frac{1}{\alpha_n} \int_{\mathbb{S}^{n-1}} \text{Ric}_p(x) d\mathbb{S}^{n-1}$$

where α_n is the volume of n -dimension unit ball in \mathbb{R}^{n+1} and $d\mathbb{S}^{n-1}$ is the area elements in \mathbb{S}^{n-1} .

Proof. Choose an orthonormal basis $\{e_1, \dots, e_n\}$ in $T_p M$ and write $x = x^i e_i$, then

$$\begin{aligned} \text{Ric}_p(x) &= \text{Ric}_p(x^i e_i) \\ &= (x^i)^2 \text{Ric}_p(e_i) \end{aligned}$$

Since $|x| = 1$, then the vector $\mu = (x^1, \dots, x^n)$ is a unit normal vector on \mathbb{S}^{n-1} . Denoting $V = (x^1 \text{Ric}_p(e_1), \dots, x^n \text{Ric}_p(e_n))$, then Stokes theorem implies

$$\begin{aligned} \frac{1}{\alpha_n} \int_{\mathbb{S}^{n-1}} (x^i)^2 \text{Ric}_p(e_i) d\mathbb{S}^{n-1} &= \frac{1}{\alpha_n} \int_{\mathbb{S}^{n-1}} \langle V, \mu \rangle d\mathbb{S}^{n-1} \\ &= \frac{1}{\alpha_n} \int_{B^n} \text{div } V d\mathbb{B}^n \\ &= \text{div } V \\ &= \sum_{i=1}^n \text{Ric}_p(e_i) \\ &= S(p) \end{aligned}$$

where B^n is unit ball in $T_p M$ with $\partial B^n = \mathbb{S}^{n-1}$. \square

Theorem 4.4.1. Let (M, g) be a Riemannian manifold, then for all $p \in M$ and r sufficiently small, the volume of the geodesic ball $B(p, r)$ is

$$\text{vol}(B(p, r)) = \alpha_n r^n \left(1 - \frac{S(p)}{6(n+2)} r^2 + O(r^3)\right)$$

where α_n is the volume of n -dimension unit ball in \mathbb{R}^{n+1} .

Proof. Note that we have

$$\sqrt{\det(g_{ij})} = \delta_{ij} - \frac{1}{6} R_{jk}(p) x^j x^k + O(|x|^3)$$

Directly computation shows

$$\begin{aligned} \text{Vol}(B(p, r)) &= \int_0^r \int_{\mathbb{S}^{n-1}(t)} \sqrt{\det g} dS dt \\ &= \int_0^r \int_{\mathbb{S}^{n-1}(t)} \left(1 - \frac{1}{6} \text{Ric}_p(x) + O(|x|^3)\right) dS dt \\ &= \alpha_n r^n - \frac{\alpha_n}{6} \int_0^r t^{n+1} dt + O(r^{n+3}) \\ &= \alpha_n r^n - \frac{\alpha_n S(p) r^{n+2}}{6(n+2)} + O(r^{n+3}) \\ &= \alpha_n r^n \left(1 - \frac{S(p)}{6(n+2)} r^2 + O(r^3)\right) \end{aligned}$$

where we use the fact $\alpha_n = \omega_{n-1}/n$. □

4.4.3. Einstein manifold.

Definition 4.4.7 (Einstein manifold). A Riemannian manifold (M, g) is called Einstein manifold, if its Ricci curvature satisfies $R_{ij} = \lambda g_{ij}$ for some $\lambda \in \mathbb{R}$.

Lemma 4.4.2 (Schur's lemma). Let (M, g) be a Riemannian manifold with $\dim M \geq 3$, suppose $R_{ij} = f g_{ij}$, where $f \in C^\infty(M)$, then (M, g) is an Einstein manifold.

Proof. If $R_{ij} = f g_{ij}$, then contracted Bianchi identity shows

$$\begin{aligned} \frac{n}{2} \nabla_i f &= g^{jk} \nabla_k f g_{ij} \\ &= \nabla_i f \end{aligned}$$

for arbitrary i , which implies f is constant, since $n \geq 3$. □

Corollary 4.4.3. For a Riemannian manifold (M, g) with $\dim M \geq 3$, it is isotropic if and only if it has constant sectional curvature.

Proof. By Remark 4.4.2, it suffices to show if M is isotropic then it has constant sectional curvature. If M is isotropic, then there exists a smooth function K such that

$$R_{ijkl} = K(g_{il}g_{jk} - g_{ik}g_{jl})$$

Consider its Ricci curvature, that is

$$R_{jk} = (n-1)Kg_{jk}$$

Then Schur's lemma implies $(n-1)K$ is constant, that is K is constant. \square

Proposition 4.4.3. Let (M, g) be an Einstein manifold of 3-dimension, then (M, g) is of constant sectional curvature.

Proof. At any point $p \in M$, we choose normal basis at this point, that is $g_{ij} = \delta_{ij}$, thus

$$R_{11} = g^{ij}R_{i11j} = R_{2112} + R_{3113} = \lambda$$

Similarly we have

$$\begin{aligned} R_{1221} + R_{3223} &= \lambda \\ R_{1331} + R_{2332} &= \lambda \end{aligned}$$

Thus we can conclude

$$R_{1221} = R_{1331} = R_{2332} = \frac{\lambda}{2}$$

that is

$$R_{ijkl} = \frac{\lambda}{2}(\delta_{il}\delta_{jk} - \delta_{ik}\delta_{jl})$$

\square

Remark 4.4.7. In fact, it's a special case of Ricci curvature controls curvature. For a n -dimensional Riemannian manifold, it's easy to see R_{jk} has $n(n+1)/2$ independent components. But for R_{ijkl} , this counting problem becomes a little bit complicated, it has

$$\frac{n^2(n^2-1)}{12}$$

independent components. Indeed, since R_{ijkl} is skew symmetric in ij and kl , this means that these pair of indices can take

$$m = \binom{n}{2} = \frac{n(n-1)}{2}$$

R_{ijkl} is also symmetric when you swap ij with kl , this means there would be

$$\frac{m(m+1)}{2} = \frac{n^4 - 2n^3 + 3n^2 - 2n}{8}$$

choices. However, these are not independent, since there is first Bianchi identity

$$R_{ijkl} + R_{jkil} + R_{kijl} = 0$$

, and it provides

$$\binom{n}{4} = \frac{n^4 - 6n^3 + 11n^2 - 6n}{24}$$

relations between these components, thus the number of independent components of R_{ijkl} is

$$\frac{n^4 - 2n^3 + 3n^2 - 2n}{8} - \frac{n^4 - 6n^3 + 11n^2 - 6n}{24} = \frac{n^4 - n^2}{12} = \frac{n^2(n^2 - 1)}{12}$$

Therefore curvature is fully determined by the Ricci curvature if and only if

$$\frac{n^2(n^2 - 1)}{12} \leq \frac{n(n + 1)}{2}$$

or in other words, $n \leq 3$.

5. EXAMPLES

Now let's compute some examples of Riemannian manifold to see their curvatures.

Example 5.0.1 (Euclidean space). Riemannian metric on Euclidean space \mathbb{R}^n is given by

$$g = \delta_{ij} dx^i \otimes dx^j$$

Thus $R_{ijkl} = 0, R_{jk} = 0$ and $S = 0$.

Example 5.0.2 (Sphere). Let $\mathbb{S}^n(K)$ denote n -dimensional sphere with radius K . There is a natural inclusion $f : \mathbb{S}^n(K) \hookrightarrow (\mathbb{R}^{n+1}, g_0)$, and we can use f to pullback g_0 to obtain a metric on $\mathbb{S}^n(K)$, denoted by $g = f^*g_0$. Given a local chart (U, φ, x^i) , we can write

$$f(x^1, \dots, x^n) = (x^1, \dots, x^n, \sqrt{K^2 - \sum_{i=1}^n (x^i)^2})$$

For any $\frac{\partial}{\partial x^i}$, we have

$$\begin{aligned} df\left(\frac{\partial}{\partial x^i}\right) &= \frac{\partial f^j}{\partial x^i} \frac{\partial}{\partial x^j} \\ &= \frac{\partial}{\partial x^i} - \frac{x^i}{\sqrt{K^2 - \sum_{i=1}^n (x^i)^2}} \frac{\partial}{\partial x^{n+1}} \end{aligned}$$

Thus for any two $\frac{\partial}{\partial x^i}, \frac{\partial}{\partial x^j}$ we have

$$\begin{aligned} g\left(\frac{\partial}{\partial x^i}, \frac{\partial}{\partial x^j}\right) &= g_0(df\frac{\partial}{\partial x^i}, df\frac{\partial}{\partial x^j}) \\ &= g_0\left(\frac{\partial}{\partial x^i} - \frac{x^i}{\sqrt{K^2 - \sum_{i=1}^n (x^i)^2}} \frac{\partial}{\partial x^{n+1}}, \frac{\partial}{\partial x^j} - \frac{x^j}{\sqrt{K^2 - \sum_{i=1}^n (x^i)^2}} \frac{\partial}{\partial x^{n+1}}\right) \\ &= \delta_{ij} + \frac{x^i x^j}{K^2 - \sum_{i=1}^n (x^i)^2} \end{aligned}$$

which implies

$$g_{ij} = \delta_{ij} + \frac{x^i x^j}{T^2}, \quad T^2 = K^2 - \sum (x^i)^2$$

Thus we have

$$\begin{aligned} g^{ij} &= \delta^{ij} - \frac{x^i x^j}{K^2} \\ \frac{\partial g_{ij}}{\partial x^k} &= \frac{\delta_{ki} x^j + \delta_{kj} x^i}{T^2} + \frac{2x^i x^j x^k}{T^4} \end{aligned}$$

So Christoffel symbol can be computed as

$$\begin{aligned}
\Gamma_{ij}^k &= \sum_l \frac{1}{2} g^{kl} \left(\frac{\partial g_{jl}}{\partial x^i} + \frac{\partial g_{il}}{\partial x^j} - \frac{\partial g_{ij}}{\partial x^l} \right) \\
&= \sum_l \frac{1}{2} \left(\delta^{kl} - \frac{x^k x^l}{K^2} \right) \left(\frac{\delta_{ij} x^l + \delta_{il} x^j}{T^2} + \frac{2x^i x^j x^l}{T^4} + \frac{\delta_{ji} x^l + \delta_{jl} x^i}{T^2} + \frac{2x^i x^j x^l}{T^4} - \frac{\delta_{li} x^j + \delta_{kj} x^i}{T^2} - \frac{2x^i x^j x^l}{T^4} \right) \\
&= \sum_l \frac{x^l}{T^2} \left(\delta_{ij} + \frac{x^i x^j}{T^2} \right) \left(\delta^{kl} - \frac{x^k x^l}{K^2} \right) \\
&= \frac{g_{ij}}{T^2} x^k \left(1 - \frac{\sum_{l=1}^n (x^l)^2}{K^2} \right) \\
&= \frac{x^k}{K^2} g_{ij}
\end{aligned}$$

Thus curvature can be written as⁴

$$\begin{aligned}
R_{ijkl} &= \frac{1}{2} (\partial_i \partial_k g_{jl} + \partial_j \partial_l g_{ik} - \partial_i \partial_l g_{jk} - \partial_j \partial_k g_{il}) + g_{rs} (\Gamma_{ik}^r \Gamma_{jl}^s - \Gamma_{jk}^r \Gamma_{il}^s) \\
&= \frac{1}{K^2} (g_{il} g_{jk} - g_{ik} g_{jl})
\end{aligned}$$

So Ricci curvature and scalar curvature can be computed as follows

$$\begin{aligned}
R_{jk} &= g^{il} R_{ijkl} \\
&= \frac{1}{K^2} g^{il} (g_{il} g_{jk} - g_{ik} g_{jl}) \\
&= \frac{1}{K^2} (n g_{jk} - \delta_k^l g_{jl}) \\
&= \frac{n-1}{K^2} g_{jk} \\
S &= g^{jk} R_{jk} \\
&= \frac{n(n-1)}{K^2}
\end{aligned}$$

Example 5.0.3 (Poincaré disk). Let $\mathbb{B}^n = \{x \in \mathbb{R}^n \mid |x| < 1\}$ with a metric

$$g = \frac{4\delta_{ij} dx^i \otimes dx^j}{(1 - |x|^2)^2}$$

$$R_{ijkl} = -(g_{il} g_{jk} - g_{ik} g_{jl})$$

Three examples we compute above all have constant sectional curvature, in fact we have

Theorem 5.0.1 (Hopf). Let (M, g) be a complete, simply-connected, n -dimensional Riemannian manifold with constant sectional curvature. Then (M, g) is isometric to either \mathbb{R}^n , S^n or \mathbb{B}^n with standard metric.

⁴Here I omit a huge computation, and I suggest you compute it by yourself. Maybe first it's quite tough for you to do this in first time, but you should try.

Part 2. Bochner's technique

6. HODGE THEORY ON RIEMANNIAN MANIFOLD

For convenience, in this section we assume (M, g) is a compact oriented Riemannian n -manifold, since we need to consider integral.

6.1. Inner product on Ω_M^k . Before we talk about Hodge theory on (M, g) , let's recall some basic facts about differential k -forms. For a k -form φ , locally it can be written as

$$\varphi = \sum_{1 \leq i_1 < \dots < i_k \leq n} \varphi_{i_1 \dots i_k} dx^{i_1} \wedge \dots \wedge dx^{i_k}$$

where $\varphi_{i_1 \dots i_k} := \varphi(\frac{\partial}{\partial x^{i_1}}, \dots, \frac{\partial}{\partial x^{i_k}})$, is skew-symmetric. If we don't want our indices are arranged, we can write

$$\varphi = \frac{1}{k!} \varphi_{i_1 \dots i_k} dx^{i_1} \wedge \dots \wedge dx^{i_k}$$

Here we mean summation runs over any arbitrary different k indices. It's clear this two expressions are same, since both $\varphi_{i_1 \dots i_k}$ and $dx^{i_1} \wedge \dots \wedge dx^{i_k}$ are skew-symmetric.

Notation 6.1.1. We always write $\varphi_{i_1 \dots i_k} dx^{i_1} \wedge \dots \wedge dx^{i_k}$ as $\varphi_I dx^I$.

Recall that we already have a induced metric g on $\bigotimes^k T^*M$, and Ω_M^k is a subbundle of $\bigotimes^k T^*M$. Thus we can define a metric on Ω_M^k as follows

Definition 6.1.1. Let φ, ψ be two k -forms, define

$$\langle \varphi, \psi \rangle := \frac{1}{k!} g(\varphi, \psi)$$

where g is induced metric on $\bigotimes^k T^*M$.

Lemma 6.1.1. For $\varphi = \varphi_I dx^I, \psi = \psi_J dx^J$, then

$$\langle \varphi, \psi \rangle = \varphi_I \psi_J g^{IJ}$$

where

$$g^{IJ} = \frac{1}{k!} g(dx^I, dx^J) = \det \begin{pmatrix} g^{i_1 j_1} & \dots & g^{i_1 j_k} \\ \dots & \dots & \dots \\ g^{i_k j_1} & \dots & g^{i_k j_k} \end{pmatrix}$$

Proof. It suffices to compute

$$g(dx^I, dx^J) = k! \det \begin{pmatrix} g^{i_1 j_1} & \dots & g^{i_1 j_k} \\ \dots & \dots & \dots \\ g^{i_k j_1} & \dots & g^{i_k j_k} \end{pmatrix}$$

Indeed, by definition we have

$$dx^I = dx^{i_1} \wedge \dots \wedge dx^{i_k} = \sum_{\sigma \in S_k} (-1)^{|\sigma|} e_{i_{\sigma(1)}} \otimes \dots \otimes e_{i_{\sigma(k)}}$$

Then

$$\begin{aligned}
g(\mathrm{d}x^I, \mathrm{d}x^J) &= \sum_{\sigma, \tau} (-1)^{|\sigma|} (-1)^{|\tau|} g(\mathrm{d}x^{i_{\sigma(1)}} \otimes \cdots \otimes \mathrm{d}x^{i_{\sigma(k)}}, \mathrm{d}x^{j_{\tau(1)}} \otimes \cdots \otimes \mathrm{d}x^{j_{\tau(k)}}) \\
&= \sum_{\sigma, \tau} (-1)^{|\sigma|} (-1)^{|\tau|} g^{i_{\sigma(1)} j_{\tau(1)}} \cdots g^{i_{\sigma(k)} j_{\tau(k)}} \\
&= \sum_{\sigma, \tau} (-1)^{|\sigma\tau^{-1}|} g^{i_{\sigma\tau^{-1}(1)} j_1} \cdots g^{i_{\sigma\tau^{-1}(k)} j_k} \\
&= \sum_{\sigma} \sum_{\rho} (-1)^{|\rho|} g^{i_{\rho(1)} j_1} \cdots g^{i_{\rho(k)} j_k} \\
&= \sum_{\sigma} \det(g^{i_p j_q}) \\
&= k! \det(g^{i_p j_q})
\end{aligned}$$

□

Remark 6.1.1. Note that here we don't assume φ_I, ψ_I is skew-symmetric, they can be arbitrary functions.

Corollary 6.1.1. For two k -forms φ, ψ , locally write them as

$$\begin{aligned}
\varphi &= \sum_{1 \leq i_1 < \cdots < i_k \leq n} \varphi_{i_1 \dots i_k} \mathrm{d}x^{i_1} \wedge \cdots \wedge \mathrm{d}x^{i_k} \\
\psi &= \sum_{1 \leq j_1 < \cdots < j_k \leq n} \psi_{j_1 \dots j_k} \mathrm{d}x^{j_1} \wedge \cdots \wedge \mathrm{d}x^{j_k}
\end{aligned}$$

with φ_I, ψ_J is skew-symmetric, then

$$\langle \varphi, \psi \rangle = \sum_{\substack{1 \leq i_1 < \cdots < i_k \leq n \\ 1 \leq j_1 < \cdots < j_k \leq n}} \varphi_{i_1 \dots i_k} \psi_{j_1 \dots j_k} \det \begin{pmatrix} g^{i_1 j_1} & \cdots & g^{i_1 j_k} \\ \vdots & \ddots & \vdots \\ g^{i_k j_1} & \cdots & g^{i_k j_k} \end{pmatrix}$$

Example 6.1.1. Let φ, ψ be two 2-forms, locally write them as

$$\varphi = \varphi_{i_1 i_2} \mathrm{d}x^{i_1} \wedge \mathrm{d}x^{i_2}, \quad \psi = \psi_{j_1 j_2} \mathrm{d}x^{j_1} \wedge \mathrm{d}x^{j_2}$$

where $i_1 < i_2, j_1 < j_2$. Then

$$\begin{aligned}
\langle \varphi, \psi \rangle &= \frac{1}{2} \varphi_{i_1 i_2} \psi_{j_1 j_2} g(\mathrm{d}x^{i_1} \wedge \mathrm{d}x^{i_2}, \mathrm{d}x^{j_1} \wedge \mathrm{d}x^{j_2}) \\
&= \frac{1}{2} \varphi_{i_1 i_2} \psi_{j_1 j_2} g(\mathrm{d}x^{i_1} \otimes \mathrm{d}x^{i_2} - \mathrm{d}x^{i_2} \otimes \mathrm{d}x^{i_1}, \mathrm{d}x^{j_1} \otimes \mathrm{d}x^{j_2} - \mathrm{d}x^{j_2} \otimes \mathrm{d}x^{j_1}) \\
&= \frac{1}{2} \varphi_{i_1 i_2} \psi_{j_1 j_2} (g^{i_1 j_1} g^{i_2 j_2} - g^{i_1 j_2} g^{i_2 j_1} - g^{i_2 j_1} g^{i_1 j_2} + g^{i_2 j_2} g^{i_1 j_1}) \\
&= \varphi_{i_1 i_2} \psi_{j_1 j_2} (g^{i_1 j_1} g^{i_2 j_2} - g^{i_1 j_2} g^{i_2 j_1}) \\
&= \varphi_{i_1 i_2} \psi_{j_1 j_2} \det \begin{pmatrix} g^{i_1 j_1} & g^{i_1 j_2} \\ g^{i_2 j_1} & g^{i_2 j_2} \end{pmatrix}
\end{aligned}$$

Definition 6.1.2 (volume form). A form vol locally looks like $\sqrt{\det g} dx^1 \wedge \cdots \wedge dx^n$, where $\sqrt{\det g} = \sqrt{\det(g_{ij})}$, is called a volume form.

Proposition 6.1.1.

$$\langle \text{vol}, \text{vol} \rangle = 1$$

Proof. Directly compute

$$\langle \sqrt{\det g} dx^1 \wedge \cdots \wedge dx^n, \sqrt{\det g} dx^1 \wedge \cdots \wedge dx^n \rangle = \det(g_{ij}) \det(g^{ij}) = 1$$

□

Definition 6.1.3 (inner product on Ω_M^k). For two k -forms φ, ψ , their inner product is defined as

$$(\varphi, \psi) := \int_M \langle \varphi, \psi \rangle \text{vol}$$

Definition 6.1.4 (formal adjoint). For a k -form φ and a $(k+1)$ -form ψ , if there exists $d^* : C^\infty(M, \Omega_M^{k+1}) \rightarrow C^\infty(M, \Omega_M^k)$ such that

$$(d\varphi, \psi) = (\varphi, d^*\psi)$$

Then d^* is called formal adjoint of d .

Remark 6.1.2. Above statement is just a formal definition, there is no guarantee for existence, but later we will see such d^* do exists.

Definition 6.1.5 (Laplace-Beltrami operator). The Laplacian operator $\Delta : C^\infty(M, \Omega_M^k) \rightarrow C^\infty(M, \Omega_M^k)$ is defined as

$$\Delta = dd^* + d^*d$$

Definition 6.1.6 (harmonic). A k -form α is called harmonic, if $\Delta\alpha = 0$. The space of all harmonic forms is denoted by $\mathcal{H}^k(M)$.

Lemma 6.1.2. A k -form α is harmonic if and only if $d\alpha = 0$ and $d^*\alpha = 0$.

Proof. Note that

$$\begin{aligned} (\alpha, \Delta\alpha) &= (\alpha, dd^*\alpha) + (\alpha, d^*d\alpha) \\ &= \|d^*\alpha\|^2 + \|d\alpha\|^2 \end{aligned}$$

□

6.2. Hodge star operator. Although we have defined an inner product on Ω_M^k , it's still quite difficult to compute. However, inner product on Ω_M^k is independent of the choice of local basis, so we can use normal coordinate to give a local basis, and define Hodge star operator on it, which will gives us an effective method to compute.

6.2.1. *Baby case.* Recall that for a \mathbb{F} -vector space V with inner product $\langle -, - \rangle$, and $\{e_1, \dots, e_n\}$ is a basis of V . For any $0 \leq k \leq n$, there is a natural basis of $\bigwedge^k V$, consisting of $\{e_I := e_{i_1} \wedge \dots \wedge e_{i_k} \mid 1 \leq i_1 < \dots < i_k \leq n\}$.

Here are two special cases:

1. For $k = 0$, we regard $\bigwedge^0 V$ as base field \mathbb{F} , and $e_I = 1$.
2. For $k = n$, we use vol to denote basis $e_1 \wedge \dots \wedge e_n$.

With respect to this basis, we can write down the induced metric on $\bigwedge^k V$ as

$$\langle e_{i_1} \wedge \dots \wedge e_{i_k}, e_{j_1} \wedge \dots \wedge e_{j_k} \rangle = \det \begin{pmatrix} \langle e_{i_1}, e_{j_1} \rangle & \dots & \langle e_{i_1}, e_{j_k} \rangle \\ \vdots & & \vdots \\ \langle e_{i_k}, e_{j_1} \rangle & \dots & \langle e_{i_k}, e_{j_k} \rangle \end{pmatrix}$$

It's clear if $\{e_1, \dots, e_n\}$ is an orthonormal basis of V , then $\{e_I\}$ is an orthonormal basis of $\bigwedge^k V$. From now on, we assume $\{e_I\}$ is an orthonormal basis of $\bigwedge^k V$.

Definition 6.2.1 (Hodge star). Hodge star operator is defined as

$$\begin{aligned} \star : \bigwedge^k V &\rightarrow \bigwedge^{n-k} V \\ e_I &\mapsto \text{sign}(I, I^c) e_{I^c} \end{aligned}$$

where I^c is $[n] - I = \{i'_1, \dots, i'_{n-k}\}$ and $\text{sign}(I, I^c)$ is the sign of the permutation $(i_1, \dots, i_k, i'_1, \dots, i'_{n-k})$.

Example 6.2.1. It's clear $\star 1 = \text{vol}$ and $\star \text{vol} = 1$.

Proposition 6.2.1.

$$\star^2 = (-1)^{k(n-k)} \text{id}, \quad \text{on } \bigwedge^k V$$

Proof. It suffices to check on basis e_I as follows

$$\begin{aligned} \star^2 e_I &= \star(\text{sign}(I, I^c) e_{I^c}) \\ &= \text{sign}(I, I^c) \text{sign}(I^c, I) e_I \\ &= (-1)^{k(n-k)} e_I \end{aligned}$$

□

Proposition 6.2.2. For $u \in \bigwedge^k V, v \in \bigwedge^{n-k} V$, we have

$$\star(u \wedge v) = (-1)^{k(n-k)} \langle u, \star v \rangle$$

Proof. It suffices to check on basis $e_I = e_{i_1} \wedge \cdots \wedge e_{i_k}, e_J = e_{j_1} \wedge \cdots \wedge e_{j_{n-k}}$. Furthermore, it's clear $e_I \wedge e_J = 0$, if $J \neq I^c$, so we may assume $J = I^c$.

$$\begin{aligned} \star(e_I \wedge e_{I^c}) &= \star(\text{sign}(I, I^c) \text{vol}) \\ &= \text{sign}(I, I^c) \\ \langle e_I, \star e_{I^c} \rangle &= \langle e_I, \text{sign}(I, I^c) e_I \rangle \\ &= \text{sign}(I, I^c) \langle e_I, e_I \rangle \\ &= \text{sign}(I, I^c) \end{aligned}$$

□

Corollary 6.2.1. For $u, v \in \bigwedge^k V$, we have

1. $u \wedge \star v = v \wedge \star u = \langle u, v \rangle \text{vol}$;
2. $\langle \star u, \star v \rangle = \langle u, v \rangle$.

Proof. For (1).

$$\star(u \wedge \star v) = (-1)^{k(n-k)} \langle u, \star^2 v \rangle = \langle u, v \rangle$$

which implies $u \wedge \star v = \langle u, v \rangle \text{vol}$. Since $\langle u, v \rangle = \langle v, u \rangle$, we obtain $u \wedge \star v = v \wedge \star u$.

For (2).

$$\begin{aligned} \langle \star u, \star v \rangle &= (-1)^{k(n-k)} \star(\star u \wedge v) \\ &= (-1)^{2k(n-k)} \star(v \wedge \star u) \\ &= (-1)^{3k(n-k)} \langle v \wedge \star^2 u \rangle \\ &= (-1)^{4k(n-k)} \langle v, u \rangle \\ &= \langle u, v \rangle \end{aligned}$$

□

Remark 6.2.1. Here are two remarks about this corollary:

1. (1) gives us a method to compute inner product, that's why we define Hodge star, some authors also use this property to denote Hodge star operator;
2. (2) implies that Hodge star operator is an isometry between $\bigwedge^k V$ and $\bigwedge^{n-k} V$.

Corollary 6.2.2 (almost self-adjoint). For $u \in \bigwedge^k V, v \in \bigwedge^{n-k} V$, we have

$$\langle u, \star v \rangle = (-1)^{k(n-k)} \langle \star u, v \rangle$$

Proof.

$$\langle u, \star v \rangle = \langle \star u, \star^2 v \rangle = (-1)^{k(n-k)} \langle \star u, v \rangle$$

□

Remark 6.2.2. This corollary implies the adjoint operator of \star is $(-1)^{k(n-k)} \star$, so here I call it almost self-adjoint.

Thanks to parallel transport, locally we always can choose an orthonormal basis $\{\frac{\partial}{\partial x^1}, \dots, \frac{\partial}{\partial x^n}\}$ of TM with dual basis $\{dx^1, \dots, dx^n\}$, which is also an orthonormal local basis of T^*M . So we can define Hodge star operator on Riemannian manifold locally as follows

$$\begin{aligned} \star : \Omega_M^k &\rightarrow \Omega_M^{n-k} \\ v_I dx^I &\mapsto v_I \text{sign}(I, I^c) dx^{I^c} \end{aligned}$$

Theorem 6.2.1. Properties of Hodge star operator:

1. $\star 1 = \text{vol}, \star \text{vol} = 1$;
2. $\star^2 = (-1)^{k(n-k)}$ on Ω_M^k ;
3. If u is a k -form and v a $(n-k)$ -form, then

$$\begin{aligned} \star(u \wedge v) &= (-1)^{k(n-k)} \langle u, \star v \rangle \\ \langle u, \star v \rangle &= (-1)^{k(n-k)} \langle \star u, v \rangle \end{aligned}$$

4. For any two k -forms u, v , then

$$\begin{aligned} u \wedge \star v &= v \wedge \star u = \langle u, v \rangle \text{vol} = \langle v, u \rangle \text{vol} \\ \langle \star u, \star v \rangle &= \langle u, v \rangle \end{aligned}$$

5. $d^* = (-1)^{nk+n+1} \star d \star$ on Ω_M^k

Remark 6.2.3. (4) allows us to give a new expression for inner product (φ, ψ) , where φ, ψ are two k -forms, that is

$$(\varphi, \psi) := \int_M \langle \varphi, \psi \rangle \text{vol} = \int_M \varphi \wedge \star \psi$$

Proof. It suffices to check (5), other cases we have already solved in the case of linear algebra. Take any $(k-1)$ -form α and k -form β , we need to show

$$(d\alpha, \beta) = (\alpha, d^* \beta)$$

that is to show

$$\int_M d\alpha \wedge \star \beta = \int_M \alpha \wedge \star d^* \beta$$

By Stokes theorem and Leibniz rule we have

$$0 = \int_M d(\alpha \wedge \star \beta) = \int_M d\alpha \wedge \star \beta + (-1)^{k-1} \int_M \alpha \wedge d \star \beta$$

Since $\star^2 = (-1)^{(n-k+1)(k-1)}$ on $(n-k+1)$ -forms, then

$$(-1)^{k-1} \int_M \alpha \wedge d \star \beta = (-1)^{k-1+(n-k+1)(k-1)} \int_M \alpha \wedge \star^2 d \star \beta$$

Therefore

$$\begin{aligned}
 (d\alpha, \beta) &= \int_M d\alpha \wedge \star \beta \\
 &= (-1)^{k+(n-k+1)(k-1)} \int_M \alpha \wedge \star \star d \star \beta \\
 &= (-1)^{nk+k+1} \int_M \alpha \wedge \star (\star d \star \beta)
 \end{aligned}$$

which implies

$$d^* \beta = (-1)^{nk+k+1} \star d \star \beta$$

□

6.2.2. General case. Although above definition gives us a neat way to compute Hodge star, it lost information about how does Hodge star depend on our Riemannian metric, and it's fatal when we not only consider computations about linear algebra, but taking derivatives.

Proposition 6.2.3. Given an arbitrary local basis $\{\frac{\partial}{\partial x^1}, \dots, \frac{\partial}{\partial x^n}\}$ of M , then

$$d^* = -g^{ij} \iota_{\frac{\partial}{\partial x^i}} \nabla_{\frac{\partial}{\partial x^j}}$$

6.2.3. Some computations.

Example 6.2.2. For a 1-form ω written as $\omega_i dx^i$ in an orthonormal frame, then

$$\begin{aligned}
 d^* \omega &= - \star d \star (\omega_i dx^i) \\
 &= - \star d \left(\sum_{i=1}^n (-1)^{i-1} \omega_i dx^1 \wedge \dots \widehat{dx^i} \wedge \dots \wedge dx^n \right) \\
 &= - \star \left(\sum_{i=1}^n \frac{\partial \omega_i}{\partial x^i} dx^1 \wedge \dots \wedge dx^n \right) \\
 &= - \sum_{i=1}^n \frac{\partial \omega_i}{\partial x^i}
 \end{aligned}$$

Example 6.2.3. For a n -form ω written as $f \text{ vol}$, then

$$\begin{aligned}
 d^* \omega &= (-1)^n \star d \star (f \text{ vol}) \\
 &= (-1)^n \star df \\
 &= (-1)^n \star \left(\frac{\partial f}{\partial x^i} dx^i \right) \\
 &= \sum_{i=1}^n (-1)^{n+i-1} \frac{\partial f}{\partial x^i} dx^1 \wedge \dots \wedge \widehat{dx^i} \wedge \dots \wedge dx^n
 \end{aligned}$$

Example 6.2.4. For a smooth function f , then

$$\begin{aligned}\Delta f &= (\mathrm{d} \mathrm{d}^* + \mathrm{d}^* \mathrm{d})f \\ &= \mathrm{d}^* \mathrm{d} f \\ &= \mathrm{d}^* \left(\frac{\partial f}{\partial x^i} \mathrm{d} x^i \right) \\ &= - \sum_{i=1}^n \frac{\partial^2 f}{\partial x^i \partial x^i}\end{aligned}$$

So as you can see, Laplace-Beltrami operator differs a sign with ordinary Laplacian.

Definition 6.2.2 (divergence). Given a Riemannian manifold (M, g) . For any vector field X , its divergence $\mathrm{div}(X)$ is defined as $\mathrm{tr} \nabla X$.

Remark 6.2.4 (local form). If we locally write X as $X^i \frac{\partial}{\partial x^i}$, then

$$\nabla X = \nabla_i X^j \mathrm{d} x^i \otimes \frac{\partial}{\partial x^j}$$

Then

$$\mathrm{tr} \nabla X = \nabla_i X^i$$

Lemma 6.2.1 (Jacobi's formula). For a function $(a_{ij}(t))$ valued in $\mathrm{GL}(n, \mathbb{R})$, we have

$$\frac{\mathrm{d}}{\mathrm{d}t} \det(a_{ij}(t)) = \det(a_{ij}(t)) a^{ij}(t) \frac{\mathrm{d}a_{ij}(t)}{\mathrm{d}t}$$

where $(a^{ij}(t))$ is the inverse matrix of $(a_{ij}(t))$.

Proposition 6.2.4.

$$\mathrm{div}(X) \mathrm{vol} = \mathcal{L}_X \mathrm{vol}$$

Proof. Cartan's magic formula shows that

$$\mathcal{L}_X = i_X \circ \mathrm{d} + \mathrm{d} \circ i_X$$

So

$$\begin{aligned}\mathcal{L}_X \mathrm{vol} &= (i_X \circ \mathrm{d} + \mathrm{d} \circ i_X) \mathrm{vol} \\ &= \mathrm{d} \circ i_X \mathrm{vol} \\ &= \mathrm{d}((-1)^{i-1} X^i \sqrt{\det g} \mathrm{d} x^1 \wedge \cdots \wedge \widehat{\mathrm{d} x^i} \wedge \cdots \wedge \mathrm{d} x^n) \\ &= \frac{1}{\sqrt{\det g}} \frac{\partial (X^i \sqrt{\det g})}{\partial x^i} \mathrm{vol} \\ &= \frac{1}{\sqrt{\det g}} \left(\frac{\partial X^i}{\partial x^i} \sqrt{\det g} + X^i \frac{\partial \sqrt{\det g}}{\partial x^i} \right) \mathrm{vol} \\ &= \left(\frac{\partial X^i}{\partial x^i} + X^i \frac{\partial \log \sqrt{\det g}}{\partial x^i} \right) \mathrm{vol} \\ &= \left(\frac{\partial X^i}{\partial x^i} + \frac{1}{2} X^i \frac{\partial \log \det g}{\partial x^i} \right) \mathrm{vol}\end{aligned}$$

Note that Jacobi's formula says

$$\frac{\partial \log \det g}{\partial x^i} = \frac{1}{\det g} \frac{\partial \det g}{\partial x^i} = g^{jk} \frac{\partial g_{jk}}{\partial x^i} = g^{jk} (\Gamma_{ij}^l g_{lk} + \Gamma_{ik}^l g_{jl}) = 2\Gamma_{ij}^j$$

Thus

$$\begin{aligned} \mathcal{L}_X \text{vol} &= \left(\frac{\partial X^i}{\partial x^i} + \frac{1}{2} X^i \frac{\partial \log \det g}{\partial x^i} \right) \text{vol} \\ &= \left(\frac{\partial X^i}{\partial x^i} + \Gamma_{ij}^j X^i \right) \text{vol} \\ &= \left(\frac{\partial X^i}{\partial x^i} + \Gamma_{ij}^i X^j \right) \text{vol} \\ &= \nabla_i X^i \text{vol} \end{aligned}$$

□

Remark 6.2.5. From the proof, we can say there is the following formula for divergence of a vector field X written as $X^i \frac{\partial}{\partial x^i}$, one has

$$\text{div}(X) = \frac{1}{\sqrt{\det g}} \frac{\partial}{\partial x^i} (\sqrt{\det g} X^i)$$

Proposition 6.2.5. If X_ω is the dual vector field of 1-form ω , then

$$d^* \omega = -\text{div}(X_\omega)$$

Proof. It suffices to check under a local orthonormal basis as follows

1. Remark 6.2.5 or direct computation shows

$$\text{div} X_\omega = \sum_{i=1}^n \frac{\partial \omega_i}{\partial x^i}$$

2. Example 6.2.2 implies

$$d^* \omega = -\sum_{i=1}^n \frac{\partial \omega_i}{\partial x^i}$$

□

6.3. Conformal Laplacian. For a smooth function u , according to Proposition 6.2.5 and Remark 6.2.5, we can write Δu as follows

$$\begin{aligned} \Delta u &= d^* du \\ &= -\text{div} \left(g^{ij} \frac{\partial u}{\partial x^i} \frac{\partial}{\partial x^j} \right) \\ &= -\frac{1}{\sqrt{\det g}} \frac{\partial}{\partial x^j} (\sqrt{\det g} g^{ij} \frac{\partial u}{\partial x^i}) \end{aligned}$$

Thus Laplace-Beltrami Δ_g with respect to g is

$$\Delta_g = -\frac{1}{\sqrt{\det g}} \frac{\partial}{\partial x^j} (\sqrt{\det g} g^{ij} \frac{\partial}{\partial x^i})$$

So if we consider conformal transformation $\tilde{g} = e^{2f}g$ for some smooth function f , we have

$$\begin{aligned}\tilde{g}_{ij} &= e^{2f}g_{ij} \\ \tilde{g}^{ij} &= e^{-2f}g^{ij} \\ \det \tilde{g} &= e^{2nf} \det g \\ \sqrt{\tilde{g}} &= e^{nf} \sqrt{\det g}\end{aligned}$$

Thus

$$\begin{aligned}\Delta_{\tilde{g}} &= -\frac{1}{e^{nf}\sqrt{\det g}} \frac{\partial}{\partial x^j} (e^{nf} \sqrt{\det g} e^{-2f} g^{ij} \frac{\partial}{\partial x^i}) \\ &= -\frac{e^{-nf}}{\sqrt{\det g}} \frac{\partial}{\partial x^j} (e^{(n-2)f} \sqrt{\det g} g^{ij} \frac{\partial}{\partial x^i}) \\ &= -\frac{e^{-2f}}{\sqrt{\det g}} \frac{\partial}{\partial x^j} (\sqrt{\det g} g^{ij} \frac{\partial}{\partial x^i}) - \frac{(n-2)e^{-2f}}{\sqrt{\det g}} \frac{\partial f}{\partial x^j} \sqrt{\det g} g^{ij} \frac{\partial}{\partial x^i} \\ &= -e^{-2f} \Delta_g - (n-2)e^{-2f} g^{ij} \frac{\partial f}{\partial x^j} \frac{\partial}{\partial x^i}\end{aligned}$$

So we have

$$\Delta_{\tilde{g}} = -e^{-2f} \Delta_g$$

when $n = 2$. It's a kind of conformal invariance. However this fails in higher dimension. Let's consider the following so-called conformal Laplacian when $n > 3$

$$\begin{aligned}L : C^\infty(M) &\rightarrow C^\infty(M) \\ u &\mapsto -\frac{4(n-1)}{n-2} \Delta_g u + Su\end{aligned}$$

where S is scalar curvature. Let's show

$$\tilde{L}u = e^{-\frac{n+2}{2}f} L(e^{\frac{n-2}{2}f} u)$$

where \tilde{L} is the conformal Laplacian after conformal transformation. Divide computations into several parts:

(1)

$$\begin{aligned}\nabla^2(e^{\frac{n-2}{2}f} u) &= \nabla \left(\frac{n-2}{2} e^{\frac{n-2}{2}f} \frac{\partial f}{\partial x^i} u dx^i + e^{\frac{n-2}{2}f} \frac{\partial u}{\partial x^i} dx^i \right) \\ &= e^{\frac{n-2}{2}f} \nabla^2 u + \frac{n-2}{2} e^{\frac{n-2}{2}f} \frac{\partial f}{\partial x^j} \frac{\partial u}{\partial x^i} dx^i \otimes dx^j \\ &\quad + \left(\frac{(n-2)^2}{4} e^{\frac{n-2}{2}f} u \frac{\partial f}{\partial x^j} \frac{\partial f}{\partial x^i} + \frac{n-2}{2} e^{\frac{n-2}{2}f} \frac{\partial f}{\partial x^i} \frac{\partial u}{\partial x^j} \right) dx^i \otimes dx^j + \frac{n-2}{2} e^{\frac{n-2}{2}f} u \nabla^2 f\end{aligned}$$

(2)

$$\begin{aligned}
\Delta_g(e^{\frac{n-2}{2}f}u) &= \text{tr}_g \nabla^2(e^{\frac{n-2}{2}f}u) \\
&= e^{\frac{n-2}{2}f} \Delta_g u + \frac{n-2}{2} e^{\frac{n-2}{2}f} g^{ij} \frac{\partial f}{\partial x^j} \frac{\partial u}{\partial x^i} \\
&\quad + g^{ij} \left(\frac{(n-2)^2}{4} e^{\frac{n-2}{2}f} u \frac{\partial f}{\partial x^j} \frac{\partial f}{\partial x^i} + \frac{n-2}{2} e^{\frac{n-2}{2}f} \frac{\partial f}{\partial x^i} \frac{\partial u}{\partial x^j} \right) + \frac{n-2}{2} e^{\frac{n-2}{2}f} u \Delta_g f
\end{aligned}$$

(3)

$$\begin{aligned}
e^{-\frac{n+2}{2}f} L(e^{\frac{n-2}{2}f}u) &= -\frac{4(n-1)}{n-2} e^{-2f} \Delta_g u - 4(n-1) e^{-2f} g^{ij} \frac{\partial f}{\partial x^j} \frac{\partial u}{\partial x^i} \\
&\quad - g^{ij} (n-2)(n-1) e^{-2f} u \frac{\partial f}{\partial x^j} \frac{\partial f}{\partial x^i} - 2(n-1) e^{-2f} u \Delta_g f + e^{-2f} S u \\
&= -\frac{4(n-1)}{n-2} e^{-2f} \Delta_g u - 4(n-1) e^{-2f} g^{ij} \frac{\partial f}{\partial x^j} \frac{\partial u}{\partial x^i} \\
&\quad - (n-2)(n-1) e^{-2f} u |df|^2 - 2(n-1) e^{-2f} u \Delta_g f + e^{-2f} S u
\end{aligned}$$

(4)

$$-\frac{4(n-1)}{n-2} \Delta_{\tilde{g}} u = -\frac{4(n-1)}{n-2} e^{-2f} \Delta_g u - 4(n-1) e^{-2f} g^{ij} \frac{\partial f}{\partial x^j} \frac{\partial u}{\partial x^i}$$

(5) Note that

$$\tilde{S} = e^{-2f} S - 2(n-1) e^{-2f} \Delta_g f - (n-2)(n-1) e^{-2f} |df|^2$$

This completes the computation. In particular, in (2) if we take $u = 1$ we have

$$-\frac{4(n-1)}{n-2} \Delta_g(e^{\frac{n-2}{2}f}) = -(n-2)(n-1) e^{\frac{n-2}{2}f} |df|^2 - 2(n-1) e^{\frac{n-2}{2}f} \Delta_g f$$

Thus we have

$$\tilde{S} = e^{-\frac{n+2}{2}f} \left(-\frac{4(n-1)}{n-2} \Delta_g e^{\frac{n-2}{2}f} + S e^{\frac{n-2}{2}f} \right) = e^{-\frac{n+2}{2}f} L(e^{\frac{n-2}{2}f})$$

So if we put $e^{2f} = \varphi^{\frac{4}{n-2}}$, we have

$$\tilde{S} = \varphi^{-\frac{n+2}{n-2}} L\varphi$$

So it's clear g is conformal to \tilde{g} with constant scalar curvature λ if and only if φ is a smooth positive solution to the Yamabe equation

$$L\varphi = \lambda \varphi^{\frac{n+2}{n-2}}$$

6.4. Hodge theorem and corollaries.

Theorem 6.4.1 (Hodge theorem). Consider the Laplace operator $\Delta : C^\infty(M, \Omega_M^k) \rightarrow C^\infty(M, \Omega_M^k)$, then

1. $\dim_{\mathbb{R}} \mathcal{H}^k(M) < \infty$;
2. There is an orthogonal decomposition

$$C^\infty(M, \Omega_M^k) = \mathcal{H}^k(M) \perp \text{im } \Delta$$

Proof. See Appendix A. \square

Corollary 6.4.1. More explicitly, we have the following orthogonal decomposition

$$C^\infty(M, \Omega_M^k) = \mathcal{H}^k(M) \oplus d(C^\infty(M, \Omega_M^{k-1})) \oplus d^*(C^\infty(M, \Omega_M^{k+1}))$$

Proof. It suffices to check $d(C^\infty(M, \Omega_M^{k-1}))$ is orthogonal to $d^*(C^\infty(M, \Omega_M^{k+1}))$. Take $d\alpha$ and $d^*\beta$, where α is a $k-1$ -form and β is a $k+1$ -form. Then

$$(d\alpha, d^*\beta) = (d^2\alpha, \beta) = 0$$

\square

Corollary 6.4.2.

$$\ker d = \mathcal{H}^k(M) \oplus d(C^\infty(M, \Omega_M^{k-1}))$$

$$\ker d^* = \mathcal{H}^k(M) \oplus d^*(C^\infty(M, \Omega_M^{k+1}))$$

Proof. Clear from above corollary. \square

Corollary 6.4.3. The natural map $\mathcal{H}^k(M) \rightarrow H^k(M, \mathbb{R})$ is an isomorphism. In other words, every element in $H^k(M, \mathbb{R})$ is represented by a unique harmonic form.

