

Cosmic Watch:

Investigating Cosmic Muons With Planes

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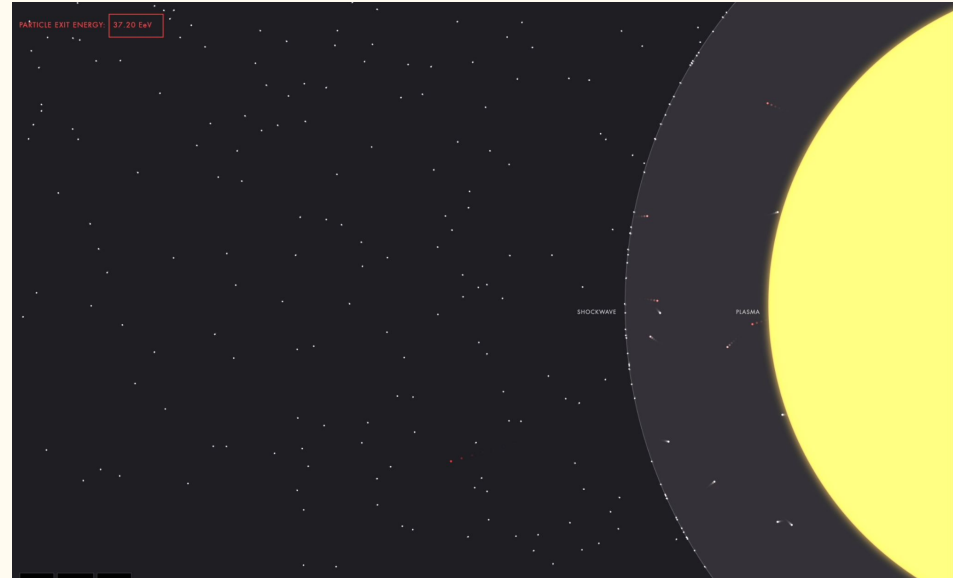
The Oh-My-God Particle

- [illegible]



What? Who? Why?

- Where do they come from?
 - Research today
 - Our galaxy
 - Sun
 - Neighboring galaxies
 - Pulsars
 - Supernovas



PARTICLE EXIT ENERGY: 37.20 EeV

SHOCKWAVE

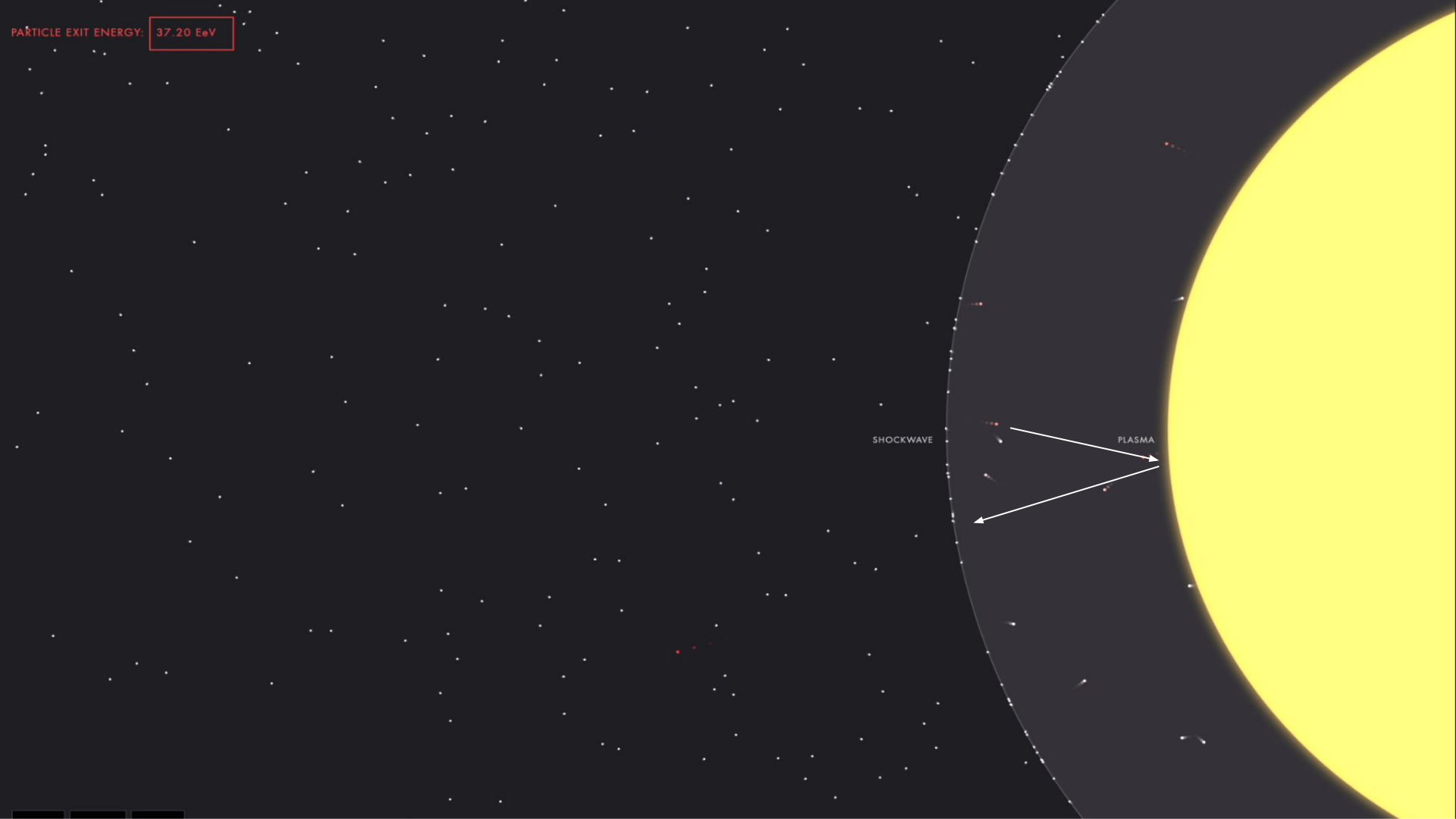
PLASMA



PARTICLE EXIT ENERGY: 37.20 EeV

SHOCKWAVE

PLASMA

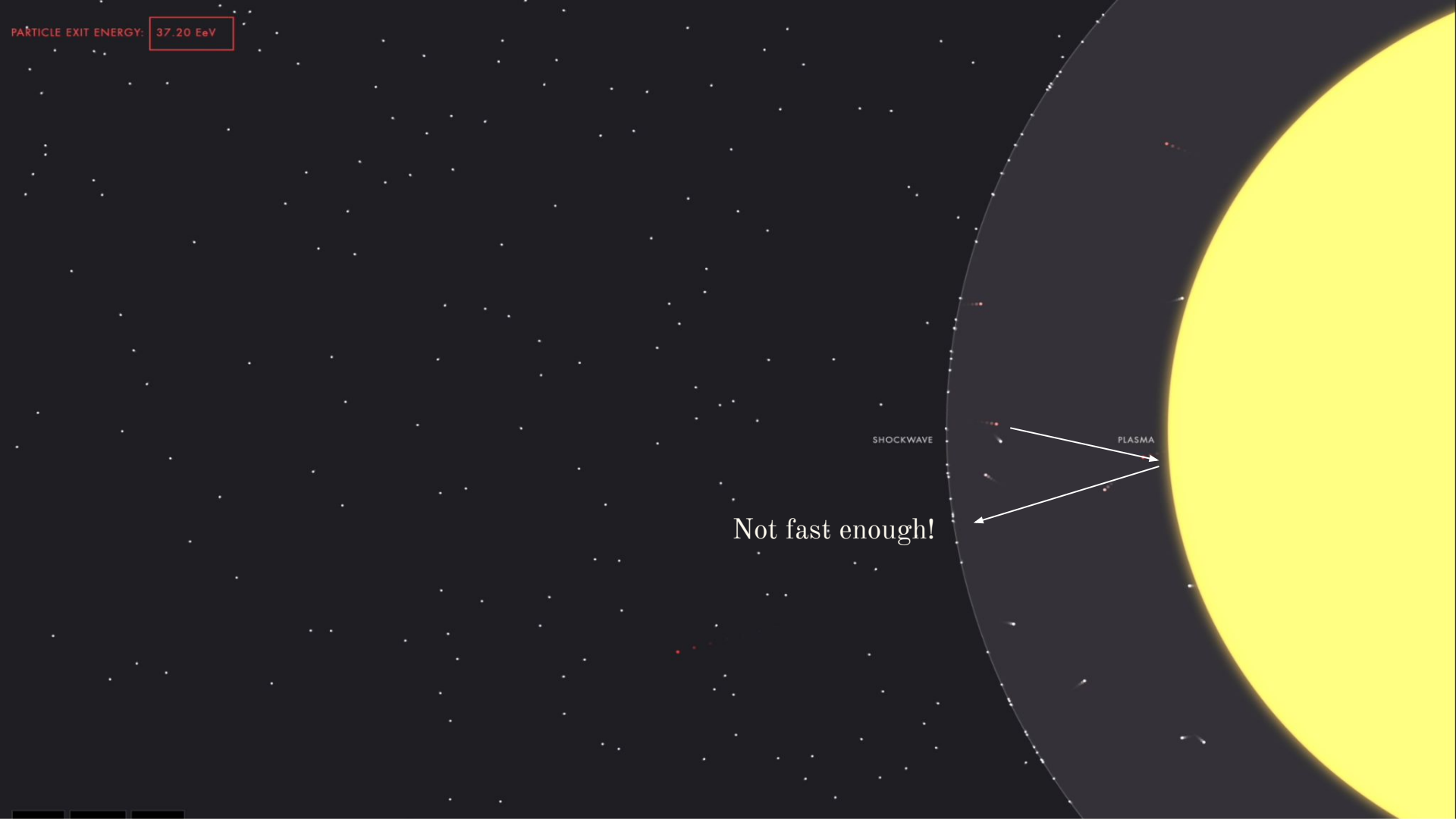


PARTICLE EXIT ENERGY: 37.20 EeV

SHOCKWAVE

PLASMA

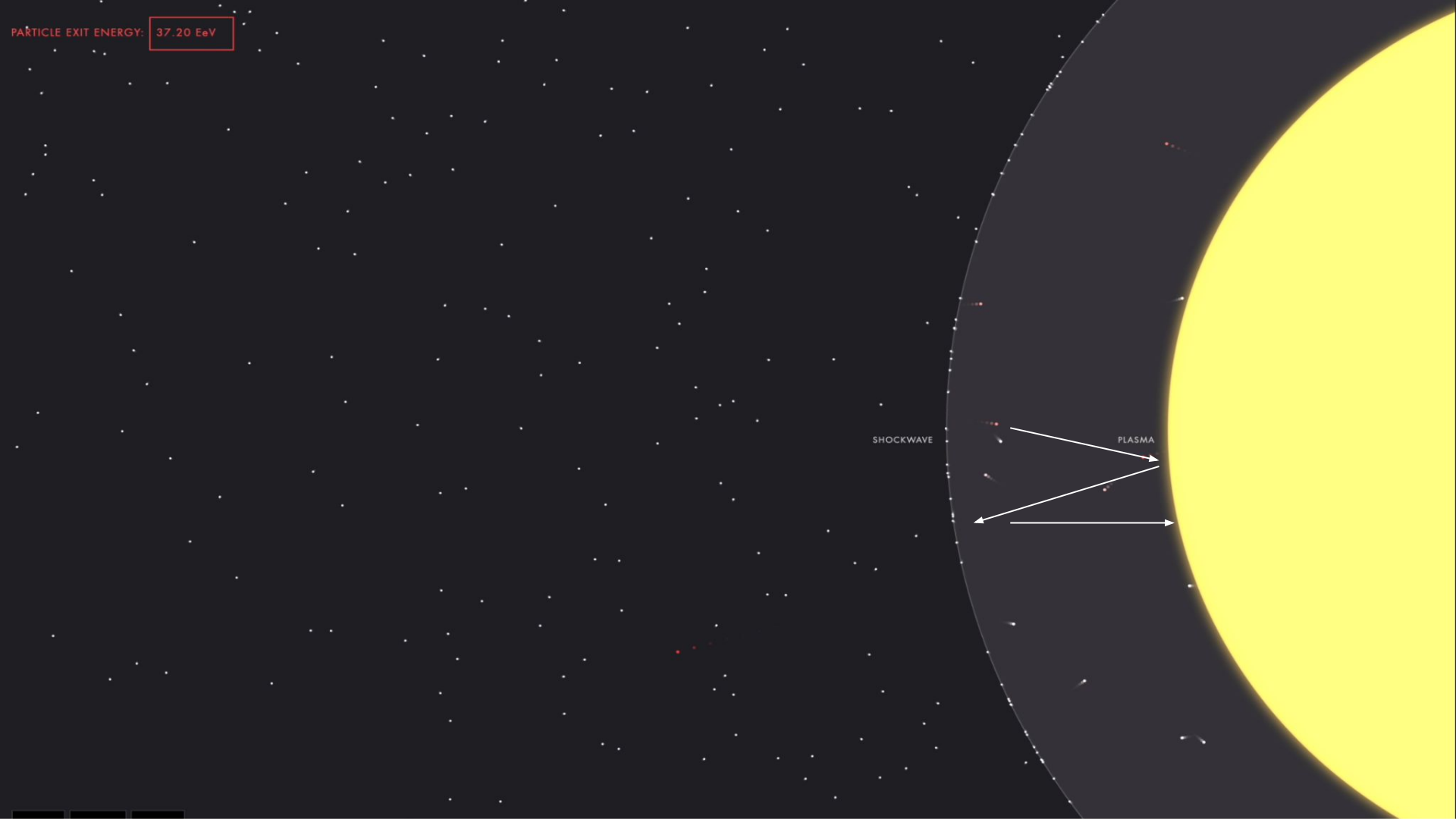
Not fast enough!



