

Hierarchical Bayesian Modelling for an Atmospheric Limb-Sounder Measuring Ozone



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Abstract

In this thesis we develop a *hierarchical* Bayesian model for an atmospheric limb-sounder targeting ozone, as in [1]. To measure most effectively, we briefly assess the informativity of the forward model for different measurement approaches and adjust the data collection accordingly. As in [2], we provide posterior distributions of hyper-parameters and ozone profiles utilising the marginal and then conditional scheme, and compare to a Tikhonov regularisation approach. Furthermore, we extend our model and the marginal and then conditional scheme to infer posterior pressure, temperature and ozone jointly. The main contribution of this work is that we provide a novel approach [3] to approximate high-dimensional posterior distributions in the tensor-train format. This allows us to generate samples from the target distribution with far fewer function evaluations than the `t-walk` sampling algorithm [4] and to quickly calculate marginals from the target distribution, which we use for quadrature to avoid sampling-based characterisation of functions.

Our results show that a hierarchical Bayesian approach, which quantifies posterior mean and variance of the parameter (ozone), provides more information about the parameter than a regularisation approach within a similar computation time. In regions where the signal strength is low and the data is noise-dominated we can not recover ozone structures from the ground truth. When including pressure and temperature describing hyper-parameters within our hierarchical Bayesian model, we show that ozone and pressure are highly correlated, whereas the model and data are uninformative about temperature. For future work, we recommend developing a more physically based model for ozone within the atmosphere.

Compared to conventional samplers tensor-train methods require a predefined grid and function outputs within computer precision, but reduce the function evaluations per independent sample significantly, once the grid and the "normalisation constant" are defined. Another advantage of the tensor-train format is that marginal distributions, useful for quadrature, can be calculated at a low computational cost. To further improve tensor-train methods, we think it is essential to be able to *automatically* lower tensor ranks, calculate "normalisation constants" to avoid numerical issues and reduce correlation structures between parameters, all of which we currently have to do by exploratory analysis. Additionally, choosing accurate interpolation schemes between grid points may be crucial to improving the effectiveness of the approximation.

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List of Abbreviations

CDF	Cumulative Distribution Function
DAG	Directed Acyclic Graph
HITRAN	High Resolution Transmission
IACT	Integrated Autocorrelation Time
IRT	Inverse Rosenblatt Transform
L	Linear
MCMC	Markov Chain Monte-Carlo
MH	Metropolis-Hastings
MLS	Microwave Limb Sounder
MTC	Marginal and Then Conditional
MVN	Multivariate Normal
MWG	Metropolis Wihtin Gibbs
NASA	National Aeronautics and Space Administration
NL	Non-Linear
RMS	Root Mean Square
RTE	Radiative Transfer Equation
RTO	Randomise Then Optimise
SIRT	Squared Inverse Rosenblatt Transform
STD	Standard Deviation
SVD	Singular Value Decomposition
TT	Tensor-Train
VMR	Volume Mixing Ratio

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1

Introduction

Here, we briefly describe the currently used standard to retrieve atmospheric trace gas concentrations from limb-sounding measurements and what motivates us to employ a hierarchical Bayesian framework addressing this inverse problem. We explain how our approach contributes to and improves upon existing methods. Lastly, we provide the reader with the thesis structure.

1.1 Motivation

The only currently operating ozone limb sounder, the Microwave Limb Sounder (MLS) on NASA’s Aura satellite, is gradually drifting from its orbit and scheduled to be phased out by 2026 [5]. A group led by Harald Schwefel has proposed an alternative approach to fill this observational gap using a much smaller platform, such as a disk-shaped resonator mounted on a 6U CubeSat [6]. The proposed system targets a narrow frequency band and converts the thermal radiation emitted by ozone from the terahertz region to the optical domain [7, 8]. This frequency conversion offers a cost-effective and energy-efficient solution, as it circumvents the need for large, energy-hungry cooling devices that are traditionally required to capture terahertz signals. Instead, signal acquisition in the optical domain can be implemented by using compact, cheap, and low-power photonic technologies.

Currently, the inverse problem to retrieve any trace gas from limb-sounding data is approached by the atmospheric physics community using optimisation and regularisation techniques developed in the 1970s [9, 10]. These methods focus on finding the “best fit to data but not the best fit to parameters” [11]. Instead, we employ a hierarchically structured Bayesian framework to provide a distribution of ozone profiles, which represents a range of feasible solutions to some given data. This probabilistic approach allows us to determine meaningful estimates and uncertainties of parameters.

1.2 Research Gap and Contribution

As already mentioned, currently the MLS retrieval algorithm [12] is based on the “optimal estimation” method from [9]. This approach provides a point estimate by fitting parameters to some data and iteratively minimising a squared residual norm, penalised against a chosen regularisation. This does not provide comprehensive information about the parameters’ underlying correlation structures, can lead to unphysical results, e.g. negative ozone concentration values [13], and biased solutions, where the bias is then removed based on empirical decisions [14, 15]. Errors are provided by a local derivative of the forward map at the optimal solution, which is inherently highly sensitive to that specific point in the parameter space. Furthermore, these regularisation methods condition on external point estimates of other parameters, such as temperature or pressure [12].

Naturally, noise (hyper-parameter) is a random process and follows a probability distribution, which we assume to know. According to that noise, as well as some other probability distributions, we can formulate explicit functions over hyper-parameters and parameters (e.g. ozone concentrations). By incorporating hyper-parameters, e.g. measurement noise and smoothness of the ozone profile, and parameters in the modelling and inversion process, we are able to provide errors and a range of feasible solutions (posterior distribution) given some data, instead of one "optimal" solution. This approach is known as *hierarchical* Bayesian modelling. Livesey et al. [12] report "unexpected spectrally correlated noise" on the MLS aura, so here is another real reason why one should include noise in the model.

To solve this inverse problem we employ a linear-Gaussian hierarchical Bayesian framework, and utilise the marginal and then conditional (MTC) method [2]. Within seconds we can evaluate the marginal posterior distribution over the hyper-parameters and the conditional posterior distribution over the parameters. This is a relatively new method within the Bayesian community, and we are the first to apply it to a forward model based on the radiative transfer equation (RTE).

Moreover, instead of using sample-based methods to characterise those posterior distributions, we are the first, to our knowledge, to approximate posterior distributions directly utilising the tensor-train (TT) format. This enables us to provide estimates and uncertainties via quadrature or the inverse Rosenblatt transform (IRT), with far fewer function evaluations compared to conventional samplers.

Since the RTE is weakly non-linear, we approximate the RTE with an affine map, which seems to be another novelty in the field of atmospheric remote sensing.

Additionally, we extend the MTC method to tackle this inverse problem by jointly inferring pressure, temperature and ozone profiles given one set of measurements.

1.3 Thesis Structure

In Chapter 2, we give a brief overview of the key methods used along with references for further reading. Chapter 3 introduces the simplified forward model based on the RTE, and discusses strategies for measuring effectively. Based on this, we simulate noisy data for an idealised limb sounder within a set atmosphere. Then, in Chapter 4, we construct our linear-Gaussian hierarchical Bayesian model and discuss some prior modelling choices. Given the simulated data, we apply the MTC method to provide posterior distributions of our Bayesian framework based on the linearised RTE to then approximate the non-linear forward model with an affine map. We compare a regularisation solution with the posterior distributions of the approximated linear forward model and a ground truth ozone profile. Further, we extend the previously build Bayesian model to include hyper-parameters corresponding to pressure and temperature and touch on some prior modelling choices. Again, using the MTC method we jointly estimate posterior ozone, pressure and temperature profiles and highlight important aspects for improving the efficiency and stability of TT approximations. Lastly, we summarise and discuss our results and provide an outlook for future work, see Chapter 5.

2

Theoretical and Technical Background

In this chapter, we introduce and briefly derive the methods used throughout this thesis, and provide references for further reading. We keep it as general as possible, as the expressions specifically tailored towards the forward map will be presented in the results Chapter 4. We begin by outlining a general hierarchical Bayesian approach to inverse problems. This is followed by some fundamentals of Markov Chain Monte Carlo (MCMC) methods, a specific example of a Metropolis-Hastings algorithm and the randomise then optimise (RTO) method. Further, we present the TT format, which allows to approximate high-dimensional functions and calculate marginal probability distributions cheaply. Lastly, we provide some background information on the Wasserstein distance, affine maps and the Tikhonov regularisation method.

2.1 Hierarchical Bayesian Inference

Assume we observe some data

$$\mathbf{y} = \mathbf{A}(\mathbf{x}) + \boldsymbol{\eta}, \quad (2.1)$$

based on a forward model $\mathbf{A}(\mathbf{x})$, which may be non-linear, a unknown parameter vector \mathbf{x} and some additive random noise vector $\boldsymbol{\eta}$. Naturally, due to the noise, which we classify through a hyper-parameter, we deal with a random process, and we wish to incorporate that in our hierarchically ordered modelling. Then define the likelihood function $\pi(\mathbf{y}|\mathbf{x}, \boldsymbol{\theta})$ according to the nature of the noise as well as all relevant information about the measurement process, captured by the model $\mathbf{A}(\mathbf{x})$. We read $\pi(\mathbf{y}|\mathbf{x}, \boldsymbol{\theta})$ as the distribution over \mathbf{y} conditioned on \mathbf{x} and $\boldsymbol{\theta}$. If we write $\boldsymbol{\eta} \sim \pi_{\boldsymbol{\eta}}(\cdot|\boldsymbol{\theta})$, \sim reads as “is distributed as”. Here $\boldsymbol{\theta}$ may account for multiple hyper-parameters, e.g. describing

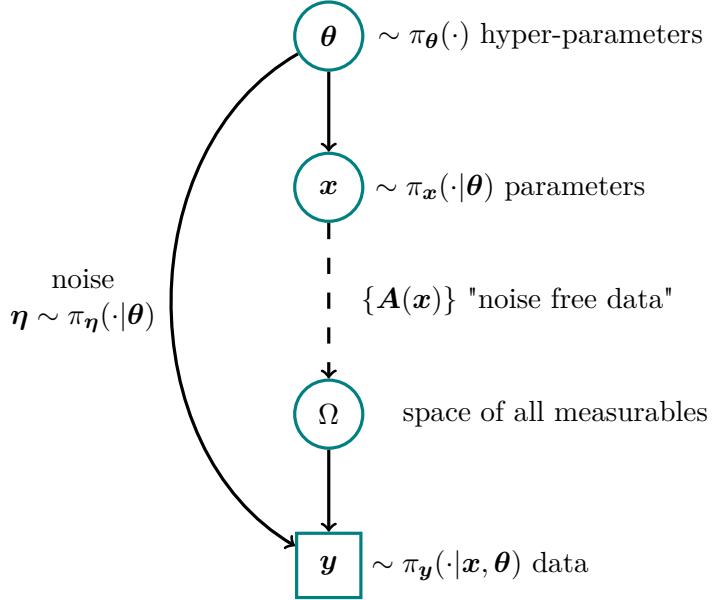


Figure 2.1: The directed acyclic graph (DAG) for an inverse problem visualises statistical dependencies as solid line arrows and deterministic dependencies as dotted arrows. The hyper-parameters θ are distributed as the hyper-prior distribution $\pi(\theta)$. The prior distribution $\pi_x(\cdot|\theta)$ for the parameter x and the noise $\eta \sim \pi_\eta(\cdot|\theta)$ are statistically dependent on some of those hyper-parameters. Then a parameter $x \sim \pi_x(\cdot|\theta)$ is mapped onto the space of all measurable $\Omega = A(x)$ deterministically through the forward model. From the space of all measurable noise-free data we observe a data set $y = A(x) + \eta$ with some additive random noise, which determines the likelihood function $\pi(y|x, \theta)$.

the distribution $\pi_\eta(\cdot|\theta)$ over the noise vector η , and describing physical properties or functional dependencies of x , e.g. smoothness, through the prior distribution $\pi(x|\theta)$. Consequently we define a hyper-prior distribution $\pi(\theta)$, where $\pi(x, \theta) = \pi(x|\theta)\pi(\theta)$. Choosing these prior distributions is ultimately a modeller's choice and crucial, as it shall not affect the posterior distribution

$$\pi(x, \theta|y) = \frac{\pi(y|x, \theta)\pi(x, \theta)}{\pi(y)} \propto \pi(y|x, \theta)\pi(x, \theta), \quad (2.2)$$

which, according to Bayes' theorem, gives us a distribution of x and θ given (conditioned on) the data, with finite $\pi(y)$. Note that here we include the hyper-parameters within the posterior distribution, which is the key idea of hierarchical Bayesian modelling, as we not only aim to quantify the posterior distribution over the parameters x , but also the posterior distribution over the hyper-parameter θ . We can visualise this hierarchically-ordered correlation structure between parameters as well as how distributions progress through a measurement process, using a directed acyclic graph (DAG), see Figure 2.1.

Usually, the objective is to calculate the expectation of a function $h(x)$, which is described

$$E_{x, \theta|y}[h(x)] = \underbrace{\int \int h(x) \pi(x, \theta|y) dx d\theta}_{\mu_{\text{int}}}. \quad (2.3)$$

If that is a high-dimensional integral and computationally not feasible to solve, we approximate

$$\mathbb{E}_{\boldsymbol{x}, \boldsymbol{\theta} | \boldsymbol{y}}[h(\boldsymbol{x})] \approx \underbrace{\frac{1}{N} \sum_{k=1}^N h(\boldsymbol{x}^{(k)})}_{\boldsymbol{\mu}_{\text{samp}}}, \quad (2.4)$$

with an unbiased sample-based Monte-Carlo estimate [16] for large enough N (law of large numbers [17, Chapter 17]). Here, the posterior samples $\{\boldsymbol{x}^{(k)}, \boldsymbol{\theta}^{(k)}\} \sim \pi_{\boldsymbol{x}, \boldsymbol{\theta}}(\cdot | \boldsymbol{y})$, for $k = 1, \dots, N$, form a sample set $\mathcal{M} = \{(\boldsymbol{x}, \boldsymbol{\theta})^{(1)}, \dots, (\boldsymbol{x}, \boldsymbol{\theta})^{(N)}\}$. The central limit theorem states that the sample mean $\boldsymbol{\mu}_{\text{samp}}^{(i)} = \mathbb{E}_{\boldsymbol{x}, \boldsymbol{\theta} | \boldsymbol{y}}[h(\boldsymbol{x})]$ of independent sample sets $\mathcal{M}^{(i)}$ for $i = 1, \dots, n$ of any distribution, converges to be normally distributed, so that

$$\sqrt{n}(\boldsymbol{\mu}_{\text{samp}}^{(i)} - \boldsymbol{\mu}_{\text{int}}) \xrightarrow{\mathcal{D}} \mathcal{N}(0, \sigma^2)[18], \quad (2.5)$$

and if $\sigma^2 < \infty$ the Monte-Carlo error $\boldsymbol{\mu}_{\text{samp}}^{(i)} - \boldsymbol{\mu}_{\text{int}}$ is bounded. In practise, we approximate the estimation error from a sample set $\mathcal{M}^{(i)}$ as

$$(\sigma^{(i)})^2 = \text{var}(\mathbb{E}_{\boldsymbol{x}, \boldsymbol{\theta} | \boldsymbol{y}}[h(\boldsymbol{x})]) \approx \frac{\text{var}(h(\boldsymbol{x}))}{N} \left(1 + 2 \underbrace{\sum_{t=1}^W \frac{\Gamma(t)}{\Gamma(0)}}_{:= 2\tau_{\text{int}}} \right) = \text{var}(h(\boldsymbol{x})) \frac{2\tau_{\text{int}}}{N}, \quad (2.6)$$

where we have to consider that the samples generated by any system or algorithm are correlated. We define the integrated autocorrelation time (IACT) τ_{int} as in [19, pp. 103-105]. Here the autocorrelation coefficient $\Gamma(t) \propto \exp\{-|t|/\tau\} \rightarrow 0$ for $t \rightarrow \infty$ at lag t decays exponentially and $\Gamma(0) = \text{var}(h(\boldsymbol{x}))$. Choosing the summation window W is crucial because it has to be large compared to the decay time τ , but for too large t the autocorrelation coefficient $\Gamma(t)$ is noise-dominated. U. Wolff [20] (and the Python implementation by D. Hesse [21]) provide a way to not only calculate the IACT safely but also to quantify the errors of the estimated IACT.

The IACT provides a good estimate of the number of steps the sampling algorithm needs to take to produce one independent sample. According to the IACT, we define the effective sample size as $2\tau_{\text{int}}/N$. We point out that for uncorrelated samples $2\tau_{\text{int}} = 1$ the error $(\sigma^{(i)})^2$ is a typical Monte-Carlo estimate. See Appendix A.2 and [19, 20, 22] for a more detailed derivation.

2.1.1 Marginal and then Conditional Method

Generating a representative sample set quickly from the posterior distribution often presents a significant challenge. This is mainly due to the strong correlations that usually exist between the parameters and hyper-parameters, as discussed by Rue and Held in [23] and illustrated in Appendix A.1. If \boldsymbol{x} cannot be parametrised directly

in terms of the hyper-parameters $\boldsymbol{\theta}$, so that $\mathbf{x}(\boldsymbol{\theta})$ is function of $\boldsymbol{\theta}$, it is beneficial to factorise the posterior distribution as

$$\pi(\mathbf{x}, \boldsymbol{\theta} | \mathbf{y}) = \pi(\mathbf{x} | \boldsymbol{\theta}, \mathbf{y}) \pi(\boldsymbol{\theta} | \mathbf{y}), \quad (2.7)$$

into the conditional posterior $\pi(\mathbf{x} | \boldsymbol{\theta}, \mathbf{y})$ over the latent field \mathbf{x} and the marginal posterior

$$\pi(\boldsymbol{\theta} | \mathbf{y}) = \frac{\pi(\mathbf{y} | \mathbf{x}, \boldsymbol{\theta}) \pi(\mathbf{x} | \boldsymbol{\theta}) \pi(\boldsymbol{\theta})}{\pi(\mathbf{x} | \boldsymbol{\theta}, \mathbf{y}) \pi(\mathbf{y})} \propto \frac{\pi(\mathbf{y} | \mathbf{x}, \boldsymbol{\theta}) \pi(\mathbf{x} | \boldsymbol{\theta}) \pi(\boldsymbol{\theta})}{\pi(\mathbf{x} | \boldsymbol{\theta}, \mathbf{y})}, \quad (2.8)$$

over the hyper-parameter $\boldsymbol{\theta}$ (see [2, Lemma 2]).

This approach, known as the MTC method, is particularly advantageous when $\mathbf{x} \in \mathbb{R}^n$ is high-dimensional, while $\boldsymbol{\theta} \in \mathbb{R}^{n_\theta}$ is low-dimensional, so that $n_\theta \ll n$ and evaluation of $\pi(\boldsymbol{\theta} | \mathbf{y})$ is cheap. Applying the law of total expectation [24], Eq. (2.3) becomes

$$\mathbb{E}_{\mathbf{x}, \boldsymbol{\theta} | \mathbf{y}}[h(\mathbf{x})] = \int \int h(\mathbf{x}) \pi(\mathbf{x} | \boldsymbol{\theta}, \mathbf{y}) d\mathbf{x} \pi(\boldsymbol{\theta} | \mathbf{y}) d\boldsymbol{\theta} \quad (2.9)$$

$$= \int \mathbb{E}_{\mathbf{x} | \boldsymbol{\theta}, \mathbf{y}}[h(\mathbf{x})] \pi(\boldsymbol{\theta} | \mathbf{y}) d\boldsymbol{\theta} \quad (2.10)$$

$$= \mathbb{E}_{\boldsymbol{\theta} | \mathbf{y}} \left[\mathbb{E}_{\mathbf{x} | \boldsymbol{\theta}, \mathbf{y}}[h(\mathbf{x})] \right], \quad (2.11)$$

where, in the case of a linear-Gaussian hierarchical Bayesian model, both the marginal distribution $\pi(\boldsymbol{\theta} | \mathbf{y})$ and the inner expectation $\mathbb{E}_{\mathbf{x} | \boldsymbol{\theta}, \mathbf{y}}[h(\mathbf{x})]$ are well defined (see next subsection). If the integral in Eq. 2.10 is expensive to calculate, we use sample based methods to produce a Markov chain $\{(\mathbf{x}, \boldsymbol{\theta})^{(1)}, \dots, (\mathbf{x}, \boldsymbol{\theta})^{(N)}\} \sim \pi(\mathbf{x}, \boldsymbol{\theta} | \mathbf{y})$ and sample from $\pi(\boldsymbol{\theta} | \mathbf{y})$ first and then draw samples from the full conditional posterior $\pi(\mathbf{x} | \boldsymbol{\theta}, \mathbf{y})$.

Linear-Gaussian hierarchical Bayesian model

In case of normally distributed noise $\boldsymbol{\eta} \sim \mathcal{N}(0, \boldsymbol{\Sigma}(\boldsymbol{\theta}))$, with zero mean and covariance $\boldsymbol{\Sigma}(\boldsymbol{\theta})$, and a linear model matrix \mathbf{A} , the data is given as

$$\mathbf{y} = \mathbf{A}\mathbf{x} + \boldsymbol{\eta}, \quad (2.12)$$

and we can derive the marginal and conditional posterior distribution explicitly. We define our hierarchical Bayesian model as

$$\mathbf{y} | \mathbf{x}, \boldsymbol{\theta} \sim \mathcal{N}(\mathbf{Ax}, \boldsymbol{\Sigma}(\boldsymbol{\theta})) \quad (2.13a)$$

$$\mathbf{x} | \boldsymbol{\theta} \sim \mathcal{N}(\boldsymbol{\mu}, \mathbf{Q}(\boldsymbol{\theta})^{-1}) \quad (2.13b)$$

$$\boldsymbol{\theta} \sim \pi(\boldsymbol{\theta}), \quad (2.13c)$$

with a Gaussian likelihood function $\pi(\mathbf{y} | \mathbf{x}, \boldsymbol{\theta})$, the prior mean $\boldsymbol{\mu}$, prior precision $\mathbf{Q}(\boldsymbol{\theta})$ and a hyper-prior distribution $\pi(\boldsymbol{\theta})$. To derive the marginal posterior and the full conditional posterior distribution, we consider the joint multivariate Gaussian distribution

$$\begin{pmatrix} \mathbf{x} \\ \mathbf{y} \end{pmatrix} \sim \mathcal{N} \left[\begin{pmatrix} \boldsymbol{\mu} \\ \mathbf{A}\boldsymbol{\mu} \end{pmatrix}, \begin{pmatrix} \mathbf{Q}(\boldsymbol{\theta}) + \mathbf{A}^T \boldsymbol{\Sigma}(\boldsymbol{\theta})^{-1} \mathbf{A} & -\mathbf{A}^T \boldsymbol{\Sigma}(\boldsymbol{\theta})^{-1} \\ \boldsymbol{\Sigma}(\boldsymbol{\theta})^{-1} \mathbf{A} & \boldsymbol{\Sigma}(\boldsymbol{\theta})^{-1} \end{pmatrix}^{-1} \right], \quad (2.14)$$

where we provide the joint precision matrix as in [25] (see also [2, 23]). Immediately, we formulate the conditional posterior as

$$\mathbf{x}|\boldsymbol{\theta}, \mathbf{y} \sim \mathcal{N}\left(\underbrace{\boldsymbol{\mu} + (\mathbf{Q}(\boldsymbol{\theta}) + \mathbf{A}^T \boldsymbol{\Sigma}(\boldsymbol{\theta})^{-1} \mathbf{A})^{-1}(\mathbf{y} - \mathbf{A}\boldsymbol{\mu})}_{\boldsymbol{\mu}_{\mathbf{x}|\boldsymbol{\theta}, \mathbf{y}}}, \underbrace{(\mathbf{Q}(\boldsymbol{\theta}) + \mathbf{A}^T \boldsymbol{\Sigma}(\boldsymbol{\theta})^{-1} \mathbf{A})^{-1}}_{\boldsymbol{\Sigma}_{\mathbf{x}|\boldsymbol{\theta}, \mathbf{y}}}\right). \quad (2.15)$$

Then the marginal posterior distribution over the hyper-parameters can be derived as in Eq. 2.8, where, as noted by Fox et al. [2], the parameter \mathbf{x} cancels and we arrive at

$$\begin{aligned} \pi(\boldsymbol{\theta}|\mathbf{y}) \propto & \sqrt{\frac{\det(\boldsymbol{\Sigma}(\boldsymbol{\theta})^{-1}) \det(\mathbf{Q}(\boldsymbol{\theta}))}{\det(\mathbf{Q}(\boldsymbol{\theta}) + \mathbf{A}^T \boldsymbol{\Sigma}(\boldsymbol{\theta})^{-1} \mathbf{A})}} \exp\left\{-\frac{1}{2}(\mathbf{y} - \mathbf{A}\boldsymbol{\mu})^T \right. \\ & \left. [\boldsymbol{\Sigma}(\boldsymbol{\theta})^{-1} - \boldsymbol{\Sigma}(\boldsymbol{\theta})^{-1} \mathbf{A}(\mathbf{Q}(\boldsymbol{\theta}) + \mathbf{A}^T \boldsymbol{\Sigma}(\boldsymbol{\theta})^{-1} \mathbf{A})^{-1} \mathbf{A}^T \boldsymbol{\Sigma}(\boldsymbol{\theta})^{-1}] (\mathbf{y} - \mathbf{A}\boldsymbol{\mu})\right\} \pi(\boldsymbol{\theta}). \end{aligned} \quad (2.16)$$

Having the marginal posterior distribution $\pi(\boldsymbol{\theta}|\mathbf{y})$ (independent of \mathbf{x}) available breaks up the correlation structure between \mathbf{x} and $\boldsymbol{\theta}$ (see Appendix A.1), and makes the MTC approach very efficient [2]. Within this scheme, we evaluate the marginal posterior first and then either condition on hyper-parameters to draw full conditional posterior samples $\mathbf{x} \sim \pi(\mathbf{x}|\boldsymbol{\theta}, \mathbf{y})$ (see Sec. 2.2.3) or evaluate the mean

$$\boldsymbol{\mu}_{\mathbf{x}|\mathbf{y}} = \int \boldsymbol{\mu}_{\mathbf{x}|\boldsymbol{\theta}, \mathbf{y}} \pi(\boldsymbol{\theta}|\mathbf{y}) d\boldsymbol{\theta} \quad (2.17)$$

and covariance matrix

$$\boldsymbol{\Sigma}_{\mathbf{x}|\mathbf{y}} = \int \boldsymbol{\Sigma}_{\mathbf{x}|\boldsymbol{\theta}, \mathbf{y}} \pi(\boldsymbol{\theta}|\mathbf{y}) d\boldsymbol{\theta} \quad (2.18)$$

of the full posterior $\pi(\mathbf{x}|\mathbf{y})$ by some quadrature rule.

2.2 Sampling Methods

In this section we present the methods that draw samples from the marginal posterior $\mathcal{M} = \{\boldsymbol{\theta}^{(1)}, \dots, \boldsymbol{\theta}^{(k)}, \dots, \boldsymbol{\theta}^{(N)}\} \sim \pi(\boldsymbol{\theta}|\mathbf{y})$ as well as the RTO method to draw samples from a normally distributed full conditional posterior $\pi(\mathbf{x}|\boldsymbol{\theta}, \mathbf{y})$. Here, \mathcal{M} denotes a Markov chain, where each new sample $\boldsymbol{\theta}^{(k)}$ is only affected by the previous one, $\boldsymbol{\theta}^{(k-1)}$. MCMC methods generate such a chain \mathcal{M} using random (Monte-Carlo) proposals $(\mathbf{x}, \boldsymbol{\theta})^{(k)} \sim q(\cdot|\boldsymbol{\theta}^{(k-1)})$ according to a proposal distribution conditioned on the previous sample (Markov), where ergodicity of the chain \mathcal{M} is a sufficient criterion for using sample-based estimates [11, 16].

The ergodicity theorem in [11] states that, if a Markov chain \mathcal{M} is aperiodic, irreducible, and reversible, then it converges to a unique stationary equilibrium distribution. In other words, the chain can reach any state from any other state (irreducibility), is not stuck in periodic cycles (aperiodicity), and satisfies the detailed balance condition [11]

(reversibility). Then the samples in that chain $\mathcal{M} \sim \pi(\boldsymbol{\theta}|\mathbf{y})$ are samples from the desired target distribution. In practice, one can inspect the trace $\pi(\boldsymbol{\theta}^{(k)}|\mathbf{y})$ for $k = 1, \dots, N$ and visually assess convergence and mixing properties of the chain to evaluate ergodicity. The sampling methods used in this thesis possess proven ergodic properties, and we therefore refer the reader to the corresponding literature for further details.

If the Markov chain over the marginal posterior $\pi(\boldsymbol{\theta}|\mathbf{y})$ is ergodic, and the full conditional samples $\mathbf{x}^{(k)} \sim \pi(\mathbf{x}|\boldsymbol{\theta}^{(k)}, \mathbf{y})$ are drawn independently, as described in Sec. 2.2.3, then the resulting joint chain $\{(\mathbf{x}, \boldsymbol{\theta})^{(1)}, \dots, (\mathbf{x}, \boldsymbol{\theta})^{(N)}\} \sim \pi(\mathbf{x}, \boldsymbol{\theta}|\mathbf{y}) = \pi(\mathbf{x}|\boldsymbol{\theta}, \mathbf{y})\pi(\boldsymbol{\theta}|\mathbf{y})$ is also ergodic [26].

2.2.1 Metropolis within Gibbs

If $\boldsymbol{\theta} \in \mathbb{R}^2$ we use a Metropolis-within-Gibbs (MWG) sampler as described in [2] to sample from $\pi(\boldsymbol{\theta}|\mathbf{y})$. With $\boldsymbol{\theta} = (\theta_1, \theta_2)$, we perform a Metropolis step in the θ_1 direction and a Gibbs step in the θ_2 direction. Ergodicity for this approach is proven in [27].

The MWG algorithm starts at the initial guess $\boldsymbol{\theta}^{(t)}$ at $t = 0$. We then propose a new sample $\theta_1 \sim q(\theta_1|\theta_1^{(t-1)})$, conditioned on the previous state, using a symmetric proposal distribution $q(\theta_1|\theta_1^{(t-1)}) = q(\theta_1^{(t-1)}|\theta_1)$, which is a special case of the Metropolis-Hastings algorithm [27]. We accept and set $\theta_1^{(t)} = \theta_1$ with the acceptance probability

$$\alpha(\theta_1|\theta_1^{(t-1)}) = \min \left\{ 1, \frac{\pi(\theta_1|\theta_2^{(t-1)}, \mathbf{y}) q(\theta_1^{(t-1)}|\theta_1)}{\pi(\theta_1^{(t-1)}|\theta_2^{(t-1)}, \mathbf{y}) q(\theta_1|\theta_1^{(t-1)})} \right\} \quad (2.19)$$

or reject and keep $\theta_1^{(t)} = \theta_1^{(t-1)}$, which we do by comparing α to a uniform random number $u \sim \mathcal{U}(0, 1)$.

Next, we perform a Gibbs step in the θ_2 direction, where Gibbs sampling is again a special case of the Metropolis-Hastings algorithm with acceptance probability equal to one, and draw the next sample $\theta_2^{(t)} \sim \pi(\cdot|\theta_1^{(t)}, \mathbf{y})$, conditioned on the current value $\theta_1^{(t)}$.

We repeat this procedure N' times and ensure convergence independently of the initial sample (irreducibility) by discarding $N_{\text{burn-in}}$ initial samples after a burn-in period, resulting in a Markov chain of length $N = N' - N_{\text{burn-in}}$.

2.2.2 T-walk Sampler as a Black Box

If the number of hyper-parameters is large, we sample from the marginal posterior $\pi(\boldsymbol{\theta}|\mathbf{y})$ using the **t-walk** sampler as by Christen and Fox [4], because it is an easy-to-use and implement sampler. The **t-walk** chooses between four different types of steps on the target distribution and is employed as a black-box algorithm in default settings, requiring the specification of the number of samples, burn-in period, support region, and the target distribution. Convergence to the target distribution is guaranteed by the construction of this algorithm.

Algorithm 1: Metropolis within Gibbs

```

1: Initialise and suppose two dimensional vector  $\boldsymbol{\theta}^{(0)} = (\theta_1^{(0)}, \theta_2^{(0)})$ 
2: for  $k = 1, \dots, N'$  do
3:   Propose  $\theta_1 \sim q(\cdot | \theta_1^{(t-1)}) = q(\theta_1^{(t-1)} | \cdot)$ 
4:   Compute

$$\alpha(\theta_1 | \theta_1^{(t-1)}) = \min \left\{ 1, \frac{\pi(\theta_1 | \theta_2^{(t-1)}, \mathbf{y}) q(\theta_1^{(t-1)} | \theta_1)}{\pi(\theta_1^{(t-1)} | \theta_2^{(t-1)}, \mathbf{y}) q(\theta_1 | \theta_1^{(t-1)})} \right\}$$

5:   Draw  $u \sim \mathcal{U}(0, 1)$ 
6:   if  $\alpha \geq u$  then
7:     Accept and set  $\theta_1^{(t)} = \theta_1$ 
8:   else
9:     Reject and keep  $\theta_1^{(t)} = \theta_1^{(t-1)}$ 
10:  end if
11:  Draw  $\theta_2^{(t)} \sim \pi(\cdot | \theta_1^{(t)}, \mathbf{y})$ 
12: end for
13: Output:  $\boldsymbol{\theta}^{(0)}, \dots, \boldsymbol{\theta}^{(k)}, \dots, \boldsymbol{\theta}^{(N)} \sim \pi(\boldsymbol{\theta} | \mathbf{y})$ 

```

2.2.3 Randomise then Optimise

If it is computationally not feasible to calculate the mean and the covariance matrix of the full posterior (see Eq. 2.17 and 2.18) via quadrature due to a large number of hyper-parameters, e.g. $\boldsymbol{\theta} \in \mathbb{R}^{n_\theta}$ and $n_\theta \geq 4$, we need an alternative way to draw samples from $\pi(\mathbf{x} | \boldsymbol{\theta}, \mathbf{y})$. For the linear-Gaussian Bayesian model, we can draw samples from the normal distribution $\pi(\mathbf{x} | \boldsymbol{\theta}, \mathbf{y})$ using the RTO [28], conditioned on $\boldsymbol{\theta}$.

The conditional distribution, as in Eq. 2.15, can be rewritten as

$$\pi(\mathbf{x} | \boldsymbol{\theta}, \mathbf{y}) \propto \pi(\mathbf{y} | \mathbf{x}, \boldsymbol{\theta}) \pi(\mathbf{x} | \boldsymbol{\theta}) \quad (2.20)$$

$$\propto \exp \left(-\frac{1}{2} (\mathbf{A}_\theta \mathbf{x} - \mathbf{y})^T \boldsymbol{\Sigma}_\theta^{-1} (\mathbf{A}_\theta \mathbf{x} - \mathbf{y}) \right) \exp \left(-\frac{1}{2} (\boldsymbol{\mu} - \mathbf{x})^T \mathbf{Q}_\theta (\boldsymbol{\mu} - \mathbf{x}) \right), \quad (2.21)$$

$$= \exp \left(-\frac{1}{2} \left\| \hat{\mathbf{A}}\mathbf{x} - \hat{\mathbf{y}} \right\|_{L^2}^2 \right), \quad (2.22)$$

where we define

$$\hat{\mathbf{A}} := \begin{bmatrix} \boldsymbol{\Sigma}_\theta^{-1/2} \mathbf{A}_\theta \\ \mathbf{Q}_\theta^{1/2} \end{bmatrix}, \quad \hat{\mathbf{y}} := \begin{bmatrix} \boldsymbol{\Sigma}_\theta^{-1/2} \mathbf{y} \\ \mathbf{Q}_\theta^{1/2} \boldsymbol{\mu} \end{bmatrix} \quad [29, 30], \quad (2.23)$$

with $\mathbf{A}(\boldsymbol{\theta}) := \mathbf{A}_\theta$, $\mathbf{Q}(\boldsymbol{\theta}) := \mathbf{Q}_\theta$ and $\boldsymbol{\Sigma}^{-1}(\boldsymbol{\theta}) := \boldsymbol{\Sigma}_\theta^{-1}$, which are all dependent on the hyper-parameters $\boldsymbol{\theta}$. A sample $\mathbf{x}^{(k)} \sim \pi(\mathbf{x} | \boldsymbol{\theta}, \mathbf{y})$ from the conditional posterior is obtained by minimising the following equation with respect to \mathbf{x} :

$$\mathbf{x}^{(k)} = \arg \min_{\mathbf{x}} \| \hat{\mathbf{A}}\mathbf{x} - (\hat{\mathbf{y}} + \mathbf{b}) \|_{L^2}^2, \quad \mathbf{b} \sim \mathcal{N}(\mathbf{0}, \mathbf{I}), \quad (2.24)$$

where we add a random perturbation \mathbf{b} . Similar to Section 2.5, this expression can be rewritten as

$$\left(\mathbf{A}_{\boldsymbol{\theta}}^T \boldsymbol{\Sigma}_{\boldsymbol{\theta}}^{-1} \mathbf{A}_{\boldsymbol{\theta}} + \mathbf{Q}_{\boldsymbol{\theta}} \right) \mathbf{x}^{(k)} = \mathbf{A}_{\boldsymbol{\theta}}^T \boldsymbol{\Sigma}_{\boldsymbol{\theta}}^{-1} \mathbf{y} + \mathbf{Q}_{\boldsymbol{\theta}} \boldsymbol{\mu} + \mathbf{v}_1 + \mathbf{v}_2, \quad (2.25)$$

with $\mathbf{v}_1 \sim \mathcal{N}(\mathbf{0}, \mathbf{A}_{\boldsymbol{\theta}}^T \boldsymbol{\Sigma}_{\boldsymbol{\theta}}^{-1} \mathbf{A}_{\boldsymbol{\theta}})$ and $\mathbf{v}_2 \sim \mathcal{N}(\mathbf{0}, \mathbf{Q}_{\boldsymbol{\theta}})$, representing independent Gaussian random variables [2, 28].

2.3 Numerical Approximation – Tensor-Train (TT)

Instead of relying on sampling-based methods to explore a target distribution $\pi(\mathbf{x})$, we can approximate this distribution on a d -dimensional grid using a TT approximation $\tilde{\pi}(\mathbf{x}) \approx \pi(\mathbf{x})$, where $\mathbf{x} \in \mathbb{R}^d$, with far fewer function evaluation compared to conventional sampling methods. In the following, we describe how to compute marginal distributions from a probability density approximated in TT format and how to generate samples using the inverse Rosenblatt transform (IRT), following the notation and procedure introduced by Cui et al. [3].

As in [3], we can define the parameter space as the product space $\mathcal{X} = \mathcal{X}_1 \times \mathcal{X}_2 \times \cdots \times \mathcal{X}_d$ with $x_k \in \mathcal{X}_k \subseteq \mathbb{R}$. The marginal density function for the k -th component is then given by

$$f_{X_k}(x_k) = \frac{1}{z} \int_{\mathcal{X}_1} \cdots \int_{\mathcal{X}_d} \lambda(\mathbf{x}) \pi(\mathbf{x}) \, dx_1 \cdots dx_{k-1} \, dx_{k+1} \cdots dx_d, \quad (2.26)$$

where we integrate over all dimensions except the k -th, and z is a normalisation constant. Here, we introduce a weight function $\lambda(x)$, which can be useful for quadrature rules [31], to which [3] refer to as a "product-form Lebesgue-measurable weighting function" and define it as

$$\lambda(\mathcal{X}) = \prod_{i=1}^d \lambda_i(\mathcal{X}_i), \quad \text{where } \lambda_i(\mathcal{X}_i) = \int_{\mathcal{X}_i} \lambda_i(x_i) \, dx_i. \quad (2.27)$$

In the TT format, the integral in Eq. 2.26 for the marginal probability can be computed at a low computational cost, as $\pi(\mathbf{x})$ is approximated by

$$\tilde{\pi}(\mathbf{x}) = \tilde{\pi}_1(x_1) \tilde{\pi}_2(x_2) \cdots \tilde{\pi}_d(x_d),$$

which is a sequence of matrix multiplications, with $\tilde{\pi}_k(x_k) \in \mathbb{R}^{r_{k-1} \times r_k}$ for a fixed point $\mathbf{x} = (x_1, \dots, x_d)$ on a predefined d -dimensional discrete univariate grid over the parameter space \mathcal{X} . We call $\tilde{\pi}_k \in \mathbb{R}^{r_{k-1} \times n \times r_k}$ a TT core with ranks r_{k-1} and r_k , where the outer ranks are $r_0 = r_d = 1$, representing each dimension on n grid points and connecting to neighbouring dimensions through its ranks. This enables us to approximate $\pi(\mathcal{X}) \approx \tilde{\pi}_1 \tilde{\pi}_2 \cdots \tilde{\pi}_d$ over a discrete parameter space \mathcal{X} using $2nr + (d-2)nr^2$ evaluation points for

fixed ranks $r = r_{k-1} = r_k$, as illustrated in Figure 2.2, instead of n^d function evaluation. Consequently, the marginal target distribution

$$\begin{aligned} f_{X_k}(x_k) \approx \frac{1}{z} & \left| \left(\int_{\mathcal{X}_1} \lambda_1(x_1) \tilde{\pi}_1(x_1) dx_1 \right) \cdots \left(\int_{\mathcal{X}_{k-1}} \lambda_{k-1}(x_{k-1}) \tilde{\pi}_{k-1}(x_{k-1}) dx_{k-1} \right) \right. \\ & \left. \lambda_k(x_k) \tilde{\pi}_k(x_k) \right. \\ & \left(\int_{\mathcal{X}_{k+1}} \lambda_{k+1}(x_{k+1}) \tilde{\pi}_{k+1}(x_{k+1}) dx_{k+1} \right) \cdots \left(\int_{\mathcal{X}_d} \lambda_d(x_d) \tilde{\pi}_d(x_d) dx_d \right) \Big| \end{aligned} \quad (2.28)$$

is computed by integrating over all TT cores except the k th core π_k , as in [32], and normalised by the constant z [3].

