

# Bayesian methods for quantifying global parameter tensions between cosmological datasets

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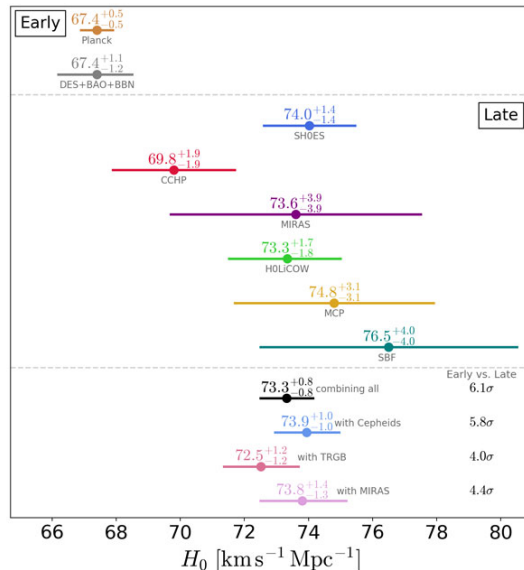


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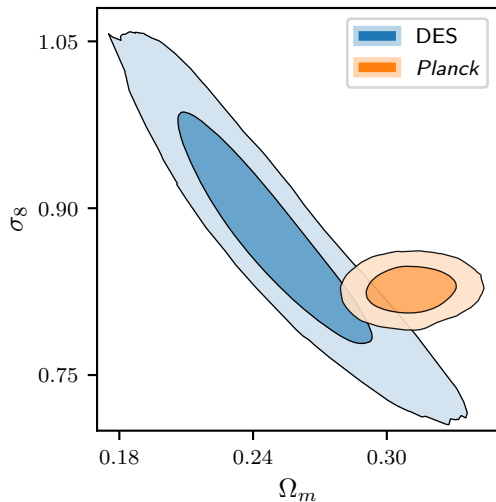
# Cosmological parameter tensions

- ▶ Measurements  $H_0$  differ between early and late time observations [1907.10625]
- ▶ “Tension” means a disagreement between different datasets on the inferred value of model parameters.
- ▶ The presence of tension indicates an error in the model and/or at least one of the datasets.
- ▶ It is statistically incorrect to combine datasets when they are in tension.



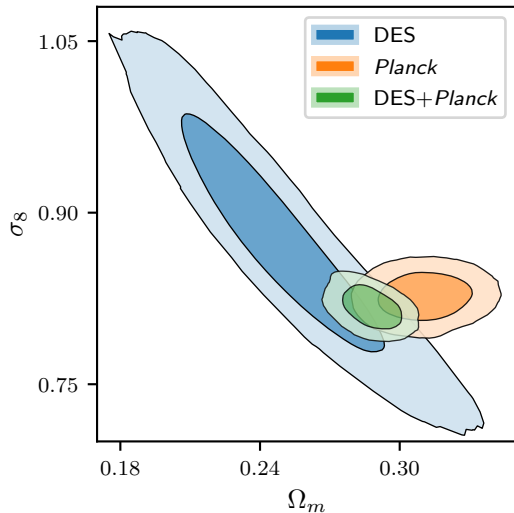
# The importance of global measures of tension

- ▶ In other situations the discrepancy doesn't exist in a single interpretable parameter
- ▶ For example: DES+*Planck* [1902.04029]
- ▶ Are these two datasets in tension?
- ▶ Can we confidently combine them?
- ▶ There are a lot more parameters – are we sure that we've chosen wisely?



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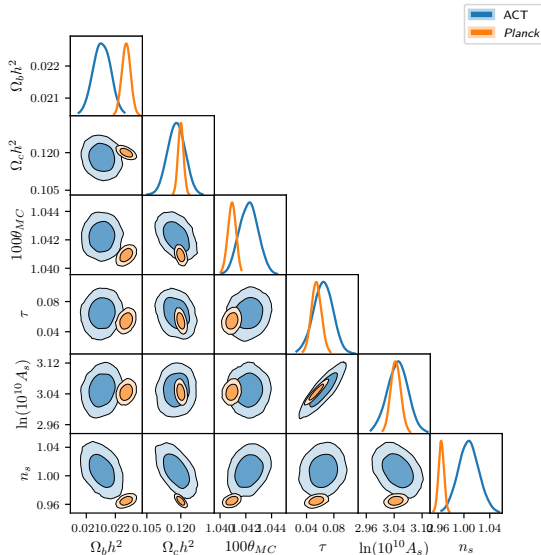


