

Infinite matter, from the electron gas to nuclear matter

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Contents

Introduction to studies of infinite matter

Studies of infinite nuclear matter play an important role in nuclear physics. The aim of this part of the lectures is to provide the necessary ingredients for performing studies of neutron star matter (or matter in β -equilibrium) and symmetric nuclear matter. We start however with the electron gas in two and three dimensions for both historical and pedagogical reasons. Since there are several benchmark calculations for the electron gas, this small detour will allow us to establish the necessary formalism. Thereafter we will study infinite nuclear matter

- at the Hartree-Fock with realistic nuclear forces and
- using many-body methods like coupled-cluster theory or in-medium SRG as discussed in our previous sections.

The infinite electron gas

The electron gas is perhaps the only realistic model of a system of many interacting particles that allows for a solution of the Hartree-Fock equations on a closed form. Furthermore, to first order in the interaction, one can also compute on a closed form the total energy and several other properties of a many-particle systems. The model gives a very good approximation to the properties of valence electrons in metals. The assumptions are

- System of electrons that is not influenced by external forces except by an attraction provided by a uniform background of ions. These ions give rise to a uniform background charge. The ions are stationary.
- The system as a whole is neutral.
- We assume we have N_e electrons in a cubic box of length L and volume $\Omega = L^3$. This volume contains also a uniform distribution of positive charge with density $N_e e / \Omega$.

The homogeneous electron gas is one of the few examples of a system of many interacting particles that allows for a solution of the mean-field Hartree-Fock equations on a closed form. To first order in the electron-electron interaction, this applies to ground state properties like the energy and its pertinent equation of state as well. The homogeneous electron gas is a system of electrons that is not influenced by external forces except by an attraction provided by a uniform background of ions. These ions give rise to a uniform background charge. The ions are stationary and the system as a whole is neutral. Irrespective of this simplicity, this system, in both two and three-dimensions, has eluded a proper description of correlations in terms of various first principle methods, except perhaps for quantum Monte Carlo methods. In particular, the diffusion Monte Carlo calculations of [Ceperley](#) and [Ceperley and Tanatar](#) are presently still considered as the best possible benchmarks for the two- and three-dimensional electron gas.

The electron gas, in two or three dimensions is thus interesting as a test-bed for electron-electron correlations. The three-dimensional electron gas is particularly important as a cornerstone of the local-density approximation in density-functional theory. In the physical world, systems similar to the three-dimensional electron gas can be found in, for example, alkali metals and doped semiconductors. Two-dimensional electron fluids are observed on metal and liquid-helium surfaces, as well as at metal-oxide-semiconductor interfaces. However, the Coulomb interaction has an infinite range, and therefore long-range correlations play an essential role in the electron gas.

At low densities, the electrons become localized and form a lattice. This so-called Wigner crystallization is a direct consequence of the long-ranged repulsive interaction. At higher densities, the electron gas is better described as a liquid. When using, for example, Monte Carlo methods the electron gas must be approximated by a finite system. The long-range Coulomb interaction in the electron gas causes additional finite-size effects that are not present in other infinite systems like nuclear matter or neutron star matter. This poses additional challenges to many-body methods when applied to the electron gas.

The infinite electron gas as a homogenous system

This is a homogeneous system and the one-particle wave functions are given by plane wave functions normalized to a volume Ω for a box with length L (the

limit $L \rightarrow \infty$ is to be taken after we have computed various expectation values)

$$\psi_{\mathbf{k}\sigma}(\mathbf{r}) = \frac{1}{\sqrt{\Omega}} \exp(i\mathbf{k}\mathbf{r})\xi_{\sigma}$$

where \mathbf{k} is the wave number and ξ_{σ} is a spin function for either spin up or down

$$\xi_{\sigma=+1/2} = \begin{pmatrix} 1 \\ 0 \end{pmatrix} \quad \xi_{\sigma=-1/2} = \begin{pmatrix} 0 \\ 1 \end{pmatrix}.$$

Periodic boundary conditions

We assume that we have periodic boundary conditions which limit the allowed wave numbers to

$$k_i = \frac{2\pi n_i}{L} \quad i = x, y, z \quad n_i = 0, \pm 1, \pm 2, \dots$$

We assume first that the electrons interact via a central, symmetric and translationally invariant interaction $V(r_{12})$ with $r_{12} = |\mathbf{r}_1 - \mathbf{r}_2|$. The interaction is spin independent.

The total Hamiltonian consists then of kinetic and potential energy

$$\hat{H} = \hat{T} + \hat{V}.$$

The operator for the kinetic energy can be written as

$$\hat{T} = \sum_{\mathbf{k}\sigma} \frac{\hbar^2 k^2}{2m} a_{\mathbf{k}\sigma}^{\dagger} a_{\mathbf{k}\sigma}.$$

Defining the Hamiltonian operator

The Hamiltonian operator is given by

$$\hat{H} = \hat{H}_{el} + \hat{H}_b + \hat{H}_{el-b},$$

with the electronic part

$$\hat{H}_{el} = \sum_{i=1}^N \frac{p_i^2}{2m} + \frac{e^2}{2} \sum_{i \neq j} \frac{e^{-\mu|\mathbf{r}_i - \mathbf{r}_j|}}{|\mathbf{r}_i - \mathbf{r}_j|},$$

where we have introduced an explicit convergence factor (the limit $\mu \rightarrow 0$ is performed after having calculated the various integrals). Correspondingly, we have

$$\hat{H}_b = \frac{e^2}{2} \int \int d\mathbf{r} d\mathbf{r}' \frac{n(\mathbf{r})n(\mathbf{r}')e^{-\mu|\mathbf{r} - \mathbf{r}'|}}{|\mathbf{r} - \mathbf{r}'|},$$

which is the energy contribution from the positive background charge with density $n(\mathbf{r}) = N/\Omega$. Finally,

$$\hat{H}_{el-b} = -\frac{e^2}{2} \sum_{i=1}^N \int d\mathbf{r} \frac{n(\mathbf{r})e^{-\mu|\mathbf{r} - \mathbf{x}_i|}}{|\mathbf{r} - \mathbf{x}_i|},$$

is the interaction between the electrons and the positive background.

