

# OpSim

Michael Johnson

<https://store.docker.com/editions/community/docker-ce-desktop-mac>

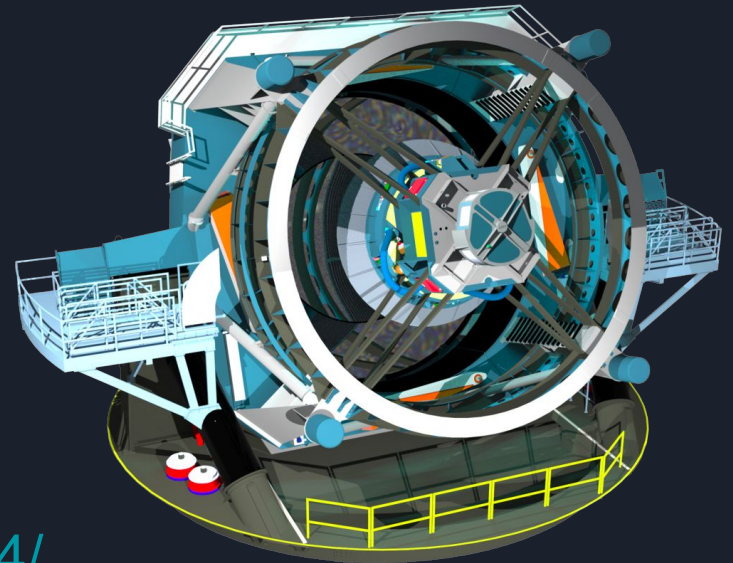
```
sudo apt-get install docker.io
```

```
sudo yum install docker.io
```

```
docker pull lsst/opsim
```

```
docker pull oboberg/opsim4:081217
```

<https://hub.docker.com/r/oboberg/opsim4/>



UNIVERSITY OF  
**Southampton**

**LSST**  
*Large Synoptic Survey Telescope*

**EPSRC**  
Engineering and Physical Sciences  
Research Council

# Introduction to Opsim

- Simulates image acquisition and field selection of LSST
- Each potential field is ranked by OpSim's algorithm
- Positive points for meeting science objectives
- Negative points for slewing, filter changes etc
- Highest ranking field is chosen and observed
- Produces a record of all observations made by LSST over the survey lifetime
- Designed in part to verify whether LSST could meet the scientific requirements



# Installation - OpSim 3.38

## LSST Stack Install

- Requires a list of dependencies to be met

## Conda Installation

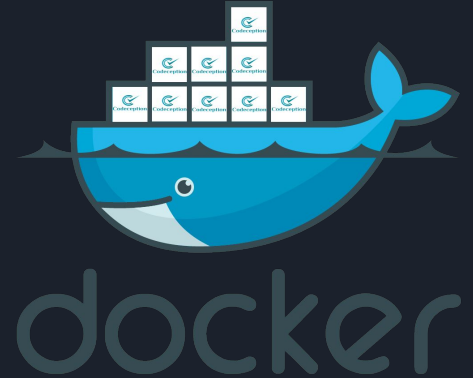
- Local installations of python can disrupt the path