# The perils of manual marginal inspection

- ▶ If you have enough parameters, then you might expect that tensions would naturally arise in some combinations by chance.
- ▶ For example, if you take ACT and *Planck*, and construct a linear combination of parameters in maximum tension:

$$t = -\Omega_b h^2 + 0.022\Omega_c h^2 + 34\theta_{MC} - 0.092\tau \\ + 0.05\ln(10^{10}A_s) + 0.067n_s$$

- ▶ In general you would expect such a parameter to be in  $\sim \sqrt{d} - \sigma$  tension [2007.08496]

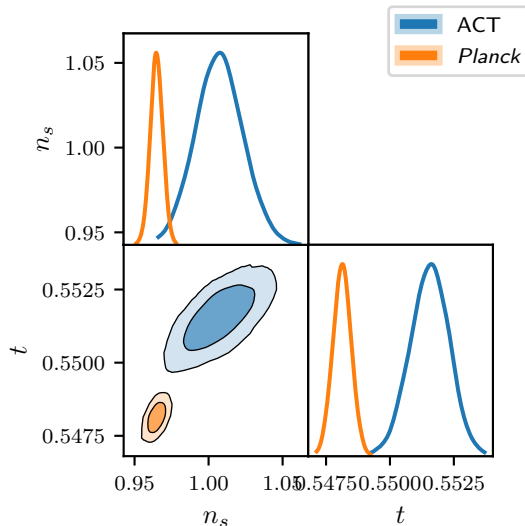


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# Bayesian language

## Notation

Datasets:  $A$  and  $B$  (e.g. *Planck* and DES)

Model:  $M$  (e.g.  $\Lambda$ CDM)

Parameters:  $\theta$  (e.g.  $(\Omega_m, \sigma_8)$ )

Likelihoods:  $\mathcal{L}$ :  $P(A|\theta)$ ,  $P(B|\theta)$

## Inference

Prior:  $\pi$ :  $P(\theta)$

Posteriors:  $\mathcal{P}$ :  $P(\theta|B)$ ,  $P(\theta|A)$ , (evaluate samples).

Bayesian evidences:  $\mathcal{Z} = \langle \mathcal{L} \rangle_\pi$ ,  $P(A)$ ,  $P(B)$  [1506.00171]

Bayes theorem:  $\mathcal{L} \times \pi = \mathcal{P} \times \mathcal{Z}$

## Anatomy

Kullback–Leibler divergence:  $\mathcal{D} = \langle \log \mathcal{P} / \pi \rangle_{\mathcal{P}} \sim \log \text{Vol}(\pi) / \text{Vol}(\mathcal{P})$  [1902.04029]

Model dimensionality:  $d = 2 \times \text{var}(\log \mathcal{L})$  [1903.06682]

Occam's razor equation:  $\log \mathcal{Z} = \langle \log \mathcal{L} \rangle_{\mathcal{P}} - \mathcal{D}$  [2102.11511]

(Released today by Hergt et al)

## Suspiciousness statistic [1902.04029] [2007.08496]

- ▶ The natural Bayesian measure of tension is the Bayes ratio

$$\mathcal{R} = \frac{\mathcal{Z}_{AB}}{\mathcal{Z}_A \mathcal{Z}_B} = \frac{P(A, B)}{P(A)P(B)} = \frac{P(A|B)}{P(A)} = \frac{P(B|A)}{P(B)} \quad (1)$$

- ▶  $\mathcal{R}$  is prior dependent, one can artificially reduce tension by drawing arbitrarily wide priors.
- ▶ Can remove this prior dependency by dividing out the KL-dependent Occam factor to give a “Suspiciousness”, computable from three MCMC chains:

$$\log S = \langle \log L_{AB} \rangle_{\mathcal{P}_{AB}} - \langle \log L_A \rangle_{\mathcal{P}_A} - \langle \log L_B \rangle_{\mathcal{P}_B} \quad (2)$$

- ▶ Can be interpreted as the maximum Bayes ratio  $\mathcal{R}$  allowed by reasonable priors.
- ▶ In the Gaussian case it is related to the usual Mahalanobis distance