Proof. Clear from above corollary. \square

Corollary 6.4.4. $\star : \mathcal{H}^k(M) \rightarrow \mathcal{H}^{n-k}(M)$ is an isomorphism.

Proof. It suffices to show \star maps harmonic forms to harmonic forms, since we already have \star maps k -forms to k -forms. By Lemma 6.1.2, we just need to show $d\star\alpha = d^*\star\alpha = 0$ for a harmonic form α . Directly compute as follows

$$d\star\alpha = (-1)^{\bullet_1} \star d\star\alpha = (-1)^{\bullet_2} \star d^*\alpha = 0$$

$$d^*\star\alpha = (-1)^{\bullet_3} \star d\star\alpha = (-1)^{\bullet_4} \star d\alpha = 0$$

Here we use \bullet, \bullet' to denote the power of (-1) , since it's not necessary for us to know what exactly it is. \square

Remark 6.4.1. In fact, above corollary follows from the following identity

$$\Delta\star = \star\Delta$$

which can be directly checked. In other words, Hodge star commutes with Laplacian Δ . Here gives a method of computation: From what we have done in the proof, we will see

$$\star d^* d = (-1)^{\bullet_2} d \star d = (-1)^{\bullet_2 + \bullet_4} d d^* \star$$

$$\star d d^* = (-1)^{\bullet_4} d^* \star d^* = (-1)^{\bullet_2 + \bullet_4} d^* d \star$$

So all we need to do is to figure out the precise number of \bullet_2, \bullet_4 and show that $\bullet_2 + \bullet_4$ is even.

Corollary 6.4.5 (Poincaré duality). $H^k(M, \mathbb{R}) \cong H^{n-k}(M, \mathbb{R})$.

Proof. Clear from Corollary 6.4.3 and Corollary 6.4.4. \square

7. BOCHNER TECHNIQUE

7.1. Bochner formula. Let (M, g) be a Riemannian manifold with Levi-Civita connection ∇ . Given a smooth function $f : M \rightarrow \mathbb{R}$, $\text{Hess } f := \nabla^2 f$ is a $(0, 2)$ -tensor.

Remark 7.1.1 (local form). Locally we have

$$\begin{aligned} \text{Hess } f &= \nabla \left(\frac{\partial f}{\partial x^j} dx^j \right) \\ &= \frac{\partial^2 f}{\partial x^i \partial x^j} dx^i \otimes dx^j + \frac{\partial f}{\partial x^j} \nabla dx^j \\ &= \frac{\partial^2 f}{\partial x^i \partial x^j} dx^i \otimes dx^j - \frac{\partial f}{\partial x^j} \Gamma_{ik}^j dx^i \otimes dx^k \\ &= \left(\frac{\partial^2 f}{\partial x^i \partial x^j} - \Gamma_{ij}^k \frac{\partial f}{\partial x^k} \right) dx^i \otimes dx^j \end{aligned}$$

that is

$$\nabla_i \nabla_j f = \frac{\partial^2 f}{\partial x^i \partial x^j} - \Gamma_{ij}^k \frac{\partial f}{\partial x^k}$$

Definition 7.1.1 (scalar Laplacian).

$$\Delta f = \text{tr}_g(\text{Hess } f) = g^{ij} \nabla_i \nabla_j f$$

Remark 7.1.2. We have already seen that scalar Laplacian differs a sign with Laplace-Beltrami. So we always use Δ to denote scalar Laplacian and Δ_g to denote Laplace-Beltrami operator.

Theorem 7.1.1. Let $f : (M, g) \rightarrow \mathbb{R}$ be a smooth function. Then

1. $p \in M$ is a local minimum(maximum), then $\nabla f(p) = 0$;
2. $p \in M$ is a local minimum, then

$$\begin{cases} \text{Hess } f(p) \geq 0 \\ \Delta f(p) \geq 0 \end{cases}$$

3. $p \in M$ is a local maximum, then

$$\begin{cases} \text{Hess } f(p) \leq 0 \\ \Delta f(p) \leq 0 \end{cases}$$

Proposition 7.1.1 (Bochner formula). Let $f : (M, g) \rightarrow \mathbb{R}$ be a smooth function, then

$$\frac{1}{2} \Delta |\nabla f|^2 = |\text{Hess } f|^2 + \text{Ric}(\nabla f, \nabla f) + g(\nabla \Delta f, \nabla f)$$

Proof. For a tensor, its norm is independent of which type it is, so we write $\nabla f = g^{ij} \nabla_i f \frac{\partial}{\partial x^j}$, then

$$\begin{aligned} |\nabla f|^2 &= g(\nabla f, \nabla f) \\ &= g(g^{ij} \nabla_i f \frac{\partial}{\partial x^j}, g^{kl} \nabla_k f \frac{\partial}{\partial x^l}) \\ &= g^{ij} g^{kl} g_{jl} \nabla_i f \nabla_k f \\ &= g^{ij} \nabla_i f \nabla_j f \end{aligned}$$

By Proposition 2.4.1, that is magic formula, we have

$$\begin{aligned} \frac{1}{2} \Delta |\nabla f|^2 &= \frac{1}{2} g^{kl} \nabla_k \nabla_l (g^{ij} \nabla_i f \nabla_j f) \\ &= \frac{1}{2} g^{kl} g^{ij} \nabla_k \nabla_l (\nabla_i f \nabla_j f) \\ &\stackrel{(1)}{=} g^{kl} g^{ij} \nabla_l \nabla_i f \cdot \nabla_k \nabla_j f + g^{kl} g^{ij} \nabla_k \nabla_l \nabla_i f \cdot \nabla_j f \\ &= |\text{Hess } f|^2 + g^{kl} g^{ij} \nabla_k \nabla_l \nabla_i f \cdot \nabla_j f \end{aligned}$$

Here we need to be careful when we're dealing with (1). Note that

$$\nabla_k \nabla_l (\nabla_i f \nabla_j f) \neq \nabla_{\frac{\partial}{\partial x^k}} \nabla_{\frac{\partial}{\partial x^l}} (\nabla_i f \nabla_j f)$$

since they differ a Christoffel symbol term. However, in above computation we do use this property, since we compute as follows

$$\begin{aligned} \nabla_k \nabla_l (\nabla_i f \nabla_j f) &= \nabla_k (\nabla_l \nabla_i f \cdot \nabla_j f + \nabla_i f \cdot \nabla_l \nabla_j f) \\ &= \nabla_k \nabla_l \nabla_i f \cdot \nabla_j f + \nabla_l \nabla_i f \cdot \nabla_k \nabla_j f \\ &\quad + \nabla_k \nabla_i f \cdot \nabla_l \nabla_j f + \nabla_i f \cdot \nabla_k \nabla_l \nabla_j f \\ &= 2 \nabla_l \nabla_i f \cdot \nabla_k \nabla_j f + 2 \nabla_k \nabla_l \nabla_i f \cdot \nabla_j f \end{aligned}$$

It doesn't matter, since it's a tensor computation, we always can do this in normal coordinate, so this computation still works.

Then the following computation completes the proof:

$$\begin{aligned} g^{kl} g^{ij} \nabla_k \nabla_l \nabla_i f \cdot \nabla_j f &\stackrel{(2)}{=} g^{kl} g^{ij} \nabla_k \nabla_i \nabla_l f \cdot \nabla_j f \\ &\stackrel{(3)}{=} g^{kl} g^{ij} (\nabla_i \nabla_k \nabla_l f - R_{kil}^s \nabla_s f) \cdot \nabla_j f \\ &= g^{ij} \nabla_i (g^{kl} \nabla_k \nabla_l f) \cdot \nabla_j f + g^{ij} R_i^s \nabla_s f \cdot \nabla_j f \\ &= g^{ij} \nabla_i (\Delta f) \cdot \nabla_j f + \text{Ric}(\nabla f, \nabla f) \\ &= g(\nabla \Delta f, \nabla f) + \text{Ric}(\nabla f, \nabla f) \end{aligned}$$

where

1. (2) holds from symmetry of Hessian;
2. (3) holds from Ricci identity.

□

7.2. Obstruction to the existence of Killing fields.

Definition 7.2.1 (Killing field). A vector field X on a Riemannian manifold (M, g) is called a Killing field, if $\mathcal{L}_X g = 0$.

Remark 7.2.1. Since vector field can generate local flows, then X is a Killing field if and only if local flows generated by X acts on M as isometries.

Theorem 7.2.1. The followings are equivalent:

1. X is a Killing field;
2. For any two vector fields Y, Z , we have

$$\langle \nabla_Y X, Z \rangle + \langle Y, \nabla_Z X \rangle = 0$$

Proof. To see (1) is equivalent to (2). Just note that

$$\begin{aligned} \mathcal{L}_X \langle Y, Z \rangle &= X \langle Y, Z \rangle - \langle \mathcal{L}_X Y, Z \rangle - \langle Y, \mathcal{L}_X Z \rangle \\ &= \langle \nabla_X Y, Z \rangle + \langle Y, \nabla_X Z \rangle - \langle [X, Y], Z \rangle - \langle Y, [X, Z] \rangle \\ &= \langle \nabla_Y X, Z \rangle + \langle Y, \nabla_Z X \rangle \end{aligned}$$

□

Remark 7.2.2. For (2) locally we have

$$g_{kj} \nabla_i X^j = -g_{ij} \nabla_k X^j$$

Thus X is a Killing vector if and only if ∇X is a skew-symmetric $(1, 1)$ -tensor, that is $\nabla_i X^j$ is skew-symmetric in i, j .

Corollary 7.2.1. If X is a Killing field, then for arbitrary vector field Y we have

$$\langle \nabla_Y X, Y \rangle = 0$$

Proof. Set $Y = Z$ in $\langle \nabla_Y X, Z \rangle + \langle Y, \nabla_Z X \rangle = 0$. □

Corollary 7.2.2. If X is parallel, that is $\nabla X = 0$, then X is Killing.

Proof. A zero matrix must be skew-symmetric. □

Corollary 7.2.3. If X is Killing, then $\operatorname{div} X = \nabla_i X^i = 0$.

Proof. The trace of a skew-symmetric matrix is zero. □

Remark 7.2.3. In fact, there're too many other properties of Killing fields, which we may prove later, such as

1. Rigidity: For a given point $p \in M$, a Killing field is uniquely determined by X_p and $(\nabla X)_p$;
2. The set of Killing fields is a Lie algebra with dimension $\leq \frac{(n+1)n}{2}$. Furthermore, if (M, g) is complete, it's Lie algebra of isometry group of M .

Lemma 7.2.1 (Bochner formula for Killing field). Let X be a Killing field, and $f = \frac{1}{2}|X|^2$. Then

1. $\nabla f = -\nabla_X X$;

2. $\text{Hess } f(Y, Y) = \langle \nabla_Y X, \nabla_Y X \rangle - R(Y, X, X, Y)$ holds for any vector field Y ;
3. $\Delta f = |\nabla X|^2 - \text{Ric}(X, X)$.

Proof. For (1). By direct computation we have

$$\begin{aligned}
 \nabla f &= \langle \nabla X, X \rangle \\
 &= \langle \nabla_k X^i dx^k \otimes \frac{\partial}{\partial x^i}, X^j \frac{\partial}{\partial x^j} \rangle \\
 &= g_{ij} X^j \nabla_k X^i dx^k \\
 &\stackrel{\text{I}}{=} -g_{ik} X^j \nabla_j X^i dx^k \\
 \nabla_X X &= X^j \nabla_j X^i \frac{\partial}{\partial x^i} \\
 &= g_{ik} X^j \nabla_j X^i dx^k
 \end{aligned}$$

where (I) holds from skew-symmetry of ∇X .

For (2). By direct computation we have

$$\begin{aligned}
 \text{Hess } f(Y, Y) &= \frac{1}{2} Y^i Y^j \nabla_i \nabla_j (g_{kl} X^k X^l) \\
 &= Y^i Y^j g_{kl} (\nabla_i X^k \nabla_j X^l + \nabla_i \nabla_j X^k \cdot X^l) \\
 &= \langle \nabla_Y X, \nabla_Y X \rangle + Y^i Y^j g_{kl} \nabla_i \nabla_j X^k \cdot X^l
 \end{aligned}$$

and

$$\begin{aligned}
 Y^i Y^j g_{kl} \nabla_i \nabla_j X^k \cdot X^l &= -Y^i Y^j g_{kj} \nabla_i \nabla_l X^k \cdot X^l \\
 &\stackrel{\text{II}}{=} -Y^i Y^j g_{kj} X^l (\nabla_l \nabla_i X^k + R_{ilm}^k X^m) \\
 &= -Y^i Y^j X^l X^m R_{ilmj} \\
 &= -R(Y, X, X, Y)
 \end{aligned}$$

where (II) holds from $g_{kj} X^l \nabla_l \nabla_i X^k = 0$, since this expression is skew symmetric in i, j .

(3) holds from (2) directly. \square

Theorem 7.2.2 (Bochner). Let (M, g) be a compact, oriented Riemannian manifold,

1. If $\text{Ric}(g) \leq 0$, then every Killing field is parallel;
2. If $\text{Ric}(g) \leq 0$ and $\text{Ric}(g) < 0$ at some point, then there is no non-trivial Killing field.

Proof. For (1). Let X be a Killing field and set $f = \frac{1}{2}|X|^2$, then

$$\begin{aligned} 0 &= \int_M \Delta f \, \text{vol} \\ &= \int_M (|\nabla X|^2 - \text{Ric}(X, X)) \, \text{vol} \\ &\geq \int_M |\nabla X|^2 \, \text{vol} \\ &\geq 0 \end{aligned}$$

Thus $|\nabla X| \equiv 0$, that is X is parallel.

For (2). From proof of (1) one can see if $\text{Ric}(g) \leq 0$ and X is a Killing field, then

$$\int_M \text{Ric}(X, X) = 0$$

which implies $\text{Ric}(X, X) \equiv 0$. So if $\text{Ric}(g) < 0$ at some point $p \in M$, then $X_p = 0$, thus $X \equiv 0$, since it's parallel. \square

7.3. Obstruction to the existence of harmonic 1-forms. To some extent, Killing field is dual to harmonic 1-form. Let's explain this in more detail.

Lemma 7.3.1. For a harmonic 1-form α , locally written as $\alpha_i dx^i$, we have

$$\begin{aligned} \nabla_i \alpha_j &= \nabla_j \alpha_i \\ g^{ij} \nabla_j \alpha_i &= 0 \end{aligned}$$

Proof. Recall α is harmonic if and only if

$$\begin{aligned} d\alpha &= 0 \\ d^* \alpha &= 0 \end{aligned}$$

It's clear

$$d(\alpha_j dx^j) = \nabla_i \alpha_j dx^i \wedge dx^j = 0$$

implies $\nabla_i \alpha_j = \nabla_j \alpha_i$. Similarly explicit expression for d^* implies the second identity. \square

Remark 7.3.1. Recall Killing field implies $g_{ij} \nabla_k X^j$ is skew-symmetric in i, k , we can see both Killing field and harmonic 1-form implies some (skew)symmetries.

Lemma 7.3.2. If α is a harmonic 1-form, then

$$\frac{1}{2} \Delta |\alpha|^2 = |\nabla \alpha|^2 + \text{Ric}(X_\alpha, X_\alpha)$$

where X_α is the dual vector field of α .

Proof. Routine computation as follows:

$$\begin{aligned}
\frac{1}{2}\Delta|\alpha|_g^2 &= \frac{1}{2}g^{kl}\nabla_k\nabla_l(g^{ij}\alpha_i\alpha_j) \\
&= |\nabla\alpha|^2 + g^{kl}g^{ij}\nabla_k\nabla_l\alpha_i \cdot \alpha_j \\
&= |\nabla\alpha|^2 + g^{kl}g^{ij}\nabla_k\nabla_i\alpha_l \cdot \alpha_j \\
&= |\nabla\alpha|^2 + g^{kl}g^{ij}(\nabla_i\nabla_k\alpha_l - R_{kil}^s\alpha_s)\alpha_j \\
&= |\Delta\alpha|^2 - g^{kl}g^{ij}R_{kil}^s\alpha_s \cdot \alpha_j \\
&= |\Delta\alpha|^2 + \text{Ric}(X_\alpha, X_\alpha)
\end{aligned}$$

□

Theorem 7.3.1 (Bochner). Let (M, g) be a compact, oriented Riemannian manifold,

1. If $\text{Ric}(g) \geq 0$, then every harmonic 1-form is parallel;
2. If $\text{Ric}(g) \geq 0$ and $\text{Ric}(g) > 0$ at some point, then there is no non-trivial harmonic 1-form.

Proof. The same as before. □

Corollary 7.3.1. Let (M, g) be a compact, oriented Riemannian manifold with $\text{Ric}(g) \geq 0$ and $\text{Ric}(g) > 0$ at some point, then $b_1(M) = 0$.

Proof. It's clear from above theorem and Corollary 6.4.3. □

Remark 7.3.2. It's a kind of vanishing theorem. In geometry, “positive” may cause “vanishing”, that's a philosophy.

Part 3. Variation formulas

8. GEODESIC II: VARIATION FORMULAS

In this section, we fix the following notations:

1. $I = [a, b] \subset \mathbb{R}$ is a closed interval;
2. For two different points $p, q \in M$, where (M, g) is a Riemannian manifold, the space of smooth curves from p to q is denoted as

$$\mathcal{L}_{p,q} = \{\text{smooth curve } \gamma : [a, b] \rightarrow M, \text{ with } \gamma(a) = p, \gamma(b) = q\}$$

3. For $\gamma \in \mathcal{L}_{p,q}$, we define $\gamma'(t) := \gamma_*\left(\frac{d}{dt}\right) \in C^\infty(I, \gamma^*TM)$. Note that γ^*TM is equipped with pullback connection $\widehat{\nabla}$ and pullback metric \widehat{g} .
4. Consider the following functionals defined on $\mathcal{L}_{p,q}$:

$$L(\gamma) = \int_a^b |\gamma'(t)| dt$$

$$E(\gamma) = \frac{1}{2} \int_a^b |\gamma'(t)|^2 dt$$

The former is called arc-length functional and the latter is called energy functional.

8.1. First variation formula.

Definition 8.1.1 (variation). Given $\gamma \in \mathcal{L}_{p,q}$, a variation of γ is a smooth map

$$\alpha : [a, b] \times (-\varepsilon, \varepsilon) \rightarrow M$$

such that

1. $\alpha(\cdot, s) \in \mathcal{L}_{p,q}$ for any $s \in (-\varepsilon, \varepsilon)$;
2. $\alpha(t, 0) = \gamma(t)$ for any $t \in [a, b]$.

Remark 8.1.1. For pullback bundle α^*TM , we use $\overline{\nabla}$ and \overline{g} to denote connection and metric pulled back from the ones on TM . By definition we have the restriction of $\overline{\nabla}$ on γ^*TM is exactly $\widehat{\nabla}$, and the restriction of \overline{g} on γ^*TM is \widehat{g} .

Definition 8.1.2 (variation vector field). For a variation α of $\gamma \in \mathcal{L}_{p,q}$, $\alpha_*\left(\frac{\partial}{\partial s}\right)\big|_{s=0} \in C^\infty(I, \gamma^*TM)$ is called variation vector field of variation α .

Remark 8.1.2. Note that

$$\begin{cases} \alpha(a, s) = p \\ \alpha(b, s) = q \end{cases}$$

for any $s \in (-\varepsilon, \varepsilon)$. Thus we have,

$$\begin{cases} \alpha_*\left(\frac{\partial}{\partial s}\right)(a, s) = 0 \\ \alpha_*\left(\frac{\partial}{\partial s}\right)(b, s) = 0 \end{cases}$$

for any $s \in (-\varepsilon, \varepsilon)$. In particular it holds for $s = 0$, that's a variation vector field vanishes at endpoints, a crucial fact we need in later computation.

Lemma 8.1.1. Let X be a smooth vector field along γ with $X(a) = X(b) = 0$. Then there exists a variation α of γ such that the variation vector field is exactly X , that is

$$\alpha_*\left(\frac{\partial}{\partial s}\right)\Big|_{s=0} = X$$

Proof. See Proposition 2.2 in Page193 of [Car92]. \square

Remark 8.1.3. Thanks to this technical lemma, we always call a vector field along γ a variation vector field, if it it vanishes at endpoints.

Theorem 8.1.1 (First variation formula). Let $\gamma : [a, b] \rightarrow (M, g)$ be a unit-speed curve, α a normal variation of γ and V the variation vector field. Then

$$\begin{aligned} \frac{d}{ds}\Big|_{s=0} L(\alpha(-, s)) &\stackrel{(1)}{=} \frac{d}{ds}\Big|_{s=0} E(\alpha(-, s)) \stackrel{(2)}{=} \int_a^b \langle \widehat{\nabla}_{\frac{d}{dt}} V, \gamma'(t) \rangle dt \\ &\stackrel{(3)}{=} - \int_a^b \langle V, \widehat{\nabla}_{\frac{d}{dt}} \gamma'(t) \rangle dt \end{aligned}$$

Proof. Note that

$$\begin{aligned} \frac{d}{ds}\Big|_{s=0} L(\alpha(-, s)) &= \int_a^b \frac{1}{2|\gamma'(t)|} \frac{\partial}{\partial s}\Big|_{s=0} |\alpha_*\left(\frac{\partial}{\partial t}\right)|^2 dt \\ &= \frac{1}{|\gamma'(t)|} \frac{d}{ds}\Big|_{s=0} E(\alpha(-, s)) \end{aligned}$$

Since γ is unit-speed, this show equality marked by (1).

Note that

$$\begin{aligned} 0 &= \int_a^b \frac{d}{dt} \langle V, \gamma'(t) \rangle dt \\ &= \int_a^b \langle \widehat{\nabla}_{\frac{d}{dt}} V, \gamma'(t) \rangle dt + \int_a^b \langle V, \widehat{\nabla}_{\frac{d}{dt}} \gamma'(t) \rangle dt \end{aligned}$$

This shows the equality marked by (3).

For equality marked by (2), we compute as follows

$$\begin{aligned}
\frac{d}{ds} E(\alpha(-, s)) &= \frac{d}{ds} \frac{1}{2} \int_a^b |\alpha_*(\frac{\partial}{\partial t})|^2 dt \\
&= \frac{1}{2} \int_a^b \frac{\partial}{\partial s} |\alpha_*(\frac{\partial}{\partial t})|^2 dt \\
&= \frac{1}{2} \int_a^b 2 \langle \bar{\nabla}_{\frac{\partial}{\partial s}} \alpha_*(\frac{\partial}{\partial t}), \alpha_*(\frac{\partial}{\partial t}) \rangle_{\bar{g}} dt \\
&\stackrel{*}{=} \int_a^b \langle \bar{\nabla}_{\frac{\partial}{\partial t}} \alpha_*(\frac{\partial}{\partial s}), \alpha_*(\frac{\partial}{\partial t}) \rangle_{\bar{g}} dt \\
&= \int_a^b \frac{\partial}{\partial t} \langle \alpha_*(\frac{\partial}{\partial s}), \alpha_*(\frac{\partial}{\partial t}) \rangle_{\bar{g}} - \langle \alpha_*(\frac{\partial}{\partial s}), \bar{\nabla}_{\frac{\partial}{\partial t}} \alpha_*(\frac{\partial}{\partial t}) \rangle_{\bar{g}} dt \\
&= - \int_a^b \langle \alpha_*(\frac{\partial}{\partial s}), \bar{\nabla}_{\frac{\partial}{\partial t}} \alpha_*(\frac{\partial}{\partial t}) \rangle_{\bar{g}} dt
\end{aligned}$$

The hallmark of above computation is the equality marked by star, which can be seen from follows

$$\bar{\nabla}_{\frac{\partial}{\partial s}} \alpha_*(\frac{\partial}{\partial t}) = B(\frac{\partial}{\partial s}, \frac{\partial}{\partial t}) + \alpha_*(\nabla_{\frac{\partial}{\partial s}} \frac{\partial}{\partial t})$$

where B is second fundamental form, and it's symmetric, which can be seen in Appendix B. Thus

$$\frac{d}{ds} \Big|_{s=0} E(\alpha(-, s)) = - \int_a^b \langle V, \hat{\nabla}_{\frac{d}{dt}} \gamma'(t) \rangle_{\hat{g}} dt$$

since $\alpha_*(\frac{\partial}{\partial s})|_{s=0} = V$ and $\alpha_*(\frac{\partial}{\partial t})|_{s=0} = \gamma'(t)$. \square

Corollary 8.1.1. Given $\gamma \in \mathcal{L}_{p,q}$. The followings are equivalent:

1. γ is a critical point of energy functional $E : \mathcal{L}_{p,q} \rightarrow \mathbb{R}$;
2. γ has constant speed $|\gamma'(t)| = c > 0$ and γ is a critical point of arc-length functional $L : \mathcal{L}_{p,q} \rightarrow \mathbb{R}$;
3. γ is a geodesic.

Proof. From (3) to (2): Firstly a geodesic must have constant speed c , and $c > 0$ since p, q are distinct points. It's also a critical point of L since first variation formula implies

$$\frac{d}{ds} \Big|_{s=0} L(\alpha(-, s)) = - \int_a^b \langle V, \hat{\nabla}_{\frac{d}{dt}} \gamma'(t) \rangle_{\hat{g}} dt = 0$$

From (2) to (1): It's clear, since from above proof we have already seen for constant speed curve, the first variation of arc-length functional and energy functional only differs a scalar.

From (1) to (3): In order to show $\hat{\nabla}_{\frac{d}{dt}} \gamma'(t) = 0$, the key point is to choose an appropriate variation vector field V to conclude. \square

8.2. Second variation formula. We already know a geodesic γ is a critical point for energy functional or arc-length functional, so it left to determine whether it's local minimum or not.

To see this, we need to consider the following 2-dimensional variation

$$\alpha : [a, b] \times (-\varepsilon_1, \varepsilon_1) \times (-\varepsilon_2, \varepsilon_2)$$

such that

1. $\alpha(t, 0, 0) = \gamma(t)$
2. $\alpha(-, s_1, s_2) \in \mathcal{L}_{p,q}$

8.2.1. *Second variation formula for energy.*

Theorem 8.2.1 (second variation formula for energy). Let $\gamma : [a, b] \rightarrow (M, g)$ be a smooth curve. If α is a 2-dimensional variation of γ with variation fields V, W . Then

$$\begin{aligned} \frac{\partial^2}{\partial s_1 \partial s_2} \Big|_{s_1=s_2=0} E(\alpha(-, s_1, s_2)) &= \int_a^b \langle \widehat{\nabla}_{\frac{d}{dt}} V, \widehat{\nabla}_{\frac{d}{dt}} W \rangle dt \\ &\quad - \int_a^b R(V, \gamma', \gamma', W) dt - \int_a^b \langle \overline{\nabla}_{\frac{\partial}{\partial s_1}} \alpha_* \left(\frac{\partial}{\partial s_2} \right) \Big|_{s_1=s_2=0}, \widehat{\nabla}_{\frac{d}{dt}} \gamma'(t) \rangle dt \end{aligned}$$

Proof. By first variation formula we have

$$\frac{\partial}{\partial s_2} E(\alpha(-, s_1, s_2)) = - \int_a^b \langle \alpha_* \left(\frac{\partial}{\partial s_2} \right), \overline{\nabla}_{\frac{\partial}{\partial t}} \alpha_* \left(\frac{\partial}{\partial t} \right) \rangle_{\overline{g}} dt$$

Thus

$$\frac{\partial^2}{\partial s_1 \partial s_2} E(\alpha(-, s_1, s_2)) = \underbrace{- \int_a^b \langle \overline{\nabla}_{\frac{\partial}{\partial s_1}} \alpha_* \left(\frac{\partial}{\partial s_2} \right), \overline{\nabla}_{\frac{\partial}{\partial t}} \alpha_* \left(\frac{\partial}{\partial t} \right) \rangle_{\overline{g}} dt}_{\text{part I}} - \underbrace{\int_a^b \langle \alpha_* \left(\frac{\partial}{\partial s_2} \right), \overline{\nabla}_{\frac{\partial}{\partial s_1}} \overline{\nabla}_{\frac{\partial}{\partial t}} \alpha_* \left(\frac{\partial}{\partial t} \right) \rangle_{\overline{g}} dt}_{\text{part II}}$$

For part II, we have

$$\overline{\nabla}_{\frac{\partial}{\partial s_1}} \overline{\nabla}_{\frac{\partial}{\partial t}} \alpha_* \left(\frac{\partial}{\partial t} \right) = R(\alpha_* \left(\frac{\partial}{\partial s_1} \right), \alpha_* \left(\frac{\partial}{\partial t} \right)) \alpha_* \left(\frac{\partial}{\partial t} \right) + \overline{\nabla}_{\frac{\partial}{\partial t}} \overline{\nabla}_{\frac{\partial}{\partial s_1}} \alpha_* \left(\frac{\partial}{\partial t} \right)$$

Thus we can write part II as

$$\begin{aligned} &- \int_a^b \langle \alpha_* \left(\frac{\partial}{\partial s_2} \right), R \left(\frac{\partial}{\partial s_1}, \frac{\partial}{\partial t} \right) \alpha_* \left(\frac{\partial}{\partial t} \right) \rangle_{\overline{g}} dt - \underbrace{\int_a^b \langle \alpha_* \left(\frac{\partial}{\partial s_2} \right), \overline{\nabla}_{\frac{\partial}{\partial t}} \overline{\nabla}_{\frac{\partial}{\partial s_1}} \alpha_* \left(\frac{\partial}{\partial t} \right) \rangle_{\overline{g}} dt}_{\text{part III}} \end{aligned}$$

For part III, we have

$$\begin{aligned} &- \int_a^b \langle \alpha_* \left(\frac{\partial}{\partial s_2} \right), \overline{\nabla}_{\frac{\partial}{\partial t}} \overline{\nabla}_{\frac{\partial}{\partial s_1}} \alpha_* \left(\frac{\partial}{\partial t} \right) \rangle_{\overline{g}} dt = - \int_a^b \langle \alpha_* \left(\frac{\partial}{\partial s_2} \right), \overline{\nabla}_{\frac{\partial}{\partial t}} \overline{\nabla}_{\frac{\partial}{\partial t}} \alpha_* \left(\frac{\partial}{\partial s_1} \right) \rangle_{\overline{g}} dt \\ &= \int_a^b \langle \overline{\nabla}_{\frac{\partial}{\partial t}} \alpha_* \left(\frac{\partial}{\partial s_2} \right), \overline{\nabla}_{\frac{\partial}{\partial t}} \alpha_* \left(\frac{\partial}{\partial s_1} \right) \rangle_{\overline{g}} dt \end{aligned}$$

Now let's evaluate at $s_1 = s_2 = 0$, then we have

1. Part I

$$- \int_a^b \left\langle \bar{\nabla}_{\frac{\partial}{\partial s_1}} \alpha_* \left(\frac{\partial}{\partial s_2} \right) \Big|_{s_1=s_2=0}, \widehat{\nabla}_{\frac{d}{dt}} \gamma'(t) \right\rangle dt$$

2. Part II

$$\int_a^b \langle \widehat{\nabla}_{\frac{d}{dt}} V, \widehat{\nabla}_{\frac{d}{dt}} W \rangle dt - \int_a^b R(V, \gamma', \gamma', W) dt$$

This completes the proof. \square

Corollary 8.2.1. Let $\gamma : [a, b] \rightarrow (M, g)$ be a geodesic, then

$$\frac{\partial^2}{\partial s_1 \partial s_2} \Big|_{s_1=s_2=0} E(\alpha(-, s_1, s_2)) = \int_a^b \langle \widehat{\nabla}_{\frac{d}{dt}} V, \widehat{\nabla}_{\frac{d}{dt}} W \rangle dt - \int_a^b R(V, \gamma', \gamma', W) dt$$

8.2.2. *Second variation formula for arc-length.*

Theorem 8.2.2 (second variation formula for arc-length). Let $\gamma : [a, b] \rightarrow (M, g)$ be a unit-speed curve. If α is a 2-dimensional variation of γ with variation fields V, W . Then

$$\begin{aligned} \frac{\partial^2}{\partial s_1 \partial s_2} \Big|_{s_1=s_2=0} L(\alpha(-, s_1, s_2)) &= \int_a^b \langle \widehat{\nabla}_{\frac{d}{dt}} V, \widehat{\nabla}_{\frac{d}{dt}} W \rangle dt - \int_a^b R(V, \gamma', \gamma', W) dt \\ &\quad - \int_a^b \left\langle \bar{\nabla}_{\frac{\partial}{\partial s_1}} \alpha_* \left(\frac{\partial}{\partial s_2} \right) \Big|_{s_1=s_2=0}, \widehat{\nabla}_{\frac{d}{dt}} \gamma'(t) \right\rangle dt \\ &\quad - \int_a^b \langle \widehat{\nabla}_{\frac{d}{dt}} V, \gamma' \rangle \langle \widehat{\nabla}_{\frac{d}{dt}} W, \gamma' \rangle dt \end{aligned}$$

Corollary 8.2.2. Let $\gamma : [a, b] \rightarrow (M, g)$ be a unit-speed geodesic. If α is a 2-dimensional variation of γ with variation fields V, W . Then

$$\begin{aligned} \frac{\partial^2}{\partial s_1 \partial s_2} \Big|_{s_1=s_2=0} L(\alpha(-, s_1, s_2)) &= \int_a^b \langle \widehat{\nabla}_{\frac{d}{dt}} V, \widehat{\nabla}_{\frac{d}{dt}} W \rangle dt - \int_a^b R(V, \gamma', \gamma', W) dt \\ &\quad - \int_a^b \langle \widehat{\nabla}_{\frac{d}{dt}} V, \gamma' \rangle \langle \widehat{\nabla}_{\frac{d}{dt}} W, \gamma' \rangle dt \\ &= \int_a^b \langle \widehat{\nabla}_{\frac{d}{dt}} V^\perp, \widehat{\nabla}_{\frac{d}{dt}} W^\perp \rangle dt - \int_a^b R(V^\perp, \gamma', \gamma', W^\perp) dt \end{aligned}$$

where

$$V^\perp = V - \langle V, \gamma' \rangle \gamma', \quad W^\perp = W - \langle W, \gamma' \rangle \gamma'$$

Remark 8.2.1. So if we want to show a geodesic γ is a (locally) minimal geodesic, it suffices to check for any 2-dimensional variation α with variation vector fields V, W , we have

$$\frac{\partial^2}{\partial s_1 \partial s_2} \Big|_{s_1=s_2=0} L(\alpha(-, s_1, s_2)) \geq 0$$

Definition 8.2.1 (index form). Suppose $\gamma : [a, b] \rightarrow (M, g)$ is a unit-speed geodesic. The index form I_γ is defined as

$$I_\gamma(V, W) = \int_a^b \langle \widehat{\nabla}_{\frac{d}{dt}} V, \widehat{\nabla}_{\frac{d}{dt}} W \rangle dt - \int_a^b R(V, \gamma', \gamma', W) dt$$

where V, W are vector fields along γ .

Thus a geodesic γ is locally minimal if and only if for any variation fields V, W we have index form $I_\gamma(V, W) \geq 0$. However, it's clear

$$I_\gamma(\gamma', \gamma') = 0$$

So if we want to obtain a kind of positive-definite property of index form, we must consider index form defined on normal vector fields along γ .

Definition 8.2.2 (normal vector field). Let $\gamma : [a, b] \rightarrow (M, g)$ be a geodesic, a vector field V along γ is called normal, if V is perpendicular to γ' .

In the following section, we will study when the index form defined on the normal vector fields along γ is positive-definite, semipositive-definite or not.

9. JACOBI FIELDS

9.1. First properties.

Definition 9.1.1 (Jacobi field). A vector field J along geodesic γ is called a Jacobi field, if it satisfies

$$\widehat{\nabla}_{\frac{d}{dt}} \widehat{\nabla}_{\frac{d}{dt}} J + R(J, \gamma') \gamma' = 0$$

Notation 9.1.1. For convenience, we sometimes use the following notations

$$\begin{aligned} J' &= \widehat{\nabla}_{\frac{d}{dt}} J \\ J'' &= \widehat{\nabla}_{\frac{d}{dt}} \widehat{\nabla}_{\frac{d}{dt}} J \end{aligned}$$

Remark 9.1.1 (local form). If we choose a parallel orthonormal vector fields $\{e_1, \dots, e_n\}$ along γ and write $J(t) = J^i(t)e_i(t)$, the condition for Jacobi fields becomes

$$\frac{d^2 J^k}{dt^2} + \langle J^j R(e_j, \gamma') \gamma', e_k \rangle = 0$$

Thus by standard results in ODEs, a Jacobi field J is completely determined by its initial conditions

$$J(0), J'(0) \in T_{\gamma(0)}M$$

Furthermore, you can see the set of Jacobi fields is a vector space with dimension $2n$.

Example 9.1.1. There is always a trivial Jacobi field along geodesic $\gamma : [a, b] \rightarrow (M, g)$, that is $J(t) = (t - a)\gamma'(t)$.

On a general Riemannian manifold, we can write down all Jacobi fields by using the following construction. However, we're more interested in Jacobi fields vanishes at one endpoint, let's write down an explicit construction for this case.

Lemma 9.1.1. Let $\gamma : [a, b] \rightarrow (M, g)$ be a geodesic and $\alpha : [a, b] \times (-\varepsilon, \varepsilon) \rightarrow (M, g)$ a variation consisting of geodesics of γ , then

$$J = \alpha_* \left(\frac{\partial}{\partial s} \right) \Big|_{s=0} \in \gamma^* TM$$

is a Jacobi field.

Proof. Note that

$$\begin{aligned} \nabla_{\frac{\partial}{\partial t}} \nabla_{\frac{\partial}{\partial t}} \alpha_* \left(\frac{\partial}{\partial s} \right) &= \nabla_{\frac{\partial}{\partial t}} \nabla_{\frac{\partial}{\partial s}} \alpha_* \left(\frac{\partial}{\partial t} \right) \\ &= R \left(\frac{\partial}{\partial t}, \frac{\partial}{\partial s} \right) \alpha_* \left(\frac{\partial}{\partial t} \right) + \nabla_{\frac{\partial}{\partial s}} \nabla_{\frac{\partial}{\partial t}} \alpha_* \left(\frac{\partial}{\partial t} \right) \\ &= R \left(\frac{\partial}{\partial t}, \frac{\partial}{\partial s} \right) \alpha_* \left(\frac{\partial}{\partial t} \right) \end{aligned}$$

Setting $s = 0$ we have

$$\widehat{\nabla}_{\frac{d}{dt}} \widehat{\nabla}_{\frac{d}{dt}} J = R(\gamma', J)\gamma' = -R(J, \gamma')\gamma'$$

which implies J is a Jacobi field. \square

Corollary 9.1.1. Let $\gamma : [0, 1] \rightarrow M$ be a geodesic with $\gamma(0) = p, \gamma'(0) = v$, where $v \in T_p M$, then for any $w \in T_p M$, consider the following variation of $\gamma(t)$ consisting of geodesics

$$\alpha(t, s) = \exp_p(t(v + sw))$$

Then $J(t) = \alpha_*(\frac{\partial}{\partial s})|_{s=0}$ is a Jacobi field along γ such that

$$J(0) = 0$$

$$J'(0) = w$$

Remark 9.1.2. In fact, for $\alpha(t, s) = \exp_p(t(v + sw))$, we can regard $t(v + sw)$ as a curve parametered by s in $T_p M$, that is it's a curve starting at tv with direction tw . So by definition we have

$$\alpha_*(\frac{\partial}{\partial s})\Big|_{s=0} = (d \exp_p)_{tv}(tw)$$

Remark 9.1.3. In normal coordinate (x^i, U, p) , we can write α explicitly as follows

$$\alpha(t, s) = (t(v^1 + sw^1), \dots, t(v^n + sw^n))$$

where $v = (v^1, \dots, v^n), w = (w^1, \dots, w^n)$ in normal coordinate. Thus Jacobi field J is given by the formula

$$J(t) = tw^i \frac{\partial}{\partial x^i} \Big|_{\gamma(t)}$$

So for arbitrary $q \in U$ and $w \in T_q M$, if we write $w = w^i \frac{\partial}{\partial x^i} \Big|_q$, then the Jacobi field $J(t) = tw^i \frac{\partial}{\partial x^i} \Big|_{\gamma(t)}$ is a Jacobi field such that $J(0) = 0, J(1) = w$.