Single-particle Hartree-Fock energy

In the first exercise below we show that the Hartree-Fock energy can be written as

$$\varepsilon_k^{HF} = \frac{\hbar^2 k^2}{2m_e} - \frac{e^2}{\Omega^2} \sum_{k' \leq k_F} \int d\mathbf{r} e^{i(\mathbf{k}' - \mathbf{k})\mathbf{r}} \int d\mathbf{r}' \frac{e^{i(\mathbf{k} - \mathbf{k}')\mathbf{r}'}}{|\mathbf{r} - \mathbf{r}'|}$$

resulting in

$$\varepsilon_k^{HF} = \frac{\hbar^2 k^2}{2m_e} - \frac{e^2 k_F}{2\pi} \left[2 + \frac{k_F^2 - k^2}{kk_F} \ln \left| \frac{k + k_F}{k - k_F} \right| \right]$$

The previous result can be rewritten in terms of the density

$$n = \frac{k_F^3}{3\pi^2} = \frac{3}{4\pi r_s^3},$$

where $n = N_e/\Omega$, N_e being the number of electrons, and r_s is the radius of a sphere which represents the volume per conducting electron. It can be convenient to use the Bohr radius $a_0 = \hbar^2/e^2 m_e$. For most metals we have a relation $r_s/a_0 \sim 2 - 6$. The quantity r_s is dimensionless.

In the second exercise below we find that the total energy $E_0/N_e = \langle \Phi_0 | \hat{H} | \Phi_0 \rangle / N_e$ for this system to first order in the interaction is given as

$$E_0/N_e = \frac{e^2}{2a_0} \left[\frac{2.21}{r_s^2} - \frac{0.916}{r_s} \right].$$

Exercise 1: Hartree-Fock single-particle solution for the electron gas

The electron gas model allows closed form solutions for quantities like the single-particle Hartree-Fock energy. The latter quantity is given by the following expression

$$\varepsilon_k^{HF} = \frac{\hbar^2 k^2}{2m} - \frac{e^2}{V^2} \sum_{k' \leq k_F} \int d\mathbf{r} e^{i(\mathbf{k}' - \mathbf{k})\mathbf{r}} \int d\mathbf{r}' \frac{e^{i(\mathbf{k} - \mathbf{k}')\mathbf{r}'}}{|\mathbf{r} - \mathbf{r}'|}$$

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a) Show first that

$$\varepsilon_k^{HF} = \frac{\hbar^2 k^2}{2m} - \frac{e^2 k_F}{2\pi} \left[2 + \frac{k_F^2 - k^2}{kk_F} \ln \left| \frac{k + k_F}{k - k_F} \right| \right]$$

Hint. Hint: Introduce the convergence factor $e^{-\mu|\mathbf{r} - \mathbf{r}'|}$ in the potential and use $\sum_{\mathbf{k}} \rightarrow \frac{V}{(2\pi)^3} \int d\mathbf{k}$

Solution. We want to show that, given the Hartree-Fock equation for the electron gas

$$\varepsilon_k^{HF} = \frac{\hbar^2 k^2}{2m} - \frac{e^2}{V^2} \sum_{p \leq k_F} \int d\mathbf{r} \exp(i(\mathbf{p} - \mathbf{k})\mathbf{r}) \int d\mathbf{r}' \frac{\exp(i(\mathbf{k} - \mathbf{p})\mathbf{r}')}{|\mathbf{r} - \mathbf{r}'|}$$

the single-particle energy can be written as

$$\varepsilon_k^{HF} = \frac{\hbar^2 k^2}{2m} - \frac{e^2 k_F}{2\pi} \left[2 + \frac{k_F^2 - k^2}{kk_F} \ln \left| \frac{k + k_F}{k - k_F} \right| \right].$$

We introduce the convergence factor $e^{-\mu|\mathbf{r}-\mathbf{r}'|}$ in the potential and use $\sum_{\mathbf{k}} \rightarrow \frac{V}{(2\pi)^3} \int d\mathbf{k}$. We can then rewrite the integral as

$$\frac{e^2}{V^2} \sum_{k' \leq k_F} \int d\mathbf{r} \exp(i(\mathbf{k}' - \mathbf{k})\mathbf{r}) \int d\mathbf{r}' \frac{\exp(i(\mathbf{k} - \mathbf{p})\mathbf{r}')}{|\mathbf{r} - \mathbf{r}'|} = \quad (1)$$

$$\frac{e^2}{V(2\pi)^3} \int d\mathbf{r} \int \frac{d\mathbf{r}'}{|\mathbf{r} - \mathbf{r}'|} \exp(-i\mathbf{k}(\mathbf{r} - \mathbf{r}')) \int d\mathbf{p} \exp(i\mathbf{p}(\mathbf{r} - \mathbf{r}')), \quad (2)$$

and introducing the abovementioned convergence factor we have

$$\lim_{\mu \rightarrow 0} \frac{e^2}{V(2\pi)^3} \int d\mathbf{r} \int d\mathbf{r}' \frac{\exp(-\mu|\mathbf{r} - \mathbf{r}'|)}{|\mathbf{r} - \mathbf{r}'|} \int d\mathbf{p} \exp(i(\mathbf{p} - \mathbf{k})(\mathbf{r} - \mathbf{r}')). \quad (3)$$