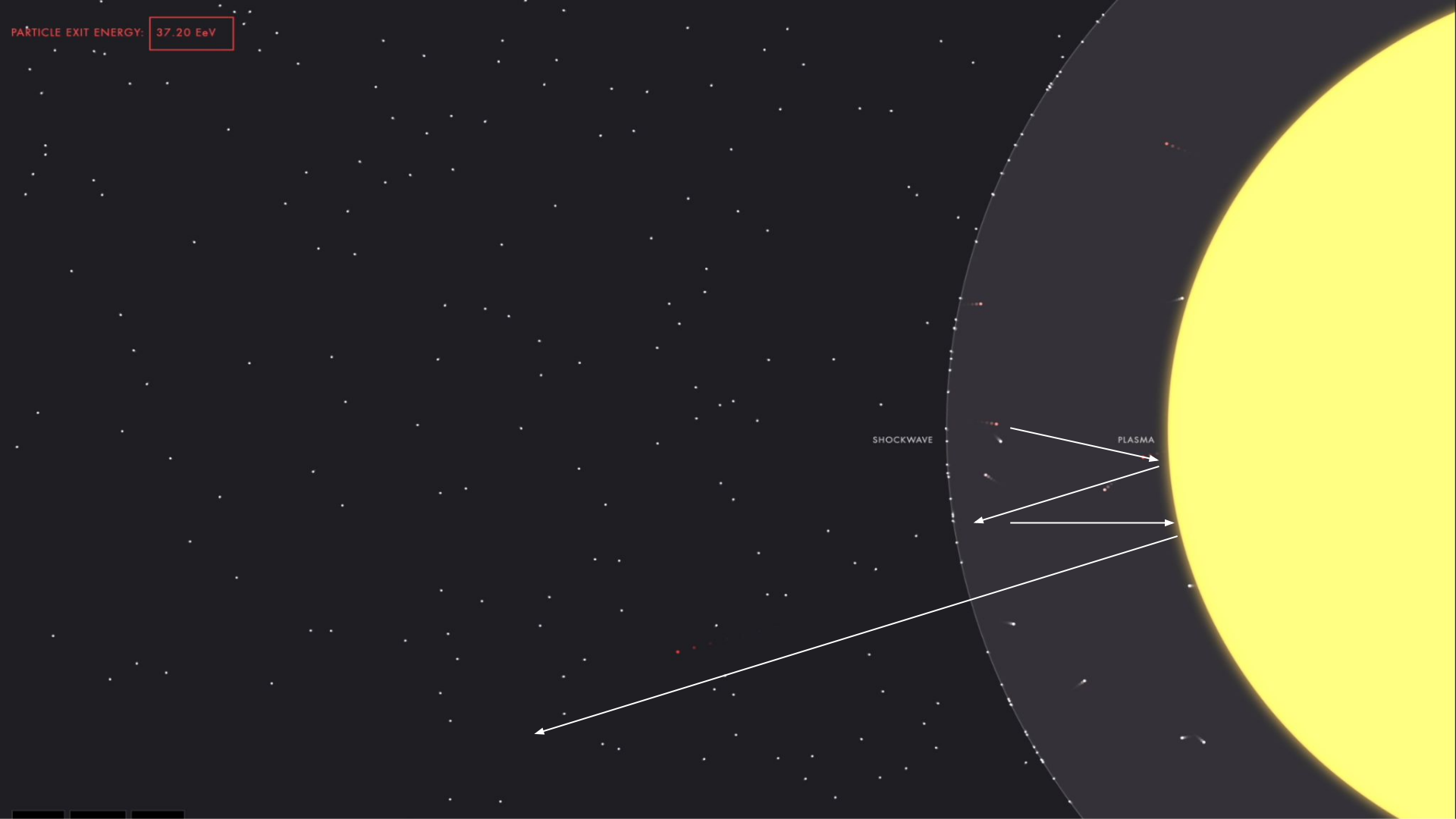
PARTICLE EXIT ENERGY: 37.20 EeV

SHOCKWAVE

PLASMA



PARTICLE EXIT ENERGY: 37.20 EeV

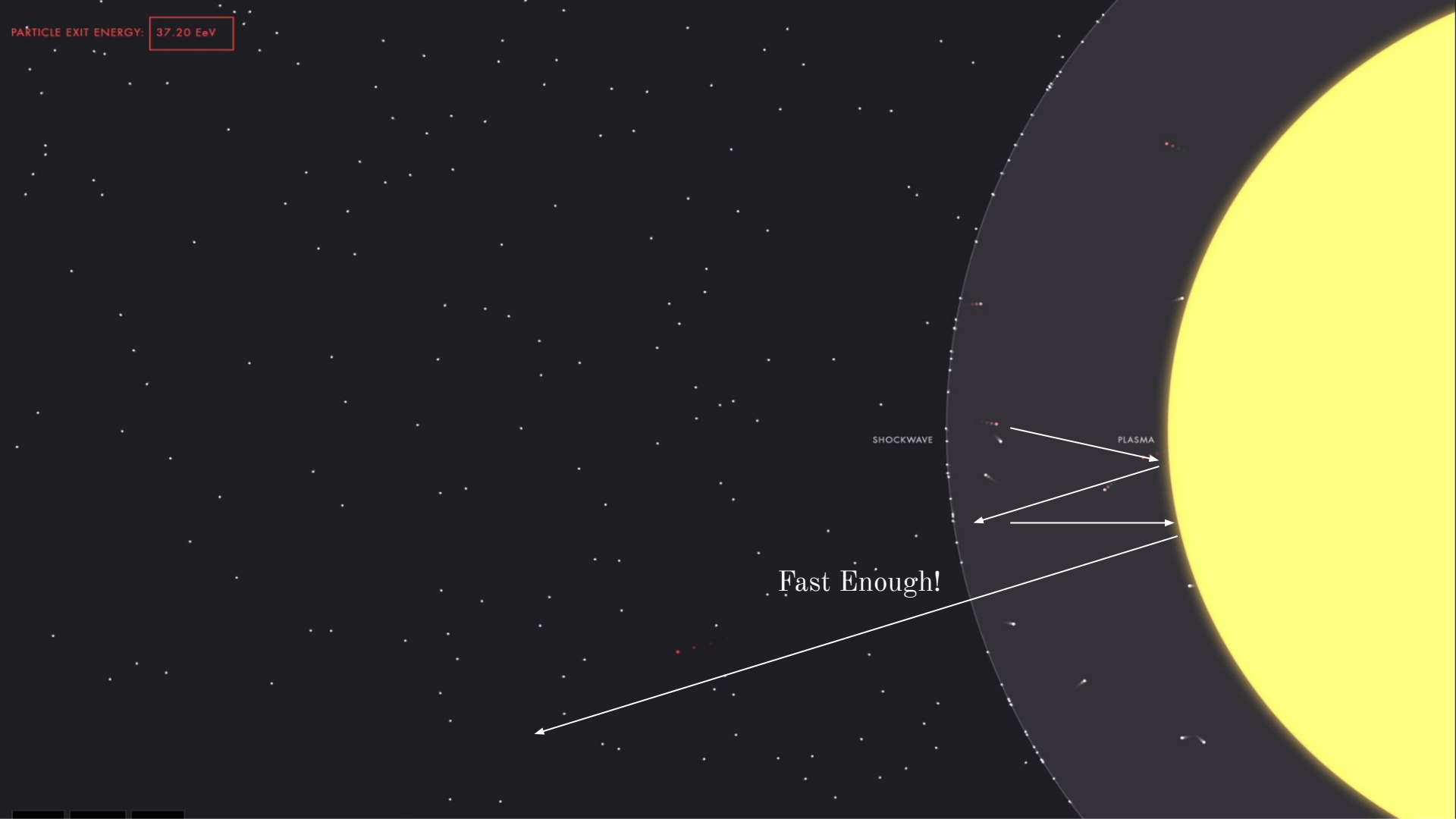


PARTICLE EXIT ENERGY: 37.20 EeV

SHOCKWAVE

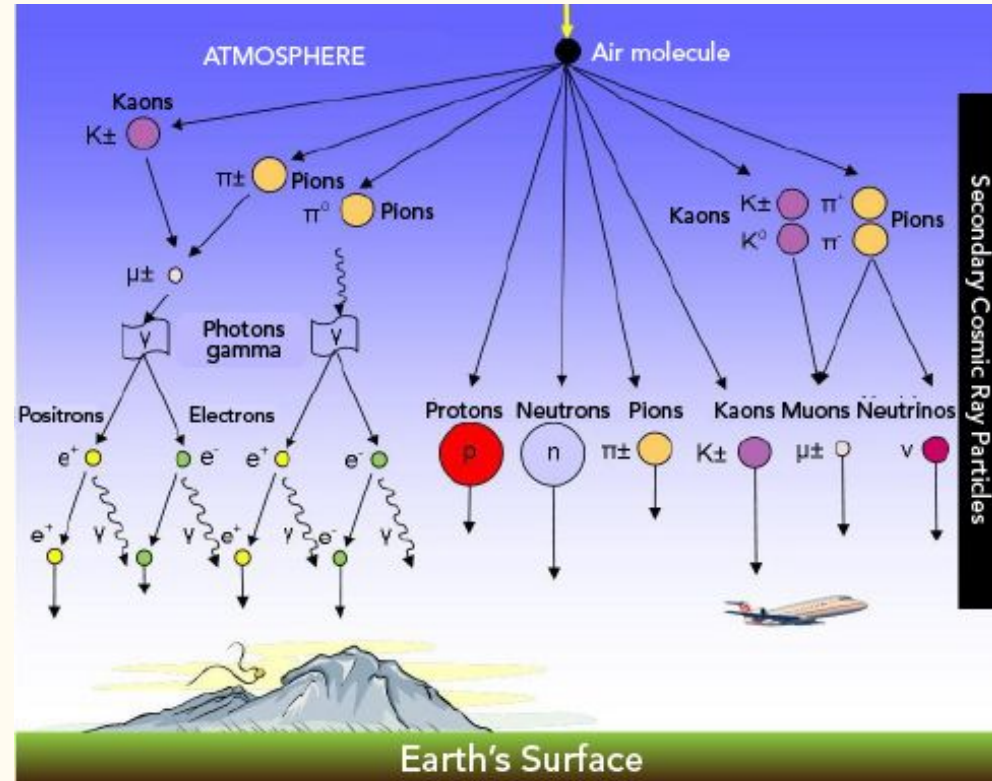
PLASMA

Fast Enough!

