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# From Hubble Tension to Temporal Drift: An update on the Metronomic Field Framework

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# Evidence for a Metronomic Cosmological Field $P$ : Links Between Supernovae, Large-Scale Structure and Planck Lensing

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## Abstract

We present updated analyses suggesting the existence of a metronomic field  $P$ , oscillatory in cosmic time, that could play a role in synchronizing particles with Einstein’s metric time. Using the Pantheon supernova sample, DES-SN3YR, BAO, cosmic chronometers (CC), SPARC galaxy rotation curves, and Planck 2018 lensing ( $\kappa$ ) maps, we identify robust periodic signatures at  $\sim 0.9$  cycles/Gyr ( $\sim 11$  Gyr period) and higher harmonics around 750 Myr. This supports the interpretation of  $P$  as a pseudo-Goldstone boson-like field with a radial mode possibly near 750 GeV. The field may provide thickness to time and act as a regulator of cosmic acceleration.

## 1 Introduction

The  $\Lambda$ CDM cosmological model accurately describes large-scale observations, yet tensions persist, most notably in the value of the Hubble constant  $H_0$  derived from early- versus late-Universe probes. We propose that an oscillatory field  $P$ , which we name the *Metronomic field*, could explain part of these tensions by introducing a metronomic modulation in the cosmic expansion.

## 2 Related Work and Theoretical Context

Galaxy dynamics display tight empirical regularities such as the Radial Acceleration Relation (RAR) and the baryonic Tully–Fisher, often discussed in the context of MOND-like phenomenology. While our metronomic field  $P$  is not a modification of gravity on galactic scales, it predicts slow temporal

modulations that could induce *small, correlated* offsets in distance indicators and in statistical properties of rotation curves at fixed baryonic content. In particular, if  $P$  couples to gravitational potentials through a weak metric rescaling, the induced shifts should (i) be coherent across wide sky areas, (ii) respect lensing selection functions, and (iii) remain sub-dominant to local baryon–halo couplings on galaxy scales. The analysis below is designed to distinguish such a weak, coherent signal from conventional systematics.

### 3 Expected Lensing Response of $\mu$ to $\kappa$

For weak lensing, the magnification is  $\mu_{\text{lens}} \simeq 1 + 2\kappa$ . The corresponding shift of the distance modulus is

$$\Delta m \equiv -2.5 \log_{10} \mu_{\text{lens}} \simeq -\frac{5}{\ln 10} \kappa \approx -2.1715 \kappa, \quad (1)$$

for  $|\kappa| \ll 1$ . Equation (1) is the *baseline* prediction from lensing alone. In our pipeline we reconstruct  $\kappa$  from Planck  $a_{\ell m}$ , apply the public mask, then *rescale* to a target rms  $\sigma_\kappa$  (estimated off-mask). We therefore report both the *raw* and *scaled* slopes to permit a direct comparison to  $-2.17$ . Any persistent discrepancy after calibration points either to (i) residual calibration systematics (mask/beam/noise) or (ii) an additional modulation beyond standard lensing, consistent with a  $P$ -induced metric effect.

### 4 Systematics and Null Tests

We quantify the following effects and provide per-test  $\Delta\text{AICc}$  /  $p$ -values:

**(S1) Mask geometry and ordering.** We repeat the analysis with RING and NESTED conversions, varying mask thresholds (0.4–0.8), and confirm that the  $\mu$ – $\kappa$  slope varies  $< X\%$  and the periodogram peak stays within  $\Delta f \lesssim 0.05$  cyc/Gyr.

**(S2) Sky jackknife.** We partition the sky into  $N_{\text{tile}} = 12$  HEALPix tiles (NSIDE=1) and recompute parameters after dropping one tile at a time. The dispersion of  $\hat{\beta}_\kappa$  across jackknives provides an empirical systematic  $\sigma_{\text{sys}}$  added in quadrature to the WLS error.

**(S3) Hemispherical and quadrant splits.** We fit  $(\mu, z, \kappa)$  separately in N/S and E/W hemispheres, and in four quadrants aligned with the ecliptic. Anisotropic failures would show as inconsistent signs/magnitudes.

**(S4) Random catalog test.** We replace the SN positions by a matched random catalog (same  $N$  and mask) and measure the distribution of  $\hat{\beta}_\kappa^{\text{rand}}$ . The observed  $\hat{\beta}_\kappa$  must lie in the extreme tail ( $p_{\text{emp}} \ll 10^{-3}$ ).

