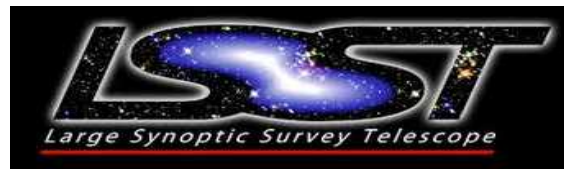


Nickel Mass Distribution of normal Type II Supernovae

Tomás Müller

Instituto de Astrofísica, PUC
Millennium Institute of Astrophysics (MAS)

LSST SN Workshop, University of Pittsburgh, Nov. 2016





Alejandro Clocchiatti
PUC

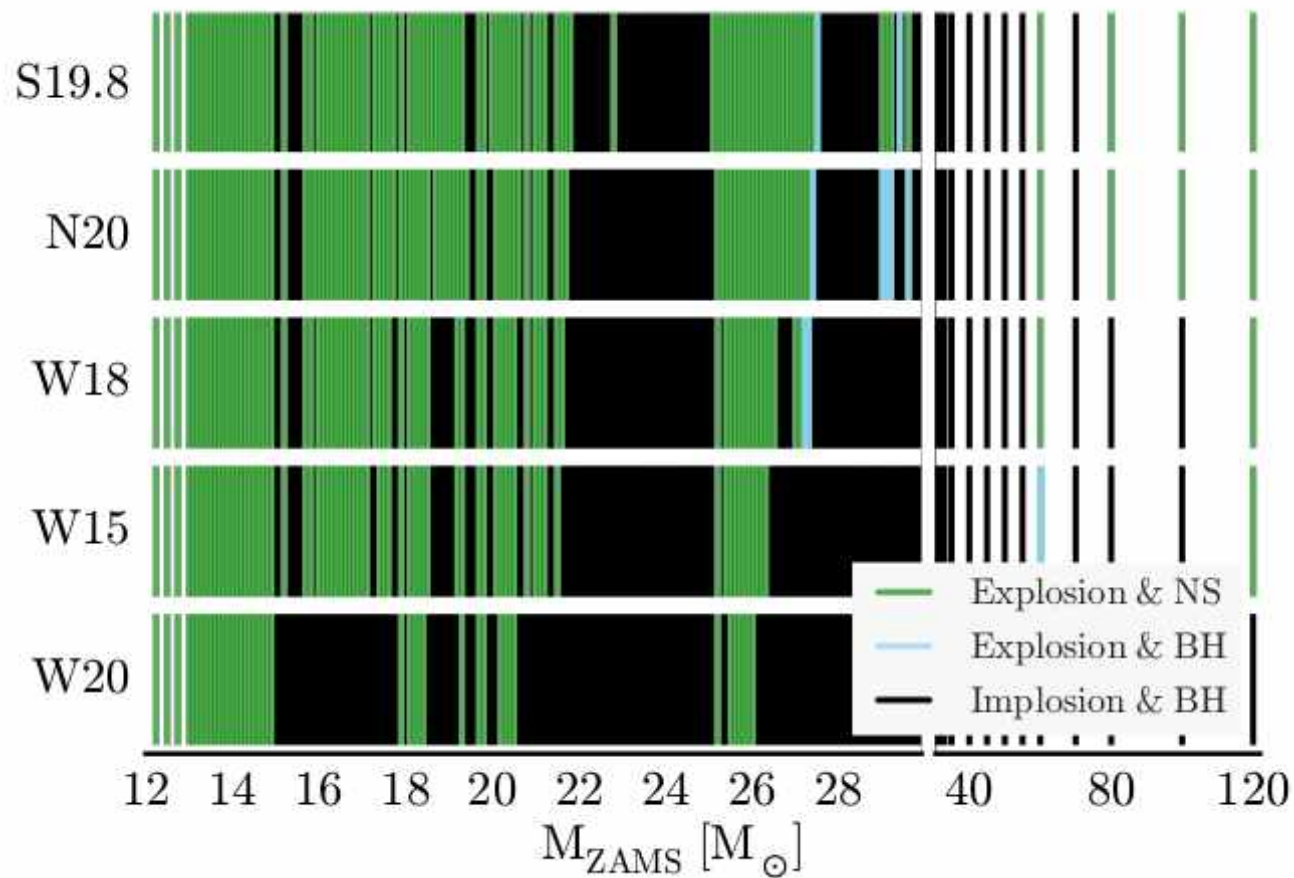


Ondrej Pejcha
Princeton University



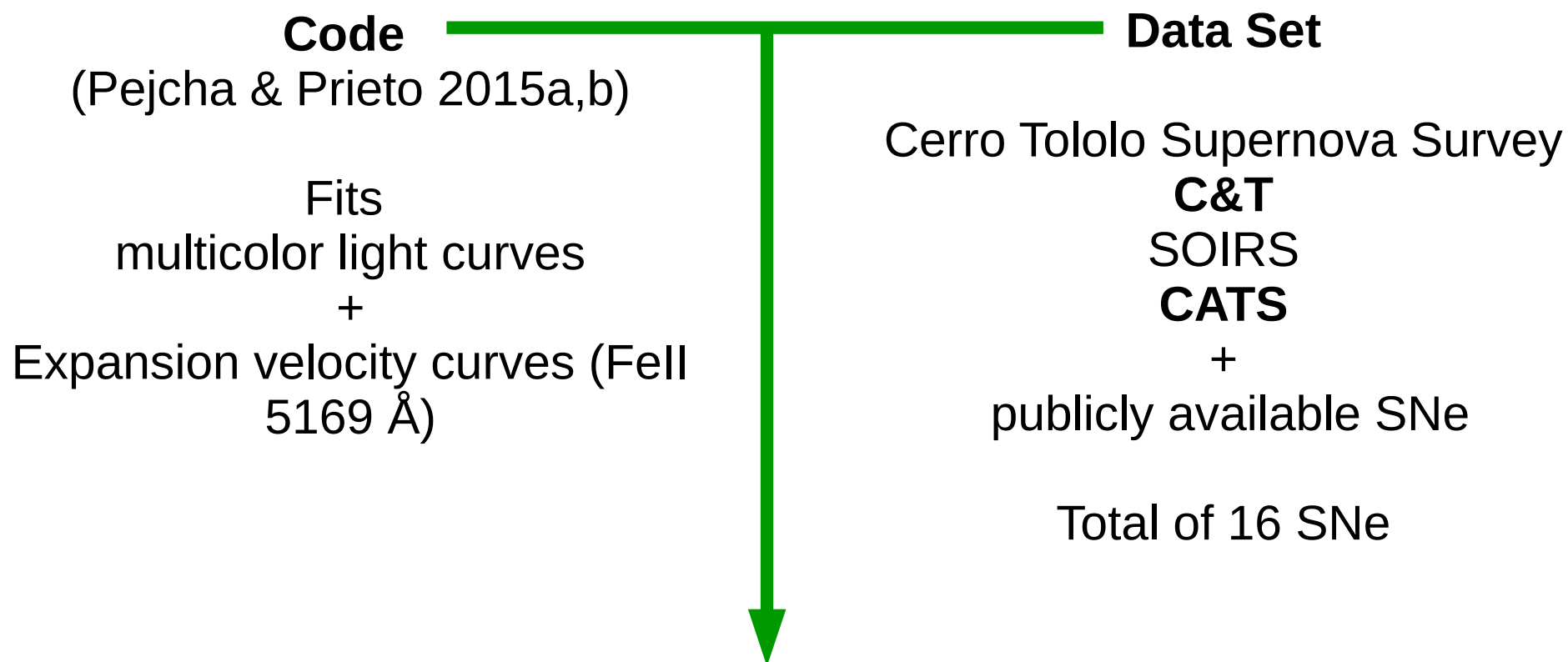
José Luis Prieto
UDP

Motivation



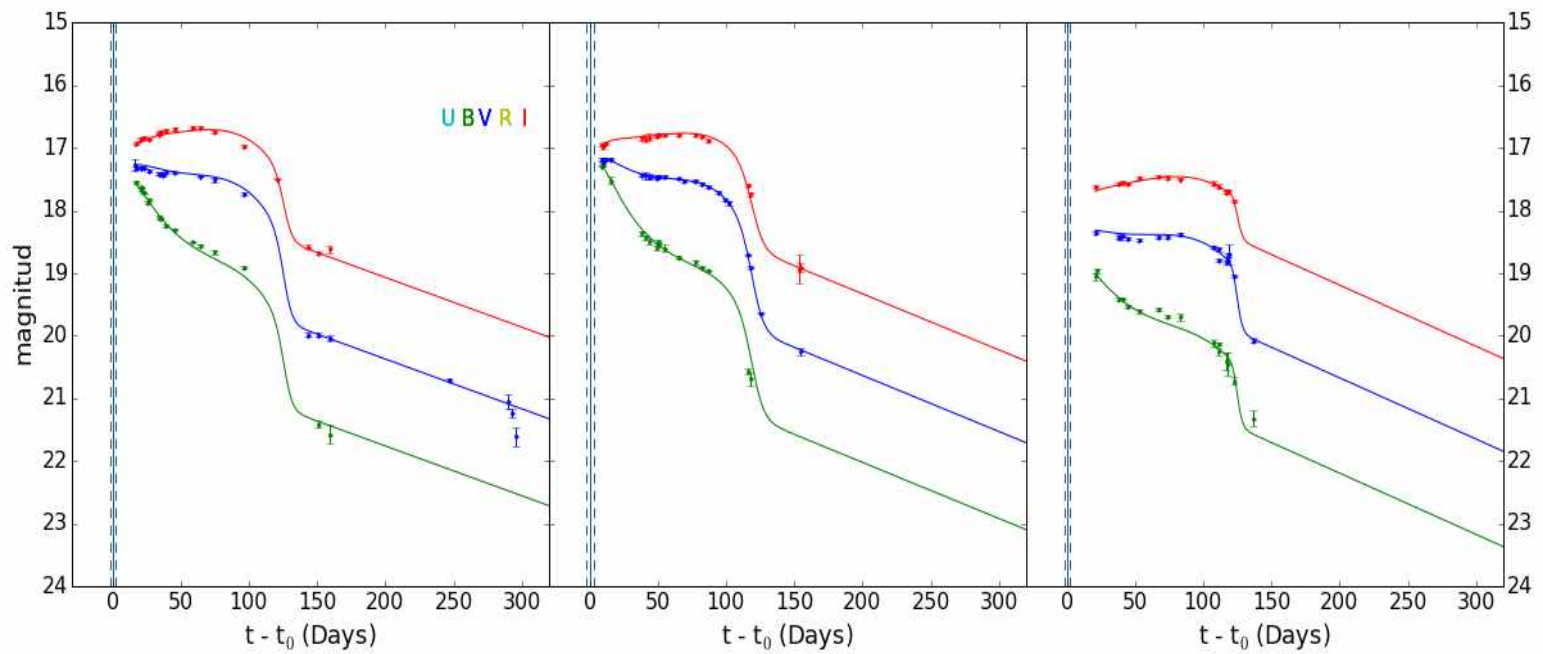
Sukhbold et al. 2016