Corollary 9.1.2. Let (M, g) be a Riemannian manifold, (x^i, U, p) is a normal coordinate containing $p \in M$. For each $q \in U \setminus \{p\}$, every vector in $T_q M$ is the value of a Jacobi field J along a radial geodesic such that J vanishes at p .

9.2. Conjugate points.

Definition 9.2.1 (conjugate points). Let $p \neq q$ be two endpoints of a geodesic γ . p and q are called conjugate along γ if there exists a non-zero Jacobi field J along γ which vanishes at endpoints.

Notation 9.2.1. The conjugate set of p , denoted by $\text{conj}(p)$ is defined as

$$\text{conj}(p) := \{q \in M \mid p \text{ and } q \text{ are conjugate along some geodesic.}\}$$

Remark 9.2.1. There are at most $n - 1$ linearly independent Jacobi fields along γ such that $J(a) = J(b) = 0$. Indeed, by Remark 9.1.1, there are at most n linearly independent Jacobi fields such that $J(a) = 0$. However, trivial Jacobi field $J(t) = (t - a)\gamma'(t)$ never vanishes at $t = b$.

Theorem 9.2.1. Let (M, g) be a Riemannian manifold, $p \in M$ and $v \in V_p \subset T_p M$. Let $\gamma_v : [0, 1] \rightarrow M$ be the geodesic $\gamma_v(t) = \exp_p(tv)$ and $q = \gamma_v(1)$. Then $(d \exp_p)_v$ is not injective if and only if q is conjugate to p along γ_v .

Proof. For any $w \in T_p M$, consider Jacobi field given by

$$J(t) = (d \exp_p)_{tv}(tw)$$

So if $w \neq 0$ lies in the kernel of $(d \exp_p)_v$, then $J(0) = J(1) = 0$, that is p is conjugate to q . Conversely, if p and q are conjugate along γ , then there exists a Jacobi field J such that $J(0) = J(1) = 0$, then it's clear

$$J(t) = (d \exp_p)_{tv}(tw)$$

where $0 \neq w = J'(0) \in T_p M$. Thus

$$(d \exp_p)_v(w) = J(1) = 0$$

which implies $(d \exp_p)_v$ is not injective. \square

Corollary 9.2.1. Let (M, g) be a complete Riemannian manifold, $p \in M$. If the conjugate locus $\text{conj}(p) = \emptyset$, then $\exp_p : T_p M \rightarrow M$ is a local diffeomorphism.

Proof. Since M is complete, then $\exp_p : T_p M \rightarrow M$ is surjective. Furthermore, since the conjugate locus $\text{conj}(p) = \emptyset$, so for arbitrary $v \in T_p M$, we have $(d \exp_p)_v$ is non-degenerated, which implies \exp_p is a local diffeomorphism at $v \in T_p M$. \square

Example 9.2.1. For $p \in \mathbb{S}^n$, we have $\text{conj}(p) = \{-p\}$.

Example 9.2.2. For $p \in S^1 \times \mathbb{R}$, we have $\text{conj}(p) = \emptyset$.

9.3. Jacobi field as a null space.

Lemma 9.3.1. Let $\gamma : [a, b] \rightarrow (M, g)$ be a unit-speed geodesic with no conjugate points, then there exist Jacobi fields J_2, \dots, J_n along γ such that

1. $J_i(a) = 0, i \geq 2$ and $\{\gamma'(b), J_2(b), \dots, J_n(b)\}$ is an orthonormal basis of $T_{\gamma(b)} M$;
2. $\langle J_i(t), \gamma'(t) \rangle \equiv 0$ for any $t \in [a, b]$;
3. $\{\gamma'(t), J_2(t), \dots, J_n(t)\}$ are linearly independent for $t \in (a, b]$.

Proof. For (1). Suppose $\{\gamma'(b), e_2, \dots, e_n\}$ is an orthonormal basis of $T_{\gamma(b)} M$, since there is no conjugate points along γ , there exists a unique Jacobi field J_i such that

$$J_i(a) = 0, J_i(b) = e_i$$

for each $i = 2, \dots, n$.

For (2). Note that

$$\begin{aligned} \frac{d^2}{dt^2} \langle J_i(t), \gamma'(t) \rangle &= \langle \widehat{\nabla}_{\frac{d}{dt}} \widehat{\nabla}_{\frac{d}{dt}} J_i, \gamma' \rangle \\ &= \langle R(J, \gamma') \gamma', \gamma' \rangle \\ &= 0 \end{aligned}$$

Thus $\langle J_i(t), \gamma'(t) \rangle = \lambda t + \mu$. Note that $\langle J_i(a), \gamma'(a) \rangle = \langle J_i(b), \gamma'(b) \rangle = 0$, which implies $\langle J_i(t), \gamma'(t) \rangle \equiv 0$ on $[a, b]$.

For (3). Suppose there exists $c \in (a, b]$ and $\lambda_i \in \mathbb{R}$ such that

$$\sum_{i=2}^n \lambda_i J_i(c) = 0$$

which implies

$$W(t) = \sum_{i=2}^n \lambda_i J_i(t) \equiv 0$$

on $(a, c]$ since there is no conjugate points. By uniqueness we have $W(t) \equiv 0$ on $(a, b]$, thus we have $\lambda_i = 0, i = 2, \dots, n$ from (1). \square

Theorem 9.3.1. Let $\gamma : [a, b] \rightarrow (M, g)$ be a unit-speed geodesic, then

1. If γ has no conjugate points, then index form I_γ is **positive-definite** on vector space consisting of normal variation fields;
2. If γ only has conjugate points as endpoints, then index form is **semipositive-definite** on vector space consisting of normal variation fields. Furthermore, Jacobi field is null space;
3. If γ has an interior conjugate point, then index form is **not positive-definite** on vector space consisting of normal variation fields.

Proof. For (1). Let $\{\gamma'(b), e_2, \dots, e_n\}$ be a orthonormal basis for $T_{\gamma(b)}M$, then there exist unique Jacobi fields J_i such that

$$J_i(a) = 0, J_i(b) = e_i$$

where $i = 2, \dots, n$. Then for any normal variation vector V along γ we write it as

$$V = \sum_{i=2}^n f_i(t) J_i(t)$$

Then it's clear $f_i(b) = 0$ since $V(b) = 0$ and $\{e_2, \dots, e_n\}$ is orthonormal. For index form we have

$$I_\gamma(V, V) = \sum_{i,j=2}^n \left\{ \underbrace{\int_a^b f_i f_j \langle J'_i, J'_j \rangle + f'_i f_j \langle J_i, J'_j \rangle + f_i f'_j \langle J'_i, J_j \rangle dt}_{\text{Part I}} + \underbrace{\int_a^b \{f'_i f'_j \langle J_i, J_j \rangle - f_i f_j R(J_i, \gamma', \gamma', J_j)\} dt}_{\text{Part II}} \right\}$$

Note that

$$\langle J'_i, J_j \rangle = \langle J_i, J'_j \rangle$$

Then Part I is

$$\int_a^b \{(f_i f_j \langle J'_i, J_j \rangle)' - f_i f_j \langle J''_i, J_j \rangle\} dt$$

Thus

$$\begin{aligned} I_\gamma(V, V) &= \sum_{i,j=2}^n f_i f_j \langle J'_i, J_j \rangle \Big|_a^b + \int_a^b f'_i f'_j \langle J_i, J_j \rangle dt \\ &= \sum_{i,j=2}^n \int_a^b f'_i f'_j \langle J_i, J_j \rangle dt \\ &\geq 0 \end{aligned}$$

Furthermore, $I_\gamma(V, V) = 0$ if and only if $\sum_{i=2}^n f'_i J(t) = 0$ if and only if $f'_i(t) = 0, t \in [a, b]$, thus $f_i(t) \equiv 0$, that is $V = 0$.

For (2). For any $c \in (a, b)$, we set $\gamma^c : [a, c] \rightarrow (M, g)$ and define I_{γ^c} . By (1) it's clear I_{γ^c} is positive-definite on the vector space consisting of normal variation fields along γ^c . By standard approximation argument we can show I_γ is semipositive-definite.

To see its null space: It's clear a normal variation Jacobi field V satisfies $I_\gamma(V, V) = 0$; Conversely, if a normal variation field V satisfies $I_\gamma(V, V) = 0$, then by a variation argument we have for arbitrary W we have

$$I_\gamma(V, W) = 0$$

Take appropriate W to see V satisfies the equation for Jacobi fields.

For (3). If $\gamma(a)$ is conjugate to $\gamma(c)$ for some $c \in (a, b)$, then there exists a non-zero normal Jacobi field J_1 along $\gamma([a, c])$ such that $J_1(a) = J_1(c) = 0$. Consider

$$J = \begin{cases} J_1(t) & t \in [a, c] \\ 0 & t \in [c, b] \end{cases}$$

It's easy to see $I_\gamma(J, J) = 0$. Note that here our J may not be smooth. Let W be a smooth normal variation vector field along γ such that $W(c) =$

$-\lim_{t \rightarrow c^-} \nabla_{\frac{d}{dt}} J_1$. It's clear $W(c) \neq 0$. Consider $J_\varepsilon = J + \varepsilon W$ and so⁵

$$I_\gamma(J_\varepsilon, J_\varepsilon) = 2\varepsilon I_\gamma(J, W) + \varepsilon^2 I_\gamma(W, W)$$

And integration by parts we have

$$I_\gamma(J, W) = \left\langle \widehat{\nabla}_{\frac{d}{dt}} J_1, W \right\rangle \Big|_a^c = -W(c)^2 < 0$$

So for sufficiently small ε we have $I_\gamma(J_\varepsilon, J_\varepsilon) < 0$, and by approximation argument we can show there exists a smooth normal variation field such that $I_\gamma(V, V) < 0$. \square

Corollary 9.3.1. Let $\gamma : [a, b] \rightarrow (M, g)$ be a unit-speed geodesic with no conjugate points, and V, W are normal vector fields satisfying $V(a) = W(a), V(b) = W(b)$. If V is a Jacobi field, then $I_\gamma(V, V) \leq I_\gamma(W, W)$, and the equality holds if and only if $V = W$.

Proof. Since V, W agree at end points, then $V - W$ is a normal variation field, thus we have

$$0 \leq I_\gamma(V - W, V - W) = I_\gamma(V, V) + I_\gamma(W, W) - 2I_\gamma(V, W)$$

Since V is a Jacobi field, then integration by parts shows

$$I_\gamma(V, V) = \left\langle \widehat{\nabla}_{\frac{d}{dt}} V, V \right\rangle \Big|_a^b = \left\langle \widehat{\nabla}_{\frac{d}{dt}} V, W \right\rangle \Big|_a^b = I_\gamma(V, W)$$

Hence we get $I_\gamma(V, V) \leq I_\gamma(W, W)$, and the equality holds if and only if $V = W$. \square

Remark 9.3.1. From second variation formula, we can conclude that a geodesic γ is a **locally minimal geodesic** if and only if it has no interior conjugate points. However, it may not be **globally minimal geodesic**. Indeed, consider $M = S^1 \times \mathbb{R}$, it's clear there is no conjugate points for any geodesic on M , thus for geodesic $\gamma : [a, b] \rightarrow M$ starting at $(x, y) \in M$, it's locally minimal, but if there exists $c \in (a, b)$ such that $\gamma(c) \in \{-x\} \times \mathbb{R}$, then γ is not globally minimal.

⁵Note that here our J and J_ε may not be smooth, so keep in mind here we already extend our index form I_γ to the one defined on piecewise smooth vector field.

10. CUT LOCUS AND INJECTIVE RADIUS

10.1. Cut locus.

Definition 10.1.1 (distance). Let (M, g) be a complete Riemannian manifold, $p, q \in M$, the distance between p and q are the length of minimal geodesic connecting p, q , denoted by $\text{dist}(p, q)$.

Definition 10.1.2 (cut time/point/locus). Let (M, g) be a complete Riemannian manifold, $p \in M$ and $v \in T_p M$.

1. The cut time of (p, v) is defined as

$$t_{\text{cut}}(p, v) = \sup\{c > 0 \mid \gamma_v|_{[0, c]} \text{ is a minimal geodesic}\}$$

2. Suppose $t_{\text{cut}}(p, v) < \infty$, the cut point of p along γ along γ_v is $\gamma_v(t_{\text{cut}}(p, v)) \in M$;
3. The cut locus of p , denoted by $\text{cut}(p)$ is the set

$$\text{cut}(p) = \{q \in M \mid \exists v \in T_p M \text{ such that } q \text{ is a cut point of } p \text{ along } \gamma_v.\}$$

Remark 10.1.1. Here are some remarks:

1. It's possibly for $t_{\text{cut}}(p, v)$ to be $+\infty$. For example, just let M be Euclidean space with standard metric;
2. The cut point (if it exists) occurs at or before the first conjugate point along every geodesic;
3. It's clear that $t_{\text{cut}}(p, v)$ depends on the $|v|$, but $\gamma_v(t_{\text{cut}}(p, v))$ is independent of $|v|$. So when we consider cut points of p along some geodesic γ , we always assume γ is unit-speed.

Theorem 10.1.1. Let (M, g) be a complete Riemannian manifold, $p \in M, v \in T_p M$ with unit length. Let $c = t_{\text{cut}}(p, v) \in (0, \infty]$, then

1. If $0 < b < c$ and b is finite, then $\gamma_v|_{[0, b]}$ has no conjugate point and it is the unique minimal unit-speed geodesic between endpoints;
2. If $c < \infty$, then $\gamma_v|_{[0, c]}$ is a minimal geodesic.
3. In the case of (2), one or both of the following holds:
 - (a) $\gamma_v(c)$ is conjugate to p along γ_v ;
 - (b) There are two or more different unit-speed minimal geodesic connecting p and $\gamma_v(c)$.

Proof. (1) and (2) is clear. For (3). Assume $\gamma_v(c)$ is not conjugate to p along γ_v , we shall prove the existence of another unit-speed minimal geodesic from p to $\gamma_v(c)$.

Firstly we choose a sequence $\{b_i\}$ descending to c . Note that $\gamma_v : [0, b_i] \rightarrow M$ is not a minimal geodesic, thus there exists a unit-speed minimal geodesic $\gamma_i : [0, a_i] \rightarrow M$ connecting p and $\gamma_v(b_i)$. In particular we have

1. $\gamma_i(a_i) = \gamma_v(b_i)$;
2. $a_i < b_i$.

If we denote $\omega_i = \gamma_i'(0) \in T_p M$, by compactness of unit sphere on $T_p M$ and the fact $\{a_i\}$ is bounded, we can find a subsequence of $\{\gamma_i\}$ such that

ω_i converging to some $w \in T_p M$ with $|w| = 1$, and $\lim_{i \rightarrow \infty} a_i = a$. For convenience we still denote this subsequence by $\{\gamma_i\}$.

On one hand $\gamma_i(a_i) = \exp_p(a_i w_i)$ converges to $\exp_p(aw)$; On the other hand, $\gamma_i(a_i) = \gamma_v(b_i)$ implies $\exp_p(cv) = \gamma_v(c) = \exp_p(aw)$. Furthermore,

$$c = \text{dist}(p, \gamma_v(c)) = \lim_{i \rightarrow \infty} \text{dist}(p, \gamma_v(b_i)) = \lim_{i \rightarrow \infty} \text{dist}(p, \gamma_i(a_i)) = \lim_{i \rightarrow \infty} a_i = a$$

So it suffices to check $v \neq w$.

By assumption we have $\gamma_v(c)$ is not conjugate to p , thus cv is not a critical point of \exp_p , that is \exp_p is injective in $B_\varepsilon(cv)$, where $\varepsilon > 0$ is sufficiently small. On one hand we have $a_i w_i \neq b_i v$ since $a_i < b_i$; On the other hand we have

$$\exp_p(b_i v) = \gamma_v(b_i) = \gamma_i(a_i) = \exp_p(a_i w_i)$$

Thus injectivity implies $a_i w_i \notin B_\varepsilon(cv)$ for sufficiently large i , since in this case $b_i v \in B_\varepsilon(cv)$. Taking limits we have

$$aw \neq cv$$

that is $w \neq v$. □

Example 10.1.1. Consider the following cases:

1. $M = \mathbb{S}^n$, then $\text{cut}(p) = \text{conj}(p) = \{-p\}$. In this case both (a), (b) hold in Theorem 10.1.1;
2. $M = \mathbb{S}^1 \times \mathbb{R}$, then $\text{cut}(p) = \{-p\} \times \mathbb{R}$. In this case (a) fails and (b) holds in Theorem 10.1.1;
- 3.

Definition 10.1.3 (injective radius). Let (M, g) be a Riemannian manifold, $p \in M$. The injective radius of p is defined as

$$\text{inj}(p) := \sup\{\rho > 0 : \exp_p \text{ is defined on } B(0, \rho) \subset T_p M \text{ and injective}\}$$

The injectivity radius of M is

$$\text{inj}(M) := \inf_{p \in M} \text{inj}(p)$$

Theorem 10.1.2. Let (M, g) be a complete Riemannian manifold, then

$$\text{inj}(p) = \begin{cases} \text{dist}(p, \text{cut}(p)) & \text{cut}(p) \neq \emptyset \\ \infty & \text{cut}(p) = \emptyset \end{cases}$$

Proof. See Proposition 10.36 in Page 312 of [Lee18]. □

Proposition 10.1.1. Let (M, g) be a complete Riemannian manifold and $p \in M$. Suppose there exists some point $q \in \text{cut}(p)$ such that $\text{dist}(p, q) = \text{dist}(p, \text{cut}(p))$, then

1. Either q is a conjugate point of p along some minimizing geodesic from p to q , or there are exactly two minimizing geodesics from p to q , say $\gamma_1, \gamma_2 : [0, b] \rightarrow M$, such that $\gamma_1'(b) = -\gamma_2'(b)$.

2. If in addition that $\text{inj}(p) = \text{inj}(M)$, and q is not conjugate to p along any minimizing geodesic, then there is a closed unit-speed geodesic $\gamma : [0, 2b] \rightarrow M$ such that $\gamma(0) = \gamma(2b) = p$ and $\gamma(b) = q$ where $b = \text{dist}(p, q)$.

Proof. For (1). Suppose q is not conjugate to p along any minimizing geodesic, then by Theorem 10.1.1 there are at least two unit-speed minimal geodesics $\gamma_1(t), \gamma_2(t)$ such that $\gamma_1(b) = \gamma_2(b) = q$. Suppose $\gamma'_1(b) \neq -\gamma'_2(b)$, then there exists unit vector $X_q \in T_q M$ such that

$$\langle X_q, \gamma'_1(b) \rangle < 0, \quad \langle X_q, \gamma'_2(b) \rangle < 0$$

Since q is not conjugate to p along γ_1 , there exists a neighborhood U_1 of $b\gamma'_1(0)$ in $T_p M$ such that $\exp_p|_{U_1}$ is diffeomorphism. Now choose a sufficiently small s and let

$$\xi_1(s) = (\exp_p|_{U_1})^{-1} \exp_q(sX_q)$$

Consider the following variation of γ_1 consisting of geodesics:

$$\alpha_1(t, s) = \exp\left(\frac{t}{b}\xi_1(s)\right)$$

It's clear $\alpha_1(t, 0) = \gamma_1(t)$, since $\xi_1(0) = (\exp_p|_{U_1})^{-1} \exp_q(0) = (\exp_p|_{U_1})^{-1}(q) = b\gamma'_1(0)$. Then the first variation formula yields

$$\left. \frac{dL(\gamma_s)}{ds} \right|_{s=0} = \langle X_q, \gamma'_1(b) \rangle < 0$$

which implies for sufficiently small s we have $L(\alpha_1(t, s)) < L(\gamma_1(t))$. For γ_2 we can do the same construction and the same argument implies for sufficiently small s we have $L(\alpha_2(t, s)) < L(\gamma_2(t))$. Thus for each sufficiently small s we have two geodesics $\alpha_1(t, s), \alpha_2(t, s)$ from p to $\exp_q(sX_q)$. However,

$$(10.1) \quad d(p, \exp_q(sX_q)) \leq L(\alpha_1(t, s)) < L(\gamma_1(t)) = \text{dist}(p, q) = \text{inj}(p)$$

A contradiction to the definition of injective radius. So any two different minimizing geodesics γ_1, γ_2 from p to q satisfy $\gamma'_1(b) = -\gamma'_2(b)$, which implies there are exactly two minimizing geodesics from p to q .

For (2). By (1) we know that there exists exactly two geodesics γ_1, γ_2 such that $\gamma_1(b) = \gamma_2(b) = q$ with $\gamma'_1(b) = \gamma'_2(b)$. Consider the loop $\gamma = \gamma_1 \circ \gamma_2^{-1}$, then it's a unit-speed geodesic such that $\gamma(0) = \gamma(2b) = p, \gamma(b) = q$, where $b = \text{dist}(p, q)$, since we have already shown $\gamma'_1(b) = -\gamma'_2(b)$. To show γ is a closed geodesic, it suffices to show $\gamma'(2b) = \gamma'(0)$, that is equivalent to show $(\gamma_1^{-1})'(b) = (\gamma_2^{-1})'(b)$. Note that in the proof of (1), condition of $\text{dist}(p, q) = \text{dist}(p, \text{cut}(p)) = \text{inj}(p)$ is used in inequality (10.1), and in fact we only need $\text{dist}(p, q) \leq \text{inj}(p)$, strict equality is not necessary. So if $\text{inj}(p) = \text{inj}(M)$, thus

$$\text{dist}(q, p) = \text{dist}(p, q) = \text{inj}(p) \leq \text{inj}(q)$$

Then (1) implies $(\gamma_1^{-1})'(b) = (\gamma_2^{-1})'(b)$. □

Part 4. Topology of Riemannian manifold

11. TOPOLOGY OF NON-POSITIVE SECTIONAL CURVATURE MANIFOLD

11.1. Cartan-Hadamard manifold.

Definition 11.1.1 (Cartan-Hadamard manifold). A simply-connected, complete Riemannian manifold with non-positive sectional curvature is called Cartan-Hadamard manifold.

11.1.1. *Expansion property of exponential map of Cartan-Hadamard manifold.* In this section we explore some properties of Cartan-Hadamard manifold using Jacobi fields.

Proposition 11.1.1. Let $p \in M$ and $\gamma : [0, 1] \rightarrow M$ be a geodesic such that $\gamma(0) = p, \gamma'(0) = v$. Then for any $w \in T_p M$ with $|w| = 1$, let $J(t)$ be the Jacobi field along γ given by

$$J(t) = (d \exp_p)_{tv}(tw)$$

Then we have the following Taylor expansions about $t = 0$

$$\begin{aligned} |J(t)|^2 &= t^2 - \frac{1}{3}R(J', \gamma', \gamma', J')(0)t^4 + O(t^4) \\ |J(t)| &= t - \frac{1}{6}R(J', \gamma', \gamma', J')(0)t^3 + O(t^3) \end{aligned}$$

Proof. For (1). Since $J(0) = 0, J'(0) = w$, the first three coefficients are given as

$$\begin{aligned} \langle J, J \rangle(0) &= 0 \\ \langle J, J' \rangle(0) &= 2\langle J, J' \rangle(0) = 0 \\ \langle J, J'' \rangle(0) &= 2\langle J', J' \rangle(0) + 2\langle J'', J \rangle(0) = 2 \\ \langle J, J''' \rangle(0) &= 6\langle J', J'' \rangle(0) + 2\langle J''', J \rangle(0) = 0 \\ &= 6\langle J', R(J, \gamma')\gamma' \rangle(0) = 0 \\ \langle J, J'''' \rangle(0) &= 8\langle J', J''' \rangle(0) + 6\langle J'', J'' \rangle(0) + 2\langle J''', J \rangle(0) \\ &= 8\langle J', J''' \rangle(0) + 6\langle R(J, \gamma')\gamma', R(J, \gamma')\gamma' \rangle(0) \\ &= 8\langle J', J''' \rangle(0) \end{aligned}$$

So we need to compute J''' . For arbitrary vector field W along γ , direct computation shows

$$\begin{aligned} \langle \widehat{\nabla}_{\frac{d}{dt}} R(J, \gamma')\gamma', W \rangle &= \frac{d}{dt} \langle R(J, \gamma')\gamma', W \rangle - \langle R(J, \gamma')\gamma, W' \rangle \\ &= \frac{d}{dt} \langle R(W, \gamma')\gamma', J \rangle - \langle R(J, \gamma')\gamma, W' \rangle \\ &= \langle R(W, \gamma')\gamma', J' \rangle - \langle \widehat{\nabla}_{\frac{d}{dt}} R(W, \gamma')\gamma', J \rangle - \langle R(J, \gamma')\gamma, W' \rangle \\ &= \langle R(J', \gamma')\gamma', W \rangle - \langle \widehat{\nabla}_{\frac{d}{dt}} R(W, \gamma')\gamma', J \rangle - \langle R(J, \gamma')\gamma, W' \rangle \end{aligned}$$

Setting $t = 0$ we obtain

$$\langle J', J''' \rangle(0) = -\langle J'(0), \widehat{\nabla}_{\frac{d}{dt}} R(J, \gamma') \gamma' \Big|_{t=0} \rangle = -R(J', \gamma', \gamma', J')(0)$$

So we have

$$|J(t)|^2 = t^2 - \frac{1}{3} R(J', \gamma', \gamma', J')(0) t^4 + O(t^4)$$

For (2). It follows directly from (1). \square

Theorem 11.1.1. Let (M, g) be a simply-connected complete Riemannian manifold. The followings are equivalent:

1. M is Cartan-Hadamard manifold;
2. For any $p \in M$ and $v, w \in T_p M$, we have

$$|(\mathrm{d} \exp_p)_v w| \geq |w|$$

3. For any $p \in M, T > 0$ and $v, w \in T_p M$, we have

$$|v - w| \leq \frac{\mathrm{dist}(\exp_p(tv), \exp_p(tw))}{t}$$

holds for arbitrary $t > 0$.

Proof. From (1) to (2). For all $p \in M$ and $v, w \in T_p M$, consider geodesic $\exp_p(tv)$ and Jacobi field

$$J(t) = (\mathrm{d} \exp_p)_{tv}(tw)$$

along it. If M has non-positive sectional curvature, direct computation shows

$$|J(t)|'' = \frac{|J|^2 |J'|^2 - \langle J, J' \rangle^2}{|J|^3} - \frac{R(J, \gamma', \gamma', J)}{|J|} \geq 0$$

for all $t > 0$. Thus consider

$$f(t) = |J(t)| - t|w|$$

It's clear $f''(t) \geq 0$ and $f'(0) = 0$, thus $f(t) \geq 0$ for all $t > 0$ since $f(0) = 0$. In particular, set $t = 1$ we have

$$|(\mathrm{d} \exp_p)_v(w)| - |w| \geq 0$$

From (2) to (1). If M has sectional curvature $K(\sigma) > 0$ at $p \in M$, where σ is the plane spanned by v, w with $|v| = |w| = 1$. Then consider geodesic $\exp_p(tv)$ and Jacobi field

$$J(t) = (\mathrm{d} \exp_p)_{tv}(tw)$$

along it. Then by Proposition 11.1.1 we have $|J(t)|'' < 0$ for sufficiently small t . If we set $f(t) = |J(t)| - t|w|$, then we can see $f(0) = 0, f'(0) = 0$ and $f''(0) < 0$ for sufficiently small t . In particular, we have

$$|(\mathrm{d} \exp_p)_{\varepsilon v}(\varepsilon w)| - |\varepsilon w| = f(\varepsilon) < 0$$

where $\varepsilon > 0$ is sufficiently small. This leads to a contradiction.

From (2) to (3). For arbitrary $t > 0$. Let $\gamma(s) : [0, 1] \rightarrow M$ be a geodesic connecting $\exp_p(tv), \exp_p(tw)$ and choose a curve $v(s) \in T_p M$ such that

$$\exp_p(v(s)) = \gamma(s)$$

for all $s \in [0, 1]$. Hence $v(0) = tv, v(1) = tw$. Then

$$\begin{aligned} \text{dist}(\exp_p(tv), \exp_p(tw)) &= \int_0^1 |\gamma'(s)| ds \\ &= \int_0^1 |(\text{d exp}_p)_{v(s)}(v'(s))| ds \\ &\geq \left| \int_0^1 v'(s) ds \right| \\ &= t|v - w| \end{aligned}$$

This shows

$$|v - w| \leq \frac{\text{dist}(\exp_p(tv), \exp_p(tw))}{t}$$

holds for arbitrary $t > 0$.

From (3) to (2). Note that

$$\begin{aligned} |(\text{d exp}_p)_v(w)| &= \lim_{t \rightarrow 0} \frac{\exp_p(v + tw) - \exp_p(v)}{t} \\ &= \lim_{t \rightarrow 0} \frac{\exp_p(tv' + tw) - \exp_p(tv')}{t} \\ &\geq |v' + w - v'| \\ &= |w| \end{aligned}$$

□

Corollary 11.1.1. Let (M, g) be a Cartan-Hadamard manifold with $a, b, c \in M$. Such points determine a unique geodesic triangle T with vertices a, b, c . Let α, β, γ be the angles of the vertices a, b, c respectively, and let A, B, C be the lengths of the side opposite the vertices a, b, c respectively. Then

1. $A^2 + B^2 - 2AB \cos \gamma \leq C^2 (< C^2, \text{ if } K < 0)$;
2. $\alpha + \beta + \gamma \leq \pi (< \pi, \text{ if } K < 0)$

Proof. See Lemma 3.1 in Page 259 of Do Carmo [Car92].

□

So you find that the exponential map of simply-connected complete Riemannian manifold with non-positive sectional curvature has a property of “expansion”.

11.1.2. *Complete Riemannian manifold with non-positive sectional curvature is $K(G, 1)$.*

Lemma 11.1.1. If (M, g) is a complete Riemannian manifold with sectional curvature $K \leq 0$, then for any $p \in M$, the conjugate locus $\text{conj}(p) = \emptyset$. In particular, $\exp_p : T_p M \rightarrow M$ is a local diffeomorphism.

Proof. Suppose q is conjugate to p along $\gamma : [0, 1] \rightarrow M$, and without loss of generality we may assume there is no conjugate point for $t \in (0, 1)$. Let J be a Jacobi field along γ with $J(0) = J(1) = 0$, then

$$\begin{aligned} \left(\frac{1}{2}|J|^2\right)' &= (g(J', J))' \\ &= g(J'', J) + g(J', J') \\ &= -R(J, \gamma', \gamma', J) + |J'|^2 \\ &\geq |J'|^2 \end{aligned}$$

Since $J'(0) \neq 0$, we have

$$\begin{aligned} g(J', J)(t) &\geq \int_0^t |J'|^2 + g(J'(0), J(0)) \\ &= \int_0^t |J'|^2 \\ &> 0 \end{aligned}$$

which implies $(\frac{1}{2}|J|^2)' = g(J', J) > 0$, a contradiction to $J(1) = 0$. \square

Lemma 11.1.2. Let M be a complete Riemannian manifold and let $f : M \rightarrow N$ be a local diffeomorphism onto a Riemannian manifold N which has the following property: For all $p \in M$ and for all $v \in T_p M$, we have $|df_p(v)| \geq |v|$. Then f is a covering map.

Proof. See Lemma 3.3 in Page150 of Do carmo[Car92]. \square

Theorem 11.1.2 (Cartan-Hadamard). If (M, g) is a complete Riemannian manifold with sectional curvature $K \leq 0$, then $\exp_p : T_p M \rightarrow M$ is a covering map.

Proof. Combine above two lemmas with Theorem 11.1.1. \square

Corollary 11.1.2. Cartan-Hadamard manifold is diffeomorphic to \mathbb{R}^n .

Corollary 11.1.3. If (M, g) is a complete Riemannian manifold with $K \leq 0$, then $\pi_k(M) = 0, k \geq 2$, that is M is $K(\pi_1(M), 1)$.

Remark 11.1.1. Thoery in topology says if a finite dimension CW complex is a $K(G, 1)$ space, then its fundamental group is torsion-free. So if M is a complete Riemannian manifold with $K \leq 0$, we have $\pi(M)$ is torsion-free. We will prove this fact later by tools of Riemannian manifold, called Cartan's torsion-free theorem.

Corollary 11.1.4. If M and N are two compact Riemannian manifold and one of them is simply-connected, then $M \times N$ has no metric with non-positive sectional curvature.

Proof. If both of M and N are simply-connected, and $M \times N$ admits a metric with non-positive sectional curvature, then it's diffeomorphic to \mathbb{R}^n for some positive integer n , a contradiction to compactness.

So suppose M is simply-connected and N is not simply-connected with universal covering \tilde{N} , then there is a universal covering

$$\pi : M \times \tilde{N} \rightarrow M \times N$$

If $M \times N$ admits a Riemannian metric g with non-positive sectional curvature, then π^*g is a complete metric of non-positive sectional curvature on $M \times \tilde{N}$, so we have $M \times \tilde{N}$ is diffeomorphic to \mathbb{R}^n for some n . M is orientable since it's simply-connected, thus $H^m(M) = \mathbb{Z}$, where $m = \dim M$, thus by Künneth formula $H^m(M \times \tilde{N}) \neq 0$, a contradiction to Poincaré lemma. \square

Remark 11.1.2. The condition simply-connected is crucial, for example $S^1 \times S^1$.

11.2. Cartan's torsion-free theorem.

Lemma 11.2.1. Let (M, g) be a Cartan-Hadamard manifold, $p \in M$ and $v \in T_p M$. For all $q \in M$ we have

$$2 \operatorname{dist}(p, q)^2 + \operatorname{dist}(p_0, p)^2 + \operatorname{dist}(p_1, p)^2 \leq \operatorname{dist}(p_0, q)^2 + \operatorname{dist}(p_1, q)^2$$

where $p_0 = \exp_p(-v)$, $p_1 = \exp_p(v)$.

Proof. Since $\exp_p : T_p M \rightarrow M$ is a diffeomorphism, there exists $w \in T_p M$ such that $q = \exp_p(w)$ with $\operatorname{dist}(p, q) = |w|$. So we have

$$\begin{aligned} \operatorname{dist}(p_0, q) &= \operatorname{dist}(\exp_p(-v), \exp_p(w)) \geq |w + v| \\ \operatorname{dist}(p_1, q) &= |w - v| \\ \operatorname{dist}(p, q)^2 &= |w|^2 \\ &= \frac{|w + v|^2 + |w - v|^2}{2} - |v|^2 \\ &\leq \frac{\operatorname{dist}(p_0, q)^2 + \operatorname{dist}(p_1, q)^2}{2} - \frac{\operatorname{dist}(p_0, p)^2 + \operatorname{dist}(p_1, p)^2}{2} \end{aligned}$$

\square

Lemma 11.2.2 (Serre). Let (M, g) be a Cartan-Hadamard manifold, $p \in M$ and $B(p, r)$ be the closed ball of radius r . If $\Omega \subset M$ is non-empty bounded set and define

$$r_\Omega = \inf\{r > 0 \mid \Omega \subset B(p, r)\}$$

There exists unique $p_\Omega \in M$ such that $\Omega \subset B(p_\Omega, r_\Omega)$.

Proof. Existence: Choose a sequence $r_i > r_\Omega$ and $p_i \in M$ such that

$$\Omega \subset B(p_i, r_i), \lim r_i = r_\Omega$$

Fix arbitrary $q \in \Omega$, one has $\operatorname{dist}(q, p_i) \leq r_i$ for each i , thus $\{p_i\}$ is bounded since we can choose $\{r_i\}$ is bounded, which has a convergent subsequence since M is complete. The limit of this convergent subsequence is p_Ω .

Uniqueness: Let $p_0, p_1 \in M$ such that

$$\Omega \subset B(p_0, r_\Omega) \cap B(p_1, r_\Omega)$$

Since \exp_{p_0} is a diffeomorphism, there exists unique v_0 such that $p_1 = \exp_{p_0} v_0$. Set $p = \exp_{p_0}(v_0/2)$, for all $q \in \Omega$ we have

$$\begin{aligned} \text{dist}(p, q)^2 &\leq \frac{\text{dist}(p_0, q)^2 + \text{dist}(p_1, q)^2}{2} - \frac{\text{dist}(p_0, p_1)^2}{4} \\ &\leq r_\Omega^2 - \frac{\text{dist}(p_0, p_1)^2}{4} \end{aligned}$$

By definition of r_Ω , we have $\text{dist}(p_0, p_1) = 0$, hence $p_0 = p_1$. \square

Theorem 11.2.1 (Cartan's fixed-point theorem). Suppose (M, g) is a Cartan-Hadamard manifold and G is a compact Lie group acting smoothly and isometrically on M , then G has a fixed-point in M .

Proof. Let $p \in M$, consider its orbit

$$\Omega = \{gp \mid g \in G\}$$

it's a bounded since M is compact. Note

$$\Omega = g\Omega \subset B(gp_\Omega, r_\Omega)$$

Then by uniqueness of p_Ω , we have p_Ω is a fixed-point of G . \square

Corollary 11.2.1. If (M, g) is a complete Riemannian manifold with $K \leq 0$, then $\pi_1(M)$ is torsion-free.

Proof. Consider the universal covering \widetilde{M} of M , it's a Cartan-Hadamard manifold, and $\pi_1(M)$ acts on \widetilde{M} freely. If there exists a torsion element φ , consider the finite group G generated by φ , it's a 0-dimension Lie group with discrete topology. By Cartan's fixed-point theorem there exists a fixed-point of G , which implies φ is identity, since $\pi_1(M)$ -action is free. \square

11.3. Preissmann's Theorem.

Definition 11.3.1 (axis). Let (M, g) be a complete Riemannian manifold, $F : M \rightarrow M$ is an isometry. A non-trivial geodesic $r : \mathbb{R} \rightarrow M$ is called an axis of F if $F \circ r$ is a non-trivial translation of r , that is there exists $c > 0$ such that

$$F(r(t)) = r(t + c)$$

Definition 11.3.2. An isometry with no fixed points that has an axis is said to be axial.

Lemma 11.3.1. $F : (M, g) \rightarrow (M, g)$ is an isometry of complete Riemannian manifold, if $\delta_F(p) = \text{dist}(p, F(p))$ has a positive minimum, then F has an axis.

Proof. Suppose δ_F attains its minimum at some $p \in M$, $\gamma(t) : [0, 1] \rightarrow M$ is a minimum geodesic connecting p and $F(p)$, then $F \circ \gamma : [0, 1] \rightarrow M$ is also a minimum geodesic connecting $F(p)$ and $F^2(p)$, since F is an isometry.

We claim these two geodesics form an angle π at point $F(p)$ and thus fit together an extension of γ to $[0, 2]$. Indeed, for any $t \in [0, 1]$,

$$\begin{aligned}
 \delta_F(p) &= \text{dist}(p, F(p)) \\
 &\leq \delta_F(\gamma(t)) \\
 &= \text{dist}(\gamma(t), F(\gamma(t))) \\
 &\leq \text{dist}(\gamma(t), \gamma(1)) + \text{dist}(\gamma(1), F(\gamma(1))) \\
 &= \text{dist}(\gamma(t), \gamma(1)) + \text{dist}(F(\gamma)(0), F(\gamma(t))) \\
 &= \text{dist}(\gamma(t), \gamma(1)) + \text{dist}(\gamma(0), \gamma(t)) \\
 &= \delta_F(p)
 \end{aligned}$$

Thus we have $(F \circ \gamma)(t) = \gamma(1 + t)$ for $0 \leq t \leq 1$. Repeating this argument to obtain a geodesic $\gamma : \mathbb{R} \rightarrow M$ with period 1, and it's an axis for F . \square

Lemma 11.3.2. Let (M, g) be a compact Riemannian manifold and $F : \widetilde{M} \rightarrow \widetilde{M}$ be a non-trivial deck transformation on the universal covering map $\pi : \widetilde{M} \rightarrow M$. Then

1. δ_F has a positive minimum and $\delta_F \geq 2 \text{inj}(M)$. In particular, F has an axis $\gamma : \mathbb{R} \rightarrow \widetilde{M}$;
2. $\pi \circ r$ is a closed geodesic in M whose length is minimal in the free homotopy class $[\pi \circ \gamma]$.

Lemma 11.3.3. Suppose (M, g) is a Cartan-Hadamard manifold with $K < 0$, if $F : M \rightarrow M$ has an axis, then it's unique up to reparametrization.

Proof. Suppose $\gamma_1, \gamma_2 : \mathbb{R} \rightarrow M$ are two axes of F , without lose of generality we may assume

$$\begin{aligned}
 F(\gamma_1(t)) &= \gamma_1(t + 1) \\
 F(\gamma_2(t)) &= \gamma_2(t + 1)
 \end{aligned}$$

Suppose γ_1, γ_2 do not intersect, then points $A = \gamma_1(0), B = \gamma_1(1) = F(A), C = \gamma_2(0)$ and $D = \gamma_2(1) = F(C)$ are all distinct. Let σ be a geodesic from A to C , then $F \circ \sigma$ is the geodesic from B to D . Furthermore, the geodesic quadrilateral $ABCD$ has angle sum 2π , since F preserves angles. However, according to Lemma 11.1.1, triangle $\triangle ABC$ and $\triangle BCD$ have angle sum strictly less than π , and

$$\begin{aligned}
 \angle ACD &\leq \angle ACB + \angle BCD \\
 \angle ABD &\leq \angle ABC + \angle CBD
 \end{aligned}$$

thus the angle sum of $ABCD$ is strictly less than 2π , a contradiction.

Hence γ_1 and γ_2 must intersect at some point $p = \gamma_1(t_1) = \gamma_2(t_2)$, then

$$\begin{aligned}
 F(p) &= F(\gamma_1(t_1)) = \gamma_1(t_1 + 1) \\
 &= F(\gamma_2(t_2)) = \gamma_2(t_2 + 1)
 \end{aligned}$$

is another intersection point. By the uniqueness of geodesic we have γ_1 is a reparametrization of γ_2 . \square

Lemma 11.3.4. If H is an additive subgroup of \mathbb{R} , then either H is dense in \mathbb{R} or $H \cong \mathbb{Z}$.

Proof. Let H be an additive subgroup of \mathbb{R} , it's clear $H \cap \mathbb{R}^{>0} \neq \emptyset$, consider

$$b := \inf\{h \in H \cap \mathbb{R}^{>0}\}$$

1. If $b > 0$: Let $h \in H$ and $k \in \mathbb{Z}$ such that

$$kb \leq |h| < (k+1)b$$

then we have $|h| - kb \in H$, and $0 \leq |h| - kb < (k+1)b - kb = b$. By the choice of b , we have $|h| - kb = 0$, which implies $h = \pm kb$. In this case $H = b\mathbb{Z}$.

2. If $b = 0$: For arbitrary $r \in \mathbb{R}^{\geq 0}$ and $\varepsilon > 0$, there exists $h \in H \cap (0, \varepsilon]$ since $b = 0$ and $k \in \mathbb{N}$ such that

$$kh \leq r \leq (k+1)h$$

Thus

$$0 \leq r - kh \leq (k+1)h - kh = h \leq \varepsilon$$

which implies $|r - kh| \leq \varepsilon$, that is H is dense in $\mathbb{R}^{\geq 0}$. For the same argument you can show H is also dense in $\mathbb{R}^{\leq 0}$.

□

Theorem 11.3.1 (Preissmann). If (M, g) is a compact Riemannian manifold with negative sectional curvature, then any non-trivial abelian subgroup of $\pi_1(M)$ is isomorphic to \mathbb{Z} .