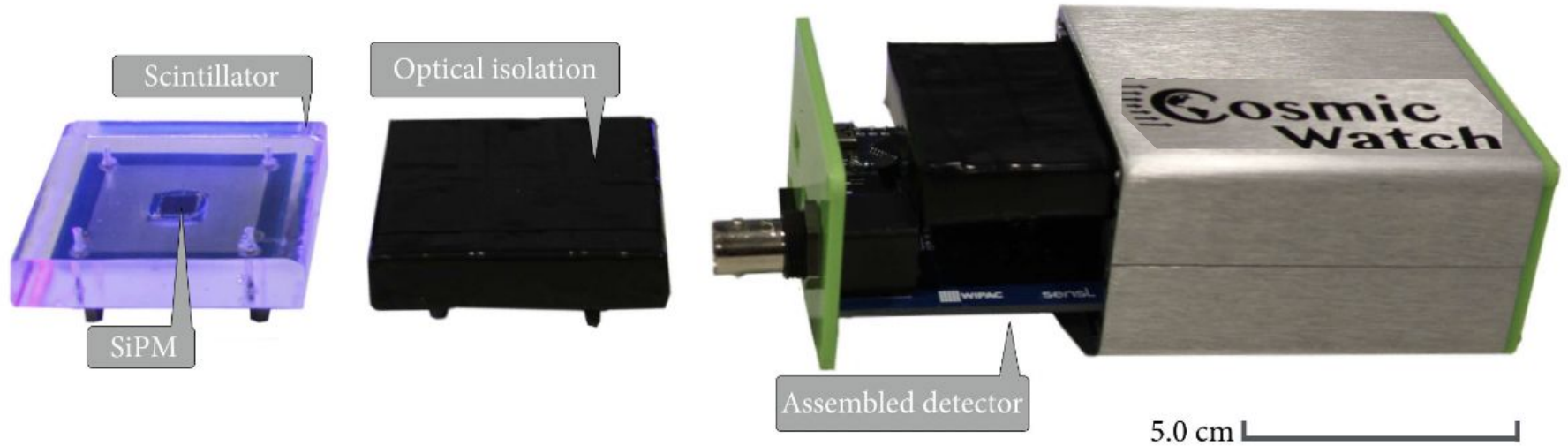


What? Who? Why?

- 74% ionized hydrogen
(free protons)
- 18% from helium nuclei
(two protons and two neutrons)
- 8% trace amounts of heavier
elements



The Cosmic Watch

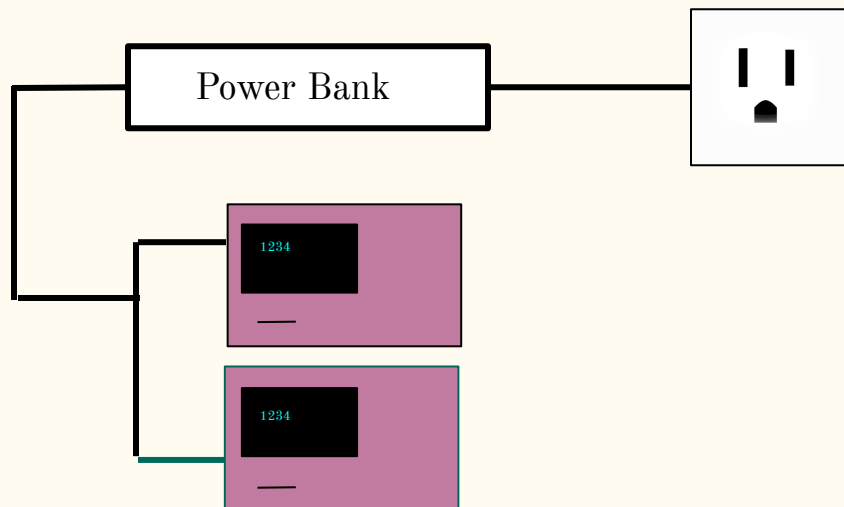


Goal

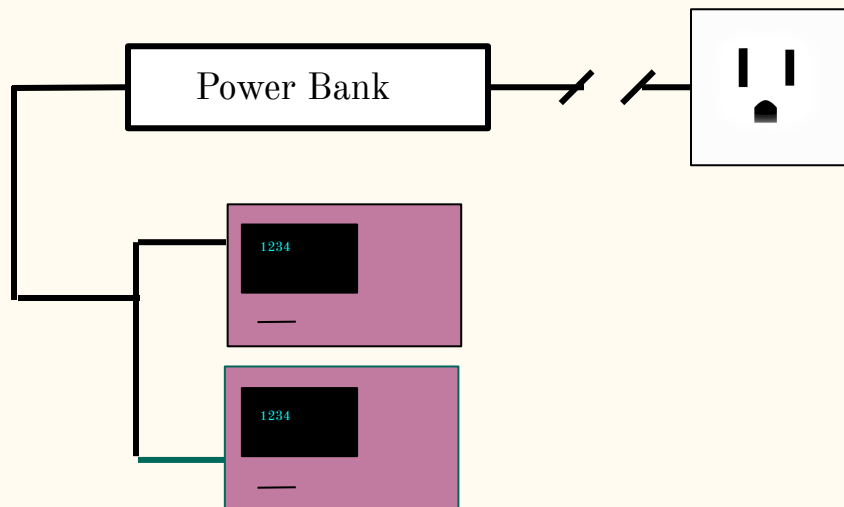
Investigate the relationship between count rate of cosmic muons as a function of altitude

1. Determine the function between altitude and counts
2. See how well the function describes the experimental data
3. Determine how energies vary across flights