SN Model + Data



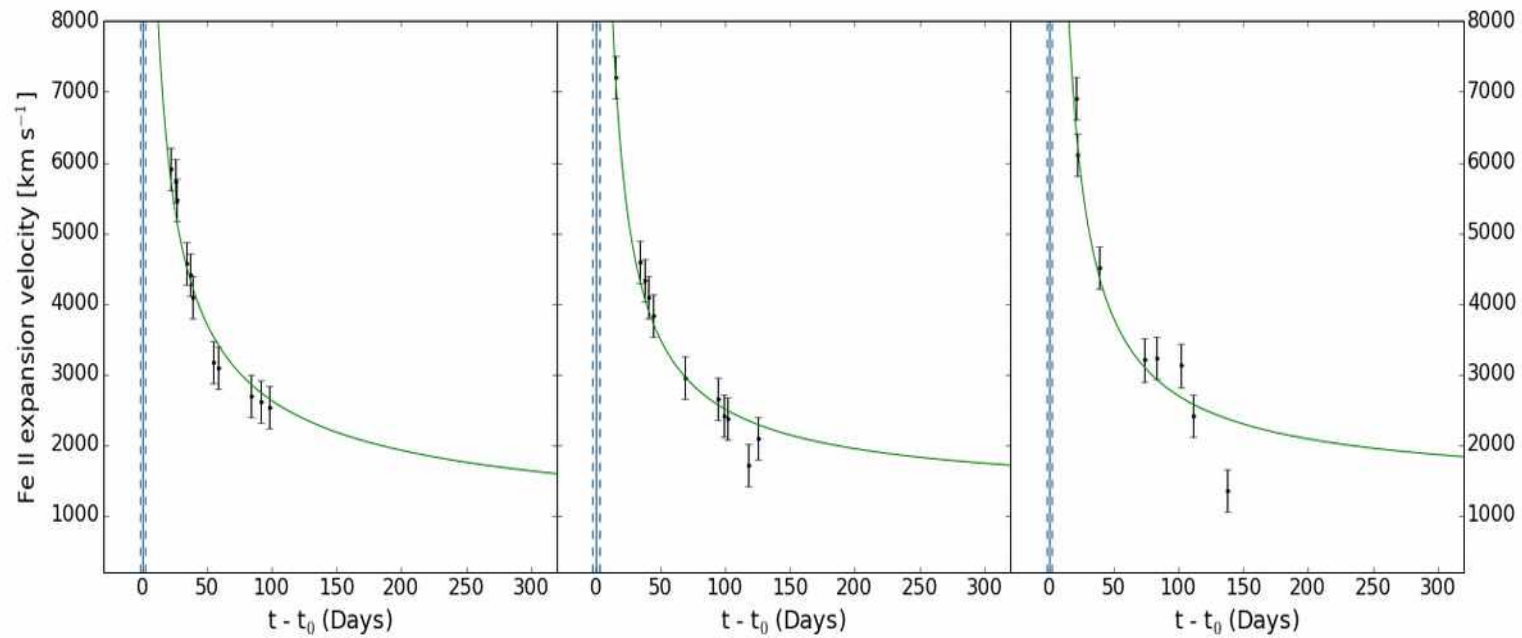
Results

SN2002gw SN2003bn SN2003hl

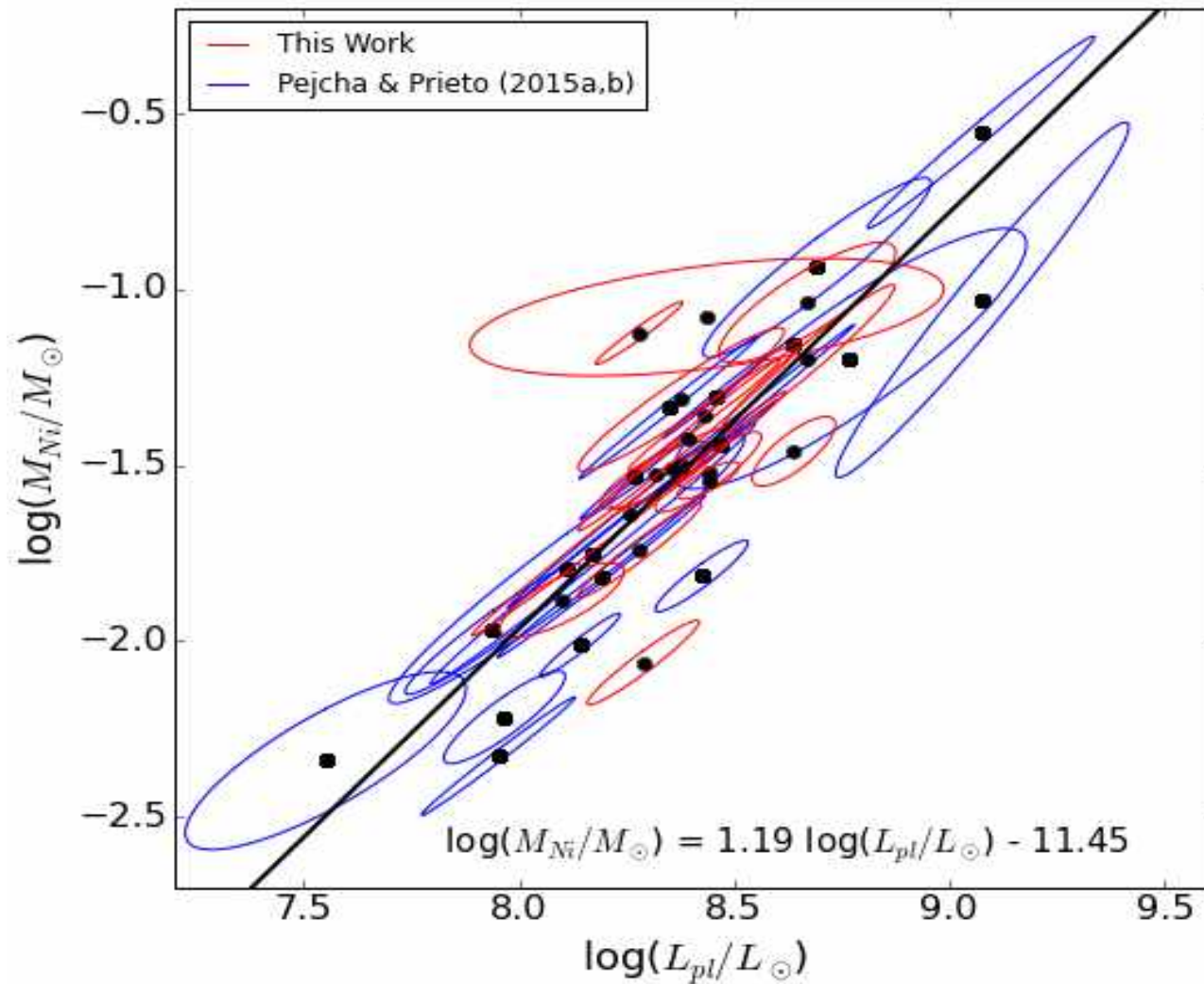


Results

SN2002gw SN2003bn SN2003hl

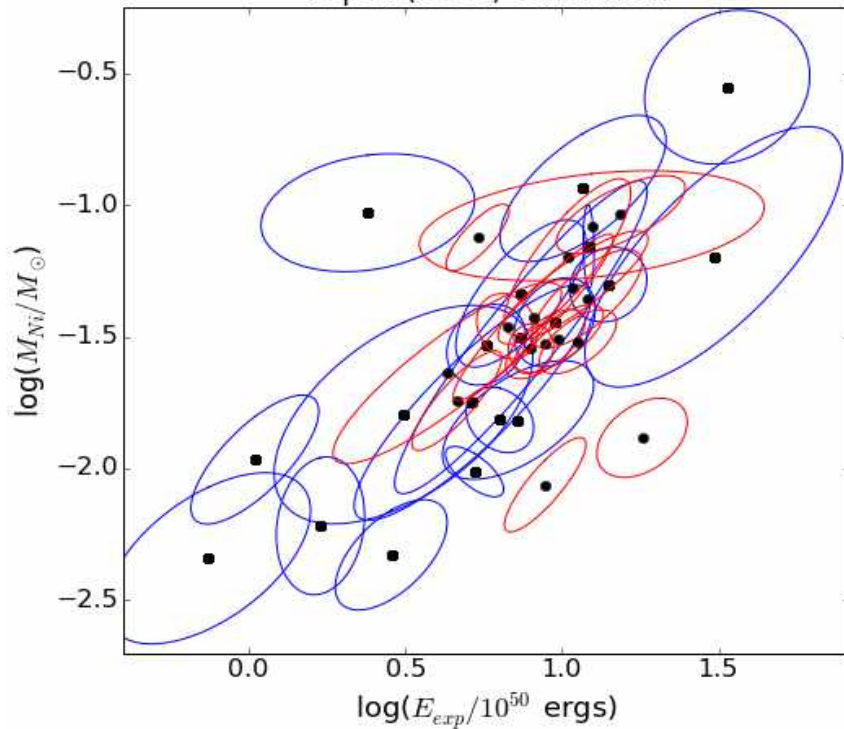


Derived Parameters