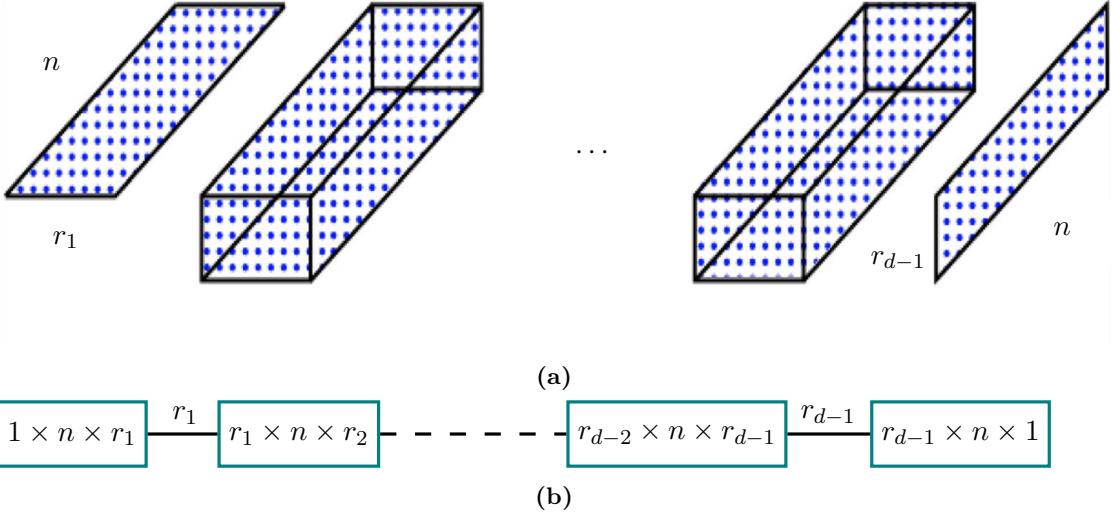


Figure 2.2: Here, we visualise the TT cores as a train of two- and three-dimensional matrices. Each core has a length n , corresponding to the number of grid points in each dimension, and the cores are connected through ranks r_k . More specifically, a core $\tilde{\pi}_k$ has dimensions $r_{k-1} \times n \times r_k$, with outer ranks $r_0 = r_d = 1$. Using the TT format enables us to represent a d -dimensional grid with only $2nr + (d-2)nr^2$ evaluation points instead of n^d grid points. Figure (a) is adapted from [33].

In practice, TT approximations may suffer from numerical instability, in particular because it is not advantageous yet to approximate the target function $\pi(\mathbf{x})$ in e.g. the logarithmic space. Hence, Cui et al. [3] approximate the square root of the probability density

$$\sqrt{\pi(\mathbf{x})} \approx \tilde{g}(\mathbf{x}) = \mathbf{G}_1(x_1), \dots, \mathbf{G}_k(x_k), \dots, \mathbf{G}_d(x_d) \quad [3, \text{Eq. 18}], \quad (2.29)$$

which ensures positivity. Here, each TT core is given by

$$G_k^{(\alpha_{k-1}, \alpha_k)}(x_k) = \sum_{i=1}^{n_k} \phi_k^{(i)}(x_k) \mathbf{A}_k[\alpha_{k-1}, i, \alpha_k], \quad k = 1, \dots, d, \quad [3, \text{Eq. 21}], \quad (2.30)$$

where $\mathbf{A}_k \in \mathbb{R}^{r_{k-1} \times n_k \times r_k}$ is the k -th coefficient tensor and $\{\phi_k^{(i)}(x_k)\}_{i=1}^{n_k}$ are the basis functions corresponding to the k -th coordinate. The approximated un-normalised

density is written as:

$$\pi(\mathbf{x}) \approx \xi + \tilde{g}(\mathbf{x})^2 \quad [3, \text{Eq. 19}], \quad (2.31)$$

where ξ is a positive constant added according to the ratio of the Lebesgue weighted L2-norm error and the Lebesgue weighting (see Eq. 2.27) such that

$$0 \leq \xi \leq \frac{1}{\lambda(\mathcal{X})} \|\tilde{g} - \sqrt{\pi}\|_{L_\lambda^2(\mathcal{X})}^2 \quad [3, \text{Eq. 35}]. \quad (2.32)$$

This leads to the normalised target function

$$f_X(\mathbf{x}) \approx \frac{1}{z} \left(\lambda(\mathbf{x}) \xi + \lambda(\mathbf{x}) \tilde{g}(\mathbf{x})^2 \right) \quad [3, \text{Eq. 19}], \quad (2.33)$$

with the normalisation constant $z = \int_{\mathcal{X}} f_X(\mathbf{x}) d\mathbf{x}$. Given the tensor train approximation of $\sqrt{\pi}$, the marginal function $f_{X_k}(x_k)$ can be expressed as

$$\begin{aligned} f_{X_k}(x_k) &\approx \frac{1}{z} \left(\xi \prod_{i=1}^{k-1} \lambda_i(\mathcal{X}_i) \prod_{i=k+1}^d \lambda_i(\mathcal{X}_i) \right. \\ &\quad + \left(\int_{\mathcal{X}_1} \lambda_1(x_1) \mathbf{G}_1^2(x_1) dx_1 \right) \cdots \left(\int_{\mathcal{X}_{k-1}} \lambda_{k-1}(x_{k-1}) \mathbf{G}_{k-1}^2(x_{k-1}) dx_{k-1} \right) \\ &\quad \lambda_k(x_k) \mathbf{G}_k^2(x_k) \\ &\quad \left. \left(\int_{\mathcal{X}_{k+1}} \lambda_{k+1}(x_{k+1}) \mathbf{G}_{k+1}^2(x_{k+1}) dx_{k+1} \right) \cdots \left(\int_{\mathcal{X}_d} \lambda_d(x_d) \mathbf{G}_d^2(x_d) dx_d \right) \right). \end{aligned} \quad (2.34)$$

2.3.1 Marginal Probability Distributions

We compute the marginal probability distributions by a procedure, to which Cui et al. [3] refer to as backwards marginalisation, see Prop. 2, and to which I add the forward marginalisation, see Prop. 1. This is similar to the left and right orthogonalisation of TT cores [34, 35]. The backwards marginalisation provides us with the coefficient matrices \mathbf{B}_k , while the forward marginalisation gives the coefficient matrices $\mathbf{B}_{\text{pre},k}$. These matrices enable the efficient evaluation of marginal functions since they integrate over the coordinates either left or right of the k -th dimension, as in [3]. In doing so, we define the mass matrix $\mathbf{M}_k \in \mathbb{R}^{n_k \times n_k}$ as

$$\mathbf{M}_k[i, j] = \int_{\mathcal{X}_k} \phi_k^{(i)}(x_k) \phi_k^{(j)}(x_k) \lambda(x_k) dx_k, \quad i, j = 1, \dots, n_k, \quad [3, \text{Eq. 22}], \quad (2.35)$$

where $\{\phi_k^{(i)}(x_k)\}_{i=1}^{n_k}$ denotes the set of basis functions for the k -th coordinate. The proposition used to compute \mathbf{B}_k , stated in Prop. 1, is adapted directly from [3].

Proposition 1 (Backwards Marginalisation as in [3]): Starting with the last coordinate $k = d$, we set $\mathbf{B}_d = \mathbf{A}_d$. The following procedure can be used to obtain the coefficient tensor $\mathbf{B}_k \in \mathbb{R}^{r_{k-1} \times n_k \times r_k}$, which we need for defining the marginal function $f_{X_k}(x_k)$ or to draw samples from $\tilde{\pi}(\mathbf{x})$ via the squared IRT scheme (see Alg. Box 3):

1. Use the Cholesky decomposition of the mass matrix, $\mathbf{L}_k \mathbf{L}_k^\top = \mathbf{M}_k \in \mathbb{R}^{n_k \times n_k}$, to construct a tensor $\mathbf{C}_k \in \mathbb{R}^{r_{k-1} \times n_k \times r_k}$:

$$\mathbf{C}_k[\alpha_{k-1}, \tau, l_k] = \sum_{i=1}^{n_k} \mathbf{B}_k[\alpha_{k-1}, i, l_k] \mathbf{L}_k[i, \tau] \quad [3, \text{Eq. (27)}]. \quad (2.36)$$

2. Unfold \mathbf{C}_k along the first coordinate and compute the thin QR decomposition, so that $\mathbf{C}_k^{(R)} \in \mathbb{R}^{r_{k-1} \times (n_k r_k)}$:

$$\mathbf{Q}_k \mathbf{R}_k = (\mathbf{C}_k^{(R)})^\top \quad [3, \text{Eq. 28}]. \quad (2.37)$$

3. Compute the new coefficient tensor:

$$\mathbf{B}_{k-1}[\alpha_{k-2}, i, l_{k-1}] = \sum_{\alpha_{k-1}=1}^{r_{k-1}} \mathbf{A}_{k-1}[\alpha_{k-2}, i, \alpha_{k-1}] \mathbf{R}_k[l_{k-1}, \alpha_{k-1}] \quad [3, \text{Eq. 29}]. \quad (2.38)$$

Proposition 2 (Forward Marginalisation): Starting with the first coordinate $k = 1$, we set $\mathbf{B}_{\text{pre},1} = \mathbf{A}_1$. The following procedure can be used to obtain $\mathbf{R}_{\text{pre},k-1} \in \mathbb{R}^{r_{k-1} \times r_{k-1}}$ for defining the marginal function $f_{X_k}(x_k)$:

1. Use the Cholesky decomposition of the mass matrix, $\mathbf{L}_k \mathbf{L}_k^\top = \mathbf{M}_k \in \mathbb{R}^{n_k \times n_k}$, to construct a tensor $\mathbf{C}_k \in \mathbb{R}^{r_{k-1} \times n_k \times r_k}$:

$$\mathbf{C}_{\text{pre},k}[\alpha_{k-1}, \tau, l_k] = \sum_{i=1}^{n_k} \mathbf{L}_k[i, \tau] \mathbf{B}_{\text{pre},k}[\alpha_{k-1}, i, l_k]. \quad (2.39)$$

2. Unfold $\mathbf{C}_{\text{pre},k}$ along the first coordinate and compute the thin QR decomposition, so that $\mathbf{C}_{\text{pre},k}^{(R)} \in \mathbb{R}^{(r_{k-1} n_k) \times r_k}$:

$$\mathbf{Q}_{\text{pre},k} \mathbf{R}_{\text{pre},k} = (\mathbf{C}_{\text{pre},k}^{(R)})^\top. \quad (2.40)$$

3. Compute the new coefficient tensor $\mathbf{B}_{\text{pre},k+1} \in \mathbb{R}^{r_{k-1} \times n_k \times r_k}$:

$$\mathbf{B}_{\text{pre},k+1}[l_{k+1}, i, \alpha_{k+1}] = \sum_{\alpha_k=1}^{r_k} \mathbf{R}_{\text{pre},k}[l_{k+1}, \alpha_k] \mathbf{A}_{k+1}[\alpha_k, i, \alpha_{k+1}]. \quad (2.41)$$

After computing the coefficient tensors $\mathbf{R}_{\text{pre},k-1}$ as in Prop. 2 and \mathbf{B}_k from Prop. 1,

the marginal PDF of k -th dimension can be expressed as

$$f_{X_k}(x_k) \approx \frac{1}{z} \left(\xi \prod_{i=1}^{k-1} \lambda_i(X_i) \prod_{i=k+1}^d \lambda_i(X_i) + \sum_{l_{k-1}=1}^{r_{k-1}} \sum_{l_k=1}^{r_k} \left(\sum_{i=1}^n \phi_k^{(i)}(x_k) \mathbf{D}_k[l_{k-1}, i, l_k] \right)^2 \right) \lambda_k(x_k), \quad (2.42)$$

where $\mathbf{D}_k \in \mathbb{R}^{r_{k-1} \times n \times r_k}$ and given as

$$\mathbf{D}_k[l_{k-1}, i, l_k] = \sum_{\alpha_{k-1}=1}^{r_{k-1}} \mathbf{R}_{\text{pre}, k-1}[l_{k-1}, \alpha_{k-1}] \mathbf{B}_k[\alpha_{k-1}, i, l_k], \quad (2.43)$$

with $\mathbf{R}_{\text{pre}, k-1} \in \mathbb{R}^{r_{k-1} \times r_{k-1}}$ and $\mathbf{B}_k \in \mathbb{R}^{r_{k-1} \times n \times r_k}$.

For the first dimension, $f_{X_1}(x_1)$ can be expressed as

$$f_{X_1}(x_1) \approx \frac{1}{z} \left(\xi \prod_{i=2}^d \lambda_i(\mathcal{X}_i) + \sum_{l_1=1}^{r_1} \left(\sum_{i=1}^n \phi_1^{(i)}(x_1) \mathbf{D}_1[i, l_1] \right)^2 \right) \lambda_1(x_1) \quad [3, \text{Eq. 30}], \quad (2.44)$$

where $\mathbf{D}_1[i, l_1] = \mathbf{B}_1[\alpha_0, i, l_1]$ and $\alpha_0 = 1$, and similarly in the last dimension

$$f_{X_d}(x_d) \approx \frac{1}{z} \left(\xi \prod_{i=1}^{d-1} \lambda_i(\mathcal{X}_i) + \sum_{l_{d-1}=1}^{r_{d-1}} \left(\sum_{i=1}^n \phi_d^{(i)}(x_d) \mathbf{D}_d[l_{d-1}, i] \right)^2 \right) \lambda_d(x_d), \quad (2.45)$$

where $\mathbf{D}_d[l_{d-1}, i] = \mathbf{B}_{\text{pre}, d}[l_{d-1}, i, \alpha_{d+1}]$ and $\alpha_{d+1} = 1$. Note that we calculate the normalisation numerically within the process of computing the marginals so that $\sum f_{X_k}(x_k) = 1$.

2.3.2 Sampling from a TT Approximation

Instead of evaluating marginal functions for quadrature, we can draw samples from the approximated function via the inverse Rosenblatt transform, as in [32], to preserve the correlation structure. Since we approximate the square root of the target function, Cui et. al. [3] call that the squared inverse Rosenblatt transform (SIRT).

Algorithm 2: Squared Inverse Rosenblatt Transform (SIRT)

```

1: Input: seeds  $\{\mathbf{u}^{(1)}, \dots, \mathbf{u}^{(N)}\} \sim \mathcal{U}(0, 1)^d$  and  $\mathbf{B}_1, \dots, \mathbf{B}_d$  from Prop. 1
2: for  $s = 1, \dots, N$  do
3:   for  $k = 1, \dots, d$  do
4:     compute normalised PDF  $f_{X_k|X_{<k}}(x_k|x_{k-1}^{(s)}, \dots, x_1^{(s)})$ , Eq. 2.47
5:     compute cumulative distribution function  $F_{X_k|X_{<k}}(x_k)$ , Eq. 2.46,
6:     project sample  $x_k^{(s)} = F_{X_k|X_{<k}}^{-1}(u_k^{(s)})$ 
7:     interpolate  $\mathbf{G}_k(x_k^{(s)})$ , Eq. 2.48
8:     update  $\mathbf{G}_{\leq k}(x_{\leq k}^{(s)}) = \mathbf{G}_{<k}(x_{<k}^{(s)}) \mathbf{G}_k(x_k^{(s)})$ 
9:   end for
10:  end for
11: Output: samples  $\{\mathbf{x}^{(1)}, \dots, \mathbf{x}^{(N)}\}$ , where each  $\mathbf{x}^{(s)} \in \mathbb{R}^d$  for  $s = 1, \dots, N$ 
```

Given the Backward marginals $\mathbf{B}_1, \dots, \mathbf{B}_d$ as in Prop. 1, we draw N uniformly distributed seeds $\{\mathbf{u}^{(1)}, \dots, \mathbf{u}^{(N)}\} \sim \mathcal{U}(0, 1)^d$, where each $\mathbf{u}^{(s)}$ is d-dimensional for $s = 1, \dots, N$. Then we calculate the first marginal $f_{X_1}(x_1)$ as in Eq. 2.44 and normalise with $z = \int_{\mathcal{X}_1} f_{X_1}(x_1) dx_1$. Next, we compute the cumulative distribution function (CDF) $F_{X_1}(x_k) = \int_{-\infty}^{x_k} f_{X_1}(\hat{x}_1) d\hat{x}_1$, which for the general case is given as:

$$F_{X_k|X_{<k}}(x_k) \approx \int_{-\infty}^{x_k} f_{X_k|X_{<k}}(\hat{x}_k|x_{k-1}, \dots, x_1) d\hat{x}_k [3, \text{Eq. 17}]; \quad (2.46)$$

and project the seed $u_k^{(s)}$ on the parameter space, so that $x_k^{(s)} = F_{X_k|X_{<k}}^{-1}(u_k^{(s)})$. The “conditional marginal” is given as:

$$\begin{aligned} f_{X_k|X_{<k}}(x_k|x_{k-1}^{(s)}, \dots, x_1^{(s)}) &\approx \frac{1}{z} \left(\xi \prod_{i=k+1}^d \lambda_i(X_i) + \right. \\ &\quad \left. \sum_{l_k=1}^{r_k} \left(\sum_{i=1}^n \phi_k^{(i)}(x_k^{(s)}) \left(\sum_{\alpha_{k-1}=1}^{r_{k-1}} \mathbf{G}_{<k}^{(\alpha_{k-1})}(x_{<k}^{(s)}) \mathbf{B}_k[\alpha_{k-1}, i, l_k] \right) \right)^2 \right) \lambda_k(x_k) [3, \text{Eq. 31}], \end{aligned} \quad (2.47)$$

where we marginalise over the dimensions $k+1, \dots, d$ via \mathbf{B}_k and condition on the previous $k-1$ samples through the product $\mathbf{G}_k(x_k^{(s)}) \in \mathbb{R}^{1 \times r_{k-1}}$ to preserve the correlation structure. Between grid points i and $i+1$, we approximate using a piecewise polynomial interpolation

$$\mathbf{G}_k(x_k^{(s)}) \approx \frac{x_k^{(s)} - x_k^{(i)}}{x_k^{(i+1)} - x_k^{(i)}} \mathbf{G}_k(x_k^{(i+1)}) + \frac{x_k^{(i+1)} - x_k^{(s)}}{x_k^{(i+1)} - x_k^{(i)}} \mathbf{G}_k(x_k^{(i)}), \quad (2.48)$$

for $x_k^{(i)} \leq x_k^{(s)} \leq x_k^{(i+1)}$ as in [32] for the next “conditional marginal”.

We repeat the procedure for each $u_k^{(s)} \in \mathbf{u}^{(s)}$ to produce samples $\mathbf{x}^{(s)} \sim f_X(\mathbf{x})$, as summarised in Alg. Box 3.

Metropolis-Hastings - correction step

Since the samples by the SIRT scheme are generated from an approximation, it is sensible to correct those using a Metropolis-Hastings (MH) importance step. In doing so, we compute the acceptance probability $\alpha = \min(w^{(s+1)}/w^{(s)}, 1)$, where

$$w(\mathbf{x}) = \frac{\pi(\mathbf{x})}{f_X(\mathbf{x})} = \frac{\pi(\mathbf{x})}{\xi + \tilde{g}(\mathbf{x})^2} \quad (2.49)$$

is the importance ratio. Note that since we calculate the ratio $w^{(s+1)}/w^{(s)}$ the normalising constants cancel. In practise we calculate the importance ratio in the log space, where $\log f_X(\mathbf{x}) = \log f_{X_1}(x_1) + \log f_{X_2|X_1}(x_2|x_1) + \dots + \log f_{X_k|X_{<k}}(x_k|x_{k-1}, \dots, x_1)$ is given as in Eq. 2.47, see [32]. We refer to this as the SIRT-MH scheme, which in theory provides the corrected chain $\{\mathbf{x}_{\text{MH}}^{(1)}, \dots, \mathbf{x}_{\text{MH}}^{(N)}\} \sim \pi(\mathbf{x})$.

Algorithm 3: MH correction step

```

1: Input: samples  $\{\mathbf{x}^{(1)}, \dots, \mathbf{x}^{(N+1)}\}$ , where each  $\mathbf{x}^{(s)} \in \mathbb{R}^d$  for  $s = 1, \dots, N + 1$ 
2: for  $s = 1, \dots, N$  do
3:   compute MH ratio  $\frac{w^{(s+1)}}{w^{(s)}} = \frac{\pi(\mathbf{x}^{(s+1)})}{\pi(\mathbf{x}^{(s)})} \frac{f_X(\mathbf{x}^{(s)})}{f_X(\mathbf{x}^{(s+1)})}$ 
4:   compute acceptance probability  $\alpha = \min(w^{(s+1)}/w^{(s)}, 1)$ 
5:   Draw  $u \sim \mathcal{U}(0, 1)$ 
6:   if  $\alpha \geq u$  then
7:     Accept and set  $\mathbf{x}_{\text{MH}}^{(s+1)} = \mathbf{x}^{(s+1)}$ 
8:   else
9:     Reject and keep  $\mathbf{x}_{\text{MH}}^{(s+1)} = \mathbf{x}^{(s)}$ 
10:  end if
11: end for
12: Output: corrected sample chain  $\{\mathbf{x}_{\text{MH}}^{(1)}, \dots, \mathbf{x}_{\text{MH}}^{(N)}\}$ , where each  $\mathbf{x}_{\text{MH}}^{(s)} \in \mathbb{R}^d$  for  $s = 1, \dots, N$ 
```

2.3.3 On the Error of a TT Approximation

A straightforward way to asses the error from the TT approximation is to calculate the relative root mean squared (RMS) error

$$\left(\frac{\int_{\mathcal{X}} (\pi(\mathbf{x}) - (\xi + \tilde{g}(\mathbf{x})^2))^2 \lambda(\mathbf{x}) d\mathbf{x}}{\int_{\mathcal{X}} \pi(\mathbf{x})^2 \lambda(\mathbf{x}) d\mathbf{x}} \right)^{1/2} = \frac{\|\pi(\mathbf{x}) - (\xi + \tilde{g}(\mathbf{x})^2)\|_{L_\lambda^2(\mathcal{X})}}{\|\pi(\mathbf{x})\|_{L_\lambda^2(\mathcal{X})}}. \quad (2.50)$$

We can approximate this integral as

$$\left(\frac{1}{N} \sum_{i=1}^N \left(\pi(\mathbf{x}^{(i)}) - (\xi + \tilde{g}(\mathbf{x}^{(i)})^2) \right)^2 \lambda(\mathbf{x}^{(i)}) \right)^{1/2} \approx \left(\int_{\mathcal{X}} (\pi(\mathbf{x}) - (\xi + \tilde{g}(\mathbf{x})^2))^2 \lambda(\mathbf{x}) d\mathbf{x} \right)^{1/2} \quad (2.51)$$

and similarly $\int_{\mathcal{X}} \pi(\mathbf{x})^2 \lambda(\mathbf{x}) d\mathbf{x}$.

Absolute Error Bound

Another way to assess the error between two distributions is to calculate the Wasserstein distance. The Kantorovich-Rubinstein duality, as in [36, 37], says that the 1-Wasserstein distance is equal to the supremum of differences in expectations over all 1-Lipschitz functions h between two probability distributions. So the 1-Wasserstein distance provides an upper error bound.

We define the 1-Wasserstein distance as

$$W_1(\pi, \tilde{\pi}) = \inf_{\nu \in \Pi(\pi, \tilde{\pi})} \int_{\mathcal{X} \times \mathcal{X}} c(\mathbf{x}, \tilde{\mathbf{x}}) \nu(\mathbf{x}, \tilde{\mathbf{x}}) d\mathbf{x} d\tilde{\mathbf{x}}, \quad (2.52)$$

where ν couples \mathbf{x} and $\tilde{\mathbf{x}}$ so that the integral over the distance $c(\mathbf{x}, \tilde{\mathbf{x}})$ weighted by the probability measures π and $\tilde{\pi}$ is the greatest lower bound of all integrals with respect

to a ν in the set of all couplings $\Pi(\pi, \tilde{\pi})$. Often ν is called a transport plan, where $c(\mathbf{x}, \tilde{\mathbf{x}})$ is the (ground) cost function, and $\nu(\mathbf{x}, \tilde{\mathbf{x}})$ is related to the mass which has to be transported and the 1-Wasserstein distance is the earth mover distance. On the other hand (Kantorovich-Rubinstein duality), we can describe the 1-Wasserstein distance

$$W_1(\pi, \tilde{\pi}) = \sup_{h(\mathbf{x}); |h(\mathbf{x}) - h(\tilde{\mathbf{x}})| \leq c(\mathbf{x}, \tilde{\mathbf{x}})} \left\{ \int_{\mathcal{X}} h(\mathbf{x}) d\pi(\mathbf{x}) - \int_{\mathcal{X}} h(\tilde{\mathbf{x}}) d\tilde{\pi}(\tilde{\mathbf{x}}) \right\} \quad (2.53)$$

$$= \sup_{h(\mathbf{x}); |h(\mathbf{x}) - h(\tilde{\mathbf{x}})| \leq c(\mathbf{x}, \tilde{\mathbf{x}})} \left\{ \mathbb{E}_{\mathbf{x} \sim \pi}[h(\mathbf{x})] - \mathbb{E}_{\tilde{\mathbf{x}} \sim \tilde{\pi}}[h(\tilde{\mathbf{x}})] \right\}. \quad (2.54)$$

as the lowest upper bound of differences in expectations over all 1-Lipschitz function $h(\mathbf{x}) \in \mathbb{R}$ in between the two distributions π and $\tilde{\pi}$, with the distance measure $c(\mathbf{x}, \tilde{\mathbf{x}}) = \|\mathbf{x} - \tilde{\mathbf{x}}\|_{L^2}$ for $\mathcal{X} \in \mathbb{R}^d$.

For two sample sets $\{\mathbf{x}^{(1)}, \dots, \mathbf{x}^{(N)}\} \sim \pi$ and $\{\tilde{\mathbf{x}}^{(1)}, \dots, \tilde{\mathbf{x}}^{(M)}\} \sim \tilde{\pi}$ the calculation of the Wasserstein distance becomes an optimisation problem, that is to find the best coupling of samples weighted by their distribution value according to an appropriate distance measure [38]. More specifically, the 1-Wasserstein distance becomes

$$W_1(\pi, \tilde{\pi}) = \min_{\nu \in \Pi(\pi, \tilde{\pi})} \sum_{j=1}^M \sum_{i=1}^N \nu_{ij} \|\mathbf{x}^{(i)} - \tilde{\mathbf{x}}^{(j)}\|_{L^2}, \quad (2.55)$$

where the transport plan $\nu \in \mathbb{R}_{\geq 0}^{N \times M}$ defines the coupling $\nu_{ij} \in \nu$ as $\nu_{ij} := \pi(\mathbf{x}^{(i)}) \tilde{\pi}(\tilde{\mathbf{x}}^{(j)})$ similar to [38, Eq. 3.166]. Additionally we require that $\sum_{i=1}^N \pi(\mathbf{x}^{(i)}) = \sum_{j=1}^M \tilde{\pi}(\tilde{\mathbf{x}}^{(j)}) = 1$. This gives us an upper bound of the absolute error in between the expected value of any 1-Lipschitz function h , e.g the upper bound of absolute differences in means related to the probability measures π and $\tilde{\pi}$.

2.4 Affine Map

The forward map, which we introduce in Ch. 3, poses a weakly non-linear forward problem, which we could tackle by treating the problem as a linear problem and then iteratively updating the non-linear part after each parameter sample. Instead, we approximate the non-linear model using an affine map $\mathbf{M} : \mathbf{A}_L \mathbf{x} \rightarrow \mathbf{A}_{NL}(\mathbf{x})$, which maps from the linear model \mathbf{A}_L to the non-linear model \mathbf{A}_{NL} , so that we define linear $\mathbf{A} = \mathbf{M} \mathbf{A}_L$, which approximates the non-linear model. Here, we give a brief introduction to affine maps.

An affine map is any linear map between two vector spaces or affine spaces, where an affine space does not need to preserve a zero origin (see [39, Def. 2.3.1]). In other words, an affine map does not need to map to the origin of the associated vector space. An affine map is a linear map on vector spaces, including a translation, or, in the words of my supervisor, C. F., a Taylor series of first order. For more information on affine spaces and maps, we refer to the books [39, 40].

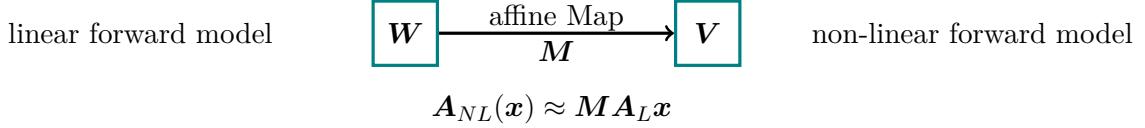


Figure 2.3: This Figure shows the schematic representation of how the affine map \mathbf{M} approximates the non-linear forward model. Here, \mathbf{V} contains values produced by the linear forward model, and \mathbf{W} contains the corresponding values from the non-linear forward model. Both \mathbf{V} and \mathbf{W} are affine subspaces over the same field. The affine map \mathbf{M} projects elements from the linear forward model space \mathbf{V} onto their counterparts in the non-linear forward model space \mathbf{W} . More specifically, the non-linear noise-free data vector $\mathbf{A}_{NL}\mathbf{x}$ is approximated by the affine map and the linear forward model so that $\mathbf{A}_{NL}(\mathbf{x}) \approx \mathbf{M}\mathbf{A}_L\mathbf{x}$.

2.5 Regularisation

As mentioned in the introduction, the currently most used method to analyse data in atmospheric physics is regularisation-based. Since we want to show that Bayesian methods provide more information than regularisation at a similar computational cost, we choose a regularisation approach closest to the linear-Gaussian Bayesian framework in Sec. 4.2.

The Tikhonov regularisation approach provides one solution \mathbf{x}_λ that minimises both the data misfit norm

$$\|\mathbf{y} - \mathbf{Ax}\|_{L^2} \quad (2.56)$$

and a regularisation semi-norm

$$\lambda \|\mathbf{T}\mathbf{x}\|_{L^2}, \quad (2.57)$$

for a given regularisation parameter $\lambda > 0$ as described in [2], with a linear forward model matrix \mathbf{A} , the data \mathbf{y} , a regularisation operator \mathbf{T} . The regularisation parameter weights the semi-norm and penalises \mathbf{x} according to that. If λ is large, then the effect of the data on the solution \mathbf{x}_λ is small or negligible and dominated by the regulariser. If λ is small, the solution \mathbf{x}_λ will be dominated by the noisy data, resulting in an overfitted \mathbf{x}_λ . We refer to [41] and [11] for a more comprehensive analysis on the effects of the regularisation parameter to the solution, e.g. due to small singular values of the forward model.

For a fixed λ , the regularised solution

$$\mathbf{x}_\lambda = \arg \min_{\mathbf{x}} \|\mathbf{y} - \mathbf{Ax}\|_{L^2}^2 + \lambda \|\mathbf{T}\mathbf{x}\|_{L^2}^2 \quad (2.58)$$

is obtained by taking the derivative with respect to \mathbf{x} :

$$\nabla_{\mathbf{x}} \left\{ (\mathbf{y} - \mathbf{Ax})^T (\mathbf{y} - \mathbf{Ax}) + \lambda \mathbf{x}^T \mathbf{T}^T \mathbf{T} \mathbf{x} \right\} = 0 \quad (2.59)$$

$$\iff \nabla_{\mathbf{x}} \left\{ \mathbf{y}^T \mathbf{y} + \mathbf{x}^T \mathbf{A}^T \mathbf{Ax} - 2 \mathbf{y}^T \mathbf{Ax} + \lambda \mathbf{x}^T \mathbf{T}^T \mathbf{T} \mathbf{x} \right\} = 0 \quad (2.60)$$

$$\iff 2 \mathbf{A}^T \mathbf{Ax} - 2 \mathbf{A}^T \mathbf{y} + 2\lambda \mathbf{T}^T \mathbf{T} \mathbf{x} = 0, \quad (2.61)$$

also known as the “regularised normal equation” $\mathbf{A}^T \mathbf{y} = \mathbf{A}^T \mathbf{A} \mathbf{x} + \lambda \mathbf{T}^T \mathbf{T} \mathbf{x}$ [42]. Solving this equation yields the regularised solution

$$\mathbf{x}_\lambda = (\mathbf{A}^T \mathbf{A} + \lambda \mathbf{L})^{-1} \mathbf{A}^T \mathbf{y}, \quad (2.62)$$

where we define $\mathbf{L} := \mathbf{T}^T \mathbf{T}$, which typically represents a discrete matrix approximation of a differential operator choice [11]. For example

$$\mathbf{T} = \frac{1}{h} \begin{bmatrix} -1 & 1 & & & \\ 0 & -1 & 1 & & \\ & \ddots & \ddots & \ddots & \\ & & 0 & -1 & 1 \\ & & & 0 & -1 \end{bmatrix}, \quad (2.63)$$

is the first order derivative with equal spacing h as in [11] then

$$\mathbf{L} = \frac{1}{h^2} \begin{bmatrix} 1 & -1 & & & \\ -1 & 2 & -1 & & \\ & \ddots & \ddots & \ddots & \\ & & -1 & 2 & -1 \\ & & & -1 & 2 \end{bmatrix}, \quad (2.64)$$

is the second order derivative with Neumann boundary conditions, see [43].

In practice, \mathbf{x}_λ is computed for a range of λ -values and evaluated based on the trade-off between the data misfit and the regularisation semi-norm. The optimal value of λ corresponds to the point of maximum curvature of the so-called L-curve [44], where the data misfit norm versus the regularisation semi-norm is plotted, see Fig. 4.13.

Additionally one can think about regularisation as a Lagrange multiplier $\mathcal{L}(\mathbf{x}, \lambda) := \lambda \sqrt{\mathbf{x}^T \mathbf{L} \mathbf{x}} + \|\mathbf{y} - \mathbf{A} \mathbf{x}\|_{L^2}$, which minimises $\sqrt{\mathbf{x}^T \mathbf{L} \mathbf{x}}$ with respect to constant λ and $\|\mathbf{y} - \mathbf{A} \mathbf{x}\|_{L^2}$ (see [2, fn. 6] and [45, Fig. 2.13]). So every solution \mathbf{x}_λ is an extremum (the most regularised solution for a given data misfit norm) and almost every sample of the posterior, which represents a feasible solution given the data, has a higher $\sqrt{\mathbf{x}^T \mathbf{L} \mathbf{x}}$ value and lays above the L-Curve.

3

The Forward Model

In this chapter, we present the forward model to which we apply all our methodology. We follow the Michelson Interferometer for Passive Atmospheric Sounding (MIPAS) handbook [1] and simulate data according to a cloud-free atmosphere in local thermodynamic equilibrium, assuming a measurement instrument with infinite spectral resolution and no pointing errors. We do not include any other instrument-specific details, such as sensor area or antenna response, as they are not available to us.

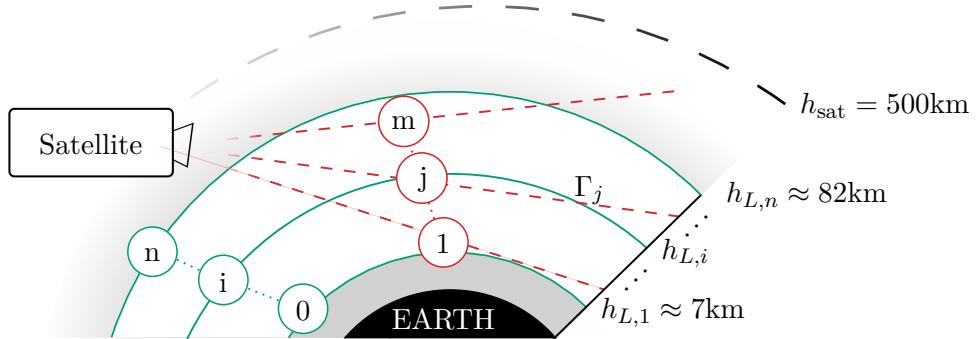


Figure 3.1: Schematic of measurement and analysis geometry, not to scale. The stationary satellite, at a constant height h_{sat} above Earth, takes m measurements along its line-of-sight defined by the line Γ_j . Each measurement has a limb height ℓ_j , $j = 1, 2, \dots, m$ defined as the closest distance of Γ_j to the Earth's surface. Between $h_{L,0} \approx 7\text{km}$ and $h_{L,n} \approx 82\text{km}$, the atmosphere is discretised into n layers as illustrated by the solid green lines.

A satellite at a constant height h_{sat} points through the atmosphere (limb-sounding) and measures thermal radiation of gas molecules along its straight line of sight Γ_j , see Figure 3.1. One measurement of the thermal radiation of one specific molecule, in our case ozone, denoted by the ozone volume mixing ratio (VMR) $x(r)$ at distance r from

the satellite, at the wave number ν , is given by the path integral

$$y_j = \int_{\Gamma_j} B(\nu, T) k(\nu, T) \frac{p(T)}{k_B T(r)} x(r) \tau(r) dr + \eta_j \quad (3.1)$$

$$\tau(r) = \exp \left\{ - \int_{r_{\text{obs}}}^r k(\nu, T) \frac{p(T)}{k_B T(r')} x(r') dr' \right\}, \quad (3.2)$$

which is the radiative transfer equation (RTE) [1]. For more information on the processes within the atmosphere for ozone, we refer to [46]. We define a tangent height h_{ℓ_j} and Γ_j for each $j = 1, 2, \dots, m$, so that the data vector $\mathbf{y} \in \mathbb{R}^m$ including some noise η_j . Within the atmosphere, the number density $p(T)/(k_B T(r))$ of molecules is dependent on the pressure $p(T)$, the temperature $T(r)$, and the Boltzmann constant k_B . The factor $\tau(r) \leq 1$ accounts for re-absorption of the radiation along the line-of-sight, which makes the RTE non-linear. The absorption constant

$$k(\nu, T) = L(\nu, T_{\text{ref}}) \frac{Q(T_{\text{ref}})}{Q(T)} \frac{\exp \{-c_2 E''/T\}}{\exp \{-c_2 E''/T_{\text{ref}}\}} \frac{1 - \exp \{-c_2 \nu/T\}}{1 - \exp \{-c_2 \nu/T_{\text{ref}}\}} \quad (3.3)$$

is dependent on the line intensity $L(\nu, T_{\text{ref}})$ at reference temperature $T_{\text{ref}} = 296K$, the lower-state energy E'' in cm^{-1} of the targeted transition and the second radiation constant $c_2 := hc/k_B \approx 1.44\text{cmK}$ as in the HITRAN database [47], with Planck's constant h and speed of light c . Since we assume that the measurement device has negligible frequency window we neglect line broadening around ν for the calculations of $L(\nu, T_{\text{ref}})$, which would normally be modelled as a convolution of the normalised Lorentz profile (collisional/pressure broadening) and the normalised Doppler (thermal broadening) profile [1]. Additionally, we target one specific molecule and calculate $k(\nu, T)$ accordingly, which usually would involve summing the individual absorption constants for multiple radiating molecules weighted by their respective VMR [1]. The total internal partition function is given as:

$$Q(T) = g' \exp \left\{ -\frac{c_2 E'}{T} \right\} + g'' \exp \left\{ -\frac{c_2 E''}{T} \right\}, \quad (3.4)$$

with the statistical weight g'' for the lower and g' for the upper energy state (also called the degeneracy factors) accounting for the molecule's non-rotational and rotational energy states (see [48]), and the upper state energy $E' = E'' + \nu$. Under the assumption of local thermodynamic equilibrium (LTE), the black body radiation acts as a source function

$$B(\nu, T) = \frac{2hc^2\nu^3}{\exp \left\{ \frac{c_2 \nu}{T} \right\} - 1}. \quad (3.5)$$

For fundamentals on the RTE, we recommend [49, Chapter 1], and for a more comprehensive model, we refer to [50].

