

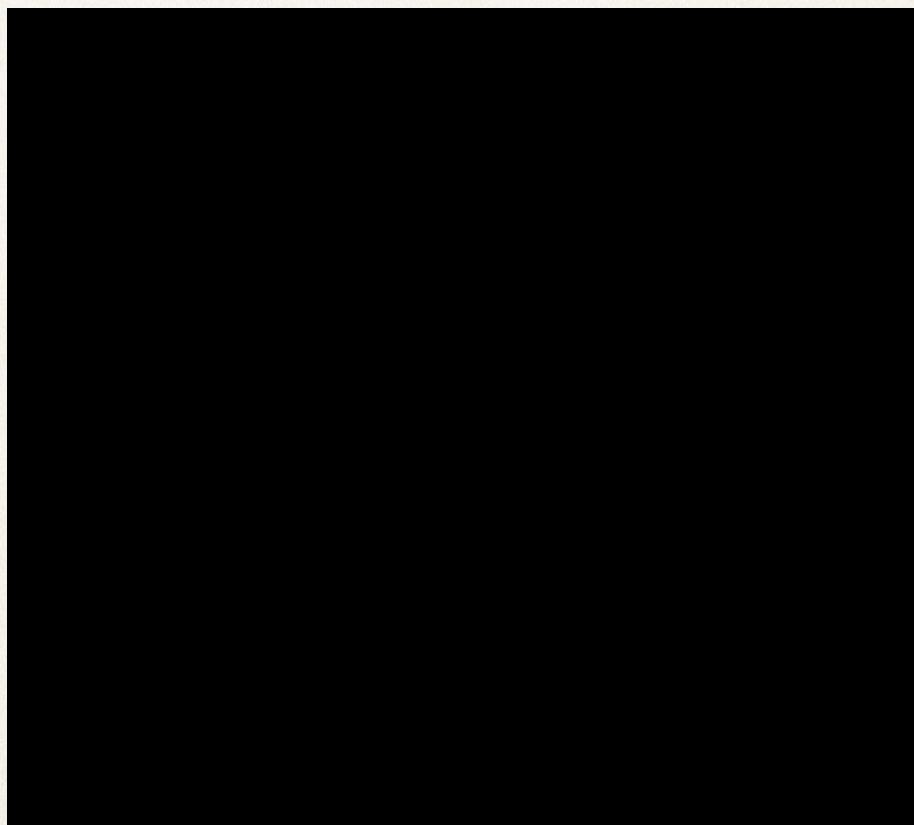


Lecture 18

लट्ठू — खिलाते भी, संभालते भी

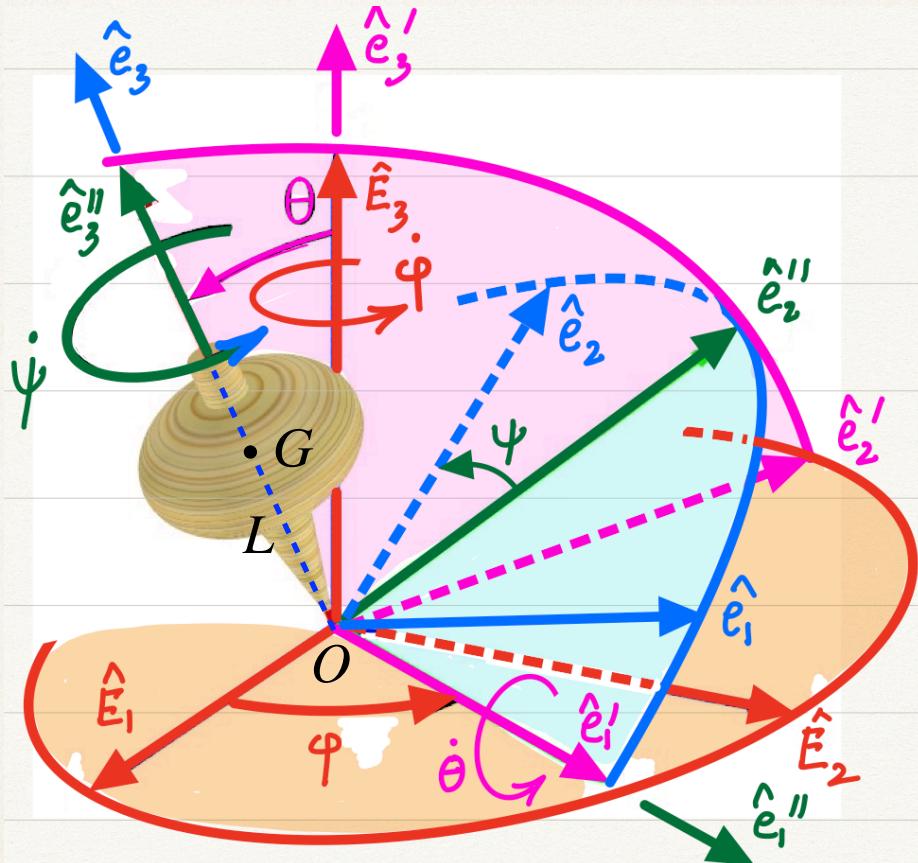
Tops & Gyroscopes

20 - 26 October, 2021



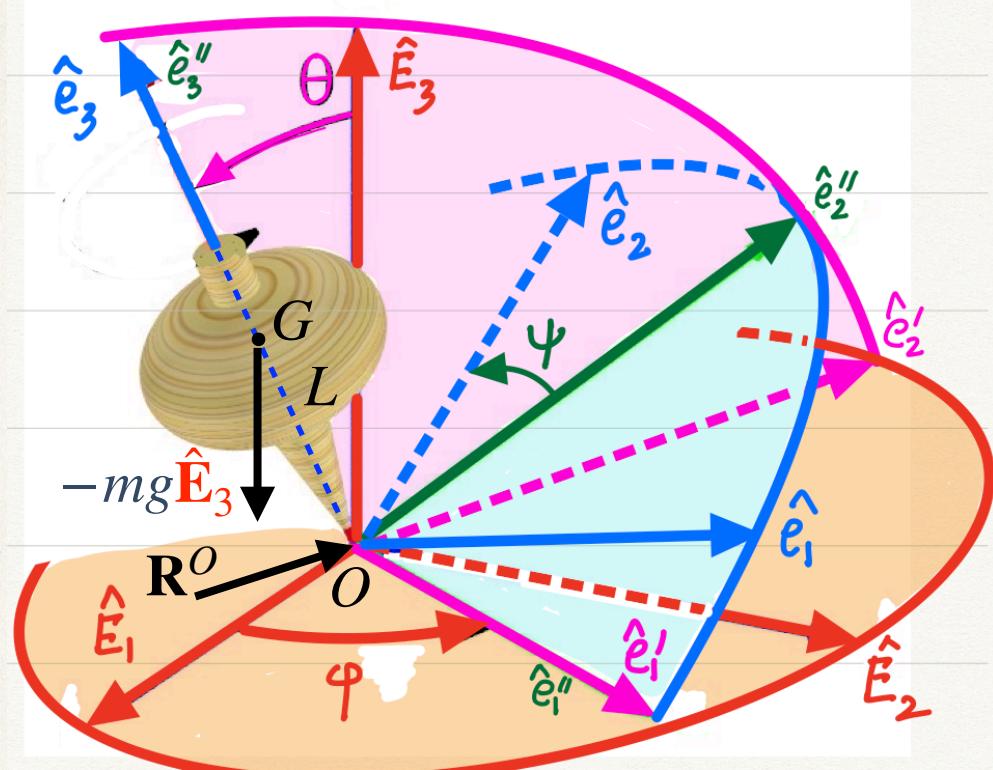
Tops: location

- I. **Orientation:** Use 3-1-3 Euler angle sequence.
- II. **CM position:** Know $\mathbf{r}^{G/O}$ once we know orientation as O is at origin of \mathcal{E}_0 .