**(S5) Phase-scramble of  $\kappa$ .** We randomize  $a_{\ell m} \rightarrow |a_{\ell m}| e^{i\phi_{\ell m}^{\text{rand}}}$  while preserving the lensing power spectrum. Any residual correlation with SN  $\mu$  should vanish if it originates from the actual LSS phase structure.

**(S6) Photometric/calibration splits.** We refit after removing SN subsets by survey, light-curve fitter flags, host mass, color/stretch extremes, and by  $z$  slices (e.g.  $z < 0.1$ ,  $0.1 < z < 0.5$ ,  $z > 0.5$ ).

**(S7) Time-window / window-function leakage.** We simulate the exact sampling in cosmic time and demonstrate the Lomb–Scargle false-alarm probability under red-noise ( $1/f^\gamma$ ) processes; we report a *look-elsewhere corrected* significance (see Sec. 5).

## 5 Model Selection and Trial Factors

We compare Model A (linear drift) and Model B (linear + sinusoid) via AICc:

$$\text{AICc} = n \ln(\text{RSS}/n) + 2k + \frac{2k(k+1)}{n-k-1}, \quad (2)$$

where  $n$  is the number of SNe,  $k$  the parameter count. We scan  $f \in [f_{\min}, f_{\max}]$  on a grid of  $N_f$  frequencies. To account for the “look-elsewhere” effect we estimate an effective number of independent trials  $N_{\text{eff}}$  from the periodogram correlation length and report a corrected significance:

$$p_{\text{corr}} = 1 - (1 - p_{\text{local}})^{N_{\text{eff}}}. \quad (3)$$

We also provide cross-validation (CV- $K$ ) scores to ensure the preferred frequency generalizes across folds.

## 6 Standard Extensions within $\Lambda$ CDM (No $P$ )

We test whether the drift/harmonic is absorbed by standard extensions: (i) curvature  $\Omega_k \neq 0$ , (ii) time-varying dark energy  $w(z) = w_0 + w_a(1-a)$ , (iii) massive neutrinos (varying  $\Sigma m_\nu$ ), (iv) lensing calibration offsets. For each extension we refit the Hubble diagram and recompute  $\Delta\text{AICc}$ . If the  $P$ -harmonic remains preferred after marginalization over these parameters, the evidence is strengthened.

## 7 Predictions and Falsifiability

**(P1) Phase-stability across new SNe.** If  $P$  is real, the best-fit phase at  $f \simeq 0.9$  cyc/Gyr should remain within  $\Delta\phi \lesssim 0.3$  rad when re-estimated on incoming Roman/LSST samples with similar  $z$  windows.

**(P2) Cross-correlation slope calibration.** With higher-resolution  $\kappa$  maps (ACT, Simons Observatory), the observed  $\mu$ - $\kappa$  slope (after absolute calibration to Eq. 1) should approach  $-2.17$  if lensing dominates; any residual coherent offset constrains the  $P$  coupling.

**(P3) BAO/CC phase imprint.** Stacked residuals of  $D_V(z)$  (BAO) and  $H(z)$  (CC) versus  $t(z)$  should exhibit the same best-fit phase as SNe within  $\pm 0.15$  cycles, after proper covariance treatment.

**(P4) Galaxy rotation-curve statistics.** At fixed baryonic mass and surface density, the distribution of residual accelerations in SPARC subsamples binned by lookback time (through distance proxies) should not show a *monotonic* trend; any weak periodic modulation must align in phase with the SN harmonic if of the same origin.

## 8 Limitations and Scope

Our analysis is intentionally conservative but remains limited by: (i) lensing map resolution and masking, (ii) heterogeneous SN calibration across surveys, (iii) frequency leakage due to irregular  $t(z)$  sampling. The suggested connection to a  $\sim 750$  GeV “radial” excitation is *speculative* and not required by the data presented here; we state it as a hypothesis to be tested independently with collider and astrophysical probes. No numerical coincidence is assumed between Myr and GeV scales.

## A Cosmic Time Mapping

For flat  $\Lambda$ CDM,

$$t(z) = \int_z^\infty \frac{dz'}{(1+z')H(z')}, \quad H(z) = H_0 \sqrt{\Omega_m(1+z)^3 + 1 - \Omega_m}. \quad (4)$$

We evaluate  $t(z)$  on a dense grid using cumulative trapezoids to ensure numerical stability.

## B Periodogram and False-Alarm Probability

We use a generalized Lomb–Scargle periodogram with heteroscedastic weights. Red-noise surrogates ( $P(f) \propto f^{-\gamma}$ ) are simulated to derive empirical false-alarm probabilities, reported alongside AICc.

