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## The Newtonian Gravitational Constant

An Index of Measurements

George T. Gillies



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# The Newtonian Gravitational Constant

## An Index of Measurements

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### Foreword

So far as our experimental knowledge extends at present, the gravitational attraction between bodies is in accordance with general relativity. In all but the strongest fields, the law of force is the Newtonian inverse square law and the gravitational attraction of a body is proportional to its inertial mass. All bodies follow geodesic paths in four dimensional space, whatever their chemical constitution, so that gravitation appears to be a strictly geometrical phenomenon. Whether or not that is a correct statement has very far reaching implications for the nature of the physical world, and from Newton's day onwards, there have been many attempts to test the validity of Newton's law of gravitation. Despite the most careful analyses, it seems there is no deviation from it that can be found in celestial mechanics. The most careful and precise experiments on gravitation are those that establish the validity of the weak equivalence principle. Other very sensitive experiments have, in the past, excluded a range of effects, such as any of temperature, or crystal structure, or gravitational shielding. The status of experimental tests

of the inverse square law is less satisfactory, while the scatter of the best measured values of the constant of gravitation itself is still considerable. The results of experiments on the weak equivalence principle and the inverse square law have recently been put to question again by the suggestion that there is a "fifth force" between baryons that would look like a massive component of gravitation, showing deviations from weak equivalence and the inverse square law and it is claimed that such deviations can be detected.

There is then considerable interest in gravitation at the present time, with many experiments planned or under way.

Experiments on gravitation are difficult, and a great deal of work done in the past, some of it very careful and still valuable, is rather inaccessible.

For all these reasons, it is most helpful to have the comprehensive Index that Dr. Gillies has compiled, and the brief but valuable summaries of the main lines of experiment and their results that he has included. The earlier edition was most useful and one can be none other than very grateful for a new version. To Dr. Gillies and to the Bureau International des Poids et Mesures, our warmest thanks are due.

A. H. Cook, F.R.S.

\* Formerly with the Bureau International des Poids et Mesures where most of this work was done

**Abstract.** The Newtonian Gravitational Constant,  $G$ , has probably been measured more often but, interestingly, with less precision than any other physical constant of fundamental importance. In an effort that has spanned more than a century to connect gravitation to the other forces of nature, over 200 experiments on  $G$  have been completed and reported; but many of them have not been reported in what would now be considered to be the open literature. This paper\* is a third, more complete attempt to carry MacKenzie's and Poynting's bibliographies forward from the 1800's to the present; and thereby include as many as possible of the experimental results on  $G$  that have been obtained since 1900.

## I. Introduction

If one were to catalogue the tools of precision measurement, an unusually high number of the listings would claim as their genesis the precision measurement of the Newtonian Gravitational Constant, herein simply referred to as  $G$ . These tools would include the torsion balance, the optical lever, the quartz fiber, synchronous detection techniques, ultra-high-precision rotations and many others. Yet  $G$  stands alone as the only fundamental constant currently known to little better than one part in a thousand although there are three measurements claiming accuracies of one part in ten thousand. In parallel with these efforts to measure the absolute value of  $G$ , there has also been a wide variety of experiments aimed at linking the gravitational force to the other forces of nature. All such efforts to date have had the striking result of demonstrating that gravity, indeed, stands alone — the last of the great classical mechanisms — in spite of its modernized presentation via general relativity.

Classical gravitational physics has been like this, and foreseeably will continue to be like this. The reason why is that, to this date, no one has

succeeded in isolating sufficiently well the gravitational interaction between laboratory masses to the point where other disturbing forces or experimental uncertainties do not dominate the measurement, or at least mask other phenomena that might be expected to occur.

It was nevertheless interesting to catalogue the large body of work done already in the hope that a thorough listing of the experimental facts concerning our knowledge of  $G$  will stimulate future work on this constant and the force that it governs.

Part of the motivation for this paper lies in the fact that much of the work on  $G$  was reported obscurely even though most experiments have been carefully designed and completed. It was, therefore, a challenge to extract from various libraries, archives and private collections the existing data that, when collectively viewed, will help to focus attention on just exactly what has been done and, more importantly, what has *not* been done in this field.

This work is meant to be a bibliography and, at present, only that. Owing to the unusually large number of references cited, any thorough discussion of all the results would have taken up more space than was available here. Nevertheless, there is a small amount of annotation provided in the following pages for each of the fourteen sections of the bibliography. The areas into which the papers have been classified are listed below:

1. Measurements of the absolute value of  $G$  and reports of important subsidiary technology.
2. Comments and reviews of measurements of  $G$ .
3. Measurements of gravitational permeability, absorption and shielding.
4. Measurements of the local directive action of the gravitational force.
5. Measurements of the dependence of  $G$  on the physical state of masses.
6. Measurements of the dependence of  $G$  on the chemical state of masses.
7. Measurements of the dependence of  $G$  on temperature.
8. Measurements of the dependence of  $G$  on the radioactivity of masses.

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\* This is the third version of this Index. The second was issued by the Bureau International des Poids et Mesures as Rapport BIPM-83/1

9. Measurements of the dependence of  $G$  on the electromagnetic energy content of masses.
10. Measurements of the dependence of  $G$  on inter-mass spacing.
11. Measurements of the dependence of  $G$  on time.
12. Measurements of spontaneous matter creation (related to 11).
13. Measurements of the dependence of  $G$  on the state of quantization of the test masses.
14. Measurements of the anisotropies of  $G$  and of inertial masses.

There are over 1200 references cited in these 14 sections. There is some duplication, as a few of the papers contain two or more experimental results each in a different area. Duplicate listings constitute about 5% (or less) of the total, however. The references are listed alphabetically in each section with a chronological sub-listing for each author in each section. Abbreviations for the journal titles follow the American Institute of Physics *Style Manual* wherever possible. For those journals not listed there, the abbreviations in the ISI\* *Current Contents* indexes have been used. Those listings that have the name(s) of the author(s) marked with an asterisk have not been consulted at the time of this writing. References to entries in the various abstracting journals have not been given except for a few special cases where the abstracted article was judged to be published obscurely.

Not all of the listings in the Mackenzie (1900) and Poynting (1894) bibliographies were repeated here. Some, like those referring to the "Fr. Berrier" Controversy of the late 1700's/early 1800's, have little scientific merit and were omitted.

Sections 11 and 12 contain several references to instrumental papers in addition to those actually quoting significant results. In particular, the efforts at Princeton University, Massachusetts Institute of Technology and the University of Virginia are listed in detail.

Finally, as of this date, there are three new experimental thrusts in this field which have just recently begun. These are the search for a spin dependence in the gravitational force, the search for a fifth force in nature, and the measurements of the force of gravity on antimatter. When the experimental investigations in these areas have had time to develop, the appropriate results will be included in updates in this index.

## II. Experimental Studies and Critical Analyses of $G$

### 1. Measurements of the Absolute Value of $G$ and Reports of Important Subsidiary Technology

The history of the measurement of the universal gravitational constant begins with geophysical studies of a related physical quantity: the mean density of the earth. These efforts started with attempts to measure the attraction of individual mountains and measurements of strata of the earth's crust as a function of depth in various mines, and they are continuing with measurements of the attraction of layers of water in large level-controllable lakes and in the oceans.

Chronologically, the torsion-balance methods came next, and these gave the most reliable results until the advent of the "torsion-pendulum" technique. The balance-beam methods were studied during the late 1800s, and today all three of these methods are being developed in experiments aimed at accuracies of one in  $10^5$ .

