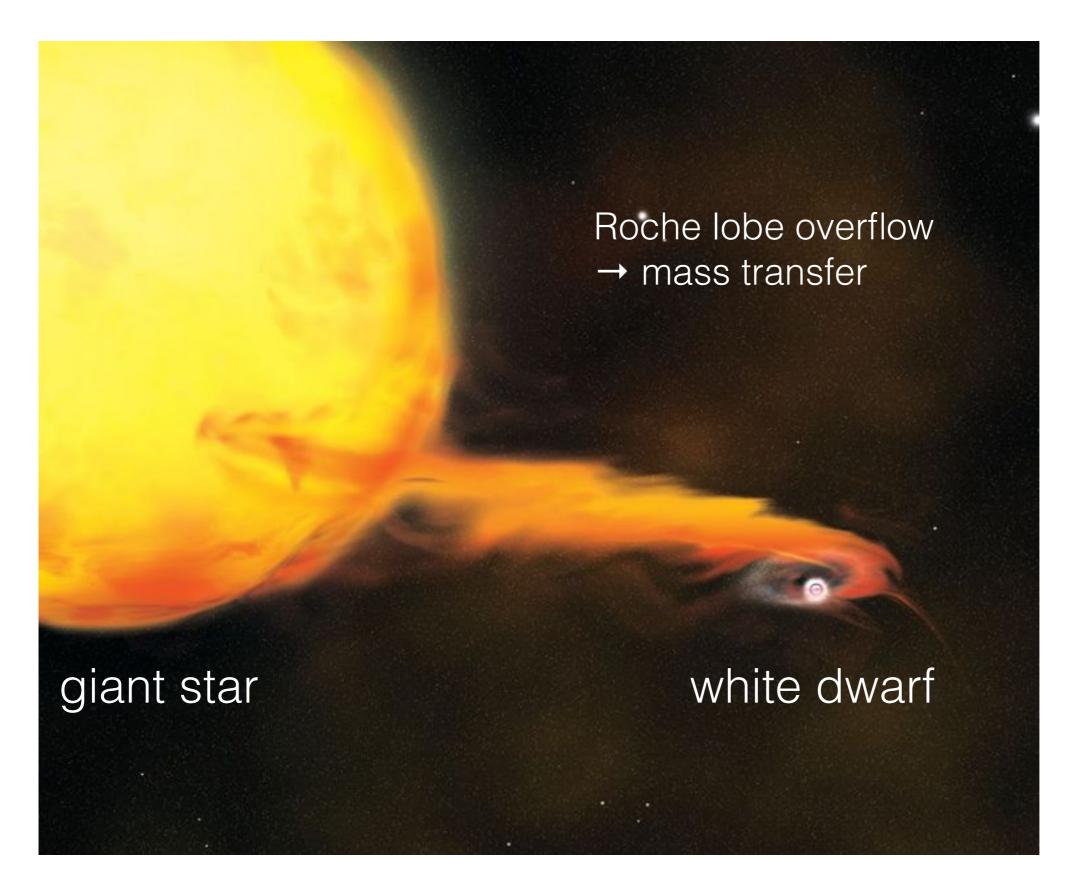
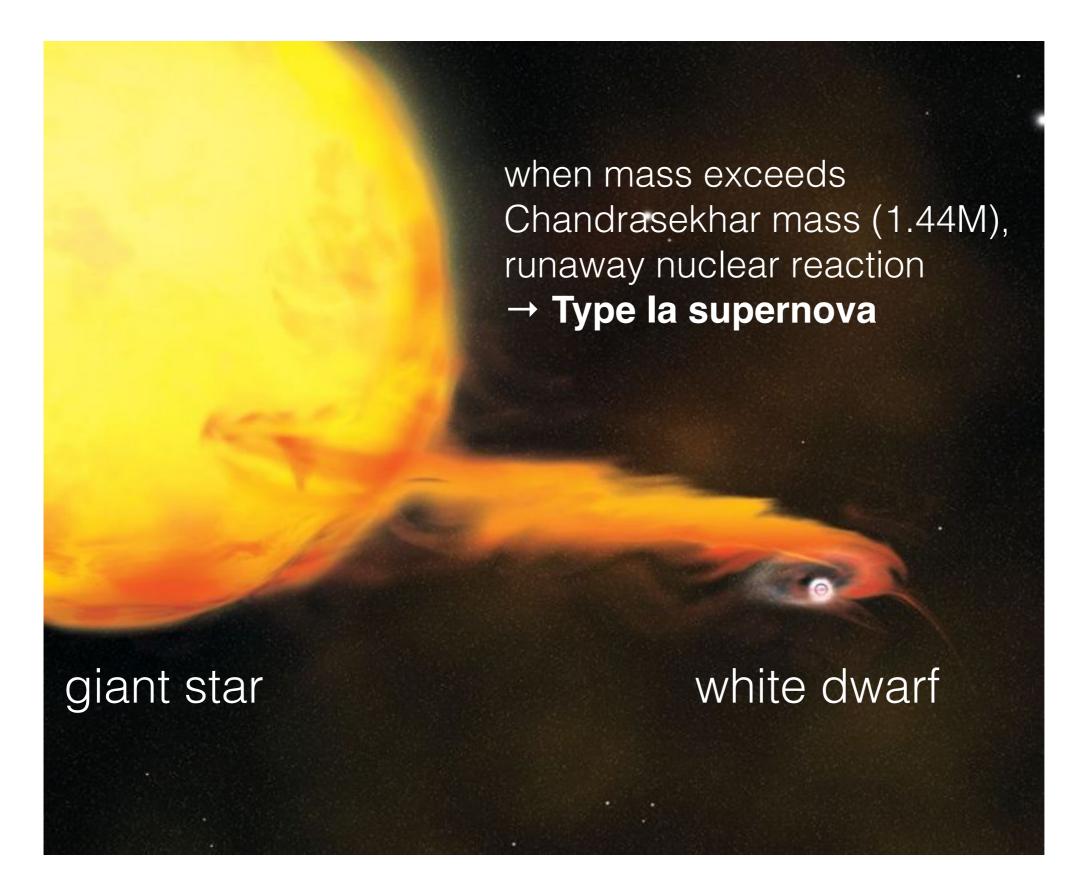
LAB 1/ Hubble's/Law