Proof. Let $\pi : \widetilde{M} \rightarrow M$ be the universal covering, then (\widetilde{M}, π^*g) is a Cartan-Hadamard manifold with negative sectional curvature. Now it suffices to show every non-trivial abelian subgroup H of $\text{Aut}_\pi(\widetilde{M})$ is isomorphic to \mathbb{Z} , since $\pi_1(M) \cong \text{Aut}_\pi(\widetilde{M})$. Let φ be a non-trivial deck transformation and $\gamma : \mathbb{R} \rightarrow \widetilde{M}$ is an axis of φ , that is there exists $c > 0$ such that

$$\varphi \circ \gamma(t) = \gamma(t + c)$$

for all $t \in \mathbb{R}$. If ψ is another non-trivial element of H , then for any $t \in \mathbb{R}$ we have

$$\varphi(\psi(\gamma(t))) = \psi\varphi(\gamma(t)) = \psi(\gamma(t + c))$$

which implies $\psi \circ \gamma$ is also an axis of φ . So by Lemma 11.3.3 we have $\psi \circ \gamma$ is a reparametrization of γ . Furthermore, they have the same speed since ψ is an isometry, thus $\psi(\gamma(t)) = \gamma(t + a)$ or $\psi(\gamma(t)) = \gamma(-t + a)$. The latter can't happen, otherwise $\psi(\gamma(\frac{a}{2})) = \gamma(\frac{a}{2})$, contradicts to $\text{Aut}_\pi(\widetilde{M})$ acts on \widetilde{M} freely.

Define $F : H \rightarrow \mathbb{R}$ by $F(\psi) = a$ such that $\psi(\gamma(t)) = \gamma(t + a)$. It's easy to see F is a group homomorphism with trivial kernel and $F(H)$ is an additive subgroup of \mathbb{R} . Consider

$$b := \inf\{h \in F(H) \cap \mathbb{R}^+\}$$

By Lemma 11.3.4, it suffices to show $b > 0$. If $b = 0$, then there exists $a \in (0, \text{inj}(M))$ and $\psi \in H$ such that

$$a = F(\psi), \quad \psi(\gamma(t)) = \gamma(t + a)$$

Since $\pi \circ \psi = \pi$, thus we have $\pi(\gamma(t)) = \pi(\gamma(t + a))$. Set $t = 0$ one has

$$\pi(\gamma(a)) = \pi(\gamma(0))$$

A contradiction to $a < \text{inj}(M)$ since $\pi \circ \gamma$ is a geodesic. \square

Corollary 11.3.1. Suppose M and N are compact smooth manifold, then $M \times N$ doesn't admit a Riemannian metric with negative sectional curvature.

Proof. If $M \times N$ admits a Riemannian metric with negative sectional curvature, Cartan's torsion-free theorem implies $\pi_1(M \times N)$ is torsion-free, thus for arbitrary $\alpha \in \pi_1(M), \beta \in \pi_1(N)$, unless either M or N is simply-connected, $\pi_1(M \times N)$ will contain an abelian subgroup $\mathbb{Z} \times \mathbb{Z}$ generated by α, β , which contradicts to Preissmann's theorem.

So we may assume M is simply-connected, then consider the universal covering $M \times \tilde{N}$ of $M \times N$, Cartan-Hadamard's theorem implies it's diffeomorphic to \mathbb{R}^n for some positive integer n , but M is orientable since it's simply-connected, so $H^m(M) = \mathbb{Z}$ where $m = \dim M$. So by Künneth formula $H^n(M \times \tilde{N}) \neq 0$, a contradiction to Poincaré lemma. \square

Lemma 11.3.5. Let (M, g) be a complete Riemannian manifold with non-positive sectional curvature and $\pi : \tilde{M} \rightarrow M$ the universal covering. If $\gamma : \mathbb{R} \rightarrow \tilde{M}$ is a common axis for all elements of $\text{Aut}_\pi(\tilde{M})$, then M is not compact.

Proof. \square

Theorem 11.3.2 (Preissmann). If (M, g) is a compact Riemannian manifold with negative sectional curvature, then $\pi_1(M)$ is not abelian.

Proof. Suppose $\pi_1(M)$ is abelian, then let γ be the axis of some deck transformation, then it's the axis of all deck transformations since $\pi_1(M)$ is abelian, which implies M is non-compact, a contradiction. \square

Theorem 11.3.3 (Byers). If (M, g) is a compact Riemannian manifold with negative sectional curvature, then any non-trivial solvable subgroup of $\pi_1(M)$ is isomorphic to \mathbb{Z} .

11.4. Other facts.

Theorem 11.4.1 (Yau). Let (M, g) be a compact Riemannian manifold with non-positive sectional curvature. If $\pi_1(M)$ is solvable, then M is flat and isometric to some \mathbb{R}^n/Γ .

Theorem 11.4.2 (Farrell-Jones). Let $(M_i, g_i), i = 1, 2$ be two compact Riemannian manifolds with non-positive sectional curvature. If $\pi_1(M_1) = \pi_1(M_2)$ then M_1 and M_2 are homeomorphic.

12. TOPOLOGY OF POSITIVE CURVATURE MANIFOLD

12.1. Myers' theorem.

Theorem 12.1.1 (Myers). Let (M, g) be a complete Riemannian manifold with $\text{Ric}(g) \geq \frac{n-1}{R^2}g$ where $n = \dim M$, then

1. $\text{diam}(M) \leq \pi R$;
2. M is compact.

Proof. For (1). If $\text{diam}(M) > \pi R$, then there exists $l > \pi R$ and a (locally) minimal geodesic $\gamma : [0, l] \rightarrow M$ of unit-speed, since M is complete. Choose a parallel orthonormal basis $\{e_1(t) = \gamma'(t), e_2(t), \dots, e_n(t)\}$ with $e_1(t), \dots, e_n(t)$ along γ and set

$$V_i(t) = \sin\left(\frac{\pi t}{l}\right)e_i(t), i = 2, \dots, n$$

It's clear $V_i(0) = V_i(l) = 0$ for $i \geq 2$. Let $\alpha : (-\varepsilon, \varepsilon) \rightarrow M$ be a variation of γ with variation field $V(t) = \sum_{i=2}^n V_i(t)$, then by second variation formula we have

$$\begin{aligned} \left. \frac{d^2 L(\alpha(t, s))}{ds^2} \right|_{s=0} &= \sum_{i=2}^n \int_0^l \langle \widehat{\nabla}_{\frac{d}{dt}} V_i, \widehat{\nabla}_{\frac{d}{dt}} V_i \rangle dt - \sum_{i=2}^n \int_0^l R(V_i, \gamma', \gamma', V_i) dt \\ &= \sum_{i=2}^n \int_0^l \left(\frac{\pi}{l}\right)^2 \cos^2\left(\frac{\pi t}{l}\right) dt - \sum_{i=2}^n \int_0^l \sin^2\left(\frac{\pi}{l}\right) R(e_i, e_1, e_1, e_i) dt \\ &\leq (n-1) \left(\frac{\pi}{l}\right)^2 \int_0^l \cos^2\left(\frac{\pi t}{l}\right) dt - \frac{(n-1)}{R^2} \int_0^l \sin^2\left(\frac{\pi t}{l}\right) dt \\ &< 0 \end{aligned}$$

A contradiction to γ is minimal. (2) follows from (1). \square

Corollary 12.1.1. Let M be a complete Riemannian manifold with positive Ricci curvature, then the universal covering of M is compact. In particular, the fundamental group $\pi_1(M)$ is finite.

Proof. Endow the universal covering \widetilde{M} with pullback metric, thus \widetilde{M} is a complete Riemannian manifold with positive Ricci curvature, thus \widetilde{M} is compact, which implies $\pi : \widetilde{M} \rightarrow M$ is a finite covering, thus $\pi_1(M)$ is finite, since $|\pi_1(M)|$ equals the number of sheets of covering. \square

Corollary 12.1.2. Let (M, g) be a complete Riemannian manifold with sectional curvature $K \geq \frac{1}{R^2}$, then M is compact and $\text{diam}(M) \leq \pi R$ and $\pi_1(M)$ is finite.

Remark 12.1.1. The estimate for the diameter given by Myers's theorem can't be improved. Indeed, the unit sphere $S^n \subset \mathbb{R}^{n+1}$ has constant sectional curvature $K = 1$ and $\text{diam}(S^n) = \pi$. A surprising theorem is that this example is unique in the following sense: Let (M, g) be a complete Riemannian manifold with dimension n , $\text{Ric}(g) \geq \frac{n-1}{R^2}g$ and $\text{diam}(M) = \pi R$, then M is isometric to sphere $S^n(R)$.

12.2. Synge's theorem.

Lemma 12.2.1. Let A be an orthogonal linear transformation of \mathbb{R}^{n-1} and suppose $\det A = (-1)^n$. Then 1 is an eigenvalue of A .

Proof. If n is even, then $\det(\lambda I - A)$ is a polynomial of odd degree, therefore A has at least a real eigenvalue, and it must be ± 1 since A is orthogonal. Furthermore, since $\det A = 1$ and the product of complex eigenvalue is positive, there is at least a real eigenvalue which equals 1.

If n is odd, then $\det A = -1$. Because the product of complex eigenvalue is positive, there are at least two real eigenvalues, and one of them is 1. \square

Theorem 12.2.1 (Synge). Let (M, g) be a compact Riemannian manifold with $K > 0$, then

1. If $\dim M$ is even and M is orientable, then M is simply-connected;
2. If $\dim M$ is odd, then M is orientable;
3. If M is even and M is not orientable, then $\pi_1(M) = \mathbb{Z}_2$.

Proof. Suppose $\pi : \widetilde{M} \rightarrow M$ is the universal covering with pullback metric $\widetilde{g} = \pi^*g$

1. If $\dim M$ is even, give \widetilde{M} the pullback orientation;
2. If $\dim M$ is odd, give \widetilde{M} arbitrary orientation.

Suppose the conclusions are not correct, thus $\pi_1(M)$ is non-trivial. Choose a non-trivial deck transformation $F : \widetilde{M} \rightarrow \widetilde{M}$ such that

1. If $\dim M$ is even, F is orientation preserving;
2. If $\dim M$ is odd, F is orientation reserving.

By Lemma, there exists an axis $\widetilde{\gamma} : \mathbb{R} \rightarrow \widetilde{M}$ for F and $\gamma = \pi \circ \widetilde{\gamma}$ is a closed geodesic in M that minimizes the length in $[\gamma]$,

$$F(\widetilde{\gamma}(t)) = \widetilde{\gamma}(t + 1)$$

\square

Example 12.1. $\mathbb{RP}^2 \times \mathbb{RP}^2$ admits no metric with positive sectional curvature, since its fundamental group is $\mathbb{Z}_2 \times \mathbb{Z}_2$.

Conjecture 12.1 (Hopf conjecture). Does $S^2 \times S^2$ admit a metric with positive sectional curvature?

12.3. Other facts.

Theorem 12.3.1. Let (M, g) be a simply-connected compact Riemannian manifold, then

1. (Hamilton) $\dim M = 3$, $\text{Ric}(g) > 0$, then M is diffeomorphism to S^3 .
2. (Hamilton) If $\dim M = 4$ with curvature operator > 0 , then M is diffeomorphism to S^4 .
3. (Böhm-Wilking) If curvature operator > 0 , then M is diffeomorphism to S^n .
4. (Brendle-Schoen) If sectional curvature satisfies $\frac{1}{4} < K \leq 1$, then M is diffeomorphism to S^n .

13. TOPOLOGY OF CONSTANT SECTIONAL CURVATURE MANIFOLD

13.1. Isometry.

Theorem 13.1.1. Let $(M, g), (\widetilde{M}, \widetilde{g})$ be two Riemannian manifold, $\varphi : M \rightarrow \widetilde{M}$ is a bijective. Then the followings are equivalent:

1. φ is an isometry, that is

$$\text{dist}_g(p, q) = \text{dist}_{\widetilde{g}}(\varphi(p), \varphi(q))$$

2. φ is a diffeomorphism and $(\varphi_*)_p : T_p M \rightarrow T_{\varphi(p)} \widetilde{M}$ is a linear isometry for all $p \in M$;

Theorem 13.1.2. $\varphi : (M, g) \rightarrow (\widetilde{M}, \widetilde{g})$ is a smooth map. Then the followings are equivalent:

1. φ is a local isometry;
2. For all $p \in M$, there exists $U \subset M$ containing p such that

$$\varphi|_U : U \rightarrow \varphi(U)$$

is an isometry.

Theorem 13.1.3 (Cartan-Ambrose-Hicks). Let (M, g) and $(\widetilde{M}, \widetilde{g})$ be two Riemannian manifold $p \in M, \widetilde{p} \in \widetilde{M}$ and $\Phi_0 : T_p M \rightarrow T_{\widetilde{p}} \widetilde{M}$ some fixed linear isometry. Suppose $\delta \in (0, \min\{\text{inj}_p(M), \text{inj}_{\widetilde{p}}(\widetilde{M})\})$. Then the followings are equivalent:

1. There exists an isometry $\varphi : B(p, \delta) \rightarrow B(\widetilde{p}, \delta)$ such that $\varphi(p) = \widetilde{p}$ and $(\varphi_*)_p = \Phi_0$;
2. If $v \in T_p M, |v| < \delta, \gamma(t) = \exp_p(tv), \widetilde{\gamma}(t) = \exp_{\widetilde{p}}(t\Phi_0 v)$ and

$$\Phi_t = P_{0,t}^{\widetilde{\gamma}} \circ \Phi_0 \circ P_{t,0}^{\gamma} : T_{\gamma(t)} M \rightarrow T_{\widetilde{\gamma}(t)} \widetilde{M}$$

Then Φ_t satisfies

$$R(u, v, w, z) = \widetilde{R}(\Phi_t u, \Phi_t v, \Phi_t w, \Phi_t z)$$

Proof. From (1) to (2). If we can show $\Phi_t = (\varphi_*)_{\gamma(t)}$, then it's clear that Φ_t preserves curvature, since φ is an isometry. By definition of Φ_t , it suffices to show the following diagram commutes

$$\begin{array}{ccc} T_p M & \xrightarrow{(\varphi_*)_p} & T_{\widetilde{p}} \widetilde{M} \\ \downarrow P_{0,t}^{\gamma} & & \downarrow P_{0,t}^{\widetilde{\gamma}} \\ T_{\gamma(t)} M & \xrightarrow{(\varphi_*)_{\gamma(t)}} & T_{\widetilde{\gamma}(t)} \widetilde{M} \end{array}$$

since $(\varphi_*)_p = \Phi_0$. Note that

$$\varphi(\gamma(t)) = \widetilde{\gamma}(t)$$

since they agree at $t = 0$, so do their derivatives. So it's tautological that

$$P_{0,t}^{\varphi \circ \gamma} \circ (\varphi_*)_p(v) = (\varphi_*)_{\gamma(t)} \circ P_{0,t}^{\gamma}(v)$$

where $v = \gamma'(0)$, since $P_{0,t}^\gamma(v) = \gamma'(t)$, $(\varphi_*)_{\gamma(t)}(\gamma'(t)) = (\varphi \circ \gamma)'(t) = P_{0,t}^{\varphi \circ \gamma} \circ (\varphi_*)_p(v)$ ⁶. Now consider $w \in T_p M$ which is not parallel to $v = \gamma'(0)$. Since both $(\varphi_*)_{\gamma(t)}$ and parallel transport preserve angles, so $P_{0,t}^{\varphi \circ \gamma} \circ (\varphi_*)_p(w)$ and $(\varphi_*)_{\gamma(t)} \circ P_{0,t}^\gamma(w)$ has the same angle with $(\varphi_*)_{\gamma(t)}(\gamma'(t))$, and they have the same length, so they're equal.

From (2) to (1). Define

$$\varphi = \exp_{\tilde{p}} \circ \Phi_0 \circ \exp_p^{-1}$$

It suffices to show for any $q \in B(p, \delta)$,

$$(\varphi_*)_q : T_q M \rightarrow T_{\varphi(q)} \widetilde{M}$$

is a linear isometry. For any $w \in T_q M$, there exists a geodesic $\gamma : [0, 1] \rightarrow M$ with $\gamma(0) = p, \gamma(1) = q$ and a Jacobi field J such that $J(0) = 0, J(1) = w$ along γ . We claim:

1. **Claim 1:** $\tilde{J}(t) = \Phi_t(J(t))$ is a Jacobi field;
2. **Claim 2:** $\tilde{J}(1) = (\varphi_*)_q(J(1))$.

From claim 2 we have

$$|(\varphi_*)_q(w)| = |\tilde{J}(1)| = |J(1)| = |w|$$

which completes the proof. Now let's give proofs of these two claims.

1. **Proof of Claim 1:** Given an orthonormal $\{e_1(0) = \frac{\gamma'(0)}{|\gamma'(0)|}, e_2(0), \dots, e_n(0)\}$ of $T_p M$ and use parallel transport to obtain a parallel frame along γ . With respect to this frame we can write $J(t) = J^i(t)e_i(t)$, then $\tilde{J}(t) = J^i(t)\tilde{e}_i(t)$, where $\tilde{e}_i(t) = \Phi_t(e_i(t))$. Furthermore, $\tilde{e}_i(t)$ is also a parallel frame by definition of Φ_t . Then $\tilde{J}(t)$ is a Jacobi field, since

$$\begin{aligned} & \frac{d^2 J^j}{dt^2} + J^i(t)|\tilde{\gamma}(t)|^2 \tilde{R}(\tilde{e}_i(t), \tilde{e}_1(t), \tilde{e}_1(t), \tilde{e}_j(t)) \\ &= \frac{d^2 J^j}{dt^2} + J^i(t)|\gamma(t)|^2 R(e_i(t), e_1(t), e_1(t), e_j(t)) \\ &= 0 \end{aligned}$$

holds for arbitrary j , where we use the fact Φ_t preserves the length and curvature, and $J(t)$ is a Jacobi field.

2. **Proof of Claim 2:**

□

Theorem 13.1.4. Let (M, g) be a connected manifold. Suppose φ and ψ are two local isometries from (M, g) to $(\widetilde{M}, \tilde{g})$. If there exists $p \in M$ such that

$$\begin{aligned} \varphi(p) &= \psi(p) \\ (d\varphi)_p &= (d\psi)_p \end{aligned}$$

Then $\varphi = \psi$.

⁶These identities hold since both γ and $\varphi \circ \gamma$ are geodesics.

Proof. Suppose $\varphi|_V, \psi|_V$ is diffeomorphism and V is a geodesic ball, then

$$f := (\varphi^{-1} \circ \psi)|_V : V \rightarrow V$$

satisfies $f(p) = p, (f_*)_p = \text{id}$. Given $q \in V$, there exists unique $v \in T_p M$ such that $\exp_p(v) = q$, then

$$f(q) = \exp_p \circ \text{id} \circ \exp_p^{-1}(q) = q$$

which implies φ agrees with ψ in V . Consider the following set

$$A = \{p \in M \mid \psi(p) = \varphi(p)\}$$

Above argument shows it's open, and it's clearly closed, then $A = M$, since M is connected. This completes the proof. \square

13.2. Hopf's theorem.

Theorem 13.2.1 (Hopf). Let (M, g) be a simply-connected complete Riemannian manifold with constant sectional curvature K , then (M, g) is isometric to $(\widetilde{M}, g_{\text{can}})$, where

$$\widetilde{M} = \begin{cases} \mathbb{S}^n(\frac{1}{\sqrt{K}}), & K > 0 \\ \mathbb{R}^n, & K = 0 \\ \mathbb{H}^n(\frac{1}{\sqrt{-K}}), & K < 0 \end{cases}$$

Proof. Let M be a simply-connected complete Riemannian manifold with constant sectional curvature K .

1. If $K \leq 0$, let $\widetilde{M} = \mathbb{R}^n$ or $\mathbb{H}^n(r)$, where $r = \frac{1}{\sqrt{-K}}$. Fix $p \in M, \tilde{p} \in \widetilde{M}$ and a linear isometry $\Phi_0 : T_p M \rightarrow T_{\tilde{p}} \widetilde{M}$, then Cartan-Ambrose-Hicks's theorem implies

$$\varphi = \exp_{\tilde{p}} \circ \Phi_0 \circ \exp_p^{-1}$$

maps $B(p, \delta)$ to $B(\tilde{p}, \delta)$ is an isometry, where $0 < \delta < \min\{\text{inj}_p(M), \text{inj}_{\tilde{p}}(\widetilde{M})\}$.

Thus it's a local isometry defined on \widetilde{M} . Furthermore, Cartan-Hadamard's theorem implies φ is a diffeomorphism, since M, \widetilde{M} are simply-connected with non-positive sectional curvature. Combine these facts together we have φ is an isometry.

2. If $K > 0$, let $\widetilde{M} = \mathbb{S}^n(r)$, where $r = \frac{1}{\sqrt{K}}$. Fix $p \in M, \tilde{p} \in \widetilde{M}$ and a linear isometry $\Phi_0 : T_p M \rightarrow T_{\tilde{p}} \widetilde{M}$, then Cartan-Ambrose-Hicks's theorem implies

$$\varphi_1 = \exp_{\tilde{p}} \circ \Phi_0 \circ \exp_p^{-1}$$

is an isometry defined on $M \setminus \{-p\}$, since $-p$ is the only cut point of p . Choose another $q \in M \setminus \{p, -p\}$, $\tilde{q} = \varphi_1(q)$ and $\Psi_0 = (d\varphi_1)_q : T_q M \rightarrow T_{\tilde{q}} \widetilde{M}$, then Cartan-Ambrose-Hicks's theorem implies

$$\varphi_2 = \exp_{\tilde{q}} \circ \Psi_0 \circ \exp_q^{-1}$$

is an isometry defined on $M \setminus \{-q\}$ by the same reason. Note that

$$\begin{aligned}\varphi_2(q) &= \tilde{q} = \varphi_1(q) \\ (\mathrm{d}\varphi_2)_q &= \Psi_0 = (\mathrm{d}\varphi_1)_q\end{aligned}$$

So by Theorem 13.1.4, we have the φ_1 agrees with φ_2 on $M \setminus \{-p, -q\}$. Thus

$$\varphi(x) = \begin{cases} \varphi_1(x), & x \in M \setminus \{-p\} \\ \varphi_2(x), & x \in M \setminus \{-q\} \end{cases}$$

is an isometry from $M \rightarrow \widetilde{M}$.

□

Notation 13.2.1. We usually use $S(n, k)$ to denote the complete, simply-connected Riemannian manifold of dimension n with constant sectional curvature k , and call them space forms.

Example 13.2.1. Let (M, g) be a complete Riemannian manifold with constant sectional curvature $K = 1$. If $\dim M = 2m$, then (M, g) is isometric to the sphere $(\mathbb{S}^{2m}, g_{\text{can}})$ or the real projective space $(\mathbb{RP}^{2m}, g_{\text{can}})$.

Proof. Note that Hopf's theorem implies (M, g) is isometric to $(\mathbb{S}^{2m}/\Gamma, g_{\text{can}})$, where Γ is the fundamental group of M , and Synge's theorem implies if $\dim M$ is even and $K > 0$, then $\pi_1(M) = \{e\}$ or $\pi_1(M) = \mathbb{Z}_2$. Combine these two facts together we have M is isometric to $(\mathbb{S}^{2m}, g_{\text{can}})$ or $(\mathbb{RP}^{2m}, g_{\text{can}})$. □

Remark 13.2.1. In general, we have no ideal about what does $\pi_1(M)$ look like.

Proposition 13.2.1. Let (M, g) be a connected, simply connected, complete Riemannian manifold. The following are equivalent:

1. (M, g) has constant sectional curvature;
2. For every pair of point $p, q \in M$ and every linear isometry $\Phi_0 : T_p M \rightarrow T_q M$, there exists an isometry $\varphi : M \rightarrow M$ such that $\varphi(p) = q$, $(\varphi_*)_p = \Phi_0$.

Proof. (1) to (2) is already shown in the proof of Hopf's theorem. For (2) to (1). Firstly let's show for $p \in M$, sectional curvature $K_p(\sigma)$ at p is independent of σ , where σ is a 2-dimensional subspace of $T_p M$. For arbitrary two 2-dimensional subspaces σ_1, σ_2 of $T_p M$. By tricks of linear algebra it's easy to find an linear isometry Φ_0 such that $\Phi_0 \sigma_1 = \sigma_2$, then there exists an isometry $\varphi : M \rightarrow M$ such that $(\varphi_*)_p = \Phi_0$. Since isometry preserves curvature, in particular we have

$$R(x, y, y, x) = R(\Phi_0 x, \Phi_0 y, \Phi_0 y, \Phi_0 x)$$

where $\{x, y\}$ is a basis of σ_1 and $\{\Phi_0 x, \Phi_0 y\}$ is a basis of σ_2 , which implies $K_p(\sigma_1) = K_p(\sigma_2)$.

Now let's show for arbitrary $p, q \in M$, sectional curvature $K_p = K_q$ ⁷. For $p, q \in M$, 2-dimensional subspace σ_1, σ_2 of $T_p M, T_q M$ respectively, we also can find a linear isometry $\Phi_0 : T_p M \rightarrow T_q M$ such that $\Phi_0 \sigma_1 = \sigma_2$, and there also exists an isometry $\varphi : M \rightarrow M$ such that $(\varphi_*)_p = \Phi_0$, then the same argument shows $K_p = K_q$. \square

⁷In fact, if $\dim M \geq 3$, then Schur's lemma implies K_p is independent of p .

Part 5. Comparison theorems

14. PREPARATIONS

In this section we select some basic tools we will use in later computations. Unless otherwise specified, (M, g) is a Riemannian manifold and (x^i, U, p) is a normal coordinate centered at $p \in M$.

14.1. Radial vector field.

Definition 14.1.1 (radial distance function). The radial distance function r defined on U is given by

$$r(q) := \sqrt{\sum_{i=1}^n (q^i)^2}$$

where $q = (q^1, \dots, q^n)$ in normal coordinate (x^i, U, p) .

Definition 14.1.2 (radial vector field). The radial vector field in $U \setminus \{p\}$ is defined as

$$\partial_r = \frac{x^i}{r} \frac{\partial}{\partial x^i}$$

Proposition 14.1.1. The geodesic starting at p with unit-speed is the integral curve of radial vector field ∂_r over $U \setminus \{p\}$.

Proof. We need to show for geodesic $\gamma : I \rightarrow U$ with $\gamma(0) = p, \gamma'(0) = v$, where $|v| = 1$, we have

$$\gamma'(b) = \partial_r|_{\gamma(b)}$$

where I is an open interval and $b \in I$.

In normal coordinate γ looks like $\gamma(t) = (tv^1, \dots, tv^n)$. If we denote $\gamma(b) = q = (q^1, \dots, q^n)$, then it's clear $v^i = q^i/b$. Furthermore, $r(q) = b$, since $|v| = 1$. Then in normal coordinate,

$$\gamma'(b) = v^i \frac{\partial}{\partial x^i} \Big|_q = \frac{q^i}{b} \frac{\partial}{\partial x^i} \Big|_q = \frac{q^i}{r(q)} \frac{\partial}{\partial x^i} \Big|_q = \partial_r|_q$$

□

Lemma 14.1.1. Given a smooth function $f : M \rightarrow \mathbb{R}$ and X is a vector field, if

1. $Xf = |X|^2$;
2. X is perpendicular to the level set of f .

then $X = \nabla f$.

Theorem 14.1.1 (Gauss lemma). Properties of radial vector fields:

1. $|\partial_r|^2 = 1$;
2. $g^{ij} \frac{\partial r}{\partial x^i} \frac{\partial}{\partial x^j} = \nabla r = \partial_r$.

Proof. For (1). It's clear, since we have already shown geodesic with unit-speed is integral curve of ∂_r .

For (2). In order to apply Lemma 14.1.1, we consider $X = \partial_r$ to $f = r$, then

$$Xr = \frac{x^i}{r} \frac{\partial r}{\partial x^i} = \sum_{i=1}^n \frac{(x^i)^2}{r^2} = 1 = |\partial_r|^2$$

This shows the first condition in above lemma. For any $q \in U \setminus \{p\}$ we write it as $q = (q^1, \dots, q^n)$ in normal coordinate with $b = r(q)$. Given $w \in T_q M$ which is tangent to the level set of r , there exists $c(s) : (-\varepsilon, \varepsilon) \rightarrow M$ such that $c(0) = q, c'(0) = w$ with $\sum_{i=1}^n (c^i(s))^2 = b$, where c^i is the coordinates of c in normal coordinate. Taking derivative with respect to s we obtain

$$\sum_{i=1}^n 2c^i(s)(c^i(s))' = 0$$

We're almost there, since $w = (c^i(0))' \frac{\partial}{\partial x^i} \Big|_q, \partial_r|_q = \frac{c^j(0)}{b} \frac{\partial}{\partial x^j} \Big|_q$ and if metric at $T_q M$ is standard, then we're done. However, we only know metric at $T_p M$ is standard, so we may use parallel transport to transport w to $T_p M$ and show they're perpendicular in $T_p M$, which implies they're perpendicular in $T_q M$, since geodesic is integral curve of ∂_r . \square

Corollary 14.1.1. The following identities hold in (x^i, U, p) :

1. $g_{ij}x^j = x^i$;
2. $g_{im} = \delta_{im} - \frac{\partial g_{ij}}{\partial x^m} x^j$;
3. $\frac{\partial g_{ij}}{\partial x^m} x^j = \frac{\partial g_{mj}}{\partial x^i} x^j$;
4. $\frac{\partial g_{ij}}{\partial x^m} x^j x^i = \frac{\partial g_{mj}}{\partial x^i} x^j x^i = 0$;
5. $\Gamma_{ij}^k x^i x^j = 0$;
6. $\nabla_{\partial_r} \partial_r = 0$ in $U \setminus \{p\}$.

Proof. For (1). On one hand by Theorem 14.1.1 we have $\partial_r = \nabla r = g^{ij} \frac{x^i}{r} \frac{\partial}{\partial x^j}$; On the other hand by definition of ∂_r we have $\partial_r = \frac{x^j}{r} \frac{\partial}{\partial x^j}$, which implies

$$g^{ij} x^i = x^j$$

This shows (1).

For (2). Take partial derivatives of (1) with respect to x^m , we have

$$\frac{\partial g_{ij}}{\partial x^m} x^j + g_{ij} \delta_{jm} = \delta_{im}$$

This shows (2).

For (3). It follows from (2), since g_{im}, δ_{im} are symmetric in i, m .

For (4). It follows from (1) and (2), since

$$\begin{aligned} \frac{\partial g_{ij}}{\partial x^m} x^j x^i &\stackrel{(2)}{=} (\delta_{im} - g_{im}) x^i = x^m - g_{im} x^i \stackrel{(1)}{=} 0 \\ \frac{\partial g_{mj}}{\partial x^i} x^j x^i &\stackrel{(2)}{=} (\delta_{mi} - g_{mi}) x^i = x^m - g_{im} x^i \stackrel{(1)}{=} 0 \end{aligned}$$

For (5). It follows from (4) and

$$\Gamma_{ij}^k = \frac{1}{2}g^{mk}\left(\frac{\partial g_{mj}}{\partial x^i} + \frac{\partial g_{im}}{\partial x^j} - \frac{\partial g_{ij}}{\partial x^m}\right)$$

For (6). Direct computation shows

$$\begin{aligned}\nabla_{\partial_r}\partial_r &= \frac{x^k}{r}\nabla_{\frac{\partial}{\partial x^k}}\left(g^{ij}\frac{x^i}{r}\frac{\partial}{\partial x^j}\right) \\ &= g^{ij}\frac{x^k}{r}\underbrace{\left\{\left(\frac{\delta_{ki}}{r} - \frac{x^k x^i}{r^3}\right)\frac{\partial}{\partial x^j}\right\}}_{\text{part I}} + \underbrace{\left\{\frac{x^i}{r}\Gamma_{kj}^m\frac{\partial}{\partial x^m}\right\}}_{\text{part II}}\end{aligned}$$

By (1) and (5) we have

$$g^{ij}\frac{x^k x^i}{r}\Gamma_{kj}^m = \frac{1}{r}\Gamma_{kj}^m x^k x^j = 0$$

which implies part II is zero. For part I, we have

$$\frac{1}{r^2}(g^{ij}x^k\delta_{ki} - \frac{(x^k)^2}{r^2}g^{ij}x^i) = \frac{1}{r^2}(g^{ij}x^i - g^{ij}x^i) = 0$$

□

Remark 14.1.1. Note that we firstly establish the fact unit-speed geodesic is integral curve of ∂_r and show $\partial_r = \nabla r$, then we obtain lots of identities. In particular we have $\nabla_{\partial_r}\partial_r = 0$, which also implies unit-speed geodesic is integral curve of ∂_r . This shows over $U \setminus \{p\}$ the following statements are equivalent:

1. The unit-speed geodesic is integral curve of ∂_r ;
2. $g^{ij}x^i = x^j$;
3. $\nabla_{\partial_r}\partial_r = 0$.

14.2. Jacobi fields on constant sectional curvature manifold.

Proposition 14.2.1. Let (M, g) be a Riemannian manifold with constant sectional curvature k and $\gamma : [0, b] \rightarrow M$ be a unit-speed geodesic. Then the normal Jacobi field with $J(0) = 0$ is of the form

$$J(t) = m \operatorname{sn}_k(t) E(t)$$

where

1. The constant m is determined by $J'(0) = mE(0)$;
- 2.

$$\operatorname{sn}_k(t) = \begin{cases} t, & k = 0 \\ \frac{\sin(\sqrt{k}t)}{\sqrt{k}}, & k > 0 \\ \frac{\sinh(\sqrt{-k}t)}{\sqrt{-k}}, & k < 0 \end{cases}$$

3. $E(t)$ is a normal parallel vector field along γ with $|E(t)| = 1$

Proof. Since (M, g) has constant sectional curvature k , thus $R_{ijkl} = k(g_{il}g_{jk} - g_{ik}g_{jl})$, so for any normal vector field J along γ we have

$$\begin{aligned} R(J, \gamma', \gamma', W) &= k(\langle J, W \rangle \langle \gamma', \gamma' \rangle - \langle J, \gamma' \rangle \langle \gamma', W \rangle) \\ &= k \langle J, W \rangle \end{aligned}$$

which implies

$$R(J, \gamma')\gamma' = kJ$$

since γ is unit-speed and J is normal. Thus equation for Jacobi field can be written as

$$0 = J'' + kJ$$

Assume $J = u(t)E(t)$, then

$$(u''(t) + ku(t))E(t) = 0$$

So if we want to find normal Jacobi fields J , it suffices to solve

$$\begin{cases} u''(t) + ku = 0 \\ u(0) = 0 \end{cases}$$

and it's clear $\text{sn}_k(t)$ is solution of this ODE. \square

14.3. Polar decomposition of metric with constant sectional curvature. Let $\pi : \mathbb{R}^n \setminus \{0\} \rightarrow \mathbb{S}^{n-1}$ given by $\pi(x) = x/|x|$. We can use π to pullback canonical metric on \mathbb{S}^{n-1} , and still use $g_{\mathbb{S}^{n-1}}$ to denote it.

Lemma 14.3.1. Let \bar{g} be the Euclidean metric on $\mathbb{R}^n \setminus \{0\}$, then

$$\bar{g} = dr \otimes dr + r^2 g_{\mathbb{S}^{n-1}}$$

where $r(x) = |x|$.

Theorem 14.3.1 (polar decomposition). Let (x^i, U, p) be a normal coordinate centered at $p \in S(n, k)$, then in U the metric g can be written as

$$g = dr \otimes dr + \text{sn}_k^2(r) g_{\mathbb{S}^{n-1}}$$

where r is radial distance function.

Proof. We use g_c to denote metric $dr \otimes dr + \text{sn}_k^2(r) g_{\mathbb{S}^{n-1}}$ and \bar{g} to denote standard metric on Euclidean space. By Theorem 14.1.1, we have

$$g(\partial_r, \partial_r) = 1 = g_c(\partial_r, \partial_r)$$

So it remains to show for each $q \in U \setminus \{p\}$ and $w_1, w_2 \in T_q M$ such that $g(w_i, \partial_r|_q) = 0, i = 1, 2$, we have

$$g(w_1, w_2) = g_c(w_1, w_2)$$

By polarization it suffices to show that $g(w, w) = g_c(w, w)$ for every such vector w .

Suppose $\text{dist}(p, q) = b$, on one hand we have

$$|w|_{g_c}^2 \stackrel{(1)}{=} \text{sn}_k^2(b) |w|_{g_{\mathbb{S}^{n-1}}}^2 \stackrel{(2)}{=} \frac{\text{sn}_k(b)}{b^2} |w|_{\bar{g}}^2$$

where

- (1) holds from definition of g_c ;
- (2) holds from polar decomposition of standard metric of Euclidean space, that is Lemma 14.3.1.

On the other hand, let $\gamma : [0, b] \rightarrow U$ be a unit-speed geodesic connecting p, q , and we can write it with respect to normal coordinate U as

$$\gamma(t) = \left(\frac{tq^1}{b}, \dots, \frac{tq^n}{b} \right)$$

where $q = (q^1, \dots, q^n)$ in normal coordinate U . Let J be a Jacobi field such that $J(0) = 0, J(b) = w$, then we have

$$|w|_g^2 = |J(b)|_g^2 \stackrel{(3)}{=} \text{sn}_k^2(b) |J'(0)|_g^2 \stackrel{(4)}{=} \text{sn}_k^2(b) |J'(0)|_g^2$$

where

- (3) holds from the fact Jacobi field on constant sectional curvature space is of form $J(t) = |J'(0)| \text{sn}_k(t) E(t)$;
- (4) holds from the metric on $T_p M$ is standard metric in normal coordinate.

Furthermore, suppose $J'(0) = a$, then we can write it as $J(t) = \alpha_* \left(\frac{\partial}{\partial s} \right) \Big|_{s=0}$, where

$$\alpha(s, t) = \exp_p(t(\gamma'(0) + sJ'(0)))$$

In normal coordinate we can write $\alpha(s, t)$ explicitly as

$$\alpha(s, t) = \left(\frac{tq^1}{b} + tsa^1, \dots, \frac{tq^n}{b} + tsa^n \right)$$

thus $J(t) = ta^i \frac{\partial}{\partial x^i} \Big|_{\gamma(t)}$. We can conclude $a^i = \frac{w^i}{b}$ by setting $t = b$, in particular we have $J'(0) = \frac{w^i}{b} \frac{\partial}{\partial x^i} \Big|_p$. Then

$$\text{sn}_k^2(b) |J'(0)|_g^2 = \text{sn}_k^2(b) \frac{|w|_g^2}{b^2} = |w|_{g_c}^2$$

□

Remark 14.3.1. Note that there are four points we need in the proof of above theorem, and the **key point** is (3), that is Jacobi field of $S(n, k)$ has the form of

$$J(t) = m \text{sn}_k(t) E(t)$$

So this motivate us that if on a normal neighborhood of some point, the Jacobi field has the above form, then metric g can be written as

$$g = dr \otimes dr + \text{sn}_k(r)^2 g_{\mathbb{S}^{n-1}}$$

in U , and thus has constant sectional curvature k .

14.4. A criterion for constant sectional curvature space. Recall that for a smooth function $f : M \rightarrow \mathbb{R}$, $\text{Hess } f$ is a $(0, 2)$ -tensor, we use \mathcal{H}_f to denote its $(1, 1)$ -type, that is

$$g(\mathcal{H}_f(X), Y) = \text{Hess } f(X, Y)$$

where X, Y are two vector fields.

In particular, if r is the radial distance function on a normal coordinate, then Hessian r is a $(2, 0)$ -tensor, that is $\nabla^2 r$, then we have

$$\mathcal{H}_r = \nabla \partial_r$$

since $(1, 0)$ -type of ∇r is ∂_r .

Proposition 14.4.1. Let (M, g) be a complete Riemannian manifold, (x^i, U, p) a normal coordinate centered at p and r the radial distance function on U . If $\gamma : [0, b] \rightarrow M$ is unit-speed geodesic with $\gamma(0) = p, \gamma'(0) = v \in T_p M$, and J is a normal Jacobi field along γ with $J(0) = 0$. Then for all $t \in (0, b]$

$$\mathcal{H}_r(J(t)) = J'(t)$$

$$\mathcal{H}_r(\gamma'(t)) = 0$$

Proof. Here we only prove the first identity, the second can be computed in the same method. Let $J'(0) = w$, then $J(t) = tw^i \frac{\partial}{\partial x^i} \Big|_{\gamma(t)}$,

$$\begin{aligned} J'(t) &= \widehat{\nabla}_{\frac{d}{dt}} \left(tw^i \frac{\partial}{\partial x^i} \Big|_{\gamma(t)} \right) \\ &= w^i \frac{\partial}{\partial x^i} \Big|_{\gamma(t)} + tw^i \widehat{\nabla}_{\frac{d}{dt}} \frac{\partial}{\partial x^i} \Big|_{\gamma(t)} \\ &= w^i \frac{\partial}{\partial x^i} \Big|_{\gamma(t)} + tw^i \Gamma_{ij}^k(\gamma(t)) \frac{d\gamma^j}{dt} \frac{\partial}{\partial x^k} \Big|_{\gamma(t)} \\ &= (w^k + tw^i v^j \Gamma_{ij}^k(\gamma(t))) \frac{\partial}{\partial x^k} \Big|_{\gamma(t)} \\ \mathcal{H}_r(J(t)) &= \nabla_{J(t)} \partial_r \\ &= \nabla_{tw^i \frac{\partial}{\partial x^i} \Big|_{\gamma(t)}} \left(\frac{x^j}{r} \frac{\partial}{\partial x^j} \right) \\ &= tw^i \nabla_{\frac{\partial}{\partial x^i} \Big|_{\gamma(t)}} \left(\frac{x^j}{r} \frac{\partial}{\partial x^j} \right) \\ &= tw^i \frac{x^j}{r} \Gamma_{ij}^k(\gamma(t)) \frac{\partial}{\partial x^k} \Big|_{\gamma(t)} + \sum_{i=1}^n tw^i \left(\frac{\delta_{ij}}{r} - \frac{x^i x^j}{r^3} \right) \frac{\partial}{\partial x^j} \Big|_{\gamma(t)} \end{aligned}$$

However, we have the following observations:

1. $r(\gamma(t)) = t$;
2. $x^i = tv^i$;
3. $\sum_{i=1}^n a^i v^i = 0$

where the last equality holds since J is a normal vector field, then

$$0 = \langle J(t), \gamma'(t) \rangle = \langle J(0), \gamma'(0) \rangle + \langle J'(0), \gamma'(0) \rangle t$$

implies $\langle J'(0), \gamma'(0) \rangle = \sum_{i=1}^n a^i v^i = 0$. \square

Corollary 14.4.1. With the same assumption as above proposition, for any vector field W along γ with $W(0) = 0$,

$$\begin{aligned} \text{Hess } r(J(s), W(s)) &\stackrel{(1)}{=} g(\mathcal{H}_r(J(s), W(s))) \\ &\stackrel{(2)}{=} g(J'(t), W(s)) \\ &\stackrel{(3)}{=} \int_0^s \langle J'(t), W(t) \rangle' dt \\ &\stackrel{(4)}{=} \int_0^s \langle J'(t), W'(t) \rangle - R(J, \gamma', \gamma', W) dt \end{aligned}$$

Corollary 14.4.2. Let $p \in U \subset S(n, k)$, where U is a normal neighborhood of p , then the following holds in $U \setminus \{p\}$

$$\mathcal{H}_r = \frac{\text{sn}'_k(r)}{\text{sn}_k(r)} \pi_r$$

where r is the radial distance function on U , and for each $q \in U \setminus \{p\}$, $\pi_r : T_q M \rightarrow T_q M$ is the orthogonal projection onto the orthogonal complement of $\partial_r|_q$.