Numerous miscellaneous methods have also been developed. These include resonant torsion pendulums, vertical pendulums, near-zone excitation of gravity-wave detectors, and long-period horizontal pendulums. There have also been several proposals for satellite determinations of  $G$ , but so far no such measurements have actually been made.

The early works on  $G$ , particularly those of Cavendish, Reich and Baily, have been reviewed frequently and the principal papers of these workers are summarized and paraphrased in most undergraduate textbooks on physics. There are, however, about 60 other determinations of  $G$  that are in the open literature. Some of these are well known too, particularly those of Boys, Poynting, Braun, Heyl and the Beams-Deslattes-Luther-Towler collaboration between the University of Virginia and the U.S. National Bureau of Standards.

Of the latest works, there are three that claim accuracies near one part in ten thousand. They are summarized in Table 1. Unfortunately, all of these results are mutually exclusive given the quoted errors. Therefore we must conclude that

**Table 1**

Author	Year	Experimental technique	$(G \pm \Delta G) \times 10^{11}$ $\text{m}^3 \text{kg}^{-1} \text{s}^{-2}$
Facy, Pontikis	1972	Resonant pendulum	$6.6714 \pm 0.0006$
Sagitov et al.	1979	Torsion pendulum	$6.6745 \pm 0.0008$
Luther, Towler	1982	Torsion pendulum	$6.6726 \pm 0.0005$

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**Table 2.**

Author	Year	Experimental technique	$(G \pm \Delta G) \times 10^{11}$ $\text{m}^3 \text{kg}^{-1} \text{s}^{-2}$
Preston	1895	Pendulum decrement	7.16
Barus	1922	Torsion balance	6.609 $\pm$ 0.020
Stern	1928	Resonant balance	6.6 $\pm$ 0.5
Renner	1974	Torsion pendulum	6.670 $\pm$ 0.008
Karagioz et al.	1976	Torsion pendulum	6.668 $\pm$ 0.002
Koldewyn, Faller	1976	Torsion pendulum	6.57 $\pm$ 0.17
Spero	1979	Nulled torsion balance	6.70 $\pm$ 0.19
Page, Geilker	1981	Torsion balance	6.1 $\pm$ 0.4
Karagioz et al.	1981	Torsion pendulum	6.5912 $\pm$ 0.0016
Speake	1983	Beam balance	6.64 $\pm$ 0.24
Oelfke	1984	Torsion balance	6.7 $\pm$ 0.2

the value of  $G$  is not, in fact, well established at the  $10^{-4}$  level.

All uncertainties quoted in Table 1 as well as that in the CODATA value below represent one standard deviation.

Even if ultimately measured in a drag-free satellite, where external horizontal gravity gradients would not influence the balance, measurements of  $G$  would still be limited by the uncertainties arising from density gradients in the materials used. This would probably occur somewhere between the  $10^{-5}$  and  $10^{-6}$  level. At that point, a totally different approach to the measurement of  $G$  will become necessary.

The 1986 CODATA value, assigned by E. R. Cohen and B. N. Taylor, is

$$G = (6.67259 \pm 0.00085) \times 10^{-11} \text{ m}^3 \text{kg}^{-1} \text{s}^{-2}.$$

Mills (1979 – see Section 10) has tabulated the results of the 22 most often cited measurements of  $G$ . There are several other works which are less well known but important nonetheless, and the results from these are entered in Table 2. The most recent experiments listed there should be considered as work in progress, and not as final results. Also, some of these measurements were not directed at  $G$  itself but at a search for some variation in  $G$ , with the quoted result being a by-product. As a rather interesting aside, it should be noted that, since so many of the earliest measurements (those by Cavendish, Reich, Baily, among others) were done for the purpose of determining the mean density of the earth, this motivation has remained the classic justification for undertaking measurements of  $G$ . With a knowledge of the earth's density and its volume, one may estimate its total mass,  $M_e$ , in kilograms. More recently, satellite ranging experiments have given us the “geocentric” gravitational constant directly:  $G_g = GM_e$ . The latest of these experiments, Ferrari et al. (1980), has an uncertainty of only about

$7 \times 10^{-8}$ . This means that any increased accuracy in our knowledge of  $G$  will automatically give us  $M_e$  with a corresponding precision, and vice versa. A few of the recent measurements are catalogued in Table 3.

## 2. Comments and Reviews of Measurements of $G$

There are four thorough reviews of the measurements of  $G$ : two dating to the late 1800s and two more recent ones. Poynting and, later, MacKenzie summarized the contemporary knowledge of  $G$  and, for that matter, all of gravitational physics, in books printed in 1894 and 1900, respectively. Sagitov (1969) recently published a similar work on  $G$ , although he omitted many references to experiments probably judged by him to be of secondary importance. Most recently, de Boer (1984) has contributed a review article which catalogues the recent major experiments and presents the results and uncertainties together on easily readable graphs. There are several other reviews of the experiments, and these are listed in the bibliography. It should be mentioned that the article on “Gravitation” in each successive edition of the Encyclopedia Britannica contains interesting and relatively thorough sections on  $G$  that are very useful.

In addition to the review articles on  $G$ , there are a large number of papers, both old and new, that comment on or discuss certain measurements

**Table 3.**

Author	Year	Satellite	$(G_g \pm \Delta G_g) (\text{km})^3 \text{s}^{-2}$
Esposito	1979	Viking I	398,600.5 $\pm$ 0.1
Esposito	1979	Viking II	398,600.65 $\pm$ 0.20
Martin, Oh	1979	ATS 6	398,600.36 $\pm$ 0.12
Ferrari et al.	1980	Lunar Orbiter 4	398,600.461 $\pm$ 0.026

of  $G$ , propose new measurement techniques, or analyze probable experimental limitations. Some of these are worth consulting, since errors have occasionally appeared in the main papers of the principal investigators. The titles of the papers in this section of the bibliography usually indicate the relevance to a certain experiment or class of experiments.

### *3. Measurements of Gravitational Permeability, Absorption and Shielding*

Although not widely known, one of the most thoroughly researched aspects of gravitational physics is the question of the existence of a gravitational analogue to magnetic permeability. The pursuit of this question, i.e., the dependence of  $G$  on the density of the matter intervening between the interacting masses, began with a null result in the late 1800s. It continued until recent times, always with null results, but with ever-increasing accuracy. We now know that if gravitational energy is, in fact, absorbed by any intervening material, it occurs at a level such that, when measured in terms of  $G$ , the result is  $(\Delta G/G)_{\text{absorp.}} \leq 10^{-16}$ . Some theoretical analyses establish a lower limit several orders of magnitude below this, too.

At first these measurements were made by employing a Cavendish balance with a cylindrical screen separating the suspended dumb-bell from the attracting masses. Each screen was made of a different type of material and all the materials had different densities. The screens were sequentially changed and the measurements of  $G$  subsequently made were analyzed for a resulting effect. Later, Majorana began a long series of experiments using a balance-beam and claimed to have found a result at the  $10^{-11}$  level, but subsequent work by himself and others disproved this. The most sensitive measurements have been made in recent times by several workers studying the period shift in horizontal pendulums during a total solar eclipse. In every case, however, the results have been null except for the results of Allais and, later, Saxl and Allen which are seldom discussed.