With a change variables to $\mathbf{x} = \mathbf{r} - \mathbf{r}'$ and $\mathbf{y} = \mathbf{r}'$ we rewrite the last integral as

$$\lim_{\mu \rightarrow 0} \frac{e^2}{V(2\pi)^3} \int d\mathbf{p} \int d\mathbf{y} \int d\mathbf{x} \exp(i(\mathbf{p} - \mathbf{k})\mathbf{x}) \frac{\exp(-\mu|\mathbf{x}|)}{|\mathbf{x}|}.$$

The integration over \mathbf{x} can be performed using spherical coordinates, resulting in (with $x = |\mathbf{x}|$)

$$\int d\mathbf{x} \exp(i(\mathbf{p} - \mathbf{k})\mathbf{x}) \frac{\exp(-\mu|\mathbf{x}|)}{|\mathbf{x}|} = \int x^2 dx d\phi d\cos(\theta) \exp(i(\mathbf{p} - \mathbf{k})x \cos(\theta)) \frac{\exp(-\mu x)}{x}.$$

We obtain

$$4\pi \int dx \frac{\sin(|\mathbf{p} - \mathbf{k}|x)}{|\mathbf{p} - \mathbf{k}|} \exp(-\mu x) = \frac{4\pi}{\mu^2 + |\mathbf{p} - \mathbf{k}|^2}. \quad (4)$$

This results gives us

$$\lim_{\mu \rightarrow 0} \frac{e^2}{V(2\pi)^3} \int d\mathbf{p} \int d\mathbf{y} \frac{4\pi}{\mu^2 + |\mathbf{p} - \mathbf{k}|^2} = \lim_{\mu \rightarrow 0} \frac{e^2}{2\pi^2} \int d\mathbf{p} \frac{1}{\mu^2 + |\mathbf{p} - \mathbf{k}|^2}, \quad (5)$$

where we have used that the integrand on the left-hand side does not depend on \mathbf{y} and that $\int d\mathbf{y} = V$.

Introducing spherical coordinates we can rewrite the integral as

$$\lim_{\mu \rightarrow 0} \frac{e^2}{2\pi^2} \int d\mathbf{p} \frac{1}{\mu^2 + |\mathbf{p} - \mathbf{k}|^2} = \frac{e^2}{2\pi^2} \int d\mathbf{p} \frac{1}{|\mathbf{p} - \mathbf{k}|^2} = \quad (6)$$

$$\frac{e^2}{\pi} \int_0^{k_F} p^2 dp \int_0^\pi d\theta \cos(\theta) \frac{1}{p^2 + k^2 - 2pk \cos(\theta)}, \quad (7)$$

and with the change of variables $\cos(\theta) = u$ we have

$$\frac{e^2}{\pi} \int_0^{k_F} p^2 dp \int_0^\pi d\theta \cos(\theta) \frac{1}{p^2 + k^2 - 2pk \cos(\theta)} = \frac{e^2}{\pi} \int_0^{k_F} p^2 dp \int_{-1}^1 du \frac{1}{p^2 + k^2 - 2pku},$$

which gives

$$\frac{e^2}{k\pi} \int_0^{k_F} p dp \{ \ln(|p+k|) - \ln(|p-k|) \}.$$

Introducing new variables $x = p + k$ and $y = p - k$, we obtain after some straightforward reordering of the integral

$$\frac{e^2}{k\pi} \left[k k_F + \frac{k_F^2 - k^2}{k k_F} \ln \left| \frac{k + k_F}{k - k_F} \right| \right],$$

which gives the abovementioned expression for the single-particle energy.

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b) Rewrite the above result as a function of the density

$$n = \frac{k_F^3}{3\pi^2} = \frac{3}{4\pi r_s^3},$$

where $n = N/V$, N being the number of particles, and r_s is the radius of a sphere which represents the volum per conducting electron.

Solution. Introducing the dimensionless quantity $x = k/k_F$ and the function

$$F(x) = \frac{1}{2} + \frac{1-x^2}{4x} \ln \left| \frac{1+x}{1-x} \right|,$$

we can rewrite the single-particle Hartree-Fock energy as

$$\varepsilon_k^{HF} = \frac{\hbar^2 k^2}{2m} - \frac{2e^2 k_F}{\pi} F(k/k_F),$$

and dividing by the non-interacting contribution at the Fermi level,

$$\varepsilon_0^F = \frac{\hbar^2 k_F^2}{2m},$$

we have

$$\frac{\varepsilon_k^{HF}}{\varepsilon_0^F} = x^2 - \frac{e^2 m}{\hbar^2 k_F \pi} F(x) = x^2 - \frac{4}{\pi k_F a_0} F(x),$$

where $a_0 = 0.0529$ nm is the Bohr radius, setting thereby a natural length scale.