To calculate the integrals in Eq. 3.1 and Eq. 3.2 numerically we discretise the atmosphere in n layers, where the i^{th} layer is defined by two spheres of radii $h_{L,i-1} < h_{L,i}$,

for $i = 1, \dots, n$, with $h_{L,0}$ and $h_{L,n}$. Then the ozone VMR $\mathbf{x} = \{x_1, x_2, \dots, x_n\} \in \mathbb{R}^n$, pressure $\mathbf{p} = \{p_1, p_2, \dots, p_n\} \in \mathbb{R}^n$ and temperature $\mathbf{T} = \{T_1, T_2, \dots, T_n\} \in \mathbb{R}^n$, as well as all other height dependent parameters, are discretised profiles with constant values in between the heights $h_{L,i-1}$ and $h_{L,i}$. Above $h_{L,n}$ and below $h_{L,0}$, the ozone VMR is set to zero, so no signal can be obtained. We evaluate the integral in Eq. (3.1) for one noise free measurement $\mathbf{A}_j(\mathbf{p}, \mathbf{T}, \mathbf{x})$, using the trapezoidal rule. Here each entry of $\mathbf{A}(\mathbf{p}, \mathbf{T}, \mathbf{x}) \in \mathbb{R}^m$ includes multiple evaluations of the integral in Eq. 3.2 to calculate the absorption $\tau(r)$. Note that, since we aim to provide estimates over pressure \mathbf{p} and temperature \mathbf{T} , we explicitly include them as parameters in our forward model. The data vector

$$\mathbf{y} = \mathbf{A}(\mathbf{x}) + \boldsymbol{\eta} \quad (3.6)$$

includes an additive noise vector $\boldsymbol{\eta} \in \mathbb{R}^m$, where we define the non-linear forward model as $\mathbf{A}(\mathbf{x}) := \mathbf{A}(\mathbf{x}, \mathbf{p}, \mathbf{T}) \in \mathbb{R}^m$ for brevity. Similarly, we define \mathbf{A}_L , which denotes the linear forward model matrix and neglects absorption (e.g. set $\tau = 1$ in Eq. (3.2)) and enables matrix-vector multiplication $\mathbf{A}_L \mathbf{x}$ to compute noise-free linear data.

Further, we classify the inverse problem as weakly non-linear (see Sec. 4.3) as neglecting the absorption changes the measurements only slightly (about 1%).

3.1 Singular Value Decomposition of the Forward Model

Before simulating some data, we provide a quick and intuitive way of assessing if the data collection is effective, how much information is passed through the forward model, and how the signal-to-noise ratio (SNR) and the measurement strategy affects that information. One way of doing this is via a singular value decomposition (SVD) of the forward model matrix

$$\mathbf{A}_L = \sum_{i=1}^r \mathbf{u}_i \sigma_i \mathbf{v}_i^T = \mathbf{U} \boldsymbol{\Sigma} \mathbf{V}^T \quad (3.7)$$

where $r = \min\{m, n\}$ for a forward model matrix $\mathbf{A}_L \in \mathbb{R}^{m \times n}$. Consider noise-free measurements $\mathbf{A}_L \mathbf{x}$ for a satellite at a fixed height of $h_{\text{sat}} = 500\text{km}$ above sea level. Our main objective is to measure ozone \mathbf{x} , so our forward model \mathbf{A}_L includes temperature and pressure, the latter is dominant, see Fig. 4.17, decreases exponentially in height and hence does affect the information passed through the model. If the pressure is high, the signal is large. If the pressure is low, the signal is low and the data is noise-dominated.

The SVD gives us information on how information is picked up from the parameter space by the forward model, described through the right singular values \mathbf{v}_i . The singular values σ_i , ordered in size from the largest σ_1 to the smallest σ_r , weight that information from the right singular values to the left singular values \mathbf{u}_i , which project onto the data space. For a large singular value, we can say that the forward model is informative about

structures in the corresponding right singular vector and vice versa. Further, for very small singular values $\sigma_i \ll \sigma_1/\text{SNR}$ below the RMS noise level or the noise standard deviation (STD), we can introduce an effective rank $r_{\text{eff}} \leq r$. Then, the parameter space spanned by $\{\mathbf{v}_{r_{\text{eff}}+1}, \dots, \mathbf{v}_r\}$ is noise dominated in the corresponding data space, see Figure 3.6. This is based on the rough assumption that if we define the SNR as

$$\text{SNR} := \frac{\max(y)}{\text{STD noise}} = \frac{\text{peak signal}}{\text{RMS noise}} [51], \quad (3.8)$$

then the maximum singular value $\max(y) \approx \sigma_1$ and the information transmitted through the forward model corresponds roughly to the singular values $\sigma_i \gtrsim \max(y)/\text{SNR}$. See [11] for a more comprehensive analysis.

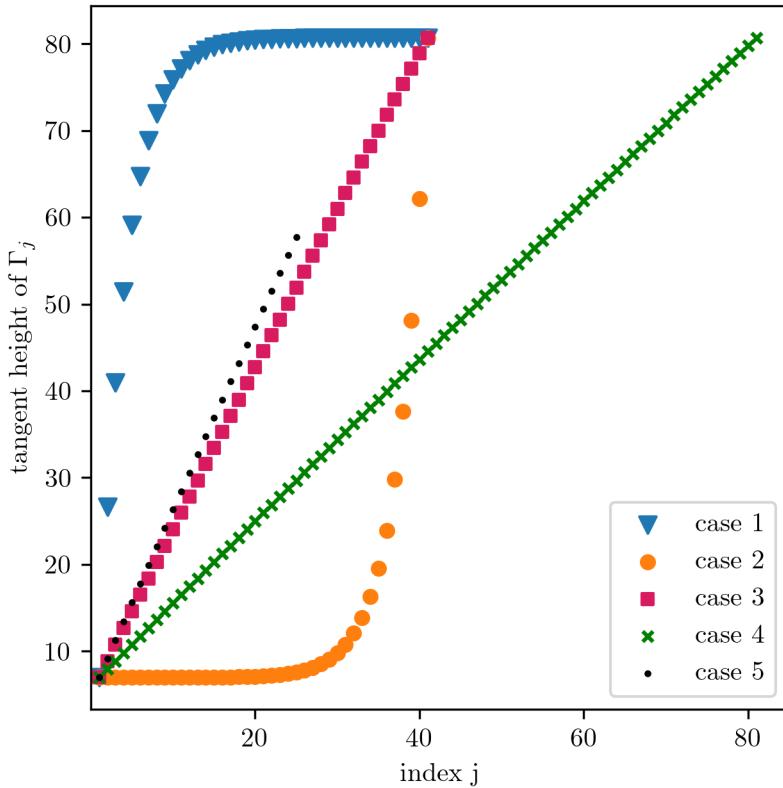


Figure 3.2: Tangent heights for five different sequences of measurements.

Next, we plot the singular values for 5 different measurement scenarios, where we either measure at equidistance spaced tangent heights or collect more data from high signal regions at low altitudes or from low signal regions at high altitudes. We assess the number of singular values above and below a certain SNR visually, to determine which of the tested cases is most effective. We start with case 3 in Fig. 3.2 where measurements are spaced according to a pointing accuracy of 150arc sec, given to us by the team of the University of New South Wales Canberra Space [52]. The pointing accuracy determines

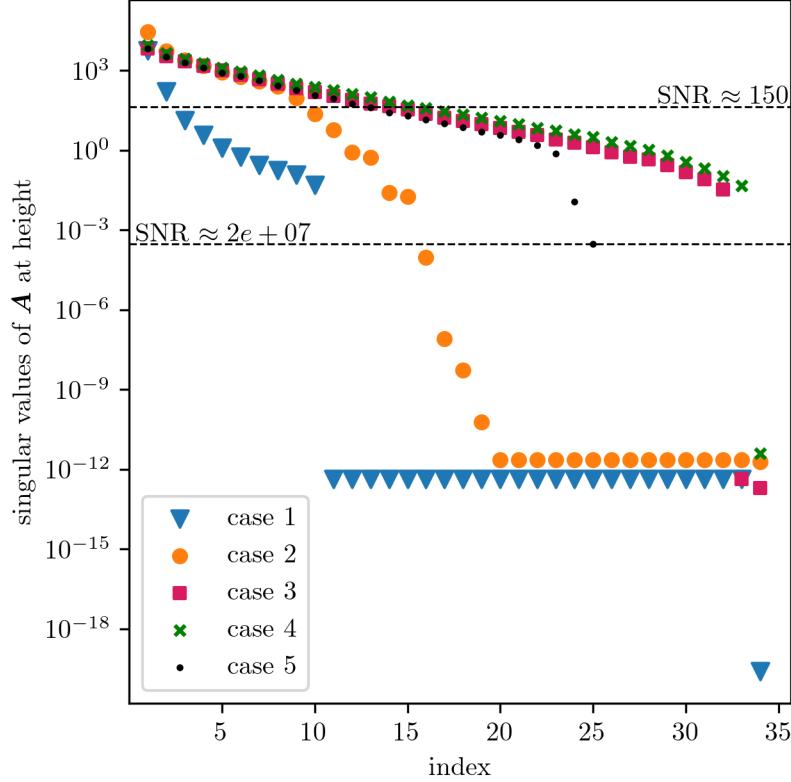


Figure 3.3: Singular values of the forward model matrix for different sequences of measurements. The corresponding tangent heights of the test cases are plotted in Fig. 3.2. The dotted vertical line marks where the singular values are dominated by noise according to a specific SNR.

how well the satellite can point in a certain direction and, hence, roughly the spacing in between two measurements. The corresponding singular values are plotted in Fig. 3.3, of which the first 25 decrease linearly in log-space and about 10-15 singular values lie above the SNR. In comparison, if we measure a lot of times in regions where the data is noise-dominated (high altitude), case 1, we do obtain more information since the singular values decrease rapidly. Measuring lots of times at low altitudes, where the data is informative, and less at higher altitudes, case 2, does not seem optimal either, as we observe one larger singular value, but the other singular values decrease faster compared to case 3. Now consider case 4, where we double the number of measurements compared to case 3, we do get slightly larger singular values, but not so significantly that it would be worth the engineering effort required to achieve that. The measurements with equidistance-spaced tangent height seem to be most effective. By exploratory analysis, we find that we can tolerate a slightly larger distance between tangent heights (pointing accuracy of 175arc sec) than required by [52], see case 5. In that case, we also stop measuring when the signal is too noisy and decrease the number of measurements taken

without losing crucial information. We note that if one wanted to obtain all information provided by the forward model, we would need a signal-to-noise ratio of roughly 10^7 .

In principle, we show that it does not depend on how one measures; one can not get more information by measuring more in regions where the information content is low or high. This contradicts the current measurement setup on the AURA MLS [12], which reports high noise in lower atmospheric regions, due to thermal radiation from the earth, and measures more in those regions.

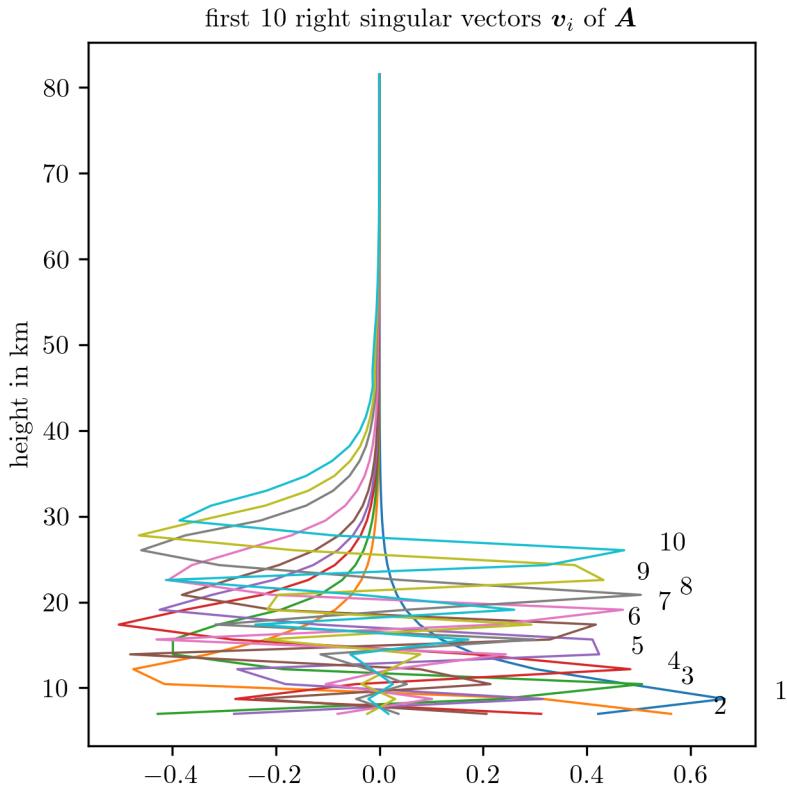


Figure 3.4: First 10 right singular vectors of the forward model matrix for measurements case 5 in Fig. 3.2. These singular vectors correspond to high singular values of the forward model in Fig. 3.3.

Consequently, we proceed with case 5 and plot the right singular vectors of the forward model versus height in the atmosphere to see where our model is most informative, or which structures of the parameter space are picked up by the model. The first 10 right singular vectors, in Fig. 3.4, corresponding to the 10 largest singular values, pick up structures in lower atmospheric regions. So we can assume that, given some data, we will be able to provide good reconstructions of the parameter in lower altitudes. Next, we plot the right singular vectors corresponding to the singular values σ_j for $j = 11, \dots, 20$ in Fig. 3.5, where the noise starts to dominate the data. These singular values lie in regions around the SNR, see Fig. 3.3, and pick up values in the middle of the atmosphere.

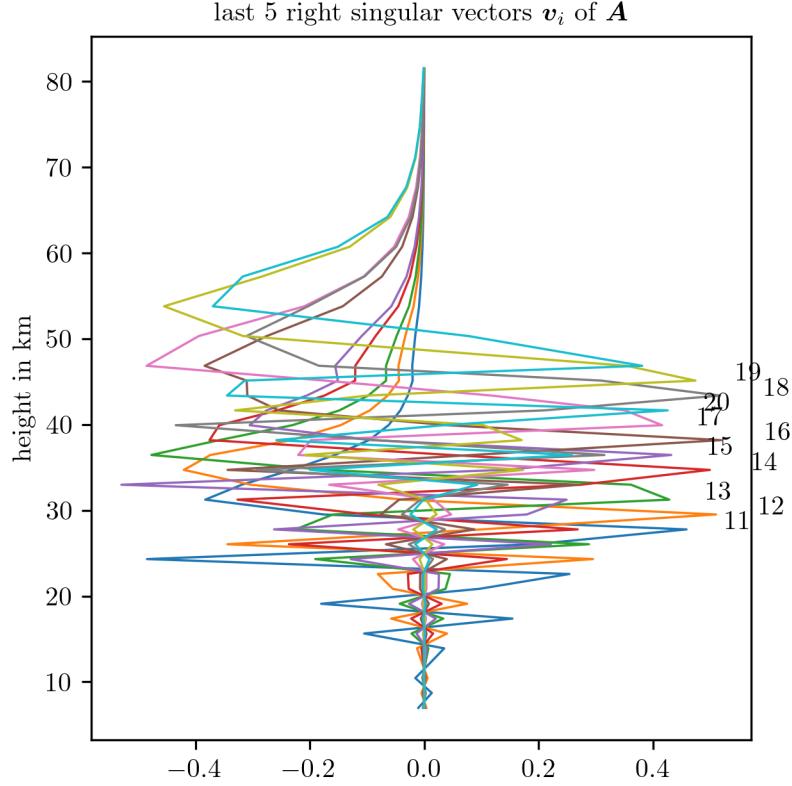


Figure 3.5: Right singular vectors with index $i = 11, \dots, 19$ of the forward model matrix for measurements case 5 in Fig. 3.2. These singular vectors correspond to singular values in Fig. 3.3 where the noise level is similar to the data.

Consequently, we expect a higher uncertainty of reconstructed parameter values in the middle atmospheric regions. The singular vectors corresponding to the last 10 singular values pick up structures in higher altitudes, but since the singular values are very small, we will not be able to retrieve those structures. More specifically, the retrieved parameter values at higher altitudes will be mostly determined by the prior or, in the case of regularisation, by the regulariser [11].

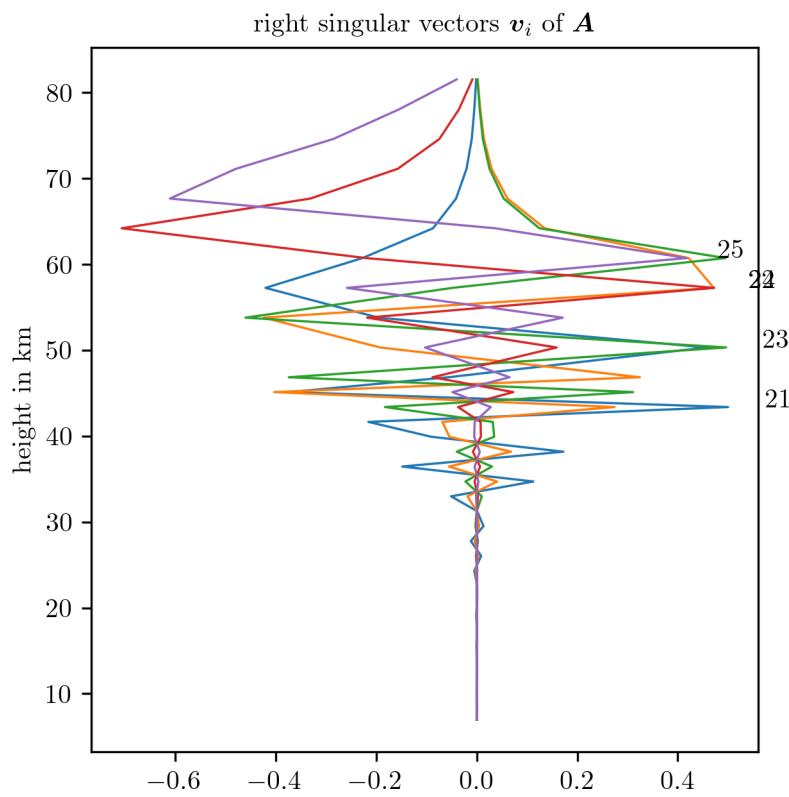


Figure 3.6: Last 10 right singular vectors of the forward model matrix for measurements case 5 in Fig. 3.2. These singular vectors correspond to small singular values of the forward model in Fig. 3.3 where the data is noise-dominated.

4

Results and Conclusions

In this chapter, we use the forward model to generate some data given an underlying ground truth and then guide the reader through the process of setting up a Bayesian framework and ultimately obtaining the posterior distributions of hyper-parameters and parameters of interest, such as ozone concentration or pressure and temperature profiles. We are using a directed acyclic graph (DAG) to visualise hierarchical and correlation structures of a Bayesian model and established a choice of prior distributions within our Bayesian model and formulated the posterior distributions. Based on the linear forward model \mathbf{A}_L , we characterise the marginal posterior for ozone and compare that to the TT approximation. Then we calculate the mean and the covariance matrix of the full posterior for ozone, which we use to find an affine map to approximate the non-linear forward model $\mathbf{A}(\mathbf{x}) \approx \mathbf{M}\mathbf{A}_L\mathbf{x}$ (see Sec. 4.3). In Sec. 4.4, we repeat the MTC scheme to provide a posterior distribution of ozone based on the approximate forward model and compare it to a regularisation approach and a ground truth. Additionally, in sec 4.5, we extend the hierarchical Bayesian model and MTC scheme to jointly provide conditional posterior distributions of ozone, temperature and pressure. Here we elaborate on some aspects of prior modelling and on our findings when using a TT to approximate the higher-dimensional marginal posterior. All programming and analysis is done in Python, and the reported computation times correspond to a MacBook Pro from 2019 with a 2.4 GHz quad-core Intel Core i5 processor.

4.1 Simulate Data based on a Ground Truth

We take a ground truth ozone profile at distinct pressure values generated from some data [13] of the MLS on the Aura satellite within the Antarctic region and with a peak in high altitude, see Fig. 4.6, which seems to be a typical nighttime profile [46].

We recursively calculate the geometric height with the hydrostatic equilibrium equation

$$\frac{dp}{p} = \frac{-gM}{R^*T} dh, \quad (4.1)$$

with the acceleration due to gravity

$$g = g_0 \left(\frac{r_0}{r_0 + h} \right), \quad (4.2)$$

where the polar radius pf the earth is $r_0 \approx 6356$ km, the gravitation at sea level is $g_0 \approx 9.81$ m/s², $R^* = 8.31432 \times 10^{-3}$ Nm/kmol/K and the mean molecular weight of the air is $M = 28.97$ kg/kmol [53]. This holds up to a geometric height of 86km, where we ignore a 0.04% non-linear change in M from 80km to 86km in geometric altitude.

Following [53] we form the temperature function

$$T(h) = \begin{cases} T_0 & , h = 0 \\ T_0 + a_0 h & , 0 \leq h < h_{T,1} \\ T_0 + a_0 h_{T,1} & , h_{T,1} \leq h < h_{T,2} \\ T_0 + a_0 h_{T,1} + a_1(h_{T,2} - h_{T,1}) + a_2(h - h_{T,2}) & , h_{T,2} \leq h < h_{T,3} \\ T_0 + a_0 h_{T,1} + a_1(h_{T,2} - h_{T,1}) \\ \quad + a_2(h_{T,3} - h_{T,2}) + a_3(h - h_{T,3}) & , h_{T,3} \leq h < h_{T,4} \\ T_0 + a_0 h_{T,1} + a_1(h_{T,2} - h_{T,1}) \\ \quad + a_2(h_{T,3} - h_{T,2}) + a_3(h_{T,4} - h_{T,3}) + a_4(h - h_{T,4}) & , h_{T,4} \leq h < h_{T,5} \\ T_0 + a_0 h_{T,1} + a_1(h_{T,2} - h_{T,1}) \\ \quad + a_2(h_{T,3} - h_{T,2}) + a_3(h_{T,4} - h_{T,3}) + a_4(h_{T,5} - h_{T,4}) \\ \quad + a_5(h - h_{T,5}) & , h_{T,5} \leq h < h_{T,6} \\ T_0 + a_0 h_{T,1} + a_1(h_{T,2} - h_{T,1}) \\ \quad + a_2(h_{T,3} - h_{T,2}) + a_3(h_{T,4} - h_{T,3}) + a_4(h_{T,5} - h_{T,4}) \\ \quad + a_5(h_{T,6} - h_{T,5}) + a_6(h - h_{T,6}) & , h_{T,6} \leq h \lesssim 86 \end{cases} \quad (4.3)$$

with gradient and height values provided by [53] (see Tab. 4.1) which acts as the ground truth temperature profile (see Fig. 4.15).

Then we can compute a data vector $\mathbf{y} = \mathbf{A}(\mathbf{x}) + \boldsymbol{\eta}$, with $m = 30$ measurements determined by the satellite pointing accuracy of 175arc sec (see case 5 in Fig. 3.2), according to the RTE as in Eq. 3.1 and 3.2 using the trapezoidal integration rule. We

subscript i	geometric height $h_{T,i}$ in km	gradient a_i
0	0	-6.5
1	11	0
2	20.1	1
3	32.2	2.8
4	47.4	0
5	51.4	-2.8
6	71.8	-2

Table 4.1: Definition of height depending temperature gradients.

assume an atmosphere between $h_{L,1} = 7\text{km}$ and $h_{L,n} = 83.3\text{km}$ with $n = 45$ equidistant layers (see Fig. 3.1). The height value $h_{L,i}$ for each layer $i = 1, \dots, n$ is defined by the pressure values from [13] and the hydrostatic equilibrium equation, see Eq. 4.1. We target Ozone at a frequency of 235.71 GHz, which lies within the region where the MLS observes ozone [14, 54]. The corresponding wave number is $\nu = 7.86\text{cm}^{-1}$. We calculate the absorption constant $k(\nu, T)$ as in Eq. 3.2, following the HITRAN database [47], which provides the line intensity $L(\nu, T_{\text{ref}})$ for the isotopologue $^{16}\text{O}_3$ with the AFGL Code 666. Lastly, we add independent and identically-distributed Gaussian noise $\boldsymbol{\eta} \sim \mathcal{N}(0, \gamma^{-1}\mathbf{I})$ so that the SNR = 150 (see Eq. 3.8) similar to [15], where a signal with a maximal spectral intensity of around 100K and a noise range of 0.4 to 1.6K is reported. We note that the methods used in this thesis will work with different SNRs or other frequencies.

When we plot the data in Fig. 4.1, which is noise-dominated in higher altitudes and hence not sensitive to structures in the higher atmospheric regions. Now, given the data, we like to determine the posterior distributions over ozone \mathbf{x} , pressure \mathbf{p} and temperature \mathbf{T} .

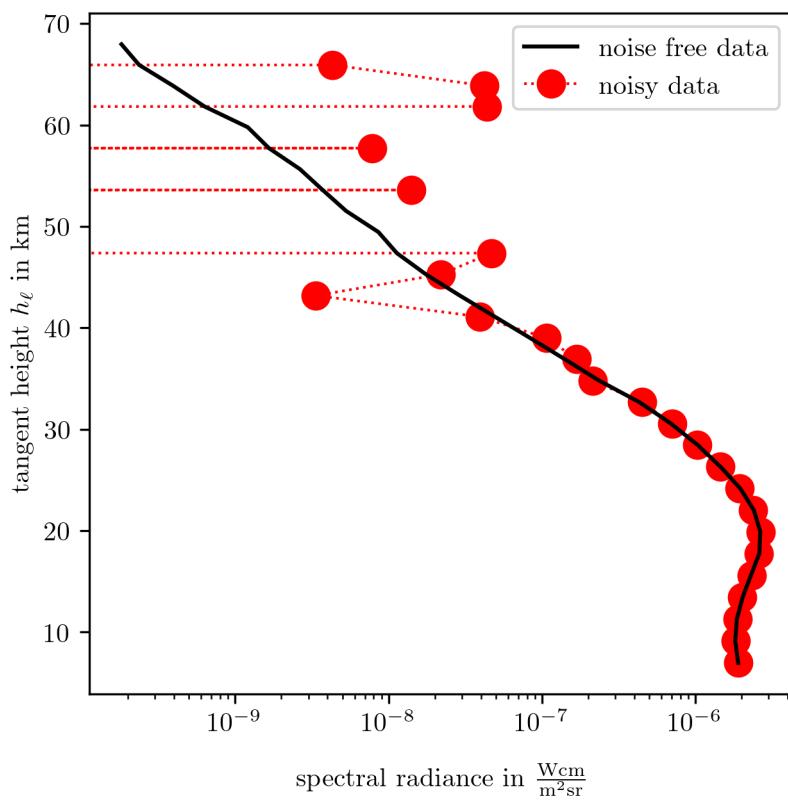


Figure 4.1: Logarithmic plot of data points at different tangent height. Note that negative values are not plotted, and noise is dominating at higher altitudes.

4.2 Hierarchical Bayesian Framework for Ozone

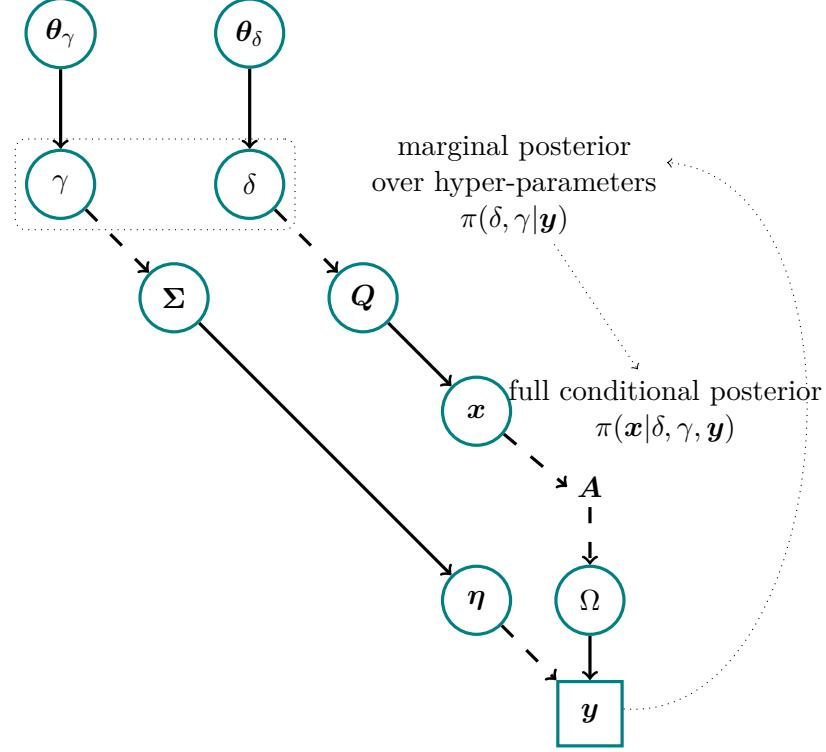


Figure 4.2: DAG for visualisation of hierarchical modelling and measuring process of ozone, including the MTC scheme. The hyper-parameter γ deterministically (dotted line) sets the noise covariance $\Sigma = \gamma^{-1} \mathbf{I}$ and hence the random (solid line) noise vector $\eta \sim \mathcal{N}(0, \gamma^{-1} \mathbf{I})$. The hyper-parameter δ determines (dotted line) the prior precision matrix $Q = \delta \mathbf{L}$ for the normally distributed (solid line) prior $x|\delta \sim \mathcal{N}(0, \delta \mathbf{L})$, where \mathbf{L} is a graph Laplacian, see Eq. 4.5. The hyper-prior distributions (solid line) $\pi(\delta, \gamma)$ are defined by θ_γ and θ_δ . Through a linear forward model \mathbf{A} , we generate a space of all measurable noise-free data \mathbf{Ax} from which we randomly observe a data set \mathbf{y} including some added noise η . Within the MTC scheme, we evaluate the marginal posterior over the hyper-parameters $\pi(\gamma, \delta | \mathbf{y})$ first and then the full conditional posterior $\pi(x | \delta, \gamma, \mathbf{y})$. This breaks the correlation structure of x and δ and γ , and allows us to evaluate the marginal posterior independent of x .

In this section, we set up the hierarchically-ordered linear-Gaussian Bayesian framework to determine the ozone posterior distribution, conditioned on ground truth temperature and pressure. Where, for now, we define the forward model matrix $\mathbf{A} := \mathbf{A}_L$ and define the distributions of that Bayesian model, similarly to a regularisation approach, as:

$$\mathbf{y} | \mathbf{x}, \gamma, \delta \sim \mathcal{N}(\mathbf{Ax}, \gamma^{-1} \mathbf{I}) \quad (4.4a)$$

$$\mathbf{x} | \delta \sim \mathcal{N}(\mathbf{0}, (\delta \mathbf{L})^{-1}) \quad (4.4b)$$

$$\delta \sim \Gamma(\alpha_\delta, \beta_\delta) \quad (4.4c)$$

$$\gamma \sim \Gamma(\alpha_\gamma, \beta_\gamma). \quad (4.4d)$$

Assuming Gaussian noise $\eta \sim \mathcal{N}(0, \gamma^{-1} \mathbf{I})$, the likelihood function is a normal distributions with mean \mathbf{Ax} and covariance matrix $\gamma^{-1} \mathbf{I}$. We define the normal prior-distribution

$\pi(\mathbf{x}|\delta)$, with zero mean and precision matrix $\delta\mathbf{L}$, where δ is a smoothness hyper-parameter and \mathbf{L} is the second order discrete derivate operator (see Eq. 4.5). Here the hyper-prior distributions $\pi(\delta)$ and $\pi(\gamma)$ are gamma distributions with shape α and rate β .

We can visualise this hierarchical structure and the correlations between different hyper-parameters and parameters through a DAG. The hyper-parameter γ sets the noise covariance deterministically (dotted line), but is itself statistically (solid line) defined by the hyper-prior distribution $\pi(\gamma)$. This is a gamma distribution, where θ_γ determines the shape and rate of $\pi(\gamma)$. Similarly θ_δ defines $\pi(\delta)$, where δ accounts for smoothness of the ozone profile and sets the prior precision $\mathbf{Q}(\delta)$. Then \mathbf{Ax} determines the space of all measurable noise-free data sets Ω , through the linear forward model, from which we observe a data set \mathbf{y} including some noise η . From that data, we then “reverse the arrows” to determine the posterior distribution $\pi(\mathbf{x}|\mathbf{y})$ over the parameter \mathbf{x} . Usually, due to underlying correlation structures, evaluating this posterior poses a significant challenge, here the MTC scheme provides the marginal posterior $\pi(\delta, \gamma|\mathbf{y})$ first and then the full conditional posterior $\pi(\mathbf{x}|\delta, \gamma, \mathbf{y})$.

4.2.1 Prior Modelling

To complete the Bayesian framework, we have to define prior distributions over the hyper-parameters and parameters. Ideally, we define the prior distributions as uninformative as possible, and include functional dependencies and physical properties.

By choosing a normally distributed prior $\pi(\mathbf{x}|\delta)$ with zero mean and no other restrictions, it is clear that our model does not take into account that ozone values can not be negative. As already mentioned, we set the precision matrix of that prior distribution to

$$\delta\mathbf{L} = \delta \begin{bmatrix} 2 & -1 & & & \\ -1 & 2 & -1 & & \\ & \ddots & \ddots & \ddots & \\ & & -1 & 2 & -1 \\ & & & -1 & 2 \end{bmatrix} \quad (4.5)$$

which is a 1-dimensional Graph Laplacian as in [2, 43] with Dirichlet boundary condition. This matrix will also act as the regulariser later in Sec. 2.5. We reduce the dimension of \mathbf{x} from 45 to 34 by discarding every second ozone VMR over a height of $\approx 47\text{km}$. Doing that, while not changing \mathbf{L} , we effectively induce a larger correlation between points at higher altitude. We plot the corresponding prior ozone profiles according to $\mathbf{x} \sim \mathcal{N}(0, (\delta\mathbf{L})^{-1})$ in Fig. B.1.

For δ and γ we pick relatively uninformative gamma distributions so that $\gamma \sim \mathcal{T}(\boldsymbol{\theta}_\gamma) \propto \gamma^{\alpha_\gamma-1} \exp(-\beta_\gamma \gamma)$ and $\delta \sim \mathcal{T}(\boldsymbol{\theta}_\delta)$, where $\boldsymbol{\theta}_\gamma = \{\alpha_\gamma, \beta_\gamma\} = \{\alpha_\delta, \beta_\delta\} = \boldsymbol{\theta}_\delta = (1, 10^{-35})$ (see

Fig. 4.10) similar to [2]. Those gamma distributions have another advantage when using the MWG algorithm to sample from the marginal posterior distribution $\pi(\delta, \gamma|\mathbf{y})$, where then $\pi(\gamma|\lambda, \mathbf{y}) \sim \mathcal{T}(\cdot)$ is a gamma distribution with $\lambda = \delta/\gamma$, and easy to sample from.

4.2.2 Posterior Distribution – linear Model

As explained in Sec. 2.1.1 we factorise the posterior

$$\pi(\mathbf{x}, \delta, \gamma|\mathbf{y}) \propto \pi(\mathbf{y}|\mathbf{x}, \delta, \gamma)\pi(\mathbf{x}, \delta, \gamma) \quad (4.6)$$

into

$$\pi(\mathbf{x}, \delta, \gamma|\mathbf{y}) = \pi(\mathbf{x}|\delta, \gamma, \mathbf{y})\pi(\delta, \gamma|\mathbf{y}) \quad (4.7)$$

the marginal posterior $\pi(\delta, \gamma|\mathbf{y})$ and full conditional posterior $\pi(\mathbf{x}|\delta, \gamma, \mathbf{y})$ (see Eq. 2.7). As discussed in Sec. 2.1.1, for the linear-Gaussian case, \mathbf{x} cancels in the marginal posterior over the hyper-parameters. Following the MTC scheme, we characterise the marginal posterior first and *then* the conditional posterior.

Marginal Posterior

Consequently, for the hierarchical model specified in Eq. 4.4, the marginal posterior distribution over the hyper-parameters is given by

$$\pi(\lambda, \gamma|\mathbf{y}) \propto \lambda^{n/2+\alpha_\delta-1} \gamma^{m/2+\alpha_\delta+\alpha_\gamma-1} \exp\left\{-\frac{1}{2}g(\lambda) - \frac{\gamma}{2}f(\lambda) - \beta_\delta \lambda \gamma - \beta_\gamma \gamma\right\}, \quad (4.8)$$

with the introduced regularisation parameter $\lambda = \delta/\gamma$, and

$$f(\lambda) = \mathbf{y}^T \mathbf{y} - (\mathbf{A}^T \mathbf{y})^T (\mathbf{A}^T \mathbf{A} + \lambda \mathbf{L})^{-1} (\mathbf{A}^T \mathbf{y}), \quad (4.9a)$$

$$\text{and } g(\lambda) = \log \det(\mathbf{A}^T \mathbf{A} + \lambda \mathbf{L}). \quad (4.9b)$$

Note that, when changing variables from $\delta = \lambda\gamma$ to λ the hyper-prior distribution changes to $\pi(\lambda) \propto \lambda^{\alpha_\delta-1} \gamma^{\alpha_\delta} \exp(-\beta_\delta \lambda \gamma)$, due to $d\delta/d\lambda = \gamma$. For each evaluation of the marginal posterior most of the computational effort lies in the calculation of $f(\lambda)$ and $g(\lambda)$, which we evaluate using the Cholesky decomposition $\mathbf{A}^T \mathbf{A} + \lambda \mathbf{L} = \mathbf{C}_\lambda \mathbf{C}_\lambda^T$. More specifically we solve for $(\mathbf{A}^T \mathbf{A} + \lambda \mathbf{L})^{-1} (\mathbf{A}^T \mathbf{y})$ via `scipy.linalg.cho_solve` and calculate the $g(\lambda) = 2 \sum \log \text{diag}(\mathbf{C}_\lambda)$. In Fig. 4.3 we see that $f(\lambda)$ and $g(\lambda)$ are well behaved within the region of interest and approximate $f(\lambda) \approx \tilde{f}(\lambda)$ with a Taylor series around the mode λ_0 of $\pi(\lambda, \gamma|\mathbf{y})$. We also note that $\tilde{g}(\lambda) \approx g(\lambda)$ behaves linearly around λ_0 in the log-space. The approximations are implicitly given by

$$f^{(r)}(\lambda_0) = (-1)^{r+1} (\mathbf{A}^T \mathbf{y})^T (\mathbf{B}_0^{-1} \mathbf{L})^r \mathbf{B}_0^{-1} \mathbf{A}_L^T \mathbf{y} \quad (4.10)$$

$$\text{and } \log \tilde{g}(\lambda) = \log g(\lambda_0) + (\log \lambda - \log \lambda_0) \frac{\log g(\lambda_{\max}) - \log g(\lambda_0)}{\log \lambda_{\max} - \log \lambda_0} \quad (4.11)$$

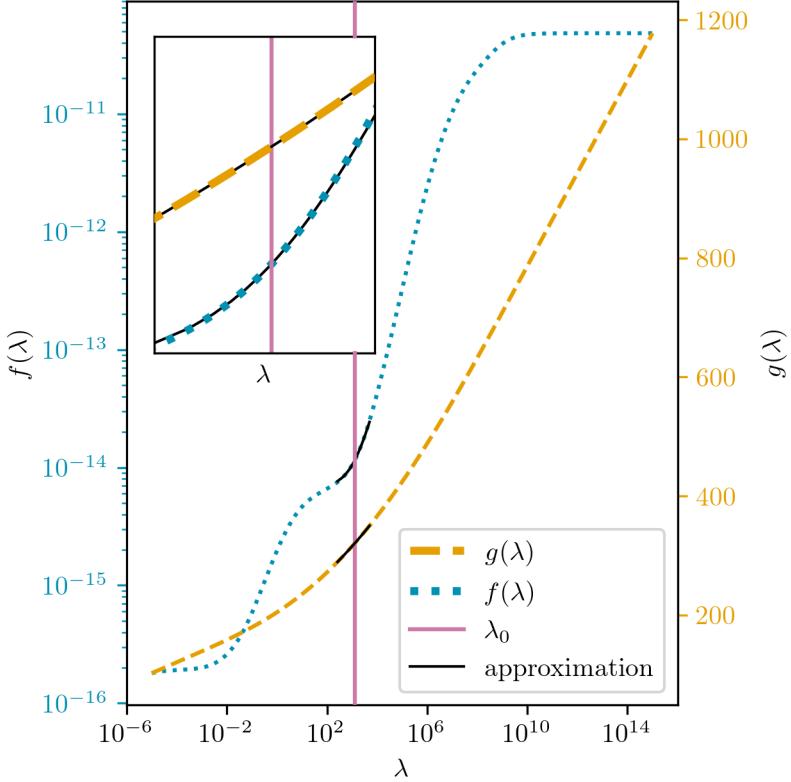


Figure 4.3: Functions $f(\lambda)$ and $g(\lambda)$ from the marginal posterior in Eq. 4.8 for a wide range of $\lambda = \delta/\gamma$. We plot the approximations (see Eq. 4.12a and Eq. 4.12b) in black around the mode of the marginal posterior (vertical line) for the sampling range of λ within the MWG algorithm.

with $\mathbf{B}_0 = \mathbf{A}^T \mathbf{A} + \lambda_0 \mathbf{L}$. We plot the approximations

$$\tilde{f}(\lambda) = \sum_{r=0}^1 f^{(r)}(\lambda_0)(\lambda - \lambda_0)^r, \quad (4.12a)$$

$$\text{and } \tilde{g}(\lambda) = \exp \log \tilde{g}(\lambda), \quad (4.12b)$$

in Fig. 4.3 and elaborate on the approximation errors in the section below. Note that usually a Taylor series includes a factor $(r!)^{-1}$, in this case it cancels in $f^{(r)}(\lambda_0)$ (see [2]).