Unless a new theory predicts a permeability effect substantially different from that empirically sought so far, it is difficult to see where the motivation would arise for new experiments in this area. This is especially so since recent gravitational analogues of electromagnetism predict that the "gravitational permeability of free space"

is only  $\frac{16\pi G}{c^2} \approx 10^{-26} \text{ m/kg}$  (Forward 1961), a very small effect indeed!

### *4. Measurements of the Local Directive Action of the Gravitational Force*

Since so many properties of crystalline materials depend upon the direction of observation (e.g. refractive index, local density distribution, thermal conductivity, etc.), it seemed reasonable to question the constancy of  $G$  within crystalline materials as well. This was first done by MacKenzie, then later by Poynting, although no anisotropy in the value of  $G$  was found in either case at  $\approx 10^{-3}$  and  $10^{-4}$  levels, respectively. Heyl did an exhaustive experiment in this area in 1924 and, by weighing crystals from each of the five non-isometric groups, was able to put a limit of  $(\Delta G/G)_{\text{anisot.}} \approx 10^{-9}$ , thereby effectively eliminating any doubt about it.

It has been suggested that one should not expect to find such anisotropies in general because they would lead to a violation of conservation of momentum. Nevertheless, the weakness of the gravitational force and the singular properties of some crystals made this, temporarily at least, an attractive area of research.

### *5. Measurements of the Dependence of $G$ on the Physical State of Masses*

This category is a rather general one, although it is possible to classify various types of experiments within it. Specifically, there have been measurements of  $G$  involving test masses and attracting masses of various geometrical shapes; for example, spheres, cylinders, rods, rings, and irregular masses have been used. In fact, a cylindrical configuration having the sphere-like field of a point source is presently under study. Attracting and test masses in the gaseous, liquid and solid state have been (or are being) studied, as well as masses which undergo a change of state during the experiment. In all cases, no departure from true constancy of  $G$  has been observed, at least within the levels of the experiments' precision. When one considers that the sun, a plasma, and its planets (gas, solid and liquid combinations) have orbits which conform precisely to the inverse square law (with only minor relativistic corrections), it is seen that this law is indeed well obeyed. The references cited here are representa-

tive examples of various types of experiments involving gases, liquids and solid masses of various shapes. Some of these citations are listed in other more appropriate sections of the bibliography, too.

#### *6. Measurements of the Dependence of $G$ on the Chemical State of Masses*

This area of research is usually interpreted as being a test of the weak equivalence principle of General Relativity, i.e. a determination of the equivalence of gravitational and inertial mass. The various searches for a non-zero Eötvös ratio which test this equivalence are catalogued by Will (1981) and are not included here, except for a few special cases given below.

The earliest measurements in this area involved pendulums of the same mass but made of different materials. Eötvös and his contemporaries expanded this to include torsion balances which had masses of different materials attached to them. His null result has been confirmed by workers at Princeton University, Moscow University and the Joint Institute for Laboratory Astrophysics, Boulder.

Several workers studied the interesting question "Does  $G$  vary while the test mass undergoes a chemical reaction?" but, in all cases, equilibrated or reacting, null results have been obtained. This area has not been without controversy, however. C. F. Brush claimed to see a difference in pendulum periods between pendulums made of bismuth and those of zinc. His observations were ultimately explained by his failure to include the buoyancy of air in his calculations.

Although not strictly belonging to this category, the experiment of L. Kreuzer (1966), at Princeton, is included here because it has been interpreted by some as providing evidence for a variation of  $G$  over the elements of the periodic table. This claim was subsequently refuted, though.

#### *7. Measurements of the Dependence of $G$ on Temperature*

Early in this century there were three attempts to measure with balance-beams a dependence of  $G$  on the temperature of the attracting masses. They all produced null results but were nonetheless open to discussion. This was so for two reasons. First, there were well-known difficulties in experiments wherein two masses at different tem-

peratures were used in a high-precision balance; and second, the unknown thermal profile of the earth immediately below the balance would seem to make an exact repetition of the experiment impossible, because any thermal dependence of gravity would presumably affect the earth's field too (and how can one control the temperature of the earth?).

Professor Shaw and his students, therefore, constructed a Boys-type balance in which the attracting masses could be heated. They measured  $G$  and analyzed their data in the form  $G = G_0(1 + \alpha T)$ . At first they found  $\alpha \approx 10^{-5} \text{ K}^{-1}$ . This result created great interest and several papers were written discussing it. A careful repetition of their original work showed, though, that after removing troublesome convection effects,  $\alpha \leq 2 \times 10^{-6} \text{ K}^{-1}$ , which effectively settled the issue in the negative. There do not seem to have been any other experiments following this one. Perhaps this is because it was realized that the large temperature difference between the earth and sun, when compared with the much smaller difference between the earth and moon, should have highlighted the existence of the effect. Instead, both orbits serve only to verify that the inverse square law is, in fact, independent of temperature. In questions of this type, however, scale factors may be important; and even though the balance-beam experiments (laboratory scale) in the earth's field produced null results at the  $10^{-9}$  level, Shaw's experiment with isolated gravitational forces (also at laboratory scale) was 2000 times less sensitive. So there may still, perhaps, be effects to be discovered.

The results of all the above experiments are given in Table 4.

**Table 4.**

Author	Year	Technique	$ \Delta G/G  \text{ K}^{-1}$
Count Rumford	1785	Beam balance	$\leq 3.3 \times 10^{-8}$
Poynting, Phillips	1905	Beam balance	$\leq 10^{-10}$
Southerns	1906	Beam balance	$\leq 10^{-8}$
Pettersson	1914	Beam balance	$\leq 10^{-9}$
Shaw	1916	Torsion balance	$\sim 10^{-5}$
Shaw, Davy	1923	Torsion balance	$\leq 2 \times 10^{-6}$

#### *8. Measurements of the Dependence of $G$ on the Radioactivity of Masses*

A very interesting area of research in this century has been the search for a connection between gravitation and what we now know to be the

nuclear forces. Studies in this area have involved a wide variety of experimental techniques. Several famous physicists have worked in this area, including G. Sagnac, P. Zeeman, A. Compton and J. J. Thomson, each lending his own special expertise to this difficult problem.

Compton, for instance, produced a large and controllable pseudogravitational (centrifugal) field by rotating samples of radium at high speeds. He concluded that this did not affect the radioactivity by more than  $10^{-3}$ . Thomson's uranium-pendulum experiment (carried out in more detail by Southern) showed no gravitational coupling to the sample's radioactivity at the level of  $5 \times 10^{-5}$ . All other measurements also produced null results, except for that of R. Geigel who claimed to have seen a small weight change in a non-radioactive sample hanging on a sensitive balance when a radium salt sample was placed nearby. He interpreted this as an absorption of radioactivity which led to an increase in the gravitational potential energy of the non-radioactive sample. W. Kaufmann shortly thereafter, however, uncovered a thermal effect which explained the apparent weight change. No one since then has repeated Geigel's experiment.

There is one modern experiment, by Whitaker, which finds a shift in the decay rate of  $^{198}\text{Au}$  when subjected to a centrifugal field of 196,000 times gravity at the earth's surface. He gives no evaluation of systematic effects in his measurement system which might mimic such an effect, however.

### *9. Measurements of the Dependence of $G$ on the Electromagnetic Energy Content of Masses*

At the same time as the searches for a radioactivity-gravitation coupling were under way, there was a parallel effort in progress aimed at finding an electromagnetism-gravitation coupling.

These searches typically involved weighing samples of steel and permanent magnets in magnetized and unmagnetized states or charged and uncharged states.