By introducing the radius r_s of a sphere whose volume is the volume occupied by each electron, we can rewrite the previous equation in terms of r_s using that the electron density $n = N/V$

$$n = \frac{k_F^3}{3\pi^2} = \frac{3}{4\pi r_s^3},$$

we have (with $k_F = 1.92/r_s$,

$$\frac{\varepsilon_k^{HF}}{\varepsilon_0^F} = x^2 - \frac{e^2 m}{\hbar^2 k_F \pi} F(x) = x^2 - \frac{r_s}{a_0} 0.663 F(x),$$

with $r_s \sim 2 - 6$ for most metals.

It can be convenient to use the Bohr radius $a_0 = \hbar^2/e^2 m$. For most metals we have a relation $r_s/a_0 \sim 2 - 6$.

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c) Make a plot of the free electron energy and the Hartree-Fock energy and discuss the behavior around the Fermi surface. Extract also the Hartree-Fock band width $\Delta\varepsilon^{HF}$ defined as

$$\Delta\varepsilon^{HF} = \varepsilon_{k_F}^{HF} - \varepsilon_0^{HF}.$$

Compare this results with the corresponding one for a free electron and comment your results. How large is the contribution due to the exchange term in the Hartree-Fock equation?

Solution. We can now define the so-called band gap, that is the scatter between the maximal and the minimal value of the electrons in the conductance band of a metal (up to the Fermi level). For $x = 1$ and $r_s/a_0 = 4$ we have

$$\frac{\varepsilon_{k=k_F}^{HF}}{\varepsilon_0^F} = -0.326,$$

and for $x = 0$ we have

$$\frac{\varepsilon_{k=0}^{HF}}{\varepsilon_0^F} = -2.652,$$

which results in a gap at the Fermi level of

$$\Delta\varepsilon^{HF} = \frac{\varepsilon_{k=k_F}^{HF}}{\varepsilon_0^F} - \frac{\varepsilon_{k=0}^{HF}}{\varepsilon_0^F} = 2.326.$$

This quantity measures the deviation from the $k = 0$ single-particle energy and the energy at the Fermi level. The general result is

$$\Delta\varepsilon^{HF} = 1 + \frac{r_s}{a_0} 0.663.$$

The following python code produces a plot of the electron energy for a free electron (only kinetic energy) and for the Hartree-Fock solution. We have chosen here a ratio $r_s/a_0 = 4$ and the equations are plotted as funtions of k/f_F .

```

import numpy as np
from math import log
from matplotlib import pyplot as plt
from matplotlib import rc, rcParams
import matplotlib.units as units
import matplotlib.ticker as ticker
rc('text',usetex=True)
rc('font',**{'family':'serif','serif':['Hartree-Fock energy']})
font = {'family' : 'serif',
        'color'   : 'darkred',
        'weight'  : 'normal',
        'size'    : 16,
        }

N = 100
x = np.linspace(0.0, 2.0,N)
F = 0.5+np.log(abs((1.0+x)/(1.0-x)))*(1.0-x*x)*0.25/x
y = x*x -4.0*0.663*F

plt.plot(x, y, 'b-')
plt.plot(x, x*x, 'r-')
plt.title(r'\bf Hartree-Fock single-particle energy for electron gas}', fontsize=20)
plt.text(3, -40, r'Parameters: $r_s/a_0=4$', fontdict=font)
plt.xlabel(r'$k/k_F$',fontsize=20)
plt.ylabel(r'$\varepsilon_k^{HF}/\varepsilon_0^{HF}$',fontsize=20)
# Tweak spacing to prevent clipping of ylabel
plt.subplots_adjust(left=0.15)
plt.savefig('hartreefockspelgas.pdf', format='pdf')
plt.show()

```

From the plot we notice that the exchange term increases considerably the band gap compared with the non-interacting gas of electrons.

We will now define a quantity called the effective mass. For $|\mathbf{k}|$ near k_F , we can Taylor expand the Hartree-Fock energy as

$$\varepsilon_k^{HF} = \varepsilon_{k_F}^{HF} + \left(\frac{\partial \varepsilon_k^{HF}}{\partial k} \right)_{k_F} (k - k_F) + \dots$$

If we compare the latter with the corresponding expression for the non-interacting system

$$\varepsilon_k^{(0)} = \frac{\hbar^2 k_F^2}{2m} + \frac{\hbar^2 k_F}{m} (k - k_F) + \dots,$$

we can define the so-called effective Hartree-Fock mass as

$$m_{HF}^* \equiv \hbar^2 k_F \left(\frac{\partial \varepsilon_k^{HF}}{\partial k} \right)_{k_F}^{-1}$$

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d) Compute m_{HF}^* and comment your results.

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e) Show that the level density (the number of single-electron states per unit energy) can be written as

$$n(\varepsilon) = \frac{V k^2}{2\pi^2} \left(\frac{\partial \varepsilon}{\partial k} \right)^{-1}$$

Calculate $n(\varepsilon_F^{HF})$ and comment the results.