W3986: Astrostatistics

Quick bit of physics... Type la supernova



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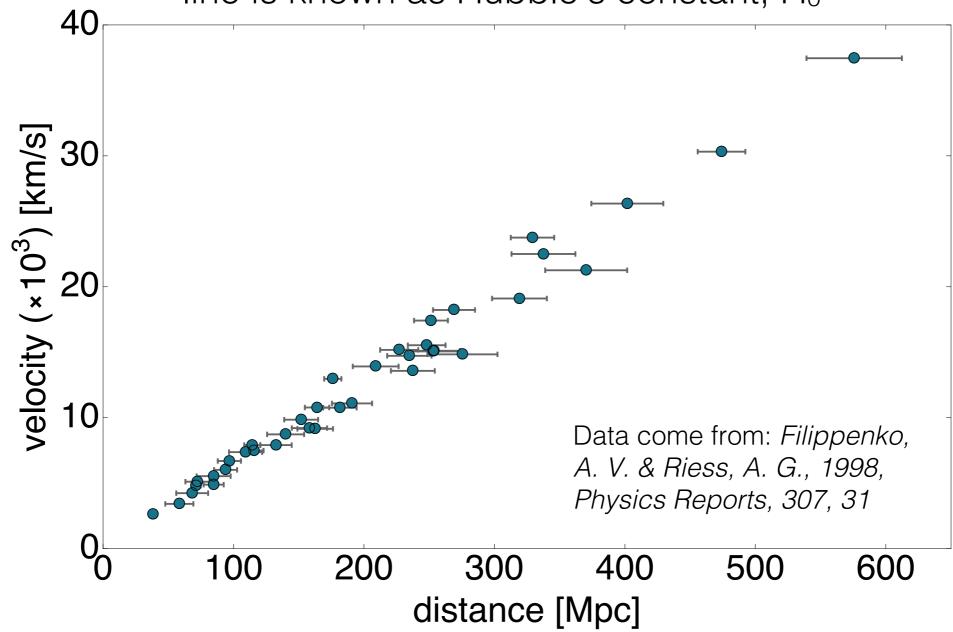