Error due to Approximation of f and g To assess the approximation error, we lay a 100-point grid over the sampling region in each dimension and compare the approximations of $f(\lambda)$, $g(\lambda)$ and $\pi(\lambda, \gamma | \mathbf{y})$ with their true function values.

Compared to a 2nd, 3rd or 4th order Taylor approximation, we find that the 1st order Taylor approximation of $f(\lambda)$ gives the smallest relative RMS error of $\approx 25\%$ in the sampling region of λ . Additionally, we find that a linear approximation of $g(\lambda)$ is sufficient with relative RMS $< 0.1\%$.

These errors then propagate into the marginal posterior $\pi(\lambda, \gamma | \mathbf{y})$, so that the relative RMS error is $\approx 7\%$ over the whole grid. When sampling, we evaluate the acceptance

ratio in the log-space, so we report a relative RMS error $< 1\%$ for $\log \pi(\lambda|\gamma, \mathbf{y})$, with a maximum relative pointwise error of $< 2\%$. We consider this good enough.

Sample from Marginal Posterior Using these approximations, we employ a Metropolis within Gibbs (MWG) sampler (see Alg. Box 1) to characterise $\pi(\lambda, \gamma|\mathbf{y})$ (see Sec. 2.2.1).

We approximate $f(\lambda)$ and $g(\lambda)$ around the mode (λ_0, γ_0) of $\pi(\lambda, \gamma|\mathbf{y})$ provided by the `scipy.optimize.fmin` function, with a limit of 25 function evaluations. Then we approximate $f(\lambda)$ with a 1st order Taylor series and $g(\lambda)$ with a linear approximation in the log-space, where for the approximation in $g(\lambda)$ we set λ_{\max} to the maximum value of λ on the TT grid (see next section).

We initialised the MWG algorithm at the mode $(\lambda^{(0)}, \gamma^{(0)}) = (\lambda_0, \gamma_0)$ and take $N = 10000$ plus $N_{\text{burn-in}} = 100$ steps in $\approx 0.3\text{s}$. The standard deviation of the normal proposal distribution is empirically set to $\sigma_\lambda = 0.8\lambda_0$ so that the acceptance rate is ≈ 0.5 as suggested in [55]. More specifically, we implement a Metropolis random walk on

$$\pi(\lambda|\gamma, \mathbf{y}) \propto \lambda^{n/2+\alpha_\delta-1} \exp\left\{-\frac{1}{2}g(\lambda) - \frac{\gamma}{2}f(\lambda) - \beta_\delta\gamma\lambda\right\}. \quad (4.13)$$

In doing so, we accept $\lambda^{k+1} = \lambda'$ or reject $\lambda^{k+1} = \lambda^k$ a proposal $\lambda' \sim \mathcal{N}(0, \sigma_\lambda)$ according to the acceptance ratio in log space

$$\log\left\{\frac{\pi(\lambda'|\gamma^{(k)}, \mathbf{y})}{\pi(\lambda^{(k)}|\gamma^{(k)}, \mathbf{y})}\right\} = \log\{\pi(\lambda'|\gamma^{(k)}, \mathbf{y})\} - \log\{\pi(\lambda^{(k)}|\gamma^{(k)}, \mathbf{y})\} \quad (4.14)$$

$$= \frac{n}{2}(\log\{\lambda'\} - \log\{\lambda^{(k)}\}) + \frac{1}{2}\Delta g + \frac{\gamma^{(k)}}{2}\Delta f + \beta_\delta\gamma^{(k)}\Delta\lambda, \quad (4.15)$$

where $\Delta\lambda = \lambda' - \lambda^{(k)}$ and $\Delta f \approx \tilde{f}(\lambda') - \tilde{f}(\lambda^{(k)}) = f^{(1)}(\lambda_0)(\Delta\lambda' - \Delta\lambda^{(k)})$, with $\Delta\lambda' = \lambda' - \lambda_0$ and $\Delta\lambda^{(k)} = \lambda^{(k)} - \lambda_0$. Similarly we approximate $\Delta g \approx \tilde{g}(\lambda') - \tilde{g}(\lambda^{(k)})$. Lastly, we do a Gibbs step on

$$\gamma^{(k+1)}|\lambda^{(k+1)}, \mathbf{y} \sim \Gamma\left(\frac{m}{2} + \alpha_\delta + \alpha_\gamma, \frac{1}{2}f(\lambda^{(k+1)}) + \beta_\gamma + \beta_\delta\lambda^{(k+1)}\right) \quad (4.16)$$

to generate marginal posterior samples $(\lambda, \gamma)^{(1)}, \dots, (\lambda, \gamma)^{(N)} \sim \pi(\lambda, \gamma|\mathbf{y})$, which we plot in Fig. 4.5 as well as the trace of the MWG to show ergodicity. We calculate the IACT with the Python implementation of [20], provided by [21], which gives us $\tau_{\text{int},\gamma} \approx 2.2 \pm 0.1$ and $\tau_{\text{int},\lambda} = 5.2 \pm 0.5$, see Fig. 4.4 and Fig. B.2.

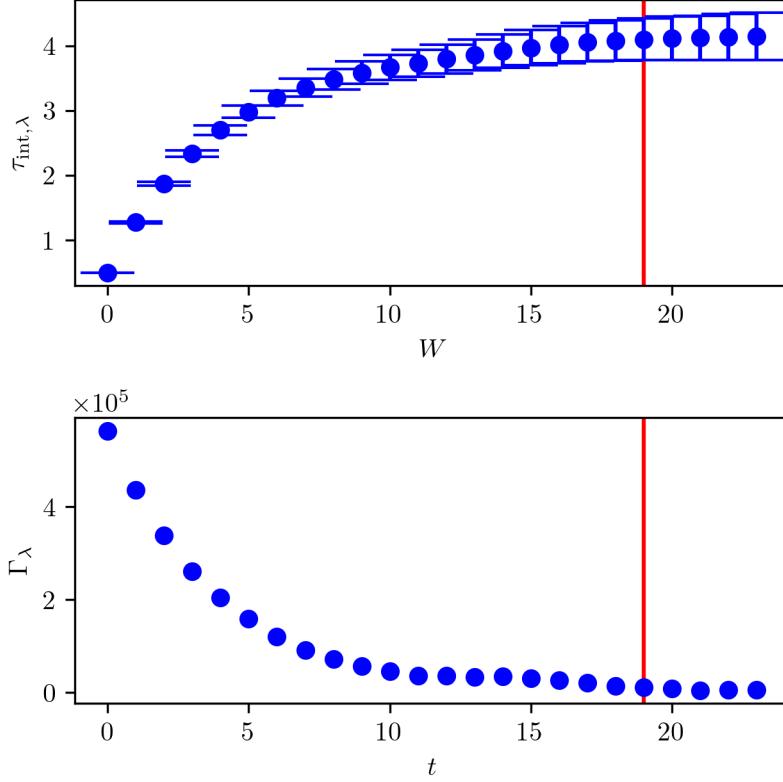


Figure 4.4: Provided by [21], the IACT $\tau_{\text{int},\lambda}$ at summation windows W is plotted as well as the autocorrelation function Γ_λ based on the samples $\lambda \sim \pi(\gamma, \lambda | \mathbf{y})$.

TT Approximation of Marginal Posterior Alternatively, we can utilise a TT approximation of the square root of the marginal posterior over a predefined grid and calculate the marginals $\pi(\gamma | \mathbf{y})$ and $\pi(\lambda | \mathbf{y})$ (see Sec. 2.3.1). We predefined the univariate grid with $n = 20$ grid points (see Fig. 4.7) over $\gamma = [0.25 \times 10^{15}, 6 \times 10^{15}]$ and $\lambda = [1, 7000]$.

We introduce a "normalisation constant" $c = 340$ to avoid underflow so that the values of $\sqrt{\pi(\lambda, \gamma | \mathbf{y})} = \exp\{0.5 \log \pi(\lambda, \gamma | \mathbf{y}) + c\}$ are within computer precision. Then we initialise the `tt.cross.rectcross.rect_cross.cross` function, based on the TT cross algorithm in [56, 57], from the Python package `ttypy` [58], with a random tensor. The number of ranks is set to constant $r = 5$, we do 1 sweep with $2n_{\text{sweeps}}2nr = 400$ function evaluations and obtain a TT approximation of $\pi(\lambda, \gamma | \mathbf{y})$ in about 0.1s. Ironically, this the same number of functions evaluations to approximate a 20 point grid in 2D, but the marginalisation over λ or γ in the TT format is computationally less expensive compared to a classical marginalisation over each individual hyper-parameter value. For larger grid sizes and higher dimensional parameter spaces the TT format is even more advantageous. To compute the marginals $\pi(\lambda | \mathbf{y})$ and $\pi(\gamma | \mathbf{y})$ we set the TT error $\xi = 1/\lambda(\mathcal{X})$ and $\lambda(x) = 1$, so that for Cartesian basis $\mathbf{M}_k = \text{diag}(\lambda_k(\mathcal{X}_k))$ in Eq. 2.35. We calculate the coefficient tensor \mathbf{B} and \mathbf{R}_{pre} as in Prop. 1 and Prop. 2 (see Sec. 2.3.1).

We plot the TT approximation as a colour map on top of the obtained samples in Fig. 4.5. The relative RMS TT approximation error over the whole grid is $\approx 7\%$ and similar to the propagation error in $\pi(\lambda, \gamma|y)$ due to the approximations of $f(\lambda)$ and $g(\lambda)$ (see further up).

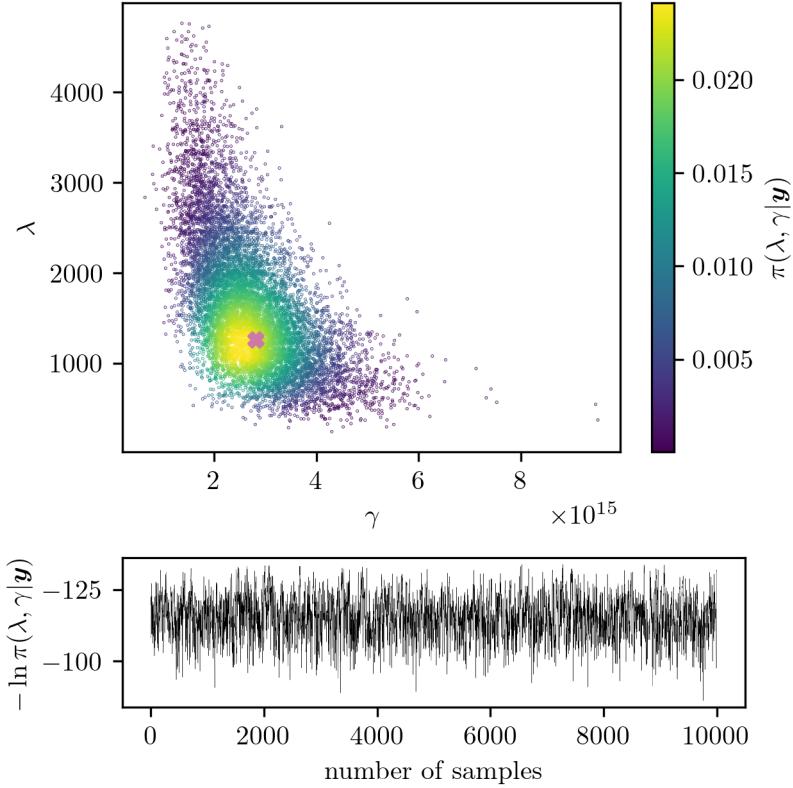


Figure 4.5: Samples from the marginal posterior $\pi(\lambda, \gamma|y)$ colour coded using the TT approximation of $\pi(\lambda, \gamma|y)$. The mode of (λ_0, γ_0) of $\pi(\lambda, \gamma|y)$ is marked with the pink cross. To show ergodicity, we plot the trace of the samples of the MWG algorithm.

Full Posterior Ozone Mean and Variance

Finally, we evaluate the normally distributed conditional posterior distribution

$$\mathbf{x}|\delta, \gamma, \mathbf{y} \sim \mathcal{N}\left(\underbrace{(\mathbf{A}^T \mathbf{A} + \delta/\gamma \mathbf{L})^{-1} \mathbf{A}^T \mathbf{y}}_{\mathbf{x}_\lambda}, \underbrace{(\gamma \mathbf{A}^T \mathbf{A} + \delta \mathbf{L})^{-1}}_{\gamma \mathbf{B}_\lambda}\right), \quad (4.17)$$

as in Eq. 2.15, with $\lambda = \delta/\gamma$. In this thesis, we compute the mean

$$\mu_{\mathbf{x}|y} = \int \mathbf{x}_\lambda \pi(\lambda|y) d\lambda \approx \sum \mathbf{x}_{\lambda_i} \pi(\lambda_i|y), \quad (4.18)$$

and covariance

$$\Sigma_{\mathbf{x}|y} = \int \gamma^{-1} \pi(\gamma|y) d\gamma \int \mathbf{B}_\lambda^{-1} \pi(\lambda|y) d\lambda \approx \sum \gamma_i^{-1} \pi(\gamma_i|y) \sum \mathbf{B}_{\lambda_i}^{-1} \pi(\lambda_i|y) \quad (4.19)$$

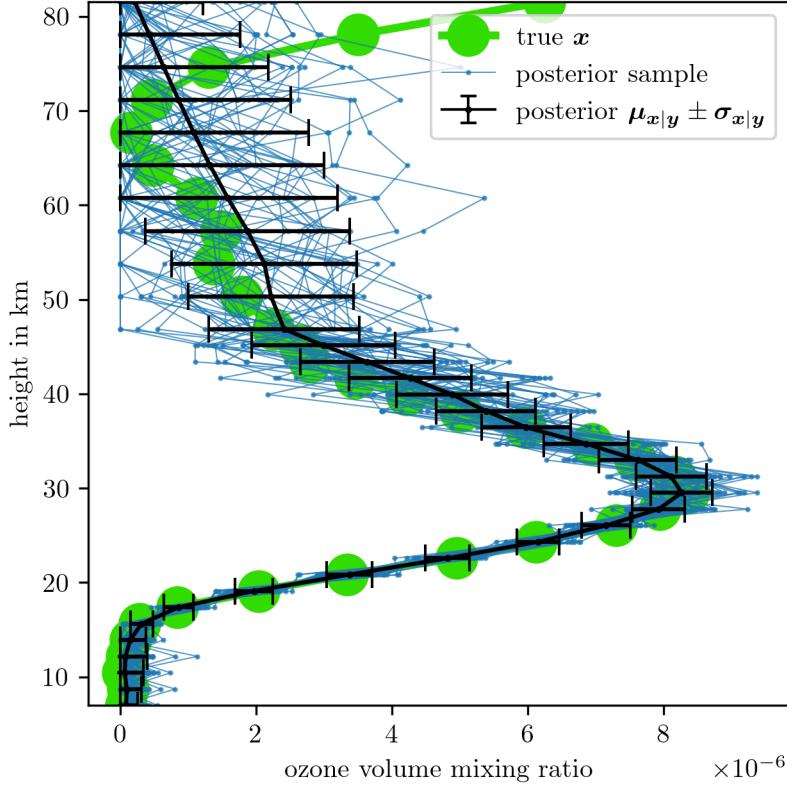


Figure 4.6: Ozone samples from the full posterior distribution $\pi(\mathbf{x}|\mathbf{y})$ after characterising full posterior mean and covariance by weighted expectations over the marginal posterior $\pi(\lambda, \gamma|\mathbf{y})$ based on the linear forward map \mathbf{A}_L . We set negative ozone VMR values to zero.

of $\pi(\mathbf{x}|\mathbf{y})$ as weighted expectations over the marginal posterior $\pi(\lambda, \gamma|\mathbf{y})$, by quadrature [59, Sec. 2.1], with $\sum \pi(\lambda_i|\mathbf{y}) = \sum \pi(\gamma_i|\mathbf{y}) = 1$. The weights $\pi(\lambda_i|\mathbf{y})$ and $\pi(\gamma_i|\mathbf{y})$ are either given by the TT approximation or by the bins for the sample-based histograms. More precisely, the heights of the sample-based histogram bars act as quadrature weights, where λ_i is defined at the centre of each bin . We use Cholesky decomposition of $\mathbf{B}_\lambda = \mathbf{A}^T \mathbf{A} + \lambda \mathbf{L}$ to invert \mathbf{B}_λ and to calculate $\mathbf{x}_\lambda = (\mathbf{A}^T \mathbf{A} + \lambda \mathbf{L})^{-1} \mathbf{A}^T \mathbf{y}$ both via `scipy.linalg.cho_solve`. It is sufficient, to evaluate \mathbf{x}_λ and invert \mathbf{B}_λ 20 times to obtain mean and covariance values of $\pi(\mathbf{x}|\mathbf{y})$ within a reasonable error (see Fig. 4.7). This takes less than 0.2s. We plot posterior samples of $\pi(\mathbf{x}|\mathbf{y})$ in Fig. 4.6 and set negative ozone values to zero, which we observe in almost every sample. The fact that we have to deal with negative ozone values is due to the poor prior choice in $\pi(\mathbf{x}|\delta)$. This indicates that we should use a different, more physically based prior or model a parametrised the ozone profile. Note that the posterior samples do not represent the ozone peak at around 80km.

If calculating the variance is too costly, the RTO method (see Sec. 2.2.3) may be a feasible alternative to draw a sample from Eq. 4.17.

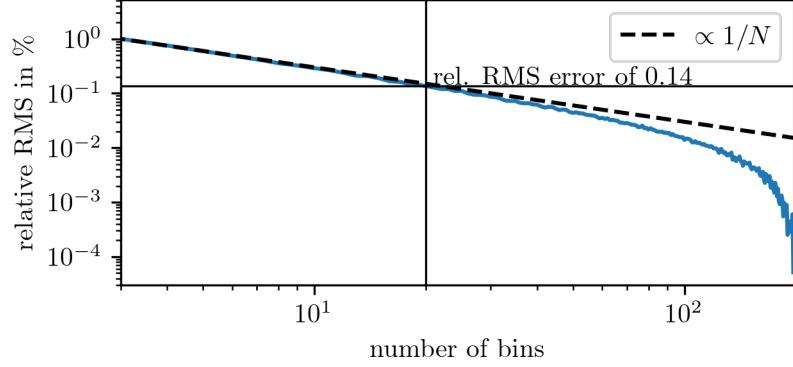


Figure 4.7: Relative RMS error of $\mu_{\mathbf{x}|\mathbf{y}}$ and covariance $\Sigma_{\mathbf{x}|\mathbf{y}}$ calculated by the weighted expectations and compared to a “ground truth” given by weighted expectations over 200 bins.

Errors of Posterior Mean and Covariance In Fig. 4.7, we plot the relative RMS error for the mean $\mu_{\mathbf{x}|\mathbf{y}}$ and covariance $\Sigma_{\mathbf{x}|\mathbf{y}}$ of $\pi(\mathbf{x}|\mathbf{y})$. We obtain those results by calculating the weighted expectation over normalised histograms of $\pi(\lambda, \gamma|\mathbf{y})$, where we vary the number of bins and compare to a solution calculated from a histogram with 200 bins. The relative error behaves roughly proportional to $1/N$ in log-space, and we consider a relative error less than 0.5% good enough, which we meet at approximately 20 bins. This sets the TT grid size and the number of evaluations of \mathbf{x}_λ in Eq. 4.18 and $(\gamma \mathbf{B}_\lambda)^{-1}$ in Eq. 4.19.

4.3 Approximate non-linear Forward Model with an Affine Map

Given the posterior distribution for ozone $\pi(\mathbf{x}|\mathbf{y})$, we can now approximate the non-linear forward model

$$\mathbf{A}(\mathbf{x}) \approx \mathbf{M}\mathbf{A}_L\mathbf{x}, \quad (4.20)$$

with an affine map \mathbf{M} (see Fig. 4.8 for the summarised strategy). Here $\mathbf{A}(\mathbf{x})$ is non-linear noise-free data and $\mathbf{A}_L\mathbf{x}$ is linear noise-free data, both with ground truth pressure and temperature.

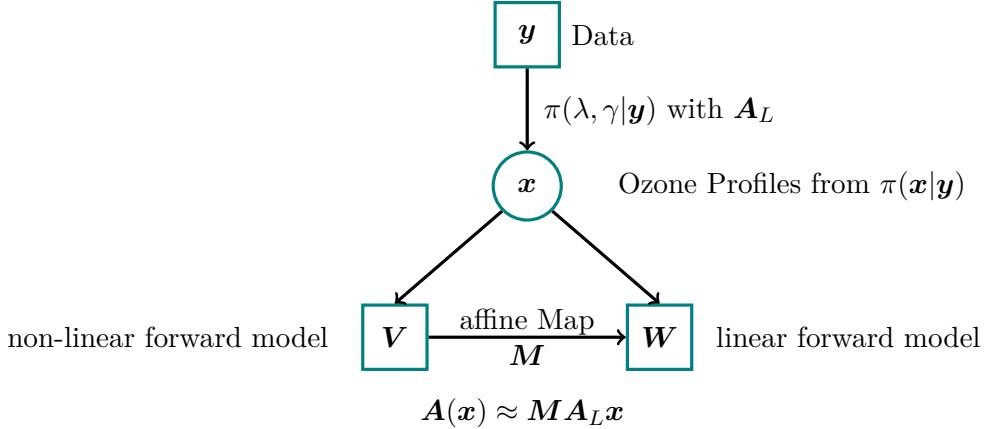


Figure 4.8: The strategy to find the affine map consists of evaluating the marginal posterior for ozone $\pi(\lambda, \gamma|\mathbf{y})$ based on the linear forward model first. Then we draw ozone samples from the full posterior. Next, we find an affine map which approximates between noise-free data based on the linear and non-linear forward model.

Using posterior ozone samples, we generate two affine subspaces and then find the mapping between those. The subspace \mathbf{W} is created by noise-free data based on the linear model and \mathbf{V} by noise-free data based on the non-linear model, given m samples $\mathbf{x}^{(j)} \sim \pi(\mathbf{x}|\mathbf{y})$ for $j = 1, \dots, m$. We report a relative RMS difference between \mathbf{W} and \mathbf{V} of about 1%, which we aim to reduce through the affine map \mathbf{M} . More specifically, the affine subspace associated with the linear forward model is

$$\mathbf{W} = \begin{bmatrix} | & | & | \\ \mathbf{A}_L\mathbf{x}^{(1)} & \dots & \mathbf{A}_L\mathbf{x}^{(j)} & \dots & \mathbf{A}_L\mathbf{x}^{(m)} \\ | & | & | \end{bmatrix} \in \mathbb{R}^{m \times m} \quad (4.21)$$

and with the non-linear forward model is

$$\mathbf{V} = \begin{bmatrix} | & & | & & | \\ \mathbf{A}(\mathbf{x}^{(1)}) & \cdots & \mathbf{A}(\mathbf{x}^{(j)}) & \cdots & \mathbf{A}(\mathbf{x}^{(m)}) \\ | & & | & & | \end{bmatrix} = \begin{bmatrix} — & v_1 & — \\ & \vdots & \\ — & v_j & — \\ & \vdots & \\ — & v_m & — \end{bmatrix} \in \mathbb{R}^{m \times m}. \quad (4.22)$$

Then we calculate the affine map

$$\mathbf{V}\mathbf{W}^{-1} = \mathbf{M} = \begin{bmatrix} — & r_1 & — \\ & \vdots & \\ — & r_j & — \\ & \vdots & \\ — & r_m & — \end{bmatrix} \in \mathbb{R}^{m \times m}. \quad (4.23)$$

by solving $v_j = r_j\mathbf{W}$ for each row r_j in \mathbf{M} , where $j = 1, \dots, m$, using the Python function `numpy.linalg.solve`. We can do that because every measurement in the data vector \mathbf{y} is independent of each other, and then every row v_j of $\mathbf{V} \in \mathbb{R}^{m \times m}$ is independent of each other as well.

We assess the affine map by calculating the relative RMS difference $\|\text{vec}(\mathbf{MW}) - \text{vec}(\mathbf{V})\|_{L^2}/\|\text{vec}(\mathbf{MW})\|_{L^2}$ between the mapped linear noise-free data and the non-linear noise-free data, which is approximately 0.001%. In Fig. 4.9, we show the mapping for one posterior ozone sample, which has not been used to create this mapping. In other words, this is an unseen event not in the training data. The relative RMS error for this approximation is roughly 0.03% and much smaller than the relative difference between noise-free linear data and non-linear data. Consequently, from here onwards, we use the approximated forward map.

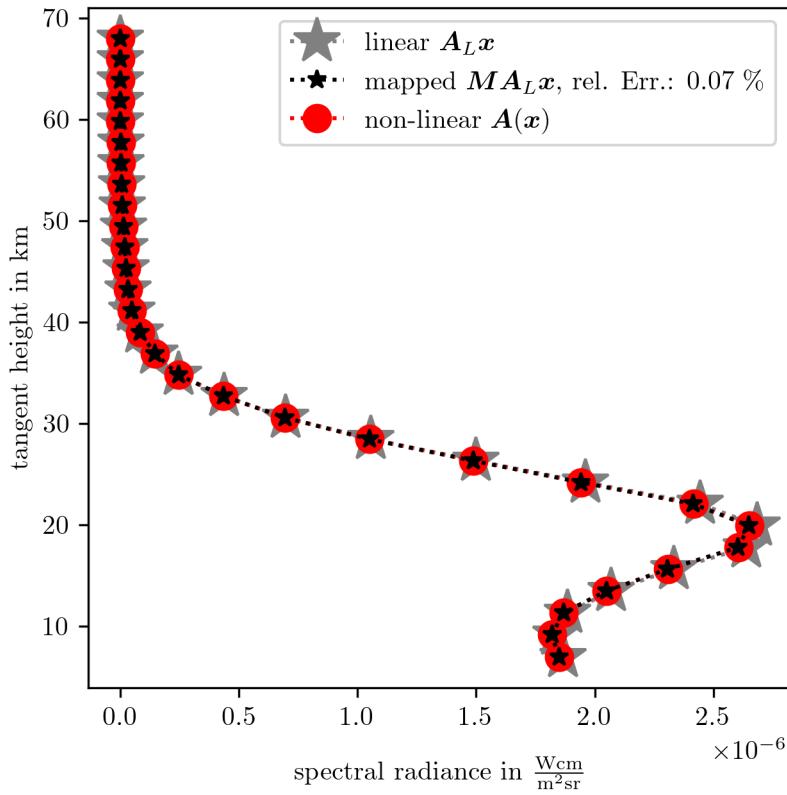


Figure 4.9: Assessment of how well we can approximate noise-free non-linear data $A(\mathbf{x})$ (red circles) with noise-free linear data $A_L \mathbf{x}$ (grey stars) and the previously calculated affine map M . The approximated noise-free data (black stars) has a relative RMS error of $\approx 0.03\%$ compared to the true non-linear noise-free data. The ozone sample to generate this noise-free data has not been used to create the affine map.

4.4 Regularisation Solution vs. Bayesian Approach – approximated Model

With the affine approximation, we define the linear forward model matrix

$$\mathbf{A} := \mathbf{M}\mathbf{A}_L \quad (4.24)$$

using the affine map \mathbf{M} . Here, we compare the posterior distribution of ozone to a regularisation approach.

4.4.1 Posterior Distribution for Ozone

Again, we use the MTC scheme and the exact same setup and procedure as in Sec. 4.2.2 to evaluate the marginal posterior and then the full conditional posterior of ozone with similar computational time.

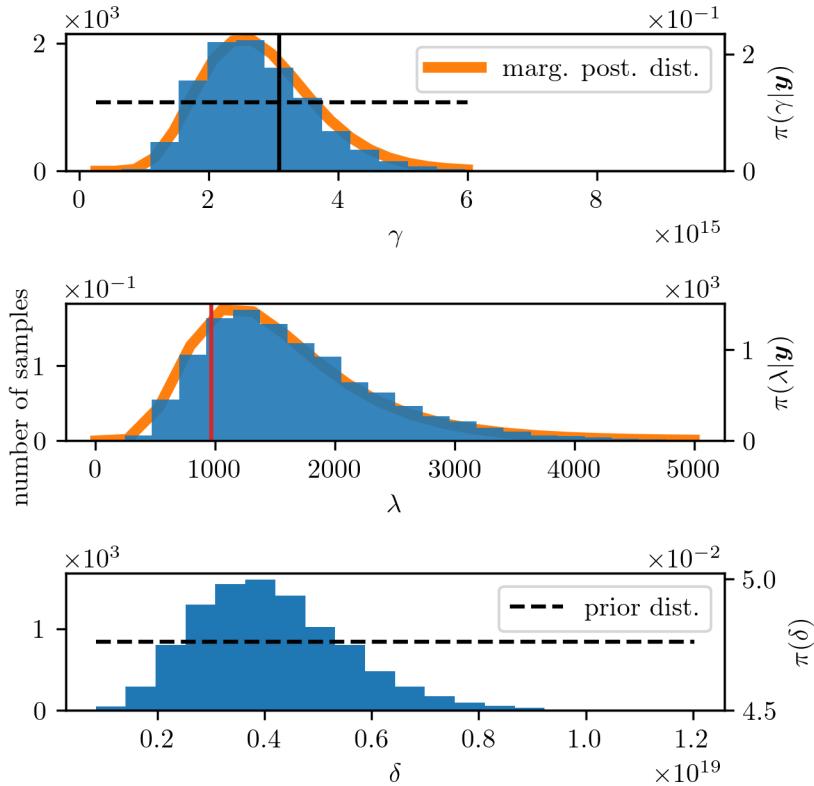


Figure 4.10: The TT approximation of the marginal posterior in orange and the samples as a histogram, as well as the prior distribution with a dotted line. We sample λ and γ using the MWG algorithm and then can calculate δ for every sample of the marginal posterior. The regularised parameter corresponding to the best regularised solution (see Fig. 4.11 and Fig. 4.13) is marked with the red vertical line. We mark the ground truth noise precision with the back vertical line.

The marginal posterior is defined as in Eq. 4.8, but with the approximated forward model. We initialise the MWG at and approximate $f(\lambda)$ and $g(\lambda)$ around the mode

of $\pi(\lambda, \gamma | \mathbf{y})$ (see Eq. 4.12a and Eq. 4.12b), and take $N = 10000$ plus $N_{\text{burn-in}} = 100$ steps. The IACT provided by [21] with $\tau_{\text{int},\gamma} \approx 2.0 \pm 0.1$ and $\tau_{\text{int},\lambda} = 4.3 \pm 0.4$ (see Fig. B.4 and Fig. B.3) are similar to previously calculated values. We plot the samples in Fig. 4.10 as well as TT approximation of the marginal posterior using 400 function evaluations (same grid; same number of ranks; see Sec. 4.2.2).

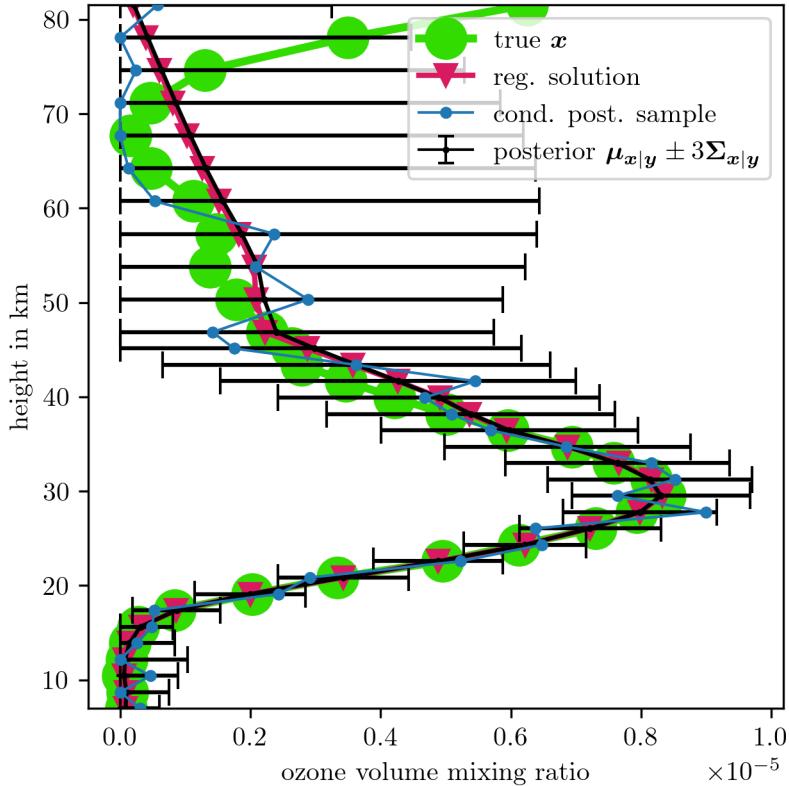


Figure 4.11: Full posterior mean and variance and one ozone sample from the full posterior. We plot the regularised solution on top of the ground truth ozone profile in green. We use the updated forward map \mathbf{MA}_L

Again, we calculate the full posterior mean $\mu_{\mathbf{x}|\mathbf{y}}$, see Eq. 4.18, and covariance matrix $\Sigma_{\mathbf{x}|\mathbf{y}}$ 4.19 as weighted expectation. We plot the results and one sample of $\pi(\mathbf{x}|\mathbf{y})$, which represents a feasible solution to this inverse problem, in Fig. 4.11, as well as the regularised solution (see next section), and one sample from the posterior. We can see that the ground truth lies within 3 times of the STD (accounting for $\approx 99\%$ of all posterior samples) around the mean, except for the peak at around 80km. Compared to the previously calculated mean and variance based on the linear forward model \mathbf{A}_L (see 4.6), the posterior distribution based on \mathbf{MA}_L does not differ significantly. This is expected since the difference between the linear and non-linear forward map of $\approx 1\%$ is small.

We can visualise how the posterior samples are affected through the prior $\delta\mathbf{L}$ and the data (forward model) $\gamma\mathbf{A}^T\mathbf{A}$ by comparing their eigenvalues to the eigenvalues of

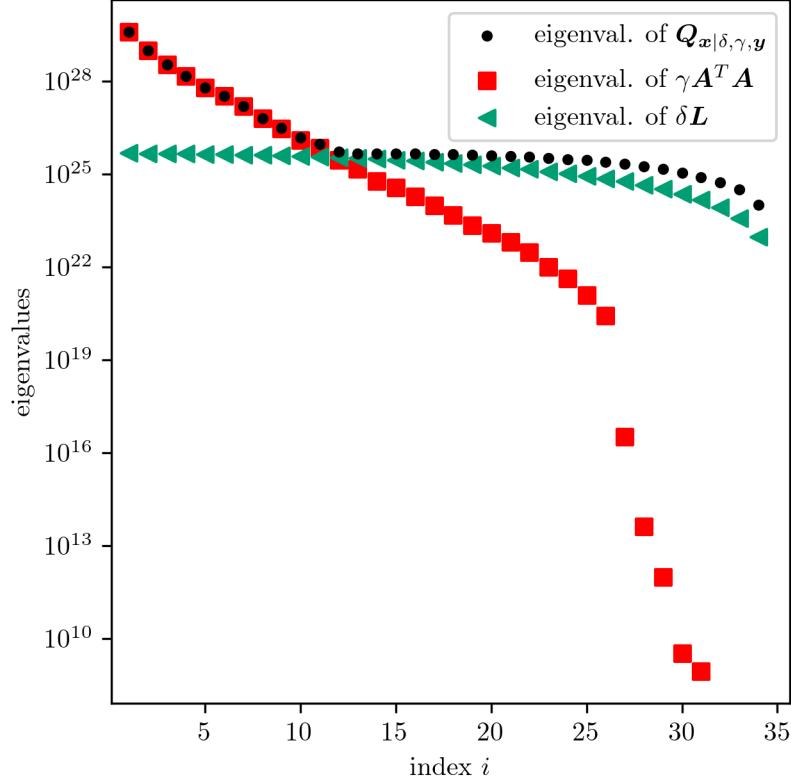


Figure 4.12: Eigenvalues of the precision matrix of $\mathbf{Q}_{\mathbf{x}|\delta,\gamma,\mathbf{y}} = \gamma \mathbf{A}^T \mathbf{A} + \delta \mathbf{L}$ of the full posterior distribution $\pi(\mathbf{x}|\delta, \gamma, \mathbf{y})$ for ozone. We see that large eigenvalues of $\gamma \mathbf{A}^T \mathbf{A}$ and $\delta \mathbf{L}$ are rather unaffected by the prior compared to small eigenvalues. The eigenbasis may differ.

the precision matrix $\mathbf{Q}_{\mathbf{x}|\delta,\gamma,\mathbf{y}} = \gamma \mathbf{A}^T \mathbf{A} + \delta \mathbf{L}$. We order the eigenvalues in size and plot those in Fig. 4.12. We observe that the larger eigenvalues of $\mathbf{Q}_{\mathbf{x}|\delta,\gamma,\mathbf{y}}$ are very much the same as the larger eigenvalues of $\mathbf{A}^T \mathbf{A}$. Once the eigenvalues of $\mathbf{A}^T \mathbf{A}$ are significantly smaller the structure of the eigenvalues of $\mathbf{Q}_{\mathbf{x}|\delta,\gamma,\mathbf{y}}$ is dominated by the structure of the eigenvalues of $\delta \mathbf{L}$. Note that the eigenvalues correspond to different eigenspaces spanned by the corresponding eigenvectors which may differ. The largest 10 eigenvalues of $\mathbf{Q}_{\mathbf{x}|\delta,\gamma,\mathbf{y}}$ to ozone profile structure in lower altitudes, where as the other eigenvalues correspond mainly to structures mainly in higher altitudes see B.25.

4.4.2 Solution by Regularisation

Since we claim that the Bayesian approach is superior to regularisation methods, we compare the MTC method to a Tikhonov regularisation, see Sec. 2.5 and [2]. This is most similar to our chosen linear-Gaussian Bayesian framework.

The regularised solution is defined as in [2, 60]

$$\mathbf{x}_\lambda = \arg \min_{\mathbf{x}} \|\mathbf{A}\mathbf{x} - \mathbf{y}\|_{L^2}^2 + \lambda \mathbf{x}^T \mathbf{L} \mathbf{x}, \quad (4.25)$$

with the regularisation parameter λ , and is typically calculated by solving the normal equations

$$\mathbf{x}_\lambda = (\mathbf{A}^T \mathbf{A} + \lambda \mathbf{L})^{-1} \mathbf{A}^T \mathbf{y} \quad (4.26)$$

(see Sec. 2.5). To find the best regularised solution, we use the L-curve method, and follow [44]. Within this method we compute \mathbf{x}_λ , for 200 different λ values in between 10^0 to 10^6 and plot the solution semi norm $\sqrt{\mathbf{x}_\lambda^T \mathbf{L} \mathbf{x}_\lambda}$ against the data misfit norm $\|\mathbf{A} \mathbf{x}_\lambda - \mathbf{y}\|_{L^2}$ (see Figure 4.13). The regularised solution corresponds to the “corner” of the L-curve at the point of maximum curvature provided by the kneedle algorithm [61] using the function `kneed.KneeLocator` in less 0.1s. We plot the corresponding regularisation parameter in Fig. 4.10, which can vary significantly compared to the λ -samples of the marginal posterior.