Exercise 2: Hartree-Fock ground state energy for the electron gas in three dimensions

We consider a system of electrons in infinite matter, the so-called electron gas. This is a homogeneous system and the one-particle states are given by plane wave function normalized to a volume Ω for a box with length L (the limit $L \rightarrow \infty$ is to be taken after we have computed various expectation values)

$$\psi_{\mathbf{k}\sigma}(\mathbf{r}) = \frac{1}{\sqrt{\Omega}} \exp(i\mathbf{k}\mathbf{r})\xi_{\sigma}$$

where \mathbf{k} is the wave number and ξ_{σ} is a spin function for either spin up or down

$$\xi_{\sigma=+1/2} = \begin{pmatrix} 1 \\ 0 \end{pmatrix} \quad \xi_{\sigma=-1/2} = \begin{pmatrix} 0 \\ 1 \end{pmatrix}.$$

We assume that we have periodic boundary conditions which limit the allowed wave numbers to

$$k_i = \frac{2\pi n_i}{L} \quad i = x, y, z \quad n_i = 0, \pm 1, \pm 2, \dots$$

We assume first that the particles interact via a central, symmetric and translationally invariant interaction $V(r_{12})$ with $r_{12} = |\mathbf{r}_1 - \mathbf{r}_2|$. The interaction is spin independent.

The total Hamiltonian consists then of kinetic and potential energy

$$\hat{H} = \hat{T} + \hat{V}.$$

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a) Show that the operator for the kinetic energy can be written as

$$\hat{T} = \sum_{\mathbf{k}\sigma} \frac{\hbar^2 k^2}{2m} a_{\mathbf{k}\sigma}^{\dagger} a_{\mathbf{k}\sigma}.$$

Find also the number operator \hat{N} and find a corresponding expression for the interaction \hat{V} expressed with creation and annihilation operators. The expression for the interaction has to be written in k space, even though V depends only

on the relative distance. It means that you need to set up the Fourier transform $\langle \mathbf{k}_i \mathbf{k}_j | V | \mathbf{k}_m \mathbf{k}_n \rangle$.

The Hamiltonian operator is given by

$$\hat{H} = \hat{H}_{el} + \hat{H}_b + \hat{H}_{el-b},$$

with the electronic part

$$\hat{H}_{el} = \sum_{i=1}^N \frac{p_i^2}{2m} + \frac{e^2}{2} \sum_{i \neq j} \frac{e^{-\mu|\mathbf{r}_i - \mathbf{r}_j|}}{|\mathbf{r}_i - \mathbf{r}_j|},$$

where we have introduced an explicit convergence factor (the limit $\mu \rightarrow 0$ is performed after having calculated the various integrals). Correspondingly, we have

$$\hat{H}_b = \frac{e^2}{2} \int \int d\mathbf{r} d\mathbf{r}' \frac{n(\mathbf{r})n(\mathbf{r}')e^{-\mu|\mathbf{r} - \mathbf{r}'|}}{|\mathbf{r} - \mathbf{r}'|},$$

which is the energy contribution from the positive background charge with density $n(\mathbf{r}) = N/\Omega$. Finally,

$$\hat{H}_{el-b} = -\frac{e^2}{2} \sum_{i=1}^N \int d\mathbf{r} \frac{n(\mathbf{r})e^{-\mu|\mathbf{r} - \mathbf{x}_i|}}{|\mathbf{r} - \mathbf{x}_i|},$$

is the interaction between the electrons and the positive background.

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b) Show that

$$\hat{H}_b = \frac{e^2}{2} \frac{N^2}{\Omega} \frac{4\pi}{\mu^2},$$

and

$$\hat{H}_{el-b} = -e^2 \frac{N^2}{\Omega} \frac{4\pi}{\mu^2}.$$

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c) Show thereafter that the final Hamiltonian can be written as

$$H = H_0 + H_I,$$

with

$$H_0 = \sum_{\mathbf{k}\sigma} \frac{\hbar^2 k^2}{2m} a_{\mathbf{k}\sigma}^\dagger a_{\mathbf{k}\sigma},$$

and

$$H_I = \frac{e^2}{2\Omega} \sum_{\sigma_1 \sigma_2} \sum_{\mathbf{q} \neq 0, \mathbf{k}, \mathbf{p}} \frac{4\pi}{q^2} a_{\mathbf{k}+\mathbf{q}, \sigma_1}^\dagger a_{\mathbf{p}-\mathbf{q}, \sigma_2}^\dagger a_{\mathbf{p}\sigma_2} a_{\mathbf{k}\sigma_1}.$$

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d) Calculate $E_0/N = \langle \Phi_0 | H | \Phi_0 \rangle / N$ for for this system to first order in the interaction. Show that, by using

$$\rho = \frac{k_F^3}{3\pi^2} = \frac{3}{4\pi r_0^3},$$

with $\rho = N/\Omega$, r_0 being the radius of a sphere representing the volume an electron occupies and the Bohr radius $a_0 = \hbar^2/e^2m$, that the energy per electron can be written as

$$E_0/N = \frac{e^2}{2a_0} \left[\frac{2.21}{r_s^2} - \frac{0.916}{r_s} \right].$$

Here we have defined $r_s = r_0/a_0$ to be a dimensionless quantity.

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e) Plot your results. Why is this system stable? Calculate thermodynamical quantities like the pressure, given by

$$P = - \left(\frac{\partial E}{\partial \Omega} \right)_N,$$

and the bulk modulus

$$B = -\Omega \left(\frac{\partial P}{\partial \Omega} \right)_N,$$

and comment your results.