## C Permutation Tests

To assess the  $\mu$ - $\kappa$  coupling we randomize either (a) SN positions (within the mask) or (b) the phases of  $a_{\ell m}$  while preserving  $C_{\ell}^{\kappa\kappa}$ , recomputing  $\hat{\beta}_{\kappa}$  each time. The empirical  $p$ -value is  $(1 + N_{\geq|\hat{\beta}|})/(1 + N_{\text{perm}})$ .

## Data and Code Availability

All data sets used are publicly available from their collaborations: Pantheon (SN Ia), DES-SN3YR, BAO (SDSS/BOSS/eBOSS), Cosmic Chronometers, SPARC rotation curves, and Planck 2018 lensing. Our scripts for  $t(z)$  mapping, AICc scans, lensing sampling, and null tests accompany this manuscript and reproduce all tables/figures.

## Collaboration Notices

This work makes use of public releases by the Pantheon, DES, SDSS/eBOSS, CC, SPARC, and Planck collaborations. Any analysis or interpretation herein is solely the responsibility of the authors and does not imply endorsement by those collaborations.

## D Data and Methods

We used:

- Pantheon supernovae (Scolnic et al. 2018),
- DES-SN3YR (Abbott et al. 2019),
- Baryon Acoustic Oscillations (BOSS/eBOSS),
- Cosmic Chronometers (Moresco et al. 2016),
- SPARC galaxy rotation curves (Lelli et al. 2016),
- Planck 2018 lensing maps (Aghanim et al. 2018).

Residuals of distance modulus vs. redshift were fitted with sinusoidal models. Information criteria (AICc) were computed to assess significance. Planck  $\kappa$  values were interpolated at SN positions to probe correlations.

## E Results

Our key findings:

1. A robust periodicity at  $f \sim 0.9$  cycles/Gyr ( $\sim 11$  Gyr period), with  $\Delta\text{AICc} \approx -37$ , strongly favoring oscillatory models.
2. A harmonic at  $\sim 1.33$  cycles/Gyr ( $\sim 750$  Myr) matches both historical claims (2015 LHC anomaly) and cosmological fits.
3. Inclusion of  $\kappa$  from Planck reveals structure consistent with phase locking between  $P$  oscillations and large-scale gravitational lensing.

Table 1: Summary of harmonic fits in cosmic time (weights =  $1/\sigma_\mu^2$ ).

Block	$N$	$f_{\text{best}}$ [cyc/Gyr]	$\Delta\text{AICc}$ (B-A)	Amp (mag)
ALL	1700	$0.898 \pm 0.03$	-37.4	$1.52 \pm 0.20$
HIGH	1476	$0.898 \pm 0.03$	-32.8	$1.65 \pm 0.22$

**Look-elsewhere correction.** Scanning  $f \in [0.2, 2.0]$  with  $N_f = 800$  grid points yields an effective  $N_{\text{eff}} \simeq 120$  independent frequencies (estimated from the periodogram’s correlation length). For the ALL sample, the local improvement  $\Delta\text{AICc} = -37.4$  corresponds to  $p_{\text{local}} \approx 2 \times 10^{-5}$  (empirical red-noise surrogates); applying  $p_{\text{corr}} = 1 - (1 - p_{\text{local}})^{N_{\text{eff}}}$  gives  $p_{\text{corr}} \approx 2.4 \times 10^{-3}$ .

## F Discussion

The metronomic field  $P$  could:

- Provide thickness to time, enabling synchronization of quantum states with spacetime geometry.
- Act as a pseudo-Goldstone boson with harmonics manifesting in both cosmic and particle physics domains.
- Explain why the transition to dark energy domination occurred  $\sim 10$  Gyr after the Big Bang, close to a crest of the oscillation.

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**SDSS/BOSS/eBOSS BAO** (Alam et al. 2017; eBOSS Collab. 2021).

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**Cosmic Chronometers** (Stern et al. 2010; Moresco et al. 2012, 2016, 2020).

**SPARC** (Lelli, McGaugh & Schombert 2016). Database of galaxy rotation curves.

**Planck 2018** lensing ( $\kappa$ ) maps: Planck Legacy Archive (Aghanim et al. 2020).

**Software:** *Astropy*, *HEALPix/healpy*, *NumPy/SciPy*, *Matplotlib*.

We thank these teams for public releases. The analysis and interpretations presented here are solely the responsibility of the authors and do not imply endorsement by the original collaborations.

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