GNR-602 PROJECT

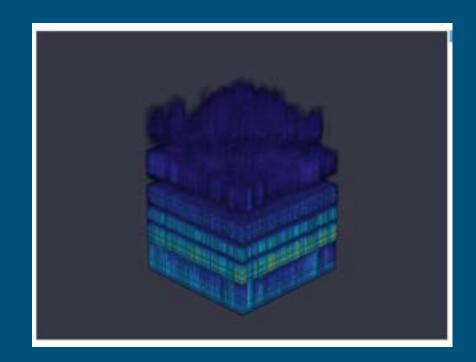
Spectral unmixing using K-Means and Fully Constrained Least squares (FCLS):

Group no.: 16

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Hyperspectral Imaging

- Hyperspectral imaging (HSI) is a technique that analyzes a wide spectrum of light instead of just assigning primary colors (red, green, blue) to each pixel.
- The narrowness and contiguous nature of measurements makes it hyperspectral.
- We collects many individual images of a scene, each taken within a different narrow wavelength band. The result is a "data cube" of the combined images.



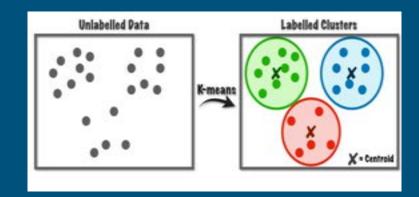
Data Cube

What is Spectral unmixing?

- Decomposition of a pixel spectrum into its constituent spectra
- Spectra of the pure materials, called endmembers, as well as their abundances in each pixel, are considered unknown
- The algorithm performs following tasks:
 - Endmember extraction :Finding the unknown endmember spectra
 - 2. Unmixing: Determining the corresponding abundances of each endmember in each pixel

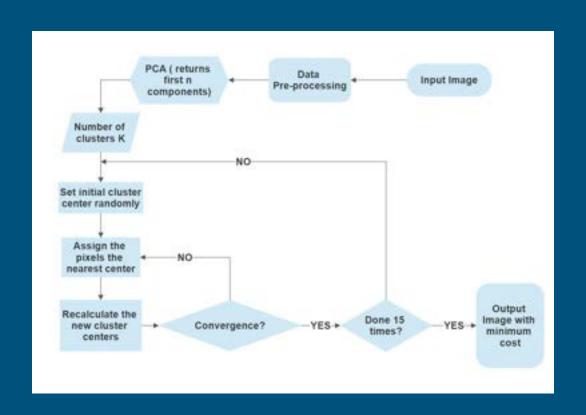
K-MEANS CLUSTERING

- k-Means clustering algorithm is used to detect the homogeneous regions in the image
- We initialize Θ randomly and report the results of the one that minimizes the cost function J of k-Means for different values of m.
- We use Squared Euclidean distance and Canberra distance



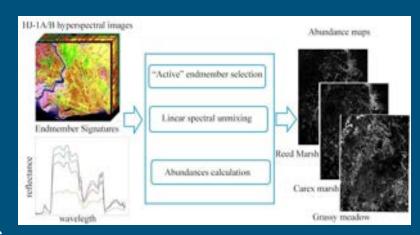
$$J(U,\Theta) = \sum_{i=1}^{N} \sum_{j=1}^{m} u_{ij} d(x_i,\theta_j)$$

K-MEANS CLUSTERING



Fully Constrained Least-squares (FCLS) Linear Unmixing

- Fully constrained least squares is least squares with the abundance sum-to-one constraint (ASC) and the abundance nonnegative constraint (ANC).
- Involves iterative computations to generate a nearly optimal solution.
- It requires a maximum of p 1 iterations and terminates when no more steering is required.



Fully Constrained Least-squares (FCLS) Linear Unmixing

Measurement model

$$\mathbf{x}_p = \mathbf{M}\mathbf{s}_p + \mathbf{n}_p$$

- Observation vector $\mathbf{x}_{tx1} = [x_1 \ x_2 \cdots x_l]^T$
- Material signature matrix $\mathbf{M}_{l \times c} = [\mathbf{m}_1 \mathbf{m}_2 \cdots \mathbf{m}_c]$
- Abundance fractions $\mathbf{s}_{e\times 1} = [s_1 \ s_2 \cdots s_e]^T$
- Nonnegative and sum-to-one constraints

$$s_j \ge 0$$
, $\sum_{j=1}^{c} s_j = 1$

$$J = \frac{1}{2} (\mathbf{r} - \mathbf{M}\alpha) (\mathbf{r} - \mathbf{M}\alpha)^T - \lambda \left(\sum_{j=1}^p \alpha_j - 1 \right).$$