We have to show first that

$$\hat{H}_b = \frac{e^2}{2} \frac{N_e^2}{\Omega} \frac{4\pi}{\mu^2},$$

and

$$\hat{H}_{el-b} = -e^2 \frac{N_e^2}{\Omega} \frac{4\pi}{\mu^2}.$$

And then that the final Hamiltonian can be written as

$$H = H_0 + H_I,$$

with

$$H_0 = \sum_{\mathbf{k}\sigma} \frac{\hbar^2 k^2}{2m_e} a_{\mathbf{k}\sigma}^\dagger a_{\mathbf{k}\sigma},$$

and

$$H_I = \frac{e^2}{2\Omega} \sum_{\sigma_1 \sigma_2} \sum_{\mathbf{q} \neq 0, \mathbf{k}, \mathbf{p}} \frac{4\pi}{q^2} a_{\mathbf{k}+\mathbf{q}, \sigma_1}^\dagger a_{\mathbf{p}-\mathbf{q}, \sigma_2}^\dagger a_{\mathbf{p}\sigma_2} a_{\mathbf{k}\sigma_1}.$$

Let us now calculate the following part of the Hamiltonian

$$\hat{H}_b = \frac{e^2}{2} \iint \frac{n(\mathbf{r})n(\mathbf{r}')e^{-\mu|\mathbf{r}-\mathbf{r}'|}}{|\mathbf{r}-\mathbf{r}'|} d\mathbf{r}d\mathbf{r}',$$

where $n(\mathbf{r}) = N_e/\Omega$, the density of the positive background charge. We define $\mathbf{r}_{12} = \mathbf{r} - \mathbf{r}'$, resulting in $d\mathbf{r}_{12} = d\mathbf{r}$, and allowing us to rewrite the integral as

$$\hat{H}_b = \frac{e^2 N_e^2}{2\Omega^2} \iint \frac{e^{-\mu|\mathbf{r}_{12}|}}{|\mathbf{r}_{12}|} d\mathbf{r}_{12} d\mathbf{r}' = \frac{e^2 N_e^2}{2\Omega} \int \frac{e^{-\mu|\mathbf{r}_{12}|}}{|\mathbf{r}_{12}|} d\mathbf{r}_{12}.$$

Here we have used that $\int \mathbf{r} = \Omega$. We change to spherical coordinates and the lack of angle dependencies yields a factor 4π , resulting in

$$\hat{H}_b = \frac{4\pi e^2 N_e^2}{2\Omega} \int_0^\infty r e^{-\mu r} dr.$$

Solving by partial integration

$$\int_0^\infty r e^{-\mu r} dr = \left[-\frac{r}{\mu} e^{-\mu r} \right]_0^\infty + \frac{1}{\mu} \int_0^\infty e^{-\mu r} dr = \frac{1}{\mu} \left[-\frac{1}{\mu} e^{-\mu r} \right]_0^\infty = \frac{1}{\mu^2},$$

gives

$$\hat{H}_b = \frac{e^2 N_e^2}{2} \frac{4\pi}{\Omega \mu^2}.$$

The next term is

$$\hat{H}_{el-b} = -e^2 \sum_{i=1}^N \int \frac{n(\mathbf{r}) e^{-\mu|\mathbf{r}-\mathbf{x}_i|}}{|\mathbf{r}-\mathbf{x}_i|} \mathbf{r}.$$

Inserting $n(\mathbf{r})$ and changing variables in the same way as in the previous integral $\mathbf{y} = \mathbf{r} - \mathbf{x}_i$, we get $d^3\mathbf{y} = d^3\mathbf{r}$. This gives

$$\hat{H}_{el-b} = -\frac{e^2 N_e}{\Omega} \sum_{i=1}^N \int \frac{e^{-\mu|\mathbf{y}|}}{|\mathbf{y}|} d^3\mathbf{y} = -\frac{4\pi e^2 N_e}{\Omega} \sum_{i=1}^N \int_0^\infty y e^{-\mu y} dy. \quad (8)$$

We have already seen this type of integral. The answer is

$$\hat{H}_{el-b} = -\frac{4\pi e^2 N_e}{\Omega} \sum_{i=1}^N \frac{1}{\mu^2},$$

which gives

$$\hat{H}_{el-b} = -e^2 \frac{N_e^2}{\Omega} \frac{4\pi}{\mu^2}.$$

Finally, we need to evaluate \hat{H}_{el} . This term reads

$$\hat{H}_{el} = \sum_{i=1}^{N_e} \frac{\hat{p}_i^2}{2m_e} + \frac{e^2}{2} \sum_{i \neq j} \frac{e^{-\mu|\mathbf{r}_i - \mathbf{r}_j|}}{|\mathbf{r}_i - \mathbf{r}_j|}.$$

The last term represents the repulsion between two electrons. It is a central symmetric interaction and is translationally invariant. The potential is given by the expression

$$v(|\mathbf{r}|) = e^2 \frac{e^{\mu|\mathbf{r}|}}{|\mathbf{r}|},$$

which we derived in connection with the single-particle Hartree-Fock derivation. More material will be added here!

Preparing the ground for numerical calculations; kinetic energy and Ewald term

The kinetic energy operator is

$$\hat{H}_{\text{kin}} = -\frac{\hbar^2}{2m} \sum_{i=1}^N \nabla_i^2, \quad (9)$$

where the sum is taken over all particles in the finite box. The Ewald electron-electron interaction operator can be written as

$$\hat{H}_{ee} = \sum_{i<j}^N v_E(\mathbf{r}_i - \mathbf{r}_j) + \frac{1}{2} N v_0, \quad (10)$$

where $v_E(\mathbf{r})$ is the effective two-body interaction and v_0 is the self-interaction, defined as $v_0 = \lim_{\mathbf{r} \rightarrow 0} \{v_E(\mathbf{r}) - 1/r\}$.