## Docker Installation

- Requires a local installation of Docker

More info: <http://ops2.lsst.org/docs/current/>

# Docker



Designed to easily ship software

Inspired by shipping containers

Each container is packed independently but handled identically once at the ship

Makes an image of the host machine with a working copy of the software

Removes the requirement to install dependencies

More lightweight and flexible than virtual machines

# Docker OpSim Installation

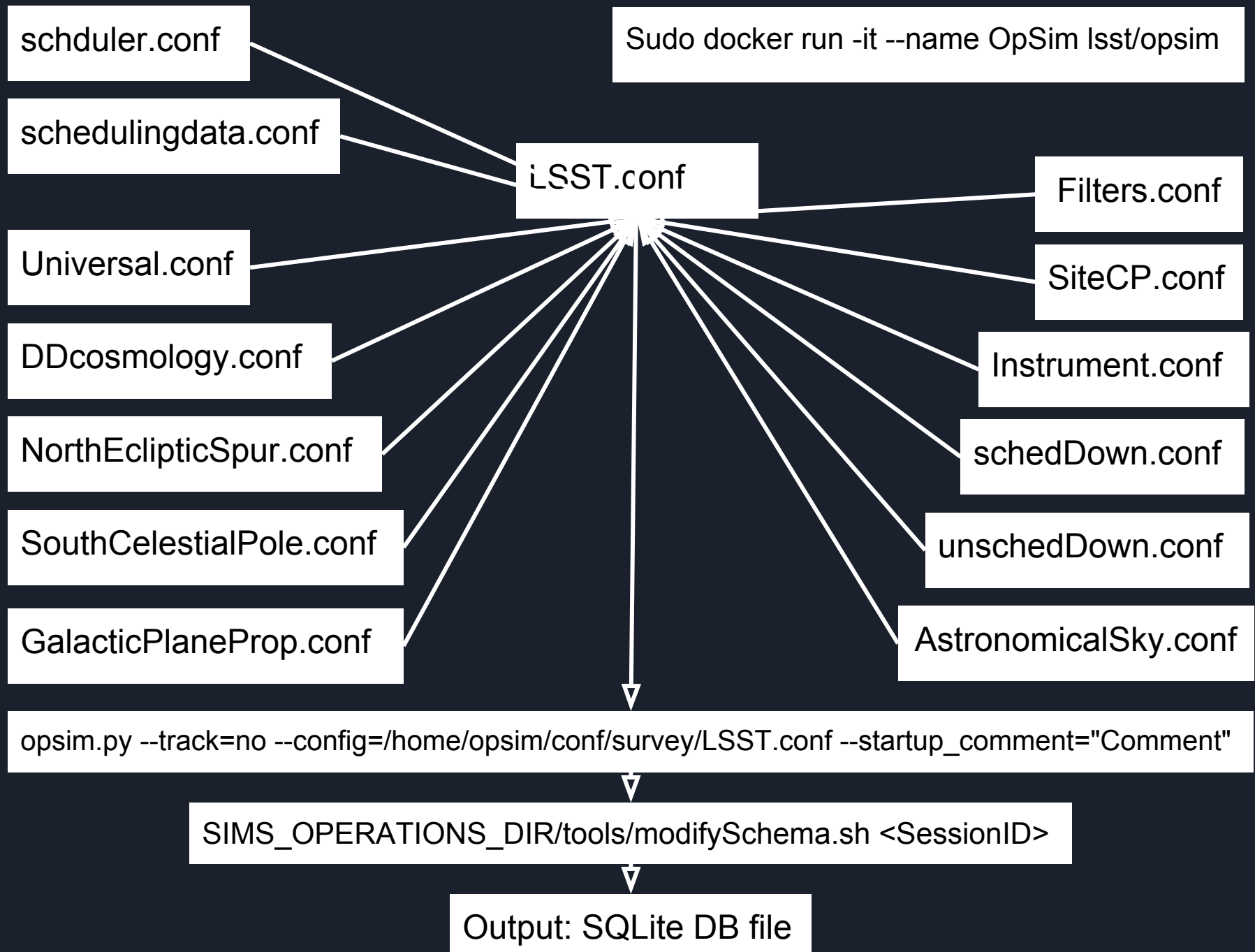
## OpSim Installation:

```
docker pull lsst/opsim
```

More info: <https://hub.docker.com/r/lsst/opsim/>

## Open an OpSim Instance:

```
docker run -it --name OpSim lsst/opsim
```

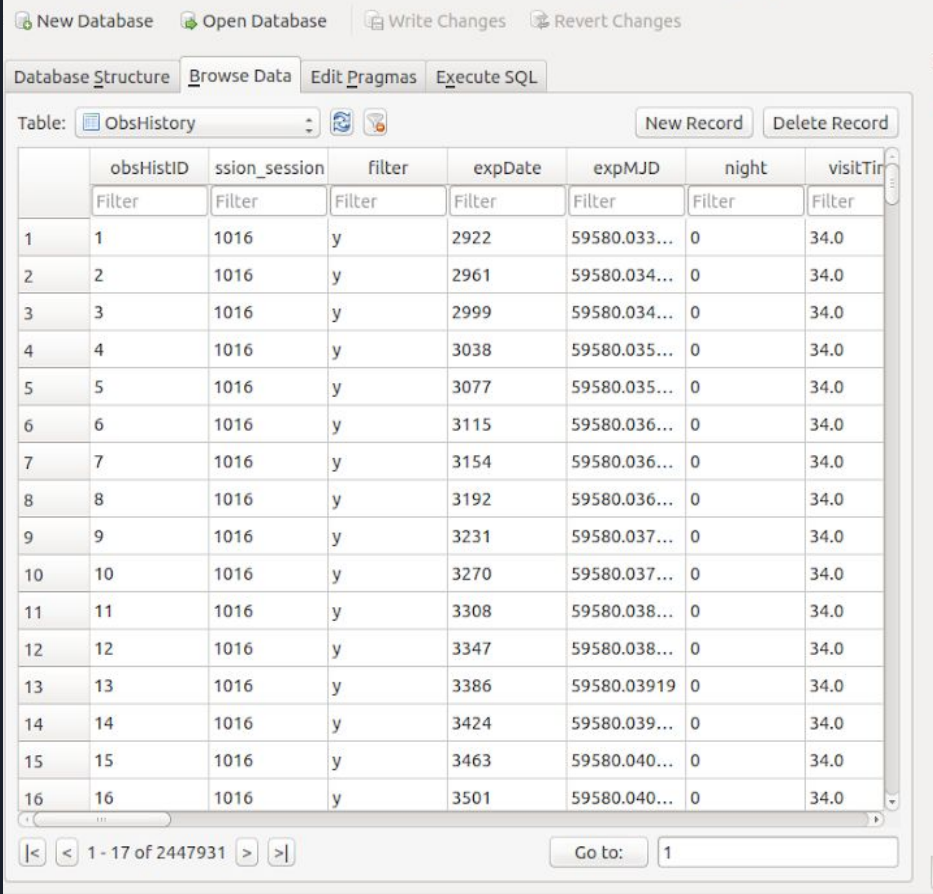


# Opsim Output Data

Relational database containing every observation LSST made during the 10 year survey