Proof. For $p \in U \setminus \{q\}$, it's clear

$$\mathcal{H}_r(\partial_r|_q) = 0 = \frac{\text{sn}'_k(r)}{\text{sn}_k(r)} \pi_r(\partial_r|_q)$$

For $w \in T_q M$ such that $g(w, \partial_r|_q) = 0$, choose a unit-speed geodesic $\gamma : [0, b] \rightarrow M$ connecting p and q and $J(t)$ is the Jacobi field such that $J(0) = 0$, $J(b) = w$. Then we must have

$$J(t) = m \text{sn}_k(t) E(t)$$

where $E(t)$ is a normal parallel vector field along γ with $|E(t)| = 1$. Then

$$\begin{aligned} m \text{sn}'_k(t) E(t) &= J'(t) \\ &= \mathcal{H}_r(J(t)) \\ &= \mathcal{H}_r(m \text{sn}_k(t) E(t)) \\ &= m \text{sn}_k(t) \mathcal{H}_r(E(t)) \end{aligned}$$

Setting $t = b$ and dividing by $\text{sn}_k(b)$ one has

$$\mathcal{H}_r(E(b)) = \frac{\text{sn}'_k(b)}{\text{sn}_k(b)} E(b)$$

Note that $w = m \text{sn}_k(b) E(b)$, this completes the proof. \square

Furthermore, the converse of above corollary still holds:

Proposition 14.4.2. Let (M, g) be a Riemannian manifold and U a normal neighborhood of $p \in M$, r radial distance function. If

$$\mathcal{H}_r = \frac{\text{sn}'_k(r)}{\text{sn}_k(r)} \pi_r$$

holds in $U \setminus \{p\}$, then (M, g) has constant sectional curvature k in U .

Proof. Let $\gamma : [0, b] \rightarrow U$ be a unit-speed geodesic $r(0) = p$, J is a normal Jacobi vector field along γ with $J(0) = 0$, then $\mathcal{H}_r(J) = J'$ implies

$$J'(t) = \frac{\text{sn}'_k(t)}{\text{sn}_k(t)} J(t)$$

holds for $t \in (0, b]$, that is

$$\left(\frac{J(t)}{\text{sn}_k(t)} \right)' = 0$$

holds for $t \in (0, b]$. So we can write every normal Jacobi fields as $J(t) = m \text{sn}_k(t) E(t)$, where E is normal a parallel vector field with $|E| = 1$ and $t \in [0, b]$. Thus by Remark 14.3.1, we have g has constant sectional curvature k in U . \square

Remark 14.4.1. For convenience, we record the exact formulas for the quotient $\frac{\text{sn}'_k}{\text{sn}_k}$ as follows

$$\frac{\text{sn}'_k(t)}{\text{sn}_k(t)} = \begin{cases} \frac{1}{t}, & k = 0 \\ \frac{1}{\sqrt{k}} \cot \frac{t}{\sqrt{k}}, & k > 0 \\ \frac{1}{\sqrt{k}} \coth \frac{t}{\sqrt{k}}, & k < 0 \end{cases}$$

and we can draw the graph as follows.

15. COMPARISON THEOREMS BASED ON SECTIONAL CURVATURE

In this section, we will see the following philosophy: “The larger curvature is, the smaller the distance is.”

15.1. Rauch comparison. Rauch comparison theorem is one of the most important comparison theorems, which gives bounds on the sizes of Jacobi fields based on sectional curvature bounds. Recall that Jacobi field is a quite useful tool, based on the following observations:

1. Corollary 9.1.2 implies that in a normal neighborhood of p , every vector field can be represented as the value of Jacobi field that vanishes at p ;
2. The zeros of Jacobi fields corresponds to conjugate points, beyond which geodesics can't be minimal.

Theorem 15.1.1 (Rauch comparison). Let (M, g) and $(\widetilde{M}, \widetilde{g})$ be two Riemannian manifold with $\dim M \leq \dim \widetilde{M}$. Suppose $\gamma : [0, b] \rightarrow M$ and $\widetilde{\gamma} : [0, b] \rightarrow \widetilde{M}$ are two unit-speed geodesics such that

1. For all $t \in [0, b]$, and any planes $\Sigma \subset T_{\gamma(t)}M$, $\gamma'(t) \in \Sigma$, $\widetilde{\Sigma} \subset T_{\widetilde{\gamma}(t)}\widetilde{M}$, $\widetilde{\gamma}'(t) \in \widetilde{\Sigma}$, we have $K_{\gamma(t)}(\Sigma) \leq K_{\widetilde{\gamma}(t)}(\widetilde{\Sigma})$;
2. $\widetilde{\gamma}(0)$ has no conjugate points along $\widetilde{\gamma}|_{[0, b]}$.

Then for any Jacobi fields $J(t)$ and $\widetilde{J}(t)$ with

1.

$$\begin{cases} J(0) = c\gamma'(0) \\ \widetilde{J}(0) = c\widetilde{\gamma}'(0) \end{cases}$$

2. $|J'(0)| = |\widetilde{J}'(0)|$;
3. $\langle J'(0), \gamma'(0) \rangle = \langle \widetilde{J}'(0), \widetilde{\gamma}'(0) \rangle$.

we have $|J(t)| \geq |\widetilde{J}(t)|$ for all $t \in [0, b]$.

Proof. Firstly we consider the following simple case:

1. $J(0) = \widetilde{J}(0) = 0$;
2. $|J'(0)| = |\widetilde{J}'(0)|$;
3. $\langle J'(0), \gamma'(0) \rangle = \langle \widetilde{J}'(0), \widetilde{\gamma}'(0) \rangle = 0$.

Since $\widetilde{\gamma}(0)$ has no conjugate points along $\widetilde{\gamma}|_{[0, b]}$, then $\frac{|J(t)|^2}{|\widetilde{J}(t)|^2}$ is well-defined for all $t \in (0, b]$, and standard calculus implies

$$\lim_{t \rightarrow 0} \frac{|J|^2}{|\widetilde{J}|^2} = \lim_{t \rightarrow 0} \frac{\langle J'(t), J(t) \rangle}{\langle \widetilde{J}'(t), \widetilde{J}(t) \rangle} = \lim_{t \rightarrow 0} \frac{|J'|^2}{|\widetilde{J}'|^2} = 1$$

So it suffices to show in $(0, b]$ we have

$$\frac{d}{dt} \left(\frac{|J|^2}{|\widetilde{J}|^2} \right) \geq 0$$

Direct computation shows above inequality is equivalent to:

$$\frac{\langle J'(t), J(t) \rangle}{|J(t)|^2} \geq \frac{\langle \tilde{J}'(t), \tilde{J}(t) \rangle}{|\tilde{J}(t)|^2}$$

holds for arbitrary $t \in (0, b]$. For arbitrary $s \in (0, b]$, we can define the following Jacobi fields by scaling $J(t)$:

$$W_s(t) = \frac{J(t)}{|J(s)|}, \quad \widetilde{W}_s(t) = \frac{\tilde{J}(t)}{|\tilde{J}(s)|}$$

Then

$$\frac{\langle J'(s), J(s) \rangle}{|J(s)|^2} = \langle W'_s(s), W_s(s) \rangle$$

So it suffices to show

$$\langle W'_s(s), W_s(s) \rangle \geq \langle \widetilde{W}'_s(s), \widetilde{W}_s(s) \rangle$$

holds for arbitrary $s \in (0, b]$. Direct computation shows:

$$\begin{aligned} \langle W'_s(s), W_s(s) \rangle &= \int_0^s (\langle W_s(t), W_s(t) \rangle)' dt \\ &= \int_0^s \langle W'_s(t), W'_s(t) \rangle dt + \int_0^s \langle W''_s(t), W_s(t) \rangle dt \\ &= \int_0^s \langle W'_s(t), W'_s(t) \rangle dt - \int_0^s R(W_s(t), \gamma'(t), \gamma'(t), W_s(t)) dt \end{aligned}$$

Choose a parallel orthonormal frame $\{e_1(t), \dots, e_n(t)\}$ with $e_1(t) = \gamma'(t)$, $e_2(t) = W_s(t)$. With respect to this frame we write

$$W_s(t) = \lambda^i(t) e_i(t)$$

Similarly we choose a parallel orthogonal frame $\{\tilde{e}_1(t), \dots, \tilde{e}_n(t)\}$ and construct the following vector field

$$\tilde{V}(t) = \lambda^i(t) \tilde{e}_i(t)$$

Then it's clear we have

$$\int_0^s \langle W'_s(t), W'_s(t) \rangle dt = \int_0^s \langle \tilde{V}'(t), \tilde{V}'(t) \rangle dt$$

and our curvature condition implies

$$\int_0^s R(W_s(t), \gamma'(t), \gamma'(t), W_s(t)) dt \leq \int_0^s \tilde{R}(\tilde{V}(t), \gamma'(t), \gamma'(t), \tilde{V}(t)) dt$$

Thus we have

$$\begin{aligned} \langle W'_s(s), W_s(s) \rangle &\leq \int_0^s \langle \tilde{V}'(t), \tilde{V}'(t) \rangle dt - \int_0^s R(\tilde{V}(t), \gamma'(t), \gamma'(t), \tilde{V}(t)) dt \\ &= \tilde{I}(\tilde{V}, \tilde{V}) \end{aligned}$$

where \tilde{I} is index form on \tilde{M} . According to Corollary 9.3.1, we have

$$\tilde{I}(\tilde{V}, \tilde{V}) \geq \tilde{I}(\tilde{W}_s, \tilde{W}_s)$$

since \widetilde{W}_s is a Jacobi field. This shows the desired result.

For general case, we consider the following decomposition

$$\begin{aligned} J(t) &= J_1(t) + \langle J(t), \gamma'(t) \rangle \gamma'(t) \\ \widetilde{J}(t) &= \widetilde{J}_1(t) + \langle \widetilde{J}(t), \widetilde{\gamma}'(t) \rangle \widetilde{\gamma}'(t) \end{aligned}$$

Then it's clear $J_1(t)$ and $\widetilde{J}_1(t)$ satisfy requirement of our simple case, that is for $t \in [0, 1]$ we have

$$|J_1(t)| \geq |\widetilde{J}_1(t)|$$

Furthermore,

$$\langle J(t), \gamma'(t) \rangle = \langle \widetilde{J}(t), \widetilde{\gamma}'(t) \rangle$$

always holds, since

$$\begin{aligned} \langle J(t), \gamma'(t) \rangle &= \langle J(0), \gamma'(0) \rangle + \langle J'(0), \gamma'(0) \rangle t \\ &\stackrel{(1)}{=} \langle \widetilde{J}(0), \widetilde{\gamma}'(0) \rangle + \langle \widetilde{J}'(0), \widetilde{\gamma}'(0) \rangle t \\ &= \langle \widetilde{J}(t), \widetilde{\gamma}'(t) \rangle \end{aligned}$$

where (1) holds from our assumption. \square

Corollary 15.1.1. Let (M, g) be a Riemannian manifold, U a normal neighborhood of $p \in M$, $\gamma : [0, b] \rightarrow U$ a unit-speed geodesic with $\gamma(0) = p$ and J a Jacobi field along γ with $J(0) = 0$.

1. If the sectional curvature $K \leq k$ in U , then $|J(t)| \geq \text{sn}_k(t)|J'(0)|$, for all $t \in [0, b_0]$, where

$$b_0 = \begin{cases} b, & k \leq 0 \\ \min\{b, \pi R\}, & k = \frac{1}{R^2} > 0 \end{cases}$$

2. If the sectional curvature $K \geq k$ in U , then

$$|J(t)| \leq \text{sn}_k(t)|J'(0)|$$

for all $t \in [0, b]$.

Proof. Apply Rauch comparison between M and constant sectional curvature Riemannian manifold \widetilde{M} . However, in order to avoid geodesic $\widetilde{\gamma}$ of \widetilde{M} from having conjugate points, we need to let $b_0 < \min\{b, \pi R\}$, when $k = \frac{1}{R^2} > 0$. \square

Remark 15.1.1. In particular, from above corollary, we immediately have the following corollary when $K \leq k$:

1. If $k \leq 0$, we have already known M has no conjugate point along any geodesic;
2. If $k = \frac{1}{R^2} > 0$, then there is no conjugate point along any geodesic with length $< \pi R$. Or in other words, the distance between two consecutive conjugate points is $\geq \pi R$.

Corollary 15.1.2 (metric comparison). Let (M, g) be a Riemannian n -manifold, U a normal neighborhood of $p \in M$. For all $k \in \mathbb{R}$, we use g_k to denote the metric $dr \otimes dr + \operatorname{sn}_k(r)g_{\mathbb{S}^{n-1}}$ in $U \setminus \{p\}$.

1. If $K \leq k$ holds for all $q \in U \setminus \{p\}$, then for $w \in T_q M$ we have

$$g(w, w) \geq g_k(w, w)$$

holds in $U_0 \setminus \{p\}$, where

$$U_0 = \begin{cases} U, & k \leq 0 \\ U \cap B(p, \pi R), & k = \frac{1}{R^2} > 0 \end{cases}$$

2. If $K \geq k$ holds for all $q \in U \setminus \{p\}$, then for $w \in T_q M$ we have

$$g(w, w) \leq g_k(w, w)$$

holds in $U \setminus \{p\}$.

Proof. If $w = \partial_r|_q$, it's clear

$$g(\partial_r|_q, \partial_r|_q) = 1 = g_c(\partial_r|_q, \partial_r|_q)$$

by Gauss lemma, then it suffices to check for $w \in T_q M$ such that $g(w, \partial_r|_q) = 0$, we have

$$g(w, w) \geq g_c(w, w)$$

Let $\gamma : [0, b] \rightarrow M$ be a unit-speed geodesic connecting p and q , and J a Jacobi field such that $J(0) = 0, J(b) = w$. In normal coordinate $J(t)$ can be written as $ta^i \frac{\partial}{\partial x^i}|_{\gamma(t)}$ for some a^i .

Since (x^i, U, p) is both normal coordinate for metric g and g_c , thus γ is also a radial geodesic for g_c , and $J(t)$ is also a Jacobi field with respect to g_c along γ . Thus we have

$$\begin{aligned} g(w, w) &= |J(b)|_g^2 \\ g_c(w, w) &= |J(b)|_{g_k}^2 \end{aligned}$$

Then by Corollary 15.1.1, this completes the proof. \square

Remark 15.1.2. The **ideal** of this proof and the proof of Theorem 14.3.1 is almost the same, that is via Corollary 9.1.2 to construct a Jacobi field valued a given vector, and then one can use Rauch comparison to compare length of given vectors.

Corollary 15.1.3. Let (M, g) and $(\widetilde{M}, \widetilde{g})$ be two Riemannian manifolds with $K \leq \widetilde{K}$. Fix $p \in M, \widetilde{p} \in \widetilde{M}$, linear isometry $\Phi_0 : T_p M \rightarrow T_{\widetilde{p}} \widetilde{M}$ and $0 \leq \delta < \min(\operatorname{inj}(p), \operatorname{inj}(\widetilde{p}))$. Then for any smooth curve $\gamma : [0, 1] \rightarrow \exp_p(B(0, \delta))$ and $\widetilde{\gamma}(t) = \exp_{\widetilde{p}} \circ \Phi_0 \circ \exp_p^{-1}(\gamma(t))$, we have

$$L(\gamma) \geq L(\widetilde{\gamma})$$

Proof. Let $c(s) = \exp_p^{-1} \circ \gamma(s)$ and $\tilde{c}(s) = \exp_{\tilde{p}}^{-1} \circ \tilde{\gamma}(s)$, then $\tilde{c}(s) = \Phi_0(c(s))$. Consider the following variations

$$\begin{aligned}\alpha(t, s) &= \exp_p(tc(s)) \\ \tilde{\alpha}(t, s) &= \exp_{\tilde{p}}(t\tilde{c}(s))\end{aligned}$$

and Jacobi fields

$$\begin{aligned}J_s(t) &= \alpha_*\left(\frac{\partial}{\partial s}\right)(t, s) \\ \tilde{J}_s(t) &= \tilde{\alpha}_*\left(\frac{\partial}{\partial s}\right)(t, s)\end{aligned}$$

A crucial observation is for arbitrary $s_0 \in [0, 1]$, we have

$$\begin{aligned}J_{s_0}(1) &= \gamma'(s_0) \\ \tilde{J}_{s_0}(1) &= \tilde{\gamma}'(s_0)\end{aligned}$$

So it suffices to prove $|J_{s_0}(1)| \geq |\tilde{J}_{s_0}(1)|$ holds for arbitrary $s_0 \in [0, 1]$, that is we need to use Rauch comparison to Jacobi fields $J_{s_0}(t)$, $\tilde{J}_{s_0}(t)$ along γ_{s_0} and $\tilde{\gamma}_{s_0}$, where $\gamma_{s_0}(t) = \alpha(t, s_0)$ and $\tilde{\gamma}_{s_0}(t) = \tilde{\alpha}(t, s_0)$. Check requirements as follows:

1. $J_{s_0}(0) = \tilde{J}_{s_0}(0) = 0$;
2. $J'_{s_0}(0) = c'(s_0)$, $\tilde{J}'_{s_0}(0) = \tilde{c}'(s_0)$, and $\tilde{c}(s_0) = \Phi_0(c(s_0))$ implies $|J'_{s_0}(0)| = |\tilde{J}'_{s_0}(0)|$, since Φ_0 is linear isometry;
3. $\langle \tilde{J}'_{s_0}(0), \tilde{\gamma}'_{s_0}(0) \rangle = \langle \Phi_0(c'(s_0)), \Phi_0(c(s_0)) \rangle = \langle c'(s_0), c(s_0) \rangle = \langle J'_{s_0}(0), \gamma'_{s_0}(0) \rangle$.

□

Corollary 15.1.4. Let (M, g) be a Riemannian n -manifold, $0 < k_1 \leq K \leq k_2$. Let γ be any geodesic in M and b the distance along γ between two consecutive conjugate points, then

$$\frac{\pi}{\sqrt{k_2}} \leq b \leq \frac{\pi}{\sqrt{k_1}}$$

Proof. Without loss of generality, we assume $\gamma : [0, b] \rightarrow M$ is a unit-speed geodesic with $\gamma(0) = p$, $\gamma(b) = q$ and p, q are two consecutive conjugate points along γ .

1. By Remark 15.1.1, we have already seen $b \geq \frac{\pi}{\sqrt{k_2}}$;
2. Apply Rauch comparison to (M, g) and $(\mathbb{S}^n(\frac{\pi}{\sqrt{k_1}}), g_{\text{can}})$, we have

$$|J(t)| \leq |\tilde{J}(t)|$$

for $t \in [0, b]$, where $J(t)$, $\tilde{J}(t)$ are defined the same as before. Suppose $b > \frac{\pi}{\sqrt{k_1}}$, then take $t = \frac{\pi}{\sqrt{k_1}}$, we have

$$0 < |J(t)| \leq |\tilde{J}(t)| = 0$$

A contradiction.

□

Theorem 15.1.2. Let (M, g) be a compact Riemannian manifold with sectional curvature $K \leq k, k > 0$. If we define

$$l(M, g) := \inf\{L(\gamma) \mid \gamma \text{ is a smooth closed geodesic in } M\}$$

Then either $\text{inj}(M) \geq \frac{\pi}{\sqrt{k}}$ or $\text{inj}(M) = \frac{l(M, g)}{2}$.

Proof. By compactness of M , there exists $p, q \in M, q \in \text{cut}(p)$ such that $\text{dist}(p, q) = \text{inj}(M) = \text{inj}(p)$. Let $\gamma : [0, b] \rightarrow M$ be a minimal geodesic connecting p and q , that is $b = \text{dist}(p, q) = \text{inj}(M)$.

1. If p and q are not conjugate along γ , then by Corollary 15.1.4 we have $\text{inj}(M) = b \geq \frac{\pi}{\sqrt{k}}$.
2. If p and q are not conjugate along γ , then by Proposition 10.1.1 there exists a unit-speed closed geodesic $\gamma : [0, 2b] \rightarrow M$ with $\gamma(0) = p, \gamma(b) = q$, where $b = \text{dist}(p, q) = \text{inj}(M)$. On one hand by definition of $l(M, g)$ one has $2b \geq l(M, g)$; On the other hand, $l(M, g) \geq 2b$, since $\text{dist}(p, q) = b$. Thus in this case $\text{inj}(M) = \frac{l(M, g)}{2}$.

□

15.2. Hessian comparison.

Theorem 15.2.1 (Hessian comparison). Let (M, g) and $(\widetilde{M}, \widetilde{g})$ be two Riemannian manifolds with the same dimension, $U \subset M, \widetilde{U} \subset \widetilde{M}$ normal neighborhoods around $p \in M$ and $\widetilde{p} \in \widetilde{M}$ respectively. Suppose

$$\gamma : [0, b] \rightarrow U, \gamma(0) = p, \gamma(b) = q$$

$$\widetilde{\gamma} : [0, b] \rightarrow \widetilde{U}, \widetilde{\gamma}(0) = \widetilde{p}, \widetilde{\gamma}(b) = \widetilde{q}$$

are two unit-speed geodesics such that

For all $t \in [0, b]$, and any planes $\Sigma \subset T_{\gamma(t)}M, \gamma'(t) \in \Sigma, \widetilde{\Sigma} \subset T_{\widetilde{\gamma}(t)}\widetilde{M}, \widetilde{\gamma}'(t) \in \widetilde{\Sigma}$, we have $K_{\gamma(t)}(\Sigma) \leq K_{\widetilde{\gamma}(t)}(\widetilde{\Sigma})$.

Then for any $v \in T_qM, \widetilde{v} \in T_{\widetilde{q}}\widetilde{M}$ with unit length and $v \perp \gamma'(b), \widetilde{v} \perp \widetilde{\gamma}'(b)$, we have

1. $\text{Hess } r(v, v) \geq \text{Hess } \widetilde{r}(\widetilde{v}, \widetilde{v})$;
2. $\Delta r(\gamma(t)) \geq \Delta \widetilde{r}(\widetilde{\gamma}(t))$ for all $t \in (0, b]$;
3. Moreover, the equality holds if and only if $K_{\Sigma}(\gamma(t)) = \widetilde{K}_{\widetilde{\Sigma}}(\widetilde{\gamma}(t))$.

Proof. For (1). Let $\{e_1(t), \dots, e_n(t)\}$ be a parallel orthonormal basis along γ such that $e_n(t) = \gamma'(t)$ and $\{\widetilde{e}_1(t), \dots, \widetilde{e}_n(t)\}$ a parallel orthonormal basis along $\widetilde{\gamma}$ such that $\widetilde{e}_n(t) = \widetilde{\gamma}'(t)$. Without loss of generality we may assume $\langle v, e_i(b) \rangle_g = \langle \widetilde{v}, \widetilde{e}_i(b) \rangle_{\widetilde{g}}$ for $i = 1, \dots, n-1$, it's just a trick of linear algebra.

Via Corollary 9.1.2 to construct Jacobi fields

$$\begin{cases} J(0) = 0, J(b) = v \\ \widetilde{J}(0) = 0, \widetilde{J}(b) = \widetilde{v} \end{cases}$$

With respect to $\{\tilde{e}_i(t)\}$ we can write $\tilde{J}(t)$ as $\tilde{J}(t) = \lambda^i(t)\tilde{e}_i(t)$, and construct $V(t) = \lambda^i(t)e_i(t)$. Then

$$\begin{aligned}
\text{Hess } r(v, v) &= \text{Hess } r(J(b), J(b)) \\
&\stackrel{\text{I}}{=} \int_0^b \langle J'(t), J'(t) \rangle - R(J, \gamma', \gamma', J) dt \\
&\stackrel{\text{II}}{\geq} \int_0^b \langle V'(t), V'(t) \rangle - R(V, \gamma', \gamma', V) dt \\
&\stackrel{\text{III}}{\geq} \int_0^b \langle \tilde{J}'(t), \tilde{J}'(t) \rangle - \tilde{R}(\tilde{J}, \tilde{\gamma}, \tilde{\gamma}, \tilde{J}) dt \\
&= \text{Hess } \tilde{r}(\tilde{J}(b), \tilde{J}(b)) \\
&= \text{Hess } \tilde{r}(\tilde{v}, \tilde{v})
\end{aligned}$$

where

I holds from Proposition 14.4.1;

II holds from Corollary 9.3.1;

III holds from our assumption on curvature and the choice of V .

For (2) and (3). They directly follow from (1) and proof of (1). \square

Corollary 15.2.1 (Hessian and Laplacian comparision). Let (M, g) be a Riemannian n -manifold and U a normal neighborhood of $p \in M$.

1. If sectional curvature $K \leq k$ in $U \setminus \{p\}$, then

$$\mathcal{H}_r \geq \frac{\text{sn}'_k(r)}{\text{sn}_k(r)} \pi_r, \quad \Delta r \geq (n-1) \frac{\text{sn}'_k(r)}{\text{sn}_k(r)}$$

holds in $U_0 \setminus \{p\}$, where

$$U_0 = \begin{cases} U, & k \leq 0 \\ U \cap B(p, \pi R), & k = \frac{1}{R^2} > 0 \end{cases}$$

2. If sectional curvature $K \geq k$ in $U \setminus \{p\}$, then

$$\mathcal{H}_r \leq \frac{\text{sn}'_k(r)}{\text{sn}_k(r)} \pi_r, \quad \Delta r \leq (n-1) \frac{\text{sn}'_k(r)}{\text{sn}_k(r)}$$

holds in $U \setminus \{p\}$;

3. Moreover, if equality holds, g has constant sectional curvature k in U_0 or U .

Proof. For (1). Apply Hessian comparision to (M, g) and space form $S(n, k)$, then we directly have

$$\text{Hess } r(v, v) \geq \text{Hess } \tilde{r}(\tilde{v}, \tilde{v})$$

for any $v \in T_q M, \tilde{v} \in T_q S(n, k)$ with unit length and $v \perp \gamma'(b), \tilde{v} \perp \tilde{\gamma}'(b)$, where

$$\gamma : [0, b] \rightarrow U, \gamma(0) = p, \gamma(b) = q$$

$$\tilde{\gamma} : [0, b] \rightarrow \tilde{U}, \tilde{\gamma}(0) = \tilde{p}, \tilde{\gamma}(b) = \tilde{q}$$

are two unit-speed geodesics, and U, \tilde{U} are normal neighborhoods of p, \tilde{p} respectively. However, we must be careful here, since if sectional curvature of M is ≤ 0 , then b can be infinite, and in this case if $k > 0$, the diameter of \tilde{U} is $< \frac{\pi}{\sqrt{k}}$. Thus we only have

$$\text{Hess } r(v, v) \geq \text{Hess } \tilde{r}(\tilde{v}, \tilde{v})$$

for $0 < b < \frac{\pi}{\sqrt{k}}$ if $k > 0$, and there is no restriction for b if $k \leq 0$. Thus by taking different geodesics and different Jacobi fields, we can show this holds for arbitrary $v \in T_q M, \tilde{v} \in T_q S(n, k)$, where $q \in U_0 \setminus \{p\}$, that is we have

$$\mathcal{H}_r \geq \frac{\text{sn}'_k(r)}{\text{sn}_k(r)} \pi_r$$

holds in $U_0 \setminus \{p\}$. By taking trace we obtain $\Delta r \geq (n-1) \frac{\text{sn}'_k(r)}{\text{sn}_k(r)}$ holds in $U_0 \setminus \{p\}$, since π_r is a projection onto a subspace with codimension 1.

For (2), the same as (1).

For (3), if

$$\mathcal{H}_r = \frac{\text{sn}'_k(r)}{\text{sn}_k(r)} \pi_r$$

holds in $U \setminus \{p\}$, then it's directly from Proposition 14.4.2. If

$$\Delta r \geq (n-1) \frac{\text{sn}'_k(r)}{\text{sn}_k(r)}$$

holds in $U \setminus \{p\}$, that is the trace of $\mathcal{H}_r - \frac{\text{sn}'_k(r)}{\text{sn}_k(r)} \pi_r$ vanishes identically in $U \setminus \{p\}$, then $\mathcal{H}_r - \frac{\text{sn}'_k(r)}{\text{sn}_k(r)} \pi_r$ vanishes identically, since it's semipositive-definite. \square

16. COMPARISON THEOREMS BASED ON RICCI CURVATURE

16.1. Local Laplacian comparison.

Theorem 16.1.1 (local Laplacian comparison). Let (M, g) be a Riemannian n -manifold and U a normal coordinate of $p \in M$. If there exists $k \in \mathbb{R}$ such that

$$\text{Ric}(g) \geq (n-1)kg$$

Then

$$\Delta r \leq (n-1) \frac{\text{sn}'_k(r)}{\text{sn}_k(r)}$$

holds in $U_0 \setminus \{p\}$, where

$$U_0 = \begin{cases} U, & k \leq 0 \\ U \cap B(p, \pi R), & k = \frac{1}{R^2} > 0 \end{cases}$$

Moreover, if equality holds, then g has constant sectional curvature in U_0 .

16.1.1. Proof via Jacobi fields.

Proof of Theorem 16.1.1 via Jacobi fields. For arbitrary $q \in U_0 \setminus \{p\}$, choose a unit-speed geodesic $\gamma : [0, b] \rightarrow M$ with $\gamma(0) = p, \gamma(b) = q$, and $\{e_1(t), \dots, e_n(t)\}$ is a parallel orthonormal frame along γ with $e_n(t) = \gamma'(t)$. Then

$$\Delta r = \sum_{i=1}^n \text{Hess } r(e_i, e_i)$$

Via Corollary 9.1.2 to construct Jacobi fields $J_i(t), i = 1, \dots, n$ such that $J_i(0) = 0, J_i(b) = e_i(b)$, then we have

$$\Delta r = \sum_{i=1}^{n-1} \text{Hess}(J_i(b), J_i(b)) \stackrel{(1)}{=} \sum_{i=1}^{n-1} I(J_i, J_i)$$

where (1) holds from Corollary 14.4.1.

Now let \widetilde{M} be the space form $S(n, k)$ and \widetilde{U} a normal coordinate of $\widetilde{p} \in \widetilde{M}$. Repeat the same process as above we have

$$(n-1) \frac{\text{sn}'_k(r)}{\text{sn}_k(r)} = \widetilde{\Delta} \widetilde{r} = \sum_{i=1}^{n-1} \widetilde{I}(\widetilde{J}_i, \widetilde{J}_i)$$

A crucial observation is that $\tilde{J}_i(t) = f(t)\tilde{e}_i(t)$, and the **key point** is that $f(t)$ is independent of i . Denote $V_i(t) = f(t)e_i(t)$, then

$$\begin{aligned}
\Delta r &= \sum_{i=1}^{n-1} I(J_i, J_i) \\
&\leq \sum_{i=1}^{n-1} I(V_i, V_i) \\
&= \sum_{i=1}^{n-1} \int_0^b \langle V'_i(t), V'_i(t) \rangle - R(V_i, \gamma', \gamma', V_i) dt \\
&\stackrel{(2)}{=} \sum_{i=1}^{n-1} \int_0^b \langle V'_i(t), V'_i(t) \rangle - f^2(t) R(e_i, e_n, e_n, e_i) dt \\
&= \sum_{i=1}^{n-1} \int_0^b \langle V'_i(t), V'_i(t) \rangle - \int_0^b f^2(t) \operatorname{Ric}(e_n, e_n) dt \\
&\leq \sum_{i=1}^{n-1} \int_0^b \langle \tilde{J}_i(t), \tilde{J}_i(t) \rangle - \int_0^b (n-1)k f^2(t) dt \\
&= \sum_{i=1}^{n-1} \tilde{I}(\tilde{J}_i, \tilde{J}_i) \\
&= (n-1) \frac{\operatorname{sn}'_k(r)}{\operatorname{sn}_k(r)}
\end{aligned}$$

the key point is used in equality marked by (2), and others are routines. \square

16.1.2. *Proof via Bochner's technique.*

Lemma 16.1.1. Let (M, g) be a Riemannian manifold, (x^i, U, p) a normal coordinate centered at p , then

$$\Delta r = \partial_r \log(r^{n-1} \sqrt{\det g})$$

in $U \setminus \{p\}$. Moreover, along any unit-speed geodesic $\gamma : [0, b] \rightarrow U$ with $\gamma(0) = p$, if we define $f(t) := \Delta r(\gamma(t))$, then

$$f(t) = \frac{n-1}{t} + O(1)$$

Proof. Direct computation shows

$$\begin{aligned}
\Delta r &= \frac{1}{\sqrt{\det g}} \frac{\partial}{\partial x^i} (g^{ij} \sqrt{\det g} \frac{\partial r}{\partial x^j}) \\
&= \frac{1}{\sqrt{\det g}} \frac{\partial}{\partial x^i} (g^{ij} \sqrt{\det g} \frac{x^j}{r}) \\
&= \frac{1}{\sqrt{\det g}} \frac{\partial}{\partial x^i} (\frac{x^i}{r} \sqrt{\det g}) \\
&= \frac{\partial}{\partial x^i} (\frac{x^i}{r}) + \frac{1}{\sqrt{\det g}} \frac{x^i}{r} \frac{\partial}{\partial x^i} (\sqrt{\det g}) \\
&= \frac{n-1}{r} + \frac{1}{\sqrt{\det g}} \partial_r (\sqrt{\det g}) \\
&= \partial_r \log(r^{n-1} \sqrt{\det g})
\end{aligned}$$

Moreover, for unit-speed geodesic $\gamma : [0, b] \rightarrow U$, we have

$$f(t) = \frac{n-1}{r(\gamma(t))} + \partial_r (\log \sqrt{\det g}) \Big|_{\gamma(t)}$$

Then note that

1. $r(\gamma(t)) = t$, since γ is unit-speed geodesic.
2. Jacobi's formula implies

$$\partial_r (\log \sqrt{\det g}) \Big|_{\gamma(t)} = \frac{1}{2} g^{ij} \frac{\partial g_{ij}}{\partial x^k} \frac{d\gamma^k}{dt} = O(1)$$

we obtain the desired results. \square

Lemma 16.1.2 (Riccati comparison theorem). If $f : (0, b) \rightarrow \mathbb{R}$ is a smooth function satisfying

1. $f(t) = \frac{1}{t} + O(1)$;
2. $f' + f^2 + k \leq 0$.

Then

$$f(t) \leq \frac{\text{sn}'_k(t)}{\text{sn}_k(t)}$$

for all $t \in (0, b)$, where $k > 0, b \leq \frac{\pi}{\sqrt{k}}$.

Proof. Consider $f_k(t) = \frac{\text{sn}'_k(t)}{\text{sn}_k(t)}$, it's a smooth function defined on $(0, b)$ satisfying

1. $f_k(t) = \frac{1}{t} + O(1)$
2. $f'_k + f_k^2 + k = 0$

Choose a smooth function $F : (0, b) \rightarrow \mathbb{R}$ satisfying

1. $F(t) = 2 \log t + O(1)$;
2. $F'(t) = f + f_k$

Then

$$\begin{aligned}\frac{d}{dt}(e^F(f - f_k)) &= e^F(f^2 - f_k^2 + f' - f'_k) \leq 0 \\ \lim_{t \rightarrow 0} e^F(f - f_k) &= 0\end{aligned}$$

Then we have $f(t) \leq f_k(t)$ holds for all $t \in (0, b)$. \square

Lemma 16.1.3.

$$|\text{Hess } r|^2 \geq \frac{(\Delta r)^2}{n-1}$$

Proof. Let $\{e_1, \dots, e_n\}$ be an orthonormal frame with $e_1 = \partial_r$. Then

$$\begin{aligned}|\text{Hess } r|^2 &= \sum_{i,j=1}^n (\langle \nabla_{e_i} \partial_r, e_j \rangle)^2 \\ &= \sum_{i,j=2}^n (\langle \nabla_{e_i} \partial_r, e_j \rangle)^2 \\ &\geq \frac{1}{n-1} \sum_{i=2}^n (\langle \nabla_{e_i} \partial_r, e_i \rangle)^2 \\ &= \frac{1}{n-1} (\Delta r)^2\end{aligned}$$

The inequality

$$|A|^2 \geq \frac{1}{k} |\text{tr}(A)|^2$$

for a $k \times k$ matrix A is a direct consequence of the Cauchy-Schwarz inequality. \square

Proof of Theorem 16.1.1 via Bochner's technique. Recall Bochner's technique says

$$\frac{1}{2} \Delta |\nabla f|^2 = |\text{Hess } f|^2 + \text{Ric}(\nabla f, \nabla f) + g(\nabla \Delta f, \nabla f)$$

Set $f = r$ we have

$$\begin{aligned}0 &= |\text{Hess } r|^2 + \text{Ric}(\nabla r, \nabla r) + g(\nabla \Delta r, \nabla r) \\ &\stackrel{(1)}{\geq} |\text{Hess } r|^2 + \partial_r(\Delta r) + (n-1)k \\ &\stackrel{(2)}{\geq} \left(\frac{\Delta r}{n-1}\right)^2 + \partial_r\left(\frac{\Delta r}{n-1}\right) + k\end{aligned}$$

where

(1) holds from $\partial_r = \nabla_r$ and lower bounded of Ricci;

(2) holds from Lemma 16.1.3 and divided by $n-1$.

Thanks to Lemma 16.1.1, we can apply Riccati comparison to $f(r) = \frac{\Delta r}{n-1}$, then we have

$$\frac{\Delta r}{n-1} \leq \frac{\text{sn}'_k(r)}{\text{sn}_k(r)}$$

This shows desired comparison.

Furthermore, if equality holds

$$\frac{\Delta r}{n-1} = \frac{\text{sn}'_k(r)}{\text{sn}_k(r)}$$

then direct computation shows

$$\left(\frac{\Delta r}{n-1}\right)^2 + \partial_r\left(\frac{\Delta r}{n-1}\right) + k = 0$$

which implies inequalities in (1) and (2) are in fact equalities. In particular one has

$$|\text{Hess } r|^2 = \frac{(\Delta r)^2}{n-1}$$

that is inequality in Cauchy-Schwarz inequality holds, which implies

$$\mathcal{H}_r = \frac{\text{sn}'_k(r)}{\text{sn}_k(r)} \pi_r$$

Then g has constant sectional curvature k in U_0 by Proposition 14.4.2. \square

16.2. Maximal principle.

Proposition 16.2.1. Let (M, g) be a Riemannian manifold and f, h be two smooth functions on M . If there is a point p such that $f(p) = h(p)$ and $f(x) \geq h(x)$ for all x near p , then

$$\nabla f(p) = \nabla h(p), \quad \text{Hess } f|_p \geq \text{Hess } h|_p, \quad \Delta f(p) \geq \Delta h(p).$$

Proof. Firstly let's consider the case $(M, g) \subset (\mathbb{R}^n, g_{\text{can}})$, it's a simple calculus since we can use Taylor expansion. To be explicit, for all x near p , we have

$$f(x) = f(p) + \nabla f(p)^T(x-p) + \frac{1}{2}(x-p)^T \text{Hess } f|_p(x-p) + O(|x|^3)$$

where ∇f is a n column vector and $\text{Hess } f$ is a $n \times n$ matrix in this case. Similarly we have

$$h(x) = h(p) + \nabla h(p)^T(x-p) + \frac{1}{2}(x-p)^T \text{Hess } h|_p(x-p) + O(|x|^3)$$

Then consider

$$f(x) - h(x) = (\nabla f - \nabla h)(p)^T(x-p) + \frac{1}{2}(x-p)^T \text{Hess}(f-h)|_p(x-p) + O(|x|^3)$$

Since $f(x) - h(x) \geq 0$ for all x near p , then we must have

$$\begin{aligned} \nabla f(p) &= \nabla h(p) \\ \text{Hess } f|_p &\geq \text{Hess } h|_p \end{aligned}$$

By taking trace we have

$$\Delta f(p) \geq \Delta h(p)$$

For general case, take $\gamma : (-\varepsilon, \varepsilon) \rightarrow M$ to be a geodesic with $\gamma(0) = p$, then use previous case on $f \circ \gamma, h \circ \gamma$ to obtain

$$\begin{aligned}\nabla_{\gamma'(0)} f(p) &= \nabla_{\gamma'(0)} h(p) \\ \text{Hess } f_p(\gamma'(0), \gamma'(0)) &\geq \text{Hess } h_p(\gamma'(0), \gamma'(0))\end{aligned}$$

Then it's clear this proposition holds if we let $v = \gamma'(0)$ run over all $v \in T_p M$. \square

Definition 16.2.1 (barrier sense). Let (M, g) be a Riemannian manifold and $f \in C^0(M)$. Suppose f_q is a C^2 function defined in a neighborhood of U of $q \in M$.

1. f_q is called a lower barrier function of f at q if

$$f_q(q) = f(q), \quad f_q(x) \leq f(x), \quad x \in U.$$

- 2.

$$\Delta f(q) \geq c$$

in the barrier sense if for all $\varepsilon > 0$, there exists a lower barrier function $f_{q,\varepsilon}$ of f at q such that

$$\Delta f_{q,\varepsilon}(q) \geq c - \varepsilon$$

- 3.

$$\Delta f(q) \leq c$$

in the barrier sense if for all $\varepsilon > 0$, there exists an upper barrier function $f_{q,\varepsilon}$ of f at q such that

$$\Delta f_{q,\varepsilon}(q) \leq c + \varepsilon$$

Definition 16.2.2 (distribution sense). Let (M, g) be a Riemannian manifold and $f \in C^0(M)$.

$$\Delta f \leq h$$

in distribution sense, if

$$\int f \Delta \varphi \leq \int h \varphi$$

holds for all $\varphi \geq 0 \in C_c^\infty(M)$

Theorem 16.2.1 (maximal principle). Let (M, g) be a Riemannian manifold and $f \in C^0(M)$.

1. If $\Delta f \geq 0$ in the barrier sense or distribution sense, then if f has a local(global) maximum, then it's local(global) constant;
2. If $\Delta f \leq 0$ in the barrier sense or distribution sense, then if f has a local(global) minimal, then it's local(global) constant;
3. $\Delta f = 0$ implies $f \in C^\infty(M)$.

Proof. Here we only prove (1) for barrier sense: First, suppose that $\Delta f(x) > 0$ everywhere. Then f can't have any local maxima at all. For if f has a local maximum at $p \in M$, then there would exist a smooth support function $f_\varepsilon(x)$ with

- (1) $f_\varepsilon(p) = f(p)$,
- (2) $f_\varepsilon(x) \leq f(x)$ for all x near p ,
- (3) $\Delta f_\varepsilon(p) > 0$.

Here (1) and (2) imply that f_ε must also have a local maximum at p . But this implies that $\nabla^2 f_\varepsilon(p) \leq 0$, which contradicts (3).