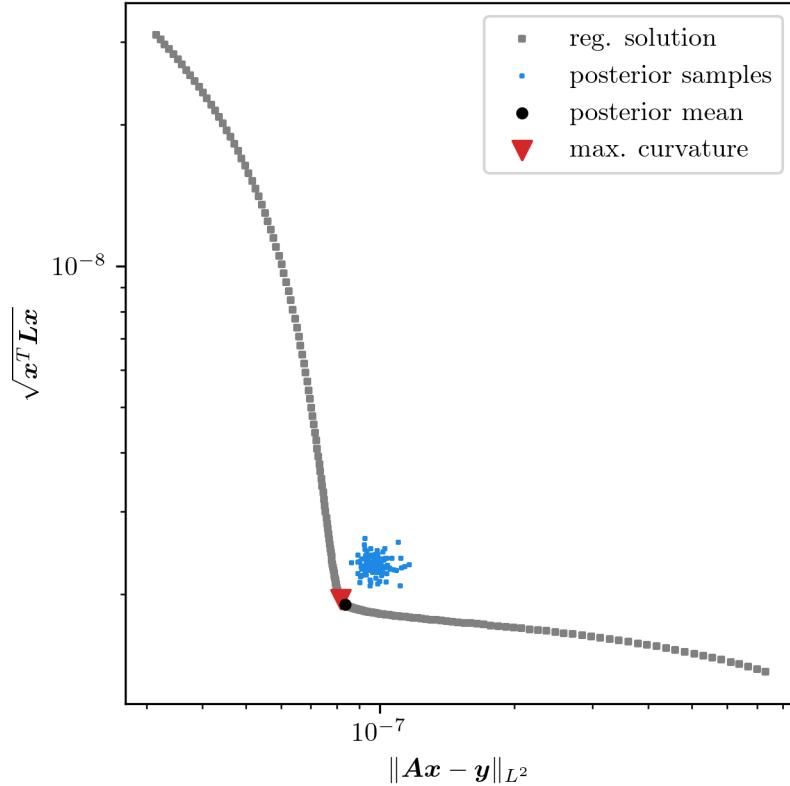


Figure 4.13: L-Curve of regularised semi norm $\sqrt{\mathbf{x}^T \mathbf{L} \mathbf{x}}$ against the data misfit norm $\|\mathbf{A} \mathbf{x}_\lambda - \mathbf{y}\|_{L^2}$ for different λ values, where \mathbf{x}_λ is calculated as in Eq. 4.26. The best regularised solution is at the point of maximum curvature (pink triangle). Additionally, we calculate the data misfit norm and the regularised norm for the mean (black circle) and samples (blue squares) of the full posterior of ozone.

The regularised solution in Fig. 4.11 is very similar to the posterior mean. It is pretty clear that the regularised solution accounts for only one possible solution and does not

provide uncertainties. The regularised solution is not similar to the samples drawn from the posterior $\pi(\mathbf{x}|\mathbf{y})$, see also Fig. 4.6. In Fig. 4.13, the samples of $\pi(\mathbf{x}|\mathbf{y})$ lie above the L-Curve, whereas the posterior mean and the regularised solution lie on the L-Curve. This does make sense, if one thinks about the mean as an (smooth) average over less-smooth samples and the regularised solution as an extremely smooth ozone profiles, see also Sec. 2.5. In contrast, the samples are less regularised and hence lie above the L-Curve, but have a similar data misfit norm, and as already mentioned, are all feasible solutions to the data. Neither the regularisation solution nor the posterior ozone profiles capture the ozone peak of the ground truth at high altitudes.

4.5 Hierarchical Bayesian Framework for Ozone, Pressure and Temperature

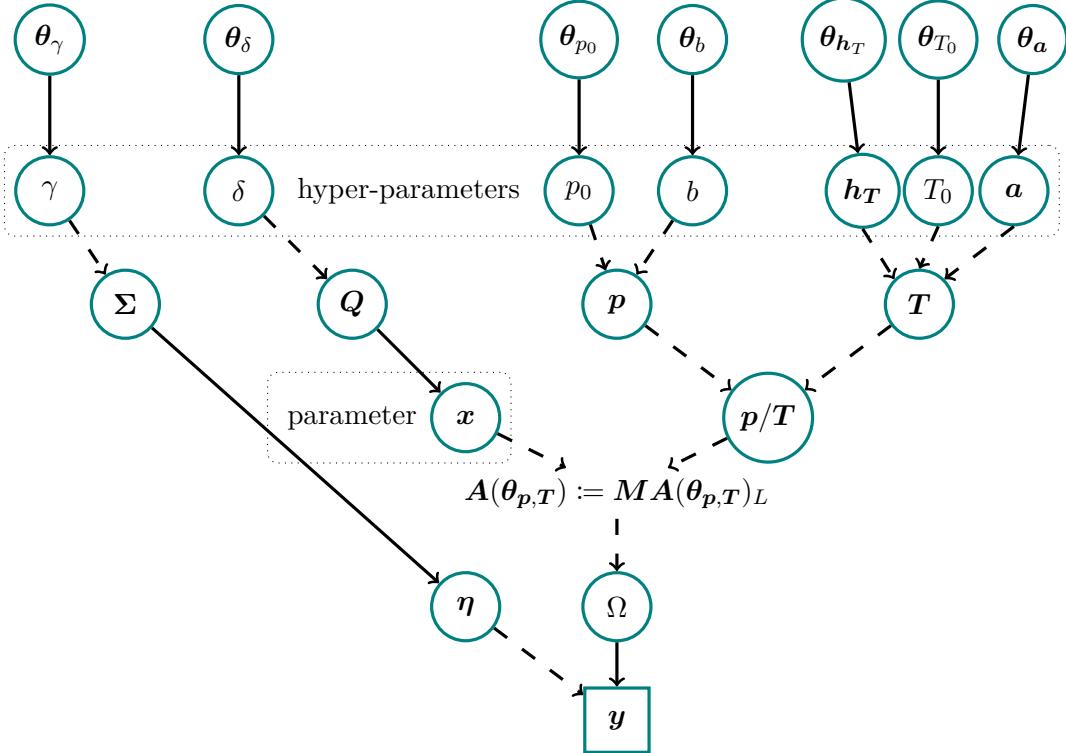


Figure 4.14: DAG of Bayesian model for ozone, pressure and temperature. The hyper-parameters $\mathbf{h}_T = \{h_{T,1}, h_{T,2}, h_{T,3}, h_{T,4}, h_{T,5}, h_{T,6}\}$, $\mathbf{a} = \{a_0, a_1, a_2, a_3, a_4, a_5, a_6\}$, T_0 , b and p_0 deterministically (dotted line) describe pressure parameter \mathbf{p} through the function in Eq. 4.28, and temperature parameter \mathbf{T} through the function in Eq. 4.3. In this case, we choose the hyper-parameters $\pi(p_0, b, T_0, \mathbf{h}_T, \mathbf{a})$ to be a normally distributed apriori, determined by $\theta_{h_T}, \theta_a, \theta_{T_0}, \theta_b, \theta_{p_0}$, which represent mean and variances. The ozone parameter \mathbf{x} is statistically (solid line) described by the prior distribution $\mathbf{x}|\delta \sim \mathcal{N}(0, \delta \mathbf{L})$. Here, the hyper-parameter δ accounts for smoothness in the ozone profile and defines the precision matrix $\mathbf{Q} = \delta \mathbf{L}$, where \mathbf{L} is the graph Laplacian as in Eq. 4.5. The noise covariance $\Sigma = \gamma^{-1} \mathbf{I}$ of the random noise vector $\boldsymbol{\eta} \sim \mathcal{N}(0, \gamma^{-1} \mathbf{I})$ is defined by the noise hyper-parameter γ . As in Sec. 4.2 described, θ_δ and θ_γ define the hyper-priors $\pi(\delta, \gamma)$. Then, we randomly observe a data set \mathbf{y} from the space of all measurables Ω through the approximated forward model $A(\theta_{p,T}) := MA(\theta_{p,T})_L$, depending on the hyper-parameter $\theta_{p,T} := \{p_0, b, \mathbf{h}_T, T_0, \mathbf{a}\}$, including some added noise. Given the data we like to determine the posterior distribution over the hyper-parameters $\pi(p_0, b, T_0, \mathbf{h}_T, \mathbf{a}, \delta, \gamma | \mathbf{y})$ first and then $\pi(\mathbf{x} | p_0, b, T_0, \mathbf{h}_T, \mathbf{a}, \delta, \gamma, \mathbf{y})$ utilising the MTC scheme.

As in Sec. 4.2, we use a DAG as in Fig. 4.14 to visualise the measurement process and correlations between pressure \mathbf{p} , temperature \mathbf{T} and ozone \mathbf{x} , which progress deterministically (dashed line) into the forward model, via $\mathbf{x} \times \mathbf{p}/\mathbf{T}$. Through their respective prior distributions, they generate a space of all possible noise-free data Ω , from which we observe some data, including some added normally distributed noise $\boldsymbol{\eta}$. This hierarchical Bayesian framework includes the hyper-parameters p_0, b for pressure (see Eq. 4.28), $\mathbf{a}, \mathbf{h}_T, T_0$ for temperature (see Eq. 4.3), δ for ozone smoothness and

model parameters	priors	TT bounds			τ_{int}	Context
		lower	upper	τ_{int}		
\boldsymbol{x}	$\mathcal{N}(0, \delta \mathbf{L})$	-	-	-	-	\boldsymbol{x}
γ	$\mathcal{T}(1, 10^{-10})$	$5 \cdot 10^{-8}$	$4.5 \cdot 10^{-7}$	9 ± 0.1	9 ± 0.1	\boldsymbol{y}
δ	$\mathcal{T}(1, 10^{-10})$	-	-	1.5 ± 0.1	1.5 ± 0.1	\boldsymbol{x}
$\lambda = \delta/\gamma$	-	500	10^4	3.5 ± 0.3	3.5 ± 0.3	\boldsymbol{x}
p_0	$\mathcal{N}(1243, 5)$	1229	1259	550 ± 9	550 ± 9	\boldsymbol{p}
T_0	$\mathcal{N}(288.15, 4.5)$	275	302	2446 ± 76	2446 ± 76	\boldsymbol{T}
$h_{T,1}$	$\mathcal{N}(11, 0.5)$	9.5	12.5	1820 ± 49	1820 ± 49	\boldsymbol{T}
b	$\mathcal{N}(0.167, 5 \cdot 10^{-4})$	0.165	0.171	2813 ± 92	2813 ± 92	\boldsymbol{p}
$h_{T,3}$	$\mathcal{N}(32.3, 2.5)$	25.2	39.8	394 ± 5	394 ± 5	\boldsymbol{T}
a_0	$\mathcal{N}(-6.5, 0.01)$	-6.53	-6.47	330 ± 4	330 ± 4	\boldsymbol{T}
$h_{T,2}$	$\mathcal{N}(20.1, 1.6)$	17.7	22.3	454 ± 7	454 ± 7	\boldsymbol{T}
a_1	$\mathcal{N}(0, 0.1)$	-0.3	0.3	508 ± 8	508 ± 8	\boldsymbol{T}
a_2	$\mathcal{N}(1, 0.01)$	0.97	1.03	341 ± 5	341 ± 5	\boldsymbol{T}
a_3	$\mathcal{N}(2.8, 0.1)$	2.5	3.1	316 ± 4	316 ± 4	\boldsymbol{T}
$h_{T,4}$	$\mathcal{N}(47.4, 5)$	45.9	48.9	324 ± 4	324 ± 4	\boldsymbol{T}
a_4	$\mathcal{N}(0, 0.1)$	-0.3	0.3	335 ± 4	335 ± 4	\boldsymbol{T}
$h_{T,5}$	$\mathcal{N}(51.4, 5)$	49.9	52.9	319 ± 4	319 ± 4	\boldsymbol{T}
a_5	$\mathcal{N}(-2.8, 0.1)$	-3.1	-2.5	335 ± 4	335 ± 4	\boldsymbol{T}
$h_{T,6}$	$\mathcal{N}(71.8, 3)$	62.5	80.8	347 ± 5	347 ± 5	\boldsymbol{T}
a_6	$\mathcal{N}(-2, 0.01)$	-2.03	-1.97	320 ± 4	320 ± 4	\boldsymbol{T}

Table 4.2: Summary of relevant parameter characteristics, bounds and sampling statistics. We denote $\mathcal{N}(\mu, \sigma)$ as the Gaussian and $\mathcal{T}(\alpha = \text{scale}, \beta = \text{rate})$ as the gamma distribution. The IACT τ_{int} is estimated as in [62] from posterior samples based on the approximated forward map.

γ for noise precision. Each of those hyper-parameters is described by the hyper-prior distribution $\pi(p_0, b, T_0, \mathbf{h}_T, \mathbf{a}, \delta, \gamma)$ defined by us. Here $\boldsymbol{\theta}_\gamma, \boldsymbol{\theta}_\delta$ determine gamma distributions, e.g. $\gamma \sim \Gamma(\alpha_\gamma, \beta_\gamma)$ with $\boldsymbol{\theta}_\gamma = \{\alpha_\gamma, \beta_\gamma\}$, and $\boldsymbol{\theta}_{p_0}, \boldsymbol{\theta}_b, \boldsymbol{\theta}_h, \boldsymbol{\theta}_{T_0}, \boldsymbol{\theta}_a$ determine a normal distribution $p_0, b, \mathbf{h}_T, \mathbf{T}_0, \mathbf{a} \sim \pi(\boldsymbol{\theta}_{p_0}, \boldsymbol{\theta}_b, \boldsymbol{\theta}_h, \boldsymbol{\theta}_{T_0}, \boldsymbol{\theta}_a)$, e.g. $b \sim \mathcal{N}(\mu_b, \sigma_b)$ and $\boldsymbol{\theta}_b = \{\mu_b, \sigma_b\}$. Here we use the approximated forward model $\mathbf{A}(\boldsymbol{\theta}_{p,T}) := \mathbf{M}\mathbf{A}(\boldsymbol{\theta}_{p,T})_L$ and denote $\boldsymbol{\theta}_{p,T} := \{p_0, b, T_0, \mathbf{h}_T, \mathbf{a}\}$, which includes all hyper-parameter related to pressure and temperature.

Then, we set up the hierarchical Bayesian framework

$$\mathbf{y}|\mathbf{x}, p_0, b, T_0, \mathbf{h}_T, \mathbf{a}, \delta, \gamma \sim \mathcal{N}(\mathbf{A}(p_0, b, T_0, \mathbf{h}_T, \mathbf{a}) \mathbf{x}, \gamma^{-1} \mathbf{I}) \quad (4.27a)$$

$$\mathbf{x}|\delta \sim \mathcal{N}(\mathbf{0}, (\delta \mathbf{L})^{-1}) \quad (4.27b)$$

$$\delta \sim \Gamma(\alpha_\delta, \beta_\delta) \quad (4.27c)$$

$$\gamma \sim \Gamma(\alpha_\gamma, \beta_\gamma) \quad (4.27d)$$

$$\mathbf{a} \sim \mathcal{N}(\boldsymbol{\mu}_{\mathbf{a}}, \boldsymbol{\Sigma}_{\mathbf{a}}) \quad (4.27e)$$

$$\mathbf{h}_T \sim \mathcal{N}(\boldsymbol{\mu}_T, \boldsymbol{\Sigma}_{\mathbf{h}_T}) \quad (4.27f)$$

$$T_0 \sim \mathcal{N}(\mu_{T_0}, \sigma_{T_0}) \quad (4.27g)$$

$$p_0 \sim \mathcal{N}(\mu_{p_0}, \sigma_{p_0}) \quad (4.27h)$$

$$b \sim \mathcal{N}(\mu_b, \sigma_b) \quad (4.27i)$$

and define a normally distributed likelihood (due to Gaussian noise) and normally distributed priors. Before we formulate the posterior distribution we carefully define $\theta_\gamma, \theta_\delta, \theta_{p_0}, \theta_b, \theta_h, \theta_{T_0}, \theta_a$, the hyper-prior scales, shapes, means and variances, which are explicitly given in Tab. 4.2.

4.5.1 Prior Modelling

We start by describing the pressure p in between $h_{L,0} \approx 7\text{km}$ and $h_{L,n} \approx 82\text{km}$ with an exponential function

$$p(h) = \exp(-b h) p_0, \quad h_{L,0} \leq h \leq h_{L,n} \quad (4.28)$$

depending on two hyper-parameters p_0, b , see Fig. 4.16. Similarly, the temperature as described in Eq. 4.3 can be parametrised with 14 hyper-parameters $\mathbf{h}_T = \{h_{T,1}, h_{T,2}, h_{T,3}, h_{T,4}, h_{T,5}, h_{T,6}\}$, $\mathbf{a} = \{a_0, a_1, a_2, a_3, a_4, a_5, a_6\}$ and T_0 (see Fig. 4.15 and Eq. 4.3). To complete the model, we have to define sensible hyper-prior variances and means of the normally distributed hyper-prior distribution for pressure and temperature-related hyper-parameters, where $\pi(\delta, \gamma)$ are the same gamma distributions as previously defined in Sec. 4.2.1. We tune the normal distribution $\pi(\mathbf{h}_T)$ so that the temperature profile maintains its structure, $h_{T,i} < h_{T,i+1}$ for $i = 1, \dots, 5$ (see Fig. B.5) and set $\pi(\mathbf{a})$ to a normal distribution as well. Similarly, we set $\pi(T_0)$ to a normal distribution so that it mimics a daily temperature variability of roughly 30K. We choose those rather informative hyper-prior distributions, because we find (see Fig. 4.17) that the data is uninformative about the temperature profile. The hyper-prior distribution $\pi(p_0, b)$ for pressure-related hyper-parameters is also normally distributed, but with rather large variance σ_b , where p_0 has a variability of around 80hPa, close to what we can observe

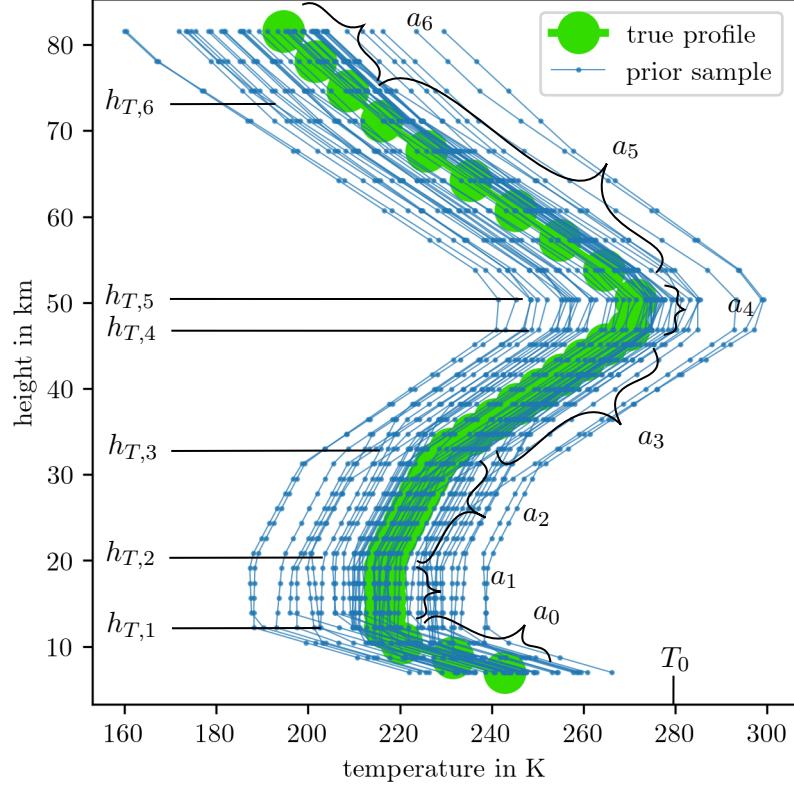


Figure 4.15: Prior samples from the hyper-prior distribution of \mathbf{h}_T , \mathbf{a} and T_0 , as defined in Tab. 4.2, where we calculate \mathbf{T} according to the function in Eq. 4.3.

when looking at weather data. We set means of the normal distribution $\pi(p_0, b, T_0, \mathbf{h}_T, \mathbf{a})$ to the ground truth values of \mathbf{p} and \mathbf{T} (see Tab. 4.2).

We plot prior samples of the pressure \mathbf{p} in Fig. 4.16, the temperature \mathbf{T} in Fig. 4.15 and the ratio \mathbf{p}/\mathbf{T} in Fig. 4.17 against the ground truth profiles. Additionally, we plot prior samples of $1/\mathbf{T}$ in Fig. B.6. Here we already observe that \mathbf{p}/\mathbf{T} inherits the structure of the pressure function and hence the model is uninformative about the temperature.

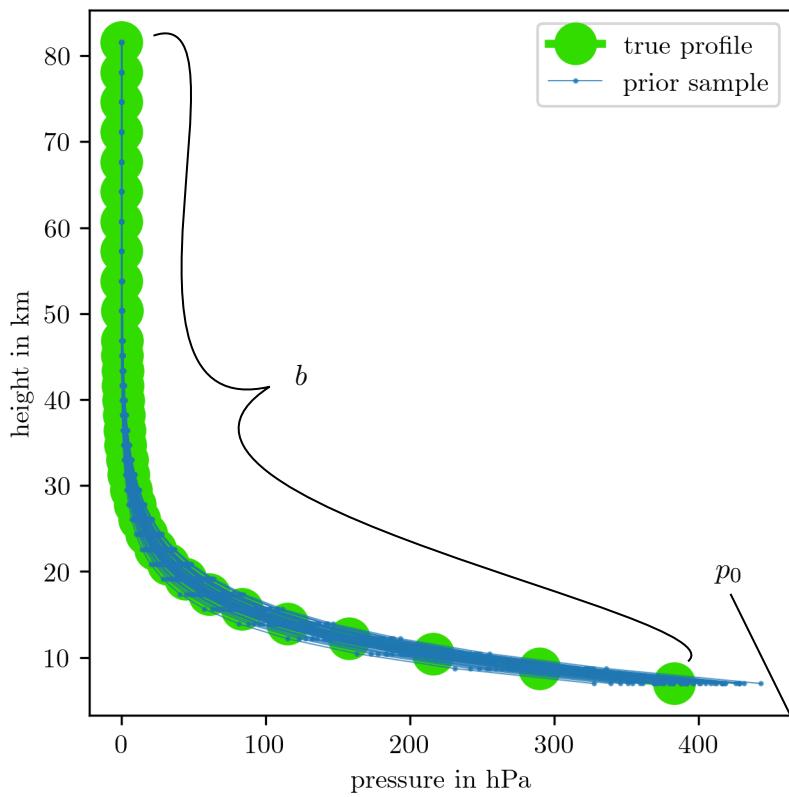


Figure 4.16: Prior samples from the hyper-prior distribution of b and p_0 as defined in Tab. 4.2 , where we calculate \mathbf{p} according to the function in Eq. 4.28.

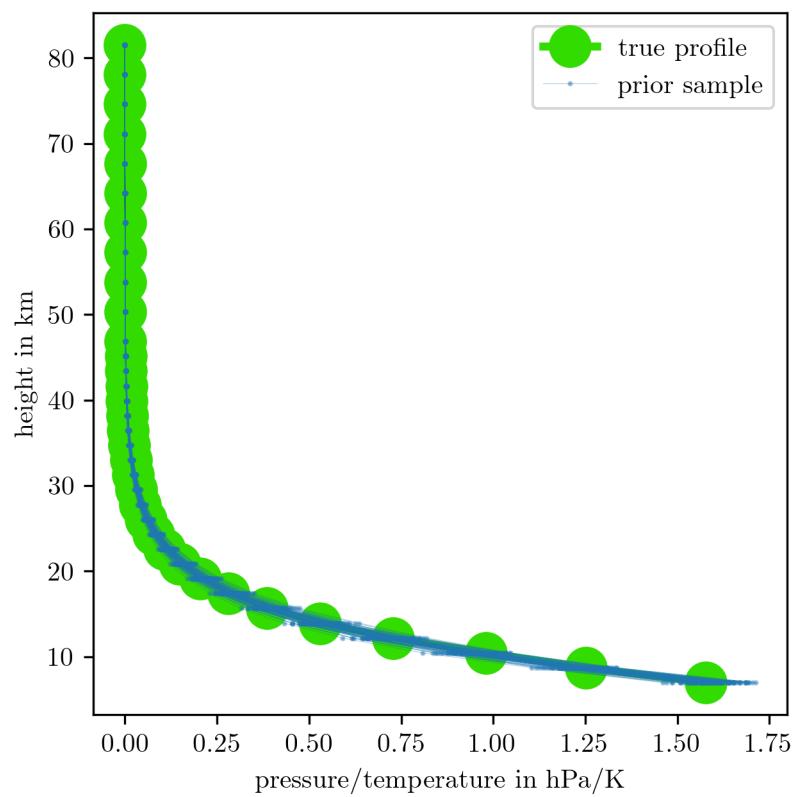


Figure 4.17: Prior samples from the hyper-prior distribution of \mathbf{h}_T , \mathbf{a} and T_0 for temperature as in Eq. 4.3 and \mathbf{b} and p_0 for pressure as in Eq. 4.28. We plot \mathbf{p}/\mathbf{T} . The hyper-priors are defined in Tab. 4.2.

4.5.2 Posterior Distribution

Here, we define the marginal and then the full conditional posterior distribution for the described Bayesian model. We either use the `t-walk` algorithm [4] to draw samples from $\pi(p_0, b, T_0, \mathbf{h}_T, \mathbf{a}, \lambda, \gamma | \mathbf{y})$ or we utilise a TT approximation on a predefined grid to generate samples via the SIRT method with an MH correction step. In doing so, we guide the reader through the procedure and point out some key aspects of how we obtain an efficient TT approximation. Lastly, we use the RTO method to draw samples from the conditional posterior $\pi(\mathbf{x} | p_0, b, T_0, \mathbf{h}_T, \mathbf{a}, \lambda, \gamma, \mathbf{y})$.

Marginal Posterior Distribution

The marginal posterior is given as

$$\pi(\boldsymbol{\theta}_{p,T}, \lambda, \gamma | \mathbf{y}) \propto \lambda^{n/2} \gamma^{m/2} \exp \left\{ -\frac{1}{2} g(\boldsymbol{\theta}_{p,T}, \lambda) - \frac{\gamma}{2} f(\boldsymbol{\theta}_{p,T}, \lambda) \right\} \pi(\boldsymbol{\theta}_{p,T}, \lambda, \gamma), \quad (4.29)$$

with $\lambda = \delta/\gamma$,

$$f(\boldsymbol{\theta}_{p,T}, \lambda) = \mathbf{y}^T \mathbf{y} - (\mathbf{A}(\boldsymbol{\theta}_{p,T})^T \mathbf{y})^T (\mathbf{A}(\boldsymbol{\theta}_{p,T})^T \mathbf{A}(\boldsymbol{\theta}_{p,T}) + \lambda \mathbf{L})^{-1} (\mathbf{A}(\boldsymbol{\theta}_{p,T})^T \mathbf{y}), \quad (4.30a)$$

$$\text{and } g(\boldsymbol{\theta}_{p,T}, \lambda) = \log \det (\mathbf{A}(\boldsymbol{\theta}_{p,T})^T \mathbf{A}(\boldsymbol{\theta}_{p,T}) + \lambda \mathbf{L}). \quad (4.30b)$$

For each evaluation of $\pi(\boldsymbol{\theta}_{p,T}, \lambda, \gamma | \mathbf{y})$ we compose $\mathbf{A}_L(\boldsymbol{\theta}_{p,T})$ as in Chapter 3, and calculate f and g directly using the Cholesky decomposition.

Sampling from Marginal Posterior For a ground truth, we run the `t-walk` [4] algorithm on $\pi(\boldsymbol{\theta}_{p,T}, \lambda, \gamma | \mathbf{y})$, where our objective is to generate 1000 independent samples from the marginal posterior. We find that the maximum IACT (see Tab. 4.2) is roughly 550, so we run the `t-walk` for $N = 1000 \times 1100$ steps with a burn-in period of $N_{\text{burn-in}} = 100 \times 1100$. We initialise the Python implementation of the `t-walk` [63] around the hyper-prior mean values and the mode of $\pi(\lambda, \gamma | \mathbf{y})$. We take a total number of $N' = N + N_{\text{burn-in}} = 1210000$ steps in ≈ 10 mins within bounds given by the iteratively defined TT grid (see Tab. 4.2). We plot the resulting histograms in Fig. 4.20 to Fig. 4.24 and the trace of the samples in Fig. B.7.

TT Approximation of Marginal Posterior The aim now is to approximate the square root of the marginal posterior

$$\pi(\lambda, \boldsymbol{\theta}_{p,T}, \gamma | \mathbf{y}) \propto \exp \left\{ \log \pi(\lambda, \boldsymbol{\theta}_{p,T}, \gamma | \mathbf{y}) + c \right\}, \quad (4.31)$$

where we introduce a "normalisation constant" $c = -100$ to avoid under or overflow and stay within computer precision. In doing so we run the `tt.cross.rectcross.rect_cross.cross` function from the `ttipy` python package [58]. We set the grid according to the results of the `t-walk`. If we compute the marginal as in Sec. 2.3, we set $\xi = 1/\lambda(\mathcal{X})$ and $\lambda(x) = 1$ so that for Cartesian basis $\mathbf{M}_k = \text{diag}(\lambda_k(\mathcal{X}_k))$. To draw samples from this TT approximation we use the SIRT-MH scheme as in Sec. 2.3.2.

Correlation structure First, we order the hyper-parameters according to their correlation structure to improve the efficiency of the Tensor-Train (TT) approximation. Specifically, we arrange the hyper-parameter space $\mathcal{X}_\gamma \times \mathcal{X}_\lambda \times \mathcal{X}_b \times \dots$ such that highly correlated hyper-parameter pairs are adjacent and directly linked through their shared TT rank. In Fig. 4.18 we plot 1000 independent samples drawn via the SIRT-MH scheme from the TT approximation of $\sqrt{\pi(\boldsymbol{\theta}_{p,T}, \lambda, \gamma | \mathbf{y})}$ and the Pearson correlation coefficient between hyper-parameter pairs. A coefficient close to 1 or -1 indicates strong correlation, while values near zero suggest weak or no correlation. We observe that the hyper-parameters λ and b , and λ and γ are highly correlated. Additionally, $h_{T,1}$ and T_0 describing the temperature at low altitudes (strong signal) are mildly correlated to b as well. This is because $h_{T,1}$ and T_0 influence "the smoothness" of p/T , which hard to see in Fig. 4.17. Interestingly, p_0 appears largely uncorrelated with other hyper-parameters, while b is the key parameter linking pressure to ozone and temperature. Hyper-parameters describing temperature at higher altitudes are very much uncorrelated and the IACTs in Tab. 4.2 agree with those results.

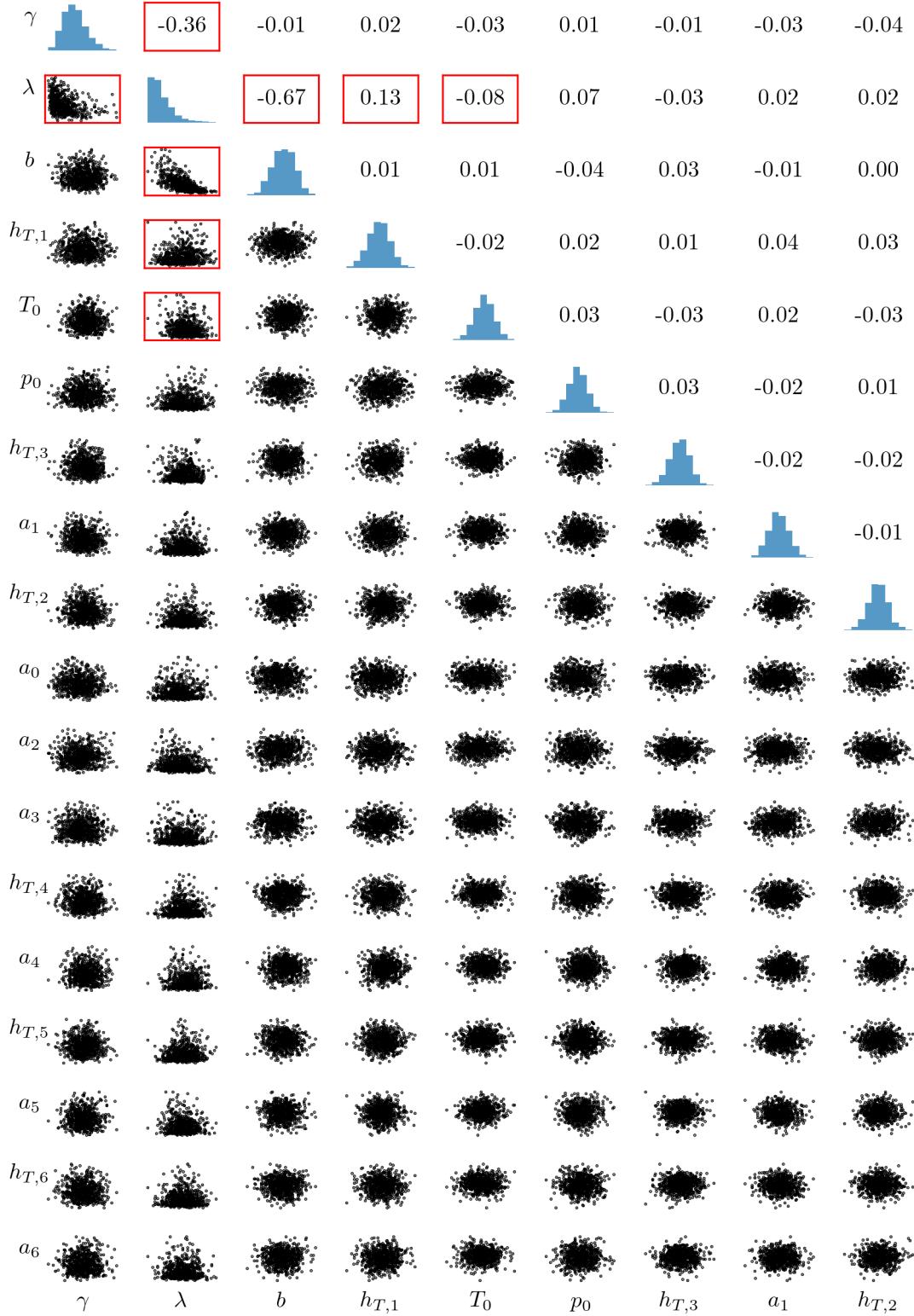
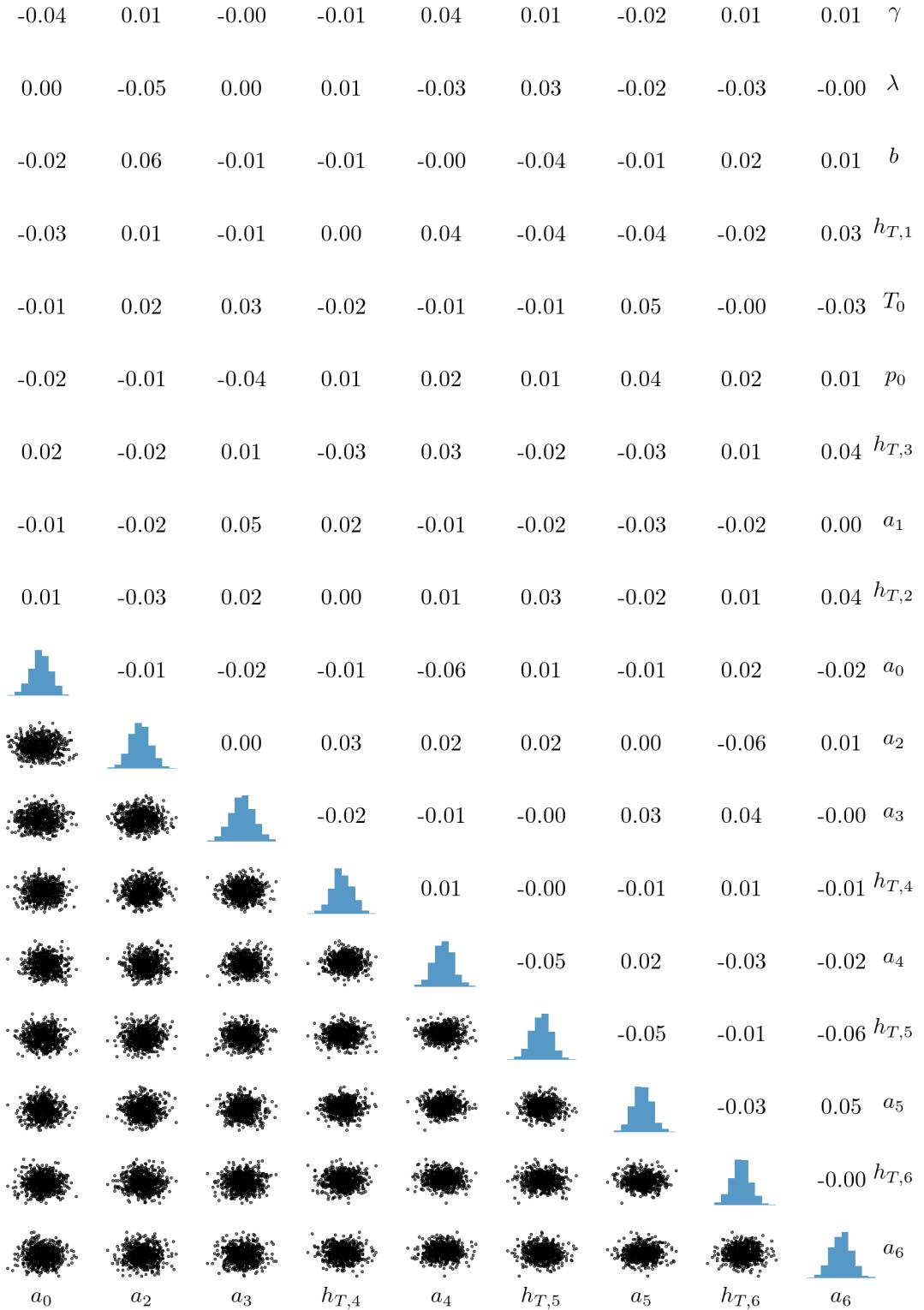


Figure 4.18: Plot of 1000 independent samples from TT approximation of $\sqrt{\pi(\theta_{p,T}, \lambda, \gamma | \mathbf{y})}$ via SIRT-MH scheme. We plot the Pearson correlation coefficient ranging from -1 to 1 for each hyper-parameter pair.



Correlation plot of samples from TT-approximation of $\sqrt{\pi(p_0, b, T_0, \mathbf{h}_T, \mathbf{a}_T | \mathbf{y}, \mathbf{x})}$ via SIRT scheme.

Find optimal Rank and Grid Size Next we aim to determine the optimal rank and grid size to accurately approximate the marginal posterior, while decreasing the number of function evaluation. We set the number of grid points to $n = 150$ and calculate different error measures for deceasing number of ranks to find the optimal number of ranks, where we compare to true marginal posterior function values and 1000 independent **t-walk** samples that distribution. Then we fix a small but tolerable rank and decrease the number of grid points until sufficient accuracy.

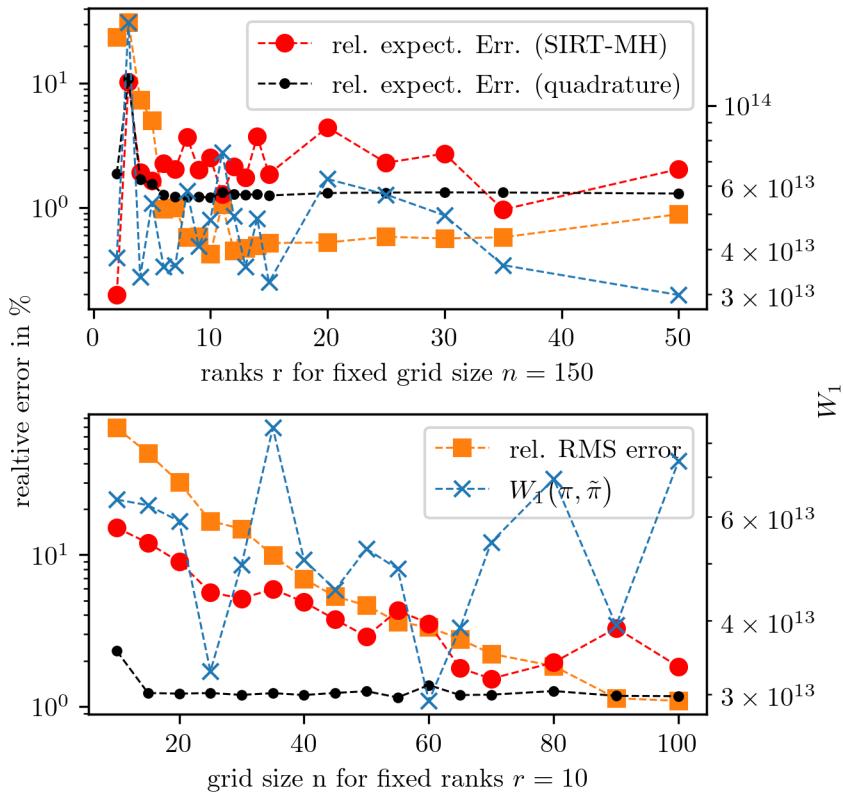


Figure 4.19: Given a TT approximation of $\sqrt{\pi(\lambda, \theta_{p,T}, \gamma | \mathbf{y})}$ we calculate the relative RMS error (orange squares) and the 1-Wasserstein distance (blue cross) between approximated values at sample points provided by the SIRT-MH and the true function values. We calculate the relative RMS error between the sample mean provided by the **t-walk** and mean values for the hyper-parameters calculated by quadrature (black dots), where we use the marginal function from the TT approximation as weights. Additionally, we plot the relative RMS error between the sample-based mean from the SIRT-MH and the **t-walk** (red circles).

