# GalPot (Galaxy potential following falcON)

This documentation is just here to explain how to compile and give parameters to GalPot. See Section 2.3 of Dehnen & Binney (1998) for a fuller explanation of how it works. See testGalPot for examples of what the code can do.

# 1 Compilation

"make" makes the library and example executables.

"make MyExecutable.exe" will compile a new executable written by the user and called "MyExecutable.cc"

The command "make GalPot" just makes the GalPot libraries.

**N.B.** In all cases here, the length unit is kpc, the time unit is Myr, and the mass unit is  $M_{\odot}$ . For reference,  $G = 4.49866 \times 10^{-12} \text{kpc}^3 \text{Myr}^{-2} \text{M}_{\odot}^{-1}$ . A velocity of 1 kpc Myr<sup>-1</sup> = 977.77km s<sup>-1</sup> (See Units.h for further).

# 2 Input

The file that is given as the input file for the potential must be of the following form (without the comment lines):

```
3
                                       # No. Disks (0 < No.disks < 3)
8.90e7
                                       # Param. for each of the three disks.
         1.8
              0.04
3.50e7
         1.2
              -0.15
                                       # N.B essential to give 5 param.
1.30e7
                                       \# No. lines of param. = No. disks
               0.25
                                       # No. Spheroids (0 \le \text{No.sph.} \le 2)
                                       \# Param. for each spheroid.
4.0e7
         0.8
              -0.5
                      1.4
                            10
2.0e7
         0.2
              -1.5
                                       # essential to give 6 parameters
```

### 2.1 Disk

The disk parameters, as given in that file are (in order)  $\Sigma_0$ ,  $R_d$ ,  $z_d$ ,  $R_0$ ,  $\epsilon$ .

 $\Sigma_0$  is the disk's central surface density (in  $M_{\odot} \text{kpc}^{-2}$  and the absence of a cutoff),  $R_d$  is the disk scale radius,  $z_d$  its scale height (though note there is a difference between negative and positive values), and  $R_0$  is an inner cutoff radius (all in kpc). The term  $\epsilon$  perturbs the disc from a pure exponential, with a peak fractional change in surface density of  $\sim \epsilon$  and a scale length of  $2R_d$ . In numerical terms (cylindrical polars), these give a surface density

$$\Sigma(R) = \Sigma_0 \exp\left(-\frac{R_0}{R} - \frac{R}{R_d} + \epsilon \cos\left(\frac{\pi R}{R_d}\right)\right),\tag{1}$$

i.e. a standard exponential disk with optional hole in the middle and/or  $\epsilon$  term modulation.

The vertical structure of the density is either an exponential in |z| (if  $z_d > 0$ ), or isothermal (if  $z_d < 0$ ), i.e.

$$\rho_d(R, z) = \begin{cases} \frac{\Sigma(R)}{2z_d} \exp\left(\frac{-|z|}{z_d}\right) & \text{for } z_d > 0\\ \frac{\Sigma(R)}{4(-z_d)} \operatorname{sech}^2\left(\frac{z}{2z_d}\right) & \text{for } z_d < 0. \end{cases}$$
 (2)

**N.B.** There is a factor of two in the denominator of the sech<sup>2</sup> profile. That is so that they tend to the same thing as  $z \to \infty$ .

#### 2.2**Spheroids**

The spheroid parameters are (in order)  $\rho_0$ , q,  $\gamma$ ,  $\beta$ ,  $r_0$ ,  $r_{cut}$ .  $\rho_0$  is a scale density (in  $M_{\odot} kpc^{-3}$ ; q is the axis ratio (q < 1 is flatter than a sphere, q > 1 is prolate);  $\gamma$  is the inner density slope,  $\beta$  the outer density slope;  $r_0$  is a scale radius and  $r_{cut}$  is a cutoff radius (both in kpc). This corresponds to a density profile

$$\rho_s = \frac{\rho_0}{(r'/r_0)^{\gamma} (1 + r'/r_0)^{\beta - \gamma}} \exp\left[-\left(r'/r_{cut}\right)^2\right],\tag{3}$$

where, in cylindrical coordinates

$$r' = \sqrt{R^2 + (z/q)^2} \tag{4}$$

### 3 Use

The example executable testGalPot gives examples of the use of the GalaxyPotential class.

# References

Dehnen W., Binney J., 1998, MNRAS, 294, 429