The very much larger size of the electromagnetic forces always makes experiments of this kind very difficult, and careful attention must be paid to the shielding of undesirable electromagnetic couplings to the laboratory which could otherwise make the results ambiguous. In spite of this, null results were always reported.

J. F. Woodward has proposed a repetition (with much higher precision) of the early work of

Faraday and Blackett in search of an electrogravitational effect, but it is not clear at this point what a positive result in these experiments would mean in terms of the value of  $G$  or the gravitational inverse square law. The reference section of his paper refers the reader to citations of earlier work along that specific line.

### *10. Measurements of the Dependence of $G$ on Inter-mass Spacing*

If a scale dependence other than pure inverse-square-law behavior were found for gravity (in the weak-field limit), such an effect would have a significant impact on current efforts to unify the four forces of nature. Because of the importance of this point, several workers have designed and executed experiments to search for non-Newtonian gravity. Although the results are routinely expressed in terms of a distance dependence for  $G$ , i.e.,  $G(R)$ , it is also standard to express such findings in terms of a departure of the force from a  $R^{-2}$  dependence of the form  $1/R^{2+\delta}$ . A unified approach to all these analyses has been made popular by Gibbons and Whiting who plot "allowed" and "disallowed" regions in a two-parameter space established by the scale factor and strength parameters in a Yukawa type of add-on gravitational potential.

The experiments to date have been done on many scales of distance, ranging from 2 cm up to several kilometers. There have been analyses of the free oscillations of the earth in terms of how they might be influenced by a non-zero  $G(R)$ ; and on the still larger scale, the motion of the planets has been seen to confirm the inverse square law to an amazingly high precision (except for the previously-mentioned relativistic corrections).

The majority of the latest experiments have been designed to be high-sensitivity null measurements, although the early work of Mackenzie and, to some extent, that of Long instead involved measurements of the absolute value of  $G$  at two or more different mass spacings. While constituting a workable scheme in principle, the absolute measurements are usually burdened with large drifts and metrological difficulties. This sometimes makes their results open to question.

Stacey and his colleagues at the University of Queensland in Australia have undertaken a series of geophysical measurements in search of a non-Newtonian component in the earth's gravitational field. A reanalysis of existing gravity data, Stacey (1983), has motivated them to do this, and their work is currently in progress.

Tests of the superposition principle as applied to gravitational fields also belong in this category. Work in this area was done by G. Luther at the U.S. National Bureau of Standards, and also by P. Czipott and J. Goodkind at the University of California, San Diego.

Finally, reviews of the  $G(R)$  measurements have been given by Hoskins et al. (1984) and by Newman (1983) wherein the results have been tabulated and thoroughly discussed. Those interested in the overall status of this type of experimental search for non-Newtonian effects are encouraged to consult these works.

### *11. Measurements of the Dependence of $G$ on Time*

Perhaps no other area of gravitation is of greater interest to theorists and cosmologists than the possibility of variations of  $G$  in time. There have been many theories calling for a time-varying  $G$ , each having its own implications on the behaviour of matter and radiation at the early moments of the universe.

Here again, the experiments fall into three categories: laboratory, geophysical and astronomical. If such an effect exists, it is agreed that it must be very, very small; probably of the order of  $\dot{G}/G \approx 10^{-11} \text{ a}^{-1}$ . Few laboratory tests of any kind have been done at this level, and it is not surprising that the existing laboratory tests of this effect are limited at the level of  $10^{-7} \leq \dot{G}/G \leq 10^{-8} \text{ a}^{-1}$ . Several experiments with test masses have been proposed for both earth-surface and orbiting laboratories which should be sensitive to

the predicted  $10^{-11} \text{ a}^{-1}$  changes, but none has yet been completed.

The geophysical tests have usually involved studies of the expansion of the earth or investigations of a paleobiologic type. While these provide corroborating evidence, they are not usually accepted as hard proof because of the many uncertain factors involved. Wesson's contributions are the most complete in this area, and his book "Cosmology and Geophysics" (1978) should be consulted for a thorough review of geophysical investigations of  $\dot{G}/G$ .

Van Flandern (1981, 1984) claimed that all three astronomical tests of  $\dot{G}/G$  are now yielding similar results; and that  $\dot{G}/G \cong -(6 \pm 2) \times 10^{-11} \text{ a}^{-1}$ . Hellings et al., however, reported in 1983 that their solar-system test of  $\dot{G}/G$  had a result of  $(0.2 \pm 0.4) \times 10^{-11} \text{ a}^{-1}$ , a value consistent with zero. These tests include: (1) lunar laser ranging, (2) radar ranging of the inner planets, and (3) lunar occultation studies and determinations of the moon's orbit. The data from these three experiments have been coming in for several years now and have been carefully analyzed by several workers. The results for  $\dot{G}/G$  from each of them have great importance, as each method poses an independent check on the other. Ritter and Gillies (1986) have pointed out, though, that such astronomical tests may be intrinsically null and that a laboratory test is, indeed, needed.

The results from other sources, including laboratory experiments and those derived from cosmological considerations, are listed in Table 5. Although these areas are, at the moment, of secondary importance to the astronomical tests listed above, it is likely that sufficiently accurate

**Table 5.**

Author	Year	Technique	$ \dot{G}/G  \text{ a}^{-1}$
Hoffmann	1962	Quartz-pendulum gravimetry	$\leq 4 \times 10^{-8}$
Curott	1965	Quartz-pendulum gravimetry	$\leq 6.2 \times 10^{-7}$
Weiss, Block	1965	Quartz-spring gravimetry (drift limit, no firm result quoted)	$\leq 3.6 \times 10^{-6}$
Stephenson	1967	Reanalysis of Heyl's $G$ measurements	$\sim 10^{-4}$ (periodic annually)
Newton	1968	Earth spin-down	$\sim 10^{-10}$
Morganstern	1972	Flat-space cosmology analysis	$\sim 10^{-11}$
Morganstern	1972	Closed-space cosmology analysis	$\sim 3 \times 10^{-11}$
Eichendorf, Reinhardt	1977	Study of variations in the surface temperature of the earth	$(2.3 \pm 0.6) \times 10^{-11}$
Barrow	1978	Cosmological considerations	$(1.5 \pm 0.7) \times 10^{-12}$
Yang et al.	1979	Analysis of nucleosynthesis data	$5 \times 10^{-13}$
Lambeck	1979	Earth spin-down	$(2.5 \pm 0.5) \times 10^{-11}$
Lapiedra, Palacios	1981	Planetary-orbit studies	$\leq 7.5 \times 10^{-13}$
Rothman, Matzner	1982	Reanalysis of nucleosynthesis data	$\leq 1.7 \times 10^{-13}$

data from all areas will have to be available and in agreement before the existence of a non-zero  $\dot{G}/G$  is accepted.

## 12. Measurements of Spontaneous Matter Creation

Most of the theories that call for a non-zero  $\dot{G}/G$  also require the spontaneous creation of matter. This is usually the result of a gauge condition; or is in response to satisfying some phenomenological requirement, such as maintaining constant density in the universe. In a Machian universe, the value of the gravitational constant and processes like matter creation are presumably coupled in such a way that the value (or existence) of one affects the other. Therefore, it seemed appropriate to include the known experimental tests, tests in progress, and proposed tests of this effect, too.

In terms of categories, the  $\dot{M}/M$  experiments are classifiable in the same way as the  $\dot{G}/G$  experiments. There are substantially fewer of them, and only one laboratory experiment, that of S. Cohen and J. King, has yielded a result which was null at the  $\dot{M}/M = 10^{-23} \text{ s}^{-1}$  level.