The negative electron charges are neutralized by a positive, homogeneous background charge. Fraser *et al.* explain how the electron-background and background-background terms, \hat{H}_{eb} and \hat{H}_{bb} , vanish when using Ewald's interaction for the three-dimensional electron gas. Using the same arguments, one can show that these terms are also zero in the corresponding two-dimensional system.

Ewald correction term

In the three-dimensional electron gas, the Ewald interaction is

$$v_E(\mathbf{r}) = \sum_{\mathbf{k} \neq \mathbf{0}} \frac{4\pi}{L^3 k^2} e^{i\mathbf{k} \cdot \mathbf{r}} e^{-\eta^2 k^2 / 4} + \sum_{\mathbf{R}} \frac{1}{|\mathbf{r} - \mathbf{R}|} \text{erfc}\left(\frac{|\mathbf{r} - \mathbf{R}|}{\eta}\right) - \frac{\pi \eta^2}{L^3}, \quad (11)$$

where L is the box side length, $\text{erfc}(x)$ is the complementary error function, and η is a free parameter that can take any value in the interval $(0, \infty)$.

Interaction in momentum space

The translational vector

$$\mathbf{R} = L(n_x \mathbf{u}_x + n_y \mathbf{u}_y + n_z \mathbf{u}_z), \quad (12)$$

where \mathbf{u}_i is the unit vector for dimension i , is defined for all integers n_x , n_y , and n_z . These vectors are used to obtain all image cells in the entire real space. The parameter η decides how the Coulomb interaction is divided into a short-ranged and long-ranged part, and does not alter the total function. However, the number of operations needed to calculate the Ewald interaction with a desired accuracy

depends on η , and η is therefore often chosen to optimize the convergence as a function of the simulation-cell size. In our calculations, we choose η to be an infinitesimally small positive number, similarly as was done by [Shepherd *et al.*](#) and [Roggero *et al.*](#).

This gives an interaction that is evaluated only in Fourier space.

When studying the two-dimensional electron gas, we use an Ewald interaction that is quasi two-dimensional. The interaction is derived in three dimensions, with Fourier discretization in only two dimensions. The Ewald effective interaction has the form

$$\begin{aligned}
v_E(\mathbf{r}) = & \sum_{\mathbf{k} \neq \mathbf{0}} \frac{\pi}{L^2 k} \left\{ e^{-kz} \operatorname{erfc} \left(\frac{\eta k}{2} - \frac{z}{\eta} \right) + \right. \\
& e^{kz} \operatorname{erfc} \left(\frac{\eta k}{2} + \frac{z}{\eta} \right) \left. \right\} e^{i\mathbf{k} \cdot \mathbf{r}_{xy}} \\
& + \sum_{\mathbf{R}} \frac{1}{|\mathbf{r} - \mathbf{R}|} \operatorname{erfc} \left(\frac{|\mathbf{r} - \mathbf{R}|}{\eta} \right) \\
& - \frac{2\pi}{L^2} \left\{ z \operatorname{erf} \left(\frac{z}{\eta} \right) + \frac{\eta}{\sqrt{\pi}} e^{-z^2/\eta^2} \right\}, \tag{13}
\end{aligned}$$

where the Fourier vectors \mathbf{k} and the position vector \mathbf{r}_{xy} are defined in the (x, y) plane. When applying the interaction $v_E(\mathbf{r})$ to two-dimensional systems, we set z to zero.

Similarly as in the three-dimensional case, also here we choose η to approach zero from above. The resulting Fourier-transformed interaction is

$$v_E^{\eta=0, z=0}(\mathbf{r}) = \sum_{\mathbf{k} \neq \mathbf{0}} \frac{2\pi}{L^2 k} e^{i\mathbf{k} \cdot \mathbf{r}_{xy}}. \tag{14}$$

The self-interaction v_0 is a constant that can be included in the reference energy.

Antisymmetrized matrix elements in three dimensions

In the three-dimensional electron gas, the antisymmetrized matrix elements are

$$\begin{aligned}
& \langle \mathbf{k}_p m_{s_p} \mathbf{k}_q m_{s_q} | \tilde{v} | \mathbf{k}_r m_{s_r} \mathbf{k}_s m_{s_s} \rangle_{AS} \\
& = \frac{4\pi}{L^3} \delta_{\mathbf{k}_p + \mathbf{k}_q, \mathbf{k}_r + \mathbf{k}_s} \left\{ \delta_{m_{s_p} m_{s_r}} \delta_{m_{s_q} m_{s_s}} (1 - \delta_{\mathbf{k}_p \mathbf{k}_r}) \frac{1}{|\mathbf{k}_r - \mathbf{k}_p|^2} \right. \\
& \quad \left. - \delta_{m_{s_p} m_{s_s}} \delta_{m_{s_q} m_{s_r}} (1 - \delta_{\mathbf{k}_p \mathbf{k}_s}) \frac{1}{|\mathbf{k}_s - \mathbf{k}_p|^2} \right\}, \tag{15}
\end{aligned}$$

where the Kronecker delta functions $\delta_{\mathbf{k}_p \mathbf{k}_r}$ and $\delta_{\mathbf{k}_p \mathbf{k}_s}$ ensure that the contribution with zero momentum transfer vanishes.