$$\underbrace{\{\mathcal{E}_0, \hat{\mathbf{E}}_i\}}_{\text{Inertial}} \xrightarrow[\dot{\varphi}\hat{\mathbf{E}}_3]{R_\varphi} \{\mathcal{E}', \hat{\mathbf{e}}'_i\} \xrightarrow[\dot{\theta}\hat{\mathbf{e}}'_1]{R_\theta} \{\mathcal{E}'', \hat{\mathbf{e}}''_i\} \xrightarrow[\dot{\psi}\hat{\mathbf{e}}''_3]{R_\psi} \underbrace{\{\mathcal{E}, \hat{\mathbf{e}}_i\}}_{\text{BFCS}}$$

Tops: kinetics



I. FBD

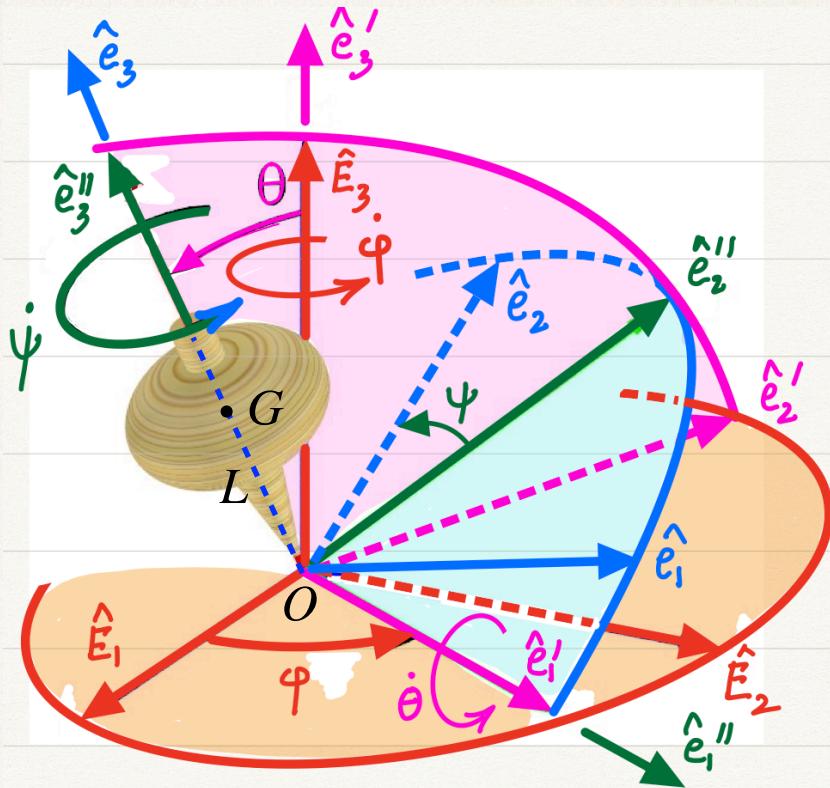
$$\text{II. LMB: } -mg\hat{\mathbf{E}}_3 + \mathbf{R}^O = m\mathbf{a}^G$$

$$\text{III. AMB}_{/O}: \mathbf{M}^O = \boldsymbol{\omega}^{\mathcal{B}} \times (\mathbf{I}^O \cdot \boldsymbol{\omega}^{\mathcal{B}}) + \mathbf{I}^O \cdot \boldsymbol{\alpha}^{\mathcal{B}}$$

$$1. \text{ In } \mathcal{E}: \quad \begin{cases} M_1 = I_1 \dot{\omega}_1 - (I_2 - I_3) \omega_2 \omega_3 \\ M_2 = I_2 \dot{\omega}_2 - (I_3 - I_1) \omega_3 \omega_1 \\ \text{Euler's eqns.} \quad M_3 = I_3 \dot{\omega}_3 - (I_1 - I_2) \omega_1 \omega_2 \end{cases}$$

$$2. \quad \left. \begin{array}{l} M_1 = mgL \sin \theta \cos \psi \\ M_2 = -mgL \sin \theta \sin \psi \end{array} \right\}, M_3 = 0$$