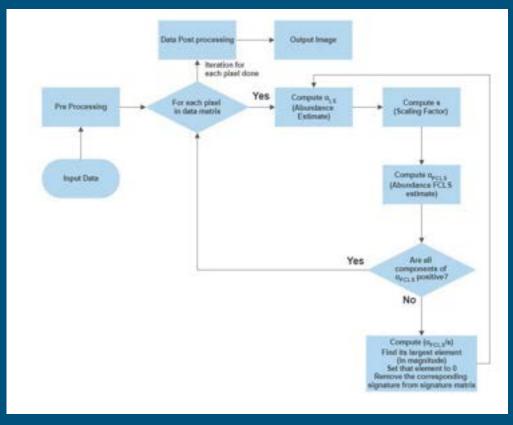


$$\frac{\partial J}{\partial \alpha}\Big|_{\hat{\alpha}_{ECLS}} = 0 \Rightarrow \hat{\alpha}_{LS} - \lambda s$$



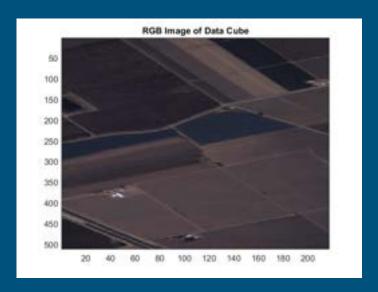
$$\lambda = (1 - 1^T \hat{\alpha}_{LS}) / (1^T s)$$
 with $s = (M^T M)^{-1} 1$

Fully Constrained Least-squares (FCLS) Linear Unmixing

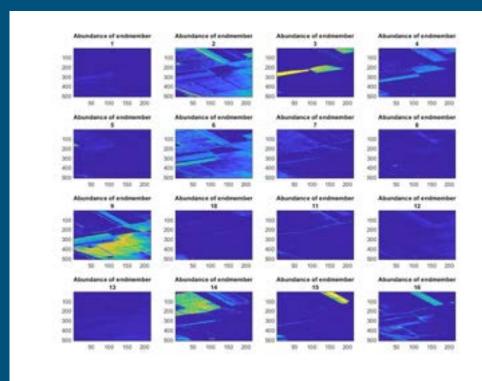


RESULTS (SALINAS)

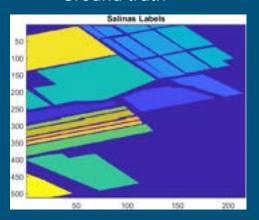
RGB Image



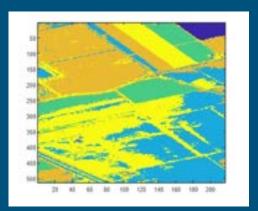
End members



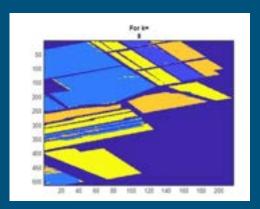
Ground truth



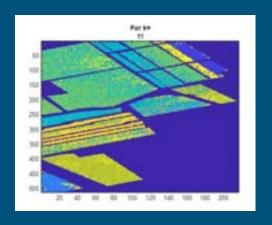
Predicted from FCLS



K-means (euclidean distance)



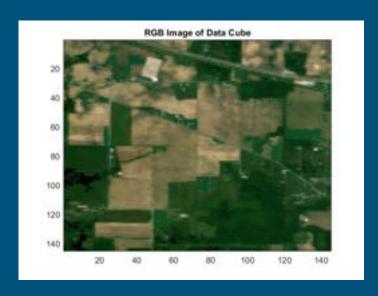
K-means (Canberra distance)



RESULTS (INDIAN PINES)

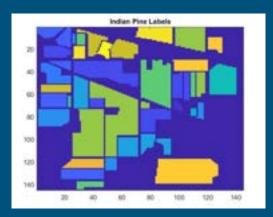
End members

RGB Image

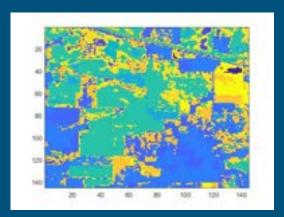




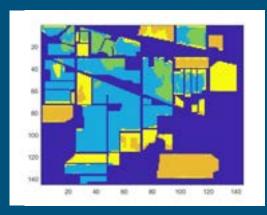
Ground truth



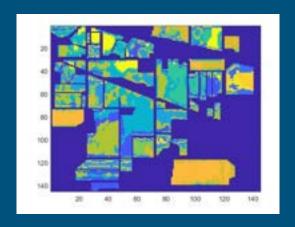
Predicted from FCLS



K-means (euclidean distance)