## 13. Measurements of the Dependence of $G$ on the State of Quantization of the Test Masses

There have been two proposals for measuring  $G$  in terms of  $h$ , the Planck constant. These experiments, if ever done, will be the first direct tests of a quantum structure of the gravitational field. An indirect test by D. Page and C. Geilker has been carried out but the result is disputed.

The force of gravity on fundamental particles has also been measured and, in particular, the behavior of neutron and atomic beams in the field of the earth has been studied. In addition, there is

an effort by Melissinos and colleagues to study Newtonian gravity for matter (particles) travelling at relativistic speeds in accelerators.

As progress is made towards a quantized theory of gravity, more experiments of the type in this category will be needed and it may eventually be possible to study the gravitational force on a microscopic basis.

## 14. Measurements of the Anisotropies of $G$ and of Inertial Mass

The Hughes-Drever class of experiments all produced results which, to very high precision, showed that matter is evenly distributed in the universe. Dicke (1961) showed that they should be null measurements in principle, however.

Somewhat later, this question of large-scale anisotropy in the universe arose again, this time, though, in the context of a variation of  $G$  with respect to direction on a universal scale. Warburton and Goodkind (1976) analyzed earth-tide data taken with a superconducting gravimeter and found an effect not inconsistent with such a variation, but could not verify it due to uncertainties in the structure of the earth tides.

A room-temperature, feedback torsion balance has been constructed at the Cavendish Laboratory for a similar investigation, but no results have yet been reported.

*Summary of Section II.* For the most part, purely theoretical papers that discuss the possibility of variations in  $G$  (but do not set limits on it) have not been included in this bibliography. The goal here was to describe instead the "hard facts" so that all workers in the field could more easily assess the state of our knowledge.

The results of the different searches for variations in  $G$  are summarized in Table 6. Usually, the

Table 6.

Effect	Author	Year	Result ( $\Delta G/G$ )
Gravitational permeability or absorption	M. Caputo	1962	$\leq 6 \times 10^{-16}$
Directive action of gravitational force	P. R. Heyl	1924	$\leq 10^{-9}$
Dependence of $G$ on physical and/or chemical states of matter	For a review, see C. Will	1981	$\leq 10^{-12}$
Temperature dependence via a Cavendish balance	P. E. Shaw, N. Davy	1922	$\leq 2 \times 10^{-6} \text{ K}^{-1}$
Temperature dependence via a common balance	J. H. Poynting, P. Phillips	1905	$\leq 10^{-10} \text{ K}^{-1}$
Gravitation/radioactivity coupling	P. Zeeman	1918	$\leq 5 \times 10^{-8}$
Dependence of $G$ on magnetization of matter	M. G. Lloyd	1909	$\leq 5 \times 10^{-12} \text{ gauss}^{-1}$
Dependence of $G$ on electrification of matter	L. Simons	1922	$\leq 1.2 \times 10^{-7} \text{ V}^{-1}$
$G(R)$	R. E. Spero et al.	1980	$(1 \pm 7) \times 10^{-5}$
$\dot{G}/G$ (Astronomical)	R. W. Hellings et al.	1983	$(0.2 \pm 0.4) \times 10^{-11} \text{ a}^{-1}$
$\dot{G}/G$ (Laboratory)	W. F. Hoffmann	1963	$\leq 4 \times 10^{-8} \text{ a}^{-1}$
$\dot{M}/M$ (Laboratory)	S. Cohen, J. King	1969	$\leq 4 \times 10^{-23} \text{ s}^{-1}$

result claiming highest precision is quoted, although where appropriate a result typical of those in its class is given instead.

Some results entered in Table 6 are dimensionless. In these cases, the authors listed had tried to set limits on some appropriate dimensionless scale factor. Their original results (where necessary) have been interpreted in terms of  $\Delta G/G$  for presentation here.

### III. Closing

As mentioned previously, this work is meant to be a resource bibliography, not a critical review of the status of each of the various areas treated here. It is hoped that the readers of this bibliography will benefit from the relatively comprehensive listing of references. Further, it is hoped that readers will respond with missing references if possible. This will be the only way that the "holes" can be filled, since the work to bring the bibliography to the present level has been more than one person should attempt alone.

The author will continue to add to this bibliography as more work on  $G$  is done in the future and as other previous reports are uncovered and made available. Updates of new material that

supplements the existing version will be submitted as regular contributions to "Metrologia".

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1858	Isenkrahe, C. Poynting, J. H.	1888	1896
Airy, G. B. Clarke, A. R. James, H.		Von Hepperger, M. Wilsing, J.	Eötvös, R. Fitzgerald, G. F. Krigar-Menzel, O.
	1880		Richarz, F. Seeliger, H. Tisserand, F.
1859	Clarke, A. R. Faye, M. Knopf, O. Mendenhall, T. C. Zanotti-Bianco, O.	1889	
Clarke, A. R. Gosselin, P. F. J. James, H. Sargent, E.		Bessel, G. F. Boys, C. V. Laska, W. Poynting, J. H. Von Jolly, P. Wilsing, J. Yarkovsky, P. O.	Preston, E. D. Ray, L. E. Richarz, F. Sanford, F. Thwing, C. B.
1861	Keller, F. Mendenhall, T. C. Struve, O.	1890	1897
	Von Jolly, P.	Cornu, A. Eötvös, R. Joly, J. Lévy, M. Tisserand, F.	Austin, L. W. Braun, C. Krigar-Menzel, O. Mendenhall, T. C. Preston, E. D.
1863	1882		
Faye, M.	Von Jolly, P. Von Sterneck, R. Wallentin, J. G.		
1864		1891	1898
Babinet, J. Pechmann, E.	1883	Bock, M. Cornu, A. Eötvös, R. Poynting, J. H.	Braun, C. Gray, P. L. Krigar-Menzel, O. Poynting, J. H. Richarz, F. Woodward, R. S.
1865			
Pratt, J. H. Scheffler, H.	1884	1892	1899
	Baille, J.-B. Hicks, W. M. König, A. Lehmann-Filhés, R. Richarz, F. Von Sterneck, R.	Isenkrahe, C. Kriechgauer, H. Poynting, J. H. Turmlitz, O.	Burgess, G. K. Gerschun, A. Gray, P. L. Krigar-Menzel, O. Lippmann, G. Poynting, J. H. Richarz, F.
1869			
Schell, A.		1893	
1872		Berget, A. Krigar-Menzel, O. Landolt, H. Poynting, J. H. Preston, E. D. Richarz, F.	1900
Folie, F. Tisserand, F.	1885		
	König, A. Mayer, A. M. Richarz, F. Von Sterneck, R. Wilsing, J.		
1873		1894	
Baille, J. Cornu, A. Thompson, B.	1886	Boys, C. V. Fresdorf, G. Hall, A. Krigar-Menzel, O. Newcomb, S. Poynting, J. H. Richarz, F.	1901
1876	Bartoli, A. Keller, F. Von Sterneck, R. Wöllner, M.		Burgess, G. K. Krigar-Menzel, O. Poynting, J. H. Richarz, F.