Tops: kinematics



$$\underbrace{\{\mathcal{E}_0, \hat{\mathbf{E}}_i\} \xrightarrow[\dot{\varphi}\hat{\mathbf{E}}_3]{R_\varphi} \{\mathcal{E}', \hat{\mathbf{e}}'_i\} \xrightarrow[\dot{\theta}\hat{\mathbf{e}}'_1]{R_\theta} \{\mathcal{E}'', \hat{\mathbf{e}}''_i\} \xrightarrow[\dot{\psi}\hat{\mathbf{e}}''_3]{R_\psi} \{\mathcal{E}, \hat{\mathbf{e}}_i\}}_{\text{Inertial}} \quad \underbrace{\{\mathcal{E}, \hat{\mathbf{e}}_i\}}_{\text{BFCS}}$$

Kinematics: $\boldsymbol{\omega}^{\mathcal{B}} = \dot{\varphi}\hat{\mathbf{E}}_3 + \dot{\theta}\hat{\mathbf{e}}'_1 + \dot{\psi}\hat{\mathbf{e}}_3$, $\boldsymbol{\alpha}^{\mathcal{B}} = \dot{\boldsymbol{\omega}}^{\mathcal{B}}$

$$\left| \begin{array}{l} \omega_1 = \dot{\varphi} \sin \theta \sin \psi + \dot{\theta} \cos \psi \\ \omega_2 = \dot{\varphi} \sin \theta \cos \psi - \dot{\theta} \sin \psi \\ \omega_3 = \dot{\varphi} \cos \theta + \dot{\psi} \end{array} \right. \quad \left| \begin{array}{l} \alpha_1 = \dot{\omega}_1 \\ \alpha_2 = \dot{\omega}_2 \\ \alpha_3 = \dot{\omega}_3 \end{array} \right.$$

Symmetric tops

I. Governing equations for $I_1 = I_2 \neq I_3$.

1. In the $\hat{\mathbf{e}}_3$ direction:

$$M_3 = 0 \implies \omega_3 = \dot{\phi} \cos \theta + \dot{\psi} = \text{const.} =: \omega_3^0$$

2. $\hat{\mathbf{e}}_1$ and $\hat{\mathbf{e}}_2$ direction equations combine:

$$I_1(\ddot{\phi} \sin \theta + 2\dot{\phi}\dot{\theta} \cos \theta) - I_3\omega_3^0\dot{\theta} = 0 ;$$

$$I_1(\ddot{\theta} - \dot{\phi}^2 \sin \theta \cos \theta) + I_3\omega_3^0\dot{\phi} \sin \theta = mgL \sin \theta$$

3. Note ψ drops out completely. **Why?**

II. Special case: Steady precession

1. $\theta \approx \text{const.} \neq 0^\circ \implies \ddot{\theta} = \dot{\theta} = 0$

$$\implies -I_1\dot{\phi}^2 \cos \theta + I_3\dot{\phi}\omega_3^0 = mgL$$

2. If $|\dot{\phi}| \ll |\dot{\psi}|$, then

i. $\dot{\psi} \approx \omega_3^0 = \text{const.}$

ii. $\dot{\phi}^2 \ll |\omega_3^0\dot{\phi}| \implies \dot{\phi} \approx \frac{mgL}{I_3\dot{\psi}} = \text{const.}$

