

## Astrophysics PST (2) - Solutions

1. The mass of the star is  $M = 5$  solar masses  $= 1 \times 10^{31}$  kg  
and its luminosity  $L = 600$  solar luminosity  $= 2.28 \times 10^{29}$  J s<sup>-1</sup>

Fraction of mass liberated per H-burning reaction =  
Mass deficit/mass of 4 protons  $= 0.0286/4.0312 = 0.0071$

The total energy that the star will be able to radiate is

$$\begin{aligned} E_{total} &= 0.0071 \times 0.1 \times M \times c^2 \\ &= 0.0071 \times 0.1 \times (1 \times 10^{31}) \times (9 \times 10^{16}) \text{ Joule} \\ &= 6.39 \times 10^{44} \text{ Joule} \end{aligned}$$

and it will radiate for

$$\frac{E_{total}}{L} = \frac{6.39 \times 10^{44}}{2.28 \times 10^{29}} \text{ s} = 2.8 \times 10^{15} \text{ s} = 8.9 \times 10^7 \text{ yr}$$

[ Btw, check understanding of what is meant by hydrostatic equilibrium!

I.e. a balance between the force of gravity inward and the pressure of hot gases pushing outward.

A balance or 'equilibrium' must be attained in order for a star to have a stable size. ]

2.

a) According to Hubble classification scheme:

*Sbc galaxy* -- galaxy has a small nucleus, with a bar-like structure  
through it. The spiral arms emerge from the ends of the bar and are loosely wound.

*Sa galaxy* -- galaxy has a relatively large nucleus, plus tightly wound spiral arms (no central bar).

To shift Ly $\alpha$  from visible (rest wavelength 121.6 nm) to visible ( $> 370$  nm)

$$\text{Redshift } z = \Delta\lambda/\lambda = (370.0 - 121.6)/121.6 = 2.04$$

For redshifts 2.04 or greater the Ly $\alpha$  line will be shifted to wavelengths longer than 370 nm.

Quasar 3C 273 distance:

$$\text{Hubble law, distance } d = v / H_0 = c z / H_0$$

$$= 3 \times 10^5 \times 0.16 / 75 = 640 \text{ Mpc}$$

3.

$$\begin{aligned}
 f_{\text{Andromeda}}/f_{\text{quasar}} &= (L_{\text{Andromeda}}/L_{\text{quasar}})(d_{\text{quasar}}^2/d_{\text{Andromeda}}^2) \\
 d_{\text{quasar}}^2/d_{\text{Andromeda}}^2 &= (f_{\text{Andromeda}}/f_{\text{quasar}})(L_{\text{quasar}}/L_{\text{Andromeda}}) \\
 d_{\text{quasar}}/d_{\text{Andromeda}} &= \sqrt{(f_{\text{Andromeda}}/f_{\text{quasar}})(L_{\text{quasar}}/L_{\text{Andromeda}})} \\
 d_{\text{quasar}}/d_{\text{Andromeda}} &= \sqrt{(10^4)(10^3 L_{\text{MW}}/3 L_{\text{MW}})} \\
 d_{\text{quasar}} &= \sqrt{10^7/3} \times 0.7 \text{ Mpc} = 1826 \times 0.7 \text{ Mpc} = 1.28 \times 10^3 \text{ Mpc}
 \end{aligned}$$

4.

$$v_{\text{rot}}^2 = \frac{GM}{r}$$

$$v_{\text{rot}}^2 = \frac{6.7 \times 10^{-11} \times 1.2 \times 10^{11} \times 2 \times 10^{30}}{8.5 \times 10^3 \times 3.1 \times 10^{16}}$$

$$v_{\text{rot}} = 247 \text{ km s}^{-1}.$$

$$\text{For circular orbit, } P(\text{orb}) = 2\pi r/v_{\text{rot}}$$

$$= \frac{2\pi \times 8.5 \times 10^3 \times 3.1 \times 10^{16}}{2.47 \times 10^5} \times \frac{1}{3.16 \times 10^7} = 2.1 \times 10^8 \text{ yr}$$

5. As the question sheet states this is a bit of a mis-use of the Hubble law!  
But it gives a useful feel for how fast the expansion is!

$$\text{Hubble law } v = H_0 \cdot d$$

$$\begin{aligned}
 \text{For given value of } H_0 &= 75 \text{ km/s/Mpc} \\
 &= 75 \times 10^3 / 3.1 \times 10^{16} \times 10^6 \\
 &= 2.419 \times 10^{-18} \text{ sec}^{-1}
 \end{aligned}$$

$$\begin{aligned}
 \text{Now for distance (Atlantic)} &= 6000 \text{ km} \\
 v &= 2.419 \times 10^{-18} \times 6000 = 1.451 \times 10^{-14} \text{ km s}^{-1} \\
 &= 1.451 \times 10^{-9} \text{ cm s}^{-1}
 \end{aligned}$$

$$\text{Distance (cosmic expansion) in a year} = 1.451 \times 10^{-9} \text{ cm s}^{-1} \times 3 \times 10^7 = 0.043 \text{ cm}$$

This compares with plate tectonic movement of continents of around 2 cm per year.