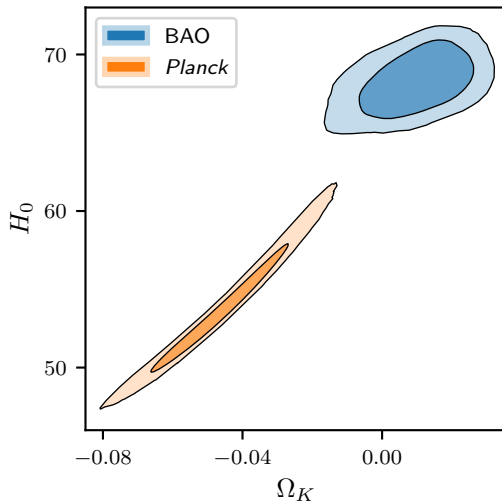
$$\log S = \frac{d}{2} - \frac{1}{2}(\mu_A - \mu_B)^T (\Sigma_A + \Sigma_B)^{-1} (\mu_A - \mu_B) \quad (3)$$

which can be used to calibrate it into a tension probability and “ $\sigma$ ” quantification.



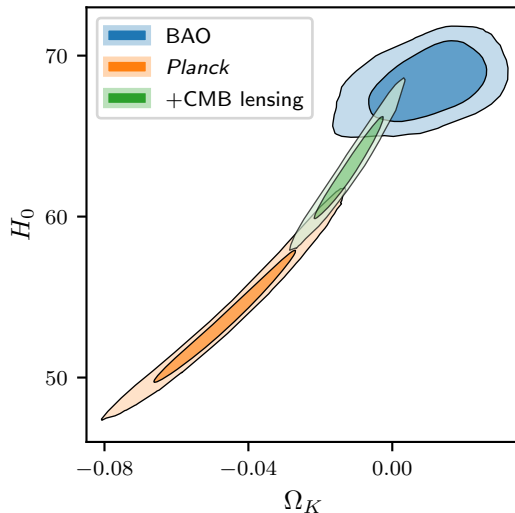
## Curvature tension $\Omega_K$

- ▶  $\Lambda$ CDM assumes the universe is flat
- ▶ If you allow  $\Omega_K \neq 0$ , *Planck* (plikTTTEEE) has a moderate preference for closed universes (50:1)
- ▶ *Planck*+CMB lensing +BAO strongly prefer a flat universe
- ▶ But, *Planck* vs lensing is  $2.5\sigma$  in tension, and *Planck* vs BAO is  $3\sigma$ .
- ▶ This is reduced if plik  $\rightarrow$  camspec
  - ▶ Di Valentino et al [1911.02087]
  - ▶ Handley [1908.09139]
  - ▶ Efsthathiou & Gratton [2002.06892]
- ▶ BAO and lensing summary statistics and compression strategy assume  $\Lambda$ CDM.



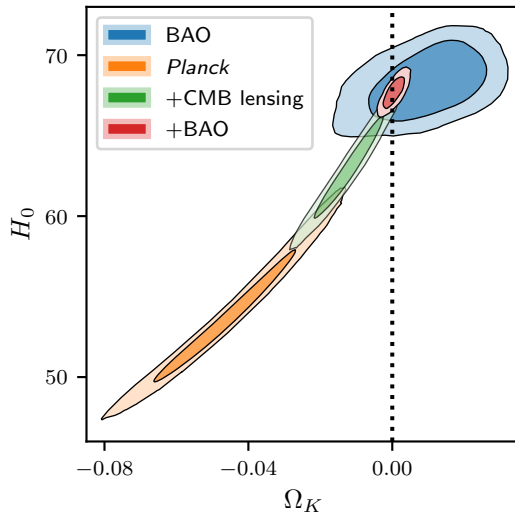
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Data	Model	Tension	Reference
DES vs Planck	$\Lambda$ CDM	$2.1\sigma$	[1902.04029]
ACT vs Planck+SPT	$\Lambda$ CDM	$2.8\sigma$	[2007.08496]
CMB lensing vs Planck	$K\Lambda$ CDM	$2.5\sigma$	[1908.09139]
BAO vs Planck	$K\Lambda$ CDM	$3\sigma$	[1908.09139]

Slides, figures and plotting code available at:

[https://github.com/williamjameshandley/talks/tree/tehran\\_2021](https://github.com/williamjameshandley/talks/tree/tehran_2021)