Hubble diagrams

- It is thought that all Type Ia supernova have the approximately same intrinsic luminosity
- That means the apparent brightness directly gives us the distance to that object (a "standard candle")
- The redshift of the spectral lines gives us the velocity
- Plotting velocity vs distance of distant objects is called a Hubble diagram, after Edwin Hubble who did this in 1929

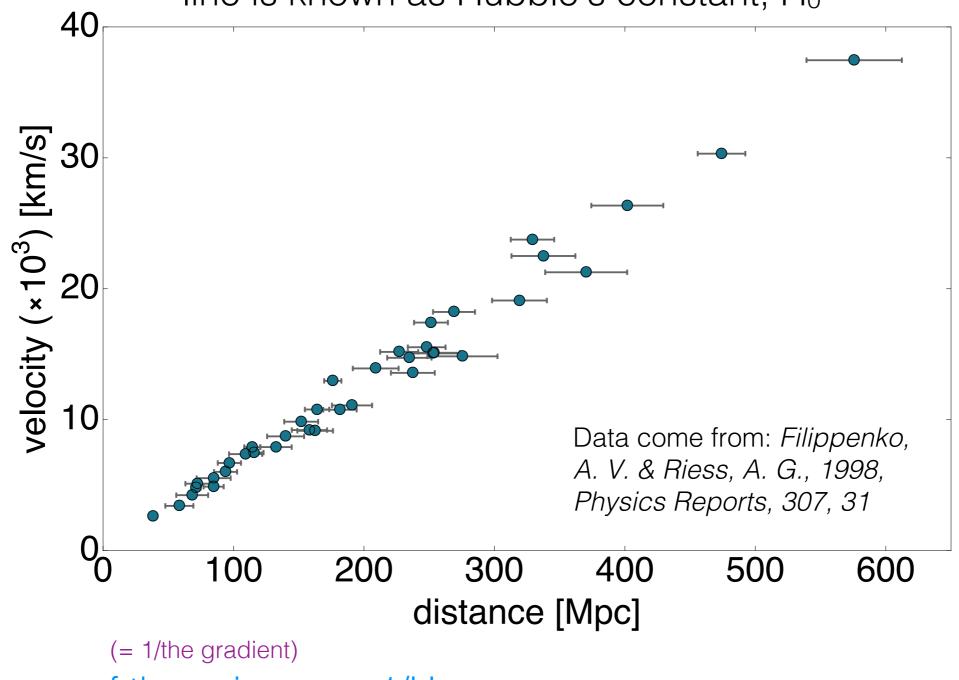


The gradient of the best-fitting straight line is known as Hubble's constant, H₀



Since H_0 = velocity/distance = units of time⁻¹ inverse time = a rate => the expansion rate of the Universe $1/rate = age of the universe = <math>1/H_0$

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age of the universe = $1/H_0$

YOUR TASK: Use this data set to calculate the age of the Universe

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1] Import data and re-arrange such that errors on are the y-axis

np.loadtxt np.genfromtxt astropy.ascii.read

check out CODING/ examples 2] Make a clear plot of the data, think about presentation

3] Do an unweighted fit of the data using numpy's polyfit np.polyfit

(3b] make a histogram of residuals/error bar- is it Gaussian?)

4] Write your own code for solving the unweighted normal equation and check that you get the same answer

SEM2

5] Use polyfit again, use weighting and also output the covariance matrix

SEM3

6] Diagonal of the covariance matrix gives you the errors, so you should now able to express age of the universe with an error

SEM3

7*] Try censoring some of the data and see how your derived uncertainty increases (anything from single example to plots of *factor by which error increases* vs *fraction of data censored*)

8*] Try perturbing each data point by random.normal(0,error_i) and make a "bootstrap" histogram of how best-fitting gradient changes over 1000 trials

9] Write up your summary slide see example summary slide

you don't need to do a "*", but grade is capped at A without