SQLite DB file

Analyse with SIMS MAF or SQL (etc)



The screenshot shows a database browser window with the following elements:

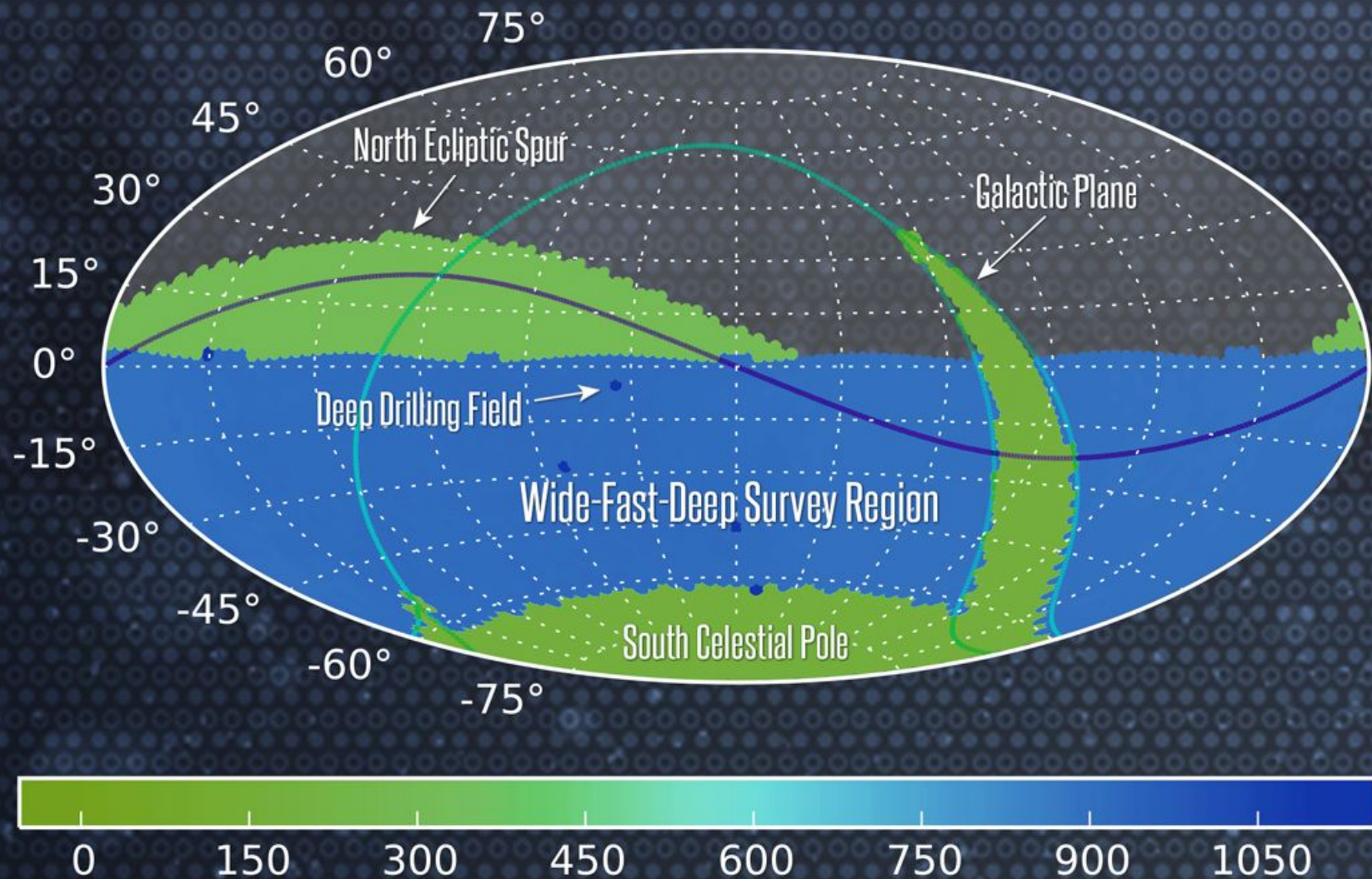
- Buttons at the top: New Database, Open Database, Write Changes, Revert Changes.
- Navigation tabs: Database Structure, Browse Data (selected), Edit Pragmas, Execute SQL.
- Table selection: Table: ObsHistory.
- Buttons: New Record, Delete Record.
- Table columns: obsHistID, ssion\_session, filter, expDate, expMJD, night, visitTir.
- Table content: 16 rows of data.
- Footer: Navigation controls (back, forward, search) and a 'Go to:' field with the value '1'.