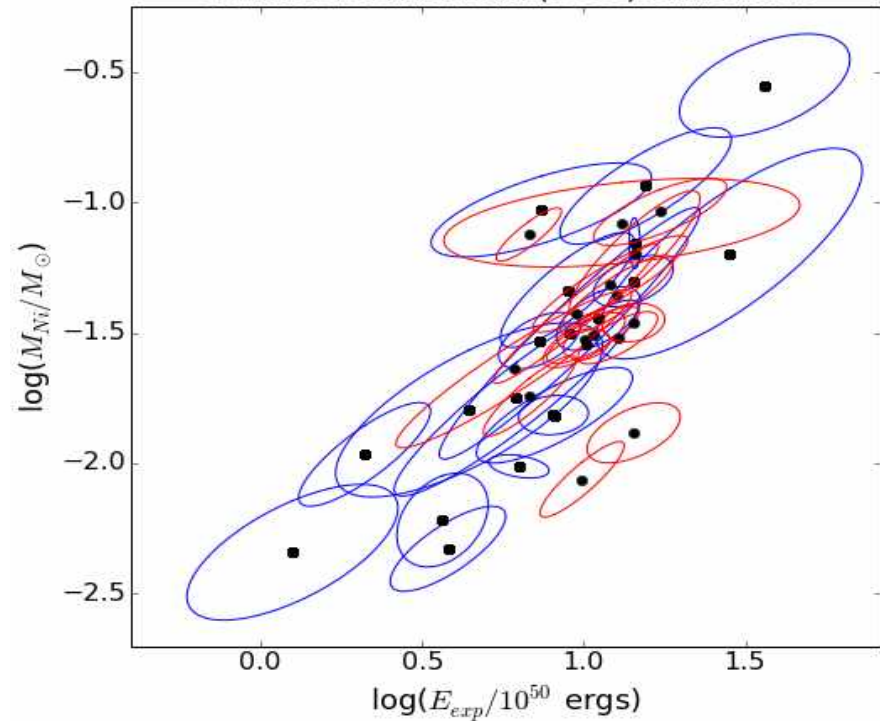


Derived Parameters

Popov (1993) calibration



Litvinova & Nadezhin (1985) calibration




Theoretical Nickel Mass Distribution

Sukhbold et al. (2016)
1D hydrodynamical models:

Prometheus Hot Bubble (P)

Kepler (K)