Tops: Simple analysis

I. For *symmetric top* ($I_1 = I_2 \neq I_3$), assume:

1. *Steady precession* ($\ddot{\theta} = \dot{\theta} = 0$);
2. *High spin* ($|\dot{\phi}| \ll |\dot{\psi}|$).
3. No moment about spin axis: $\mathbf{M}^O \cdot \hat{\mathbf{e}}_3 = 0$

$$\Rightarrow \begin{cases} \omega_1 = \dot{\phi} \sin \theta \sin \psi + \dot{\theta} \cos \psi = \dot{\phi} \sin \theta \sin \psi \\ \omega_2 = \dot{\phi} \sin \theta \cos \psi - \dot{\theta} \sin \psi = \dot{\phi} \sin \theta \cos \psi \\ \omega_3 = \dot{\phi} \cos \theta + \dot{\psi} \approx \dot{\psi} \end{cases} \Rightarrow |\omega_3| \gg |\omega_1|, |\omega_2|$$

II. $\mathbf{h}^O = \sum I_i \omega_i \hat{\mathbf{e}}_i \approx I_3 \dot{\psi} \hat{\mathbf{e}}_3 \Rightarrow \overset{\circ}{\mathbf{h}}{}^O = I_3 \dot{\psi} \hat{\mathbf{e}}_3$

III. AMB_O: $\mathbf{M}^O = \dot{\mathbf{h}}^O = \boldsymbol{\omega}^{\mathcal{B}} \times \mathbf{h}^O + \overset{\circ}{\mathbf{h}}{}^O$

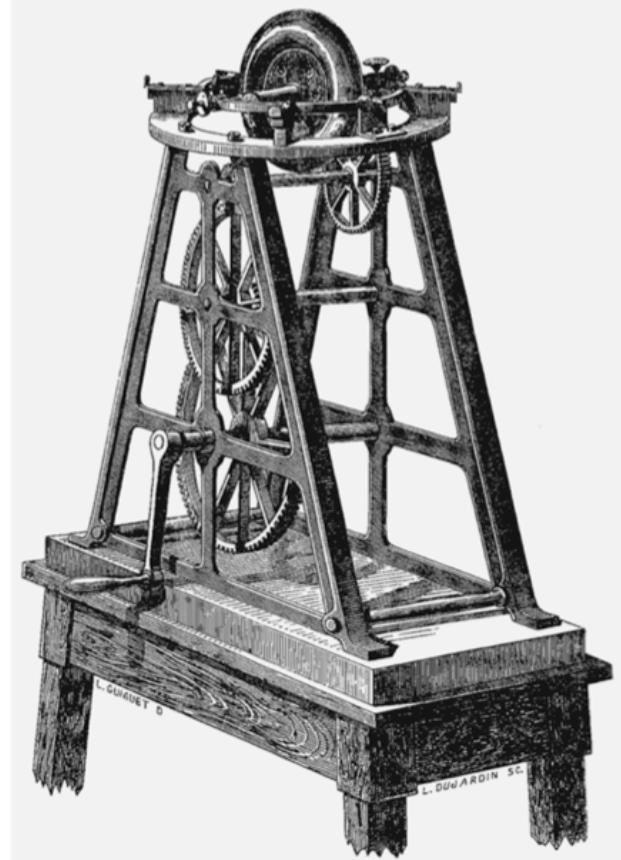
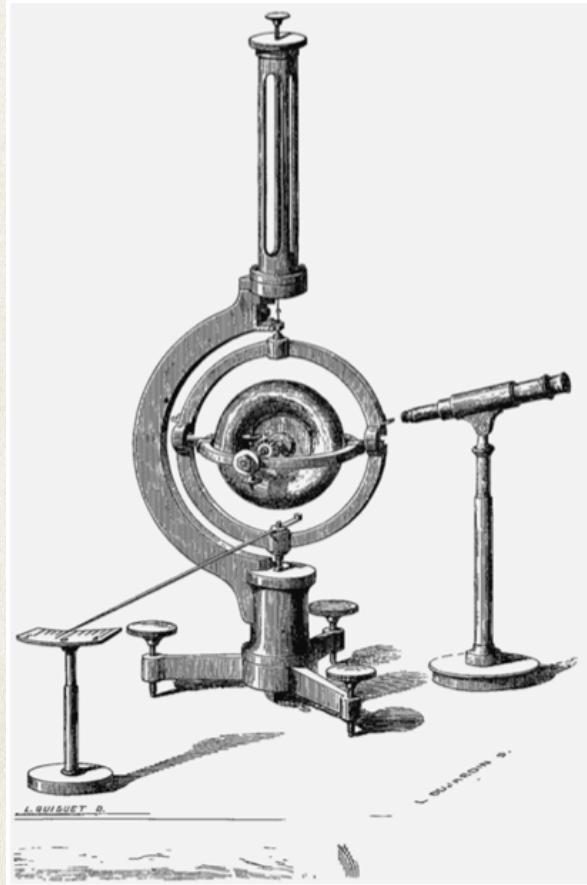
1. $\Rightarrow \mathbf{M}^O \approx I_3 \dot{\phi} \hat{\mathbf{E}}_3 \times \dot{\psi} \hat{\mathbf{e}}_3 + I_3 \cancel{\dot{\psi}} \hat{\mathbf{e}}_3^0$