	obsHistID	ssion_session	filter	expDate	expMJD	night	visitTir
	Filter	Filter	Filter	Filter	Filter	Filter	Filter
1	1	1016	y	2922	59580.033...	0	34.0
2	2	1016	y	2961	59580.034...	0	34.0
3	3	1016	y	2999	59580.034...	0	34.0
4	4	1016	y	3038	59580.035...	0	34.0
5	5	1016	y	3077	59580.035...	0	34.0
6	6	1016	y	3115	59580.036...	0	34.0
7	7	1016	y	3154	59580.036...	0	34.0
8	8	1016	y	3192	59580.036...	0	34.0
9	9	1016	y	3231	59580.037...	0	34.0
10	10	1016	y	3270	59580.037...	0	34.0
11	11	1016	y	3308	59580.038...	0	34.0
12	12	1016	y	3347	59580.038...	0	34.0
13	13	1016	y	3386	59580.03919	0	34.0
14	14	1016	y	3424	59580.039...	0	34.0
15	15	1016	y	3463	59580.040...	0	34.0
16	16	1016	y	3501	59580.040...	0	34.0



# Number of Visits

(all-band, 10 years)

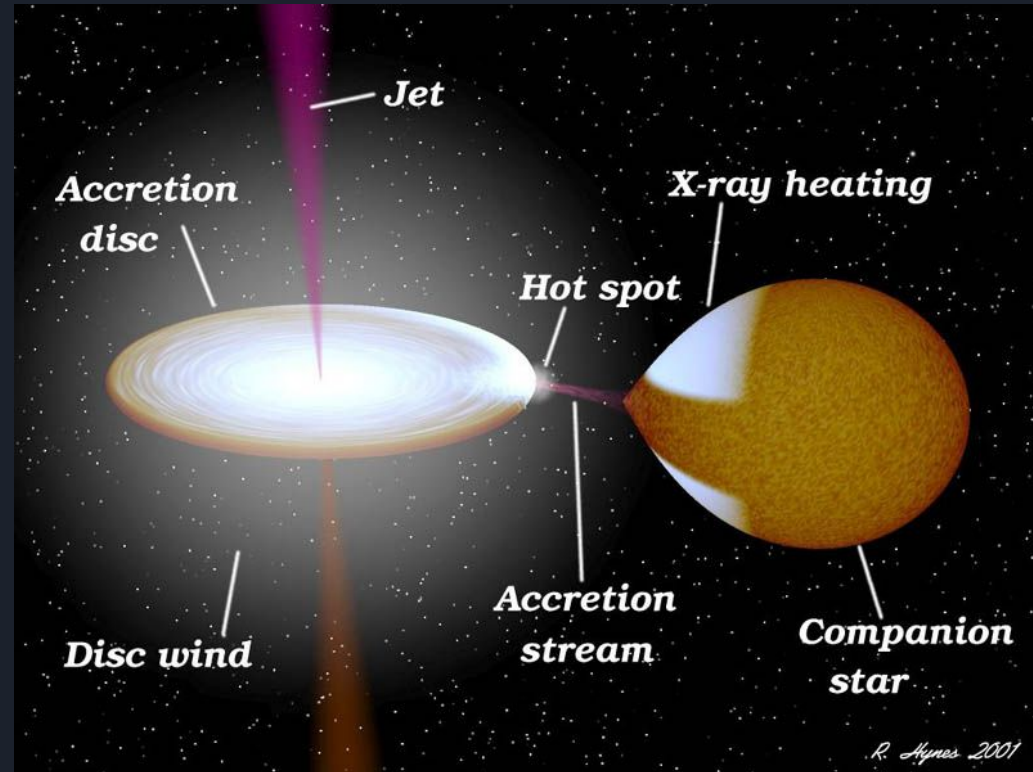


# Low Mass X-ray Binaries

Low mass companion star ( $<1M_{\odot}$ ) in orbit around a black hole or neutron star

Compact object accretes matter via Roche Lobe overflow, forming an accretion disc

Many LMXBs exhibit very bright, rare X-ray outbursts separated by decades of quiescence



Credit: NASA/R. Hynes



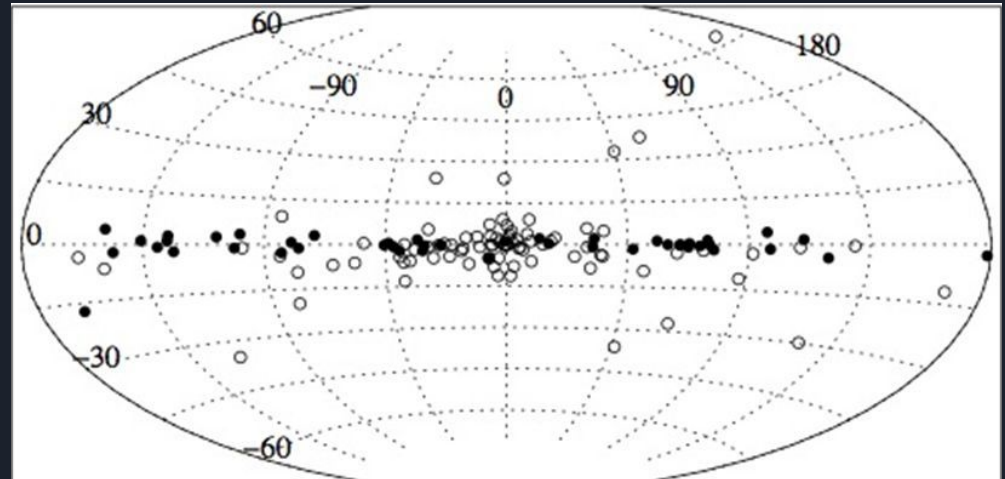
# LMXB Population

Estimated to be thousands of LMXBs within the Milky Way however, there are only ~200 known systems.