Experimental Setup



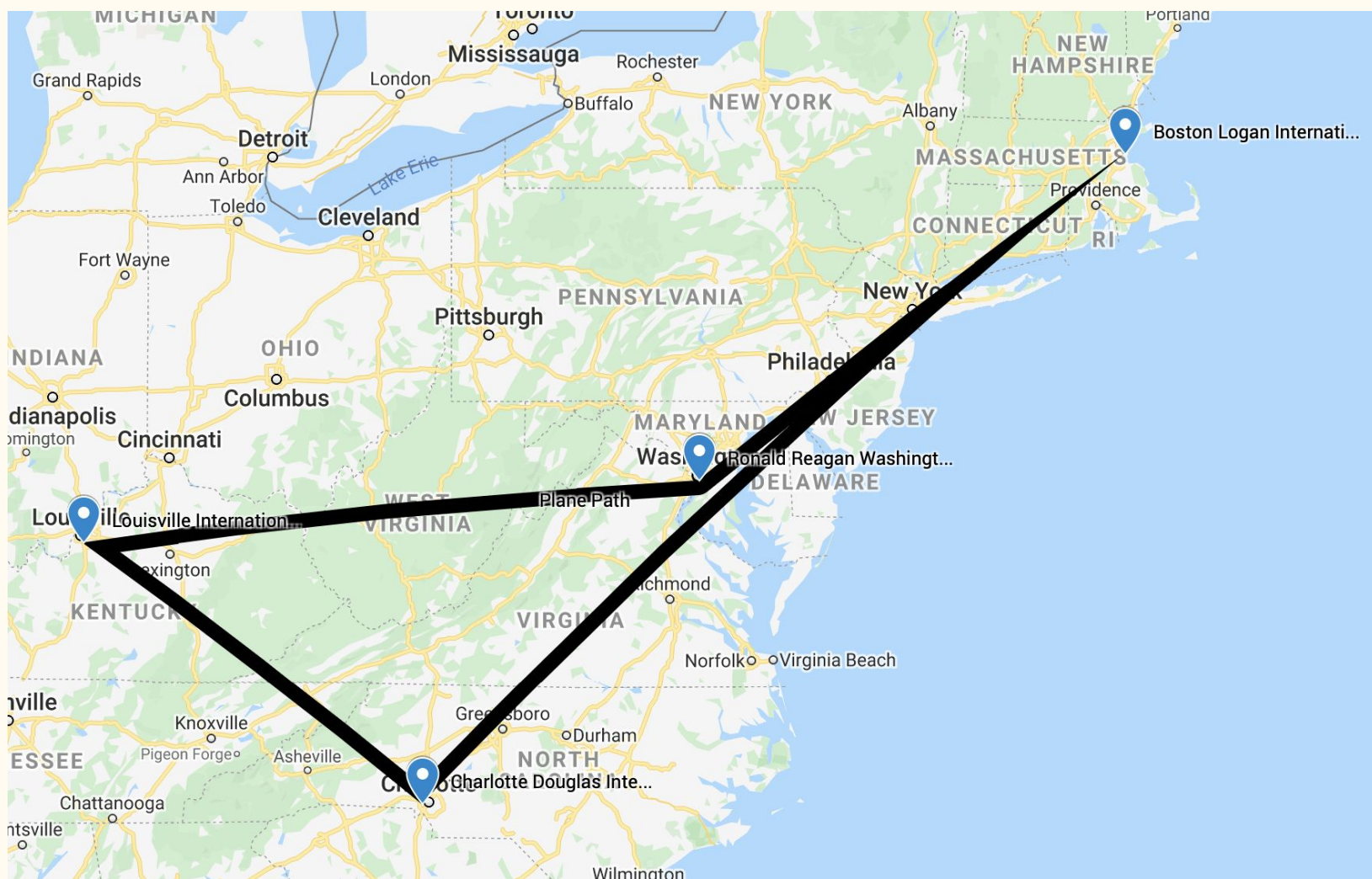
Experimental Setup

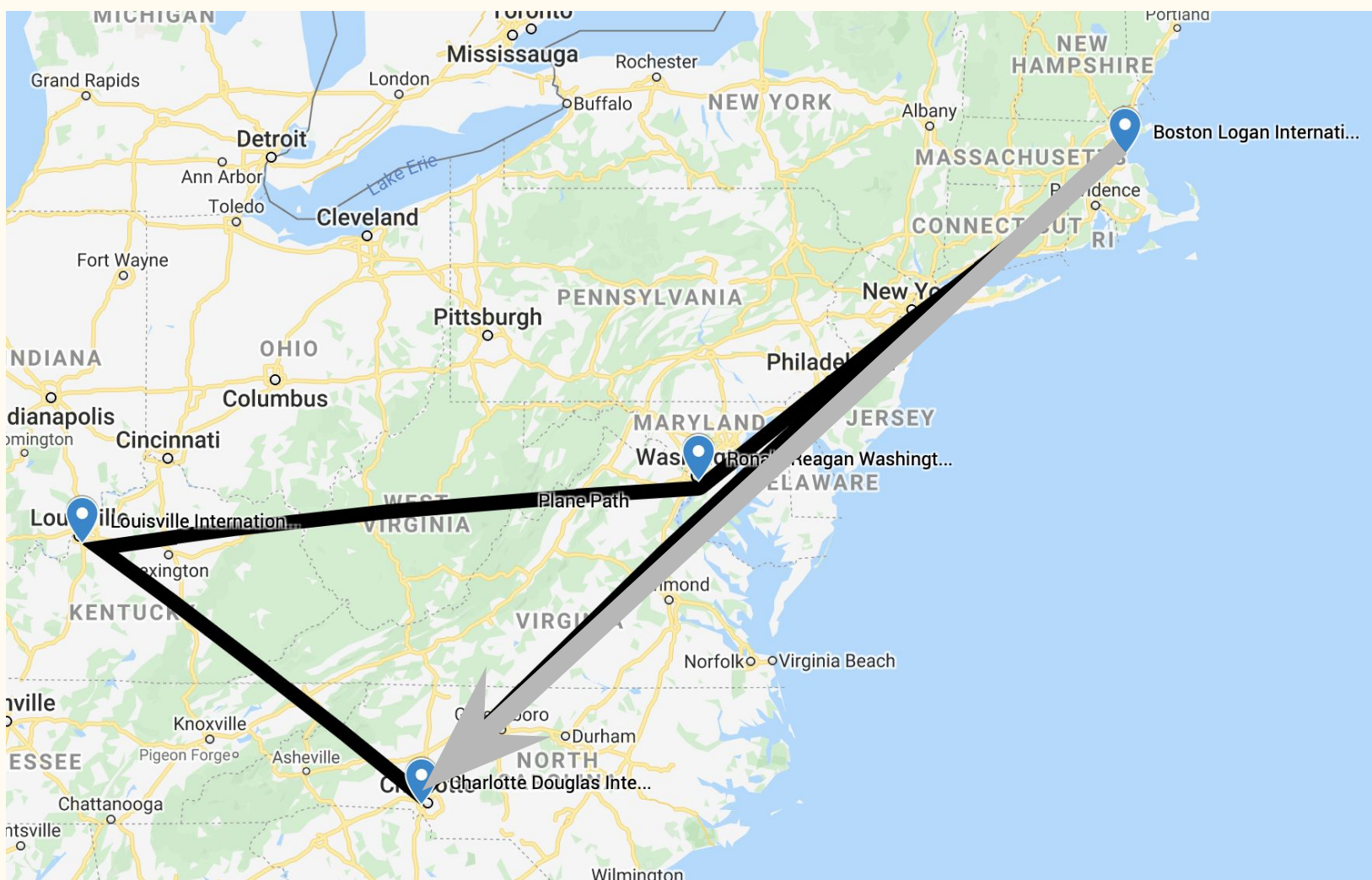


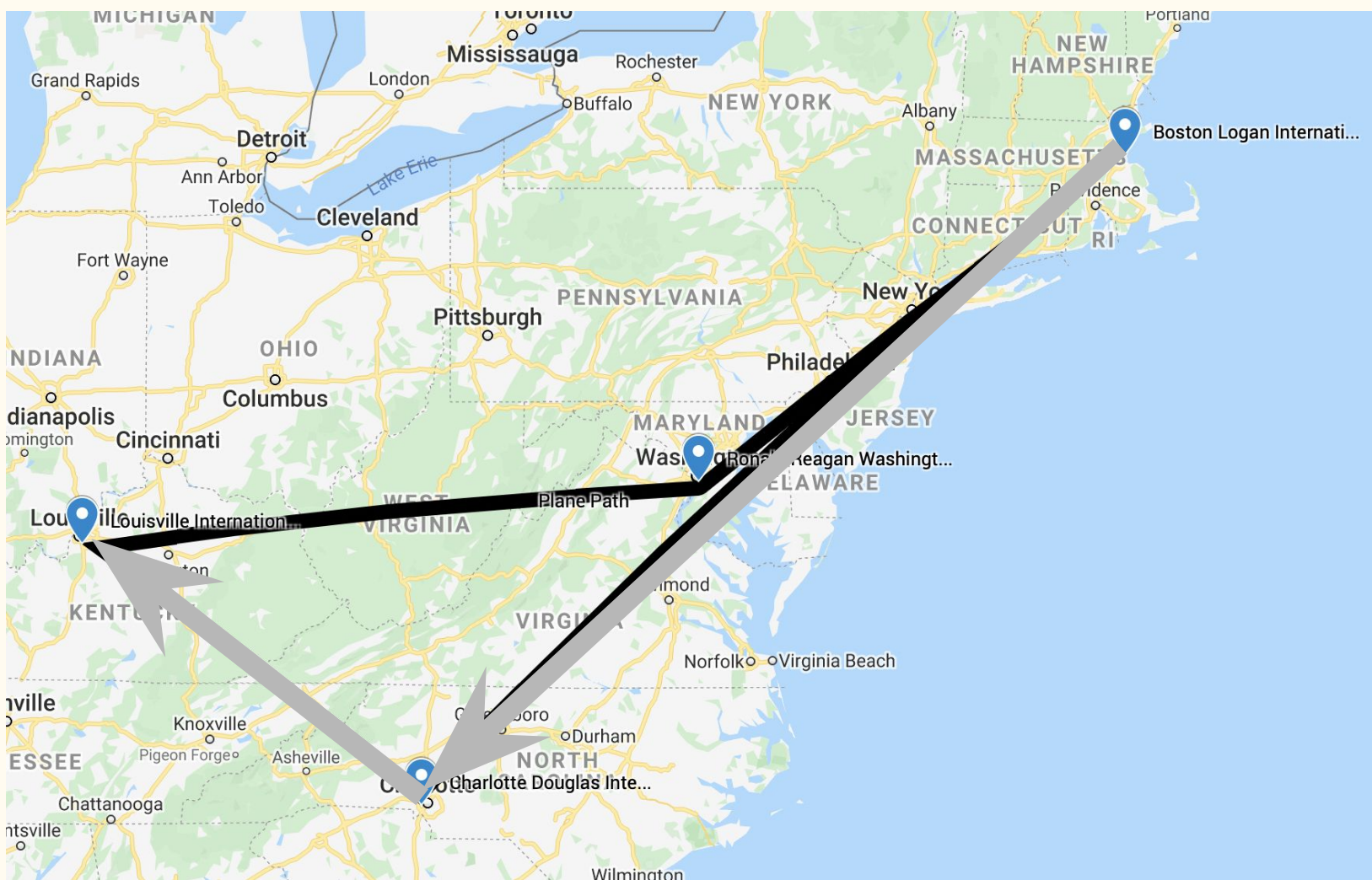
American Airlines Flights

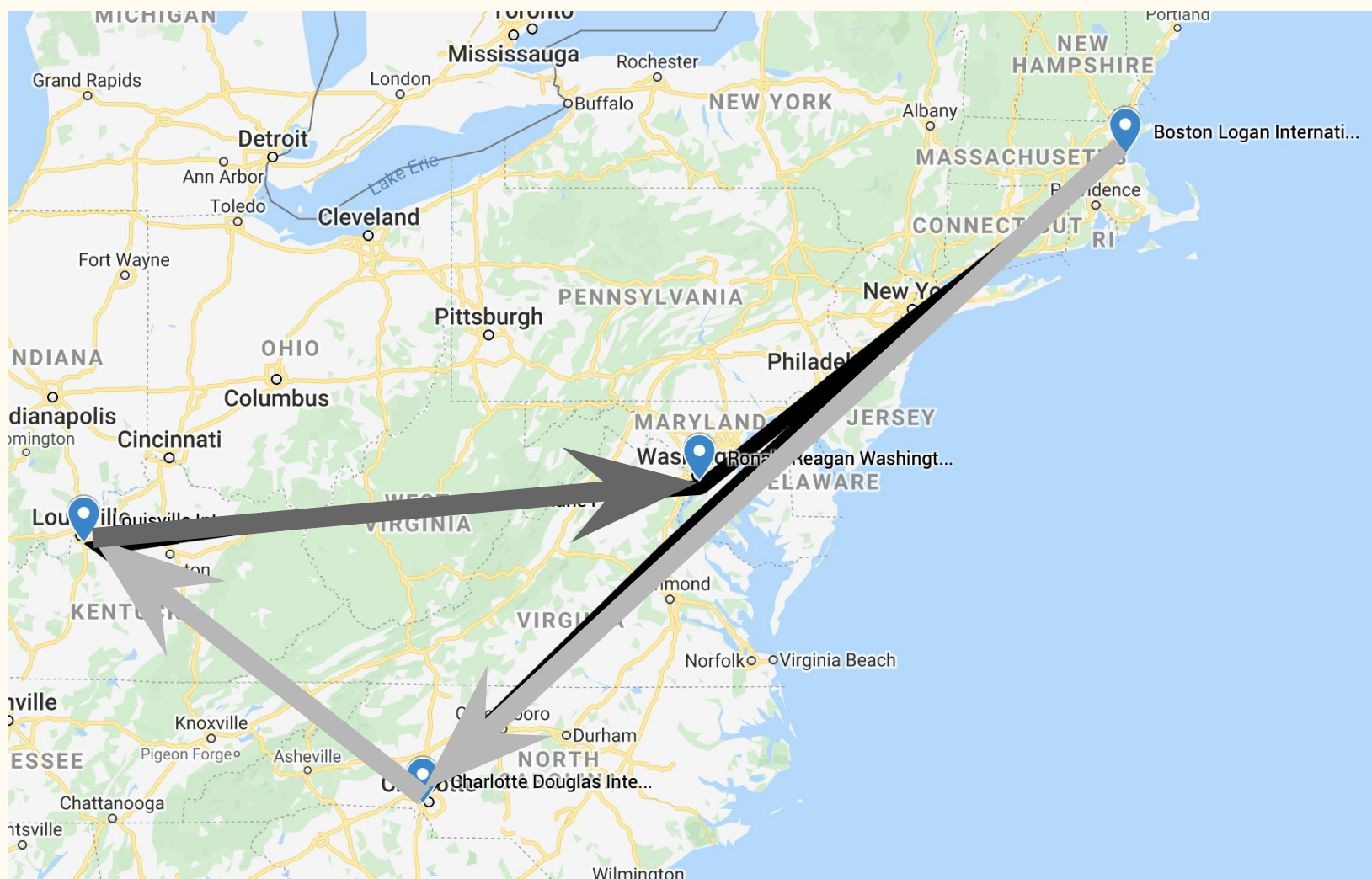
Flight Number, Date and Time	Max Altitude (km)	Duration (s)