Next just assume that $\Delta f \geq 0$ and let $p \in M$ be a local maximum for f . For sufficiently small $r < \text{inj}(p)$ we therefore have a function $f : (B(p, r), g) \rightarrow \mathbb{R}$ with $\Delta f \geq 0$ and a global maximum at p . If f is constant on $B(p, r)$, then we are done; otherwise, we can assume (by possibly decreasing r) that f is not equal to $f(p)$ on $S(p, r) = \{x \in M : \text{dist}(p, x) = r\}$. Then define $V = \{x \in S(p, r) : f(x) = f(p)\}$. Now construct a smooth function $h = e^{\alpha\varphi} - 1$ such that

$$\begin{aligned} h &< 0 && \text{on } V \\ h(p) &= 0 \\ \Delta h &> 0 && \text{on } \bar{B}(p, r) \end{aligned}$$

This function is found by first selecting an open disc $U \subset S(p, r)$ that contains V . We can then find φ such that

$$\begin{aligned} \varphi(p) &= 0 \\ \varphi &< 0 && \text{on } U \\ \nabla\varphi &\neq 0 && \text{on } \bar{B}(p, r) \end{aligned}$$

In an appropriate coordinate system (x^1, \dots, x^n) we can simply assume that U looks like the lower half-plane: $x^1 < 0$ and then define $\varphi = x^1$. Then choose α so big that $\Delta h = \alpha e^{\alpha\varphi} (\alpha |\nabla\varphi|^2 + \Delta\varphi) > 0$ on $\bar{B}(p, r)$.

Now consider the function $f_\delta = f + \delta h$ on $\bar{B}(p, r)$. Provided that δ is very small, this function has a local maximum in the interior $B(p, r)$, since

$$f_\delta(p) = f(p) > \max \{f(x) + \delta h(x) = f_\delta(x) : x \in \partial B(p, r)\}$$

On the other hand, we can also show that f_δ has positive Laplacian, thus giving a contradiction with the first part of the proof. To see that the Laplacian is positive, select f_ε a support function from below for f at $q \in B(p, r)$. Then $f_s + \delta h$ is a support function from below for f_δ at q . The Laplacian of this support function is estimated by

$$\Delta(f_\varepsilon + \delta h)(q) \geq -\varepsilon + \delta \Delta h(q),$$

which for given δ must become positive as $\varepsilon \rightarrow 0$. □

16.3. Global Laplacian comparison.

16.3.1. In the barrier sense.

Lemma 16.3.1. Let (M, g) be a complete Riemannian manifold, $p, q \in M$ and $\gamma : [0, b] \rightarrow M$ a unit-speed minimal geodesic connecting p, q . Then for any $0 < \varepsilon < b$, $\gamma|_{[\varepsilon, b]}$ is the unique minimal geodesic connecting $\gamma(\varepsilon)$ and q .

Proposition 16.3.1. Let (M, g) be a complete Riemannian manifold and $p, q \in M$. Let $\gamma : [0, b] \rightarrow M$ be a unit-speed minimal geodesic with $\gamma(0) = p$ and $\gamma(b) = q$. For any small $\varepsilon > 0$, we define

$$r_\varepsilon(x) = \varepsilon + \text{dist}(\gamma(\varepsilon), x), \quad x \in M.$$

Then

1. $q \notin \text{cut}(\gamma(\varepsilon))$ and in particular, r_ε is smooth at q .
2. r_ε is an upper barrier function of $r(x) = \text{dist}(p, x)$ at point q .

Proof. For (1). If $q \in \text{cut}(\gamma(\varepsilon))$, then by definition there exists a minimal geodesic $\tilde{\gamma}$ connecting $\gamma(\varepsilon)$ and q which is no longer minimizing after q . By Lemma 16.3.1, we have $\tilde{\gamma}$ is exactly $\gamma|_{[\varepsilon, b]}$. By Theorem 10.1.1, there are two cases:

1. q is conjugate to $\gamma(\varepsilon)$ along γ . That's impossible, since γ is the minimal geodesic connecting p and q ;
2. There exists at least two different minimal geodesics connecting $\gamma(\varepsilon)$ and q , by Lemma 16.3.1, that's also impossible.

This shows $q \notin \text{cut}(\gamma(\varepsilon))$. In particular, r_ε is smooth at q .

For (2). Firstly note that $\gamma(b) = q$, then

$$r(q) = \text{dist}(p, q) = \text{dist}(\gamma(0), \gamma(b)) \stackrel{\text{I}}{=} \text{dist}(\gamma(0), \gamma(\varepsilon)) + \text{dist}(\gamma(\varepsilon), \gamma(b)) \stackrel{\text{II}}{=} r_\varepsilon(q)$$

where

I holds since γ is a minimal geodesic;

II holds since γ is unit-speed minimal geodesic, then $\text{dist}(\gamma(0), \gamma(\varepsilon)) = \varepsilon$.

By triangle inequality, one has

$$r(q') = \text{dist}(p, q') \leq \varepsilon + \text{dist}(\gamma(\varepsilon), q') = r_\varepsilon(q')$$

for all q' near q . Combining these two facts together we have r_ε is an upper barrier function of r . \square

Theorem 16.3.1 (global Laplacian comparison). Let (M, g) be a complete Riemannian manifold with

$$\text{Ric}(g) \geq (n-1)kg$$

Then for $q \in M$

$$\Delta r(q) \leq (n-1) \frac{\text{sn}'_k(r(q))}{\text{sn}_k(r(q))}$$

in the barrier sense.

Proof. We consider the following three cases:

1. If $q \in M \setminus \{p\} \cup \text{cut}(p)$, it's exactly smooth case we have proven;
2. If $q = p$, it's clear, since the right hand is infinite;
3. For arbitrary $q \in \text{cut}(p)$, there exists a unit-speed $\gamma : [0, b] \rightarrow M$ with $\gamma(0) = p, \gamma(b) = q$. Then for each $\gamma > 0$, define

$$\gamma_\varepsilon(x) = \varepsilon + \text{dist}(\gamma(\varepsilon), x)$$

Then by Proposition 16.3.1 we have $\gamma_\varepsilon(x)$ is an upper barrier of $r(x)$ and γ_ε is smooth at q . Thus we have

$$\begin{aligned}\Delta\gamma_\varepsilon(q) &= \Delta \operatorname{dist}(\gamma_\varepsilon, q) \\ &\leq (n-1) \frac{\operatorname{sn}'_k(\gamma_\varepsilon(q) - \varepsilon)}{\operatorname{sn}_k(\gamma_\varepsilon(q) - \varepsilon)} \\ &= (n-1) \frac{\operatorname{sn}'_k(\gamma(q) - \varepsilon)}{\operatorname{sn}_k(\gamma(q) - \varepsilon)}\end{aligned}$$

which descends to $(n-1) \frac{\operatorname{sn}'_k(\gamma(q))}{\operatorname{sn}_k(\gamma(q))}$ as $\varepsilon \rightarrow 0$ by monotonicity. This completes the proof. \square

16.3.2. In the distribution sense.

Proposition 16.1. Let (M, g) be a Riemannian manifold and $f : M \rightarrow \mathbb{R}$ be a Lipschitz function. Then for any $\varphi \in C_0^\infty(M, \mathbb{R})$, one has

$$-\int_M \langle \nabla \varphi, \nabla f \rangle d\operatorname{vol}_g = \int_M \Delta \varphi \cdot f d\operatorname{vol}_g.$$

Proof. Let $f : M \rightarrow \mathbb{R}$ be Lipschitz function, then by a partition-of-unity procedure one may express f as a locally finite sum $\sum_{\alpha \in I} f_\alpha$ subordinate to open covering $\{U_\alpha\}_{\alpha \in I}$, that is f_α has compact support in U_α . Without loss of generality, one chooses locally finite open covering U_α by geodesic balls. Then each f_α can be considered to be a function with compact support on a euclidean space.

Rademacher's theorem says if U is an open subset of \mathbb{R}^n and $f : U \rightarrow \mathbb{R}$ is a Lipschitz function, then f is differentiable almost everywhere in U . So in this viewpoint, we can see Lipschitz functions on a Riemannian manifold is almost everywhere differentiable.

The followings are routine calculus to show integration by parts holds:

$$\begin{aligned}\operatorname{div}(f \nabla \varphi) &= \nabla_k (f \nabla \varphi)^k \\ &= \frac{\partial (f \nabla \varphi)^k}{\partial x^k} + \Gamma_{ks}^k (f \nabla \varphi)^s \\ &= \frac{\partial (f g^{ik} \frac{\partial \varphi}{\partial x^i})}{\partial x^k} + \Gamma_{ks}^k f g^{is} \frac{\partial \varphi}{\partial x^i} \\ &= \underbrace{g^{ik} \frac{\partial f}{\partial x^k} \frac{\partial \varphi}{\partial x^i}}_{\text{part I}} + \underbrace{f \left(\frac{\partial g^{ik}}{\partial x^k} \frac{\partial \varphi}{\partial x^i} + g^{ik} \frac{\partial^2 \varphi}{\partial x^k \partial x^i} + g^{is} \Gamma_{ks}^k \frac{\partial \varphi}{\partial x^i} \right)}_{\text{part II}}\end{aligned}$$

We have the following observations:

1. Part I equals

$$\begin{aligned} g^{ik} \frac{\partial f}{\partial x^k} \frac{\partial \varphi}{\partial x^i} &= g_{lj} g^{lk} \frac{\partial f}{\partial x^k} g^{ji} \frac{\partial \varphi}{\partial x^i} \\ &= \langle g^{lk} \frac{\partial f}{\partial x^k} \frac{\partial}{\partial x^l}, g^{ji} \frac{\partial \varphi}{\partial x^i} \frac{\partial}{\partial x^j} \rangle \\ &= \langle \nabla f, \nabla \varphi \rangle \end{aligned}$$

2. Note

$$\begin{aligned} \frac{\partial g^{ik}}{\partial x^k} + g^{is} \Gamma_{ks}^k \frac{\partial \varphi}{\partial x^i} &= -g^{is} g^{kt} \frac{\partial g_{st}}{\partial x^k} + \frac{1}{2} g^{is} g^{kt} \left(\frac{\partial g_{kt}}{\partial x^s} + \frac{\partial g_{st}}{\partial x^k} - \frac{\partial g_{ks}}{\partial x^t} \right) \\ &= -\frac{1}{2} g^{is} g^{kt} \left(\frac{\partial g_{ks}}{\partial x^t} + \frac{\partial g_{st}}{\partial x^k} - \frac{\partial g_{kt}}{\partial x^s} \right) \\ &= -g^{kt} \Gamma_{kt}^i \end{aligned}$$

where $\frac{\partial g^{ik}}{\partial x^k} = -g^{is} g^{kt} \frac{\partial g_{st}}{\partial x^k}$ holds from the fact $g^{ik} g_{kt} = \delta_t^i$, then take partial derivative with respect to x^k to conclude.

3. From (2) and local expression of Δ , it's clear part II equals $f\Delta\varphi$.

Thus we have

$$\operatorname{div}(f\nabla\varphi) = \langle \nabla\varphi, \nabla f \rangle + f\Delta\varphi$$

Then divergence theorem completes the proof. \square

Theorem 16.3.2 (global Laplacian comparision II). Let (M, g) be a complete Riemannian manifold with

$$\operatorname{Ric}(g) \geq (n-1)kg$$

Then for $x \in M$

$$\Delta r(x) \leq (n-1) \frac{\operatorname{sn}'_k(r(x))}{\operatorname{sn}_k(r(x))}$$

in the distribution sense.

Proof. For fixed $p \in M$, the domain $\Sigma(p)$ of injective radius $\operatorname{inj}(p)$ is a star-shaped open subset of $T_p M$ and $M = \exp_p(\Sigma(p)) \cup \operatorname{cut}(p)$. The boundary of $\Sigma(p)$ is locally a graph of continuous function and so there exists a family of star-shaped domains $\{U_j\}$ with smooth boundaries such that

$$U_j \subset U_{j+1} \subset \cdots \subset \Sigma(p), \quad \Sigma(p) = \bigcup U_j$$

If we set $\Omega = \exp_p(\Sigma(p))$, then $\Omega = \bigcup \Omega_j$, where $\Omega_j = \exp_p(U_j)$. Since each U_j is star-shaped, by Gauss lemma, on each boundary $\partial\Omega_j$, one has $\frac{\partial r}{\partial v} = g(\nabla r, v) \geq 0$ where v is the outer normal vector on $\partial\Omega_j$.

Therefore for each $\varphi \in C_c^\infty(M)$ with $\varphi \geq 0$, one has

$$\begin{aligned}
\int_M r \Delta \varphi \, \text{vol} &\stackrel{(1)}{=} - \int_M \langle \nabla r, \nabla \varphi \rangle \, \text{vol} \\
&\stackrel{(2)}{=} - \lim_j \int_{\Omega_j \setminus \{p\}} \langle \nabla r, \nabla \varphi \rangle \\
&\stackrel{(3)}{=} \lim_j \left(\int_{\Omega_j \setminus \{p\}} \Delta r \varphi \, \text{vol} - \int_{\partial \Omega_j} \varphi \frac{\partial r}{\partial v} \right) \\
&\stackrel{(4)}{\leq} \lim_j \int_{\Omega_j \setminus \{p\}} \Delta r \varphi \, \text{vol} \\
&\stackrel{(5)}{\leq} \lim_j \int_{\Omega_j \setminus \{p\}} (n-1) \frac{\text{sn}'_k(r)}{\text{sn}_k(r)} \varphi \, \text{vol} \\
&\stackrel{(6)}{=} \int_{\Omega \setminus \{p\}} (n-1) \frac{\text{sn}'_k(r)}{\text{sn}_k(r)} \, \text{vol} \\
&\stackrel{(7)}{=} \int_M (n-1) \frac{\text{sn}'_k(r)}{\text{sn}_k(r)} \varphi \, \text{vol}
\end{aligned}$$

where

- (1) holds from the fact r is Lipschitz and Proposition 16.1;
- (2) and (6) holds from dominated convergence theorem;
- (3) holds from Stokes theorem;
- (4) holds from $\varphi \geq 0$ and $\frac{\partial r}{\partial v} \geq 0$;
- (5) holds from Local Laplacian comparison theorem, that is Theorem 16.1.1;
- (7) holds from the fact $\text{cut}(p)$ is zero-measure.

□

16.4. Volume comparison.

Lemma 16.4.1. Let (M, g) be a complete, connected Riemannian manifold and $p \in M$. For any $\delta \in \mathbb{R}^+$

$$\exp_p(B(0, \delta) \cap \Sigma(p)) \subset B(p, \delta) \subset \exp_p(B(0, \delta) \cap \Sigma(p)) \cup \text{cut}(p)$$

In particular, under the map $\Phi : \mathbb{R}^+ \times \mathbb{S}^{n-1} \rightarrow T_p M \setminus \{0\}$ given by $\Phi(\rho, \omega) = \rho \omega$

$$\begin{aligned}
\text{Vol}(B(p, \delta)) &= \text{Vol}(\exp_p(B(0, \delta)) \cap \Sigma(p)) \\
&= \int_{B(0, \delta) \cap \Sigma(p)} \exp_p^* \, \text{vol} \\
&= \int_{B(0, \delta)} \chi_{\Sigma(p)} \exp_p^* \, \text{vol} \\
&= \int_{\mathbb{S}^{n-1}} \int_0^\delta \chi_{\Sigma(p)} \sqrt{\det g} \circ \Phi(\rho, \omega) \rho^{n-1} \, d\rho \, \text{vol}_{\mathbb{S}^{n-1}}
\end{aligned}$$

Corollary 16.4.1. Let $p \in S(n, k)$

1. If $k \leq 0$, then for any $\delta \in \mathbb{R}^+$

$$\text{Vol}(B(p, \delta)) = \int_{\mathbb{S}^{n-1}} \int_0^\delta \text{sn}_k^{n-1}(\rho) d\rho \text{vol}_{\mathbb{S}^{n-1}}$$

2. If $k = \frac{1}{R^2} \geq 0$, then for any $\delta \in \mathbb{R}^+$

$$\text{Vol}(B(p, \delta)) = \int_{\mathbb{S}^{n-1}} \int_0^\delta \chi_{B(0, \pi R)} \text{sn}_k^{n-1}(\rho) d\rho \text{vol}_{\mathbb{S}^{n-1}}$$

Lemma 16.4.2. Let (M, g) be a Riemannian manifold, and (x^i, U, p) be a geodesic ball chart of radius b around $p \in M$.

1. If $K \leq k$, then for each fixed $\omega \in \mathbb{S}^{n-1}$ the volume density ratio

$$\lambda(\rho, \omega) = \frac{\rho^{n-1} \sqrt{\det g} \circ \Phi(\rho, \omega)}{\text{sn}_k^{n-1}(\rho)}$$

is non-decreasing in $\rho \in (0, b_0)$ where

$$b_0 = \begin{cases} b, & k \leq 0 \\ \min\{b, \pi R\}, & k = \frac{1}{R^2} \end{cases}$$

Moreover, $\lim_{\rho \rightarrow 0} \lambda(\rho, \omega) = 1$.

2. If $K \geq k$ or $\text{Ric}(g) \geq (n-1)kg$, then for each fixed $\omega \in \mathbb{S}^{n-1}$ the volume density ratio $\lambda(\rho, \omega)$ is non-increasing in $\rho \in (0, b)$ and $\lim_{\rho \rightarrow 0} \lambda(\rho, \omega) = 1$.

Proof. By Corollary 15.2.1 and Lemma 16.1.1

$$\partial_r \log(r^{n-1} \sqrt{\det g}) = \Delta r \geq (n-1) \frac{\text{sn}'_k(r)}{\text{sn}_k(r)} = \partial_r \log(\text{sn}_k^{n-1}(r))$$

Hence $\log\left(\frac{r^{n-1} \sqrt{\det g}}{\text{sn}_k^{n-1}(r)}\right)$ is a non-decreasing function of r along each radial geodesic γ , that is

$$\frac{d}{dt} \left(\log \left(\frac{r^{n-1} \sqrt{\det g}}{\text{sn}_k^{n-1}(r)} \right) \circ \gamma(t) \right) \geq 0$$

Hence, $f(r) = \frac{r^{n-1} \sqrt{\det g}}{\text{sn}_k^{n-1}(r)}$ is a non-decreasing function of r along each radial geodesic γ . It is easy to see that $r \circ \Phi = \rho$ (the exponential map is used in normal coordinate). Hence,

$$\lambda(\rho, \omega) = f \circ \Phi(\rho, \omega)$$

is nondecreasing in ρ for any fixed $\omega \in \mathbb{S}^{n-1}$. It is obvious that

$$\lim_{\rho \rightarrow 0} \sqrt{\det g} = \lim_{\rho \rightarrow 0} \frac{\rho^{n-1}}{\text{sn}_k^{n-1}(\rho)} = 1.$$

The proof of (2) is similar. □

Lemma 16.4.3. Let $f : [0, +\infty) \rightarrow [0, +\infty)$, $g : [0, +\infty) \rightarrow (0, +\infty)$ be two integrable functions. If

$$\lambda(t) = \frac{f(t)}{g(t)} : [0, +\infty) \rightarrow [0, +\infty)$$

is non-increasing, then

$$F(t) = \frac{\int_0^t f(\tau) d\tau}{\int_0^t g(\tau) d\tau} : [0, +\infty) \rightarrow [0, +\infty)$$

is non-increasing. Moreover, if there exists $0 < t_1 < t_2$ such that

$$F(t_1) = F(t_2),$$

then $\lambda(t) \equiv \lambda(t_1)$ for almost all $t \in [0, t_2]$.

Proof. We can assume $f(t) > 0$ for all $t \in [0, +\infty)$, otherwise we replace it by $f(t) + \varepsilon g(t)$ for some $\varepsilon > 0$. Given $0 < t_1 < t_2$, we need to show

$$\int_0^{t_1} f(\tau) d\tau \int_0^{t_2} g(\tau) d\tau - \int_0^{t_2} f(\tau) d\tau \int_0^{t_1} g(\tau) d\tau \geq 0.$$

Indeed,

$$\begin{aligned} & \int_0^{t_1} f(\tau) d\tau \int_0^{t_2} g(\tau) d\tau - \int_0^{t_2} f(\tau) d\tau \int_0^{t_1} g(\tau) d\tau \\ &= \int_0^{t_1} f(\tau) d\tau \int_0^{t_2} g(\tau) d\tau - \int_0^{t_1} f(\tau) d\tau \int_0^{t_1} g(\tau) d\tau - \int_{t_1}^{t_2} f(\tau) d\tau \int_0^{t_1} g(\tau) d\tau \\ &= \int_0^{t_1} f(\tau) d\tau \int_{t_1}^{t_2} g(\tau) d\tau - \int_{t_1}^{t_2} f(\tau) d\tau \int_0^{t_1} g(\tau) d\tau \\ &\stackrel{(1)}{\geq} \int_0^{t_1} \frac{f(t_1)}{g(t_1)} g(\tau) d\tau \int_{t_1}^{t_2} \frac{g(t_1)}{f(t_1)} f(\tau) d\tau - \int_{t_1}^{t_2} f(\tau) d\tau \int_0^{t_1} g(\tau) d\tau \\ &= 0 \end{aligned}$$

where (1) holds from $\lambda(t)$ is non-increasing. It is clear that if $F(t_1) = F(t_2)$, then the inequality marked by (1) is an equality, which implies for almost all $t \in [0, t_2]$, $\lambda(t) \equiv \lambda(t_1)$. \square

Remark 16.4.1. For any $0 \leq \delta_1 < \delta_2 \leq \delta_3 < \delta_4$, we can slightly adapt above proof to show

$$\frac{\int_{\delta_3}^{\delta_4} f(\tau) d\tau}{\int_{\delta_3}^{\delta_4} g(\tau) d\tau} \leq \frac{\int_{\delta_1}^{\delta_2} f(\tau) d\tau}{\int_{\delta_1}^{\delta_2} g(\tau) d\tau}$$

Indeed, just note that

$$\begin{aligned}
& \int_{\delta_3}^{\delta_4} f(\tau) d\tau \int_{\delta_2}^{\delta_1} g(\tau) d\tau - \int_{\delta_1}^{\delta_2} f(\tau) d\tau \int_{\delta_3}^{\delta_4} g(\tau) d\tau \\
& \leq \int_{\delta_3}^{\delta_4} \frac{f(\delta_3)}{g(\delta_3)} g(\tau) d\tau \int_{\delta_2}^{\delta_1} g(\tau) d\tau - \int_{\delta_1}^{\delta_2} \frac{f(\delta_2)}{g(\delta_2)} g(\tau) d\tau \int_{\delta_3}^{\delta_4} g(\tau) d\tau \\
& = \left(\frac{f(\delta_3)}{g(\delta_3)} - \frac{f(\delta_2)}{g(\delta_2)} \right) \int_{\delta_2}^{\delta_1} g(\tau) d\tau \int_{\delta_3}^{\delta_4} g(\tau) d\tau \\
& \leq 0
\end{aligned}$$

Theorem 16.4.1 (Bishop-Gromov). Let (M, g) be a complete Riemannian manifold and $p \in M$. Let $B(p, \delta)$ be the metric ball centered at p with radius δ and g_k be the metric with constant sectional curvature k on $B(p, \delta) \setminus \{p\}$.

1. Suppose $K \leq k$, then the volume ratio $\frac{\text{Vol}_g(B(p, \delta))}{\text{Vol}_{g_k}(B(p, \delta))}$ is non-decreasing for any $0 < \delta \leq \delta_0$ where $\delta_0 = \text{inj}(p)$ if $k \leq 0$, and $\delta_0 = \min\{\text{inj}(p), \pi/\sqrt{k}\}$ if $k > 0$. Moreover,

$$\lim_{\delta \rightarrow 0} \frac{\text{Vol}_g(B(p, \delta))}{\text{Vol}_{g_k}(B(p, \delta))} = 1.$$

In particular,

$$\text{Vol}_g(B(p, \delta)) \geq \text{Vol}_{g_k}(B(p, \delta)),$$

2. If $K \geq k$ or $\text{Ric}(g) \geq (n-1)kg$, then the volume ratio $\frac{\text{Vol}_g(B(p, \delta))}{\text{Vol}_{g_k}(B(p, \delta))}$ is non-increasing for $\delta \in \mathbb{R}^+$. Moreover,

$$\lim_{\delta \rightarrow 0} \frac{\text{Vol}_g(B(p, \delta))}{\text{Vol}_{g_k}(B(p, \delta))} = 1$$

In particular,

$$\text{Vol}_g(B(p, \delta)) \leq \text{Vol}_{g_k}(B(p, \delta)),$$

3. Furthermore, if there exists $\delta_1 < \delta_2$ such that

$$\frac{\text{Vol}_g(B(p, \delta_1))}{\text{Vol}_{g_k}(B(p, \delta_1))} = \frac{\text{Vol}_g(B(p, \delta_2))}{\text{Vol}_{g_k}(B(p, \delta_2))}$$

then $\text{Vol}_g(B(p, \delta)) = \text{Vol}_{g_k}(B(p, \delta))$ for any $\delta \in [0, \delta_2]$ and g has constant sectional curvature k on $B(p, \delta_2)$.

Proof. For (1). By the assumption, we know the metric ball $B(p, \delta)$ is actually a geodesic ball. We have the expression

$$\begin{aligned}
\frac{\text{Vol}_g(B(p, \delta))}{\text{Vol}_{g_k}(B(p, \delta))} & \stackrel{\text{I}}{=} \frac{\int_{\mathbb{S}^{n-1}} \int_0^\delta \rho^{n-1} \sqrt{\det g} \circ \Phi(\rho, \omega) d\rho d \text{Vol}_{\mathbb{S}^{n-1}}}{\int_{\mathbb{S}^{n-1}} \int_0^\delta \text{sn}_k^{n-1}(\rho) d\rho d \text{Vol}_{\mathbb{S}^{n-1}}} \\
& \stackrel{\text{II}}{=} \frac{1}{\text{Vol}(\mathbb{S}^{n-1})} \int_{\mathbb{S}^{n-1}} \left(\frac{\int_0^\delta \rho^{n-1} \sqrt{\det g} \circ \Phi(\rho, \omega) d\rho}{\int_0^\delta \text{sn}_k^{n-1}(\rho) d\rho} \right) d \text{Vol}_{\mathbb{S}^{n-1}}.
\end{aligned}$$

where

I holds from Lemma 16.4.1;

II holds from Fubini's theorem.

By Lemma 16.4.2, one has $\lambda(\rho, \omega) = \frac{\rho^{n-1} \sqrt{\det g} \circ \Phi(\rho, \omega)}{\text{sn}_k^{n-1}(\rho)}$ is non-decreasing in ρ , then by Lemma 16.4.3 we have $\frac{\text{Vol}_g(B(p, \delta))}{\text{Vol}_{g_k}(B(p, \delta))}$ is non-decreasing in ρ . On the other hand,

$$\lim_{\delta \rightarrow 0} \frac{\text{Vol}_g(B(p, \delta))}{\text{Vol}_{g_k}(B(p, \delta))} = 1$$

Hence, for any $0 < \delta \leq \delta_0$, $\text{Vol}_g(B(p, \delta)) \geq \text{Vol}_{g_k}(B(p, \delta))$

For (2). Let's divide into the following two cases:

(a) If $k \leq 0$, for any $\delta \in \mathbb{R}^+$, we get

$$\begin{aligned} \frac{\text{Vol}_g(B(p, \delta))}{\text{Vol}_{g_k}(B(p, \delta))} &= \frac{\int_{\mathbb{S}^{n-1}} \int_0^\delta \chi_{\Sigma(p)} \rho^{n-1} \sqrt{\det g} \circ \Phi(\rho, \omega) d\rho d \text{Vol}_{\mathbb{S}^{n-1}}}{\int_{\mathbb{S}^{n-1}} \int_0^\delta \text{sn}_k^{n-1}(\rho) d\rho d \text{Vol}_{\mathbb{S}^{n-1}}} \\ &= \frac{1}{\text{Vol}(\mathbb{S}^{n-1})} \int_{\mathbb{S}^{n-1}} \left(\frac{\int_0^\delta \chi_{\Sigma(p)} \rho^{n-1} \sqrt{\det g} \circ \Phi(\rho, \omega) d\rho}{\int_0^\delta \text{sn}_k^{n-1}(\rho) d\rho} \right) d \text{Vol}_{\mathbb{S}^{n-1}}. \end{aligned}$$

where these two equalities hold from the same reasons. So in this case we consider

$$\tilde{\lambda}(\rho, \omega) := \chi_{\Sigma(p)} \lambda(\rho, \omega)$$

It's clear $\tilde{\lambda}$ is also non-increasing in ρ , since $\chi_{\Sigma(p)}$ is just a cut-off function, then the same argument implies for arbitrary $\delta \in \mathbb{R}^+$, one has $\text{Vol}_g(B(p, \delta)) \leq \text{Vol}_{g_k}(B(p, \delta))$.

(b) If $k = \frac{1}{R^2} > 0$, for any $\delta \in \mathbb{R}^+$, we get

$$\begin{aligned} \frac{\text{Vol}_g(B(p, \delta))}{\text{Vol}_{g_k}(B(p, \delta))} &= \frac{\int_{\mathbb{S}^{n-1}} \int_0^\delta \chi_{\Sigma(p)} \rho^{n-1} \sqrt{\det g} \circ \Phi(\rho, \omega) d\rho d \text{Vol}_{\mathbb{S}^{n-1}}}{\int_{\mathbb{S}^{n-1}} \int_0^\delta \chi_{B(0, \pi R)} \text{sn}_k^{n-1}(\rho) d\rho d \text{Vol}_{\mathbb{S}^{n-1}}} \\ &= \frac{1}{\text{Vol}(\mathbb{S}^{n-1})} \int_{\mathbb{S}^{n-1}} \left(\frac{\int_0^\delta \chi_{\Sigma(p)} \rho^{n-1} \sqrt{\det g} \circ \Phi(\rho, \omega) d\rho}{\int_0^\delta \chi_{B(0, \pi R)} \text{sn}_k^{n-1}(\rho) d\rho} \right) d \text{Vol}_{\mathbb{S}^{n-1}}. \end{aligned}$$

So in this case we consider⁸

$$\tilde{\lambda}(\rho, \omega) := \frac{\chi_{\Sigma(p)}}{\chi_{B(0, \pi R)}} \lambda(\rho, \omega)$$

Then the same argument shows the result.

For (3).

□

⁸Be careful, our notation here is a little bit ambiguous, since it's nonsense if $\chi_{B(0, \pi R)} = 0$. However, Myers' theorem implies $\text{diam}(M, g) \leq \pi R$, hence $\Sigma(p) \subset B(0, \pi R)$, so here the explicit means of $\frac{\chi_{\Sigma(p)}}{\chi_{B(0, \pi R)}}$ is as follows

$$\frac{\chi_{\Sigma(p)}}{\chi_{B(0, \pi R)}} = \begin{cases} 1, & \delta \in \Sigma(p) \\ 0, & \text{otherwise} \end{cases}$$

Corollary 16.4.2. Let (M, g) be a complete Riemannian n -manifold with $\text{Ric}(g) \geq 0$. Then the volume growth of (M, g) satisfies

$$\text{Vol}_g(B(p, r)) \leq c_n r^n$$

where c_n is a constant > 0 depending only on n .

Proof. Consider $k = 0$ and use Theorem 16.4.1, one has

$$\text{Vol}_g(B(p, r)) \leq \text{Vol}_{g_0}(B(p, r)) = \frac{\text{Vol}_{g_1}(\mathbb{S}^{n-1})r^n}{n}$$

where \mathbb{S}^{n-1} is the unit sphere. Thus we just set $c_n = \text{Vol}_{g_1}(\mathbb{S}^{n-1})/n$ to conclude. \square

Corollary 16.4.3. Let (M, g) be a complete Riemannian n -manifold with $\text{Ric}(g) \geq 0$. If

$$\lim_{r \rightarrow \infty} \frac{\text{Vol}_g(B(p, r))}{r^n} \geq \frac{\text{Vol}_{g_1}(\mathbb{S}^{n-1})}{n}$$

where \mathbb{S}^{n-1} is unit sphere, then (M, g) is isometric to $(\mathbb{R}^n, g_{\text{can}})$.

Proof. Note that $\text{Vol}_{g_0}(B(p, r)) = \frac{\text{Vol}_{g_1}(\mathbb{S}^{n-1})r^n}{n}$, then our assumption is equivalent to say

$$\lim_{r \rightarrow \infty} \frac{\text{Vol}_g(B(p, r))}{\text{Vol}_{g_0}(B(p, r))} = 1$$

However, by Theorem 16.4.1 we know volume ratio $\frac{\text{Vol}_g(B(p, r))}{\text{Vol}_{g_0}(B(p, r))}$ is non-increasing, with

$$\lim_{r \rightarrow 0} \frac{\text{Vol}_g(B(p, r))}{\text{Vol}_{g_0}(B(p, r))} = 1$$

which implies $\frac{\text{Vol}_g(B(p, r))}{\text{Vol}_{g_0}(B(p, r))} = 1$ holds for arbitrary $r > 0$. By rigidity of volume comparison, we conclude g has constant sectional curvature 0 on $B(p, r)$ for arbitrary $r > 0$. Since $\overline{B(p, \infty)} = M$, we deduce (M, g) has constant sectional curvature 0.

Thanks to Hopf's theorem, now it suffices to show M is simply-connected, suppose $\pi : \mathbb{R}^n \rightarrow M$ is the universal covering, one deduces that

$$|\pi_1(M)| = \frac{\text{Vol}_{g_0}(\mathbb{R}^n)}{\text{Vol}_g(M)} = 1$$

which implies M is simply-connected. \square

Corollary 16.4.4. Let (M, g) be a complete Riemannian n -manifold with $\text{Ric}(g) \geq (n-1)kg$ for some constant $k > 0$. Then

$$\text{Vol}_g(M) \leq \text{Vol}_{g_k}(\mathbb{S}^n(\frac{1}{\sqrt{k}}))$$

If the equality holds, then (M, g) is isometric to $\mathbb{S}^n(1/\sqrt{k})$.

Proof. Let $k = 1/R^2$, then Myers's theorem implies $\text{diam}(M, g) \leq \pi R$, thus compact. Hence, for any $p \in M$ one has $\Sigma(p) \subset B(0, \pi R)$. Therefore

$$\text{Vol}_g(B(p, \pi R)) = \text{Vol}_g(M)$$

where $B(p, \pi R)$ is a metric ball in M . On the other hand, it is obvious that

$$\text{Vol}_{g_k}(B(p, \pi R)) = \text{Vol}_{g_k}(\mathbb{S}^n(R))$$

Hence by Theorem 16.4.1, one has

$$\text{Vol}_g(M) \leq \text{Vol}_{g_k}(\mathbb{S}^n(R))$$

Furthermore, if the equality holds, g has constant sectional curvature on $B(p, \pi R)$. Then use the argument in Corollary 16.4.3 completes the proof. \square

Corollary 16.4.5. Let (M, g) be a complete Riemannian manifold and $p \in M$. Let $B(p, \delta)$ be the metric ball centered at p with radius δ and g_k be the metric with constant sectional curvature k on $B(p, \delta) \setminus \{p\}$. If $\text{Ric}(g) \geq (n-1)kg$, then for any $0 \leq \delta_1 < \delta_2 \leq \delta_3 < \delta_4$

$$\frac{\text{Vol}_g(B(p, \delta_4)) - \text{Vol}_g(B(p, \delta_3))}{\text{Vol}_g(B(p, \delta_2)) - \text{Vol}_g(B(p, \delta_1))} \leq \frac{\text{Vol}_{g_k}(B(p, \delta_4)) - \text{Vol}_{g_k}(B(p, \delta_3))}{\text{Vol}_{g_k}(B(p, \delta_2)) - \text{Vol}_{g_k}(B(p, \delta_1))}$$

Proof. Just note that volume density ratio is non-decreasing, then by Remark 16.4.1, one has

$$\frac{\text{Vol}_g(B(p, \delta_4)) - \text{Vol}_g(B(p, \delta_3))}{\text{Vol}_{g_k}(B(p, \delta_4)) - \text{Vol}_{g_k}(B(p, \delta_3))} \leq \frac{\text{Vol}_g(B(p, \delta_2)) - \text{Vol}_g(B(p, \delta_1))}{\text{Vol}_{g_k}(B(p, \delta_2)) - \text{Vol}_{g_k}(B(p, \delta_1))}$$

This gives desired result. \square

Theorem 16.4.2 (Cheng). Let (M, g) be a complete Riemannian n -manifold with $\text{Ric}(g) \geq (n-1)kg$ for some constant $k > 0$. If $\text{diam}(M) = \pi/\sqrt{k}$, then (M, g) is isometric to $\mathbb{S}^n(1/\sqrt{k})$.

Proof. Let $k = 1/R^2$. There exist points $p, q \in M$ and $\text{dist}(p, q) = \pi R$, then for any $\delta \in (0, \pi R)$

$$B(p, \delta) \cap B(q, \pi R - \delta) = \emptyset$$

By Theorem 19

$$\begin{aligned} \text{Vol}_g(M) &\stackrel{(1)}{\geq} \text{Vol}_g(B(p, \delta)) + \text{Vol}_g(B(q, \pi R - \delta)) \\ &\stackrel{(2)}{\geq} \text{Vol}_{g_k}(B(p, \delta)) \frac{\text{Vol}_g(B(p, \pi R))}{\text{Vol}_{g_k}(B(p, \pi R))} + \text{Vol}_{g_k}(B(q, \pi R - \delta)) \frac{\text{Vol}_g(B(q, \pi R))}{\text{Vol}_{g_k}(B(q, \pi R))} \\ &\stackrel{(3)}{=} \text{Vol}_g(M) \end{aligned}$$

where

(1) holds from $B(p, \delta) \cap B(q, \pi R - \delta) = \emptyset$;

(2) holds from Theorem 16.4.1.

(3) holds since for any $x, y \in M$, $\text{Vol}_g(B(x, \pi R)) = \text{Vol}_g(M)$ and

$$\text{Vol}_{g_k}(B(x, \pi R)) = \text{Vol}_{g_k}(\mathbb{S}^n(R)), \quad \text{Vol}_{g_k}(B(x, \delta)) + \text{Vol}_{g_k}(B(y, \pi R - \delta)) = \text{Vol}_{g_k}(\mathbb{S}^n(R))$$

Hence, for any $0 < \delta < \pi R$.

$$\frac{\text{Vol}_g(B(p, \delta))}{\text{Vol}_{g_k}(B(p, \delta))} = \frac{\text{Vol}_g(B(p, \pi R))}{\text{Vol}_{g_k}(B(p, \pi R))} = \frac{\text{Vol}_g(M)}{\text{Vol}_{g_k}(\mathbb{S}^n(R))}.$$

Let $\delta \rightarrow 0$, and we deduce $\text{Vol}_g(M) = \text{Vol}_{g_k}(\mathbb{S}^n(R))$. By Proposition 16.4.4, (M, g) is isometric to $\mathbb{S}^n(R)$. \square

Theorem 16.4.3 (Bishop-Yau). Let (M, g) be a complete non-compact Riemannian n -manifold with $\text{Ric}(g) \geq 0$. Then the volume growth of (M, g) satisfies

$$c_n \text{Vol}_g(B(p, 1)) \cdot r \leq \text{Vol}_g(B(p, r))$$

for $r \geq 1$, where c_n is a positive constant depending only on n .

Proof. Let $x \in \partial B(p, 1+r)$, then

$$B(p, 1) \subset B(x, 2+r) \setminus B(x, r), \quad B(x, r) \subset B(p, 1+2r)$$

By Corollary 16.4.5, one has

$$\begin{aligned} \text{Vol}_g(B(p, 1)) &\leq \text{Vol}_g(B(x, 2+r)) - \text{Vol}_g(B(x, r)) \\ &\leq \text{Vol}_g(B(x, r)) \cdot \frac{\text{Vol}(B(x, 2+r)) - \text{Vol}(B(x, r))}{\text{Vol}(B(x, r))} \\ &\leq \text{Vol}_g(B(p, 1+2r)) \cdot \frac{(2+r)^n - r^n}{r^n} \\ &\leq \text{Vol}_g(B(p, 1+2r)) \cdot \frac{1}{r} c_n \end{aligned}$$

where $r \geq 1$. By changing variable, we obtain the lower bound. \square

Proposition 16.4.1. Let (M, g) be a Cartan-Hadamard manifold with $\text{Ric}(g) \leq -kg$ for some $k > 0$. Then for any $p \in M$

$$\text{Vol}_g(B(p, r)) \geq c_n e^{\sqrt{kr}}$$

where c_n is a positive constant depending only on n .

Proposition 16.4.2 (Cheeger-Colding). For each integer $n \geq 2$, there exists a real number $\delta(n) \in (0, 1)$ with the following property: if (M, g) is a compact Riemannian manifold of dimension n with $\text{Ric}(g) \geq (n-1)g$ and

$$\text{Vol}(M, g) \geq (1 - \delta(n)) \text{Vol}(\mathbb{S}^n)$$

then M is diffeomorphic to \mathbb{S}^n .

17. SPLITTING THEOREM

17.1. Geodesic rays.

Definition 17.1.1 (geodesic ray). A geodesic ray is a unit-speed geodesic $\gamma : [0, \infty) \rightarrow M$ such that for any $s, t \geq 0$,

$$\text{dist}(\gamma(s), \gamma(t)) = |s - t|$$

Lemma 17.1.1. Let (M, g) be a complete Riemannian manifold. then the following are equivalent:

1. M is non-compact;
2. For any $p \in M$, there exists a geodesic ray $\gamma : [0, \infty) \rightarrow M$ starting from p .

Proof. For (1) to (2). If M is non-compact, for any $p \in M$, there is a sequence of points $\{p_i\}$ such that $\text{dist}(p, p_i) = i$. Let $\gamma_i(t) = \exp_p(tv_i)$ be a unit-speed minimal geodesic connecting p and p_i , that is $\gamma_i(0) = p$ and $\gamma_i(i) = p_i$. By possibly passing to a subsequence, we may assume $v_i \rightarrow v \in T_p M$. Then

$$\gamma(t) = \exp_p(tv), \quad t \in [0, +\infty)$$

is a unit-speed geodesic ray. Indeed, for any $s, t \geq 0$, and for any $k > \max\{s, t\}$, one has

$$\text{dist}(\gamma_k(s), \gamma_k(t)) = |s - t|.$$

By continuity of exponential map \exp_p , one obtains

$$\text{dist}(\gamma(s), \gamma(t)) = \lim_{k \rightarrow +\infty} \text{dist}(\gamma_k(s), \gamma_k(t)) = |s - t|$$

Hence γ is a geodesic ray.

For (2) to (1). It's trivial. □

17.2. Buseman function.

Definition 17.2.1. Let (M, g) be a complete Riemannian manifold, $p \in M$ and $\gamma : [0, \infty) \rightarrow M$ be a geodesic ray starting from p . For any $t \geq 0$, $b_\gamma^t : M \rightarrow \mathbb{R}$ as

$$b_\gamma^t(x) := \text{dist}(x, \gamma(t)) - t$$

Proposition 17.2.1. Let (M, g) be a complete non-compact Riemannian manifold, $p \in M$ and γ be a geodesic ray starting from p . The function $b_\gamma^t(x) : M \rightarrow \mathbb{R}$ has the following properties:

1. For any fixed $x \in M$, $b_\gamma^t(x)$ is non-increasing in t .
2. For any $x \in M$ and $t \geq 0$, $|b_\gamma^t(x)| \leq \text{dist}(x, \gamma(0))$.
3. For any $x, y \in M$ and $t \geq 0$, $|b_\gamma^t(x) - b_\gamma^t(y)| \leq \text{dist}(x, y)$.