~60 of which are thought to contain a black hole

~20 dynamically confirmed black hole masses

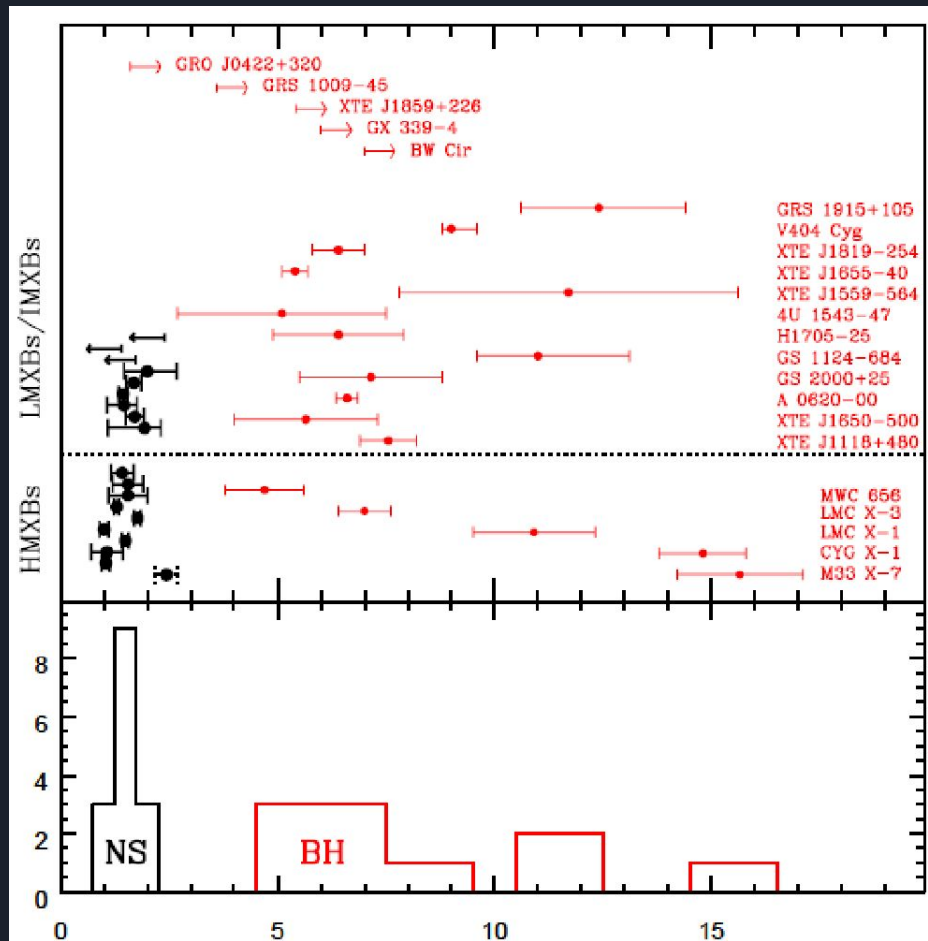
~90% of known systems reside in the Galactic Plane



Grimm et al. 2002

Extinction from Galactic gas and dust means that the vast majority are too faint to observe routinely in the optical.

# NS/BH mass Distribution



Casares et al. 2017

Observations of the periodic variability of LMXBs can be used in conjunction with spectral information to determine the masses of the binary components.