653 11/18 9:45 am	10.97 ± 0.88	7250 ± 580
5326 11/18 1:21 pm	7.92 ± 0.63	3438 ± 275
4917 11/21 6:21 am	10.67 ± 0.85	3668 ± 293
2152 11/21 9:00 am	9.46 ± 0.76	3995 ± 319

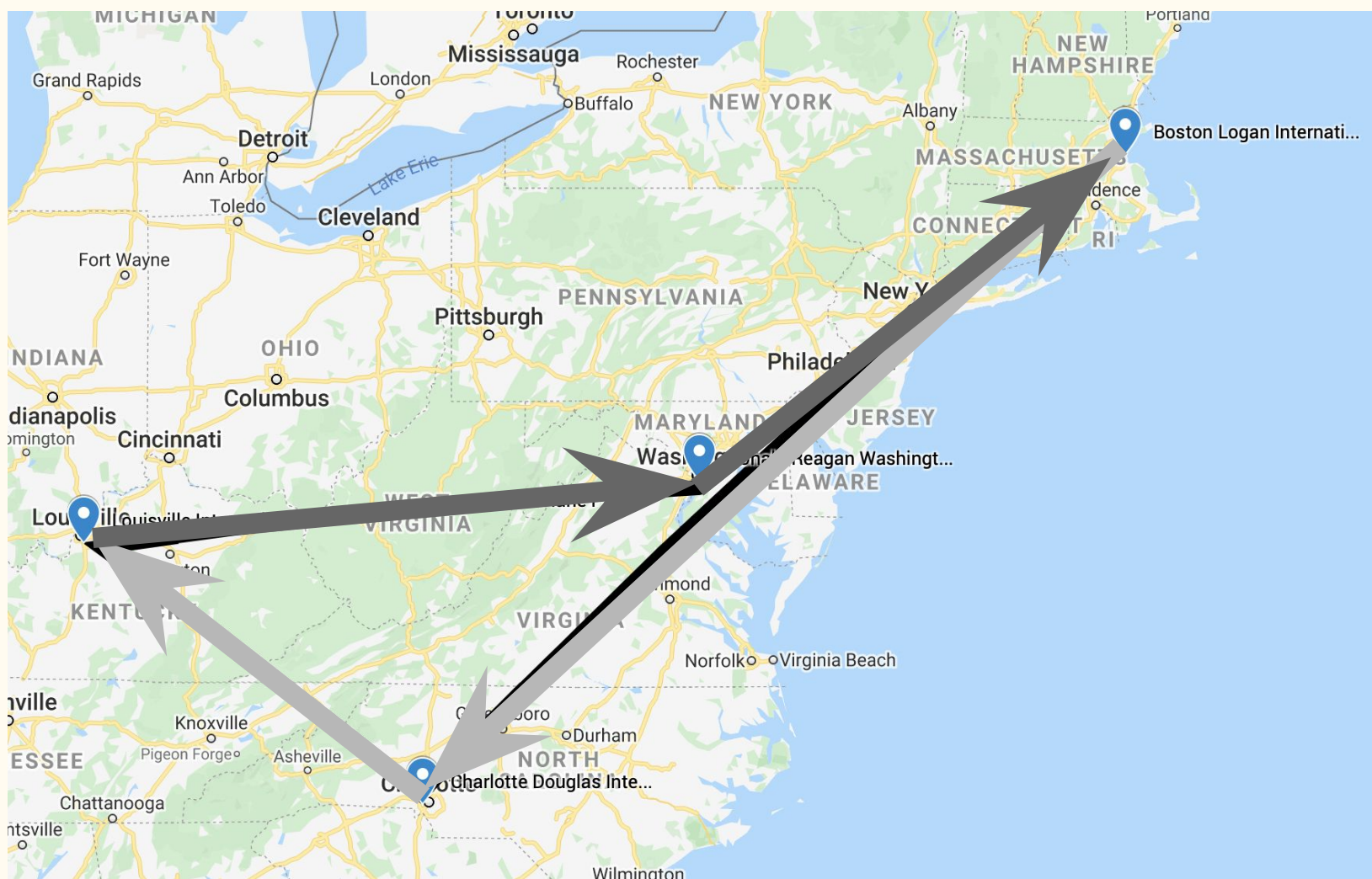
Flight data is extracted from FlightAware.com