1902	1911	1920	1927
Burgess, G. K.	Erismann, T.	Barus, C.	Brush, C. F.
Gore, J. H.		Compton, A. H.	Burgess, G. K.
Poynting, J. H.		Majorana, Q.	Heyl, P. R.
Thomson, J. J.	1912	Nipher, F. E.	Kunz, J.
	Bottlinger, K. F.	Pettersson, H.	Schlomka, T.
1903	De Sitter, W.	Poynting, J. H.	
Geigel, R.			1928
Kaufmann, W.	1913		Brush, C. F.
	Cremieu, V.		Kogbetliantz, E.
1904	Nipher, F. E.		Stern, T. E.
Haupt, E.	1914		
Laager, F.	Bottlinger, K. F.		1929
Take, E.	Burton, C. V.		Brush, C. F.
	Pettersson, H.		
1905		1922	1930
Cremieu, V.	1916	Barus, C.	Chazy, J.
Kleiner, A.	Barton, E. H.	Brush, C. F.	Heyl, P. R.
Phillips, P.	Davison, C.	Davy, N.	Kogbetliantz, E.
Poynting, J. H.	Kennard, E. H.	Eddington, A. S.	Kunz, J.
	Larmor, J.	Eichelberger, W. S.	Majorana, Q.
1906	Nipher, F. E.	Eötvös, R.	Zahradnicek, J.
	Sampson, R. A.	Fekete, E.	
Cremieu, V.	Shaw, P. E.	Heyl, P. R.	
Sagnac, G.		Majorana, Q.	
Southern, L.	1917	Morgan, H. R.	
	Barton, E. H.	Pekar, D.	1932
1907	Boys, C. V.	Potter, H. H.	Zahradnicek, J.
	Burton, C. V.	Richardson, O. W.	
Bauer, L. A.	Cremieu, V.	Shaw, P. E.	
Cremieu, V.	Hayes, C.	Simons, L.	1933
	Larmor, J.	Wilson, H. A.	Harrington, E. A.
1908	Lindemann, F. A.	Wulf, T.	Zahradnicek, J.
	Lodge, O.		
Bauer, L. A.	Loring, F. H.	1923	1935
Erismann, T.	Paterson, C. C.	Barus, C.	Fekete, E.
	Shaw, P. E.	Boys, C. V.	
	Smith, F. E.	Brush, C. F.	
	Thomas, J. S. G.	Davy, N.	
1909	Todd, G. W.	Popovici, C.	1937
Arrhenius, S.	1918	Potter, H. H.	Armellini, G.
Bauer, L. A.		Shaw, P. E.	
Cremieu, V.	Barus, C.	1924	1938
Lloyd, M. G.	Lodge, O.	Brush, C. F.	Armellini, G.
Poynting, J. H.	Tangi, K.	Heyl, P. R.	Bauer, L.
Thomson, J. J.	Zeeman, P.		Heyl, P. R.
Todd, G. W.		1925	1939
Von Seeliger, H.		Barus, C.	Chazy, J.
Walker, C.	1919	Heyl, P. R.	Jeffreys, H.
	Barus, C.		Popovici, C.
1910	Compton, A. H.	1926	1940
	Cremieu, V.	Brush, C. F.	
Cremieu, V.	Donnan, F. G.	Tangl, K.	Chazy, J.
Poynting, J. H.	Majorana, Q.		
Southern, L.	Rutherford, E.		

1942	1958	1962	Paya, D. Roos, P. Sagitov, M. U. Slichter, L. B. Weiss, R.
Chrzanowski, P. Heyl, P. R. Langevin, P.	Allais, M. F. C. Cocconi, G. Salpeter, E. E.	Banna, P. Caputo, M. Chernikov, Y. A. Chernyshov, V. I. Cook, A. H. Dicke, R. H. Finzi, A.	
1943	1958/1959		1966
Freed, S. Jaffey, A. H. Schultz, M. L.	Banna, P.	Grushevsky, N. P. Hoffmann, W. F. Kagal'nikova, N. N. Parker, R. L. Peebles, P. J. Radzievskii, V. V. Sagitov, M. U. Shuvalov, V. V. Thüring, B. Tiron, M.	Beams, J. W. Cameron, A. G. W. Demarque, P. R. Dicke, R. H. Ezer, D. Forward, R. L. Hughes, V. W. Kreuzer, L. B. Kulthau, A. R. Lewis, S. A. Lowry, R. A. Miller, L. R. Parker, H. M. Roeder, R. C. Senter, J. P. Weber, J. Williams, W. L.
1947	1959		
Estermann, I. Simpson, O. C. Stern, O.	Allais, M. F. C. Carrelli, A. Champion, F. C. Davy, N.		
1948	1960	1963	
Armellini, G. Teller, E.	Beltran-Lopez, V. Cocconi, G. Dicke, R. H. Epstein, S. T. Flint, F. V. Frauenfelder, H. Garwin, E. L. Hoffmann, W. F. Hughes, V. W. Kagal'nikova, I. I. Krotkov, R. Lier, R. H. Lüscher, E. Margulies, S. Peacock, R. N. Radzievskii, V. V. Robinson, H. G. Salpeter, E. E. Sherwin, C. W.	Braginsky, V. B. Butler, R. Egyed, L. Finzi, A. Groten, E. Harrison, J. C. Hoffmann, W. F. Kostko, O. Kronberg, P. MacDougall, J. Mark, R. Mason, J. Neiman, M. Rudenko, V. N. Rukman, G. I. Sandquist, A. Tomaschek, R.	
1950			1967
Bell Duane, B. H. Jacobson Kolm			Bahcall, J. N. Beams, J. W. Berman, D. Bond, A. Forward, R. L. Gamow, G. Koester, L. Kulthau, A. R. Long, D. R. Lowry, R. A. Miller, L. R. Parker, H. M. Senter, J. P. Shaviv, G. Southwell, W. H. Stephenson, L. M. Towler, W. R.
1951		1964	
McReynolds, A. W.			
1953			
Bleksley, A. E. H. Eötvös, R. Klutz, H. Pekar, D.			
1954	1961		
Heyl, P.R. Majorana, Q.	Beltran-Lopez, V. Caputo, M. Dicke, R. H. Dobrokhotov, Y. S. Drever, R. W. P. Eberhard, O. Forward, R. L. Grushevsky, N. P. Hoffmann, W. F. Hoffmann, B. Hughes, V. W. Krotkov, R. Lysenko, V. I. Pariisky, N. N. Radzievskii, V. V. Robinson, H. G. Sagitov, M. U. Schneiderov, A. J. Sigl, R. Thüring, B. Tomaschek, R.		1968
1955		1965	
Majorana, Q. Nijgh, G. J. Tomaschek, R. Wapstra, A. H.			Adamuti, I. A. Alpher, R. A. Berman, D. Birch, F. Braginsky, V. B. Cohen, S. A. Cook, A. H. Counselman, C. C. Finzi, A. Forward, R. L. Gamow, G. Ivankin, L. G. King, J. G. Kreuzer, L. B. Lowry, R. A. Martynov, V. K. Nash, J. H. Neeley, A. C.
1957			
Allais, M. F. C. Brein, R. Dicke, R. H. Majorana, Q. Meissner, H.			