For stable and comparable results, we do 5 sweeps in the `tt.cross.rectcross.rect_cross.cross` python function initialised with a random TT. Then we draw 1000 independent samples from the TT approximation of the marginal posterior via the SIRT-MH scheme. To calculate the 1-Wasserstein distance, as in Eq. 2.55, between SIRT-MH samples weighted with the TT approximation of marginal posterior, and the **t-walk** samples, weighted by the true marginal posterior values, we use

the `SamplesLoss("sinkhorn", p=1, blur=0.05, scaling=0.8)` function with default settings from the Python package `geomloss` [64]. This provides the unbiased Sinkhorn divergence, which converges towards the Wasserstein distance and can be understood as the generalised Quicksort algorithm [38]. Here, $p = 1$ defines the distance measure $\|\mathbf{x} - \tilde{\mathbf{x}}\|_{L^2}$, the blur parameter can be understood as an entropic penalty and the scaling parameter specifies the trade-off between speed ($\text{scaling} < .4$) and accuracy ($\text{scaling} > .9$) [64]. Additionally, we use the marginal functions of each TT approximation to calculate the means $\boldsymbol{\mu}_{\text{TT}} \in \mathbb{R}^{18}$ of each hyper-parameter by weighted expectations. Then we obtain relative RMS difference $\|\boldsymbol{\mu}_{\text{TT}} - \boldsymbol{\mu}_{\text{t-walk}}\|_{L^2}/\|\boldsymbol{\mu}_{\text{t-walk}}\|_{L^2}$. Here $\boldsymbol{\mu}_{\text{t-walk}}$ denotes the "true sample-based means" from the `t-walk`. Further, we calculate the relative RMS between the SIRT-MH sample-based mean $\boldsymbol{\mu}_{\text{SIRT-MH}}$ and $\boldsymbol{\mu}_{\text{t-walk}}$, and the relative RMS error at SIRT-MH samples compared to ground truth function values (Eq. 2.51).

We plot all of these measures in Fig. 4.19, where we observe that a rank $r = 10$ is sufficient because the error measures are relatively stable for $r \geq 10$. Then we decrease the grid size and decide that for $n \geq 20$ the relative differences of sample-based means to $\boldsymbol{\mu}_{\text{t-walk}}$ (red circles in Fig. 4.19) and the RMS error at the SIRT-MH samples (orange squares in Fig. 4.19) around 10% to 20% is good enough. We relate a more accurate interpolation in between grid points to those decreasing sample-based relative errors for an increasing number of grid points, since the chosen linear interpolation (see Eq. 2.48) is a rather rudimentary choice. The quadrature-based relative expectation error (black dots in Fig. 4.19) is almost constant for ranks ≥ 7 and grid sizes ≥ 15 . Since the hyper-parameters have different length scales, we are only interested in the trend of the 1-Wasserstein distance (blue crosses in Fig. 4.19), which is quite fluctuant but decreases with increasing ranks and stays within a similar range for decreasing grid size.

Further, we decrease the number of function evaluations and define ranks $r = [1, 10, 10, 10, 10, 10, 5, 5, 5, 5, 3, 2, 2, 2, 2, 2, 2, 1]$ harvesting the correlation structure of $\pi(\boldsymbol{\theta}_{p,T}, \lambda, \gamma | \mathbf{y})$ even more. We do one sweep in the `tt.cross.rectcross.rect_cross.cross`, reducing the computation time to ≈ 5 s and the number of function evaluations to 24120, where we initialise at a previously calculated approximation. We report an average IACT (from [20, 21]) of $\approx 0.6 \pm 0.1$ for the samples drawn via the SIRT-MH scheme, which means that we need 2 function evaluations per independent sample once the TT approximation is available. To draw 1000 independent samples, including generating a TT approximation, takes ≈ 0.5 min, and we report a relative RMS error of $\approx 25\%$ evaluated over those 1000 independent samples. We plot the marginals for each hyper-parameter in Fig. 4.20 to Fig. 4.25 and samples in Fig. 4.18. We observe that, besides λ and γ , only the marginal posterior of the b hyper-parameter is seriously affected by the data and has significantly changed compared to the hyper-prior distribution.

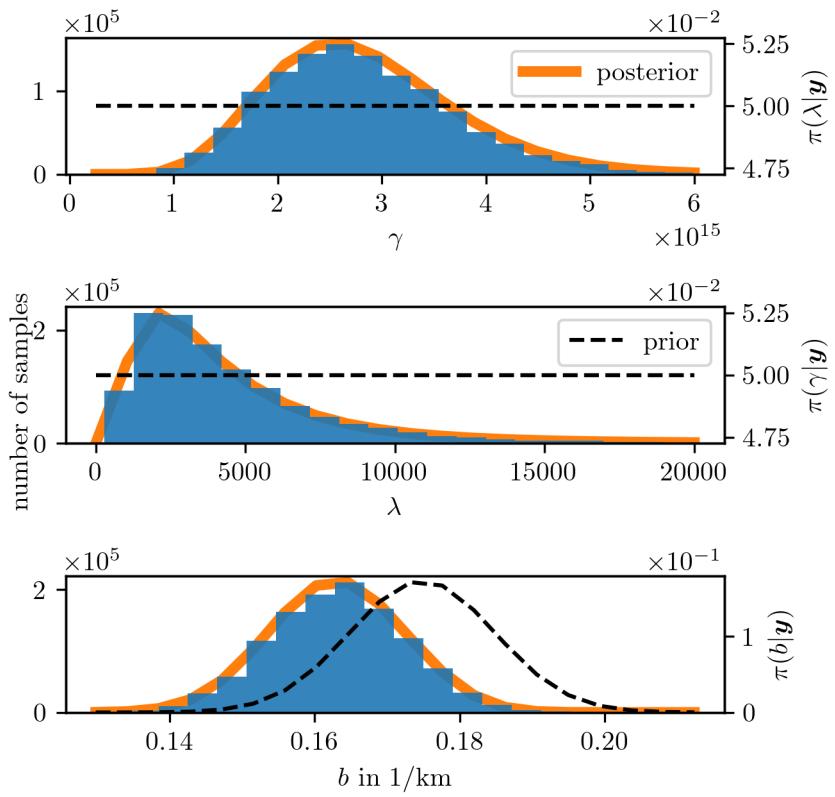


Figure 4.20: TT approximation of the marginal posterior in orange and the samples as a histogram as well as the prior distribution with a dotted line.

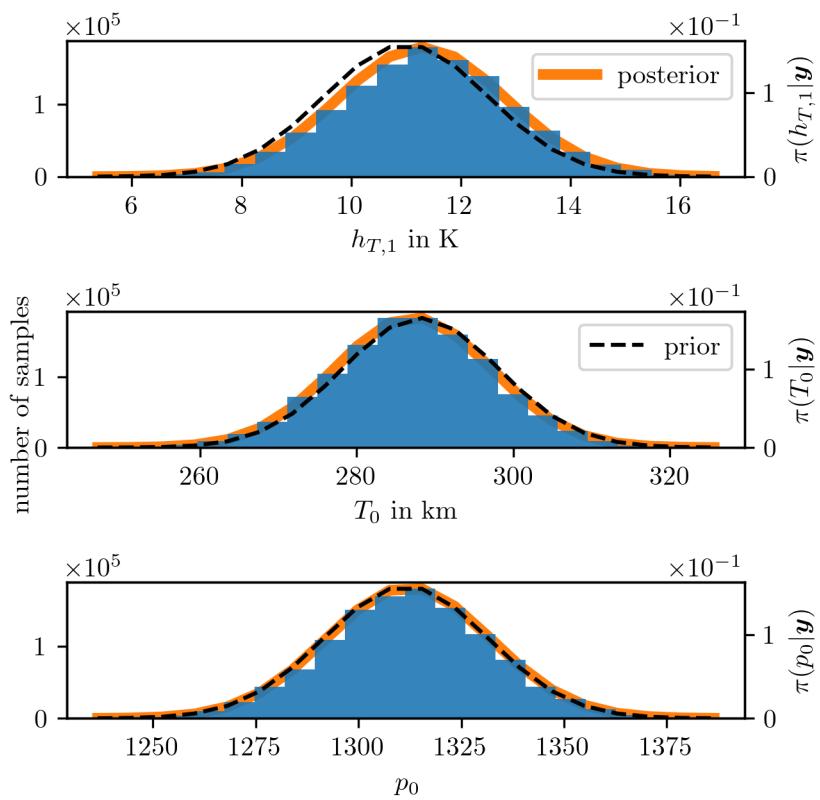


Figure 4.21: TT approximation of the marginal posterior in orange and the samples as a histogram as well as the prior distribution with a dotted line.

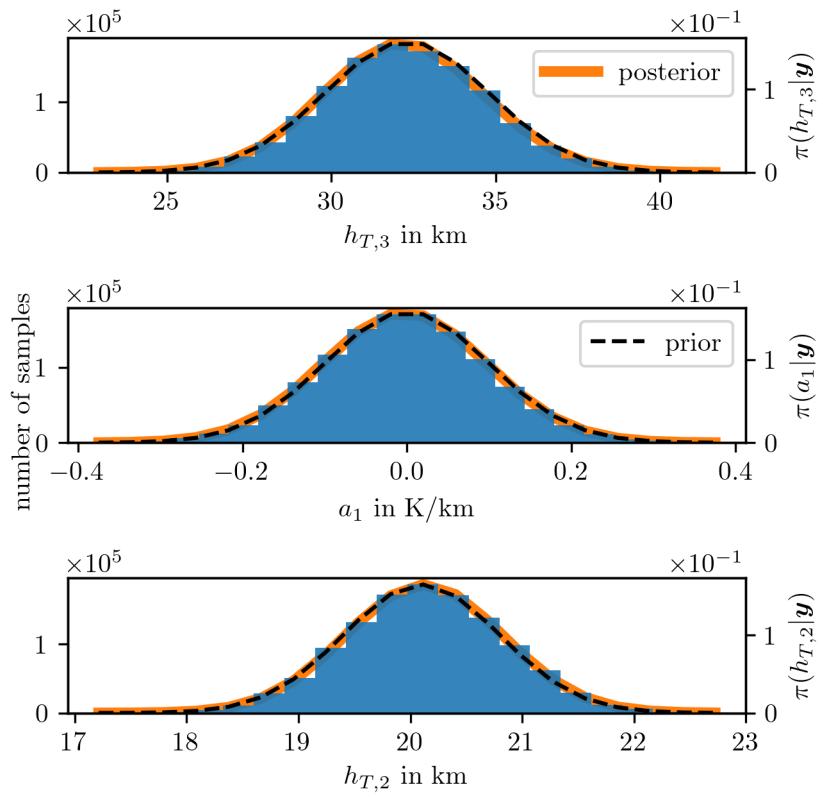


Figure 4.22: TT approximation of the marginal posterior in orange and the samples as a histogram as well as the prior distribution with a dotted line.

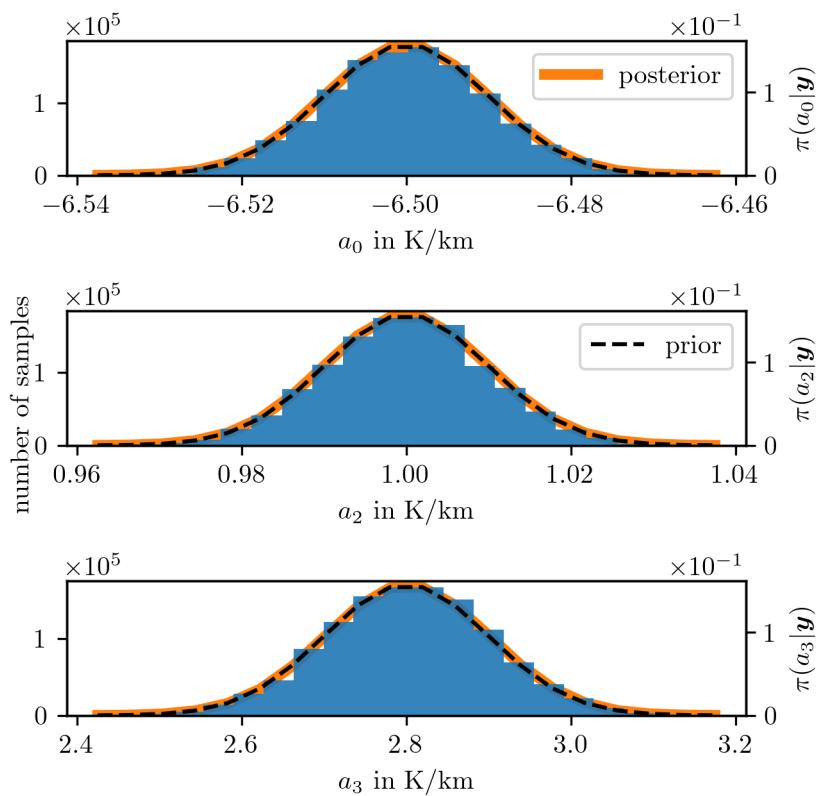


Figure 4.23: TT approximation of the marginal posterior in orange and the samples as a histogram as well as the prior distribution with a dotted line.

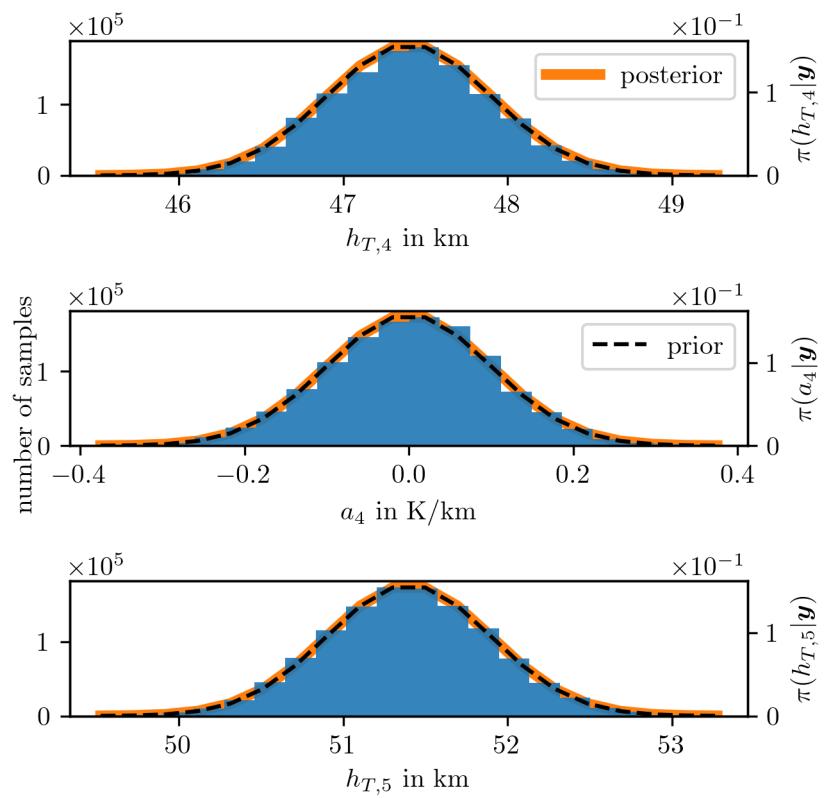


Figure 4.24: TT approximation of the marginal posterior in orange and the samples as a histogram as well as the prior distribution with a dotted line.

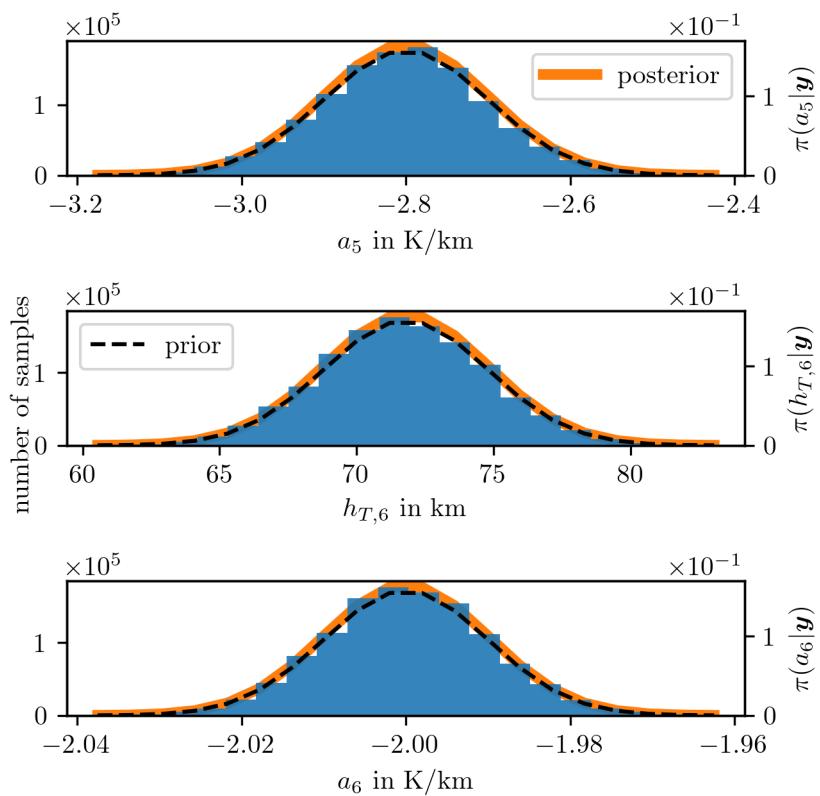


Figure 4.25: TT approximation of the marginal posterior in orange and the samples as a histogram as well as the prior distribution with a dotted line.

Conditional Posterior Distribution

We use the RTO method (see Sec. 2.2.3) to obtain ozone samples from the full conditional posterior and sample hyper-parameter samples directly from the marginal posterior to compute posterior temperature and pressure profiles according to their respective function (see Eq. 4.3 and Eq. 4.28). We plot posterior profiles of ozone in Fig. 4.28, temperature in Fig. 4.26 and pressure in Fig. 4.27.

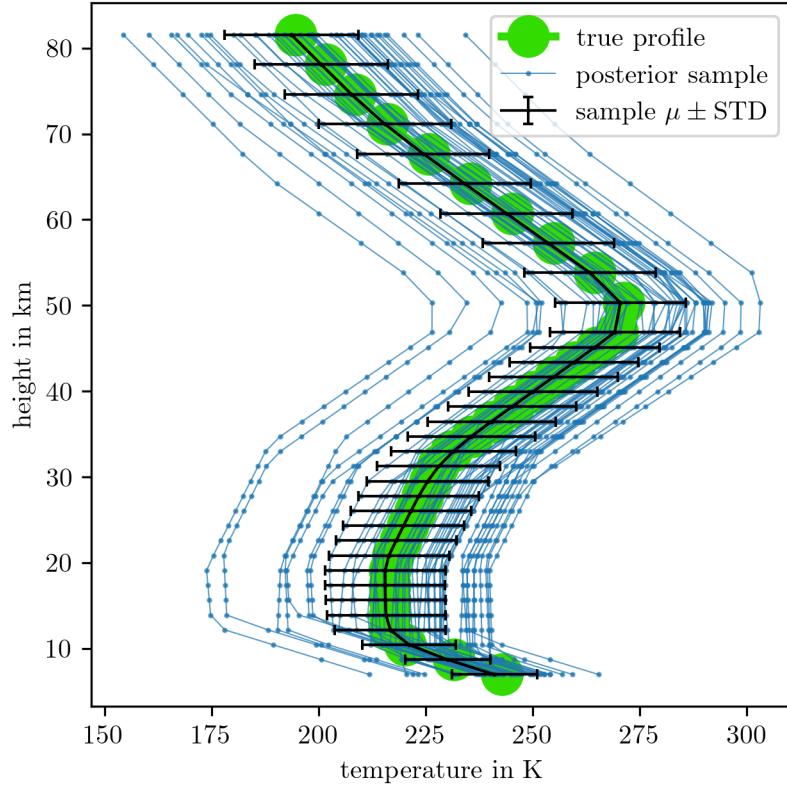


Figure 4.26: According to the hyper-parameter samples from the marginal posterior distribution (see Fig. 4.21 to Fig. 4.25) we plot the corresponding posterior temperature profile as given by Eq. 4.3.

We observe that pressure and ozone are highly correlated. Since the hyper-parameter b is smaller than its ground truth value, the posterior pressure profile does not exponentially decrease as strong as the ground truth pressure (see Fig. 4.20). This results in posterior pressure values which are slightly larger than the ground truth, causing posterior ozone values to be much smaller, but still with a similar structure, compared to the ground truth. Again, we are not able to recover the peak at large altitudes. The posterior temperature profiles look (as expected) similar to the prior temperature profiles.

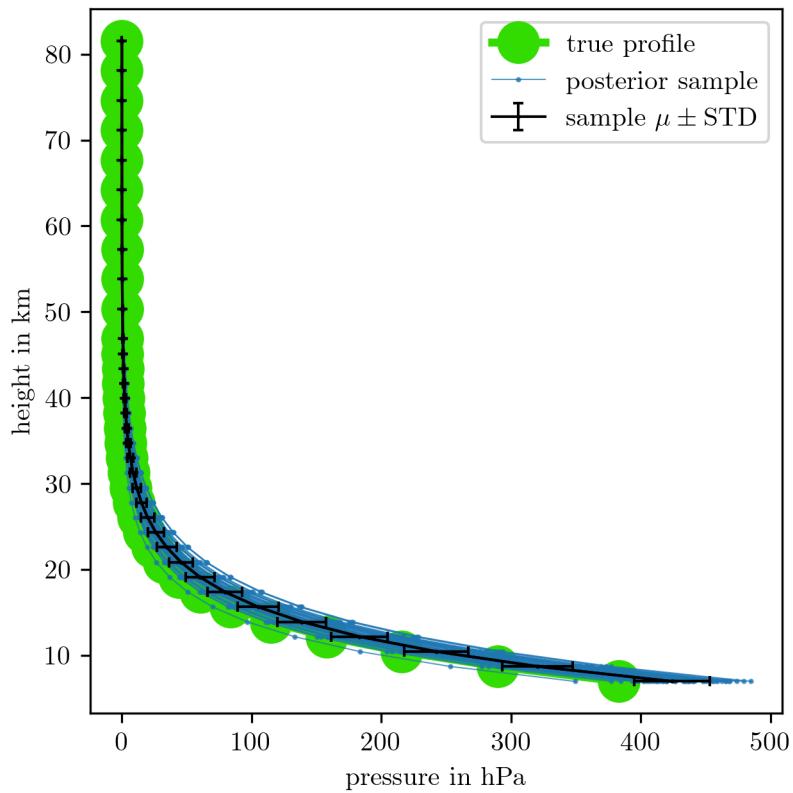


Figure 4.27: According to the hyper-parameter samples from the marginal posterior distribution (see Fig. 4.20 and Fig. 4.21) we plot the corresponding posterior pressure profile as given by Eq. 4.28.

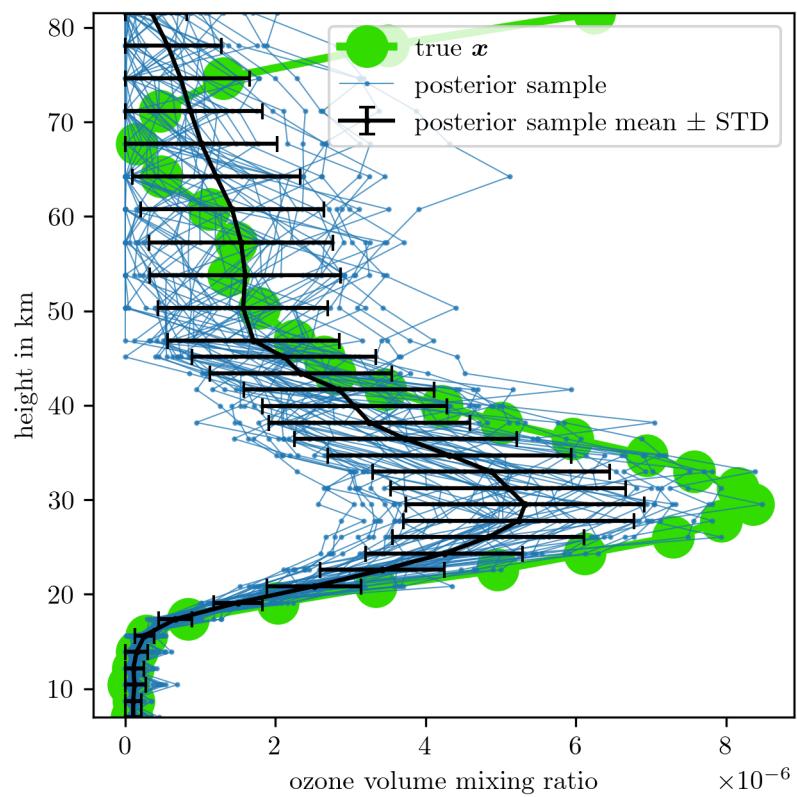


Figure 4.28: According to the hyper-parameter samples from the marginal posterior distribution (see Fig. 4.20) we plot the corresponding ozone sample from the conditional posterior using the RTO method.

5

Summary and Outlook

In this chapter, we summarise the key results and conclusions of our work and provide an outlook for future research. We compare the Bayesian approach to a regularisation approach and elaborate on the differences between sampling-based methods and the TT approximation. Lastly, we situate our results within the broader context of atmospheric modelling and discuss the implications for the future development of an atmospheric limb sounder.

5.1 Regularisation Solution vs. Hierarchical Bayesian Approach

Using a regularisation approach, we need 200 solves for \boldsymbol{x}_λ to obtain one solution of this inverse problem. In contrast, the hierarchical Bayesian approach involves generating around 10000 samples from the marginal posterior or 400 function evaluations to approximate $\pi(\lambda, \gamma | \mathbf{x})$, followed by 20 evaluations of \boldsymbol{x}_λ and \mathbf{B}_λ^{-1} to characterise the full posterior. Regardless, either method has a runtime of less than a second on a basic laptop.

While regularisation yields a single optimal solution, a Bayesian framework provides a distribution of ozone profiles, which are all feasible solutions to the inverse problem, and hence true errors. Moreover, within the hierarchical Bayesian approach, we can include prior knowledge about the noise, ozone profile and many more physical processes through hyper-parameters offering an arbitrarily flexible and informative inference framework.

5.2 Sampling Methods vs. TT Approximation

Using the TT approximation, requires far fewer function evaluations of the target distribution compared to sample-based methods, but requires to predefine a grid and a

normalisation constant, which, for now, we have to find iteratively. Relying solely on TT approximations may lead to a substantial amount of trial and error, and dealing with numerical issues. Nevertheless, once properly configured, we have shown the potential and advantages in computational cost of TT methods.

More specifically, the TT approximation of the 2 dimensional marginal is as fast as the MWG sampler (both within 1s). While the sampler takes 10000 steps, the TT approximation only needs 400 function evaluations; this is a factor of 25. Alternatively, for low-dimensional distributions, it may be preferable to approximate integrals directly using existing freely available quadrature libraries and packages such as `quadpy`.

In higher dimensions, such as the 18 dimensional marginal posterior considered in this thesis, TT methods outperform samplers like the `t-walk`, once a grid and normalisation constant have been defined. Although the `t-walk` may not be the best sampler for this specific marginal posterior and the underlying correlation structure, it is robust and easy to implement. To illustrate the efficiency of TT approximations we compare the number of function evaluations per 1000 independent samples. For 1000 independent samples with a maximum IACT of 550 and a burn-in period of 100 independent samples, the `t-walk` needs 1210000 function evaluations. In contrast, 24120 function evaluations are enough to approximate the marginal posterior in the TT format. Then drawing 1000 independent samples via the SIRT-MH scheme requires another 2000 function evaluations with an IACT of ≈ 0.6 . So the cost per independent sample for the `t-walk` is 1210 and for the TT approach is 27 (rounded up) function evaluations, including the burn-in period and TT approximation via `rect_cross.cross`. Or, after the burn-in phase, the `t-walk` requires around 1100 function evaluations per independent sample, while with an approximation of a probability density in the TT format, only 2 function evaluations are needed per independent sample.

For future application, we suggest improving the efficiency of the TT approximation by, e.g. reducing the correlation structure through a coordinate system rotation or using better interpolators in between grid points to reduce the approximation error. This may be particularly important when the CDF in the SIRT scheme is not smooth due to poor approximations of the target density at previous samples. Moreover, using a different reference measure for integration as in [3], such as a Gaussian measure instead of the current Lebesgue measure, may increase numerical stability. Currently, we have to predefined a normalisation constant and lower ranks manually, bounding the ranks automatically would be helpful (see e.g. [65]).

5.3 Atmospheric Physics

Here we summarise results within the context of our simplified physical atmospheric limb-sounding model. We demonstrated that the underlying non-linear forward model can be approximated with an affine map and the linear model, making this a linear inverse problem. For future application, we wish to include more measurement device-specific hyper-parameters in the forward model, e.g. uncertainty in pointing accuracy or an antenna response function.

In Sec. 3.1, we showed that we do not gain more information if we measure more frequently or collect more data in noise-dominated regions and that we need a SNR of $\approx 10^7$ to produce data, which is informative about ozone at higher altitudes.

Fig. 4.28 and Fig. 4.27 illustrate that pressure and ozone are highly correlated. One has to consider that when conditioning on pressure or inferring ozone and pressure jointly, as a slight change in pressure does skew the ozone VMR significantly. We could fix that by choosing a more restrictive prior for the pressure-related hyper-parameter b , but that would not be objective. By explanatory analysis, we found that data with a SNR of ≈ 10000 recovers an ozone (without a peak in higher altitudes) and pressure profile close to the ground truth. As previously mentioned in the prior analysis (see Fig. 4.17), the model as well as the data are uninformative about temperature and dominated by the exponentially decreasing pressure.

All the samples plotted in Fig. 4.11, Fig. 4.6, and Fig. 4.28 present valid solutions to the inverse problem, but consistently fail to capture the ozone peak at higher altitudes. This is due to noise-dominated data (see Fig. 4.1) and low signal strength in upper atmospheric regions, where the variability of the posterior ozone is large and primarily determined by the prior. For a more physical-based prior, e.g. a truncated multivariate Gaussian $\mathbf{x}|\delta \sim \mathcal{N}(\boldsymbol{\mu}_{\text{Trunc}}, \mathbf{Q}_{\text{Trunc}}^{-1}(\delta))$ with $\mathbf{a} \leq \mathbf{x} \leq \mathbf{b}$ between some truncation bounds \mathbf{a} and \mathbf{b} the joint multivariate Gaussian in Eq. 2.13 has the same form. In that case, one would have to replace the non-truncated mean and precision matrix with the truncated version as in [66, pp. 204-205]. Currently, we are not able to give analytic expression for the truncated mean and precision matrix, but [67] provide that numerically and [68] provides Python code to quickly sample from a truncated multivariate Gaussian. We conclude that the main objective for future research is to develop a more accurate (prior) model for ozone in the atmosphere, which captures physical properties and chemical processes of ozone.

References

- [1] C. Readings. *Envisat MIPAS An Instrument for Atmospheric Chemistry and Climate Research*. Noordwijk: ESA Publications Division, 2000.
- [2] Colin Fox and Richard A Norton. “Fast sampling in a linear-Gaussian inverse problem”. In: *SIAM/ASA Journal on Uncertainty Quantification* 4.1 (2016), pp. 1191–1218.
- [3] Tiangang Cui and Sergey Dolgov. “Deep composition of tensor-trains using squared inverse rosenblatt transports”. In: *Foundations of Computational Mathematics* 22.6 (2022), pp. 1863–1922.
- [4] J. Andrés Christen and Colin Fox. “A general purpose sampling algorithm for continuous distributions (the t-walk)”. In: *Bayesian Analysis* 5.2 (2010), pp. 263 –281.
- [5] Bryan Duncan. *Aura at 20 Years*.
<https://science.nasa.gov/science-research/earth-science/aura-at-20-years/>. [Online; accessed 31/08/25]. NASA’s Goddard Space Flight Center (GSFC), 2024.
- [6] Susan L Ustin and Elizabeth McPhee Middleton. “Current and near-term Earth-observing environmental satellites, their missions, characteristics, instruments, and applications”. In: *Sensors* 24.11 (2024), p. 3488.
- [7] Mallika Irene Suresh et al. “Multichannel upconversion of terahertz radiation in an optical disk resonator”. In: *Opt. Express* 33.5 (2025), pp. 10302–10311.
- [8] Florian Sedlmeir et al. “Detecting THz in the telecom range: All resonant THz up-conversion in a whispering gallery mode resonator”. In: *2014 39th International Conference on Infrared, Millimeter, and Terahertz waves (IRMMW-THz)*. 2014, pp. 1–2.
- [9] Clive D Rodgers. “Retrieval of atmospheric temperature and composition from remote measurements of thermal radiation”. In: *Reviews of Geophysics* 14.4 (1976), pp. 609–624.
- [10] Nathaniel J. Livesey et al. *Earth Observing System (EOS) Microwave Limb Sounder (MLS) Version 5.0x Level 2 and 3 data quality and description document*. Version 5.0-1.1a. NASA Goddard Earth Sciences Data and Information Services Center, 2022.
- [11] Sze M Tan, Colin Fox, and Geoff K. Nicholls. *Course notes for ELEC 445 – Inverse Problems and Imaging*.
<https://coursesupport.physics.otago.ac.nz/wiki/pmwiki.php/ELEC445/HomePage>. [Online; accessed 10/12/23]. Physics Department, University of Otago, 2016.
- [12] Nathaniel J Livesey et al. “Retrieval algorithms for the EOS Microwave limb sounder (MLS)”. In: *IEEE Transactions on Geoscience and Remote Sensing* 44.5 (2006), pp. 1144–1155.
- [13] Schwartz M. et al. *MLS/Aura Level 2 Ozone (O3) Mixing Ratio V005*.
https://disc.gsfc.nasa.gov/datasets/ML203_005/summary?keywords=mls%20o3. [Online; accessed 25/04/24]. NASA Goddard Earth Sciences Data and Information Services Center, 2020.
- [14] N. J. Livesey et al. “Validation of Aura Microwave Limb Sounder O3 and CO observations in the upper troposphere and lower stratosphere”. In: *Journal of Geophysical Research: Atmospheres* 113.D15 (2008).
- [15] L. Froidevaux et al. “Validation of Aura Microwave Limb Sounder stratospheric ozone measurements”. In: *Journal of Geophysical Research: Atmospheres* 113.D15 (2008).

- [16] Gareth O. Roberts and Jeffrey S Rosenthal. “General state space Markov chains and MCMC algorithms”. In: *Probability Surveys* 1 (2004), pp. 20–71.
- [17] S. P. Meyn and R.L. Tweedie. *Markov Chains and Stochastic Stability. 2nd Edition*. New York: Cambridge University Press, 2009.
- [18] Charles J Geyer. “Practical markov chain monte carlo”. In: *Statistical science* (1992), pp. 473–483.
- [19] U. Wolff and B. Bunk. *Lecture Notes on Computational Physics II [in german]*. www-com.physik.hu-berlin.de/comphys/comphys.htm. [Online; accessed 29/08/25]. Humboldt University, Berlin, 2016.
- [20] Ulli Wolff. “Monte Carlo errors with less errors”. In: *Computer Physics Communications* 156.2 (2004), pp. 143–153.
- [21] Drik Hesse. *py-uwerr; Python implementation of Monte Carlo error analysis a la Wolff*. <https://github.com/dhesse/py-uwerr>. [Online; accessed 09/09/25].
- [22] A. Sokal. “Monte Carlo Methods in Statistical Mechanics: Foundations and New Algorithms”. In: *Functional Integration: Basics and Applications*. Ed. by Cecile DeWitt-Morette, Pierre Cartier, and Antoine Folacci. Boston, MA: Springer US, 1997, pp. 131–192.
- [23] Havard Rue and Leonhard Held. *Gaussian Markov random fields: theory and applications*. London: CRC press, 2005.
- [24] Charles W. Champ and Andrew V. Sills. “The Generalized Law of Total Covariance”. In: *preprint* (2022).
- [25] Daniel Simpson, Finn Lindgren, and Håvard Rue. “Think continuous: Markovian Gaussian models in spatial statistics”. In: *Spatial Statistics* 1 (2012), pp. 16–29.
- [26] Dootika Vats et al. “Understanding Linchpin Variables in Markov Chain Monte Carlo”. In: *preprint* (2022). URL: <https://arxiv.org/pdf/2210.13574.pdf>.
- [27] Gareth O. Roberts and Jeffrey S Rosenthal. “Harris recurrence of Metropolis-within-Gibbs and trans-dimensional Markov chains”. In: *The Annals of Applied Probability* 16.4 (2006), 2123–2139.
- [28] Johnathan M Bardsley. “MCMC-based image reconstruction with uncertainty quantification”. In: *SIAM Journal on Scientific Computing* 34.3 (2012), A1316–A1332.
- [29] Johnathan M Bardsley et al. “Randomize-then-optimize: A method for sampling from posterior distributions in nonlinear inverse problems”. In: *SIAM Journal on Scientific Computing* 36.4 (2014), A1895–A1910.
- [30] John Bardsley and Tiangang Cui. “A Metropolis-Hastings-Within-Gibbs Sampler for Nonlinear Hierarchical-Bayesian Inverse Problems”. In: *2017 MATRIX Annals*. Vol. 2. MATRIX Book Series. Switzerland: Springer, 2019, pp. 2–12.
- [31] Philip J Davis and Philip Rabinowitz. *Methods of numerical integration*. San Diego, CA: Academic Press, Inc., 1984.
- [32] Sergey Dolgov et al. “Approximation and sampling of multivariate probability distributions in the tensor train decomposition”. In: *Statistics and Computing* 30 (2020), pp. 603–625.
- [33] Colin Fox et al. “Grid methods for Bayes-optimal continuous-discrete filtering and utilizing a functional tensor train representation”. In: *Inverse Problems in Science and Engineering* 29.8 (2021), pp. 1199–1217.
- [34] Ivan V Oseledets. “Tensor-train decomposition”. In: *SIAM Journal on Scientific Computing* 33.5 (2011), pp. 2295–2317.
- [35] Ivan Oseledets. “DMRG Approach to Fast Linear Algebra in the TT-Format”. In: *Computational Methods in Applied Mathematics* 11.3 (2011), pp. 382–393.

- [36] John Thickstun. *Kantorovich-rubinstein duality*. https://courses.cs.washington.edu/courses/cse599i/20au/resources/L12_duality.pdf. [Online; accessed 31/08/25]. University of Washington, 2019.
- [37] Luigi Ambrosio, Elia Brué, and Daniele Semola. *Lectures on Optimal Transport*. Cham: Springer Nature Switzerland, 2024.
- [38] Jean Feydy. “Analyse de données géométriques, au delà des convolutions”. Thesis. Université Paris-Saclay, July 2020. URL: <https://theses.hal.science/tel-02945979>.
- [39] Marcel Berger. *Geometry I. 4th Edition*. Berlin Heidelberg: Springer-Verlag, 2009.
- [40] Katsumi Nomizu and Takeshi Sasaki. *Affine differential geometry*. Cambridge: Cambridge University Press, 1994.
- [41] Per Christian Hansen. “Regularization, GSVD and truncated GSVD”. In: *BIT numerical mathematics* 29.3 (1989), pp. 491–504.
- [42] Per Christian Hansen. “The L-Curve and its Use in the Numerical Treatment of Inverse Problems”. English. In: *Computational Inverse Problems in Electrocardiology*. Ed. by P. Johnston. WIT Press, 2001, pp. 119–142.
- [43] Yu-Xiang Wang et al. “Trend Filtering on Graphs”. In: *Journal of Machine Learning Research* 17.105 (2016), pp. 1–41.
- [44] Per Christian Hansen and Dianne Prost O’Leary. “The use of the L-curve in the regularization of discrete ill-posed problems”. In: *SIAM Journal on Scientific Computing* 14.6 (1993), pp. 1487–1503.
- [45] KC Santosh, Nibaran Das, and Swarnendu Ghosh. “Chapter 3 - Deep learning models”. In: *Deep Learning Models for Medical Imaging*. Ed. by KC Santosh, Nibaran Das, and Swarnendu Ghosh. Primers in Biomedical Imaging Devices and Systems. Academic Press, 2022, pp. 65–97.
- [46] Jae N. Lee and Dong L. Wu. “Solar Cycle Modulation of Nighttime Ozone Near the Mesopause as Observed by MLS”. In: *Earth and Space Science* 7.4 (2020).
- [47] Iouli E Gordon et al. “The HITRAN2020 molecular spectroscopic database”. In: *Journal of Quantitative Spectroscopy and Radiative Transfer* 277 (2022), p. 107949.
- [48] Marie Šimečková et al. “Einstein A-coefficients and statistical weights for molecular absorption transitions in the HITRAN database”. In: *Journal of Quantitative Spectroscopy and Radiative Transfer* 98.1 (2006), pp. 130–155.
- [49] George B. Rybicki and Alan P. Lightman. *Radiative processes in Astrophysics*. Weinheim: Wiley-VCH, 2004.
- [50] W.G. Read et al. “The clear-sky unpolarized forward model for the EOS aura microwave limb sounder (MLS)”. In: *IEEE Transactions on Geoscience and Remote Sensing* 44.5 (2006), pp. 1367–1379.
- [51] Colin Fox. *Blokkurs on computing MCMC for inverse problems*. unpublished. Physics Department, University of Otago, 2025.
- [52] Australian National Concurrent Design Facility. *CubeSat Microwave Radiometer Mission to Support Global Ozone Layer Monitoring. Concept Study - Summary Report*. unpublished, internal report. Canberra BC: UNSW Canberra Space, 2023.
- [53] U.S. Standard Atmosphere, 1976. Washington, D.C.: United States. National Oceanic and Atmospheric Administration, United States Committee on Extension to the Standard Atmosphere, 1976.
- [54] Joe W Waters et al. “The earth observing system microwave limb sounder (EOS MLS) on the Aura satellite”. In: *IEEE Transactions on Geoscience and Remote Sensing* 44.5 (2006), pp. 1075–1092.