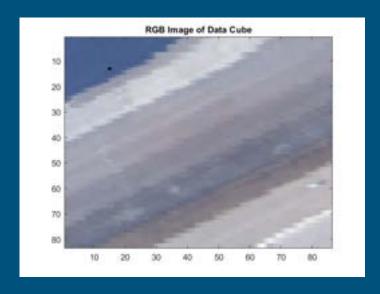


K-means (Canberra distance)

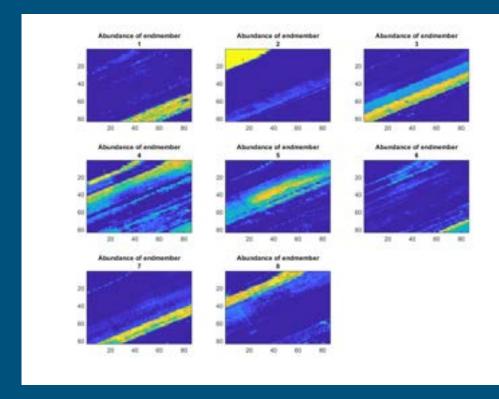


RESULTS (SALINAS-A)

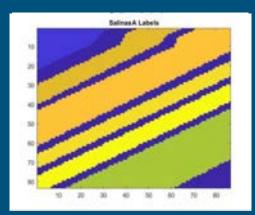
RGB Image



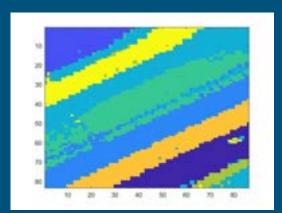
End members



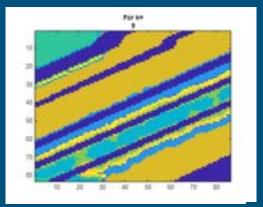
Ground truth



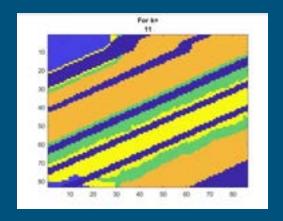
Predicted from FCLS



K-means (euclidean distance)

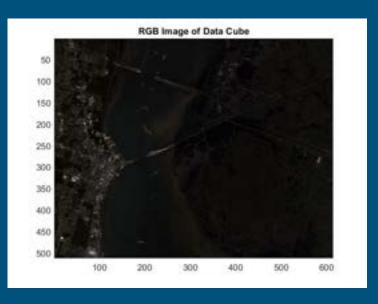


K-means (Canberra distance)

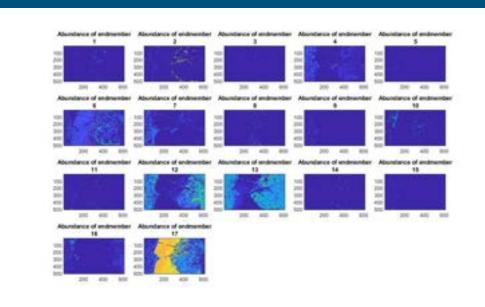


RESULTS (KSC)

RGB Image

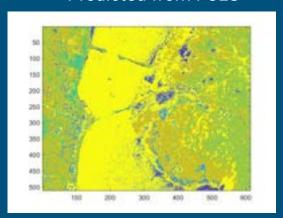


End members

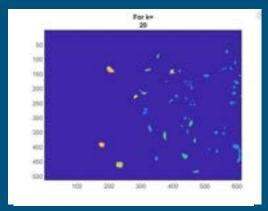


Ground truth

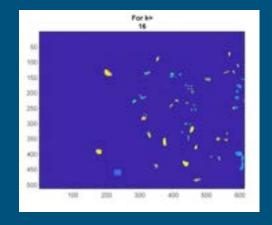
Predicted from FCLS



K-means (euclidean distance)

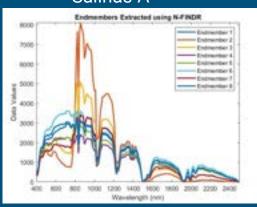


K-means (Canberra distance)

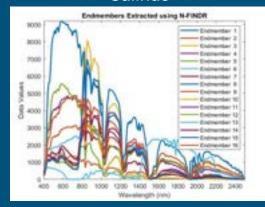


Endmembers extracted from each dataset

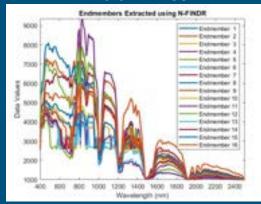
Salinas-A



Salinas



Indian Pines



Kennedy Space Centre