Newton, R. R.	Grotens, E.	Gugel, L. G.	Caves, C. M.
Senter, J. P.	Ingalls, R. P.	Hill, H. A.	Chin, C.-W.
Shapiro, I. I.	Izmaylov, V. P.	Kalinnikov, I. I.	Fujii, Y.
Steger, P. J.	Karagioz, O. V.	Kolosnitsyn, N. I.	Gillies, G. T.
Tarakanov, Y. A.	Kaula, W. M.	Lowry, R. A.	Gittus, J. H.
Thompson, B.	Kezhutin, N. G.	Morrison, D.	Goodkind, J. M.
	Kocheryan, E. G.	Morrison, L. V.	Gugel, L. G.
1969	Kulthau, A. R.	Sagitov, M. U.	Izmaylov, V. P.
	Likhoshersnyh, D. U.	Tadzhidinov, K. G.	Karagioz, O. V.
	Lowry, R. A.	Taylor, B. N.	Kezhutin, N. G.
Adamuti, I. A.	MacDonald, G. J. F.	Towler, W. R.	Kropotkin, P. N.
Allen, M.	Milyukov, V. K.	Wesson, P. S.	Long, D. R.
Bahcall, J. N.	Morganstern, R. E.		Lowry, R. A.
Beams, J. W.	Mulholland, J. D.		Malin, S.
Blitch, M. G.	Parker, H. M.		Ritter, R. C.
Chapman, P. K.	Pettengill, G. H.		Smalley, L. L.
Cohen, S. A.	Plotkin, H. E.		Stothers, R.
Hovorka, J.	Pontikis, C.		Thorne, K. S.
Hulett, M. J.	Sagitov, M. U.		Towe, K. M.
King, J. G.	Saxl, E. J.		Van Flandern, T. C.
Kulthau, A. R.	Shapiro, I. I.		Voronkov, V. V.
Lowry, R. A.	Silverberg, E. C.		Warburton, R. J.
McVey, E. S.	Smith, W. B.		
Parker, H. M.	Tarakanov, Y. A.		
Rose, R. D.	Thyssen-Bornemisza, S.		
Sagitov, M. U.	Towler, W. R.		
Saxl, E. J.	Wilk, L. S.		
Shaviv, G.	Wilkinson, D. T.		
Towler, W. R.			
Wilk, L. S.			
1970			
Blood, B. E.	Beams, J. W.		1976
Douglass, Jr., D. H.	Bryzzhev, L. D.		Adams, P. J.
Facy, L.	Buchneva, L. G.		Adamuti, I. A.
Jeffreys, H.	Bullerwell, W.		Agafonov, N. I.
Kezhutin, N. G.	Cook, A. H.		Barrell, S. S.
Lee, W. N.	Faller, J. E.		Beams, J. W.
Long, D. R.	Fujii, Y.		Bleyer, U.
Pontikis, C.	Gaskell, T. H.		Campbell, D. B.
Renner, Y.	Gilvarry, J. J.		Canuto, V.
Sagitov, M. U.	Grotens, E.		Chesnokova, T. S.
Stewart, A. D.	Izmaylov, V. P.		Chin, C.-W.
Tarakanov, Y. A.	Karagioz, O. V.		Deslattes, R. D.
Vinti, J. P.	Kocheryan, E. G.		Dmitrieva, T. I.
1971	Koldewyn, W. A.		Faulkner, D. J.
Agafonov, N. I.	Kulthau, A. R.		Gillies, G. T.
Allen, M.	Lowry, R. A.		Gittus, J. H.
Alley, C. O.	Marussi, A.		Goodkind, J. M.
Ash, M. B.	Morganstern, R. E.		Gugel, L. G.
Beams, J. W.	Muller, P. M.		Hooge, F. N.
Bender, P. L.	O'Hanlon, J.		Hsieh, S.-H.
Blood, B. E.	Parker, H. M.		Izmaylov, V. P.
Bocchio, F.	Pontikis, C.		John, R. W.
Boulanger, Y. D.	Sagitov, M. U.		Kalinnikov, I. I.
Buchneva, L. G.	Sugimoto, D.		Karagioz, O. V.
Chesnokova, T. S.	Thyssen-Bornemisza, S.		Kocheryan, E. G.
Colavita, P. A.	Towler, W. R.		Koester, L.
Cook, A. H.	Vinti, J. P.		Koldewyn, W. A.
Currie, D. G.	Wyckoff, R. D.		Lewis, B. M.
Dicke, R. H.			Liebscher, D.-E.
Facy, L.			Long, D. R.
Faller, J. E.			Lowry, R. A.
Fujii, Y.			Luther, G. G.
			Malin, S.
			Mansfield, V. N.
			Milyukov, V. K.
			Monakhov, E. A.
			Nazarenko, V. S.
			Paik, H. J.
			Pettengill, G. H.
			Poulis, J. A.
			Reasonberg, R. D.
			Ritter, R. C.
			Roxburgh, I. W.
			Sagitov, M. U.
1973			
	Alpher, R. A.		
	Anikin, V. L.		
	Beams, J. W.		
	Cohen, R. E.		
	Dirac, P. A. M.		
1975			
	Agafonov, N. I.		
	Beams, J. W.		
	Beaumont, C.		
	Cohen, R. E.		
	Braginsky, V. B.		
	Bullen, K. E.		

Shapiro, I. I.	1978	Carignan, C.	Gillies, G. T.
Steigman, G.		Carr, B. J.	Goldman, T.
Stephenson, L. M.		Caves, C. M.	Haars, H.
Stothers, R.		Chesnokova, T. S.	Hirakawa, H.
Tadzhidinov, K. G.		De Boer, H.	Hoskins, J. K.
Taranakov, Y. A.		Dmitrieva, T. I.	Hsieh, S.-H.
Towe, K. M.		Easthope, P.	Jones, Jr., G. R.
Towler, W. R.		Eötvös, R.	Levitt, L. S.
Tsiang, E.		Esposito, P. B.	Leyh, C. H.
Van Dijk, H. H. J.		Estabrook, F. B.	Long, D. R.
Van Flandern, T. C.		Farinella, P.	Michaelis, W.
Warburton, R. J.		Feinberg, G.	Milani, A.
Will, C. M.		Fialovszky, L.	Mulholland, J. D.
1977		Frontov, V. N.	Narlikar, J. V.
Anderson, J. D.		Fujimoto, M.-K.	Neville, D. E.
Bagmet, A. L.		Hirakawa, H.	Newman, R.
Blake, G. M.		Hoskins, J. K.	Ni, W.-T.
Bleyer, U.		Hsieh, S.-H.	Nieto, M. M.
Braginsky, V. B.		Hu, C.-C.	Nobili, A. M.
Canuto, V. M.		Lambeck, K.	Oide, K.
Caputo, M.		Liu, F.-H.	Paik, H. J.
Caves, C. M.		Liu, W.-N.	Panov, V. I.
Davies, P. C. W.		Martin, C. F.	Pellam, J.
De Boer, H.		Metherell, A. J. F.	Rana, N. C.
Desabbata, V.		Milani, A.	Ritter, R. C.
Eichendorf, W.		Mills, Jr., A. P.	Schlimgme, E.
Estabrook, F. B.		Milyukov, V. K.	Schultz, J.
Everitt, C. W. F.		Monakhov, E. A.	Shapiro, I. I.
Fitch, J. P.		Nazarenko, V. S.	Sinclair, W. S.
Frontov, V. N.		Newman, R.	Sjogren, W. L.
Hirakawa, H.		Ni, W.-T.	Spero, R. E.
Hiramatsu, S.		Nobili, A. M.	Suzuki, T.
Hooge, F. N.		Oh, I. H.	Tsubono, K.
Izmaylov, V. P.		Owen, J. R.	Wesson, P. S.
John, R. W.		Paik, H. J.	Williams, J. G.
Karagioz, O. V.		Panov, V. I.	Woodward, J. F.
Kocheryan, E. G.		Rees, M. J.	1981
Koldewyn, W. A.		Rood, R. T.	Alpher, R. A.
Liebscher, D.-E.		Rudenko, V. N.	Augustin, R.
Lodenquai, J.		Sagitov, M. U.	Barrow, J. D.
Lyttleton, R. A.		Schramm, D. N.	Canuto, V. M.
Mansfield, V.		Shey, Y.-C.	Carey, S. W.
Marchant, A.		Sirois, A.	Chen, Y. T.
Mikkelsen, D. R.		Spero, R. E.	Chow, T. L.
Mucket, J. P.		Stegna, L.	Cowsik, R.
Newman, M. J.		Steigman, G.	Dannehold, T. L.
Newman, R.		Sucher, J.	Davies, R. D.
Ogawa, Y.		Tadzhidinov, K. G.	De Boer, H.
Panov, V. I.		Treder, H.-J.	Fialovszky, L.
Pellam, J.		Williams, E.	Fujii, Y.
Philberth, K.		Woodward, J. F.	Geilker, C. D.
Poulis, J. A.		Yang, C.-H.	Gibbons, G. W.
Reasonberg, R. D.		Yang, J.	Gillies, G. T.
Reinhardt, M.		Yourgrau, W.	Goodson, R. E.
Rizzati, P.		Yu, H.-T.	Haars, H.
Roxburgh, I. W.	1979	1980	Hirakawa, H.
Schultz, J.		Canuto, V. M.	Holding, S. C.
Shapiro, I. I.		Chan, H. A.	Hut, P.
Spero, R. E.		Cheung, W. S.	Izmaylov, V. P.
Thorne, K. S.		Cook, A. H.	Jones, G. R. Jr.
Treder, H.-J.		Dannehold, T.	Karagioz, O. V.
Tsubono, K.		De Boer, H.	Kimura, S.
Wahlquist, H. D.		Dirac, P. A. M.	Kononenko, I. V.
Yamaguchi, Y.		Farinella, P.	Kuroda, K.
		Ferrari, A. J.	Lapiedra, R.
		Frontov, V. N.	Long, D. R.