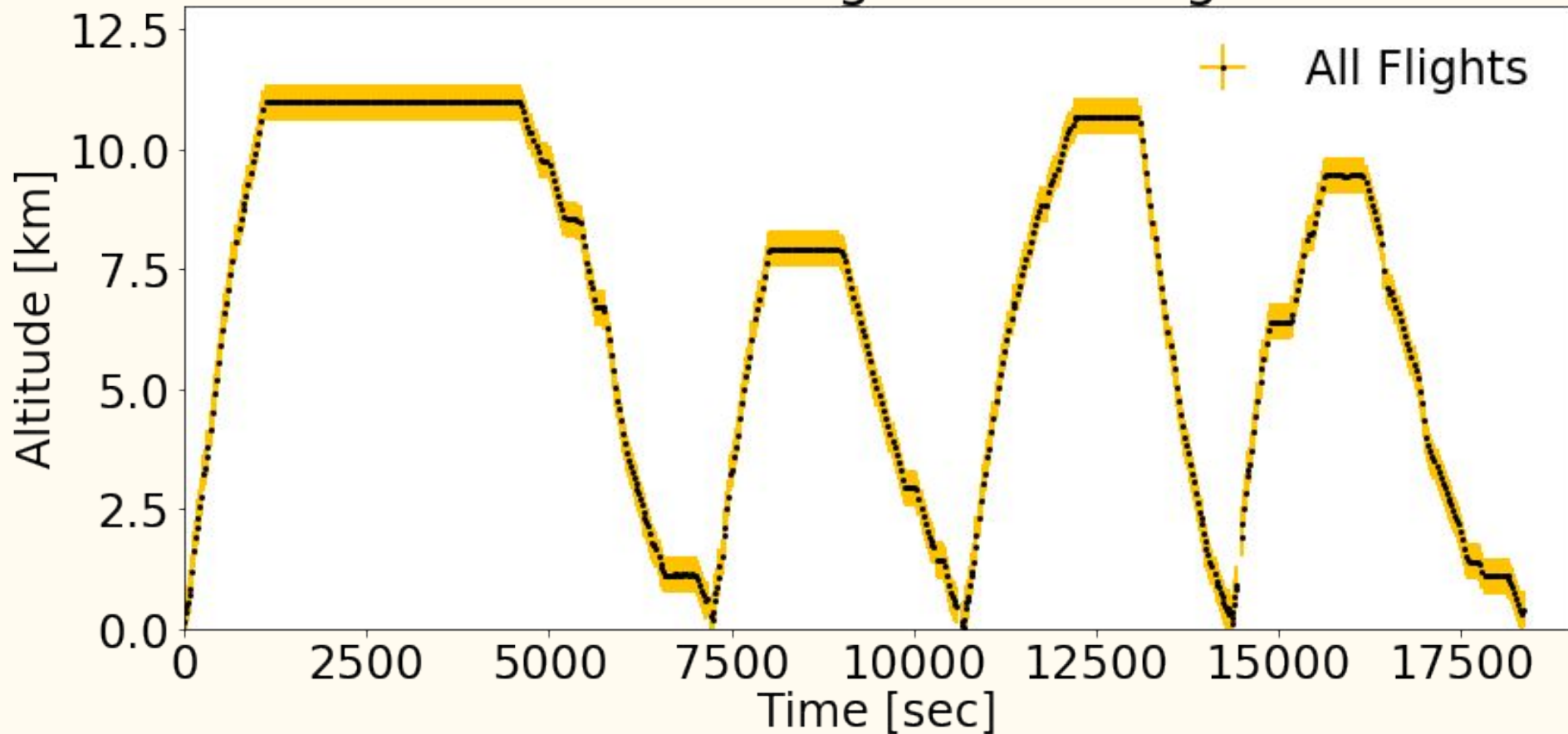




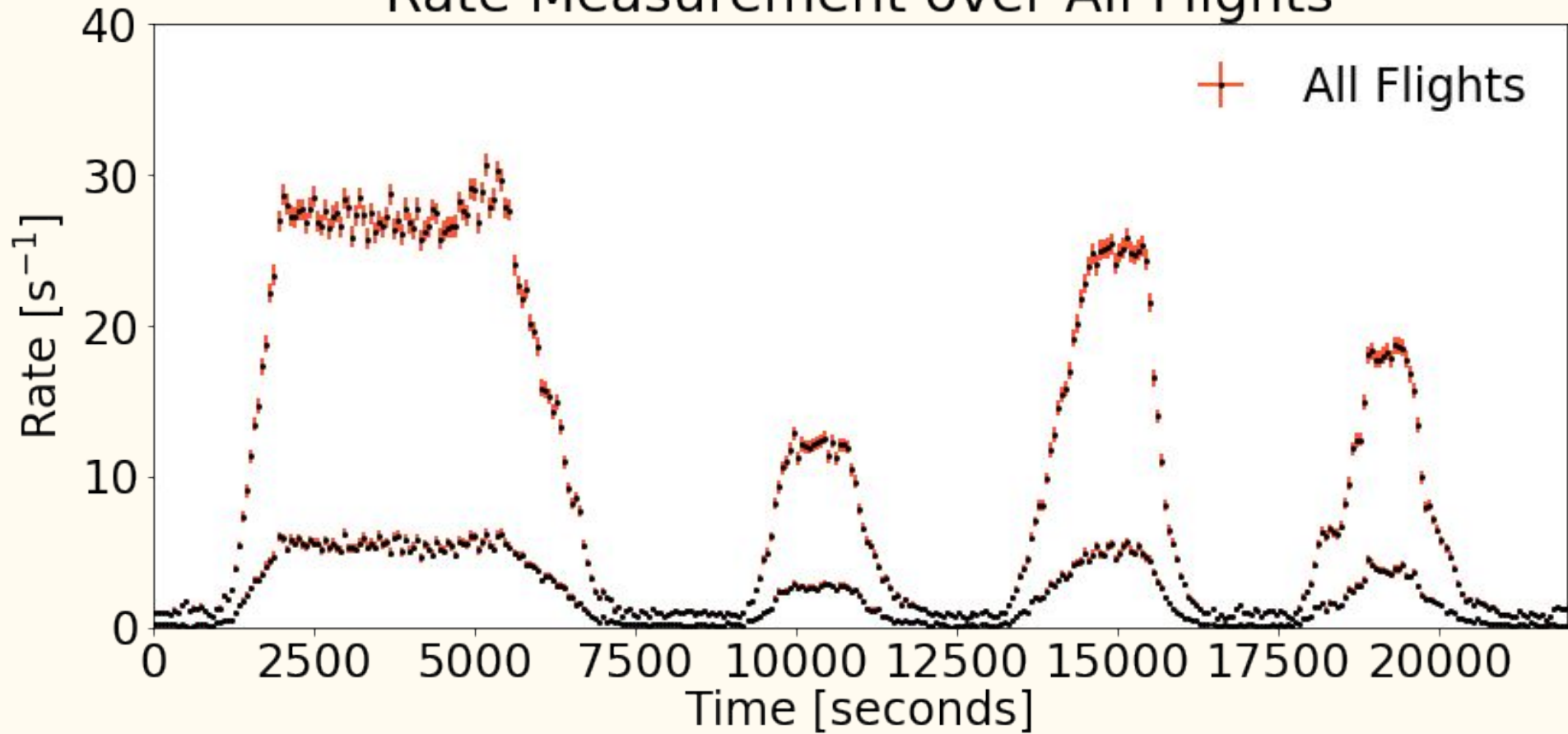




Altitude of Four Flights from Flight Data



Rate Measurement over All Flights



Total Count rate: 15.0772 \pm 0.044Hz

Coincident Count rate: 3.0732 \pm 0.0199Hz

4.8833 \pm 0.0347Hz

1.1214 \pm 0.0166Hz

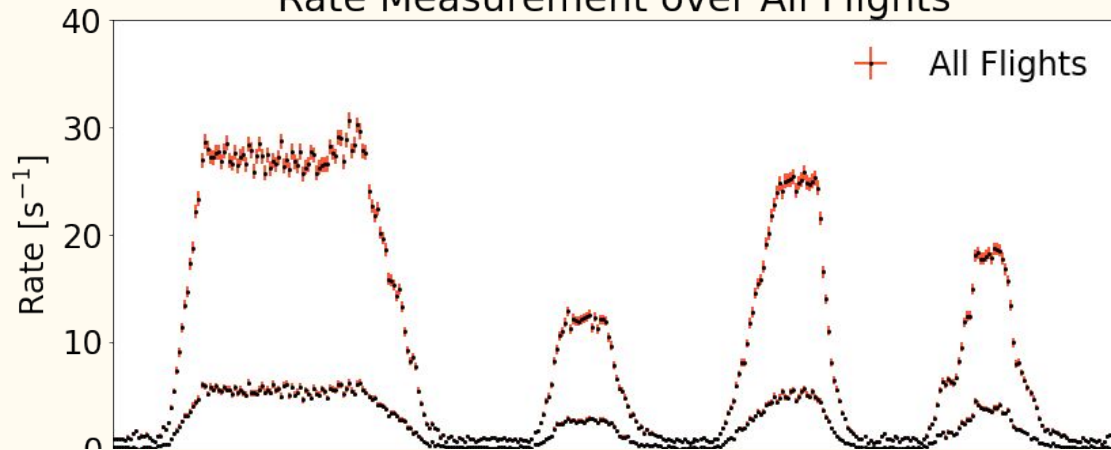
9.8344 \pm 0.0487Hz

2.0905 \pm 0.0224Hz

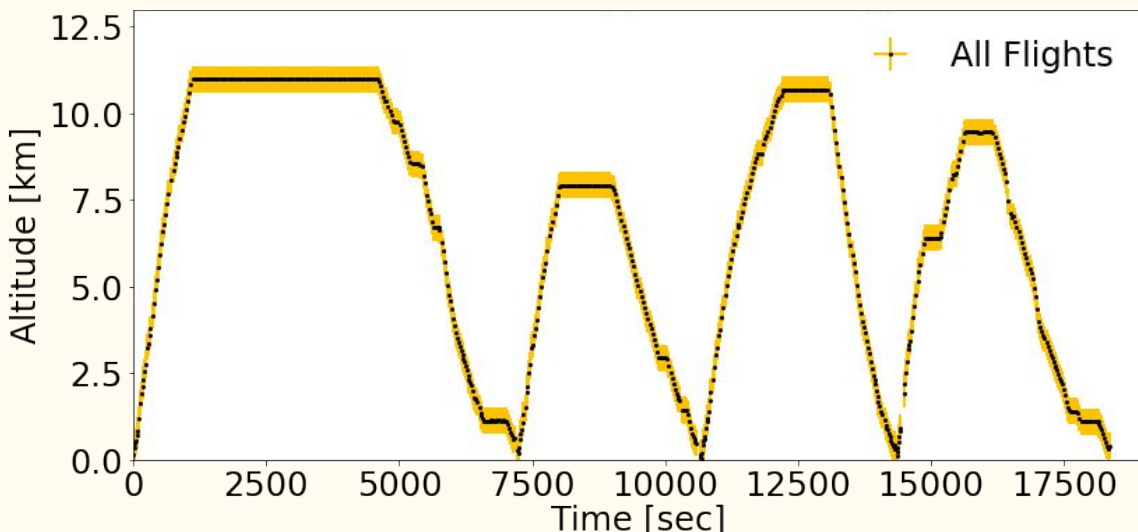
5.7157 \pm 0.0351Hz

1.2579 \pm 0.0165Hz

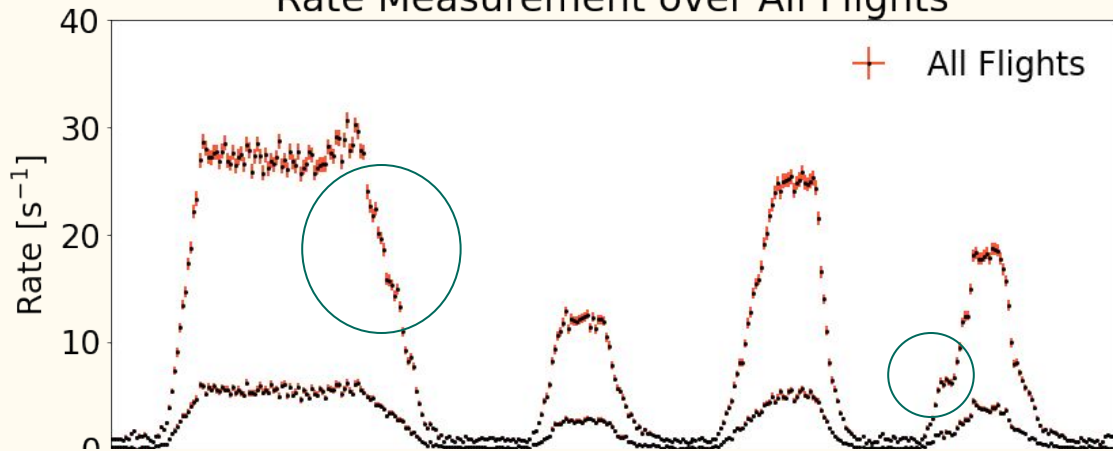
Rate Measurement over All Flights



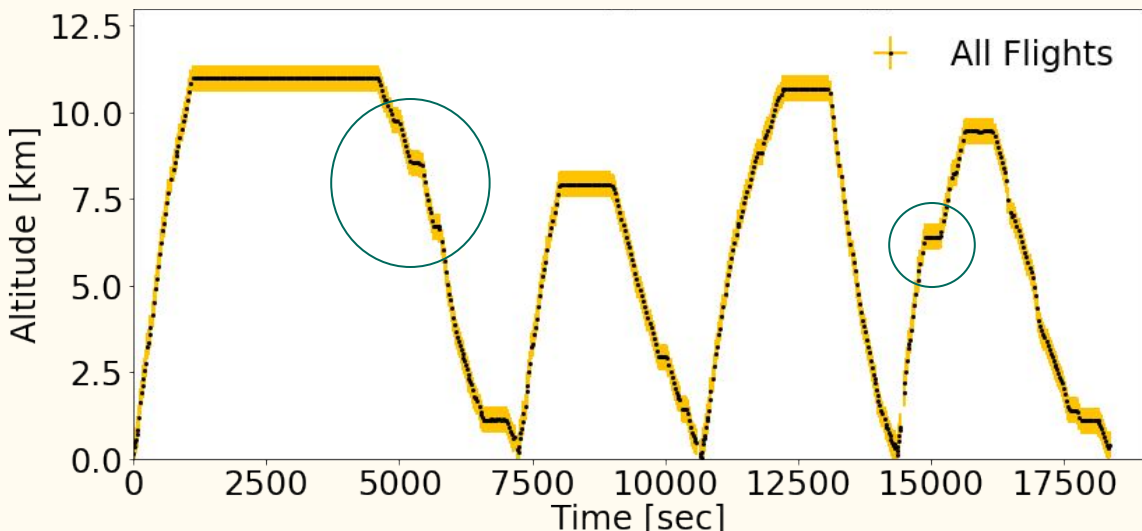
- Background radiation
- Decay altitude
- Similarities



Rate Measurement over All Flights



- Background radiation
- Decay altitude
- Similarities



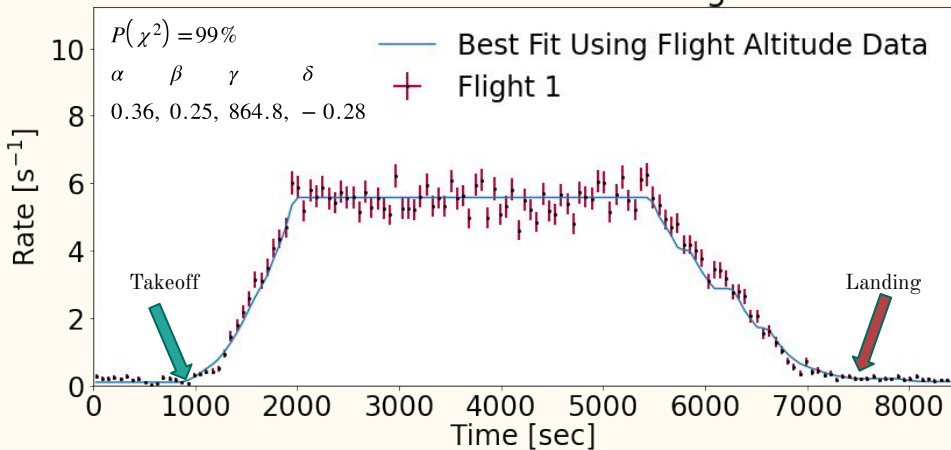
Connecting Altitude and Rate

$$Rate [Hz] = \alpha e^{\beta \cdot Altitude[t-\gamma]} + \delta$$

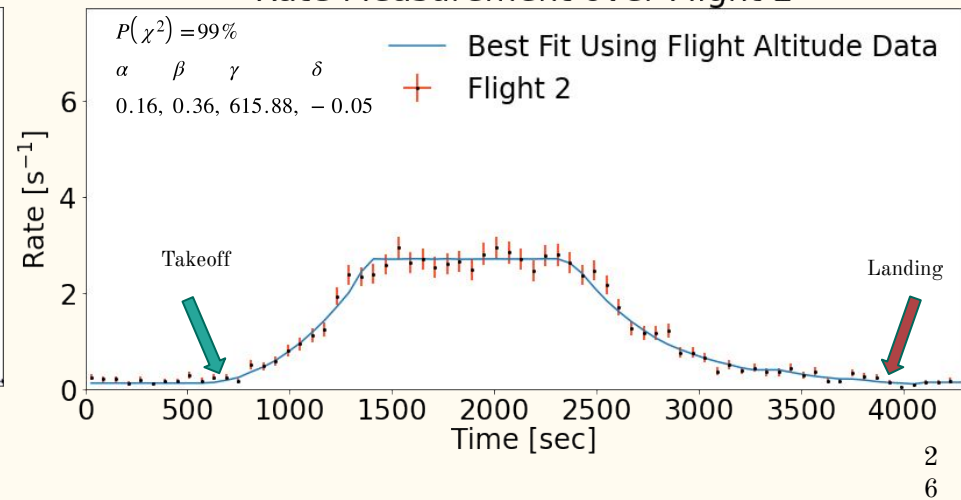
Altitude[t- γ] : Flight data taken from flightaware.com, taken at 30 second intervals, interpolated with cubic spline