- [55] Gareth Roberts. *ST911 Fundamentals of Statistical Inference - Part III*. Department of Statistics, University of Warwick, 2015.
- [56] Ivan Oseledets and Eugene Tyrtyshnikov. "TT-cross approximation for multidimensional arrays". In: *Linear Algebra and its Applications* 432.1 (2010), pp. 70–88.
- [57] Sergey Dolgov and Robert Scheichl. "A Hybrid Alternating Least Squares–TT-Cross Algorithm for Parametric PDEs". In: *SIAM/ASA Journal on Uncertainty Quantification* 7.1 (2019), pp. 260–291.
- [58] I. V. Oseledets et al. *tpty - a Python implementation of the TT-toolbox*. <https://github.com/oseledets/tpty>. [accessed 23/06/25]. 2018.
- [59] Josef Dick, Frances Y. Kuo, and Ian H. Sloan. "High-dimensional integration: The quasi-Monte Carlo way". In: *Acta Numerica* 22 (2013), 133–288.
- [60] Per Christian Hansen. *Discrete Inverse Problems: Insight and Algorithms*. Philadelphia: SIAM, 2010.
- [61] Ville Satopää et al. "Finding a "Kneedle" in a Haystack: Detecting Knee Points in System Behavior". In: *2011 31st International Conference on Distributed Computing Systems Workshops*. IEEE. 2011, pp. 166–171.
- [62] Ulli Wolff. *UWerr.m Version6*. <https://www.physik.hu-berlin.de/de/com/ALPHAssoft>. [Online; accessed 5/11/23]. Humboldt-Universität to Berlin, 2004.
- [63] J. Andrés Christen and Colin Fox. *The t-walk software*. <https://www.cimat.mx/~jac/twalk/>. [Online; accessed 25/11/24]. CIMAT, Mexico, and University of Otago, New Zealand.
- [64] Jean Feydy. *GeomLoss – Geometric Loss functions between sampled measures, images and volumes*. <https://www.kernel-operations.io/geomloss/api/pytorch-api.html>. [Online; accessed 12/09/25].
- [65] Paul B. Rohrbach et al. "Rank Bounds for Approximating Gaussian Densities in the Tensor-Train Format". In: *SIAM/ASA Journal on Uncertainty Quantification* 10.3 (2022), pp. 1191–1224.
- [66] Samuel Kotz, N. Balakrishnan, and Norman L. Johnson. *Computational methods for inverse problems*. New York: John Wiley and Sons, Inc, 2000.
- [67] B. G. Manjunath and Stefan Wilhelm. "Moments Calculation for the Doubly Truncated Multivariate Normal Density". In: *Journal of Behavioral Data Science* 1.1 (2021), 17–33.
- [68] Z. I. Botev. "The Normal Law Under Linear Restrictions: Simulation and Estimation via Minimax Tilting". In: *Journal of the Royal Statistical Society Series B: Statistical Methodology* 79.1 (2016), pp. 125–148.
- [69] M. Capiński and P.E. Kopp. *Measure, Integral and Probability*. Springer Undergraduate Mathematics Series. London: Springer-Verlag London, 2004.
- [70] M. Simonnet. *Measures and Probabilities*. New York: Springer-Verlag, 1996.
- [71] Vesa Kaarnioja. *Inverse Problems. Eighth lecture*. <https://vesak90.userpage.fu-berlin.de/ip23/week8.pdf>. [Online; accessed 10/04/25]. Freie Universität Berlin, Department of Mathematics and Computer Science, 2023.
- [72] Greg Lawler. *Notes on probability*. <https://www.math.uchicago.edu/~lawler/probnotes.pdf>. [Online; accessed 10/04/25]. The University of Chicago, Department of Mathematics, 2016.

Appendices

A

Theoretical and technical background

A.1 Correlation Structure

In the book Gaussian Markov Random Fields [23], Rue and Held demonstrate that a strong correlation between the hyper-parameter μ and the latent field \mathbf{x} can significantly slow down convergence when using samplers, in particular Gibbs samplers. They consider the hierarchical model

$$\mu \sim \mathcal{N}(0, 1) \quad (\text{A.1a})$$

$$\mathbf{x}|\mu \sim \mathcal{N}(\mu \mathbf{1}, \mathbf{Q}^{-1}), \quad (\text{A.1b})$$

and apply a Gibbs sampler based on the full conditional distributions

$$\mu^{(k)} | \mathbf{x}^{(k)} \sim \mathcal{N}\left(\frac{\mathbf{1}^T \mathbf{Q} \mathbf{x}^{(k-1)}}{1 + \mathbf{1}^T \mathbf{Q} \mathbf{1}}, \left(1 + \mathbf{1}^T \mathbf{Q} \mathbf{1}\right)^{-1}\right) \quad (\text{A.2})$$

$$\mathbf{x}^{(k)} | \mu^{(k)} \sim \mathcal{N}(\mu^{(k)} \mathbf{1}, \mathbf{Q}^{-1}). \quad (\text{A.3})$$

As illustrated in Figure A.1, when the sampler is restricted to steps only in the μ -direction (horizontal axis) or the \mathbf{x} -direction (vertical axis), it requires many iterations to adequately explore the parameter space. This inefficiency arises from the high correlation between μ and \mathbf{x} , visible in Figure A.1 as a 'squeeze' of the distribution.

A solution to the slow mixing problem is to update (μ, \mathbf{x}) jointly. Since μ is one dimensional, effectively only the marginal density of μ is needed.

$$\mu^* \sim q(\mu^* | \mu^{(k-1)}) \quad (\text{A.4})$$

$$\mathbf{x}^{(k)} | \mu^* \sim \mathcal{N}(\mu^* \mathbf{1}, \mathbf{Q}^{-1}) \quad (\text{A.5})$$

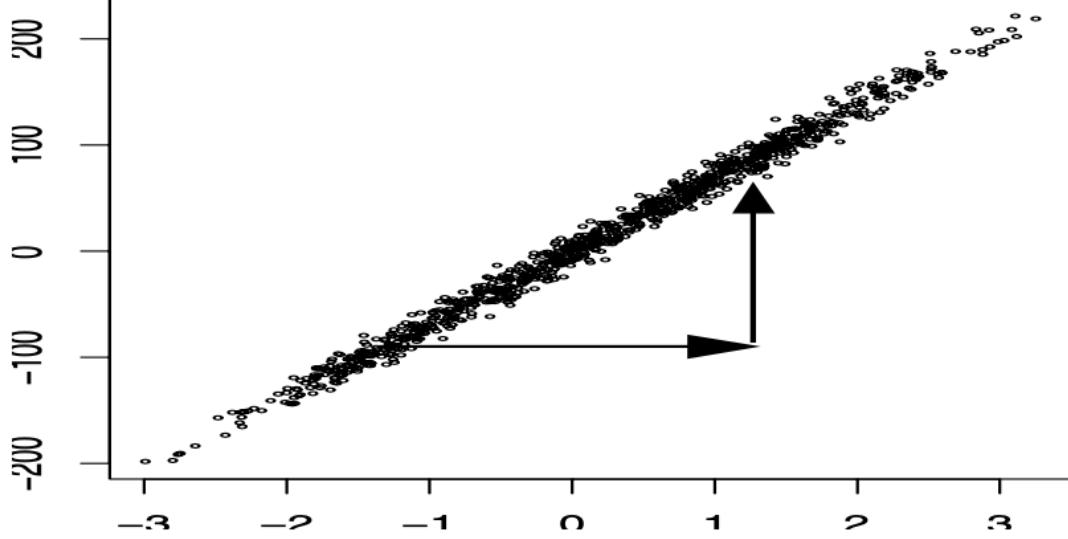


Figure A.1: The figure taken from [23, Figure 4.1 (b)], shows samples from the chain of μ (x-axis) and $\mathbf{1}^T \mathbf{Q} \mathbf{x}^{(k)}$ (y-axis) for over 1000 iterations, based on the hierarchical model in Eq. A.1, with an autoregressive process encoded in \mathbf{Q} . The algorithm updates μ and \mathbf{x} successively from their full conditional distributions (see Eq. A.2 and Eq. A.3). The plot displays $(\mu^{(k)}, \mathbf{1}^T \mathbf{Q} \mathbf{x}^{(k)})$, with $\mu^{(k)}$ on the horizontal axis and $\mathbf{1}^T \mathbf{Q} \mathbf{x}^{(k)}$ on the vertical axis. The slow mixing and convergence of μ result from its strong dependence on $\mathbf{1}^T \mathbf{Q} \mathbf{x}^{(k)}$, while the sampler permits only axis-aligned (horizontal and vertical) and does not allow diagonal moves, as illustrated by the arrows.

With a simple MCMC algorithm targeting μ one can explore the sample space efficiently and only draw a corresponding sample for \mathbf{x} from its full conditional once, for instance, the proposal μ^* has been accepted.

A.2 On the Monte-Carlo Error and Integrated Autocorrelation time

To assess the error $(\sigma^{(i)})^2$ of a samples-based estimate

$$\boldsymbol{\mu}^{(i)} := \text{E}_{\mathbf{x}|\mathbf{y}}[h(\mathbf{x})] = \frac{1}{N} \sum_{k=1}^N h(\mathbf{x}^{(k)}), \quad (\text{A.6})$$

from the chain $\mathcal{M}^{(i)} = \{\mathbf{x}^{(1)}, \dots, \mathbf{x}^{(k)}, \dots, \mathbf{x}^{(s)}, \dots, \mathbf{x}^{(N)}\} \sim \pi(\mathbf{x}|\mathbf{y})$, we ignore systematic error due to initialisation bias (burn-in period), but we have to take into account that samples produced by any system or algorithm are correlated. To derive the IACT, we follow Ulli Wolff's lecture notes [19] (or alternatively [20]).

In general, the error of a Monte-Carlo estimate is:

$$(\sigma^{(i)})^2 = \text{var}(\boldsymbol{\mu}^{(i)}) = \text{var}(\text{E}_{\mathbf{x}|\mathbf{y}}[h(\mathbf{x})]) = \left(\frac{1}{N} \sum_{k=1}^N h(\mathbf{x}^{(k)}) - \boldsymbol{\mu}^{(i)} \right)^2. \quad (\text{A.7})$$

Expanding this summation, we see that

$$(\sigma^{(i)})^2 = \frac{1}{N^2} \sum_{k,s=1}^N \Gamma(k-s) \quad (\text{A.8})$$

with the auto correlation coefficient $\Gamma(k-s) = (h(\mathbf{x}^{(k)}) - \boldsymbol{\mu}^{(i)})(h(\mathbf{x}^{(s)}) - \boldsymbol{\mu}^{(i)})$. Next we rewrite

$$\sum_{k,s=1}^N \Gamma(k-s) = \text{var}(h(\mathbf{x})) \sum_{k,s=1}^N \frac{\Gamma(k-s)}{\Gamma(0)} = \text{var}(h(\mathbf{x})) \sum_{k,s=1}^N \rho(k-s), \quad (\text{A.9})$$

with the normalised auto correlation coefficient $\rho(k-s) = \Gamma(k-s)/\Gamma(0)$ at lag $k-s$ and $\Gamma(0) = \text{var}(h(\mathbf{x}))$ for $k=s$. Typically $\Gamma(t)$ decays exponentially so that, for $N \gg \tau$, $\Gamma(t) \xrightarrow{t \rightarrow \infty} \exp\{-|t|/\tau\}$ and we can approximate

$$\sum_{k,s=1}^N \rho(k-s) = N \sum_{t=-(N-1)}^{N-1} \left(1 - \frac{t}{N}\right) \rho(t) \approx N \sum_{t=-\infty}^{\infty} \rho(t), \quad (\text{A.10})$$

see [22, p. 137], and define the IACT as in [19, pp. 103-105]. If $\tau \gg 1$

$$\sum_{t=-\infty}^{\infty} \rho(t) = 1 + 2 \sum_{t=1}^{\infty} (e^{-1/\tau})^t = 1 + 2 \frac{e^{-1/\tau}}{1 - e^{-1/\tau}} \approx 1 + 2 \frac{1 - 1/\tau}{1/\tau} = 2\tau - 1 \approx 2\tau_{int}, \quad (\text{A.11})$$

so that $\tau_{int} \approx \tau$. Here we use the geometric power series $\sum_{n=0}^{\infty} x^n = 1/(1-x)$ and the Taylor series $e^x \approx 1+x$ for small x . In practise, the estimate for the Monte-Carlo error is:

$$(\sigma^{(i)})^2 \approx \frac{\text{var}(h(\mathbf{x}))}{N} \sum_{t=-\infty}^{\infty} \rho(t) \approx \frac{\text{var}(h(\mathbf{x}))}{N} \underbrace{\left(1 + 2 \sum_{t=1}^W \rho(t)\right)}_{:=2\tau_{int}} = \text{var}(h(\mathbf{x})) \frac{2\tau_{int}}{N}, \quad (\text{A.12})$$

where W is the summation window. The IACT provides a good estimate of how many steps the sampling algorithm needs to take to produce one independent sample. More specifically, the effective sample size $\frac{2\tau_{int}}{N}$ gives an estimate of how efficient a sampler is.

A.3 Measure theory

Assume that the triple $(\Omega, \mathcal{F}, \mathbb{P})$ defines a probability space, where Ω denotes the complete sample space, \mathcal{F} is a σ -algebra consisting of a collection of countable subsets $\{A_n\}_{n \in \mathbb{N}}$ with $A_n \subseteq \Omega$, and \mathbb{P} is a probability measure on \mathcal{F} . The formal conditions for \mathbb{P} to be a probability measure, and for \mathcal{F} to be a σ -algebra over Ω , are given in Appendix A.3. We denote

$$\mathbb{P}(A) = \int_A d\mathbb{P} \quad (\text{A.13})$$

as the probability of an event $A \in \mathcal{F}$. By applying the Radon-Nikodym theorem [69], we can change variables

$$\mathbb{P}(A) = \int_A \frac{d\mathbb{P}}{dx} dx = \int_A \pi(x) dx, \quad (\text{A.14})$$

where dx is a reference measure on the same probability space, commonly referred to as the Lebesgue measure. The Radon-Nikodym derivative $\frac{d\mathbb{P}}{dx}$ of \mathbb{P} with respect to x is often interpreted as the probability density function (PDF) $\pi(x)$. Thus, we say that \mathbb{P} has a density $\pi(x)$ with respect to x [70, Chapter 10].

Now, let $X : \Omega \rightarrow \mathbb{R}^d$ be a d -dimensional random variable mapping from the probability space $(\Omega, \mathcal{F}, \mathbb{P})$ to the measurable space $(\mathbb{R}^d, \mathcal{X})$, where \mathcal{X} is a collection of subsets in \mathbb{R}^d [71]. Then the associated PDF $\pi(x)$ is a joint density of X , induced by the probability measure on Ω [69, 71].

Recall the probability space $(\Omega, \mathcal{F}, \mathbb{P})$, where Ω denotes the sample space, and \mathcal{F} is a collection of countable subsets $\{A_n\}_{n \in \mathbb{N}}$ of Ω . Each $A_n \subseteq \Omega$ is called an event, and a map $\mathbb{P} : \mathcal{F} \rightarrow \mathbb{R}$ is referred to as a measure. In the following, we describe the conditions required for \mathcal{F} to be a σ -algebra, and for \mathbb{P} to qualify as a probability measure. We refer to [72] [69] for further reading.

A.3.1 Probability Measure

For a probability measure, we require:

- $\mathbb{P}(\Omega) = 1$ and $\mathbb{P}(\emptyset) = 0$
- $\mathbb{P}(A) \in [0, 1]$
- $\mathbb{P}(\bigcup_{j \in \mathbb{N}} A_j) = \sum_{j \in \mathbb{N}} \mathbb{P}(A_j)$ if we have pairwise disjoint sets or $A_i \cap A_j = \emptyset$ for $i \neq j$

In other words, the probability assigned to the entire sample space must be equal to one, $\mathbb{P}(\Omega) = 1$, and the probability of the empty set must be zero, $\mathbb{P}(\emptyset) = 0$. For any subset $A \subseteq \Omega$, the probability $\mathbb{P}(A)$ must lie between zero and one, i.e., $\mathbb{P}(A) \in [0, 1]$. If e.g. two subsets A and B are disjoint (i.e., $A \cap B = \emptyset$), then the probability of their union satisfies $\mathbb{P}(A \cup B) = \mathbb{P}(A) + \mathbb{P}(B)$. This property must also hold for a countable sequence of disjoint sets $\{A_j\}_{j \in \mathbb{N}}$, such that $\mathbb{P}\left(\bigcup_{j \in \mathbb{N}} A_j\right) = \sum_{j \in \mathbb{N}} \mathbb{P}(A_j)$.

A.3.2 σ -Algebra

A collections of subsets \mathcal{F} is called σ -algebra if:

- $\emptyset, \Omega \in \mathcal{F}$,
- if $A \in \mathcal{F}$ then $A^C := A/\Omega \in \mathcal{F}$
- if $A_1, A_2, \dots \in \mathcal{F}$ then $\bigcup_{j \in \mathbb{N}} A_j \in \mathcal{F}$

In other words, the empty set \emptyset and the entire sample space Ω must always be elements of \mathcal{F} . If a set $A \in \mathcal{F}$, then its complement $A^C = \Omega \setminus A$ must also be in \mathcal{F} . If, in terms of a probability measure, we are able to assign a probability $\mathbb{P}(A)$ to an event A , we must also be able to assign a probability to the event “not A ”, i.e., $\mathbb{P}(A^C)$. Finally, if a countable collection of sets $A_1, A_2, \dots \in \mathcal{F}$, then their union $\bigcup_{j \in \mathbb{N}} A_j$ must also be in \mathcal{F} . These three properties define the requirements for \mathcal{F} to be a σ -algebra.

A.4 Python Code

18.0pt plus 2.0pt minus 1.0pt

```

1 def MargBack(TTCore, univarGrid):
2     ''' Backward marginalisation (see Prop. 1) as in SIRT from Cui et al. [3] '''
3
4     dim = len(univarGrid)
5     B = dim * [None] # coeffTensor
6     B[-1] = TTCore[-1]
7     R = [None] * dim
8     C = [None] * dim
9
10    for k in range(dim - 1, 0, -1):
11        r_kmin1, n, r_k = np.shape(TTCore[k])
12        # Eq. 2.35, [3, Eq. 22] !! we set Lebesgue Measure to const = one
13        M = np.identity(n) * (univarGrid[k][-1] - univarGrid[k][0]) # Mass matrix
14        L = scy.linalg.cholesky(M)
15
16        # construct Tensor C Eq. 2.36, [3, Eq. 27]
17        C[k] = np.zeros((r_kmin1, n, r_k))
18        for alpha in range(0, r_kmin1):
19            for l in range(0, r_k):
20                C[k][alpha, :, l] = B[k][alpha, :, l] @ L[:, :]
21
22        # unfold along first coordinate and compute thin QR decomposition of C^T
23        # Eq. 2.37, [3, Eq. 28]
24        Q, R[k] = np.linalg.qr(C[k].reshape((r_kmin1, n * r_k), order='C').transpose(), mode='reduced')
25
26        # compute next coefficient tensor Eq. 2.38, [3, Eq. 29]
27        r_kmin2, n, r_kmin1 = np.shape(TTCore[k - 1])
28        B[k - 1] = np.zeros(np.shape(TTCore[k - 1]))
29        for alpha_2 in range(0, r_kmin2):
30            for l_1 in range(0, r_kmin1):
31                B[k - 1][alpha_2, :, l_1] = TTCore[k - 1][alpha_2, :, :] @ R[k][l_1, :]
32
33 return B

```

Listing A.1: Pyhton code to calculate Backward marginals, as in Prop. 1 and [3].

```

1 def MargForw(TTCore, univarGrid):
2     ''' Forward marginalisation (see Prop. 2)
3         similar to backward marginalisation as in Cui et al. [3] '''
4
5     # compute pre marginal coefficients starting at dim = 1, k = 0
6     BPre = dim * [None] # coeffTensor
7     LebLam = 1 # !! Lebesgue Measure
8     BPre[0] = TTCore[0]
9     RPre = [None] * dim
10    CPre = [None] * dim
11
12    for k in range(0, dim-1):
13        r_kmin1, n, r_k = np.shape(TTCore[k])
14        # Eq. 2.35, [3, Eq. 22] !! we set Lebesgue Measure to const = one
15        M = np.identity(n) * (univarGrid[k][-1] - univarGrid[k][0]) # Mass matrix
16        L = scy.linalg.cholesky(M)
17
18        # construct Tensor C Eq. 2.39
19        CPre[k] = np.zeros((r_kmin1, n, r_k))
20        for alpha in range(0, r_kmin1):
21            for l in range(0, r_k):
22                CPre[k][alpha, :, l] = BPre[k][alpha, :, l] @ L[:, :]
23
24        # unfold along first coordinate and compute thin QR decomposition of C
25        # Eq. 2.40
26        Q, RPre[k] = np.linalg.qr(CPre[k].reshape((r_kmin1 * n, r_k)), order='C'), mode='reduced')
27
28        # compute next coefficient tensor Eq. 2.41
29        r_k, n, r_kpls1 = np.shape(TTCore[k + 1])
30        BPre[k + 1] = np.zeros(np.shape(TTCore[k + 1]))
31        for alpha_1 in range(0, r_kpls1):
32            for l_1 in range(0, r_k):
33                BPre[k + 1][l_1, :, alpha_1] = RPre[k][l_1, :] @ TTCore[k + 1][:, :, alpha_1]
34
35
36    return BPre
37

```

Listing A.2: Python code to calculate forward marginals, as in Prop. 2.

```

1 def SIRT(seeds, SQTT, univarGrid, BackMarg, absError):
2     ''' do squared inverse rosenblatt transform (SIRT) as in Cui et al. [3] '''
3
4     dim, numbSampl = seeds.shape
5     sampls = np.zeros(seeds.shape) # samples from approximated PDF
6     probVal = np.zeros(seeds.shape) # PDF values, for MH-correction step
7     Approx = np.zeros(seeds.shape[1]) # TT-Approx., to compare to true function
8
9     # Lebesgue measure for quadrature Eq. 2.27
10    WholeLebLam = np.zeros(dim)
11    for k in range(0, dim):
12        WholeLebLam[k] = (univarGrid[k][-1] - univarGrid[k][0])
13    lamX = np.ones(dim)
14    for k in range(1, dim):
15        lamX[k - 1] = np.prod(WholeLebLam[k:])
16
17
18    gamError = absError / np.prod(WholeLebLam) # error as in Eq. 2.32 [3, Eq. (35)]
19
20    # sample from first dimension [3, Eq. (30)]
21    firstMarg = gamError * lamX[0] + np.sum(BackMarg[0][0, :, :] ** 2, 1)
22    # cumulative distribution function, normalised numerically Eq. 2.46 [3, Eq. (17)]
23    firstCDF = np.cumsum(firstMarg / np.sum(firstMarg))
24    # draw samples as 'inverse transform'
25    sampls[0] = np.interp(seeds[0], firstCDF, univarGrid[0])
26    probVal[0] = np.interp(sampls[0], univarGrid[0], firstMarg / np.sum(firstMarg))
27
28    # sample from other dimensions
29    for n in range(0, numbSampl):
30        # piecew. poly. interpol in first dimension Eq. 2.48 [32]
31        CurrApprCore = LinInterPolTT(SQTT[0], univarGrid[0], sampls[0][n])
32        for d in range(1, dim):
33            # marginal function, conditioned on previous samples
34            rank_min, gridSize, rank_pls = BackMarg[d].shape
35            MargDep = np.zeros((BackMarg[d].shape))
36            for r in range(0, rank_min):
37                # condition on previous samples
38                MargDep[r, :, :] = CurrApprCore[0, r] * BackMarg[d][r, :, :]
39
40            currMarg = gamError * lamX[d] + np.sum(np.sum(np.copy(MargDep), axis=0)** 2,
41                axis=1) # Eq. 2.47 [3, Eq. (31)]
42
43            currCDF = np.cumsum(currMarg / np.sum(currMarg)) # Eq. 2.46 [3, Eq. (17)]
44
45            # draw sample as 'inverse transform'
46            sampls[d][n] = np.interp(seeds[d][n], currCDF, univarGrid[d])
47            probVal[d][n] = np.interp(sampls[d][n], univarGrid[d],
48                currMarg / np.sum(currMarg))
49            # piecew. poly. interpol., Eq. 2.48 [32], cond. on samp. for next PDF
50            CurrApprCore = np.copy(CurrApprCore) @ LinInterPolTT(SQTT[d], univarGrid[d],
51                sampls[d][n])
52
53            Approx[n] = gamError + CurrApprCore ** 2
54
55    return sampls, probVal, Approx

```

Listing A.3: Pyhton code to draw samples via SIRT, as in Alg. Box 3.

B

Additional Figures

B.1 Ozone

B.1.1 Ozone Prior

B.1.2 Integrated Autocorrelation plots

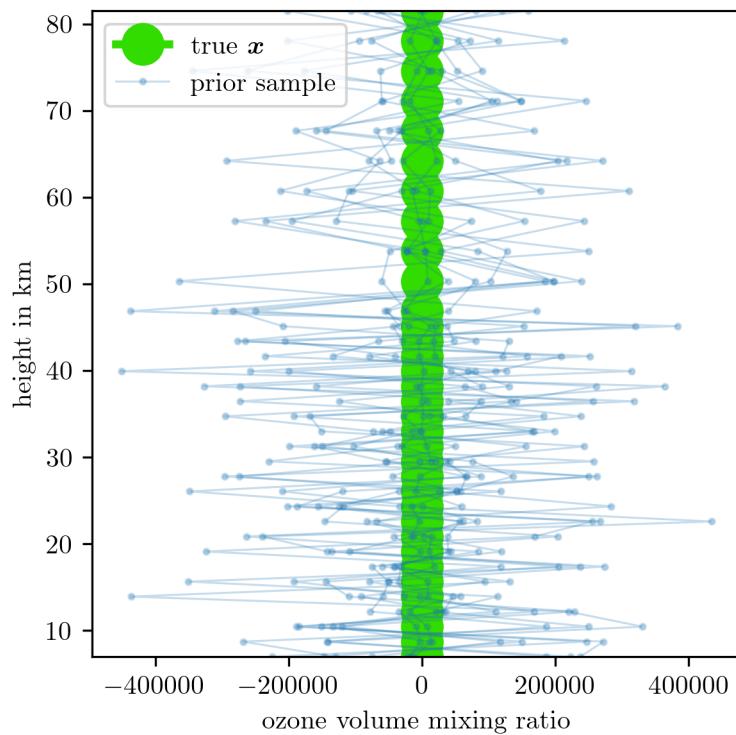


Figure B.1: We draw samples from ozone prior distribution $\mathbf{x} \sim \mathcal{N}(0, \delta \mathbf{L})$ after generating a sample from the hyper-prior distribution $\delta \sim \mathcal{T}(1, 10^{-10})$. Note that since the variance of prior samples is very large compared to the ozone volume mixing ratios, the ozone profile appears to be constant, which is not the case, see e.g. Fig. 4.6.

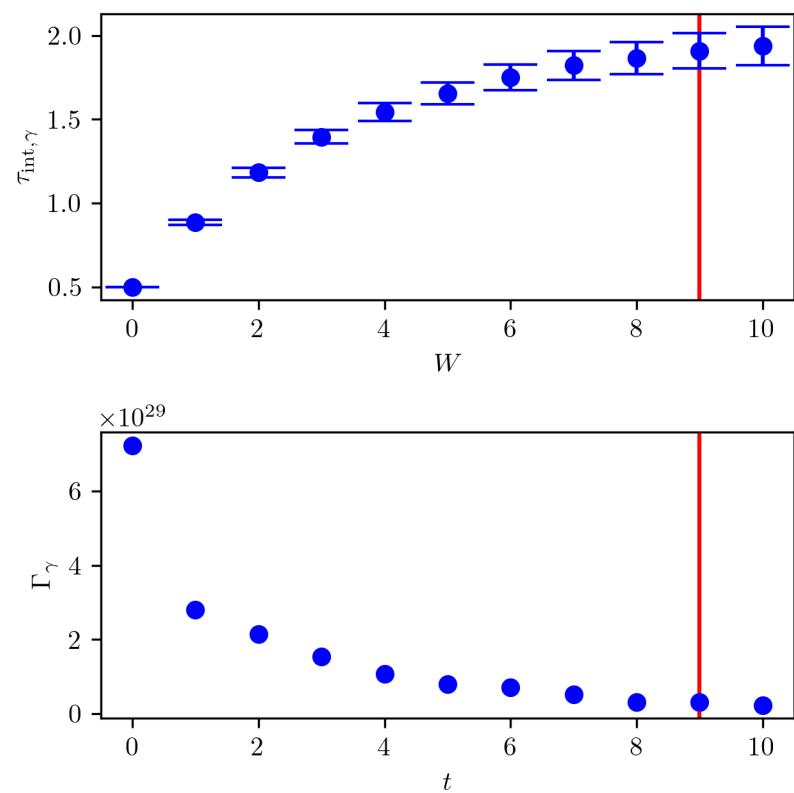


Figure B.2: Here the autocorrelation function Γ_γ at different lags W is plotted as well as the IATC $\tau_{\text{int},\gamma}$ for the samples from $\pi(\gamma, \lambda | \mathbf{y})$ based on the linear forward model.

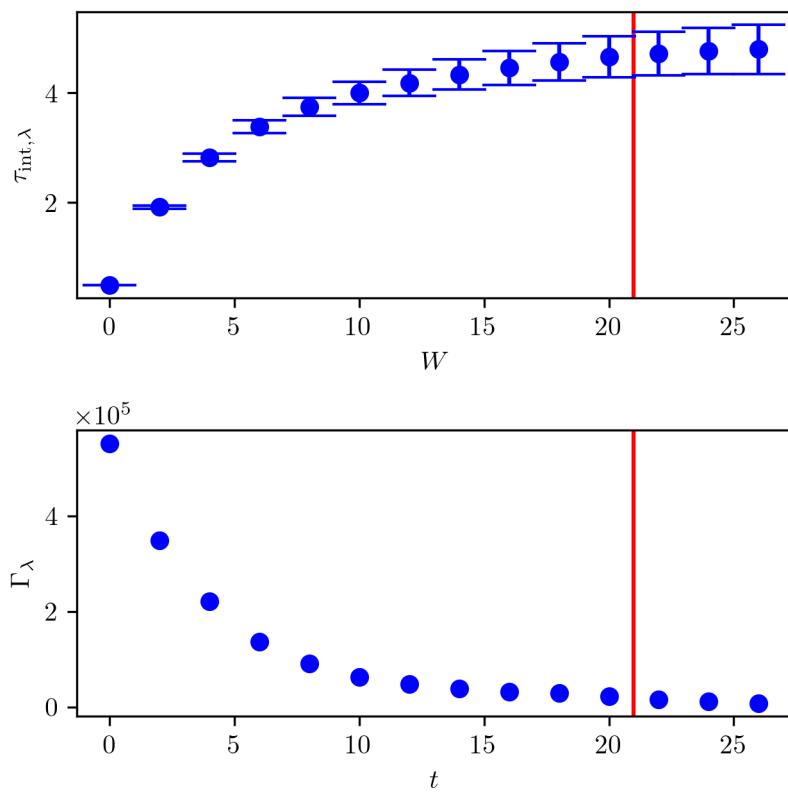


Figure B.3: IACT for samples $\lambda \sim \pi(\cdot|\gamma, \mathbf{y})$ based on the approximated forward model.

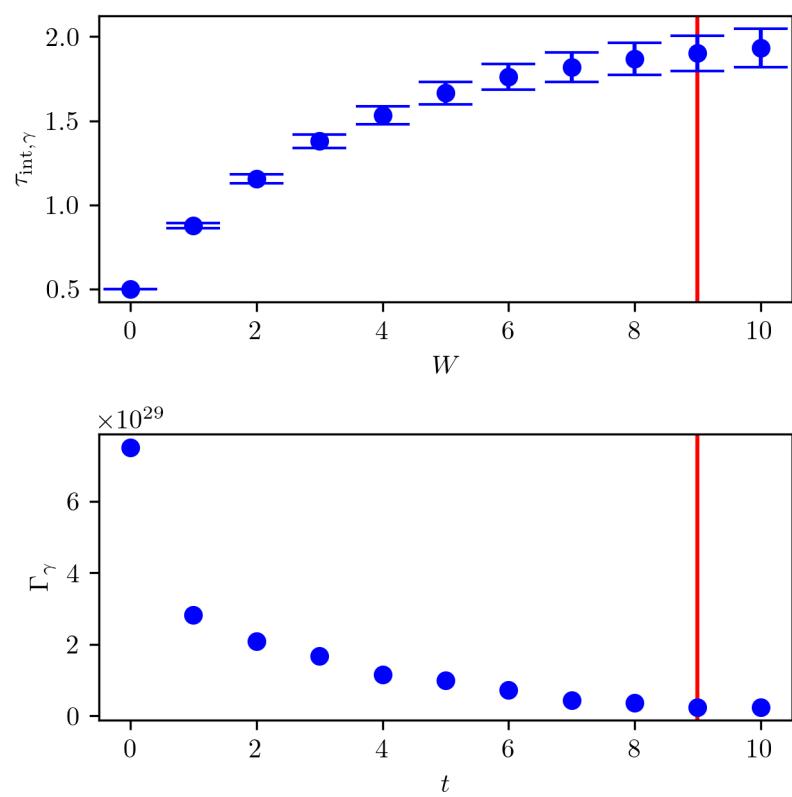


Figure B.4: IACT for samples $\gamma \sim \pi(\cdot | \lambda, \mathbf{y})$ based on the approximated forward model.

B.2 Pressure and Temperature

B.2.1 Priors

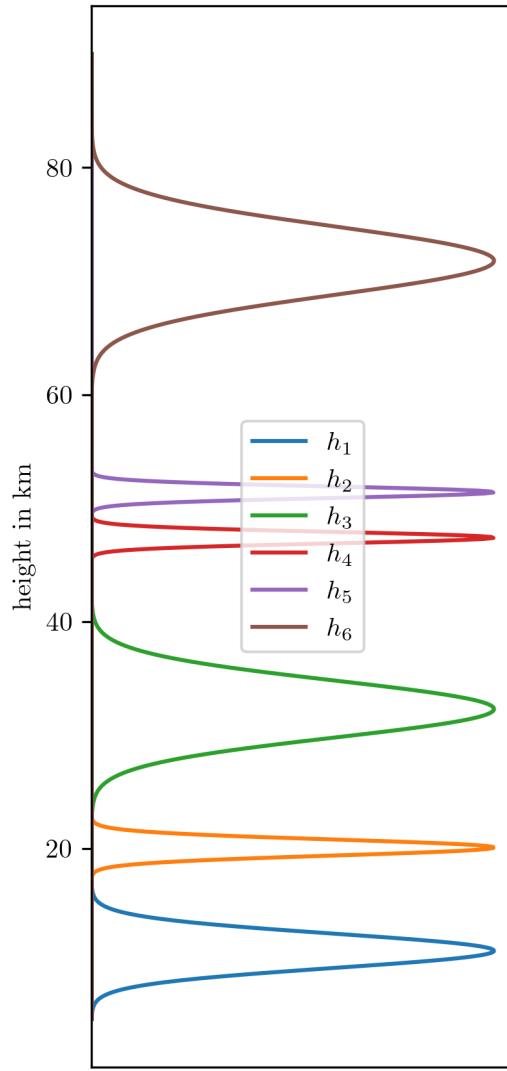


Figure B.5: Prior distributions $\pi(\mathbf{h}_T)$, which we choose so that they do not overlap and not conflict with the temperature function 4.3

B.2.2 T-walk Trace

B.2.3 Integrated Autocorrelation Plots

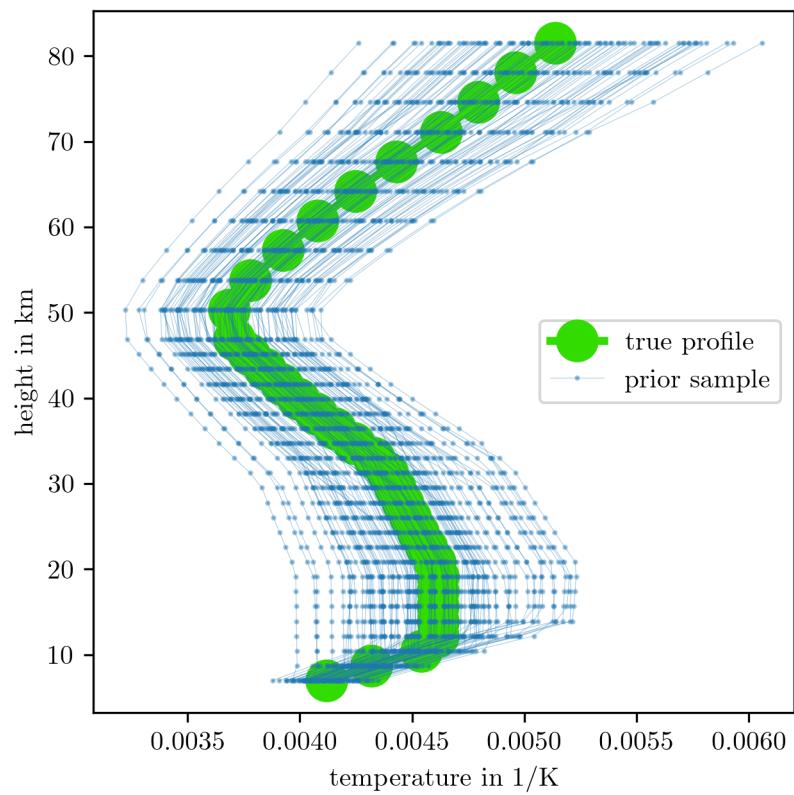


Figure B.6: Prior samples of the inverted temperature profile.

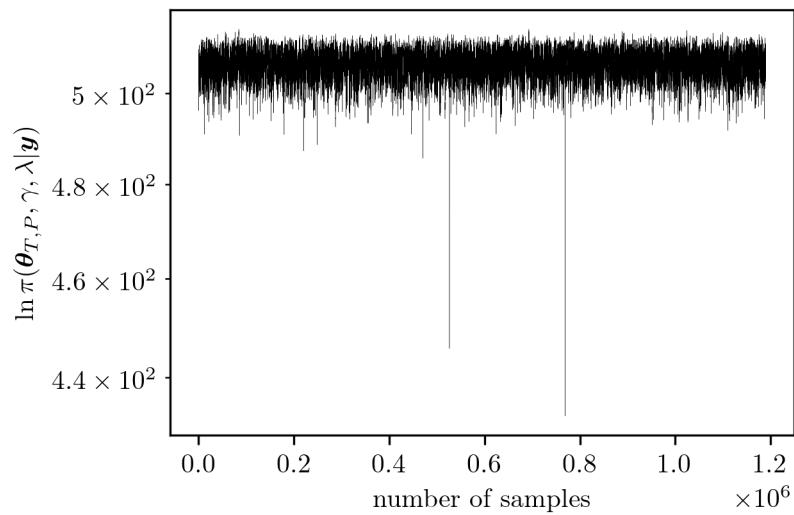


Figure B.7: Output trace of the t-walk on the posterior distribution $\pi(p_0, b, \mathbf{h}_T, \mathbf{a} | \gamma, \mathbf{y})$.