M_{prog}  M_{Ni}

$\text{IMF} * M_{\text{prog}}$  M_{Ni} Distribution

IMF lower limit: $9 M_{\odot}$

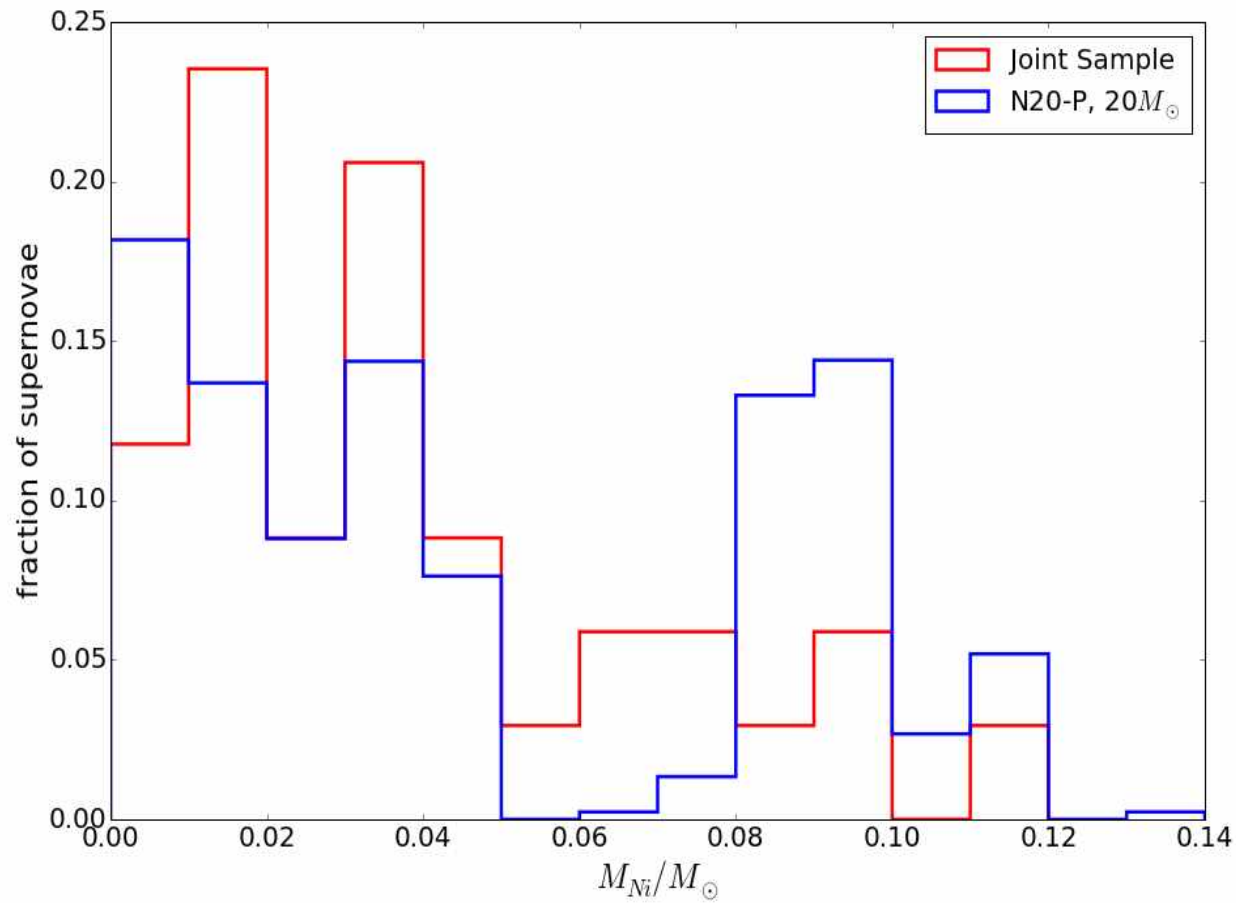
IMF upper limit: $15 - 30 M_{\odot}$

Neutrino mechanism

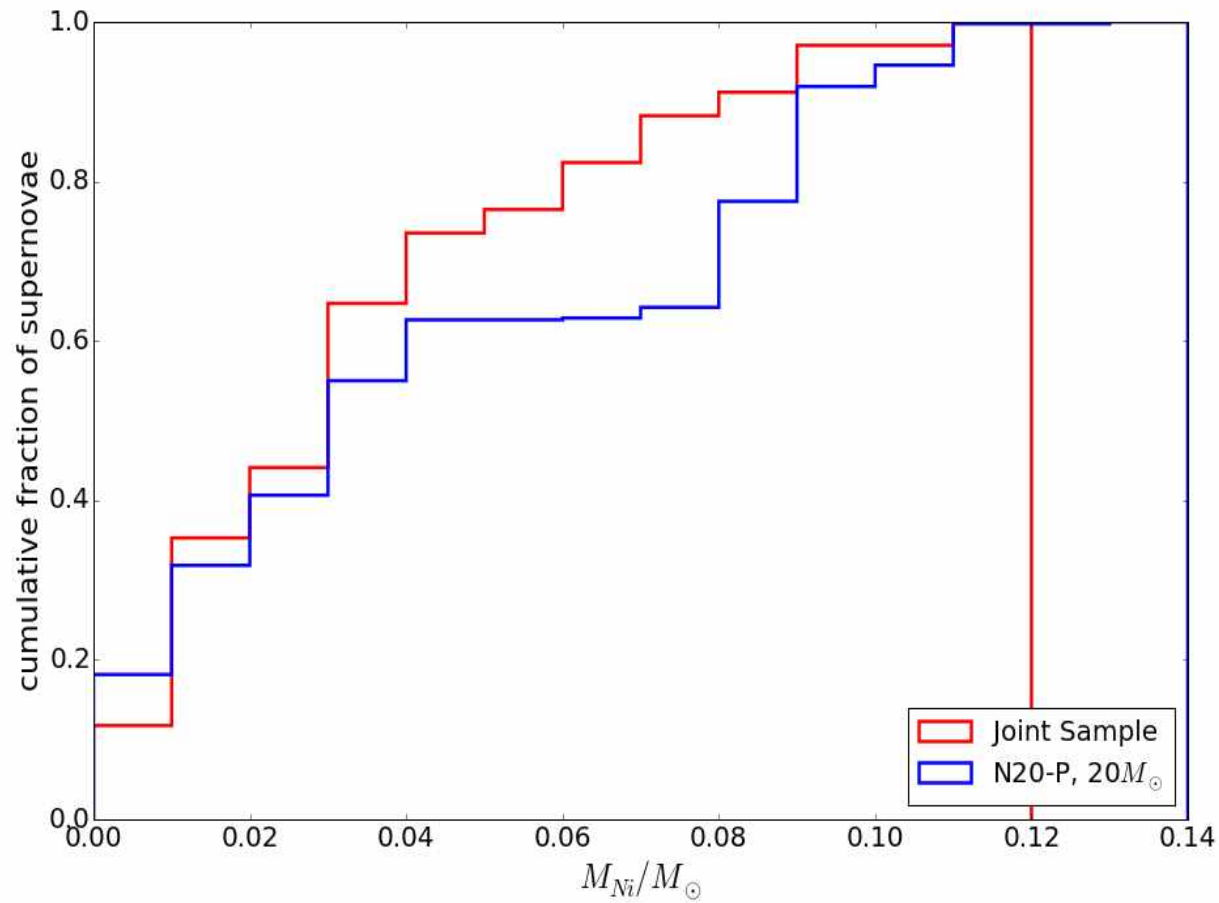
TABLE 8. EXPLOSION RESULTS FOR THE N20 AND W18 ENGINES