Maher, A. R.	Ritter, R. C.	Wahr, J. M.	Itano, W. M.
Malamud, H.	Rothman, T.	Wang, Q.	Keesey, M. S.
Mapoles, E. R.	Thiessen, P. A.	Yilmaz, H.	Kolosnitsyn, N. I.
Maskelyne, N.	Towler, W. R.	Zhang, P.	Kuroda, K.
McQueen, H. W. S.	Treder, H.-J.		Lau, E. K.
Melissons, A. C.	Tsubono, K.		Lauterbach, R.
Metherell, A. J. F.	Whitaker, D. H.	1984	Mel'nikov, V. N.
Michaelis, W.	Worden, Jr., P. W.	Anandan, J.	Milyukov, V. K.
Milyukov, V. K.	Zhang, P.	Bender, P. L.	Mulholland, J. D.
Monakhov, E. A.		Bleyer, U.	Newhall, X. X.
Morris, D.		Chan, H. A.	Newman, R.
Nazarenko, V. S.		Chen, Y. T.	Prestage, J. D.
Page, D. N.		Cheung, W. S.	Radynov, A. G.
Palacios, J. A.		Cook, A. H.	Reiner, P. J.
Ponomorov, I. S.		De Boer, H.	Ritter, R. C.
Ritter, R. C.		De Bra, D.	Schultz, J.
Runcorn, S. K.		Deslattes, R. D.	Schwede, H. F.
Sagitov, M. U.		Fairbank, W. M.	Spero, R.
Shafranovskaya, M. M.		Faller, J. E.	Standish, E. M.
Shipman, H. L.		Gillies, G. T.	Stedman, G. E.
Silin, A. A.		Hirakawa, H.	Williams, J. G.
Speake, C. C.		Holding, S. C.	Wineland, D. J.
Stacey, F. D.		Hoskins, J. K.	Winkler, L. I.
Suzuki, T.		Hu, C.-C.	
Tadzhidinov, K. G.		John, R. W.	1986
Tsubono, K.		Kuroda, K.	Aciu, A.
Tuck, G. J.		Leyh, C. H.	Alexandrescu, M.
Van Flandern, T. C.		Liebscher, D.-E.	Ashby, N.
Veselov, K. E.		Liu, F.-H.	Avron, Y.
Wagner, A. R.		Liu, W.-N.	Bartlett, D. F.
Wesson, P. S.		Long, D. R.	Bender, P. L.
Whiting, B. F.		Loper, D. E.	Cristea, O.
Will, C. M.		Luther, G. G.	De Boer, H.
		Maddox, J.	Farinella, P.
1982		Metherell, A. J. F.	Fowler, W. B.
Adamuti, I. A.		Mio, N.	Gillies, G. T.
Aspden, H.		Newman, R.	Goodwin, B. D.
Augustin, R.		Ni, W.-T.	Haars, H.
Ballentine, L. E.		Oelfke, W. C.	Hills, J. G.
Bernard, B. E.		Paik, H. J.	Hinze, W. J.
Canuto, V. M.		Pinotsis, A. D.	Holding, S. C.
Chan, H. A.		Reiner, P. J.	Karim, M.
Chen, Y. T.		Ritter, R. C.	Kim, Y. E.
Cheung, W. S.		Sagitov, M. U.	Klepacki, D. J.
Cook, A. H.		Sanders, R. H.	Kolosnitsyn, N. I.
Dahlen, F. A.		Schultz, J.	Liakhovets, V. D.
Dannehold, T.		Speake, C. C.	Livio, M.
De Boer, H.		Spero, R.	Lupu, D.
Everitt, C. W. F.		Stacey, F. D.	Maddox, J.
Gillies, G. T.		Steenbeck, M.	Melissinos, A. C.
Goldman, I.		Towler, W. R.	Metherell, A. J. F.
Haars, H.		Treder, H.-J.	Michaelis, W.
Hawkins, B.		Tsubono, K.	Milani, A.
Hirakawa, H.		Tuck, G. J.	Minti, H.
Jones, G. R. Jr.		Van Flandern, T. C.	Moore, G. I.
Kuroda, K.		Yang, C.-H.	Nobili, A. M.
Liu, H.		Yu, H.-T.	Piso, M.
Luther, G. G.			Popovici, L.
Matzner, R.		1985	Quinn, T. J.
Michaelis, W.			Reece, C. E.
Moody, M. V.		Anderson, J. D.	Reiner, P. J.
Neville, D. E.		Bernard, B. E.	Ritenour, R. L.
Ogawa, Y.		Bollinger, J. J.	Ritter, R. C.
Page, D. N.		Gillies, G. T.	Riveros, C.
Paik, H. J.		Helling, R. W.	Rogers, J.
Qin, R.		Hirakawa, H.	Saru, D.
		Hoskins, J. K.	Semertzidis, J.

Simaciu, I.	Vucetich, H.	Zhou, R.	Gillies, G. T.
Stacey, F. D.	Wahr, J. M.		Marussi, A.
Stancu, D.	Wang, Z.		Nieto, M. M.
Toohey, W. J.	Winkler, L. I.	In Press	Ritter, R. C.
Tuck, G. J.	Wu, Y.-S.		Speake, C. C.
Van Buren, D.	Wuensch, W.	Bernard, B. E.	Winkler, L. I.
Vincent, M. A.	Zaitsev, N. A.	Clarke, R. G. S.	