γ factor to allow for fit to match up takeoff and landing time

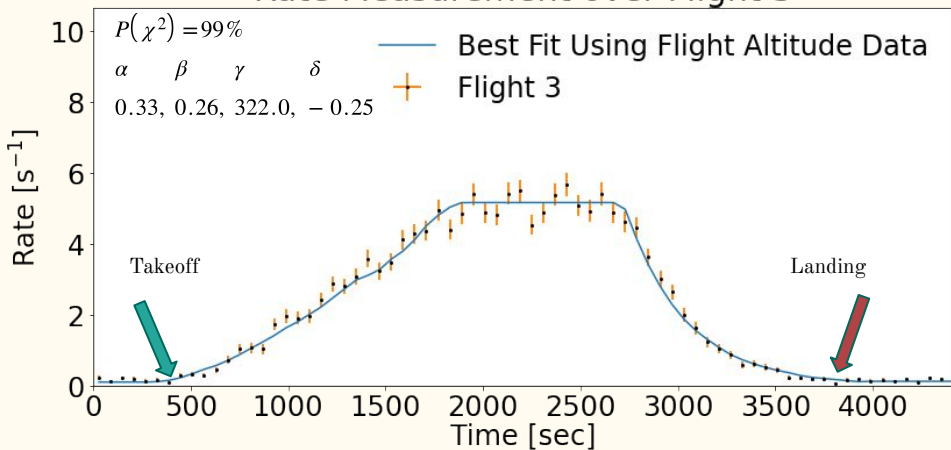
Rate Measurement over Flight 1



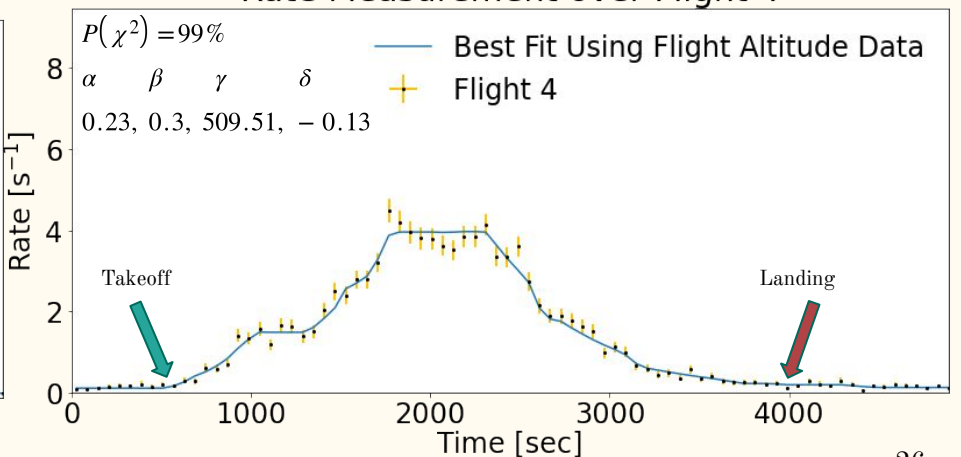
Rate Measurement over Flight 2



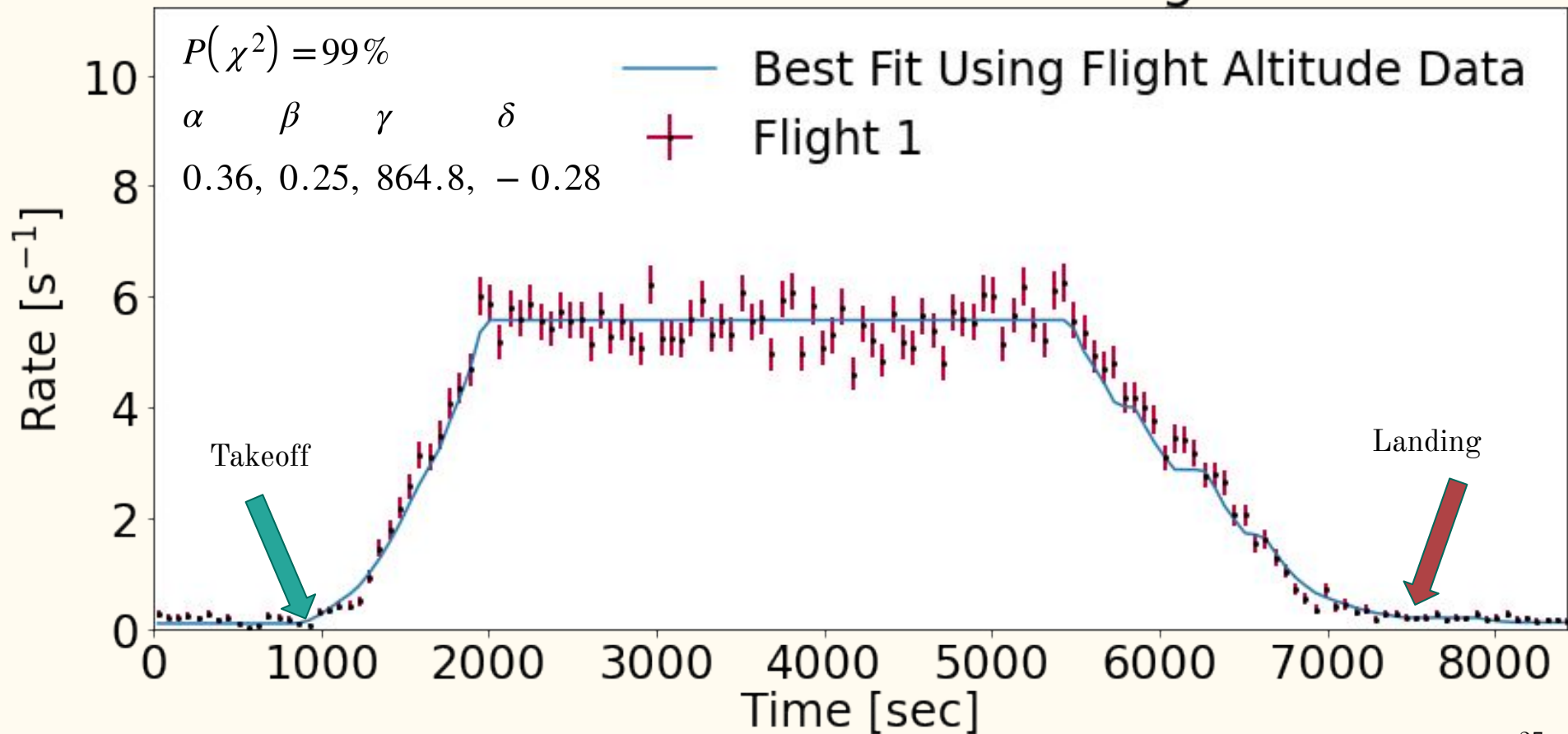
Rate Measurement over Flight 3



Rate Measurement over Flight 4



Rate Measurement over Flight 1

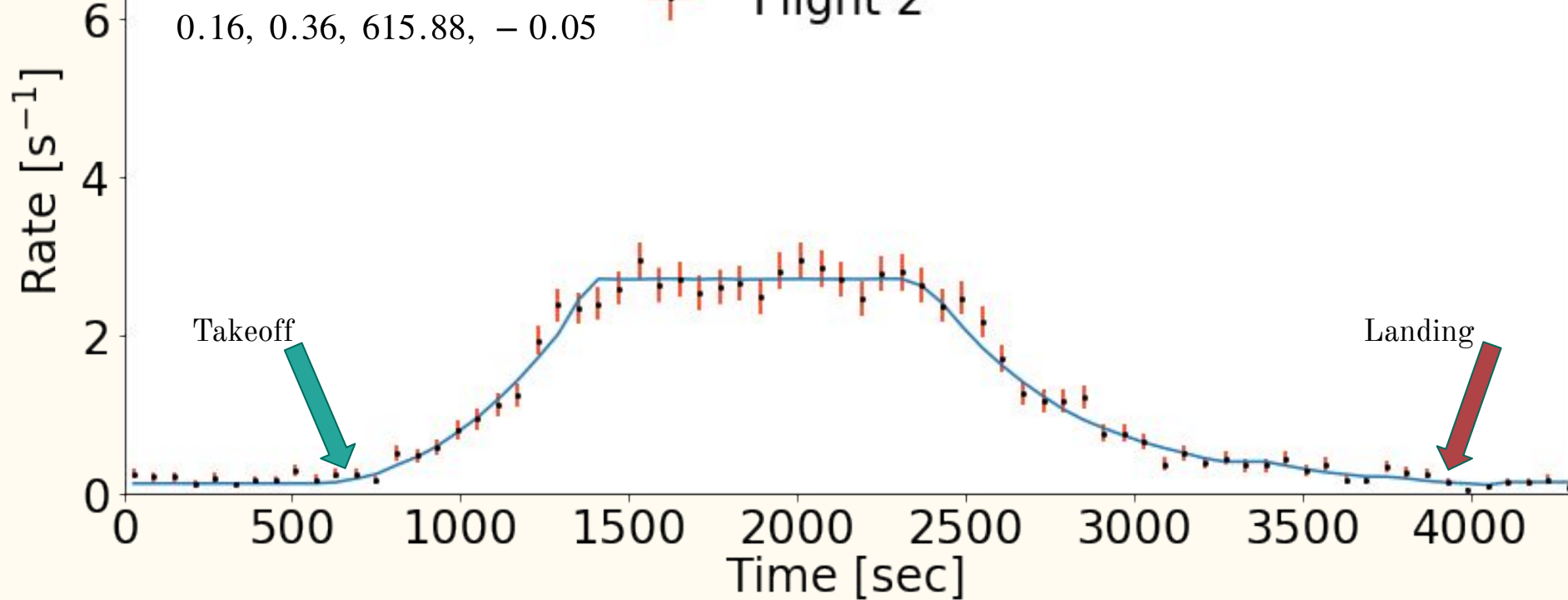


Rate Measurement over Flight 2

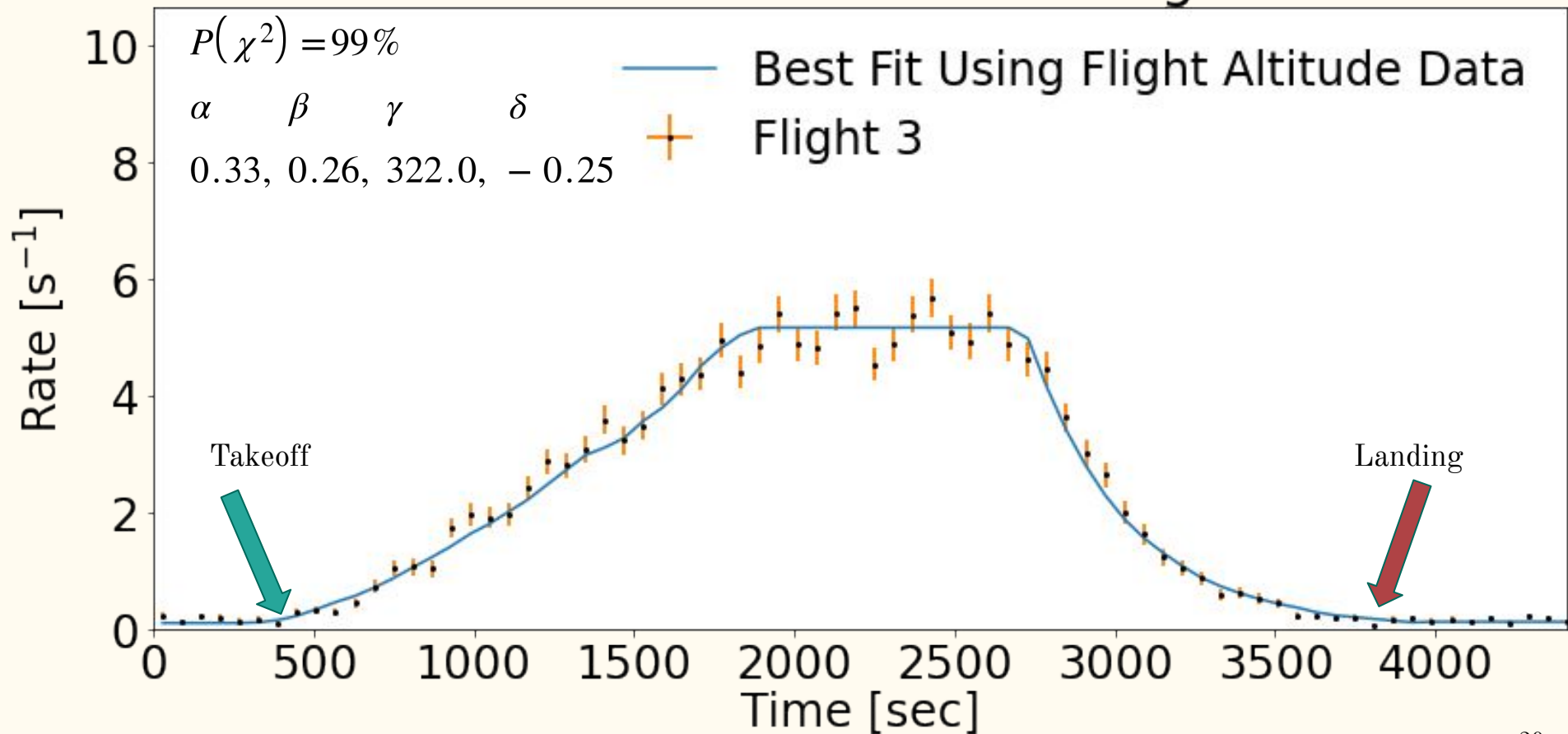
$$P(\chi^2) = 99\%$$

α β γ δ
0.16, 0.36, 615.88, -0.05

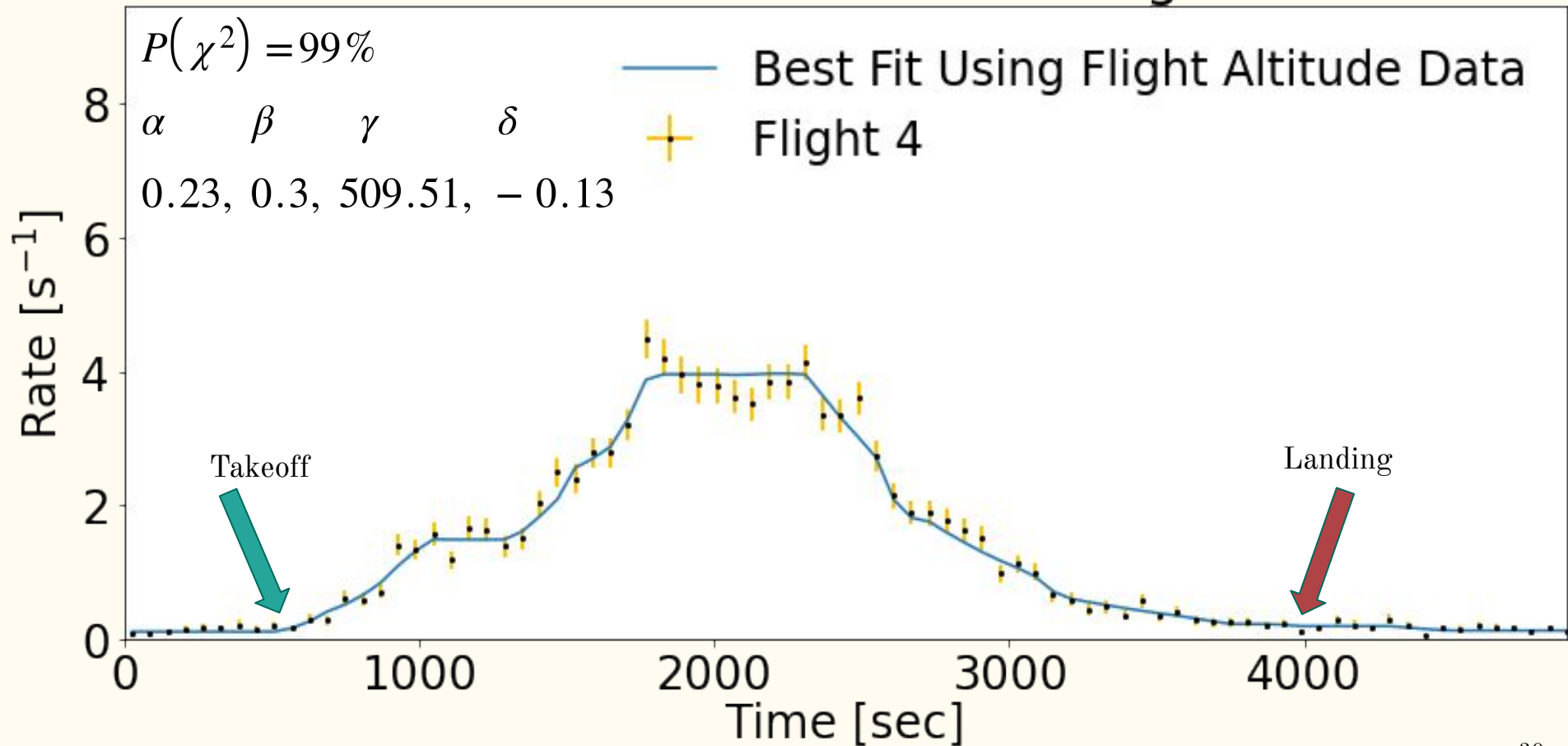
— Best Fit Using Flight Altitude Data
+ Flight 2