Progenitor	$M_{\text{Ni}} [M_{\odot}]$				$E_{\text{expl.}} [B]$				$M_{\text{remnant}} [M_{\odot}]$	
	N20		W18		N20		W18		N20	W18
	K. P.		K. P.		K. P.		K. P.		P.	P.
12.25	0.055	0.089	0.063	0.086	1.44	1.44	1.36	1.36	1.56	1.56
12.5	0.059	0.092	0.059	0.088	1.44	1.44	1.35	1.35	1.58	1.58
12.75	0.062	0.087	0.060	0.082	1.29	1.29	1.20	1.20	1.63	1.64
13.0	0.070	0.094	0.065	0.083	1.32	1.32	1.18	1.18	1.66	1.68
13.1	0.061	0.086	0.058	0.080	1.22	1.22	1.11	1.11	1.59	1.60
13.2	0.061	0.088	0.058	0.082	1.26	1.26	1.14	1.14	1.59	1.60
13.3	0.061	0.086	0.059	0.081	1.22	1.22	1.12	1.12	1.60	1.61
13.4	0.063	0.086	0.061	0.081	1.21	1.21	1.12	1.12	1.61	1.62
13.5	0.064	0.093	0.062	0.087	1.33	1.33	1.23	1.23	1.61	1.62
13.6	0.073	0.104	0.070	0.097	1.51	1.51	1.38	1.38	1.62	1.64
13.7	0.071	0.103	0.068	0.096	1.48	1.48	1.35	1.35	1.63	1.64
13.8	0.072	0.101	0.069	0.096	1.43	1.43	1.33	1.33	1.65	1.66
13.9	0.071	0.097	0.069	0.091	1.36	1.36	1.27	1.27	1.66	1.67
14.0	0.070	0.097	0.068	0.091	1.36	1.36	1.27	1.27	1.67	1.68
14.1	0.069	0.094	0.067	0.089	1.30	1.30	1.24	1.23	1.69	1.69
14.2	0.067	0.090	0.066	0.086	1.25	1.25	1.20	1.19	1.69	1.70
14.3	0.072	0.096	0.068	0.089	1.31	1.31	1.21	1.21	1.70	1.71
14.4	0.070	0.090	0.069	0.088	1.22	1.22	1.19	1.19	1.72	1.72
14.5	0.077	0.089	0.077	0.088	1.09	1.09	1.07	1.07	1.76	1.76
14.6	0.072	0.090	0.071	0.086	1.17	1.17	1.13	1.13	1.75	1.75
14.7	0.079	0.089	0.078	0.086	1.07	1.07	1.01	1.01	1.77	1.78
14.8	0.072	0.085	0.071	0.083	1.07	1.07	1.05	1.05	1.78	1.78
14.9	0.076	0.088	0.075	0.085	1.07	1.07	1.04	1.04	1.78	1.78
15.2	0.070	0.082	0.071	0.079	0.94	0.94	0.83	0.83	1.55	1.57
15.7	0.075	0.086	0.075	0.081	0.95	0.95	0.81	0.81	1.57	1.59
15.8	0.085	0.097	0.074	0.074	1.06	1.06	0.65	0.65	1.56	1.64
15.9	0.079	0.079			0.70	0.70			1.67	
16.0	0.094	0.110	0.075	0.079	1.26	1.26	0.78	0.78	1.54	1.62
16.1	0.075	0.084	0.078	0.087	0.89	0.89	0.97	0.97	1.59	1.59
16.2	0.095	0.111	0.076	0.081	1.23	1.23	0.79	0.79	1.55	1.62
16.3	0.097	0.113	0.078	0.083	1.23	1.23	0.80	0.80	1.55	1.62

