

Data processing software for RHESSI and STIX instruments

Technical Design Document

**OVERVIEW**

Object Spectral Executive (OSPEX) is an object-oriented interface for X-ray spectral analysis of solar data written in IDL.

Official source: <https://hesperia.gsfc.nasa.gov/ssw/packages/spex/doc/ospex_explanation.htm>

Our software is oriented to reconstruct some elements of OSPEX package and provide a tool to process the RHESSI data (Reuven Ramaty High Energy Solar Spectroscopic Imager, NASA Small Explorer Mission).

In addition to this, it is planned to include observed data from the X-ray Spectrometer(STIX) aboard Solar Orbiter(launch date is February 2020).

The project is presented on GitHub repository:

<https://github.com/LAbdrakhmanovaOBSPM/OSPEX-Object-Spectral-Executive-in-Python>

**PURPOSE OF THIS DOCUMENT**

This document is a Technical Design Document for use OSPEX written in Python. It provides guidance and materials which are intended to assist the relevant management or technical staff. It is also useful background reading for anyone involved in developing or using the application.

**SOFTWARE DEVELOPMENT TOOLS**

This section lists the tools chosen to assist software development including testing.

Recommendations tools:

*Python 3.6*

*Libraries:*

*Astropy*

*Numpy*

*Pandas*

*Matplotlib*

*Tkinter*

Following modules should come by default with Python after installation but we recommend you to check if they work correctly. If not, please, import them manually:

*Webbrowser*

*Datetime*

*Re*

*Copy*

*Importlib*

To run:

*Linux, Mac, Windows:* When imported the libraries you should call main script from terminal*:*

1. Open the terminal*:* ***cmd***

2. Go to the directory where your program is located: ***cd*** *directory location*

3. Run the software by calling the main script*:* ***python3 main.py***

If you want to work with content and make modifications:

1. Go to the directory where your program is located:

*cd directory location*

2. Call Python:

*python3*

3. Upload file content from the plotting.py script :

*from plotting import \**

4. Choose desired file(s), "Data" folder:

*File = Input(“filename.fits”)*

Now you are able to call any parameters and functions of the .fits file

5. Some examples:

a) to load the parameter RATE from DATA:

*File.rate*

b) to plot Spectrum for Flux:

*File.plot\_spectrum\_flux()*

**CLASSES, COMPONENT DESCRIPTION AND ACRONYMS**

For the software implementation and test, this and the previous section provide sufficient information for a programmer to produce the software, and for a user, who is maybe not a developer.

The logical and physical structure of the application and components will be defined in detail:

* meaning of the classes;
* description of each parameter;
* relationships between the elements, i.e. the structure;
* initial values of each element.

|  |  |
| --- | --- |
| ***Classes*** | |
| SecondWindow | Used to describe the procedures to create an interface for Select Input part |
| Input | Provides the methods to load the parameters from input data and plot Spectrum, Time Profile and Spectrogram |
| Fitting | Performs a spectrum fitting. Creates a new window called “SPEX Fit Options” |
| BackgroundWindow | A window to edit time intervals and create background plots |
| backround\_plot | Main function to create background plots |
| ***Component description*** | |
| Rate | Count Rate data in each energy channel |
| Time | Array of Low energy ranges |
| Time\_del | Accumulation time |
| E\_min | Array of Low energy ranges |
| E\_max | Array of High energy ranges |
| deltaE | Energy range from E\_min to E\_max(3-250 keV) |
| TimeNew | Array of converted time from sec to hours:min:sec |

Additional information can also be found in the code comments.

**FUNCTIONS TO FIT AND INPUT PARAMETERS**

**Category: spectral fitting**

1. One dimensional Power Law.

Parameters:

* amplitude – model amplitude at the reference energy;
* x – reference energy;
* alpha – power law index.

2. Broken Power Law function - One dimensional power law model with a break.

Parameters:

* amplitude – model amplitude at the break energy;
* x\_break – break energy;
* alpha 1 – power law index for x<x\_break;
* alpha 2 – power law index for x>x\_break.

3. One dimensional Gaussian model.

Parameters:

* amplitude – amplitude of the Gaussian;
* mean – mean of the Gaussian;
* stddev – standard deviation of the Gaussian.

4. 1D Polynomial model.

Parameters:

* degree – degree of the series.

5. Exponential function.

Parameters:

* t0 – normalization;
* t1 – pseudo temperature.

6. Single power – law times an exponential.

Parameters:

p - first 3 parameters describe the single power – law function, e - exponential;

* p0 – normalization at epivot for power – law;
* p1 – negative power - law index;
* p2 – epivot (keV) for power – law;
* e3 – normalization for exponential;
* e4 – pseudo temperature for exponential.

**APPLICATION ARCHITECTURE**