21 realistic masses of BHs  
5-16  $M_{\odot}$

Typical errors 30%

Goals: improve statistics and  
reduce errors to 10%

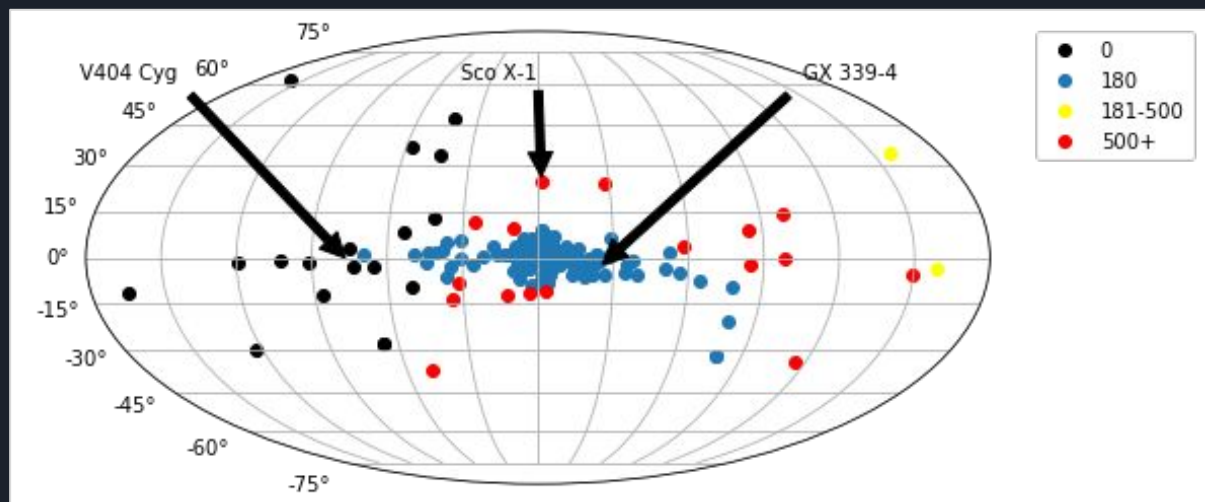
# The Baseline Strategy

## Minion\_1016

Location of 180 LMXBs referenced with the results of the current baseline strategy, Minion\_1016

Each dot represents a field that contains at least one LMXB

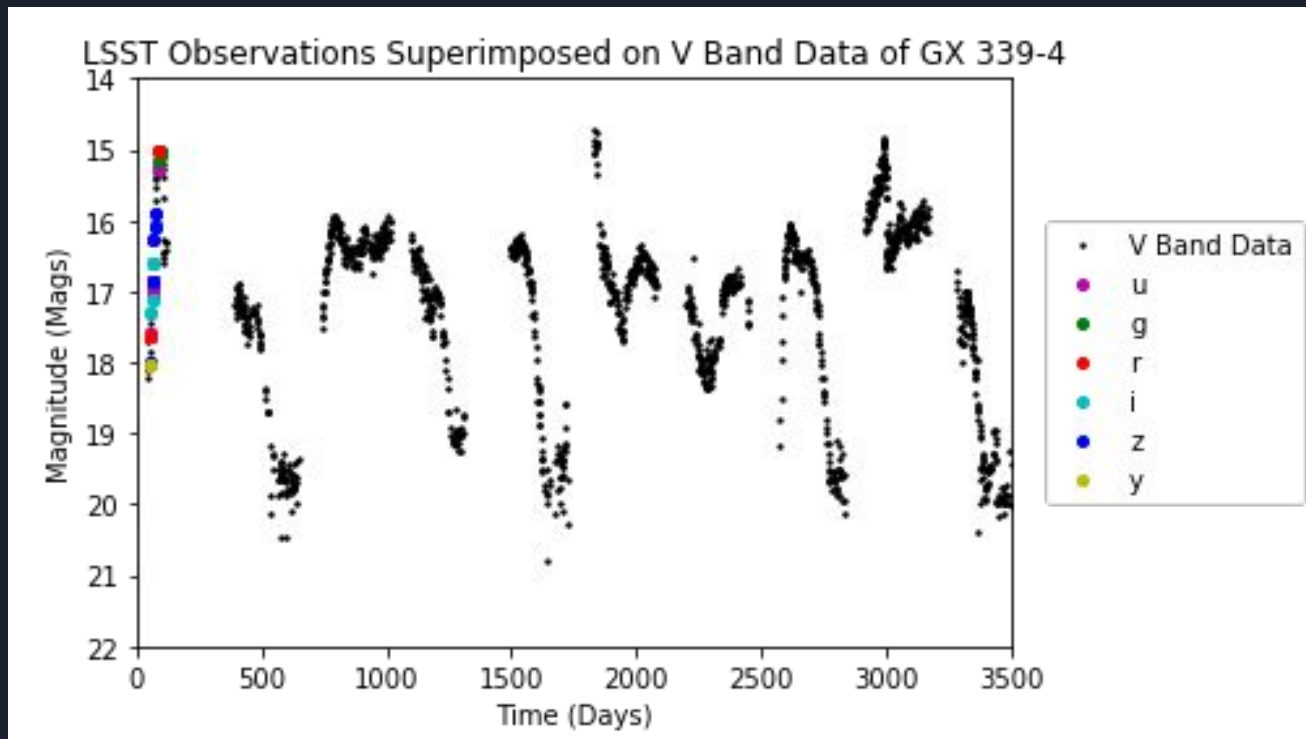
Numbers represent the total observations made in all filters over the full 10 year survey



# Galactic Plane Observations in Minion\_1016

Observations taken throughout the survey of the region containing GX 339-4 were superimposed onto the optical SMARTS data