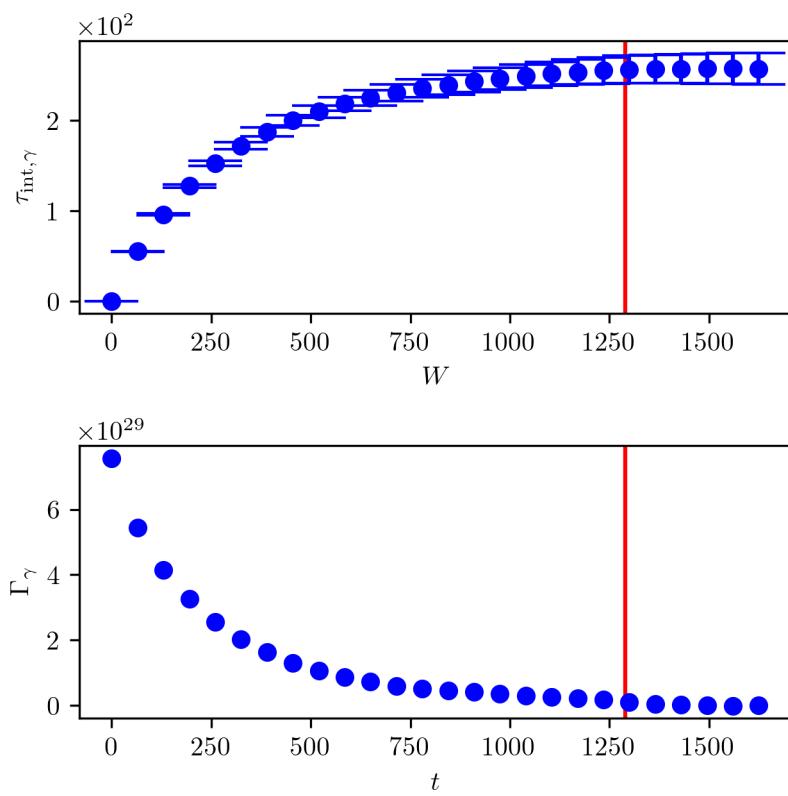


Figure B.8: IACT and autocorrelation function for samples $h_1 \sim \pi(\cdot | h_{T,2}, h_{T,3}, h_{T,4}, h_{T,5}, h_{T,6}, a_0, a_1, a_2, a_3, a_4, a_5, a_6, T_0, b, p_0, \mathbf{y})$

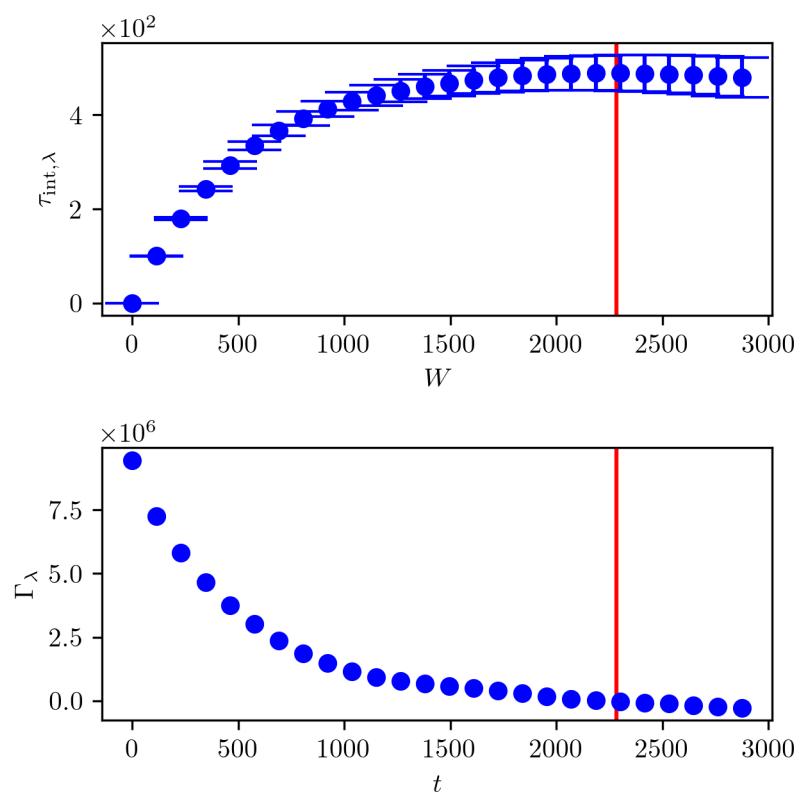


Figure B.9: IACT and autocorrelation function for samples $h_2 \sim \pi(\cdot | h_1, h_3, h_4, h_5, h_6, a_0, a_1, a_2, a_3, a_4, a_5, a_6, T_0, b, p_0, \mathbf{y})$

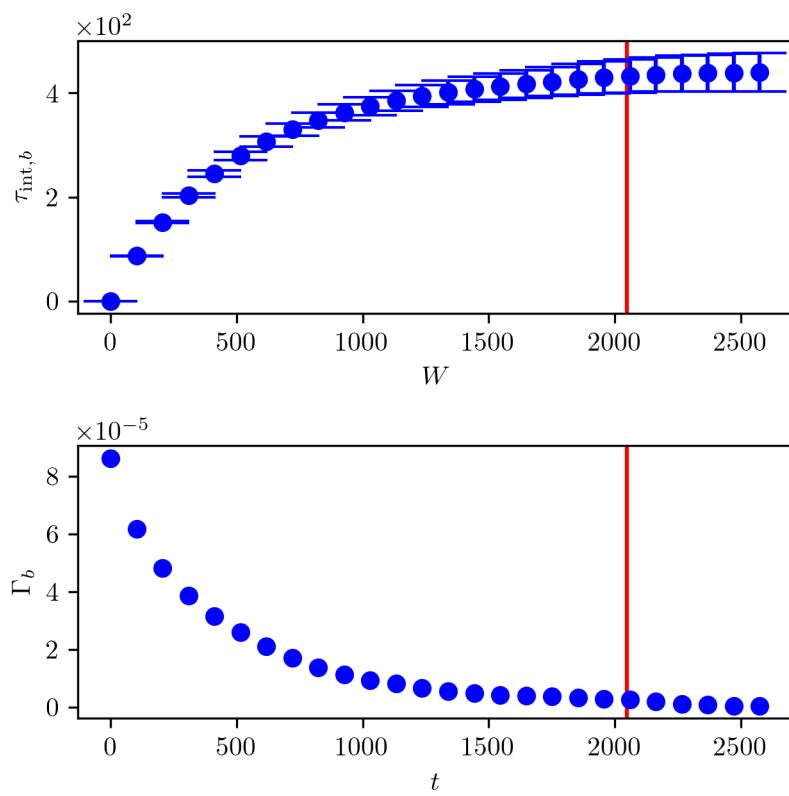


Figure B.10: IACT and autocorrelation function for samples $h_3 \sim \pi(\cdot | h_1, h_2, h_4, h_5, h_6, a_0, a_1, a_2, a_3, a_4, a_5, a_6, T_0, b, p_0, y)$

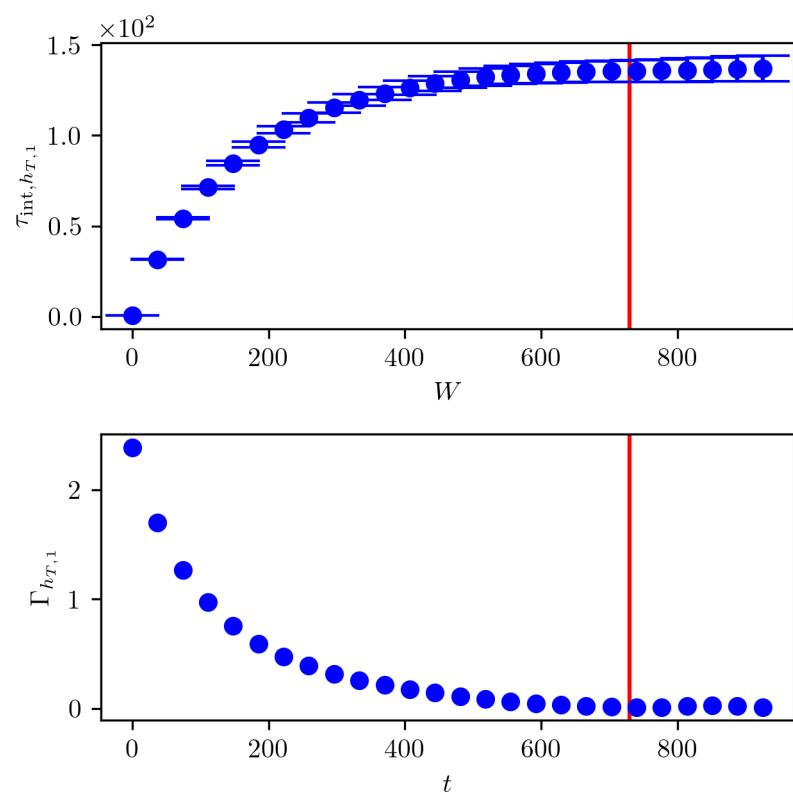


Figure B.11: IACT and autocorrelation function for samples $h_4 \sim \pi(\cdot | h_1, h_2, h_3, h_5, h_6, a_0, a_1, a_2, a_3, a_4, a_5, a_6, T_0, b, p_0, \mathbf{y})$

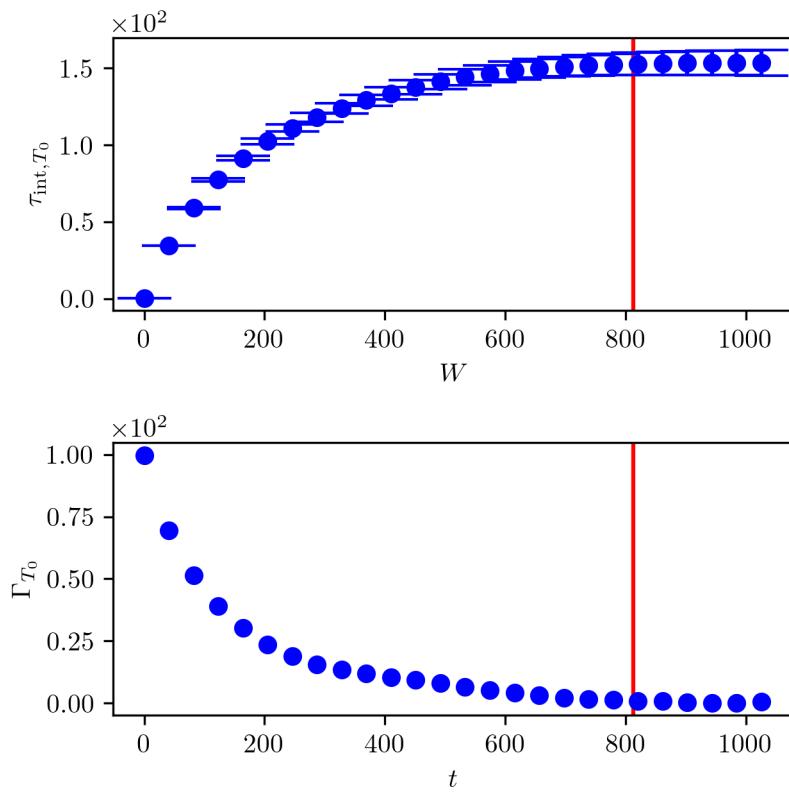


Figure B.12: IACT and autocorrelation function for samples $h_5 \sim \pi(\cdot | h_1, h_2, h_3, h_4, h_6, a_0, a_1, a_2, a_3, a_4, a_5, a_6, T_0, b, p_0, y)$

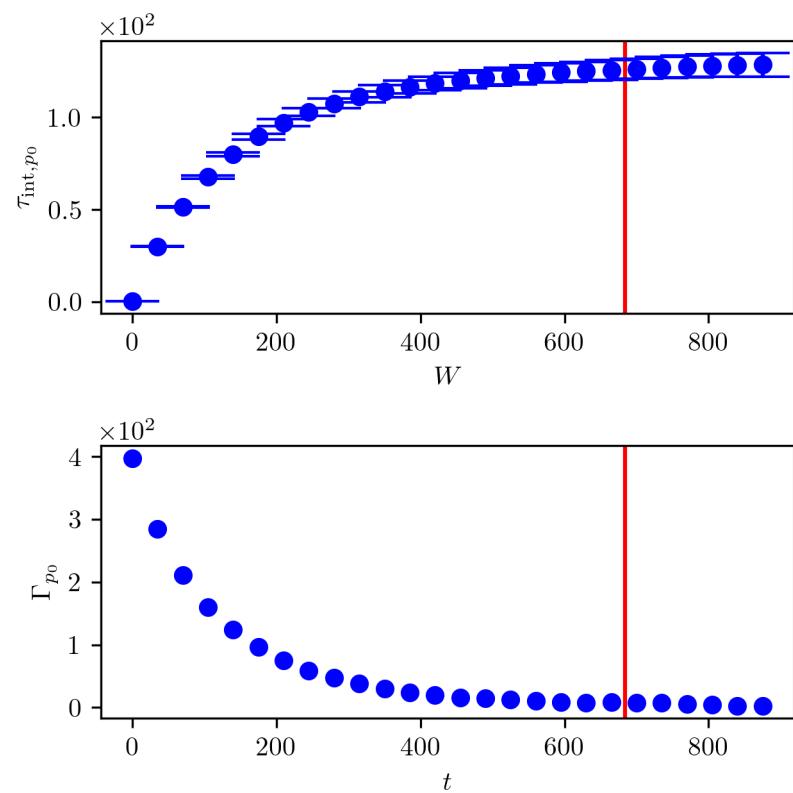


Figure B.13: IACT and autocorrelation function for samples $h_6 \sim \pi(\cdot | h_1, h_2, h_3, h_4, h_5, a_0, a_1, a_2, a_3, a_4, a_5, a_6, T_0, b, p_0, \mathbf{y})$

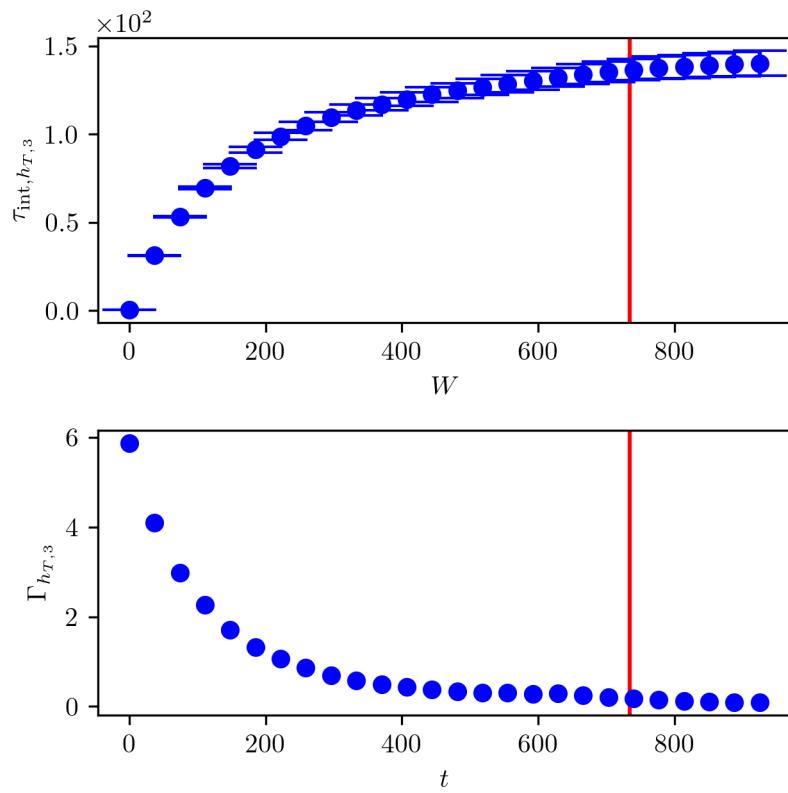


Figure B.14: IACT and autocorrelation function for samples $a_0 \sim \pi(\cdot | h_1, h_2, h_3, h_4, h_5, h_6, a_1, a_2, a_3, a_4, a_5, a_6, T_0, b, p_0, y)$

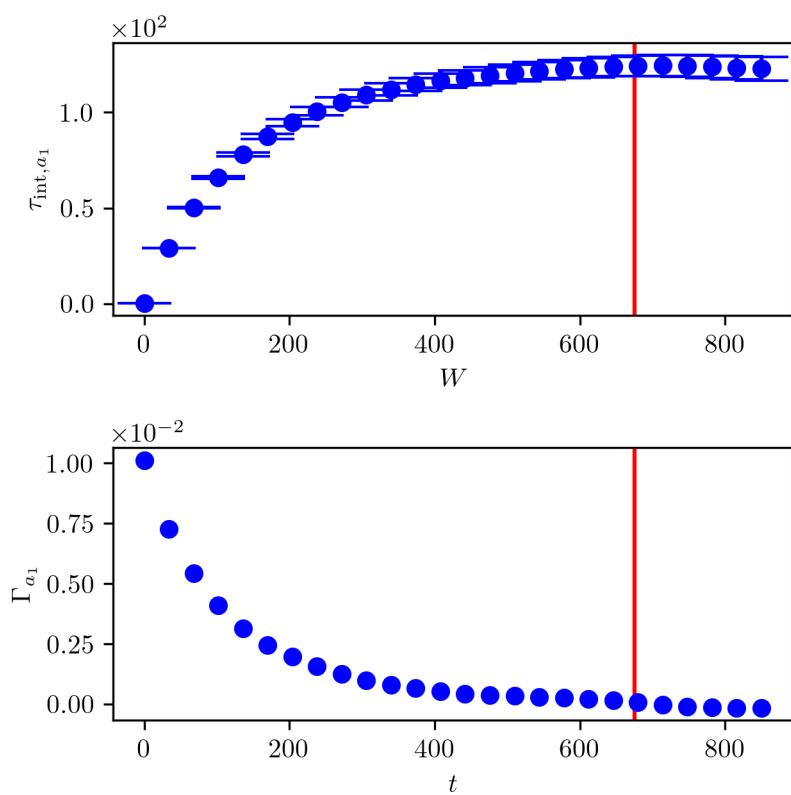


Figure B.15: IACT and autocorrelation function for samples $a_1 \sim \pi(\cdot | h_1, h_2, h_3, h_4, h_5, h_6, a_0, a_2, a_3, a_4, a_5, a_6, T_0, b, p_0, \mathbf{y})$

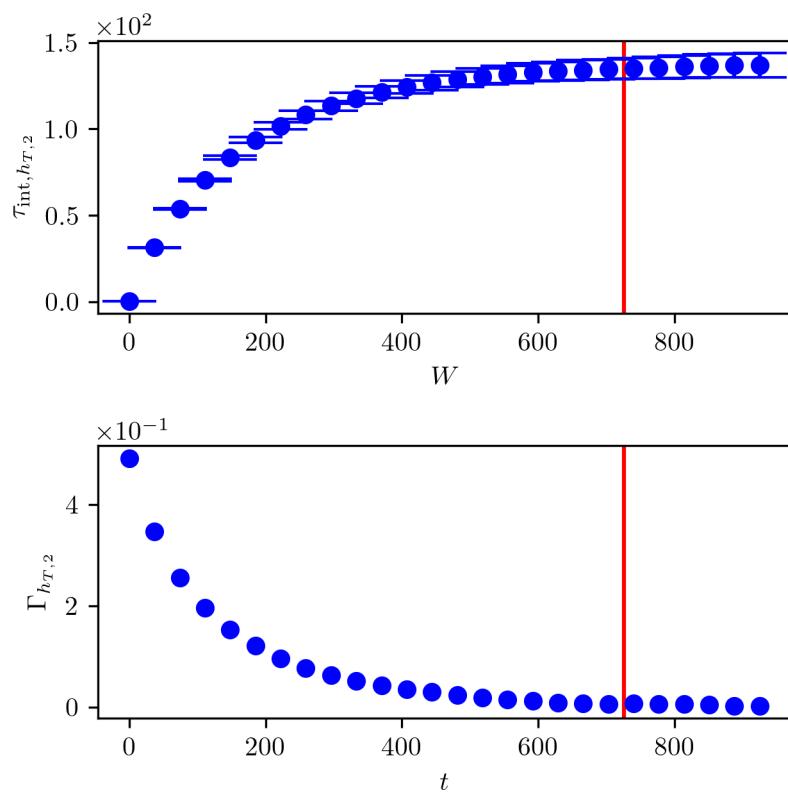


Figure B.16: IACT and autocorrelation function for samples $a_2 \sim \pi(\cdot | h_1, h_2, h_3, h_4, h_5, h_6, a_0, a_1, a_3, a_4, a_5, a_6, T_0, b, p_0, \mathbf{y})$

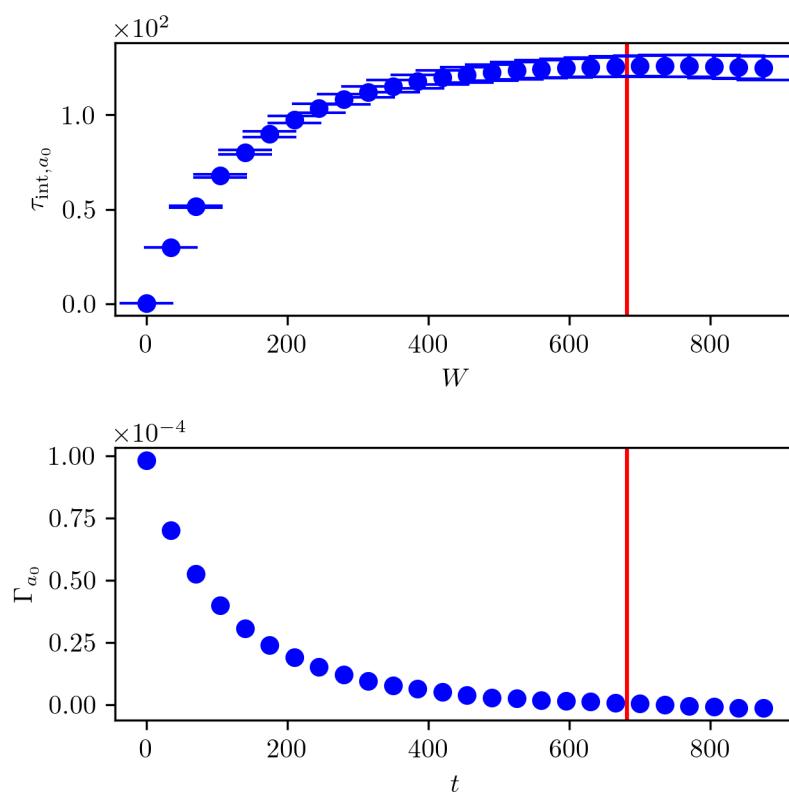


Figure B.17: IACT and autocorrelation function for samples $a_3 \sim \pi(\cdot | h_1, h_2, h_3, h_4, h_5, h_6, a_0, a_1, a_2, a_4, a_5, a_6, T_0, b, p_0, \mathbf{y})$

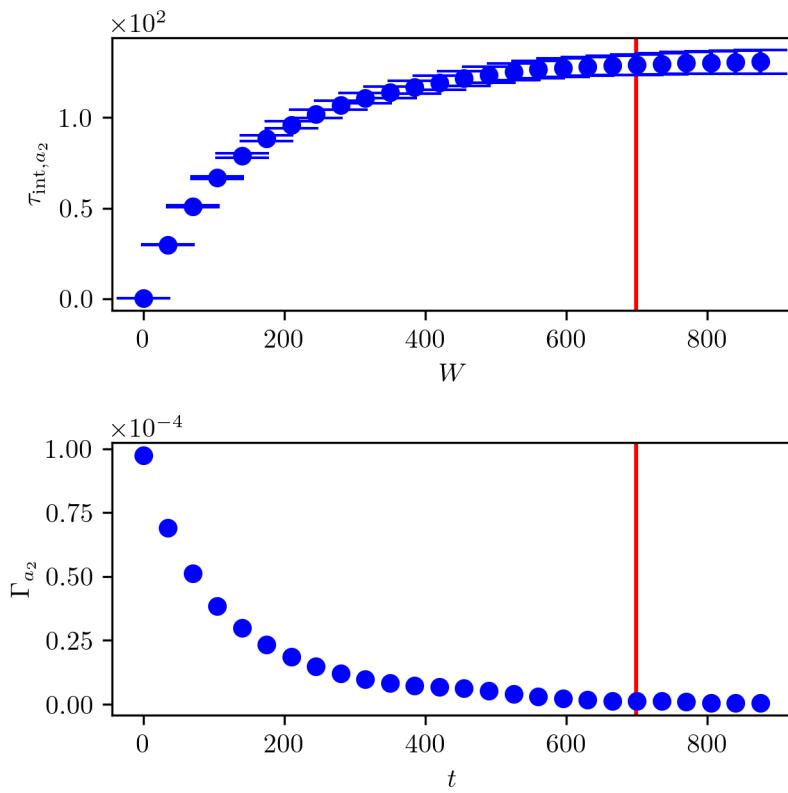


Figure B.18: IACT and autocorrelation function for samples $a_4 \sim \pi(\cdot | h_1, h_2, h_3, h_4, h_5, h_6, a_0, a_1, a_2, a_4, a_5, a_6, T_0, b, p_0, \mathbf{y})$

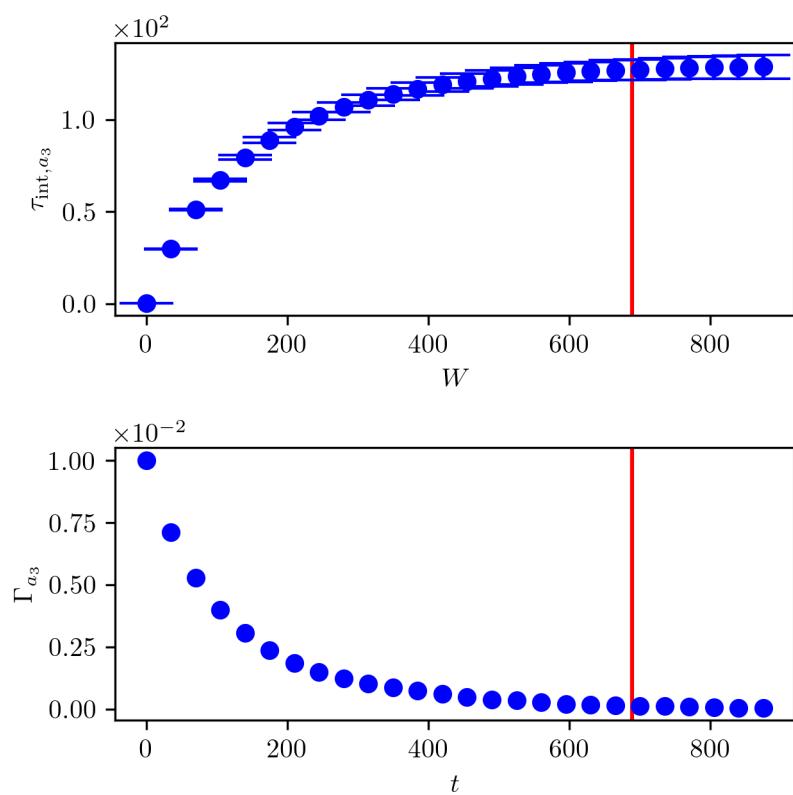


Figure B.19: IACT and autocorrelation function for samples $a_5 \sim \pi(\cdot | h_1, h_2, h_3, h_4, h_5, h_6, a_0, a_1, a_2, a_3, a_4, a_5, T_0, b, p_0, \mathbf{y})$

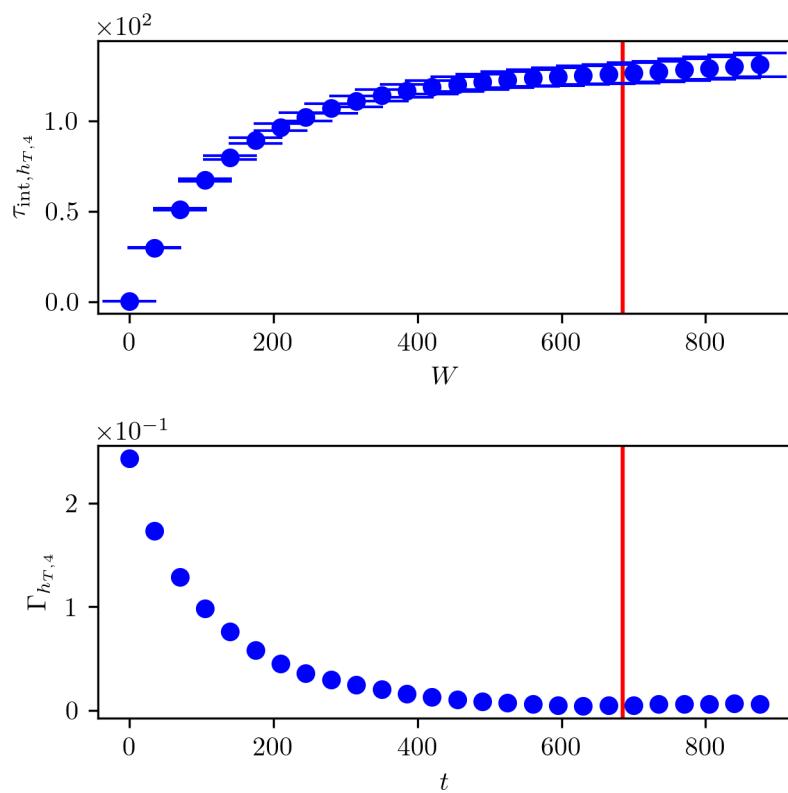


Figure B.20: IACT and autocorrelation function for samples $a_6 \sim \pi(\cdot | h_1, h_2, h_3, h_4, h_5, h_6, a_0, a_1, a_2, a_3, a_4, a_5, T_0, b, p_0, \mathbf{y})$

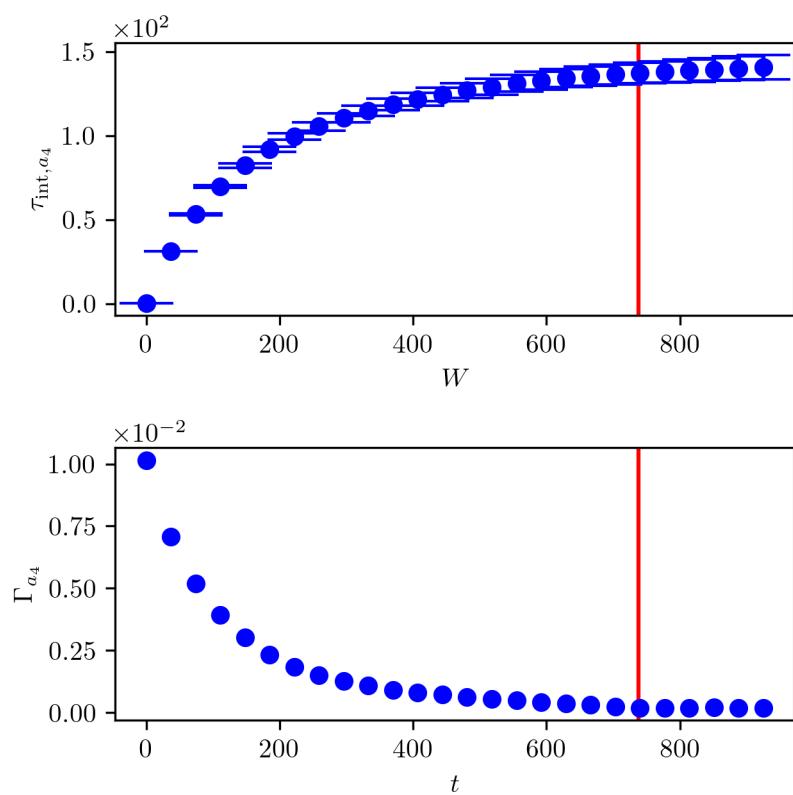


Figure B.21: IACT and autocorrelation function for samples $T_0 \sim \pi(\cdot | h_1, h_2, h_3, h_4, h_5, h_6, a_0, a_1, a_2, a_3, a_4, a_5, a_6, b, p_0, \mathbf{y})$

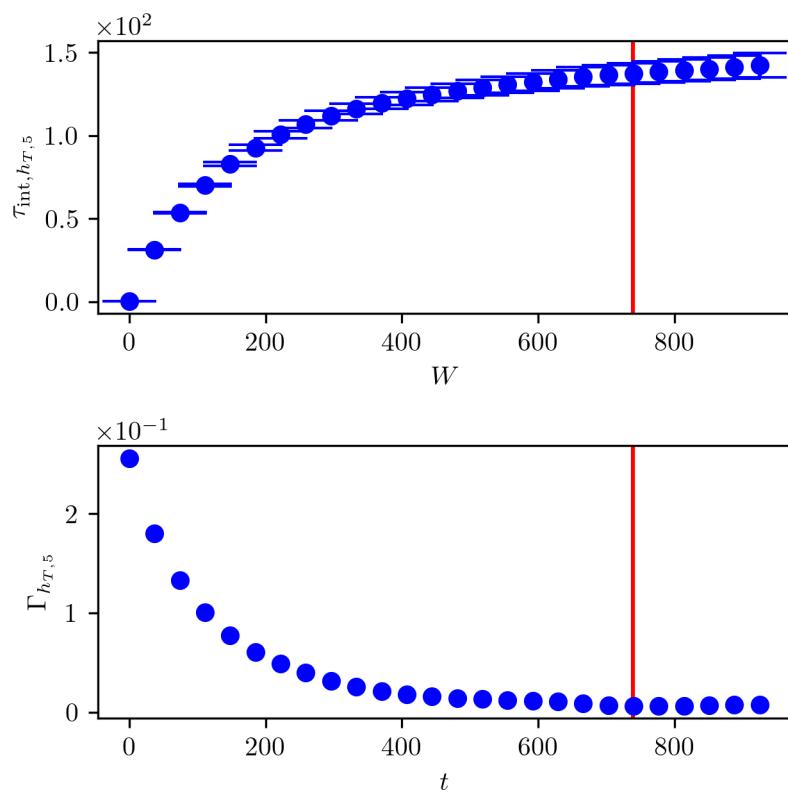


Figure B.22: IACT and autocorrelation function for samples $b \sim \pi(\cdot | h_1, h_2, h_3, h_4, h_5, h_6, a_0, a_1, a_2, a_3, a_4, a_5, a_6, T_0, p_0, \mathbf{y})$

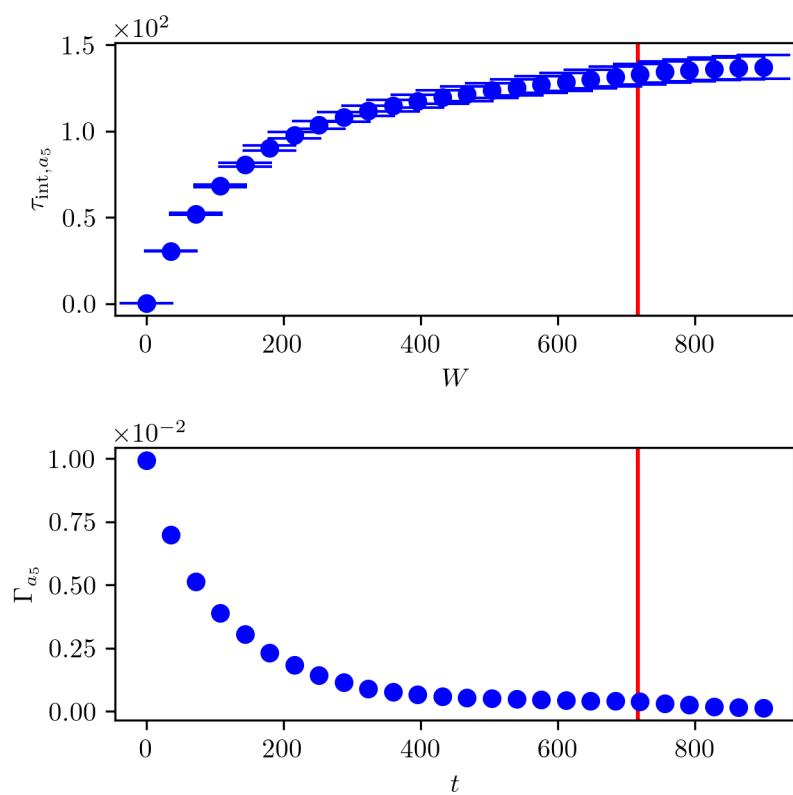


Figure B.23: IACT and autocorrelation function for samples $p_0 \sim \pi(\cdot | h_1, h_2, h_4, h_5, h_6, a_0, a_1, a_2, a_3, a_4, a_5, a_6, T_0, b, \mathbf{y})$

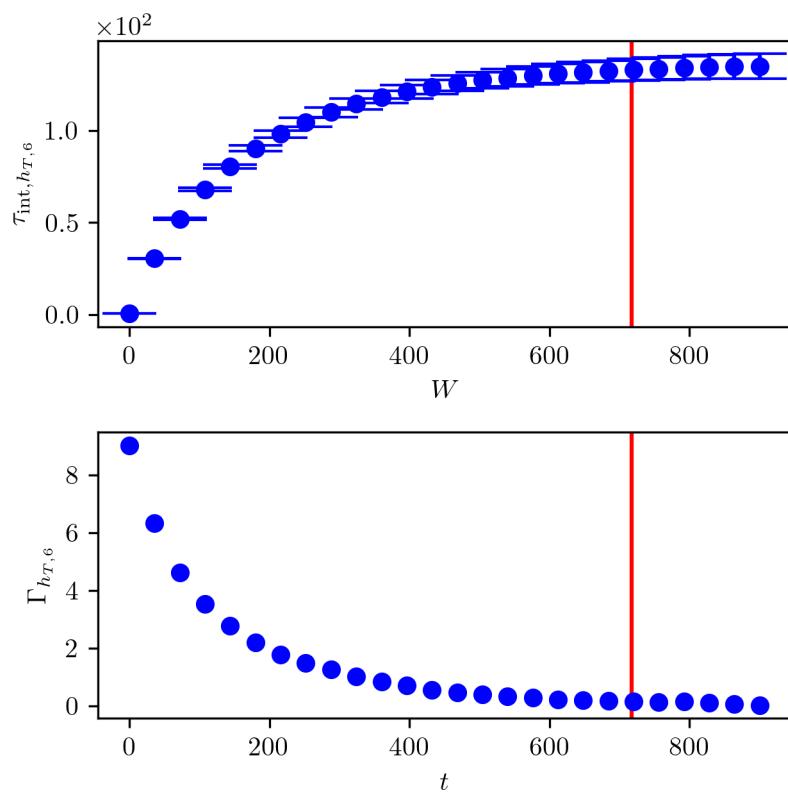


Figure B.24: IACT and autocorrelation function for samples $p_0 \sim \pi(\cdot | h_1, h_2, h_4, h_5, h_6, a_0, a_1, a_2, a_3, a_4, a_5, a_6, T_0, b, \mathbf{y})$

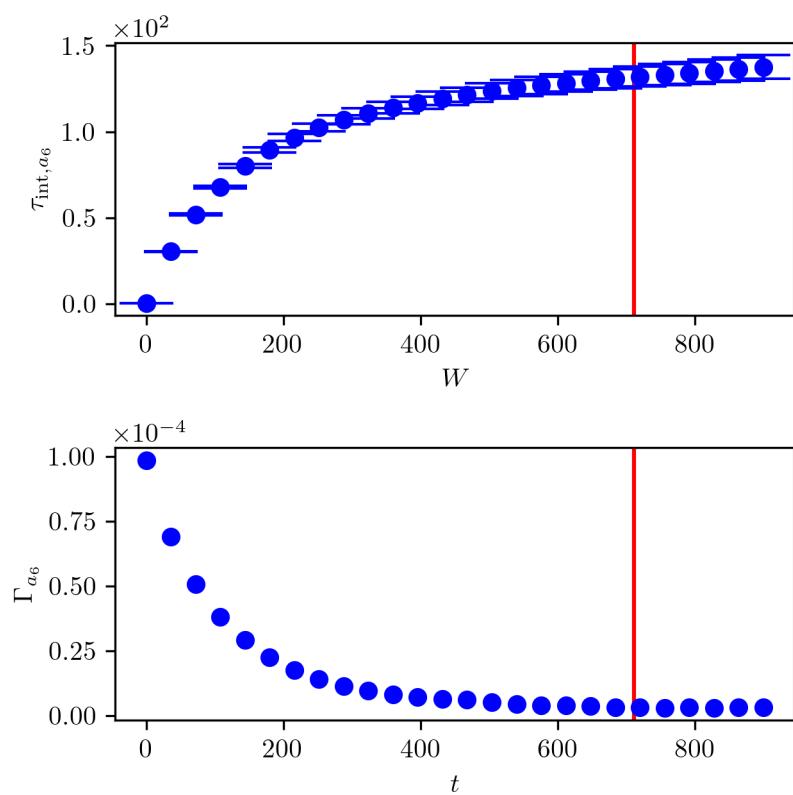


Figure B.25: IACT and autocorrelation function for samples $p_0 \sim \pi(\cdot | h_1, h_2, h_4, h_5, h_6, a_0, a_1, a_2, a_3, a_4, a_5, a_6, T_0, b, \mathbf{y})$