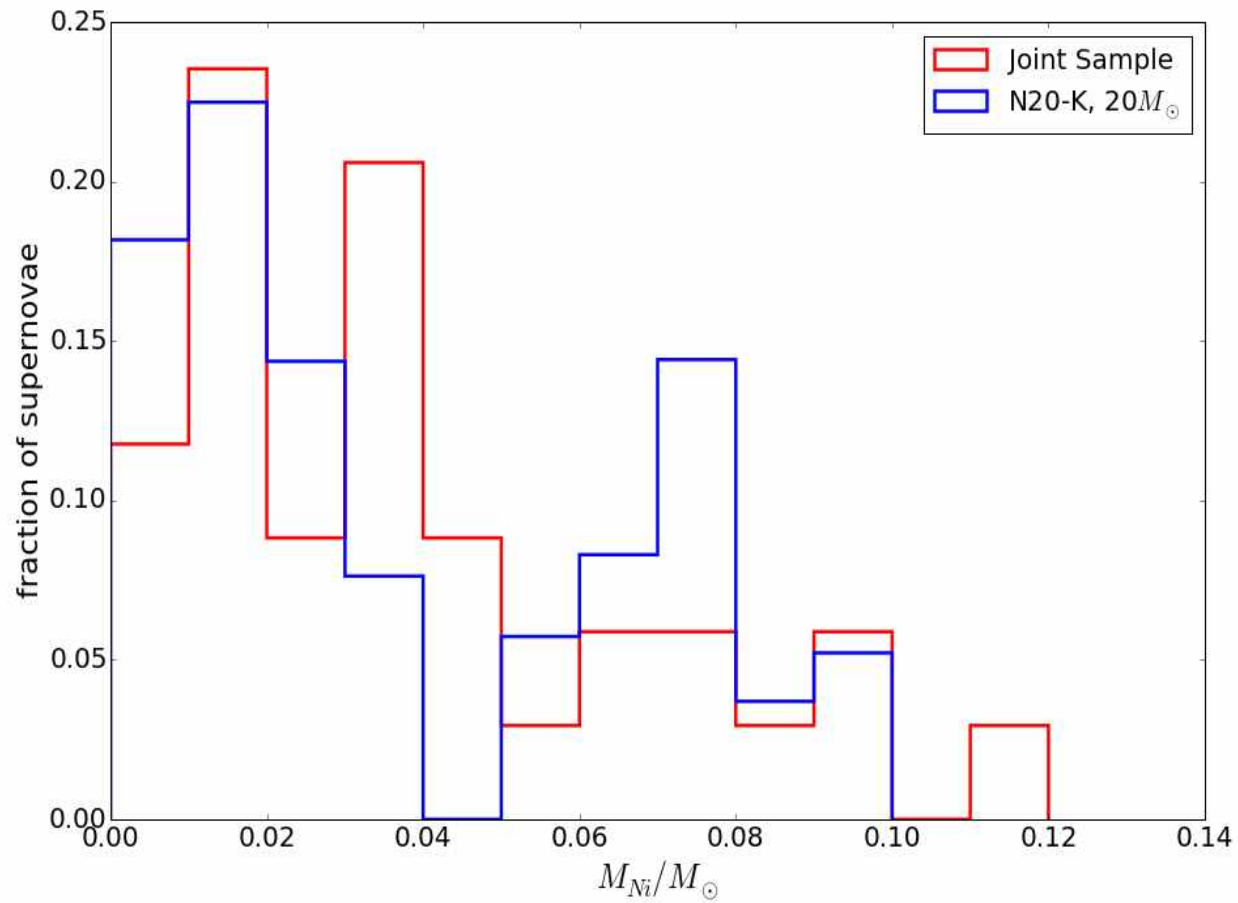
Nickel Mass Distributions



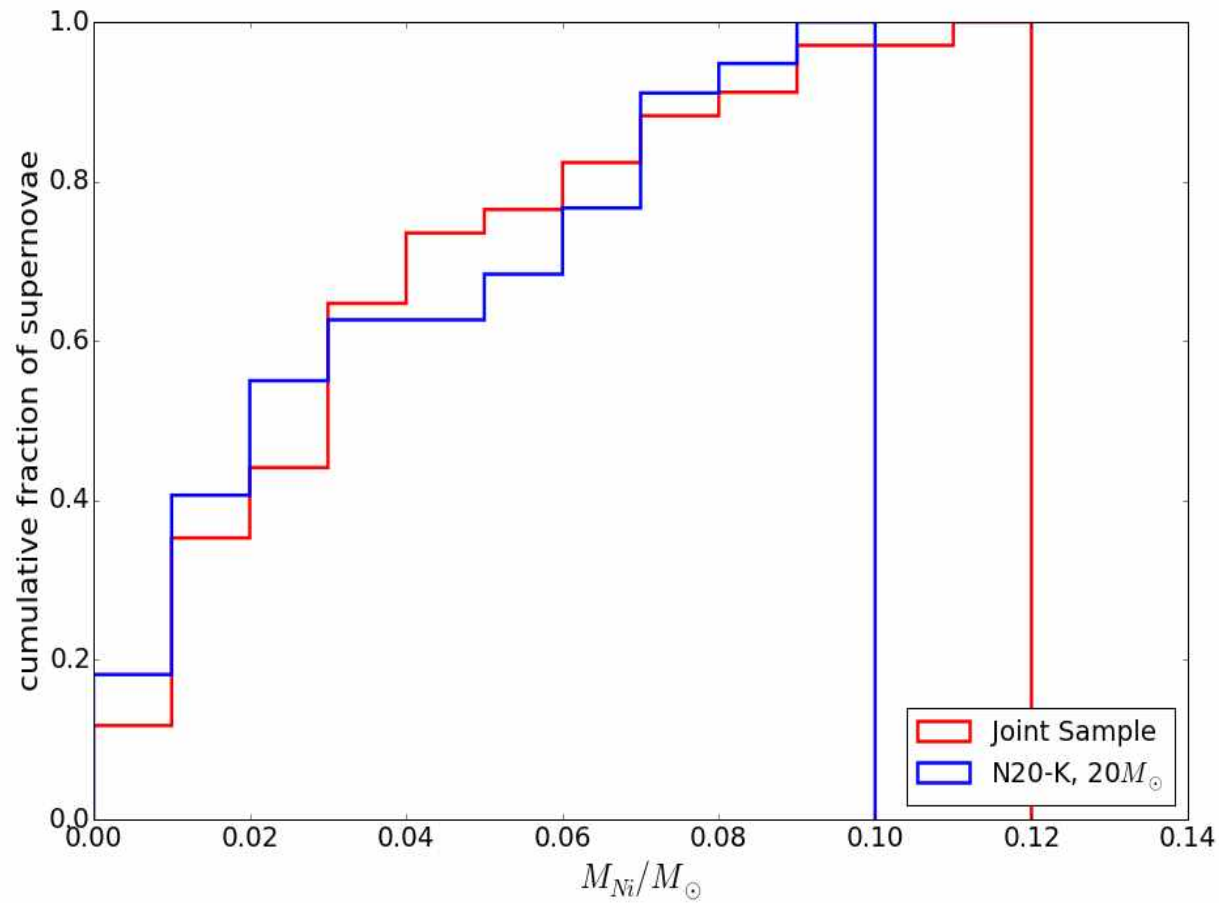
Nickel Mass Distributions



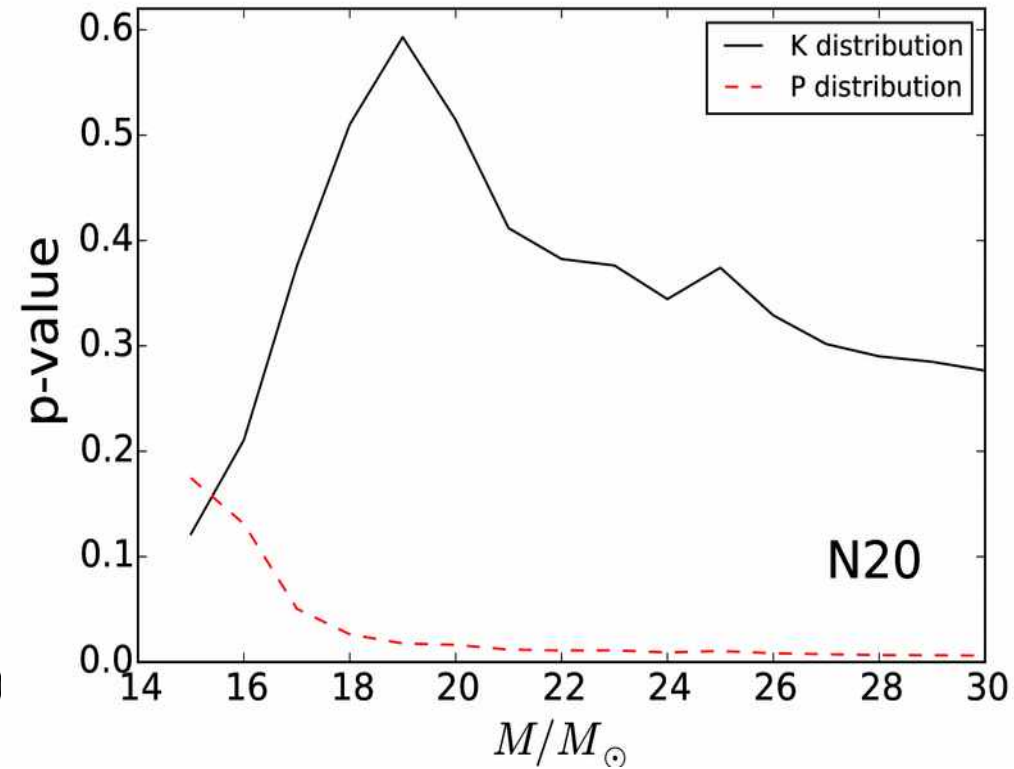
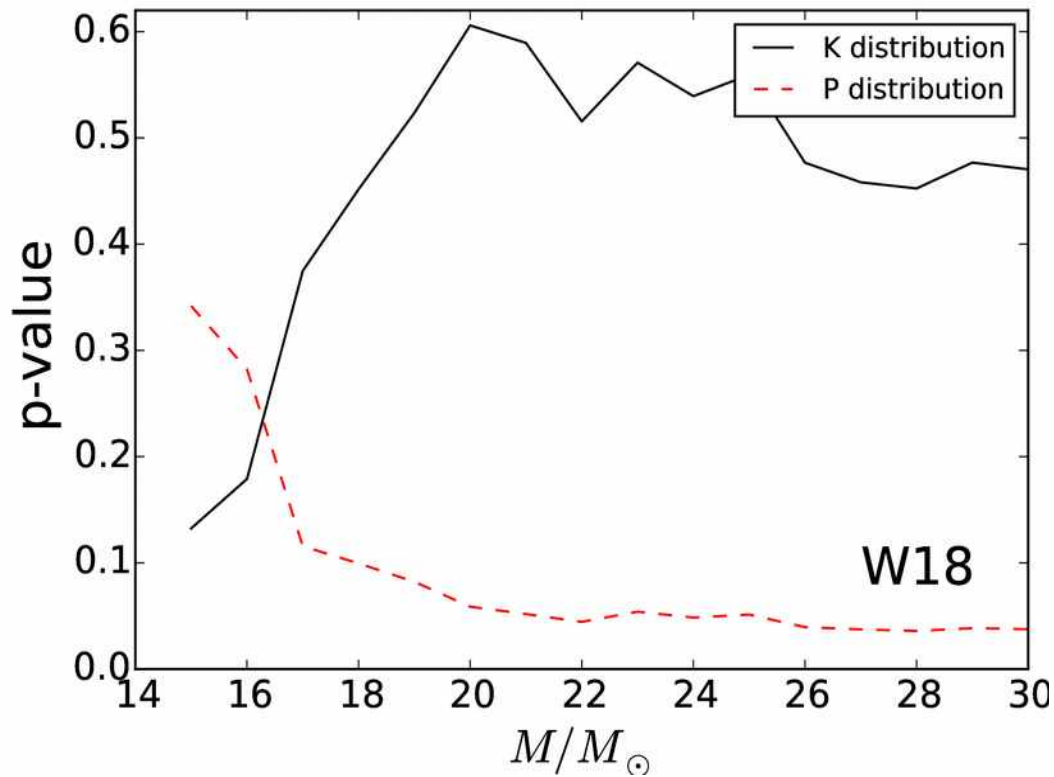
Nickel Mass Distributions



Nickel Mass Distributions



Comparing Models



Summary

- We retrieved known correlations between nickel mass and bolometric luminosity, and nickel mass and explosion energy
- Comparison of observations with neutrino mechanism models (nickel mass)
- The KEPLER model seems to adjust better the observations than the Prometheus Hot Bubble model
- We would like to shed more light into the conclusion by Pejcha & Thompson (2015) and Sukhbold et al. (2015) that there is no single mass below which all stars explode turning into a neutron star and above which black holes form, but rather there is a more complex behavior