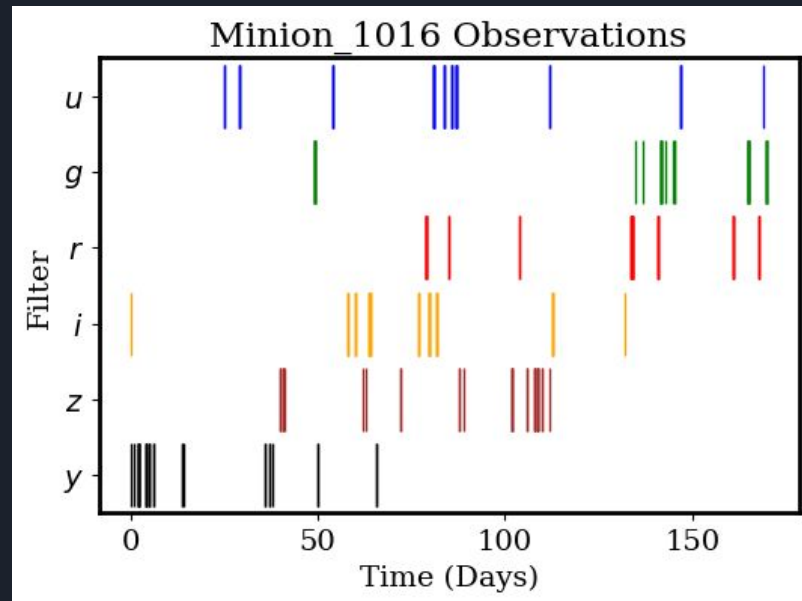
All Galactic Plane observations are taken during the first 7-10 months



# Galactic Plane Observations in Minion\_1016

Each line represents an observation of GX 339-4 taken by LSST, colour denotes the filter

The observations are very clustered with it being common for  $\frac{1}{6}$  of all observations per filter per field within the 10 year survey to occur within 2 hours



# Alternative Strategies

Minion\_1016 - baseline strategy

Minion\_1016j - baseline strategy with random time jitter between  $\pm 1$  day for each observation

Minion\_1020 - observations distributed evenly amongst all fields

astro\_lsst\_01\_1004 - baseline strategy with Galactic Plane included in the main survey region



# LMXB light curves

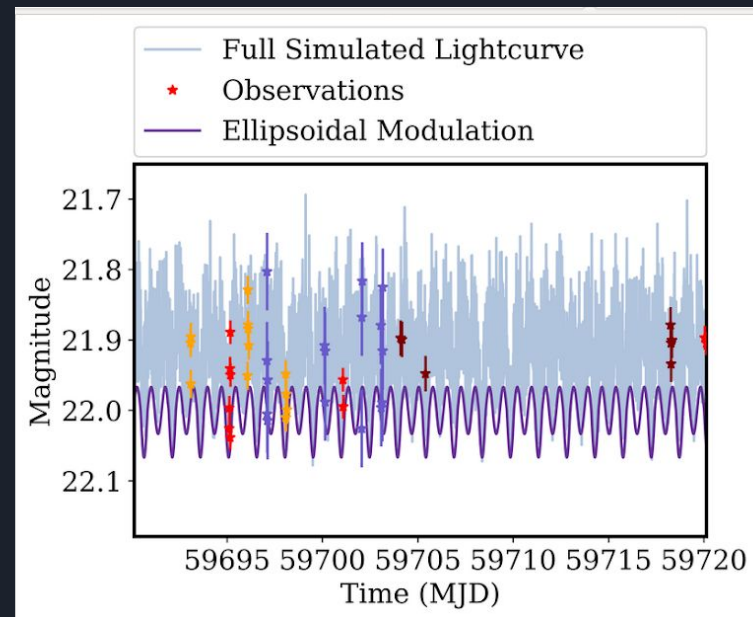
10 year simulated light curves composed of a base magnitude corresponding to the donor star, plus optical variation.

Light curves sampled by LSST observations (shown as \*s)

Optically variability composed of two parts:

Ellipsoidal Modulation - due to orbital motion

Flaring - red noise thought to be due to instabilities in the accretion flow



Johnson et al. Submitted

# Period Recovery - Method

A Multi-band Lomb-Scargle algorithm was used to recover the periods (VanderPlas and Ivezić 2015)

To determine the significance:

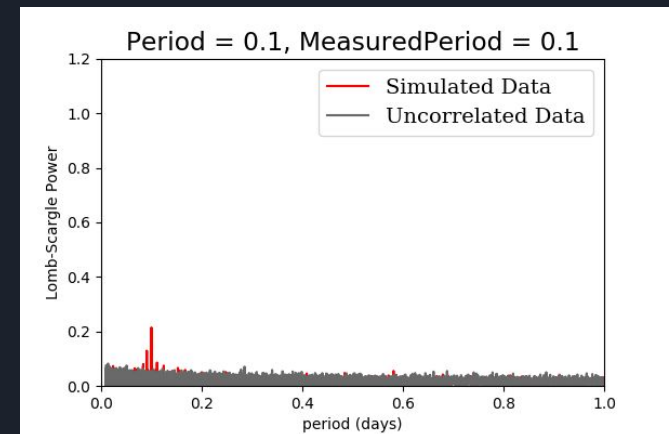
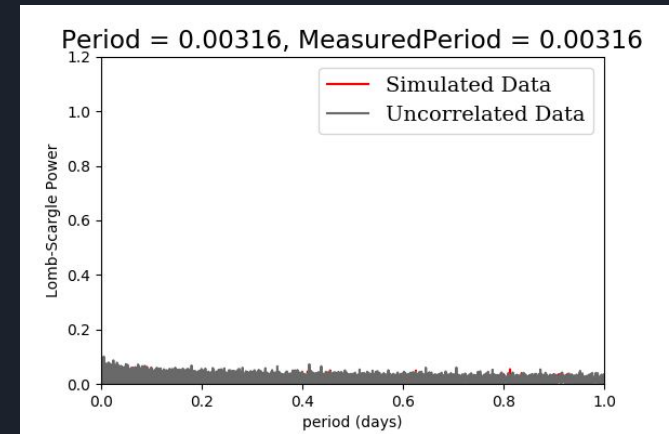
Magnitudes and dates in the light curve were shuffled, creating an uncorrelated data set