Rate Measurement over Flight 3



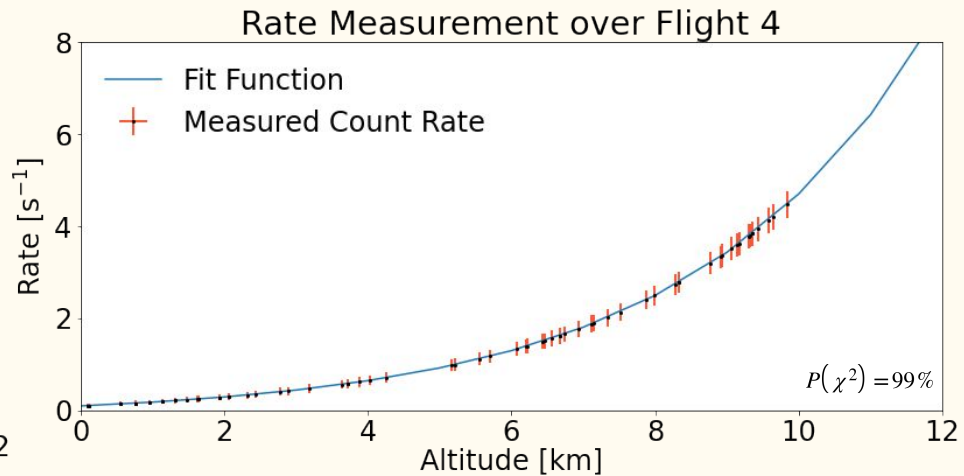
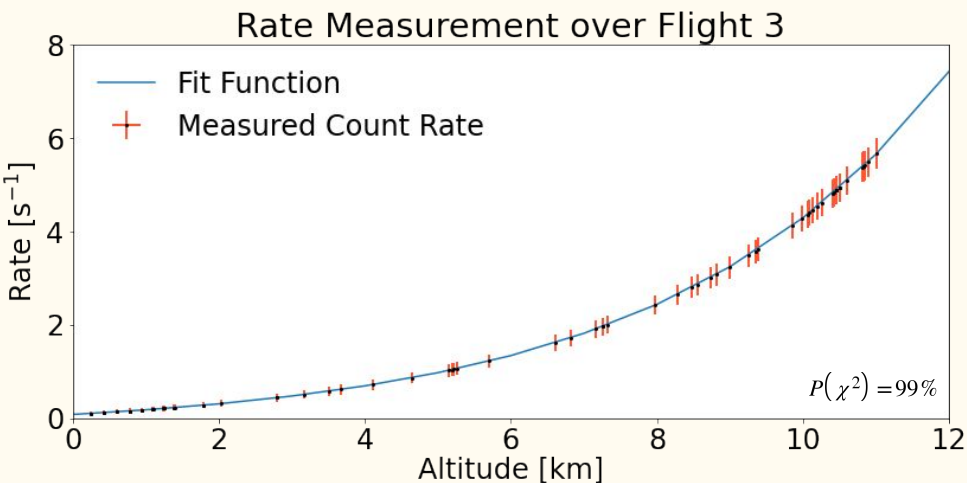
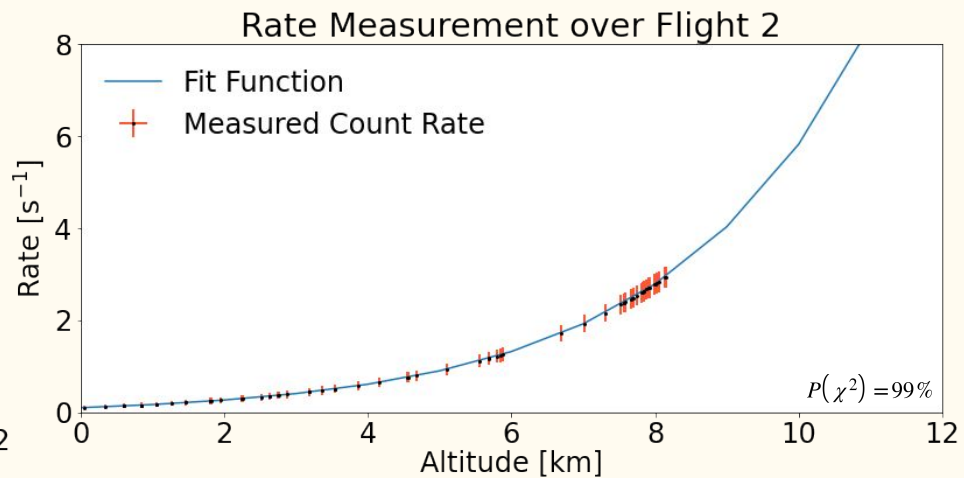
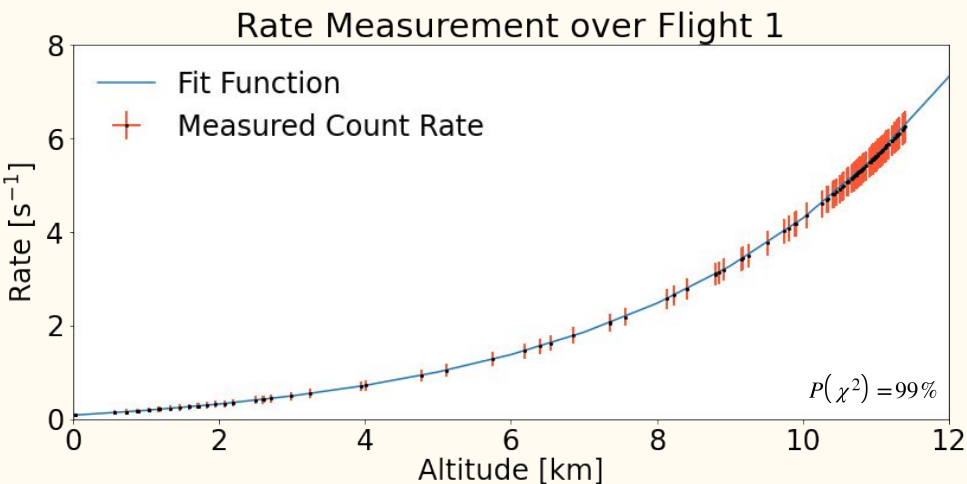
Rate Measurement over Flight 4



How well does the fit describe the data?

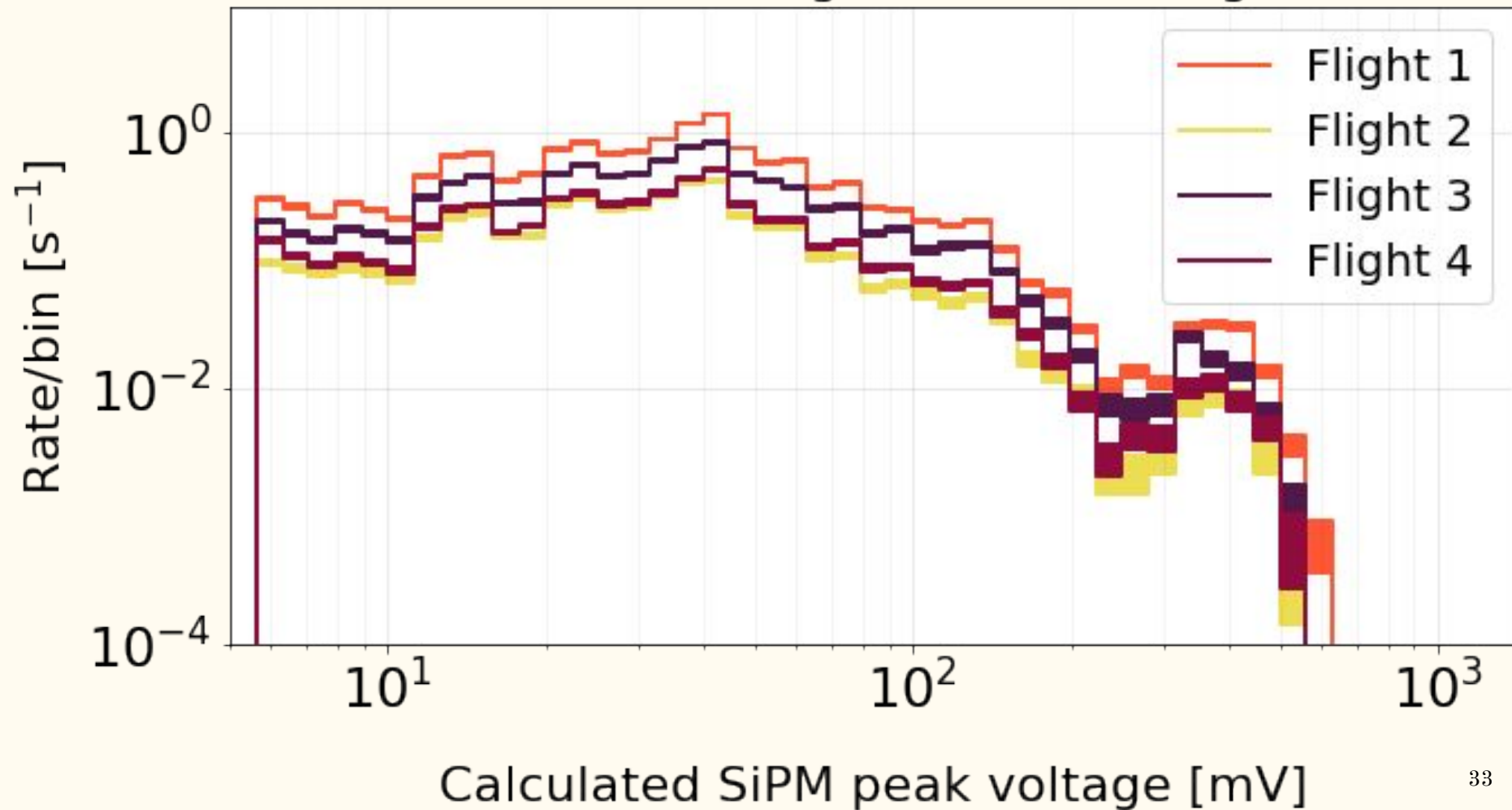
$$Rate [Hz] = \alpha e^{\beta \cdot Altitude[t-\gamma]} + \delta$$

- We have found α β γ δ for each flight
- Use our Cosmic Watch data to go from rate to altitude
- Discover the exponential relationship!



The count rate uncertainties from square root of the sum all the events measured at a given altitude

SiPM Peak Voltages over Four Flights



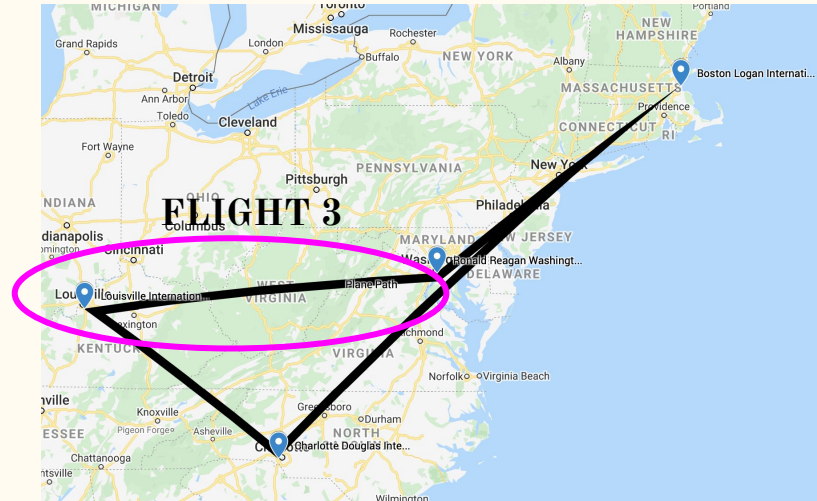
Uncertainties

The Latitude Effect

Uncertainties

The Latitude Effect

- Magnetic field behaves like dipole magnet
- Uniform across latitudes



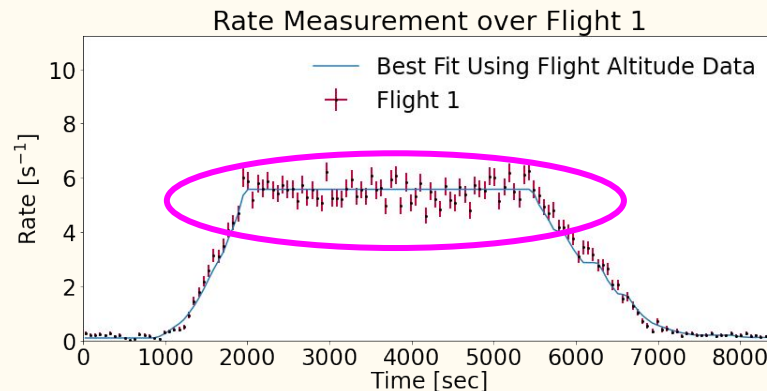
Uncertainties

The Latitude Effect

- Magnetic field behaves like dipole magnet
- Uniform across latitudes

Uncertainty Value: 5%

- Deviation of counts at constant altitude



Uncertainties

Weather

- Varies based on flight number, flight location, and flight altitude
- Pressure, temperature, etc.
- Alters the presence of background radiation

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- Varies based on flight number, flight location, and flight altitude
- Pressure, temperature, etc.
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Uncertainty Value: 6%

- Largest deviations in pressure along flight path

Conclusion

- Experimentally confirms that at higher altitudes we find more cosmic rays

(Flight 3)
$$Rate [Hz] = \alpha e^{\beta \cdot Altitude} + \delta$$

$$\alpha = 0.33 \pm 0.04 \quad \beta = 0.26 \pm 0.03 \quad \gamma = 322 \pm 41 \quad \delta = -0.25 \pm 0.03$$

- Four flights, two detectors, two days
- Relationship between counts, energy, and altitude

Future experiments:

- Investigate flight paths through repeat flights
- Use Cosmic Watch data to uncover magnetic anomalies

**Thank
You**