$\Rightarrow \mathbf{M}^O = I_3 \dot{\phi} \times \dot{\psi}$

$\dot{\phi}$: Precessional rate; $\dot{\psi}$: Spin rate

2. $\mathbf{M}^O = mgL \hat{\mathbf{e}}_1'' \Rightarrow mgL = I_3 \dot{\phi} \dot{\psi}$.

Gyroscopes



Foucault's gyroscope (left) and its launching device (right) for measuring the Earth's rotation

Gyroscope

Fast spinning rotor, generally heavy.

1. *Symmetric rotor:* ($I_1 = I_2 \neq I_3$);
2. *Steady precession* ($\ddot{\theta} = \dot{\theta} = 0$);
3. *High spin* ($|\dot{\varphi}| \ll |\dot{\psi}|$);
4. *No moment about spin axis:* $\mathbf{M}^O \cdot \hat{\mathbf{e}}_3 = 0$
5. **Gyroscope equation:** $\mathbf{M}^O = I_3 \dot{\varphi} \times \dot{\psi}$
6. *Spin stops falling and causes precession.*
7. *Frictional losses cause rotor to fall.*



Application I

Spin stabilization of projectiles



Application II

Stabilization of ships and space crafts



As vessel rolls, the gyro naturally moves fore and aft, delivering a torque to PORT and STBD change orientation, it stabilizes the



Application III

Self-assembly robots

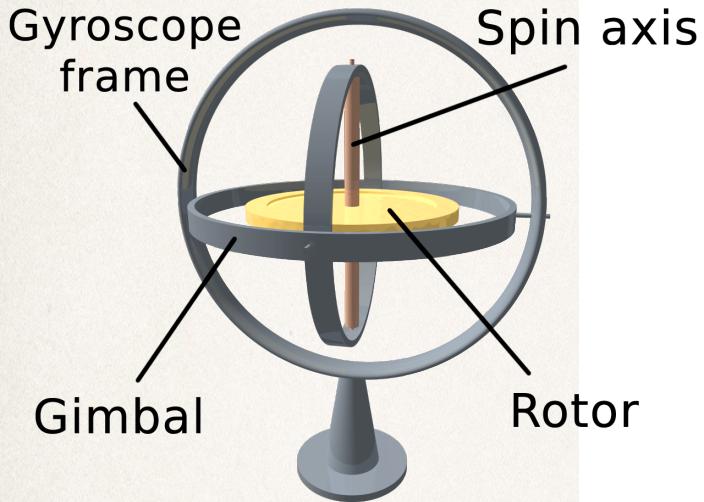
The Cubli

Building a cube that can jump up and balance



Application IV

Inertial navigation systems



DESPITE YEARS OF STUDYING
PHYSICS, I STILL FIND
GYROSCOPES A LITTLE FREAKY.



GREETINGS,
HUMAN.

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