LS algorithm was evaluated over uncorrelated data set and the maximum peak recorded

This was repeated 10,000 times

$x$  = no. of times the highest peak in the correlated data set was greater than that of the uncorrelated set

$$\text{Sig} = x / 10,000$$



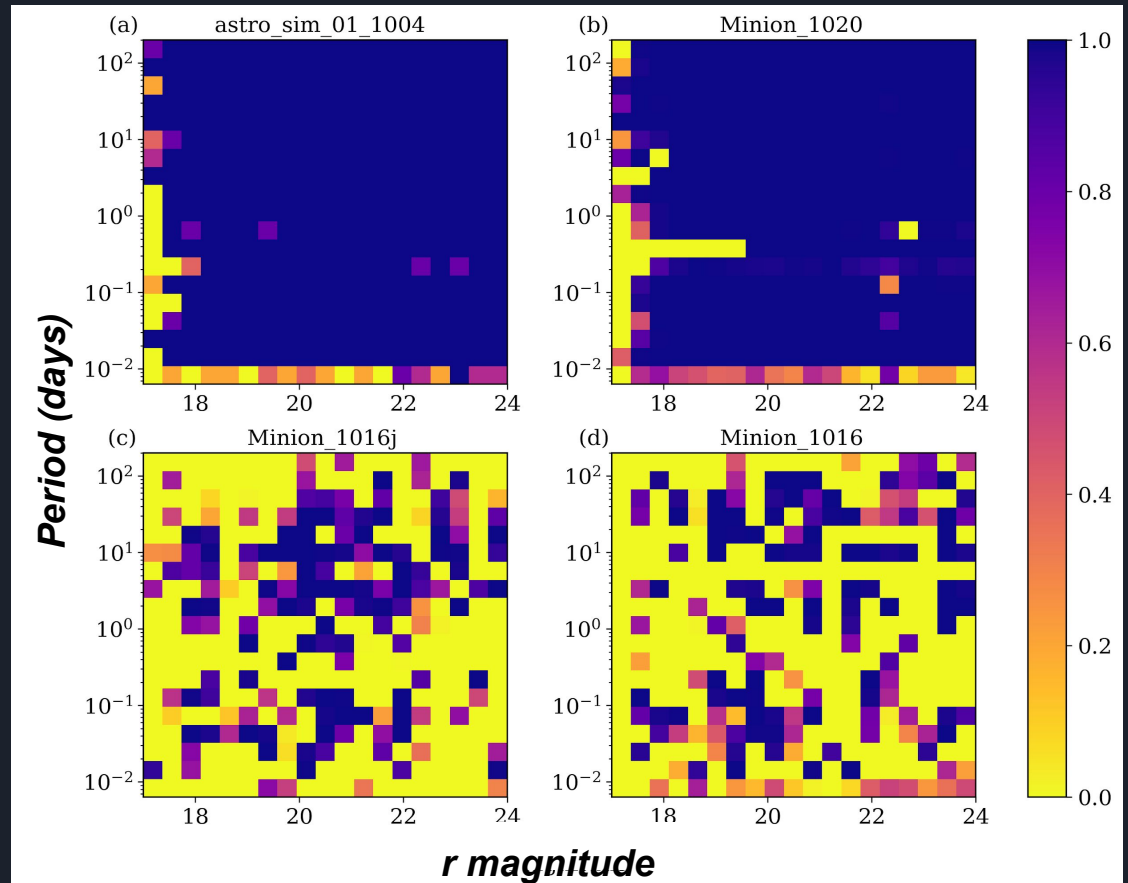
# Period Recovery - Results

The multi-band Lomb-Scargle algorithm was applied to the results from the sampled lightcurve to determine what periods could be recovered.

Period ranges from  $\sim 9$  minutes to 200 days to encompass all known LMXBs

Magnitude range from 17-24 mags in  $r$ , motivated by the visible range for LSST

Colour denotes significance



# Summary

OpSim produces a relational database including every observations made by LSST throughout the 10 year survey

OpSim was used to quantise the period recovery of different observing strategies, demonstrating a factor of  $\sim 2$  increase when not including a reduced cadence on the Galactic Plane