white dwarf

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1 Master's Degree in Particle Physics and Cosmos

1.1 Cosmos Problem 9.3 White Dwarf

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For a white dwarf, the non-relativistic equation of state is given as:

$$P = \alpha \rho^{5/3}$$

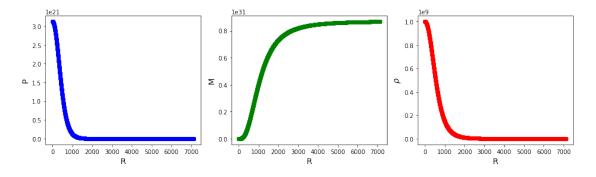
And if the white dwarf is in equilibrium,

$$\frac{dP}{dr} = -\rho(r)\frac{GM(r)}{r^2}$$

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In [2]: import numpy as np
        from matplotlib import pyplot as plt
        %matplotlib inline
In [55]: def dP(a, b):
             alpha = 3123161.171
             G = 6.67E-11
             return ((4 / 3) * np.pi * G * a * (r0 ** 2) * (b ** 2))
         def white_dwarf(r0, rho_0 ,dr=1):
             Builds a white dwarf
             alpha = 3123161.171
             G = 6.67E-11
             m0 = (4 / 3) * np.pi * rho_0 * (r0 ** 3)
             P0 = alpha * (rho_0 ** (5/3))
             # lists
             R = [dr]
             M = [mO]
             P = [P0]
             rho = [rho_0]
             while rho[dr - 1] > 1E4:
                 dr = dr + 1
```

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new_P = P[dr - 2] - dP(R[dr - 2], rho[dr - 2])
    new\_rho = (new\_P / alpha) ** (3/5)
    new_M = M[dr - 2] + (4 * np.pi * new_rho * ((r0 * dr)**3 - (r0 * (dr - 1))**3
    R.append(dr)
    P.append(new_P)
    rho.append(new_rho)
    M.append(new_M)
#plot
plt.figure(figsize=(16,4))
plt.subplot(1, 3, 1)
plt.plot(R,P,'ob')
plt.xlabel("R", fontsize=14)
plt.ylabel("P", fontsize=14)
plt.subplot(1, 3, 2)
plt.plot(R,M,'og')
{\tt plt.xlabel("R", fontsize=14)}
plt.ylabel("M", fontsize=14)
plt.subplot(1, 3, 3)
plt.plot(R,rho,'or')
plt.xlabel("R", fontsize=14)
plt.ylabel(r"$\rho$", fontsize=14)
plt.show()
```

In [56]: white_dwarf(1E4, 1E9)



It is obtained $0.10R_{solar}$ and $1.45M_{solar}$