Classification

Now that we have done analysis over globular clusters, weh know the distribution of flux from globular cluster x-ray sources, we can use that to indentify the quiscent sources, or at least say with certain confidence which sources (or which observation) are not in burst state.

We will feed only those observation for training data.

we have the confirmed classification of sources which are both in quiscent state, for those sources, we will pick out the catalogue in which observations are made when the fluxs are less than certain threshold (mean of globular cluster sources flux)

We can consider each observation as distinct source for training.

Classes Considered

```
LMXRB Black Holes
LMXRB Neutron Star
CV
Pulsars
```

LMXRB BH

All classes - LMXRB x-ray binary Black Hole

```
Number of sources - 123
Number of Observations - 1100
observation Meeting critrion - 824
```

Catalogues

RITTERLMXB
NGC3115CX0
INTREFCAT
XRBCAT
WGACAT
SAXWFCCAT
SAXWFCCAT2

LMXRB NS

All classes - LMXRB X-ray binary pulsars

```
Number of observations - 1268
observations meeting criteria - 971
```

Catalogues

XMMSSCLWBS IBISCAT

RITTERLMXB

INTREFCAT

RASS2MASS

XRBCAT

WGACAT

SMCWINGCXO

SAXWFCCAT

CV

```
Number of sources - 322
Number of observations - 3609
observation meeting criteria - 2604
```

Catalogues

GC47TUCCX0

XMMSLEWCLN

M31CFCX0

RASS2MASS

XMMSSCLWBS

IBISCAT

INTREFCAT

WGACAT

NGC6791CX0

 ${\tt CHICAGOCXO}$

M31CX02

INTVARCAT

INTIBISASS

INTIBISGAL

RASS6DFGS

RASSUSNOID

RBS

M83XMM

RASSCNDINS

Data Pre-procesing

All filters are average combined for flux

Dropped all empty rows and all empty columns

\$-log_{10}{flux}\$ is used

All data columns are normalised

Data columns distributions are normalised \$(x_i-mean)/var\$

Nan Values are imputed with 0

Data Filtering

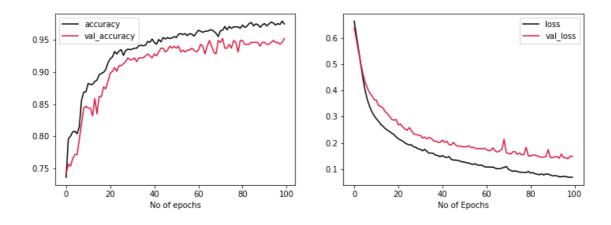
Filter	criteria
streak_src_flag	False
pileup_flag	False
mstr_streak_src_flag	False
Flux value	< \$10^{-12} erg/cm^2/s\$
Significance	> 20

Parameters Used

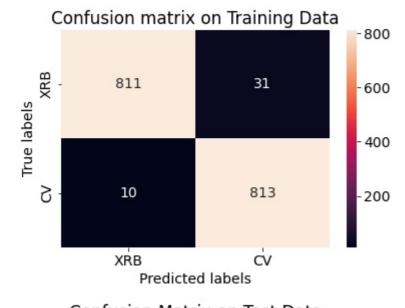
- FLux
- · Variability
- Hardness
- Model Fit

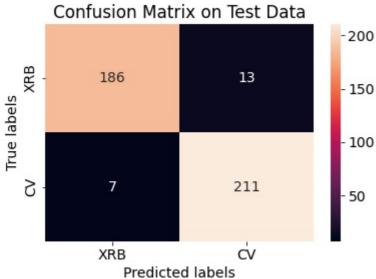
Result

X-ray binary (NS+BH) Vs CV

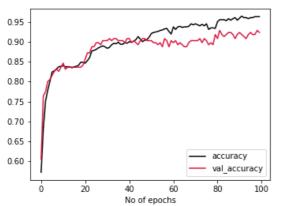


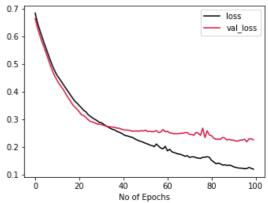
Training Accuracy - 97.9 % Test Accuracy - 95.20 %





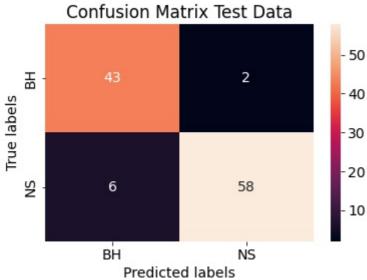
NS vs BH



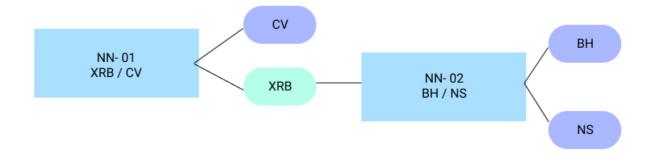


Training Accuracy - 96.5 % Test Accuracy - 92.3 %





Combined NN



Test data