Proof. For (1). Note that for $t > s > 0$, one has

$$\begin{aligned} b_\gamma^t(x) - b_\gamma^s(t) &= \text{dist}(x, \gamma(t)) - \text{dist}(x, \gamma(s)) + s - t \\ &\leq \text{dist}(\gamma(t), \gamma(s)) + s - t \\ &= |t - s| + s - t \\ &= 0 \end{aligned}$$

For (2),(3). Directly from triangle inequality. \square

Definition 17.2.2 (Buseman function). The Buseman function with respect to the geodesic ray is defined as

$$b_\gamma := \lim_{t \rightarrow \infty} b_\gamma^t(x)$$

Example 17.2.1 (Buseman function on hyperbolic plane). Note that geodesics on $\mathbb{H} = \{(x, y) \in \mathbb{R}^2 \mid y \geq 0\}$ are

1. Semicircles centered on \mathbb{R} ;
2. Straight lines perpendicular to \mathbb{R} .

Given $x \in \mathbb{H}$, in order to compute Buseman function

$$b_\gamma(x) = \lim_{t \rightarrow \infty} \text{dist}(x, \gamma(t)) - \text{dist}(\gamma(0), \gamma(t))$$

It suffices to solve the following calculus: Fix $z_1, z_2 \in \mathbb{H}$ and $\alpha \in \partial\mathbb{H} = \mathbb{R} \cup \{\infty\}$, solve

$$(17.1) \quad \lim_{q \rightarrow \alpha} \text{dist}(q, z_1) - \text{dist}(q, z_2) = ?$$

then we can set $q = \gamma(t), \alpha = \gamma(\infty), z_1 = x, z_2 = \gamma(0)$ to conclude. Let's divide into several steps:

Step one: For arbitrary $r > s > 0$, the distance between ri, si in \mathbb{H} is $\ln \frac{r}{s}$, where i is imaginary number. Indeed, since metric on this line is exactly $\frac{dy \otimes dy}{y^2}$.

Step two: In hyperbolic planes, it's possible to use isometry to translate any two points to the positive imaginary axis. To be explicit, consider the Möbius transformation V mapping Poincaré disk \mathbb{D} to \mathbb{H} with inverse V^{-1} , given by

$$\begin{aligned} z &= V(w) = \frac{-iw + 1}{w - i} \\ w &= V^{-1}(z) = \frac{iz + 1}{z + i} \end{aligned}$$

Now for arbitrary $z_1, z_2 \in \mathbb{H}$, firstly use V^{-1} to send z_1, z_2 to $w_1, w_2 \in \mathbb{D}$ respectively, then let $S(w) = e^{i\theta} \frac{w - w_1}{1 - \bar{w}_1 w}$ be transformation in \mathbb{D} that send w_1 to 0, with θ chosen carefully so that w_2 get sent to the positive imaginary axis, that is, w_2 get sent to the point ki , where $k = |S(w_2)|$. Finally apply V to this situation, 0 gets sent to i and ki get sent to $\frac{1+k}{1-k}i$.

Step three: Combine step one and two, one can conclude that for arbitrary $z_1, z_2 \in \mathbb{H}$, the distance between them are

$$\text{dist}(z_1, z_2) = \ln \frac{1+k}{1-k}$$

If we express k in terms of z_1, z_2 , one has

$$\text{dist}(z_1, z_2) = \ln \frac{|z_1 + z_2| + |z_1 - z_2|}{|z_1 + z_2| - |z_1 - z_2|}$$

Step four: Consider a special case of (17.1), that is we assume $z_1 = ri, z_2 = i$, where $\ln r = \text{dist}(z_1, z_2)$. Now we choose a sequence $q_n = u_n + iv_n$ such that $u_n \rightarrow \alpha$ and $v_n \rightarrow v$, where $v = 0$ as $n \rightarrow \infty$. Then

$$\begin{aligned} \lim_{q \rightarrow \alpha} (\text{dist}(q, ri) - \text{dist}(q, i)) &= \lim_{n \rightarrow \infty} (\ln \left(\frac{|q_n + ri| + |q_n - ri|}{|q_n + ri| - |q_n - ri|} \right) - \ln \left(\frac{|q_n + i| + |q_n - i|}{|q_n + i| - |q_n - i|} \right)) \\ &= \lim_{n \rightarrow \infty} (\ln \left(\frac{|q_n + ri| + |q_n - ri|}{|q_n + i| + |q_n - i|} \right) + \ln \left(\frac{|q_n + i| - |q_n - i|}{|q_n + ri| - |q_n - ri|} \right)) \\ &= \lim_{n \rightarrow \infty} \ln \frac{\sqrt{u_n^2 + (v_n + r)^2} + \sqrt{u_n^2 + (v_n - r)^2}}{\sqrt{u_n^2 + (v_n + 1)^2} + \sqrt{u_n^2 + (v_n - 1)^2}} \\ &\quad + \lim_{n \rightarrow \infty} \ln \frac{\sqrt{u_n^2 + (v_n + 1)^2} - \sqrt{u_n^2 + (v_n - 1)^2}}{\sqrt{u_n^2 + (v_n + r)^2} - \sqrt{u_n^2 + (v_n - r)^2}} \\ &= \lim_{v_n \rightarrow 0} \ln \underbrace{\frac{\sqrt{\alpha^2 + (v_n + r)^2} + \sqrt{\alpha^2 + (v_n - r)^2}}{\sqrt{\alpha^2 + (v_n + 1)^2} + \sqrt{\alpha^2 + (v_n - 1)^2}}}_{\text{part I}} \\ &\quad + \lim_{v_n \rightarrow 0} \ln \underbrace{\frac{\sqrt{\alpha^2 + (v_n + 1)^2} - \sqrt{\alpha^2 + (v_n - 1)^2}}{\sqrt{\alpha^2 + (v_n + r)^2} - \sqrt{\alpha^2 + (v_n - r)^2}}}_{\text{part II}} \end{aligned}$$

It's clear Part I is $\frac{\sqrt{\alpha^2 + r^2}}{\sqrt{\alpha^2 + 1}}$, and apply L'Hospital's rule to Part II one has

$$\lim_{v_n \rightarrow 0} \frac{\sqrt{\alpha^2 + (v_n + 1)^2} - \sqrt{\alpha^2 + (v_n - 1)^2}}{\sqrt{\alpha^2 + (v_n + r)^2} - \sqrt{\alpha^2 + (v_n - r)^2}} = \lim_{v_n \rightarrow 0} \frac{\frac{v_n + 1}{\sqrt{\alpha^2 + (v_n + 1)^2}} - \frac{v_n - 1}{\sqrt{\alpha^2 + (v_n - 1)^2}}}{\frac{v_n + r}{\sqrt{\alpha^2 + (v_n + r)^2}} - \frac{v_n - r}{\sqrt{\alpha^2 + (v_n - r)^2}}} = \frac{\sqrt{\alpha^2 + r^2}}{r\sqrt{\alpha^2 + 1}}$$

which implies

$$\lim_{q \rightarrow \alpha} \text{dist}(q, ri) - \text{dist}(q, i) = \ln \frac{\alpha^2 + r^2}{\alpha^2 + 1} - \ln r$$

Step five: In order to solve general case of (17.1), we can use processes in step two to translate z_1, z_2 to the positive imaginary axis. However, α is also translated into a new point α' , that is

$$\alpha' = V \circ S \circ V^{-1}(\alpha)$$

where V, V^{-1} and S are defined in step two. Thus from step four one has

$$\lim_{q \rightarrow \alpha} \text{dist}(q, z_1) - \text{dist}(q, z_2) = \ln \frac{(\alpha')^2 + r^2}{(\alpha')^2 + 1} - \ln r$$

where $\ln r = \text{dist}(z_1, z_2)$.

Proposition 17.2.2. Let (M, g) be a complete non-compact Riemannian manifold, $p \in M$ and γ be a geodesic ray starting from p . The Busemann function $b_\gamma : M \rightarrow \mathbb{R}$ is Lipschitz continuous with $\text{Lip}(b_\gamma) \leq 1$

Proof. It follows from Arzela-Ascoli lemma. \square

Proposition 17.2.3. Let (M, g) be a complete non-compact Riemannian manifold, and γ be a geodesic ray starting from $p \in M$. If $\text{Ric}(g) \geq 0$, then

$$\Delta b_\gamma \leq 0$$

in the sense of distribution.

Proof. For any non-negative smooth function $\varphi \in C_0^\infty(M)$, one has

$$\begin{aligned} \int_M \Delta \varphi b_\gamma^t \text{vol} &= \int_M \Delta \varphi (\text{dist}(x, \gamma(t)) - t) \text{vol} \\ &\stackrel{(1)}{=} \int_M \Delta \varphi \text{dist}(x, \gamma(t)) \text{vol} \\ &\stackrel{(2)}{\leq} \int_M \frac{(n-1)\varphi}{\text{dist}(x, \gamma(t))} \text{vol} \end{aligned}$$

where

(1) holds from Stokes' theorem;

(2) holds from Theorem 16.3.2.

Then Lebesgue's dominated convergence implies

$$\int_M \Delta \varphi b_\gamma \text{vol} \leq 0$$

\square

Definition 17.2.3 (geodesic line). A geodesic line is a unit-speed geodesic $\gamma : (-\infty, \infty) \rightarrow M$ such that for any $s, t \in \mathbb{R}$,

$$\text{dist}(\gamma(s), \gamma(t)) = |s - t|$$

Lemma 17.2.1. Let (M, g) be a connected, non-compact Riemannian manifold. If M contains a compact subset K such that $M \setminus K$ has at least two unbounded components⁹, then there is a geodesic line passing through K .

Proof. Since $M \setminus K$ has at least two unbounded components, there are two unbounded sequences of points $\{p_i\}$ and $\{q_i\}$ such that any curve from p_i to q_i passes through K . Let $\gamma_i : [-a_i, b_i] \rightarrow M$ be minimal geodesics connecting p_i and q_i with $\gamma_i(-a_i) = p_i$, $\gamma_i(b_i) = q_i$ and $\gamma_i(0) \in K$. Hence, $a_i \rightarrow +\infty$

⁹Some authors use “ends” to call such unbounded components.

and $b_i \rightarrow +\infty$. By possibly passing to subsequences, $\{\gamma_i\}$ converges to a geodesic line $\gamma_\infty : (-\infty, +\infty) \rightarrow M$. \square

Proposition 17.2.4. Let (M, g) be a complete non-compact Riemannian manifold with $\text{Ric}(g) \geq 0$. If (M, g) contains a geodesic line γ , then $b_{\gamma_\pm} : M \rightarrow \mathbb{R}$ are smooth harmonic functions with

$$|\nabla b_{\gamma_\pm}| = 1, \quad \text{Hess } b_{\gamma_\pm} = 0$$

where $\gamma_\pm(t) = \gamma(\pm t) : [0, +\infty) \rightarrow M$.

Proof. Let $b(x) = b_{\gamma_+}(x) + b_{\gamma_-}(x)$. By the triangle inequality

$$\begin{aligned} b(x) &= \lim_{s \rightarrow +\infty} \text{dist}(x, \gamma_+(s)) + \text{dist}(x, \gamma_-(s)) - 2s \\ &= \lim_{s \rightarrow +\infty} \text{dist}(x, \gamma(s)) + \text{dist}(x, \gamma(-s)) - 2s \\ &\geq 0 \end{aligned}$$

By Proposition 17.2.3, $\Delta b \leq 0$ in the sense of distributions. On the other hand,

$$b(\gamma(t)) = \lim_{s \rightarrow +\infty} \text{dist}(\gamma(t), \gamma(s)) + \text{dist}(\gamma(t), \gamma(-s)) - 2s = 0$$

Hence the subharmonic function b attains its absolute minimum, by Theorem 16.2.1, $b \equiv 0$, that is $b_{\gamma_+} = -b_{\gamma_-}$. Hence $\Delta b_{\gamma_+} = \Delta b_{\gamma_-} = 0$, and by Wely's lemma one has b_{γ_\pm} are smooth.

Bochner's formula says

$$\frac{1}{2} \Delta |\nabla f|^2 = |\text{Hess } f|^2 + \text{Ric}(\nabla f, \nabla f) + g(\nabla \Delta f, \nabla f)$$

Let $f = b_{\gamma_+}$, then

$$\frac{1}{2} \Delta |\nabla b_{\gamma_+}|^2 \geq |\text{Hess } b_{\gamma_+}|^2 \geq 0$$

since b_{γ_+} is harmonic and $\text{Ric}(g) \geq 0$, thus $|\nabla b_{\gamma_+}|^2$ is superharmonic. On the other hand, by Proposition 17.2.2, $\text{Lip}(b_{\gamma_+}) \leq 1$, and so $|\nabla b_{\gamma_+}| \leq 1$. Note that

$$b_{\gamma_+}(\gamma_+(t)) = \lim_{s \rightarrow +\infty} \text{dist}(\gamma_+(t), \gamma_+(s)) - s = \lim_{s \rightarrow +\infty} |t - s| - s = -t$$

For any $x = \gamma_+(t_0)$

$$|\nabla b_{\gamma_+}|(x) \stackrel{(1)}{=} |\nabla b_{\gamma_+}| |\gamma'_+(t_0)| \stackrel{(2)}{\geq} |\langle \nabla b_{\gamma_+}(x), \gamma'_+(t_0) \rangle| = 1$$

where

(1) holds from the trivial fact γ_+ is unit-speed;

(2) holds from Cauchy-Schwarz inequality.

Hence, the superharmonic function $|\nabla b_{\gamma_+}|^2$ attains its absolute maximum in M , hence $|\nabla b_{\gamma_+}|^2 \equiv 1$ on M . Again by the Bochner formula, one has $\text{Hess } b_{\gamma_+} = 0$. The same argument holds for b_{γ_-} , this completes the proof. \square

Lemma 17.2.2. Let (M, g) be a complete Riemannian manifold, and V a smooth vector field with $|V|_g \leq C$ for some constant C . Then V is a complete vector field.

Proof. We need to show the integral curve of V is globally defined, that is defined on \mathbb{R} . Suppose $\gamma : (a, b) \rightarrow M$ is an integral curve of M and $b < \infty$. For arbitrary $t, s \in (a, b)$, we have

$$\gamma(t) = \gamma(s) + \int_s^t V(\gamma(\tau)) d\tau$$

By using the boundedness of V , we can conclude that

$$|\gamma(t) - \gamma(s)| \leq C|t - s|$$

which implies $\gamma(t)$ is uniformly continuous on (a, b) , thus it's possible to extend γ to $(a, b]$ since $b < \infty$, a contradiction. \square

Proposition 17.2.5. Let (M, g) be a complete Riemannian manifold. Suppose $f \in C^\infty(M, \mathbb{R})$ satisfies

$$|\nabla f| = 1 \quad \text{and} \quad \text{Hess } f = 0.$$

Let Σ denote $f^{-1}(0)$, with induced metric $h := g|_\Sigma$.

1. (Σ, h) is a totally geodesic submanifold of (M, g) .
2. The map

$$F : (\mathbb{R} \times \Sigma, g_\mathbb{R} \oplus h) \rightarrow (M, g), \quad F(t, p) = \exp_p(t \nabla_p f)$$

is an isometry.

Proof. For (1). Recall that (Σ, h) is a totally geodesic submanifold of (M, g) if the second fundamental form of Σ vanishes, and facts in basic differential geometry says the second fundamental form of a hyperplane Σ with induced metric is given by

$$\mathbf{II}(v, w) := \langle \nabla_v n, w \rangle$$

where n is the normal vector of Σ . In this case, if we consider $\Sigma = f^{-1}(0)$, then the normal vector of Σ is exactly ∇f , and thus

$$\mathbf{II}(v, w) := \langle \nabla_v \nabla f, w \rangle$$

Then $\text{Hess } f = \nabla^2 f = 0$ implies the second fundamental form of Σ vanishes, that is Σ is a totally geodesic submanifold of (M, g) .

For (2). For a fixed p , let $X = \nabla f$, and consider $\gamma(t) = \exp_p(tX_p)$. Since $\nabla X = 0$, we have $E(t) = X(\gamma(t))$ and $\gamma'(t)$ are two parallel vector fields along γ with the same initial value. Hence

$$\gamma'(t) = X(\gamma(t))$$

that is γ is exactly the integral curve of X . Furthermore, since $|X| = 1$, by Lemma 17.2.2 one has γ is globally defined, and one can deduce F is a global flow of X , thus it's a diffeomorphism.

Now it remains to prove that F is an isometry. For $v \in T_p \Sigma$, let J be the Jacobi field along γ with $J(0) = 0$ and $J'(0) = v$. By the radial curvature equation

$$R(-, \nabla f, \nabla f, -) = \text{Hess}\left(\frac{1}{2}|\nabla f|^2\right)(-, -) - (\nabla_{\nabla f} \text{Hess } f)(-, -) - \text{Hess } f(\nabla_- \nabla f, -)$$

one has $R(-, \nabla f, \nabla f, -) = 0$, thus Jacobi equation

$$J''(t) + R(J, \gamma')\gamma' = 0$$

reduces to $J''(t) = 0$. It implies that $J'(t)$ is a parallel vector field and in particular, $|J'(t)| \equiv |J'(0)| = |v|$. By uniqueness of Jacobi fields, we deduce

$$J(t) = tJ'(t)$$

Then F is an isometry holds as follows:

- (a) It is easy to see that $(dF)_{(1,p)}v = J(1)$, thus $|(dF)_{(1,p)}v| = |J(1)| = |J'(1)| = |v|$;
- (b) $|(dF)_{(0,p)}\partial_t| = |\nabla f| = 1 = |\partial_t|$.

□

17.3. Splitting theorem and its corollaries.

Theorem 17.3.1 (splitting theorem). Let (M, g) be a complete Riemannian n -manifold with $\text{Ric}(g) \geq 0$. If there is a geodesic line in M , then (M, g) is isometric to $(\mathbb{R} \times N, g_{\mathbb{R}} \oplus g_N)$, where $\text{Ric}(g_N) \geq 0$

Proof. Directly from Proposition 17.2.4 and Proposition 17.2.5. □

Corollary 17.3.1. Let (M, g) be a complete Riemannian n -manifold with $\text{Ric}(g) \geq 0$

1. (M, g) is isometric to $(\mathbb{R}^k \times N, g_{\mathbb{R}^k} \oplus g_N)$, where N does not contain a geodesic line and $\text{Ric}(g_N) \geq 0$.
2. The isometry group splits

$$\text{Iso}(M, g) \cong \text{Iso}(\mathbb{R}^k) \times \text{Iso}(N, g_N)$$

Theorem 17.3.2 (structure theorem for manifold with $\text{Ric} \geq 0$). Let (M, g) be a compact Riemannian manifold with $\text{Ric}(g) \geq 0$, and $\pi : (\widetilde{M}, \widetilde{g}) \rightarrow (M, g)$ is its universal covering with the pullback metric.

1. There exists some interger $k \geq 0$ and a compact Riemannian manifold (N, g_N) with $\text{Ric}(g_N) \geq 0$ such that $(\widetilde{M}, \widetilde{g})$ is isometric to $(\mathbb{R}^k \times N, g_{\text{can}} \oplus g_N)$.
2. The isometry group splits

$$\text{Iso}(\widetilde{M}, \widetilde{g}) \cong \text{Iso}(\mathbb{R}^k) \times \text{Iso}(N, g_N)$$

Proof. For (1). Suppose to the contrary that N is non-compact, then fix a point $x_0 \in N$, there exists a geodesic ray $\gamma : [0, \infty) \rightarrow N$ starting from x_0 . Since M is compact, there exists a compact subset $\widetilde{K} \subset \widetilde{M}$ such that

$$\text{Aut}_{\pi}(\widetilde{M})\widetilde{K} = \widetilde{M}$$

□

Corollary 17.3.2. $\mathbb{S}^n \times \mathbb{S}^1$ doesn't admit any Ricci flat metrics when $n = 2, 3$.

Proof. If $\mathbb{S}^n \times \mathbb{S}^1$ admits a Ricci flat metric, after splitting its universal covering we obtain a Ricci flat metric on \mathbb{S}^p . However, \mathbb{S}^n doesn't admit such a metric when $n = 2, 3$. Indeed, since any Einstein manifold with dimension 2 or 3 has constant sectional curvature, thus if $\mathbb{S}^n, n = 2, 3$ admit a Ricci flat metric, then it has constant sectional curvature 0, and it's also simply-connected, so Hopf's theorem implies it's diffeomorphic to \mathbb{R}^n , a contradiction. □

Remark 17.3.1. It's clear $\mathbb{S}^1 \times \mathbb{S}^1$ admits a Ricci flat metric, and when $n \geq 4$, we don't know whether \mathbb{S}^n admit a Ricci flat metric or not.

Corollary 17.3.3. Let (M, g) be a compact Riemannian manifold with $\text{Ric}(g) \geq 0$, and $(\widetilde{M}, \widetilde{g})$ is its universal covering equipped with pullback metric.

1. If \widetilde{M} is contractible, then $(\widetilde{M}, \widetilde{g})$ is isometric to $(\mathbb{R}^n, g_{\text{can}})$ and (M, g) is flat;
2. If $(\widetilde{M}, \widetilde{g})$ doesn't contain a geodesic line, then $\pi_1(M)$ is finite and $b_1(M) = 0$;

Proof. For (1). If $\widetilde{M} \cong N \times \mathbb{R}^k$ is contractible, we must have N is just a point, since it's compact,

For (2). If \widetilde{M} doesn't contain a geodesic line, then \widetilde{M} is compact, which implies $|\pi_1(M)|$ is finite. Furthermore, since there is a natural Hurwicz surjective

$$h : \pi_1(M) \rightarrow H_1(M; \mathbb{Z})$$

thus $H_1(M; \mathbb{Z})$ can't have free part, otherwise h can't be surjective, since there is no surjective map from a finite group to an infinite one. So we have $b_1(M) = 0$. □

Corollary 17.3.4. Let (M, g) be a compact Riemannian manifold with $\text{Ric}(g) > 0$. If there exists a point $p \in M$ such that $\text{Ric}(g) > 0$ on $T_p M$, then $\pi_1(M)$ is finite and $b_1(M) = 0$.

Proof. Since $\text{Ric}(g) > 0$ on the whole tangent space $T_p M$, the universal covering $(\widetilde{M}, \widetilde{g})$ can't split into a product $(\mathbb{R}^k \times N, g_{\text{can}} \oplus g_N)$, since metric on \widetilde{M} is pullback metric, and g_{can} on \mathbb{R}^k has vanishing Ricci curvature. Thus \widetilde{M} is compact, consequently we have $|\pi_1(M)|$ is finite and $b_1(M) = 0$. □

Corollary 17.3.5. Let (M, g) be a compact Riemannian n -manifold with $\text{Ric}(g) > 0$, then $b_1(M) \leq n$. Moreover, if $b_1(M) = n$ if and only if (M, g) is isometric to some flat torus $(\mathbb{T}^n, g_{\mathbb{T}^n})$.

Proof.

□

Part 6. Symmetric spaces

18. WHAT IS SYMMETRIC SPACE?

18.1. Basic settings.

Definition 18.1.1 (symmetric space). A Riemannian manifold (M, g) is called a symmetric space if for each $p \in M$ there exists an isometry $\varphi : M \rightarrow M$ such that $\varphi(p) = p$ and $(d\varphi)_p = -\text{id}$.

Remark 18.1.1. Here φ is called a symmetry at point p . Note that Theorem 13.1.4, that is rigidity property of isometry implies if symmetry at point p exists, it's unique.

Definition 18.1.2 (local symmetric space). Riemannian manifold (M, g) is called a local symmetric space if each $p \in M$ has a neighborhood U on which there exists an isometry $\varphi : U \rightarrow U$ such that $\varphi(p) = p$ and $(d\varphi)_p = -\text{id}$.

Lemma 18.1.1. The following are equivalent:

1. (M, g) is a symmetric space;
2. For each $p \in M$, there exists an isometry $\varphi : M \rightarrow M$ such that $\varphi^2 = -\text{id}$ and p is an isolated fixed point of φ .

Proof. For (1) to (2). Note that $(d\varphi^2)_p = (d\varphi)_p \circ (d\varphi)_p = \text{id}$ and $\varphi^2(p) = p$, then by Theorem, one has $\varphi^2 = \text{id}$. If p is not an isolated fixed point, then there exists a sequence $p_i \rightarrow p$ such that $\varphi(p_i) = p_i$. Let $\delta \in (0, \text{inj}(M))$, $q \in B(0, \delta)$ and $v = \exp_p^{-1}(q)$. Note that $\varphi(\exp_p(tv))$ and $\exp_p(tv)$ are two geodesics connecting p and q , since φ is an isometry, thus

$$\varphi(\exp_p(tv)) = \exp_p(tv)$$

by uniqueness. In particular, one has $v = (d\varphi)_p v$, a contradiction.

For (2) to (1). From $\varphi^2 = \text{id}$ we have $(d\varphi)_p^2 = \text{id}$, so only possibly eigenvalues of $(d\varphi)_p$ are ± 1 . Now it suffices to show all eigenvalues of $(d\varphi)_p$ are -1 . Otherwise if it has an eigenvalue 1 , there exists some non-zero $v \in T_p M$ such that $(d\varphi)_p v = v$. Then

$$\varphi(\exp_p(tv)) = \exp_p(tv)$$

are the same geodesics in a sufficiently small neighborhood, since both $\varphi(\exp_p(tv))$ and $\exp_p(tv)$ are geodesics and they have the same starting point and direction. In particular, p is not an isolated fixed point, a contradiction. \square

Example 18.1.1. Consider Euclidean space $(\mathbb{R}^n, g_{\text{can}})$, for each $p \in \mathbb{R}^n$, then the reflection $s_p(x) = 2p - x$ is a symmetry at point p .

Example 18.1.2. Consider the unit sphere $(\mathbb{S}^n, g_{\text{can}})$, for each $p \in \mathbb{S}^n$, $s_p(x) = 2\langle x, n \rangle p - x$ is a symmetry at point p , where n is outer normal vector at point p .

18.2. Riemannian homogeneous space. Let's recall some basic facts in differential geometry.

Definition 18.2.1 (smooth Lie group action). A Lie group G acts on a smooth manifold M smoothly, if the following conditions are satisfied:

1. Every $g \in G$ induces a diffeomorphism of M , denoted by $x \rightarrow gx$, where $x \in M$.
2. The map $G \times M \rightarrow M$ given by $(g, x) \mapsto gx$ is smooth.
3. For $g_1, g_2 \in G$ and $x \in M$, $(g_1 g_2)x = g_1(g_2 x)$.

Definition 18.2.2 (G -homogeneous space). A smooth manifold M endowed with a transitive smooth G -action is called a homogeneous G -space, where G is a Lie group.

Here are some tools which can be used to construct homogeneous manifolds. In fact, the most interesting examples of homogeneous space comes from this construction.

Theorem 18.2.1. Let G be a Lie group and H be a closed subgroup of G . Then

1. The left coset space G/H is a topological manifold of dimension $\dim G - \dim H$;
2. G/H admits a smooth structure, such that the quotient map $\pi : G \rightarrow G/H$ is a smooth submersion;
3. The left action

$$\begin{aligned} G \times G/H &\rightarrow G/H \\ (g_1, g_2 H) &\mapsto (g_1 g_2) H \end{aligned}$$

turns G/H into a homogeneous G -space.

Now let's consider the same things under stage of Riemannian manifold.

Definition 18.2.3 (Riemannian G -homogeneous space). If Lie group G acts smoothly and transitively as isometries on a Riemannian manifold (M, g) , then (M, g) is called a Riemannian G -homogeneous space.

Theorem 18.2.2 (Myers-Steenrod). Let G denote the isometry group of a Riemannian manifold (M, g) . Then

1. G is a Lie group with respect to compact-open topology;
2. G acts on M smoothly;
3. For each $p \in M$, the isotropy group G_p is compact;
4. If M is compact, then G is compact.

Proof. See [MS39]. □

Remark 18.2.1. Thanks to this theorem, we say a Riemannian manifold (M, g) is a Riemannian homogeneous space, if its isometry group G acts on it transitively.

Lemma 18.2.1. Let (M, g) be a Riemannian homogeneous space. If there exist a point $p \in M$ and an isometry $\varphi : M \rightarrow M$ such that $\varphi(p) = p$ and $(d\varphi)_p = -\text{id}$, then (M, g) is a Riemannian symmetric space.

Proof. Since (M, g) is a Riemannian homogeneous space, then isometry group acts on M transitively, thus for any $q \in M$, there exists an isometry $\psi : M \rightarrow M$ such that $\psi(p) = q$, then

$$\varphi_q := \psi \circ \varphi \circ \psi^{-1}$$

is an isometry such that $\varphi_q(q) = q$ and $(d\varphi_q)_q = -\text{id}$. \square

Remark 18.2.2. Thanks to this remark, if we want to show a homogeneous space is a symmetric space, it suffices to find the symmetry at some point, then we can construct symmetry at every point.

18.3. The relations between symmetric, local symmetric and homogeneous space. In this section, we will explain the relations between the following three spaces:

1. Symmetric space;
2. Local symmetric space;
3. Riemannian homogeneous space.

Firstly, let's give another characterization of locally symmetric space via curvature, which follows from the following lemma.

Lemma 18.3.1. Let (M, g) be a Riemannian manifold and $\gamma : (a, b) \rightarrow M$ be a smooth curve. Let

$$P_{s,t}^\gamma : T_{\gamma(s)}M \rightarrow T_{\gamma(t)}M$$

be the corresponding parallel transport. Then for any $s \in (a, b)$ with $v = \gamma'(s)$, one has

$$\nabla_v R = \left. \frac{d}{dt} \right|_{t=s} ((P_{s,t}^\gamma)^* R_{\gamma(t)}).$$

In particular, if $\nabla R = 0$, then

$$(P_{s,t}^\gamma)^* R_{\gamma(t)} = R_{\gamma(s)}$$

Proof. Let $v_1 = v$, then for any $v_2, v_3, v_4, v_5 \in T_{\gamma(s)}M$ we set $X_i(t) = P_{s,t}^\gamma v_i \in T_{\gamma(t)}M$, $i = 2, \dots, 5$. Hence

$$\begin{aligned} \nabla R(v_1, \dots, v_5) &= \lim_{t \rightarrow s} \nabla R(X_1, \dots, X_5) \\ &= \lim_{t \rightarrow s} X_1 R(X_2, \dots, X_5) - \sum_{i=2}^5 \widehat{R}(X_2, \dots, \widehat{\nabla_{\frac{d}{dt}} X_i}, \dots, X_5) \\ &= \lim_{t \rightarrow s} X_1 R(X_2, \dots, X_5) \\ &= \left. \frac{d}{dt} \right|_{t=s} R_{\gamma(t)}(X_2(t), \dots, X_5(t)) \\ &= \left. \frac{d}{dt} \right|_{t=s} (P_{s,t}^\gamma)^* R_{\gamma(t)}(v_2, \dots, v_5) \end{aligned}$$

□

Theorem 18.3.1. Let (M, g) be a complete Riemannian manifold, the following are equivalent:

1. (M, g) is a local symmetric space;
2. $\nabla R = 0$.

Proof.

□

Theorem 18.3.2. Let (M, g) be a complete, simply-connected local symmetric space, then (M, g) is a symmetric space.

Proof.

□

Corollary 18.3.1. Let (M, g) be a complete local symmetric space, then it's isometric to \widetilde{M}/Γ , where \widetilde{M} is a symmetric space and Γ is a discrete Lie group that acts freely, properly and isometrically on \widetilde{M} .

Proof. Let $(\widetilde{M}, \widetilde{g})$ be the universal covering of (M, g) with pullback metric, then

□

Theorem 18.3.3. Let (M, g) be a symmetric space, then

1. (M, g) is complete;
2. For any isometry $\varphi : M \rightarrow M$ with $(d\varphi)_p = -\text{id}$ and $\varphi(p) = p$, if $v \in T_p M$, then

$$\varphi(\exp_p(v)) = \exp_p(-v)$$

3. The isometry group $\text{Iso}(M, g)$ acts transitively on M .

Proof. For (1). For arbitrary geodesic $\gamma : [0, 1] \rightarrow M$ with $\gamma(0) = p, \gamma'(0) = v$. the curve $\beta(t) = \varphi(\gamma(t)) : [0, 1] \rightarrow M$ is also a geodesic with $\beta(0) = p$ and $\beta'(0) = -v$. Now we obtain a smooth extension $\gamma' : [0, 2] \rightarrow M$ of γ , given by

$$\gamma'(t) = \begin{cases} \beta(1-t), & t \in [0, 1] \\ \gamma(t-1), & t \in [1, 2] \end{cases}$$

Repeat above process to extend γ to a geodesic defined on \mathbb{R} , this shows completeness.

For (2). Just consider geodesics $\varphi(\exp_p(tv)) = \exp_p(-tv)$.

For (3). Let p, q be any two points in M and $\gamma : [0, 1] \rightarrow M$ be a geodesic with $\gamma(0) = p, \gamma(1) = q$. Let $m = \gamma(\frac{1}{2})$ and $\varphi_m : M \rightarrow M$ the symmetry at m . Consider $\beta(t) = \varphi_m(\gamma(\frac{1}{2}-t))$, then $\beta(0) = m, \beta'(0) = \gamma'(\frac{1}{2})$, which implies $\beta(t) = \gamma(\frac{1}{2}+t)$. Therefore $q = \gamma(1) = \beta(\frac{1}{2}) = \varphi_m(\gamma(0)) = \varphi_m(p)$. □

Corollary 18.3.2. Symmetric space (M, g) is a Riemannian homogeneous space.

Finally, we introduce an interesting topological property of symmetric space.

Proposition 18.3.1. Let (M, g) be a symmetric space, then $\pi_1(M)$ is abelian.

Proof. The key points are the following observations:

1. Any homotopy class in $\pi_1(M)$ can be represented by a closed geodesic of minimal length, by a shortening process;
2. For a closed geodesic γ , if we consider symmetry s at point $p = \gamma(0)$, then $s \circ \gamma = \gamma^{-1}$. Furthermore $s \circ (\gamma_1 \circ \gamma_2) = \gamma_1^{-1} \circ \gamma_2^{-1}$, where $\gamma_i, i = 1, 2$ are two closed geodesics with $\gamma_i(0) = p$;
3. The inversion is a group homomorphism if and only if group is abelian.

From (1), we may use closed geodesics to represent elements in $\pi_1(M)$, then inversion is a group homomorphism on $\pi_1(M)$, since from (2) we know it's exactly symmetry s , so (3) implies $\pi_1(M)$ is abelian. \square

Remark 18.3.1. Later we will show that any Lie group is a symmetric space. In particular, fundamental group of a Lie group is abelian.

18.4. Killing field as a Lie algebra of isometry group.

Proposition 18.1. Let (M, g) be a Riemannian manifold and X be a Killing field.

1. If γ is a geodesic, then $J(t) = X \circ \gamma(t)$ is a Jacobi field;
2. For any two vector fields Y, Z ,

$$\nabla_Y \nabla_Z X - \nabla_{\nabla_Y Z} X + R(X, Y)Z = 0$$

Proof. For (1). It's known that X is a Killing field if and only if it generates a flow φ_s by local isometries. Thus for a geodesic γ , we can obtain a variation of geodesics via $\alpha(s, t) = \varphi_s(\gamma(t))$, thus

$$X \circ \gamma(t) = \left. \frac{\partial \varphi_s(\gamma(t))}{\partial s} \right|_{s=0}$$

is a Jacobi field.

For (2). Without lose of generality we may consider normal coordinate $\{x^i\}$ centered at p and assume $X = X^i \frac{\partial}{\partial x^i}, Y = \frac{\partial}{\partial x^j}, Z = \frac{\partial}{\partial x^k}$. Then

$$\begin{aligned} \nabla_Y \nabla_Z X - \nabla_{\nabla_Y Z} X + R(X, Y)Z &= \nabla_j \nabla_k X + X^i R_{ijk}^l \frac{\partial}{\partial x^l} \\ &= \left(\frac{\partial^2 X^l}{\partial x^j \partial x^k} + X^i \frac{\partial \Gamma_{ki}^l}{\partial x^j} + X^i R_{ijk}^l \right) \frac{\partial}{\partial x^l} \\ &= \left(\frac{\partial^2 X^l}{\partial x^j \partial x^k} + X^i \frac{\partial \Gamma_{jk}^l}{\partial x^i} \right) \frac{\partial}{\partial x^l} \end{aligned}$$

since

$$R_{ijk}^l = \frac{\partial \Gamma_{jk}^l}{\partial x^i} - \frac{\partial \Gamma_{ik}^l}{\partial x^j} + \Gamma_{jk}^s \Gamma_{is}^l - \Gamma_{ik}^s \Gamma_{js}^l$$

Now it suffices to show

$$\frac{\partial^2 X^l}{\partial x^j \partial x^k} + X^i \frac{\partial \Gamma_{jk}^l}{\partial x^i} \equiv 0$$

In order to show this, for arbitrary $p \in M$, consider a geodesic γ starting at p and consider Jacobi field $J(t) = X \circ \gamma(t)$. Direct computation shows

$$\begin{aligned}
 J'(t) &= \left(\frac{\partial X^i}{\partial x^k} \frac{d\gamma^k}{dt} + X^i \Gamma_{ki}^l \frac{d\gamma^k}{dt} \right) \frac{\partial}{\partial x^l} \Big|_{\gamma(t)} \\
 J''(0) &= \left(\frac{\partial^2 X^l}{\partial x^j \partial x^k} \frac{d\gamma^j}{dt} \frac{d\gamma^k}{dt} + X^i \frac{\partial \Gamma_{ki}^l}{\partial x^j} \frac{d\gamma^j}{dt} \frac{d\gamma^k}{dt} \right) \frac{\partial}{\partial x^l} \Big|_p \\
 &= \left(\frac{\partial^2 X^l}{\partial x^j \partial x^k} + X^i \frac{\partial \Gamma_{ki}^l}{\partial x^j} \right) \frac{d\gamma^j}{dt} \frac{d\gamma^k}{dt} \frac{\partial}{\partial x^l} \Big|_p \\
 &= \left(\frac{\partial^2 X^l}{\partial x^j \partial x^k} + X^i \frac{\partial \Gamma_{jk}^l}{\partial x^i} + X^i \frac{\partial \Gamma_{ki}^l}{\partial x^j} - X^i \frac{\partial \Gamma_{jk}^l}{\partial x^i} \right) \frac{d\gamma^j}{dt} \frac{d\gamma^k}{dt} \frac{\partial}{\partial x^l} \Big|_p \\
 &= \left(\frac{\partial^2 X^l}{\partial x^j \partial x^k} + X^i \frac{\partial \Gamma_{jk}^l}{\partial x^i} \right) \frac{d\gamma^j}{dt} \frac{d\gamma^k}{dt} \frac{\partial}{\partial x^l} \Big|_p - R(X, \gamma') \gamma'
 \end{aligned}$$

which implies

$$\frac{\partial^2 X^l}{\partial x^j \partial x^k} + X^i \frac{\partial \Gamma_{jk}^l}{\partial x^i} = 0$$

holds at point p . Since p is arbitrary, this completes the proof. \square

Lemma 18.4.1. Killing field on a complete Riemannian manifold (M, g) is complete.

Proof. Let X be a Killing field, we need to show the flow $\varphi_t : M \rightarrow M$ generated by X is defined for $t \in \mathbb{R}$. Otherwise, we assume φ_t is defined on (a, b) . Note that for each $p \in M$, curve $\varphi_t(p)$ is a curve defined on (a, b) having finite constant speed, since φ_t is isometry. Then we have $\varphi_t(p)$ can be extended to the one defined on \mathbb{R} , since M is complete. \square

Theorem 18.4.1. Let (M, g) be a complete Riemannian manifold and \mathfrak{g} the space of Killing fields, then \mathfrak{g} is isomorphic to the Lie algebra of $G = \text{Iso}(M, g)$.

Proof. Since $[\mathcal{L}_X, \mathcal{L}_Y] = \mathcal{L}_{[X, Y]}$, we know \mathfrak{g} is a Lie algebra. Now let's construct correspondence:

1. Given a Killing field X , by Lemma 18.4.1, one deduces that the flow $\varphi : \mathbb{R} \times M \rightarrow M$ generated by X is a one parameter subgroup $\gamma : \mathbb{R} \rightarrow G$, and $\gamma'(0) \in T_e G$;
2. Given $v \in T_e G$, consider the one-parameter subgroup $\gamma(t) = \exp(tv) : \mathbb{R} \rightarrow G$ which gives a flow by

$$\begin{aligned}
 \varphi : \mathbb{R} \times M &\rightarrow M \\
 (t, p) &\mapsto \exp(tv) \cdot p
 \end{aligned}$$

Then the vector field X generated by this flow is a Killing field.

This gives an one to one correspondence between Killing fields and Lie algebra of G , and it's a Lie algebra isomorphism in fact. \square

Theorem 18.4.2. Let (M, g) be a symmetric space and G the isometry group. For any $p \in M$, the Lie algebra of the isotropy subgroup G_p is isomorphic to

$$\mathfrak{l} = \{X \in \mathfrak{g} \mid X_p = 0\}$$

where \mathfrak{g} is the Lie algebra of G , that is the space of Killing fields.

Proof. Let $X \in \mathfrak{g}$ with $X_p = 0$, and $\varphi_t : M \rightarrow M$ the flow of X . It suffices to show $\varphi_t(p) = p$ for all $t \in \mathbb{R}$. We use $\gamma_p(t)$ to denote $\varphi_t(p)$, then for any smooth function $f : M \rightarrow \mathbb{R}$ and $s \in \mathbb{R}$, we have

$$\begin{aligned} \gamma'_p(s)f &= \left. \frac{d}{dt} \right|_{t=s} f \circ \gamma_p(t) \\ &= \left. \frac{d}{dt} \right|_{t=0} f \circ \gamma_p(t+s) \\ &= \left. \frac{d}{dt} \right|_{t=0} f \circ \varphi_s \circ \varphi_t(p) \\ &= \left. \frac{d}{dt} \right|_{t=0} (f \circ \varphi_s)(\gamma_p(t)) \\ &= \gamma'_p(0)(f \circ \varphi_s) \\ &= X_p(f \circ \varphi_s) \\ &= 0 \end{aligned}$$

Hence $\gamma'_p(s) = 0$ for all $s \in \mathbb{R}$, thus $\gamma_p(s)$ is constant, which implies $\gamma_p(s) = \gamma_p(0) = p$. \square

Proposition 18.4.1. Let (M, g) be a complete Riemannian manifold and $p \in M$. Then a Killing field is determined by the values X_p and $(\nabla X)_p$. In particular,

$$\dim \mathfrak{g} \leq \frac{n(n+1)}{2}$$

Proof. It suffices to show if $X_p = 0$ and $(\nabla X)_p = 0$, then $X \equiv 0$. For arbitrary $q \in M$, let $\gamma : [0, 1] \rightarrow M$ be a geodesic connecting p and q with $\gamma'(0) = v$. Since $J(t) = X \circ \gamma(t)$ is a Jacobi field, a direct computation shows

$$\nabla_v X(p) = J'(0)$$

thus $J(t) \equiv 0$, since Jacobi field is determined by two initial values. In particular, $X(q) = J(1) = 0$. Since ∇X is skew-symmetric, then

$$\dim \mathfrak{g} \leq n + \frac{n(n-1)}{2} = \frac{n(n+1)}{2}$$

\square

Definition 18.4.1 (transvection). Let (M, g) be a Riemannian symmetric space and γ a geodesic. The transvection along γ is defined as

$$T_t = s_{\gamma(\frac{t}{2})} \circ s_{\gamma(0)}$$

where s_p is the symmetry at point p .

Proposition 18.4.2. Let (M, g) be a Riemannian symmetric space and γ a geodesic. Then

1. For any $a, t \in \mathbb{R}$, $s_{\gamma(a)}(\gamma(t)) = \gamma(2a - t)$;
2. T_t translates the geodesic, that is $T_t(\gamma(s)) = \gamma(t + s)$;
3. $(dT_t)_{\gamma(s)} : T_{\gamma(s)}M \rightarrow T_{\gamma(t+s)}M$ is the parallel transport $P_{s,t+s}^\gamma$;
4. T_t is one-parameter subgroup of $\text{Iso}(M, g)$.