Similarly, the matrix elements for the two-dimensional electron gas are

$$\begin{aligned}
& \langle \mathbf{k}_p m_{s_p} \mathbf{k}_q m_{s_q} | v | \mathbf{k}_r m_{s_r} \mathbf{k}_s m_{s_s} \rangle_{AS} \\
&= \frac{2\pi}{L^2} \delta_{\mathbf{k}_p + \mathbf{k}_q, \mathbf{k}_r + \mathbf{k}_s} \left\{ \delta_{m_{s_p} m_{s_r}} \delta_{m_{s_q} m_{s_s}} (1 - \delta_{\mathbf{k}_p \mathbf{k}_r}) \frac{1}{|\mathbf{k}_r - \mathbf{k}_p|} \right. \\
&\quad \left. - \delta_{m_{s_p} m_{s_s}} \delta_{m_{s_q} m_{s_r}} (1 - \delta_{\mathbf{k}_p \mathbf{k}_s}) \frac{1}{|\mathbf{k}_s - \mathbf{k}_p|} \right\}, \tag{16}
\end{aligned}$$

where the single-particle momentum vectors $\mathbf{k}_{p,q,r,s}$ are now defined in two dimensions.

Fock operator

In the calculations, the self-interaction constant is included in the reference energy. We therefore get the Fock-operator matrix elements

$$\langle \mathbf{k}_p | f | \mathbf{k}_q \rangle = \frac{\hbar^2 k_p^2}{2m} \delta_{\mathbf{k}_p, \mathbf{k}_q} + \sum_{\mathbf{k}_i} \langle \mathbf{k}_p \mathbf{k}_i | v | \mathbf{k}_q \mathbf{k}_i \rangle_{AS}. \tag{17}$$

In work of Shepherd *et al.* the matrix elements were defined with the self-interaction constant included in the two-body interaction. This gives Fock-operator matrix elements with a gap constant.

Periodic boundary conditions and single-particle states

When using periodic boundary conditions, the discrete-momentum single-particle basis functions

$$\phi_{\mathbf{k}}(\mathbf{r}) = e^{i\mathbf{k} \cdot \mathbf{r}} / L^{d/2}$$

are associated with the single-particle energy

$$\varepsilon_{n_x, n_y} = \frac{\hbar^2}{2m} \left(\frac{2\pi}{L} \right)^2 (n_x^2 + n_y^2) \tag{18}$$

for two-dimensional systems and

$$\varepsilon_{n_x, n_y, n_z} = \frac{\hbar^2}{2m} \left(\frac{2\pi}{L} \right)^2 (n_x^2 + n_y^2 + n_z^2) \tag{19}$$

for three-dimensional systems.

We choose the single-particle basis such that both the occupied and unoccupied single-particle spaces have a closed-shell structure. This means that all single-particle states corresponding to energies below a chosen cutoff are included in the basis. We study only the unpolarized spin phase, in which all orbitals are occupied with one spin-up and one spin-down electron.

The table illustrates how single-particle energies fill energy shells in a two-dimensional electron box. Here n_x and n_y are the momentum quantum numbers, $n_x^2 + n_y^2$ determines the single-particle energy level, $N_{\uparrow\downarrow}$ represents the cumulated number of spin-orbitals in an unpolarized spin phase, and $N_{\uparrow\uparrow}$ stands for the cumulated number of spin-orbitals in a spin-polarized system.

Magic number for the two-dimensional electron gas

$n_x^2 + n_y^2$	n_x	n_y	$N_{\uparrow\downarrow}$	$N_{\uparrow\uparrow}$
0	0	0	2	1
1	-1	0		
	1	0		
	0	-1		
	0	1	10	5
2	-1	-1		
	-1	1		
	1	-1		
	1	1	18	9
4	-2	0		
	2	0		
	0	-2		
	0	2	26	13
5	-2	-1		
	2	-1		
	-2	1		
	2	1		
	-1	-2		
	-1	2		
	1	-2		
	1	2	42	21

Hartree-Fock energies

Finally, a useful benchmark for our calculations is the expression for the reference energy E_0 per particle. Defining the $T = 0$ density ρ_0 , we can in turn determine in three dimensions the radius r_0 of a sphere representing the volume an electron occupies (the classical electron radius) as

$$r_0 = \left(\frac{3}{4\pi\rho} \right)^{1/3}.$$

In two dimensions the corresponding quantity is

$$r_0 = \left(\frac{1}{\pi\rho} \right)^{1/2}.$$

One can then express the reference energy per electron in terms of the dimensionless quantity $r_s = r_0/a_0$, where we have introduced the Bohr radius $a_0 = \hbar^2/e^2m$. The energy per electron computed with the reference Slater determinant can then be written as (using hereafter only atomic units, meaning that $\hbar = m = e = 1$)

$$E_0/N = \frac{1}{2} \left[\frac{2.21}{r_s^2} - \frac{0.916}{r_s} \right],$$

for the three-dimensional electron gas. For the two-dimensional gas the corresponding expression is (show this)

$$E_0/N = \frac{1}{r_s^2} - \frac{8\sqrt{2}}{3\pi r_s}.$$

For an infinite homogeneous system, there are some particular simplifications due to the conservation of the total momentum of the particles. By symmetry considerations, the total momentum of the system has to be zero. Both the kinetic energy operator and the total Hamiltonian \hat{H} are assumed to be diagonal in the total momentum \mathbf{K} . Hence, both the reference state Φ_0 and the correlated ground state Ψ must be eigenfunctions of the operator $\hat{\mathbf{K}}$ with the corresponding eigenvalue $\mathbf{K} = \mathbf{0}$. This leads to important simplifications to our different many-body methods. In coupled cluster theory for example, all terms that involve single particle-hole excitations vanish.