Proof. For (1). It follows from the uniqueness of geodesics with given initial value.

For (2). Note that

$$\begin{aligned} T_t(\gamma(s)) &= s_{\gamma(\frac{t}{2})} \circ s_{\gamma(0)}(\gamma(s)) \\ &= s_{\gamma(\frac{t}{2})}(\gamma(-s)) \\ &= \gamma(t + s) \end{aligned}$$

For (3). S

For (4). In order to show $T_{t+s} = T_t \circ T_s$, it suffices to check they're same at some point, so do their derivatives, since isometry can be determined by these two values. Note that

$$\begin{aligned} T_{t+s}(\gamma(0)) &= \gamma(t + s) \\ &= T_t \circ T_s(\gamma(0)) \\ (dT_{t+s})_{\gamma(0)} &= P_{0,t+s}^\gamma \\ &= P_{s,t+s}^\gamma \circ P_{0,s}^\gamma \\ &= (dT_t)_{\gamma(s)} \circ (dT_s)_{\gamma(0)} \\ &= (d(T_t \circ T_s))_{\gamma(0)} \end{aligned}$$

This completes the proof. \square

Definition 18.4.2 (infinitesimal transvection). Let (M, g) be a Riemannian symmetric space. For any point $p \in M$ and any $v \in T_p M$, there is a geodesic $\gamma_v : \mathbb{R} \rightarrow M$ with $\gamma_v(0) = p$ and $\gamma'_v(0) = v$. The infinitesimal generator X of transvections T_t along γ is given by

$$X_q = \left. \frac{d}{dt} \right|_{t=0} T_t(q).$$

This Killing field X is called an infinitesimal transvection.

Part 7. Appendix

APPENDIX A. HODGE THEOREM

In this section, we mainly follow the Chapter 6 of [War10].

A.1. Introduction and proof of Hodge theorem. We shall use Δ^* to denote the adjoint of Laplace-Beltrami operator on Ω_M^k . This operator is precisely Δ itself, since Laplace-Beltrami operator is self-adjoint, and we usually make no distinction between Δ and Δ^* . However, this distinction will be important for the form of the following definition.

An important question is to find a necessary and sufficient condition for there to exist a solution ω of equation $\Delta\omega = \alpha$, where α is a given k -form. Suppose ω is a solution, then

$$(\Delta\omega, \varphi) = (\alpha, \varphi)$$

holds for all k -forms φ . Equivalently we have

$$(\omega, \Delta^*\varphi) = (\alpha, \varphi)$$

holds for all k -forms φ . In this viewpoint, we can regard a solution of $\Delta\omega = \alpha$ as a certain type of linear functional on $C^\infty(M, \Omega_M^k)$, namely solution ω determines a bounded linear functional l on $C^\infty(M, \Omega_M^k)$ by

$$l(\varphi) = (\omega, \varphi), \quad \varphi \in C^\infty(M, \Omega_M^k)$$

such that

$$l(\Delta^*\varphi) = (\alpha, \varphi)$$

holds for all k -forms φ .

Definition A.1.1 (weak solution). A linear functional l on $C^\infty(M, \Omega_M^k)$ is called a weak solution of $\Delta\omega = \alpha$, if

$$l(\Delta^*\varphi) = (\alpha, \varphi)$$

holds for all k -forms φ .

We have seen that each ordinary solution of $\Delta\omega = \alpha$ determines a weak solution of it, it turns out that the major effort of this section will be to prove a regularity theorem which says that the converse of this is true, that is each weak solution determines an ordinary solution. The main step is to show if l is a weak solution of $\Delta\omega = \alpha$, then there exists a smooth form ω such that

$$l(\varphi) = (\alpha, \varphi), \quad \varphi \in C^\infty(M, \Omega_M^k)$$

Then ω is an ordinary solution follows from

$$(\Delta, \varphi) = (\omega, \Delta^*\varphi) = l(\Delta^*\varphi) = (\alpha, \varphi)$$

holds for all k -forms φ , which implies $\Delta\omega = \alpha$.

The key theorems we will prove are listed as follows:

Theorem A.1.1 (regularity theorem). Let $\alpha \in C^\infty(M, \Omega_M^k)$, and l be a weak solution of $\Delta\omega = \alpha$, then there exists $\omega \in C^\infty(M, \Omega_M^k)$ such that

$$l(\varphi) = (\omega, \varphi)$$

holds for every k -forms φ . In particular, $\Delta\omega = \alpha$.

Theorem A.1.2. Let $\{\alpha_n\}$ be a sequence of smooth k -forms on M such that $\|\alpha_n\| \leq c$ and $\|\Delta\alpha_n\| \leq c$ for all n and for some constant $c > 0$. Then a subsequence of $\{\alpha_n\}$ is a Cauchy sequence in $C^\infty(M, \Omega_M^k)$.

Corollary A.1.1. There exists a constant $c > 0$ such that

$$\|\psi\| \leq c\|\Delta\psi\|$$

holds for all $\psi \in (\mathcal{H}^k)^\perp$

Proof. Suppose the contrary, then there exists a sequence $\psi_j \in (\mathcal{H}^k)^\perp$ with $\|\psi_j\| = 1$ and $\|\Delta\psi_j\| \rightarrow 0$. By Theorem A.1.2, there exists a subsequence of $\{\psi_j\}$ which for convenience we can assume to be $\{\psi_j\}$ itself, is Cauchy. Thus for each $\varphi \in C^\infty(M, \Omega_M^k)$, $\lim_{j \rightarrow \infty} (\psi_j, \varphi)$ exists. Consider the linear functional l on $C^\infty(M, \Omega_M^k)$ defined by

$$l(\varphi) := \lim_{j \rightarrow \infty} (\psi_j, \varphi), \quad \varphi \in C^\infty(M, \Omega_M^k)$$

It's clear l is bounded, and

$$l(\Delta\varphi) = \lim_{j \rightarrow \infty} (\psi, \Delta\varphi) = \lim_{j \rightarrow \infty} (\Delta\psi_j, \varphi) = 0$$

holds for all $\varphi \in C^\infty(M, \Omega_M^k)$, which implies l is a weak solution of $\Delta\psi = 0$. By Theorem A.1.1, there exists a k -form ψ such that $l(\varphi) = (\psi, \varphi)$, where $\varphi \in C^\infty(M, \Omega_M^k)$. Consequently $\psi_j \rightarrow \psi$, and $\psi \in (\mathcal{H}^k)^\perp$ with $\|\psi\| = 1$. However, Theorem A.1.1 implies $\psi \in \mathcal{H}^k$, a contradiction. \square

Holding above results, we can prove Hodge theorem.

Theorem A.1.3 (Hodge theorem). Consider the Laplace operator $\Delta : C^\infty(M, \Omega_M^k) \rightarrow C^\infty(M, \Omega_M^k)$, then

1. $\dim_{\mathbb{R}} \mathcal{H}^k < \infty$;
2. There is an orthogonal direct sum decomposition

$$C^\infty(M, \Omega_M^k) = \mathcal{H}^k \oplus \text{im } \Delta$$

Proof. For (1). If \mathcal{H}^k is not finite dimensional, then there exists an infinite orthonormal sequence. By Theorem A.1.2, this orthonormal sequence contains a Cauchy sequence, which is impossible. Thus \mathcal{H}^k is finite dimensional.

For (2). Note that we naturally have the following orthogonal decomposition

$$C^\infty(M, \Omega_M^k) = (\mathcal{H}^k)^\perp \oplus \mathcal{H}^k$$

The theorem will be proved by showing that $(\mathcal{H}^k)^\perp = \text{im } \Delta$. We use \mathcal{H} to denote the projection from $C^\infty(M, \Omega_M^k)$ to \mathcal{H}^k , that is $\mathcal{H}(\alpha)$ is the harmonic part of α .

It's easy to see $\text{im } \Delta \subset (\mathcal{H}^k)^\perp$, since for all $\omega \in C^\infty(M, \Omega_M^k)$ and $\alpha \in \mathcal{H}^k$, we have

$$(\Delta\omega, \alpha) = (\omega, \Delta\alpha) = 0$$

To see converse, for $\alpha \in (\mathcal{H}^k)^\perp$, we define a linear functional l on $\text{im } \Delta$ by setting

$$l(\Delta\varphi) := (\alpha, \varphi)$$

for all $\varphi \in C^\infty(M, \Omega_M^k)$.

1. l is well-defined, since if $\Delta\varphi_1 = \Delta\varphi_2$, then $\varphi_1 - \varphi_2 \in \mathcal{H}^k$, then $(\alpha, \varphi_1 - \varphi_2) = 0$;
2. l is bounded. Indeed, for $\varphi \in C^\infty(M, \Omega_M^k)$, let $\psi = \varphi - \mathcal{H}(\varphi)$. Then

$$\begin{aligned} |l(\Delta\varphi)| &= |l(\Delta\psi)| \\ &= |(\alpha, \psi)| \\ &\leq \|\alpha\| \|\psi\| \\ &\stackrel{*}{\leq} c \|\alpha\| \|\Delta\psi\| \\ &= c \|\alpha\| \|\Delta\varphi\| \end{aligned}$$

where $*$ holds from Corollary A.1.1.

By Hahn-Banach theorem, l extends to a bounded linear functional on $C^\infty(M, \Omega_M^k)$, thus l is a weak solution of $\Delta\omega = \alpha$. By Theorem A.1.1, there exists a k -form ω such that $\Delta\omega = \alpha$. Hence

$$(\mathcal{H}^k)^\perp = \text{im } \Delta$$

This completes the proof of Hodge theorem. □

APPENDIX B. SECOND FUNDAMENTAL FORM

B.1. Pullback connection. In this section, we fix the following notations:

1. Vector bundle E equipped with a metric g over a smooth manifold M , endowed with Levi-Civita connection ∇^E ;
2. $f : M \rightarrow N$ is a smooth map;
3. $\{dx^i\}$ is used to denote a local basis of TM , $\{dy^m\}$ is used to denote a local basis of TN and $\{e_\alpha\}$ is used to denote a local basis of E .

Definition B.1.1 (pullback vector bundle). The pullback vector bundle f^*E over M is defined by the set

$$\widehat{E} = f^*E := \{(p, v) \in M \times E \mid f(p) = \pi(v)\}$$

endowed with subspace topology.

Remark B.1.1 (local form). A local basis of \widehat{E} can be written as

$$\widehat{e}_\alpha(x) := f^*e_\alpha(x) = e_\alpha(f(x))$$

where $x \in M$.

Definition B.1.2 (pullback connection). The pullback connection $\widehat{\nabla}$ over $\widehat{E} \rightarrow M$ is defined as:

$$\begin{aligned} \widehat{\nabla} : C^\infty(M, \widehat{E}) &\rightarrow C^\infty(M, T^*M \otimes \widehat{E}) \\ f^*(s) &\mapsto f^*(\nabla s) \end{aligned}$$

where $s \in C^\infty(M, E)$.

Remark B.1.2 (local form). If we take a local basis $\{\widehat{e}_\alpha\}$ of \widehat{E} , then

$$\begin{aligned} \widehat{\nabla} \widehat{e}_\alpha &= f^*(\nabla e_\alpha) \\ &= f^*(\Gamma_{m\alpha}^\beta dy^m \otimes e_\beta) \\ &= \Gamma_{m\alpha}^\beta(f) \frac{\partial f^m}{\partial x^i} dx^i \otimes \widehat{e}_\beta \end{aligned}$$

Note that we can also use f to pullback metric g on E to obtain a metric on \widehat{E} , denoted by \widehat{g} . Locally we can write

$$\begin{aligned} \widehat{g}_{\alpha\beta} \widehat{e}^\alpha \otimes \widehat{e}^\beta &:= f^*(g_{\alpha\beta} e^\alpha \otimes e^\beta) \\ &= g_{\alpha\beta}(f) \widehat{e}^\alpha \otimes \widehat{e}^\beta \end{aligned}$$

that is $\widehat{g}_{\alpha\beta} = g_{\alpha\beta}(f)$.

Lemma B.1.1. The pullback connection $\widehat{\nabla}$ is compatible with \widehat{g} , that is for any vector field X of M and section s, t of \widehat{E} , we have

$$X\widehat{g}(s, t) = \widehat{g}(\widehat{\nabla}_X s, t) + \widehat{g}(s, \widehat{\nabla}_X t)$$

Proof. Locally we take $X = \frac{\partial}{\partial x^i}$, $s = \hat{e}_\alpha$, $t = \hat{e}_\beta$, then

$$\begin{aligned}\frac{\partial}{\partial x^i} \hat{g}_{\alpha\beta} &= \frac{\partial}{\partial x^i} g_{\alpha\beta}(f) \\ &= \frac{\partial f^m}{\partial x^i} \frac{\partial}{\partial y^m} g_{\alpha\beta}(f) \\ &= \frac{\partial f^m}{\partial x^i} (\Gamma_{m\alpha}^\gamma(f) g_{\gamma\beta}(f) + \Gamma_{m\beta}^\gamma(f) g_{\alpha\gamma}(f)) \\ \hat{g}(\hat{\nabla}_{\frac{\partial}{\partial x^i}} \hat{e}_\alpha, \hat{e}_\beta) &= \Gamma_{m\alpha}^\gamma(f) \frac{\partial f^m}{\partial x^i} g_{\gamma\beta}(f) \\ \hat{g}(\hat{e}_\alpha, \hat{\nabla}_{\frac{\partial}{\partial x^i}} \hat{e}_\beta) &= \Gamma_{m\beta}^\gamma(f) \frac{\partial f^m}{\partial x^i} g_{\alpha\gamma}(f)\end{aligned}$$

This completes the proof. \square

Definition B.1.3. The curvature tensor \hat{R} of pullback connection $\hat{\nabla}$ on vector bundle $\hat{E} \rightarrow M$ is given by

$$\hat{R}(X, Y, s, t) = \hat{g}(\hat{\nabla}_X \hat{\nabla}_Y s - \hat{\nabla}_Y \hat{\nabla}_X s, t)$$

Remark B.1.3 (local form).

$$\hat{R}_{ij\alpha\beta} = R_{mn\alpha\beta} \frac{\partial f^m}{\partial x^i} \frac{\partial f^n}{\partial x^j}$$

where $R_{mn\alpha\beta}$ is curvature of ∇^E .

Proof.

$$\begin{aligned}\hat{R}_{ij\alpha\beta} &= \hat{R}(\frac{\partial}{\partial x^i}, \frac{\partial}{\partial x^j}, \hat{e}_\alpha, \hat{e}_\beta) \\ &= \hat{g}(\hat{R}(\frac{\partial}{\partial x^i}, \frac{\partial}{\partial x^j}) \hat{e}_\alpha, \hat{e}_\beta) \\ &= \hat{g}(\hat{\nabla}_{\frac{\partial}{\partial x^i}} \hat{\nabla}_{\frac{\partial}{\partial x^j}} \hat{e}_\alpha - \hat{\nabla}_{\frac{\partial}{\partial x^j}} \hat{\nabla}_{\frac{\partial}{\partial x^i}} \hat{e}_\alpha, \hat{e}_\beta)\end{aligned}$$

So it suffices to compute

$$\begin{aligned}\hat{\nabla}_{\frac{\partial}{\partial x^i}} \hat{\nabla}_{\frac{\partial}{\partial x^j}} \hat{e}_\alpha &= \hat{\nabla}_{\frac{\partial}{\partial x^i}} (\Gamma_{m\alpha}^\gamma(f) \frac{\partial f^m}{\partial x^j} \hat{e}_\gamma) \\ &= \frac{\partial}{\partial x^i} (\Gamma_{m\alpha}^\gamma(f) \frac{\partial f^m}{\partial x^j}) \hat{e}_\gamma + \Gamma_{m\alpha}^\gamma(f) \frac{\partial f^m}{\partial x^j} \hat{\nabla}_{\frac{\partial}{\partial x^i}} \hat{e}_\gamma \\ &= (\frac{\partial \Gamma_{m\alpha}^\gamma}{\partial y^n} \frac{\partial f^n}{\partial x^i} \frac{\partial f^m}{\partial x^j} + \Gamma_{m\alpha}^\gamma(f) \frac{\partial^2 f^m}{\partial x^i \partial x^j}) \hat{e}_\gamma + \frac{\partial f^m}{\partial x^j} \frac{\partial f^n}{\partial x^i} \Gamma_{m\alpha}^\gamma \Gamma_{n\gamma}^\delta \hat{e}_\delta \\ &= \frac{\partial f^m}{\partial x^j} \frac{\partial f^n}{\partial x^i} (\frac{\partial \Gamma_{m\alpha}^\gamma}{\partial y^n} + \Gamma_{m\alpha}^\delta \Gamma_{n\delta}^\gamma) \hat{e}_\gamma + \Gamma_{m\alpha}^\gamma \frac{\partial^2 f^m}{\partial x^i \partial x^j} \hat{e}_\gamma\end{aligned}$$

Thus

$$\hat{\nabla}_{\frac{\partial}{\partial x^i}} \hat{\nabla}_{\frac{\partial}{\partial x^j}} \hat{e}_\alpha - \hat{\nabla}_{\frac{\partial}{\partial x^j}} \hat{\nabla}_{\frac{\partial}{\partial x^i}} \hat{e}_\alpha = \frac{\partial f^m}{\partial x^j} \frac{\partial f^n}{\partial x^i} R_{mn\alpha}^\gamma \hat{e}_\gamma$$

that is

$$\begin{aligned}\widehat{R}_{ij\alpha\beta} &= \widehat{g}\left(\frac{\partial f^m}{\partial x^j} \frac{\partial f^n}{\partial x^i} R_{mn\alpha}^\gamma \widehat{e}_\gamma, \widehat{e}_\beta\right) \\ &= \frac{\partial f^m}{\partial x^j} \frac{\partial f^n}{\partial x^i} R_{mn\alpha}^\gamma g_{\gamma\beta} \\ &= \frac{\partial f^m}{\partial x^j} \frac{\partial f^n}{\partial x^i} R_{mn\alpha\beta}\end{aligned}$$

□

B.2. Second fundamental form. In this section, we fix the following notations:

1. $f : (M, g^M, \nabla^M) \rightarrow (N, g^N, \nabla^N)$ is a smooth map between two Riemannian manifolds.
2. Γ_{ij}^k is used to denote Christoffel symbol of ∇^M and Γ_{mn}^l is used to denote Christoffel symbol of ∇^N .
3. $\widehat{\nabla}$ is the connection on f^*TN induced by ∇^N .

Definition B.2.1 (second fundamental form). The second fundamental form $B \in C^\infty(M, T^*M \otimes T^*M \otimes f^*TN)$ of f is defined as

$$B(X, Y) := \widehat{\nabla}_X f_* Y - f_*(\nabla_X^M Y) \in C^\infty(M, f^*TN)$$

where $X, Y \in C^\infty(M, TM)$.

Remark B.2.1 (local form). Suppose that $X = \frac{\partial}{\partial x^i}, Y = \frac{\partial}{\partial x^j}$, then one has

$$f_*(\nabla_{\frac{\partial}{\partial x^i}} \frac{\partial}{\partial x^j}) = \Gamma_{ij}^k f_* \frac{\partial}{\partial x^k} = \Gamma_{ij}^k \frac{\partial f^m}{\partial x^k} \frac{\partial}{\partial y^m}$$

And

$$\begin{aligned}\widehat{\nabla}_{\frac{\partial}{\partial x^i}} \left(\frac{\partial f^m}{\partial x^j} \frac{\partial}{\partial y^m} \right) &= \frac{\partial^2 f^m}{\partial x^i \partial x^j} \frac{\partial}{\partial y^m} + \frac{\partial f^m}{\partial x^j} \widehat{\nabla}_{\frac{\partial}{\partial x^i}} \frac{\partial}{\partial y^m} \\ &= \left(\frac{\partial^2 f^l}{\partial x^i \partial x^j} + \frac{\partial f^m}{\partial x^j} \frac{\partial f^n}{\partial x^i} \Gamma_{nm}^l \right) \frac{\partial}{\partial y^l}\end{aligned}$$

Therefore

$$\begin{aligned}B_{ij} &:= B\left(\frac{\partial}{\partial x^i}, \frac{\partial}{\partial x^j}\right) \\ &= \left(\frac{\partial^2 f^l}{\partial x^i \partial x^j} + \frac{\partial f^m}{\partial x^j} \frac{\partial f^n}{\partial x^i} \Gamma_{mn}^l - \Gamma_{ij}^k \frac{\partial f^l}{\partial x^k} \right) \frac{\partial}{\partial y^l}\end{aligned}$$

Remark B.2.2 (Geodesic). Consider a smooth regular curve $\gamma : [a, b] \rightarrow M$, we can regard it as $\gamma : [a, b], g, \nabla \rightarrow (M, g^M, \nabla^M)$. Thus our second fundamental form in this case is

$$B_{ij} = \left(\frac{d^2 \gamma^k}{dt^2} + \frac{d\gamma^i}{dt} \frac{d\gamma^j}{dt} \Gamma_{ij}^k \right) \frac{\partial}{\partial x^k}$$

So condition for geodesic is exactly second fundamental form is zero.

Remark B.2.3 (Hessian). Consider smooth function f , we can regard it as $f : (M, g^M, \nabla^M) \rightarrow (\mathbb{R}, g, \nabla)$, where metric and connection on \mathbb{R} are trivial. Thus our second fundamental form in this case is

$$B_{ij} = \left(\frac{\partial^2 f}{\partial x^i \partial x^j} - \Gamma_{ij}^k \frac{\partial f}{\partial x^k} \right) \frac{\partial}{\partial y}$$

since $\Gamma_{mn}^l = 0$. That's exactly our Hess f , so second fundamental form generalizes our Hessian of smooth function;

Since Hessian of a smooth function is exactly $\nabla(\nabla f)$, where $\nabla f \in C^\infty(M, T^*M)$.

In fact we can express our second fundamental form B as $\tilde{\nabla} df$, where $df \in C^\infty(M, T^*M \otimes f^*TN)$ and $\tilde{\nabla}$ is the connection on $T^*M \otimes f^*TN$ induced by ∇^M and pullback connection on f^*TN . Indeed, note that locally we have

$$df = \frac{\partial f^m}{\partial x^i} dx^i \otimes \frac{\partial}{\partial y^m}$$

Then

$$\begin{aligned} \tilde{\nabla} df &= \tilde{\nabla} \left(\frac{\partial f^m}{\partial x^i} dx^i \otimes \frac{\partial}{\partial y^m} \right) \\ &= \frac{\partial^2 f^m}{\partial x^j \partial x^i} dx^j \otimes dx^i \otimes \frac{\partial}{\partial y^m} - \frac{\partial f^m}{\partial x^i} \Gamma_{jk}^i dx^j \otimes dx^k \otimes \frac{\partial}{\partial y^m} + \frac{\partial f^m}{\partial x^i} \frac{\partial f^n}{\partial x^j} \Gamma_{mn}^l dx^i \otimes dx^j \otimes \frac{\partial}{\partial y^l} \\ &= \left(\frac{\partial^2 f^l}{\partial x^i \partial x^j} - \frac{\partial f^l}{\partial x^k} \Gamma_{ij}^k + \frac{\partial f^m}{\partial x^i} \frac{\partial f^n}{\partial x^j} \Gamma_{mn}^l \right) dx^i \otimes dx^j \otimes \frac{\partial}{\partial y^l} \\ &= B \end{aligned}$$

as desired.

APPENDIX C. HARMONIC MAP AND ITS VARIATION

In this section we fix a smooth map $f : (M, g) \rightarrow (N, h)$ between Riemannian manifolds with second fundamental form B . Keep in mind we regard df as a section of $T^*M \otimes f^*TN$ and B as a section of $T^*M \otimes T^*M \otimes f^*TN$.

C.1. Harmonic map and totally geodesic.

Definition C.1.1 ((scalar) Laplacian). The (scalar) Laplacian of f is defined as

$$\Delta_g f := \text{tr}_g B \in C^\infty(M, f^*TN)$$

Definition C.1.2 (harmonic map). f is called harmonic map if its scalar Laplacian $\Delta_g f = 0$.

Definition C.1.3 (totally geodesic). f is called totally geodesic, if its second fundamental form $B = 0$.

Example C.1.1. For a geodesic $\gamma : [a, b] \rightarrow (M, g)$, if we endow $[a, b]$ with canonical metric, then γ is totally geodesic, thus it's harmonic.

Remark C.1.1. If γ is regular, that is $\gamma'(t) \neq 0$ for each $t \in [a, b]$, or in other words γ is immersion, there is an induced metric g_0 on $[a, b]$ given by

$$g_0 = g_{ij} \frac{d\gamma^i}{dt} \frac{d\gamma^j}{dt} dt \otimes dt$$

Furthermore, if γ is unit-speed, then

$$1 = |\gamma'(t)|^2 = g_{ij} \frac{d\gamma^i}{dt} \frac{d\gamma^j}{dt}$$

which implies g_0 is canonical metric on $[a, b]$.

Example C.1.2. For a smooth function $f : (M, g) \rightarrow \mathbb{R}$, if we endow \mathbb{R} with canonical metric, then f is a harmonic map if and only if it's a harmonic function.

Lemma C.1.1. Let $\gamma : [a, b] \rightarrow M$ be a smooth curve and $\tilde{\gamma} = f \circ \gamma$. If we use $\hat{\nabla}$ and $\tilde{\nabla}$ to denote the induced connection on γ^*TM and $\tilde{\gamma}^*TN$ respectively, then

$$\tilde{\nabla}_{\frac{d}{dt}} \tilde{\gamma}_* \left(\frac{d}{dt} \right) = f_* \left(\hat{\nabla}_{\frac{d}{dt}} \gamma_* \left(\frac{d}{dt} \right) \right) + \gamma^* B$$

Proof. Directly compute

$$\begin{aligned} \tilde{\nabla}_{\frac{d}{dt}} \tilde{\gamma}_* \left(\frac{d}{dt} \right) &= \left(\frac{d^2 \tilde{\gamma}^l}{dt^2} + \Gamma_{mn}^l(\tilde{\gamma}) \frac{d\tilde{\gamma}^m}{dt} \frac{d\tilde{\gamma}^n}{dt} \right) \frac{\partial}{\partial y^l} \\ &= \left\{ \frac{\partial f^l}{\partial x^k} \frac{d^2 \gamma^k}{dt^2} + \left(\frac{\partial^2 f^l}{\partial x^i \partial x^j} + \Gamma_{mn}^l \frac{\partial f^m}{\partial x^i} \frac{\partial f^n}{\partial x^j} \right) \frac{\partial \gamma^i}{dt} \frac{\partial \gamma^j}{dt} \right\} \frac{\partial}{\partial y^l} \\ &= \left\{ \frac{\partial f^l}{\partial x^k} \left(\frac{d^2 \gamma^k}{dt^2} + \Gamma_{ij}^k \frac{d\gamma^i}{dt} \frac{d\gamma^j}{dt} \right) + \left(\frac{\partial^2 f^l}{\partial x^i \partial x^j} + \frac{\partial f^m}{\partial x^i} \frac{\partial f^n}{\partial x^j} \Gamma_{mn}^l - \Gamma_{ij}^k \frac{\partial f^l}{\partial x^k} \right) \frac{d\gamma^i}{dt} \frac{d\gamma^j}{dt} \right\} \frac{\partial}{\partial y^l} \\ &= f_* \left(\hat{\nabla}_{\frac{d}{dt}} \gamma_* \left(\frac{d}{dt} \right) \right) + \gamma^* B \end{aligned}$$

□

Theorem C.1.1. The followings are equivalent:

1. f is totally geodesic;
2. f maps geodesics in M to geodesics in N .

Proof. Clear from above lemma. □

C.2. First variation of smooth map.

Definition C.2.1 (energy functional). The energy density of smooth function $f : (M, g) \rightarrow (N, h)$ is

$$e(f) = |\mathrm{d}f|^2$$

The energy of f is

$$E(f) = \frac{1}{2} \int_M e(f) \mathrm{vol}$$

Remark C.2.1 (local form). We can locally write energy density as

$$\begin{aligned} \langle \mathrm{d}f, \mathrm{d}f \rangle &= \left\langle \frac{\partial f^m}{\partial x^i} \mathrm{d}x^i \otimes \frac{\partial}{\partial y^m}, \frac{\partial f^n}{\partial x^j} \mathrm{d}x^j \otimes \frac{\partial}{\partial y^n} \right\rangle \\ &= g^{ij} h_{mn}(f) \frac{\partial f^m}{\partial x^i} \frac{\partial f^n}{\partial x^j} \end{aligned}$$

Theorem C.2.1. The Euler-Lagrange equation of $E(f)$ is

$$\widehat{\nabla}^* \mathrm{d}f = 0$$

where $\widehat{\nabla}^*$ is the formal adjoint operator of $\widehat{\nabla}$.

Proof. We fix the following notations in the proof:

1. Consider a smooth variation $F : M \times \mathbb{R} \rightarrow N$ of f , we also write $f_t(-) = F(-, t)$ for convenience;
2. Set $\overline{M} = M \times \mathbb{R}$ and there is a natural metric $\overline{g} = g \times g_{\mathbb{R}}$ on \overline{M} ;
3. The pullback F^*TN bundle is denoted by W , and induced connection on W is denoted by ∇^W ;
4. Fix $t \in \mathbb{R}$, $f_t : M \rightarrow N$, then $\mathrm{d}f_t$ is a section of $T^*M \otimes f_t^*TN$, and we can regard it as a section of $T^*\overline{M} \otimes W$.

Holding above notations, we have

$$\begin{aligned} \frac{\mathrm{d}}{\mathrm{d}t} E(f_t) &= \frac{1}{2} \frac{\mathrm{d}}{\mathrm{d}t} \int_M |\mathrm{d}f_t|^2 \mathrm{vol} \\ &= \int_M \langle \nabla_{\frac{\partial}{\partial t}}^{T^*\overline{M} \otimes W} \mathrm{d}f_t, \mathrm{d}f_t \rangle \mathrm{vol} \end{aligned}$$

Here we claim

$$\langle \nabla_{\frac{\partial}{\partial t}}^{T^*\overline{M} \otimes W} \mathrm{d}f_t, \mathrm{d}f_t \rangle \stackrel{1}{=} \langle \nabla_{\frac{\partial}{\partial t}}^{T^*\overline{M} \otimes W} \mathrm{d}F, \mathrm{d}f_t \rangle \stackrel{2}{=} \langle \nabla^W F_* \left(\frac{\partial}{\partial t} \right), \mathrm{d}f_t \rangle$$

1. For equation marked 1: Note that

$$\begin{aligned} dF - df_t &= \frac{\partial F^m}{\partial x^i} dx^i \otimes \frac{\partial}{\partial y^m} + \frac{\partial F^m}{\partial t} dt \otimes \frac{\partial}{\partial y^m} - \frac{\partial f_t^m}{\partial x^i} dx^i \otimes \frac{\partial}{\partial y^m} \\ &= \frac{\partial F^m}{\partial t} dt \otimes \frac{\partial}{\partial y^m} \end{aligned}$$

since $\frac{\partial F^m}{\partial x^i} = \frac{\partial f_t^m}{\partial x^i}$. So we have

$$\begin{aligned} \nabla^{T^*\overline{M} \otimes W}(dF - df_t) &= \frac{\partial^2 F^l}{\partial t^2} dt \otimes dt \otimes \frac{\partial}{\partial y^l} + \frac{\partial F^m}{\partial t} dt \otimes \left(\frac{\partial F^n}{\partial t} \Gamma_{mn}^l dt \otimes \frac{\partial}{\partial y^l} + \frac{\partial F^n}{\partial x^i} \Gamma_{mn}^l dx^i \otimes \frac{\partial}{\partial y^l} \right) \\ &= \left(\frac{\partial^2 F^l}{\partial t^2} + \frac{\partial F^m}{\partial t} \frac{\partial F^n}{\partial t} \Gamma_{mn}^l \right) dt \otimes dt \otimes \frac{\partial}{\partial y^l} + \frac{\partial F^m}{\partial t} \frac{\partial F^n}{\partial x^i} \Gamma_{mn}^l dx^i \otimes dt \otimes \frac{\partial}{\partial y^l} \end{aligned}$$

Thus we have

$$\nabla_{\frac{\partial}{\partial t}}^{T^*\overline{M} \otimes W}(dF - df_t) = \left(\frac{\partial^2 F^l}{\partial t^2} + \frac{\partial F^m}{\partial t} \frac{\partial F^n}{\partial t} \Gamma_{mn}^l \right) dt \otimes \frac{\partial}{\partial y^l}$$

From above expression it's clear

$$\langle \nabla_{\frac{\partial}{\partial t}}^{T^*\overline{M} \otimes W}(dF - df_t), df_t \rangle = 0$$

since there is no dt in df_t , which implies equation marked 1 holds.

2. For equation marked 2: For arbitrary $X \in C^\infty(M, TM) \subset C^\infty(\overline{M}, T^*\overline{M})$, since second fundamental form is symmetric, thus

$$\begin{aligned} (\nabla_{\frac{\partial}{\partial t}}^{T^*\overline{M} \otimes W} dF)(X) &= (\nabla_X^{T^*\overline{M} \otimes W} dF)\left(\frac{\partial}{\partial t}\right) \\ &= \nabla_X^W F_*\left(\frac{\partial}{\partial t}\right) - F_*\left(\nabla_X^{\overline{M}} \frac{\partial t}{\partial t}\right) \\ &= \nabla_X^W F_*\left(\frac{\partial}{\partial t}\right) \end{aligned}$$

Now let v be an arbitrary variation vector field, that is

$$v = F_*\left(\frac{\partial}{\partial t}\right)\Big|_{t=0} \in C^\infty(M, f^*TN)$$

Hence when $t = 0$ we have

$$\left(\nabla^W F_*\left(\frac{\partial}{\partial t}\right)\right)\Big|_{t=0} = \widehat{\nabla} v$$

where $\widehat{\nabla}$ is the induced connection on f^*TN . So we have first variation formula

$$\begin{aligned} \frac{d}{dt}\Big|_{t=0} E(f_t) &= \int_M \langle \widehat{\nabla} v, df \rangle \text{vol} \\ &= \int_M \langle v, \widehat{\nabla}^* df \rangle \text{vol} = 0 \end{aligned}$$

where $\widehat{\nabla}^*$ is the formal adjoint operator of $\widehat{\nabla}$. since v is arbitrary, we deduce $\widehat{\nabla}^* df = 0$. \square

C.3. Second variation formula of harmonic map. Consider the following variation map of f

$$F : M \times (-\varepsilon_1, \varepsilon_1) \times (-\varepsilon_2, \varepsilon_2) \rightarrow N$$

with variation fields

$$\begin{aligned} v &= F_* \left(\frac{\partial}{\partial t} \right) \Big|_{s=t=0} \in C^\infty(M, f^*TN) \\ w &= F_* \left(\frac{\partial}{\partial s} \right) \Big|_{s=t=0} \in C^\infty(M, f^*TN) \end{aligned}$$

For convenience we denote $F(-, s, t) = f_{s,t}(-)$.

Theorem C.3.1 (second variation formula). If $f : (M, g) \rightarrow (N, h)$ is a harmonic map, then the second variation of the harmonic map f along v, w is

$$\frac{\partial^2}{\partial s \partial t} \Big|_{s=t=0} E(f_{s,t}) = \int_M \langle \widehat{\nabla} v, \widehat{\nabla} w \rangle \text{vol} - \int_M g^{ij} R_{pmnq} v^p w^q \frac{\partial f^m}{\partial x^i} \frac{\partial f^n}{\partial x^j} \text{vol}$$

Proof. In this proof, we still use the notations in proof of first variation formula. By first variation formula, we have

$$\frac{\partial}{\partial t} E(f_{s,t}) = \int_M \langle \nabla^W F_* \left(\frac{\partial}{\partial t} \right), df_{s,t} \rangle \text{vol}$$

So

$$\begin{aligned} \frac{\partial^2}{\partial s \partial t} E(f_{s,t}) &= \underbrace{\int_M \langle \nabla^{T^* \overline{M} \otimes W} \nabla^W F_* \left(\frac{\partial}{\partial t} \right), df_{s,t} \rangle}_{\text{part I}} \text{vol} \\ &\quad + \underbrace{\int_M \langle \nabla^W F_* \left(\frac{\partial}{\partial t} \right), \nabla^{T^* \overline{M} \otimes W} df_{s,t} \rangle}_{\text{part II}} \text{vol} \end{aligned}$$

Note that

$$\begin{aligned} \nabla_{\frac{\partial}{\partial s}}^{T^* \overline{M} \otimes W} df_{s,t} &= \nabla_{\frac{\partial}{\partial s}}^{T^* \overline{M} \otimes W} \left(\frac{\partial F^m}{\partial x^i} dx^i \otimes \frac{\partial}{\partial y^m} \right) \\ &= \frac{\partial^2 F^m}{\partial s \partial x^i} dx^i \otimes \frac{\partial}{\partial y^m} + \frac{\partial F^m}{\partial x^i} \frac{\partial F^n}{\partial s} \Gamma_{mn}^l dx^i \otimes \frac{\partial}{\partial y^l} \\ &= \left(\frac{\partial^2 F^l}{\partial s \partial x^i} + \frac{\partial F^m}{\partial x^i} \frac{\partial F^n}{\partial s} \Gamma_{mn}^l \right) dx^i \otimes \frac{\partial}{\partial y^l} \\ \widehat{\nabla} w &= \widehat{\nabla} \frac{\partial}{\partial x^i} \left(\frac{\partial F^n}{\partial s} \Big|_{t=s=0} \right) dx^i \otimes \frac{\partial}{\partial y^n} + \frac{\partial F^m}{\partial s} \frac{\partial F^n}{\partial x^i} \Big|_{t=s=0} \Gamma_{mn}^l dx^i \otimes \frac{\partial}{\partial y^l} \\ &= \left(\frac{\partial^2 F^l}{\partial s \partial x^i} + \frac{\partial F^m}{\partial x^i} \frac{\partial F^n}{\partial s} \Big|_{t=s=0} \Gamma_{mn}^l \right) dx^i \otimes \frac{\partial}{\partial y^l} \end{aligned}$$

which implies setting $t = s = 0$ we have part II is

$$\int_M \langle \widehat{\nabla} v, \widehat{\nabla} w \rangle \text{vol}$$

For part I, take arbitrary $X \in C^\infty(M, TM) \subset C^\infty(\overline{M}, T^*\overline{M})$, we have
Hence we obtain

$$\nabla_{\frac{\partial}{\partial s}}^{T^*\overline{M} \otimes W} \nabla^W F_* \left(\frac{\partial}{\partial t} \right) (X) = (\nabla^{T^*\overline{M} \otimes W} \nabla^W F_* \left(\frac{\partial}{\partial t} \right) (X)) \left(\frac{\partial}{\partial s}, X \right)$$

Setting $t = s = 0$ we have

Hence

$$\begin{aligned} \frac{\partial^2}{\partial s \partial t} \Big|_{t=s=0} E(f_{s,t}) &= \int_M \langle \widehat{\nabla}(\nabla_{\frac{\partial}{\partial s}}^W F_* \left(\frac{\partial}{\partial t} \right) \Big|_{s=t=0}), df \rangle \text{vol} \\ &\quad + \int_M g^{ij} R_{pmqn} v^p w^q \frac{\partial f^m}{\partial x^i} \frac{\partial f^n}{\partial x^j} \text{vol} + \int_M \langle \widehat{\nabla} w, \widehat{\nabla} v \rangle \text{vol} \end{aligned}$$

If f is harmonic, that is $\widehat{\nabla}^* df = 0$, we obtain the desired formula. \square

C.4. Bochner formula for harmonic map. Recall that for a smooth function $f : (M, g) \rightarrow \mathbb{R}$,

$$\frac{1}{2} \Delta |df|^2 = |\text{Hess } f|^2 + \text{Ric}(\nabla f, \nabla f)$$

In this section we generalize this formula to smooth map $f : (M, g) \rightarrow (N, h)$ between Riemannian manifolds, to get similar Bochner's theorem we have proven before.

Theorem C.4.1. Let $f : (M, g) \rightarrow (N, h)$ be a smooth map between Riemannian manifolds, then

$$\frac{1}{2} \Delta_g |df|^2 = |\widehat{\nabla} df|^2 + \langle \widehat{\nabla}(df), df \rangle + g^{ik} g^{jl} R_{ij} \frac{\partial f^m}{\partial x^k} \frac{\partial f^n}{\partial x^l} h_{mn} - g^{kl} g^{ij} R_{mnpq} \frac{\partial f^m}{\partial x^i} \frac{\partial f^n}{\partial x^j} \frac{\partial f^p}{\partial x^k} \frac{\partial f^q}{\partial x^l}$$

Theorem C.4.2. Let $f : (M, g) \rightarrow (N, h)$ be a harmonic map between Riemannian manifolds. If

1. M is connected compact with positive Ricci curvature;
2. N has non-positive sectional curvature.

Then f is constant.

Proof. Suppose $|df|^2$ attains its maximum at some point $p \in M$, we have

$$\Delta_g |df|^2(p) \leq 0$$

Thus

$$\frac{1}{2} \Delta_g |df|^2 \geq g^{ik} g^{jl} R_{ij} \frac{\partial f^m}{\partial x^k} \frac{\partial f^n}{\partial x^l} h_{mn} - g^{kl} g^{ij} R_{mnpq} \frac{\partial f^m}{\partial x^i} \frac{\partial f^n}{\partial x^j} \frac{\partial f^p}{\partial x^k} \frac{\partial f^q}{\partial x^l}$$

WLOG we may assume $g_{ij}(p) = \delta_{ij}$, $h_{\alpha\beta}(f(p)) = \delta_{mn}$ by choosing normal coordinate. Then

$$\begin{aligned} \frac{1}{2} \Delta_g |df|^2 &\geq \sum_{i,j,m} R_{ij} \frac{\partial f^m}{\partial x^i} \frac{\partial f^m}{\partial x^j} - \sum_{i,j} R_{mnpq} \frac{\partial f^m}{\partial x^i} \frac{\partial f^n}{\partial x^i} \frac{\partial f^p}{\partial x^j} \frac{\partial f^q}{\partial x^j} \\ &\geq 0 \end{aligned}$$

since R_{ij} is positive, which implies $df \equiv 0$, thus f is constant since M is connected. \square

Corollary C.4.1. If (M, g) be a connected Riemannian manifold and $f : (M, g) \rightarrow (N, h)$ is totally geodesic, then $df = 0$.

Corollary C.4.2. $f : (M, g) \rightarrow (N, h)$ is a harmonic map, then

- 1.
- 2.

APPENDIX D. TOPOLOGY

D.1. The universal covering.

Definition D.1.1 (deck transformation). \widetilde{M} is the universal covering of M , the deck transformation group is defined as

$$\text{Aut}_\pi(\widetilde{M}) = \{F : \widetilde{M} \rightarrow \widetilde{M} \text{ is diffeomorphism} \mid \pi \circ F = \pi\}$$

Furthermore, we have the following facts

1. $\text{Aut}_\pi(\widetilde{M}) \cong \pi_1(M)$;
2. $\text{Aut}_\pi(\widetilde{M})$ acts on \widetilde{M} smoothly, freely and isometrically;
3. $\widetilde{M} / \text{Aut}_\pi(\widetilde{M}) \cong M$.

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